# Lifetime Difference in D<sup>0</sup> – D<sup>0</sup> Mixing Within R-parity Violating Supersymmetry

Gagik Yeghiyan

Wayne State University, Detroit, MI, USA

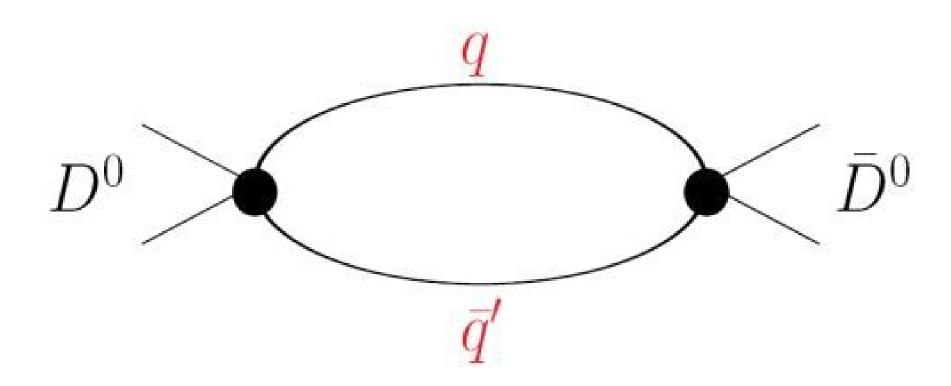
**DPF 2009** 

Based on

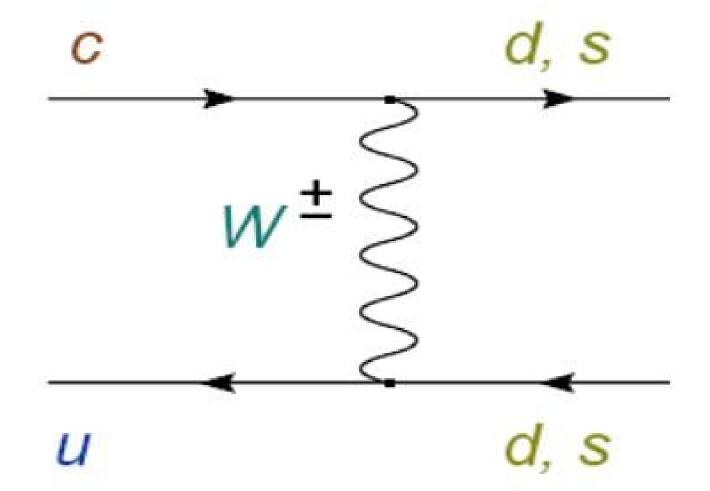
A. A. Petrov, G. K. Yeghiyan, Phys. Rev. D 77, 034018 (2008)

#### Outline of the Talk:

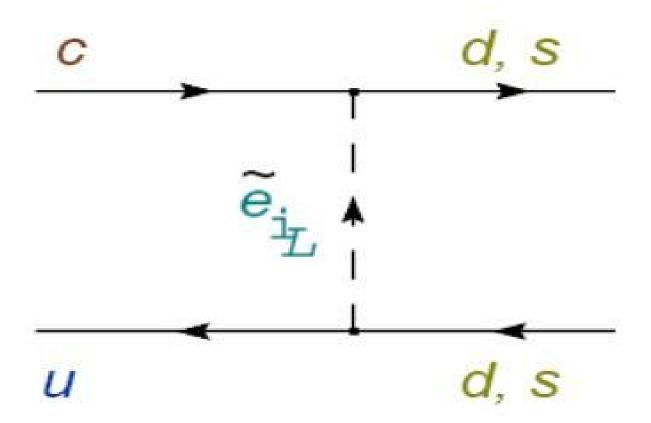
- 1. Issues of possible large NP contribution to  $y_D = \Delta \Gamma_D / (2\Gamma_D)$
- 2. y<sub>D</sub> within RPV SUSY models, subtleties of analysis
- 3. Numerical results
- 4. Conclusions



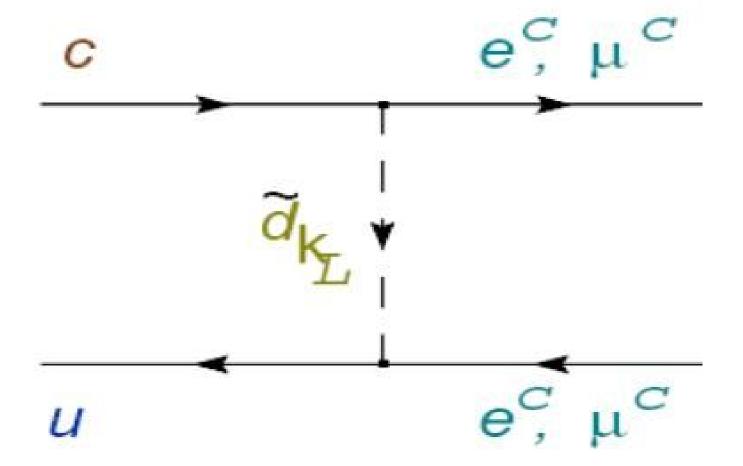
If  $|n\rangle$  is a charmless state,  $\Delta C=1 \ D^0 \rightarrow n \ decay amplitude is:$  $A[D^0 \rightarrow n] = A_n^{(SM)} + A_n^{(NP)}$ 



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#### Let

$$A[D^0 \rightarrow n] = A_n^{(SM)} + A_n^{(NP)}$$
, | n > is a charmless state

# Then for $y_D = \Delta \Gamma_D / (2\Gamma_D)$

$$y_D = \sum_n \frac{\rho_n}{\Gamma_D} A_n^{(SM)} \bar{A}_n^{(SM)} + 2 \sum_n \frac{\rho_n}{\Gamma_D} A_n^{(NP)} \bar{A}_n^{(SM)} + \sum_n \frac{\rho_n}{\Gamma_D} A_n^{(NP)} \bar{A}_n^{(NP)}$$

An (NP) should be smaller than theoretical and experimental uncertainties in predictions for D-meson decays.

Is y<sub>D</sub> ≈ y<sub>SM</sub>? y<sub>NP</sub> may be essential as well! Golowich, Pakvasa, Petrov, PRL 98, 181801 (2007)

$$y_D = \sum_n \frac{\rho_n}{\Gamma_{\rm D}} A_n^{\rm (SM)} \bar{A}_n^{\rm (SM)} + 2\sum_n \frac{\rho_n}{\Gamma_{\rm D}} A_n^{\rm (NP)} \bar{A}_n^{\rm (SM)} + \sum_n \frac{\rho_n}{\Gamma_{\rm D}} A_n^{\rm (NP)} \bar{A}_n^{\rm (NP)}$$

The SM contribution vanishes in the exact flavor SU(3) limit.

Short distance: LO  $y_{SM} \sim m_s^6/m_c^6$   $y_{SM} \sim 10^{-8}$  NLO  $y_{SM} \sim m_s^4/m_c^4$   $y_{SM} \sim 10^{-6}$ 

Long distance:  $\sim m_s^2/m_c^2$ ,  $y_{SM} \sim 10^{-4} \div 10^{-2}$ 

- may explain  $y_D^{exp} = (7.3 \pm 1.8) \cdot 10^{-3}$ 

### Key point:

 If non-vanishing in the exact flavor SU(3) limit

 Or if vanishing, but being (like long-distance SM contribution) ~m<sub>s</sub><sup>2</sup>/m<sub>c</sub><sup>2</sup> when SU(3) is broken,

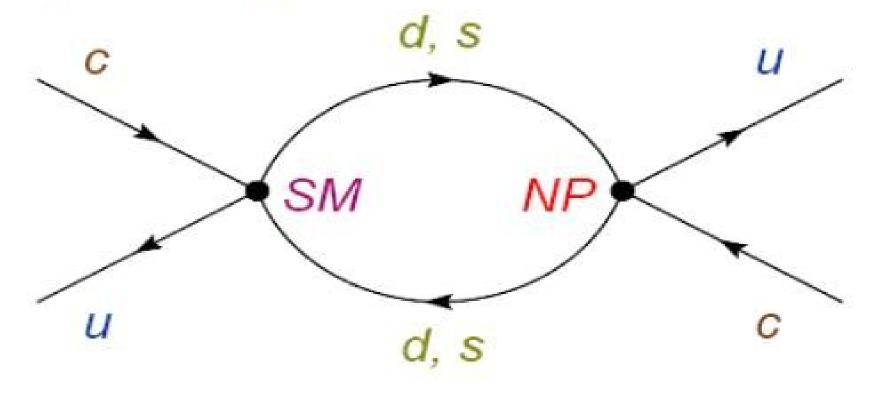
NP contribution to y<sub>D</sub> may be essential or even dominant (depending on M<sub>NP</sub> and NP couplings)

$$y_D = \sum_n \frac{\rho_n}{\Gamma_D} A_n^{(SM)} \bar{A}_n^{(SM)} + 2 \sum_n \frac{\rho_n}{\Gamma_D} A_n^{(NP)} \bar{A}_n^{(SM)} + \sum_n \frac{\rho_n}{\Gamma_D} A_n^{(NP)} \bar{A}_n^{(NP)}$$

# The second term (interference of the SM and NP \( \Delta C=1 \) transitions) -

within the most popular SM extensions

 $y_{SM, NP} \le 10^{-4} - fails to explain$  $y_{D}^{exp} = (7.3 \pm 1.8) \cdot 10^{-3}$ 



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Golowich, Pakvasa, Pertov, PRL 98, 181801 (2007)

S. L. Chen, X. G. He, A. Hovhannissyan, H.S. Tsai JHEP 09, 044 (2007)

Ch. H. Chen, Ch. Q. Geng, S. H. Nam, PRL 99, 019101 (2007)

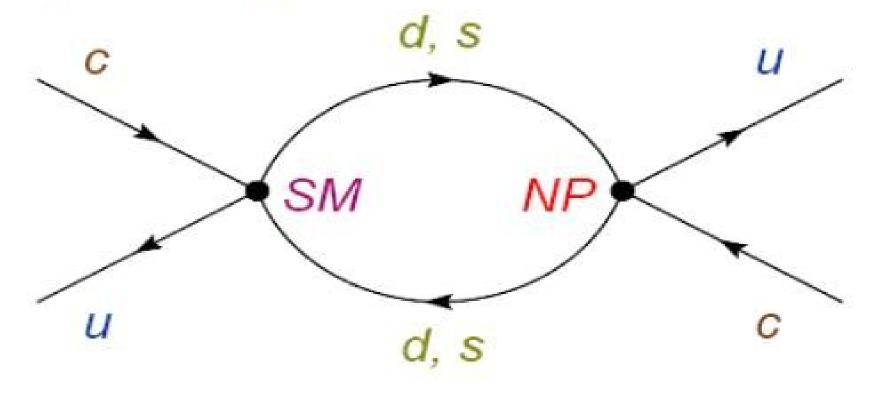
G. K. Yeghiyan, PRD 76, 117701 (2007)

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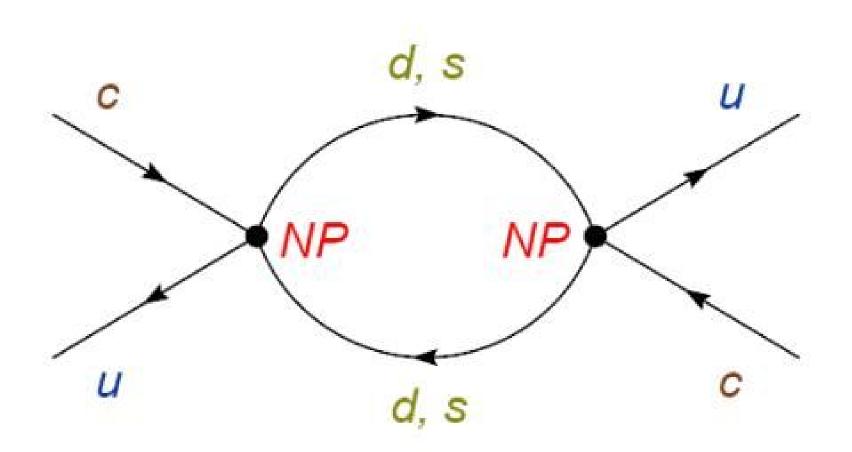
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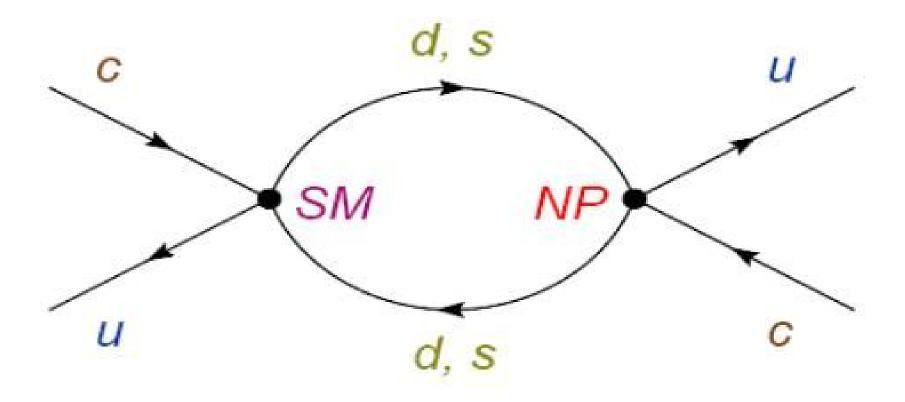
## Exact flavor SU(3) limit:

y<sub>SM</sub> = 0

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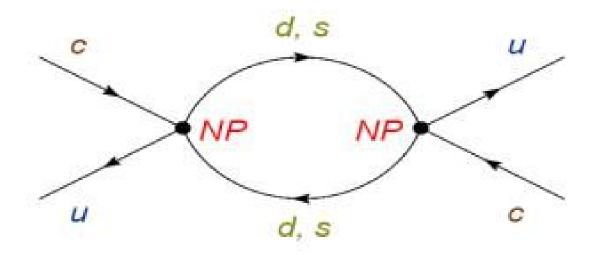
- y<sub>SM</sub> = 0
- In many popular models (also in RPV SUSY)
   y<sub>SM. NP</sub> = 0



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Flavor SU(3) is broken: use power counting: Powers of m<sub>s</sub>/m<sub>c</sub> vs. M<sub>W</sub><sup>4</sup>/M<sub>NP</sub><sup>4</sup>

Approximate rule: M<sub>W</sub><sup>4</sup>/M<sub>NP</sub><sup>4</sup> > m<sub>s</sub><sup>2</sup>/m<sub>c</sub><sup>2</sup>

### Purpose of our work:

- Revisit the problem of the NP contribution to y<sub>D</sub> within RPV SUSY models
- 2. Provide constraints on RPV couplings, insensitive or weakly sensitive to assumptions on R-conserving sector

## Purpose of our work:

- Revisit the problem of the NP contribution to y<sub>D</sub> within RPV SUSY models
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- In contrast to studies of  $x_D = \Delta m_D/\Gamma_D$  have strong predictive power (Golowich, Hewett, Pakvasa, Petrov, PRD 76, 095009 (2007)), if only one may neglect R-conserving sector contribution

For y<sub>D</sub>, R-conserving SUSY sector contributes at two-loop level, by dipenguin diagrams, and expected to be small

Revisit the problem within RPV SUSY: take into account the transformation of RPV couplings from the weak isospin basis to the quark mass eigenbasis

We considered a general low-energy SUSY scenario with no assumptions made on a SUSY breaking mechanism at unification scales.

## Superpotential:

$$W_{\mathbb{R}} = \sum_{i,j,k} \left[ \frac{1}{2} \lambda_{ijk} L_i L_j E_k^c + \lambda'_{ijk} L_i Q_j D_k^c + \frac{1}{2} \lambda''_{ijk} U_i^c D_j^c D_k^c \right]$$

# To avoid rapid proton decay, we put λ"=0. In the quark mass eigenbasis

$$\mathcal{L}_{R} = -\sum_{i,j,k} \widetilde{\lambda}'_{ijk} \left[ \widetilde{e}_{iL} \overline{d}_{kR} u_{jL} + \widetilde{u}_{jL} \overline{d}_{kR} e_{iL} + \widetilde{d}^*_{kR} \overline{e}^c_{iR} u_{jL} \right] +$$

$$+\sum_{i,j,k}\lambda'_{ijk}\left[\tilde{\nu}_{iL}\bar{d}_{kR}d_{jL}+\tilde{d}_{jL}\bar{d}_{kR}\nu_{iL}+\tilde{d}^*_{kR}\bar{\nu}^c_{iR}d_{jL}\right]+h.c.$$

where

$$\widetilde{\lambda}'_{irm} = V_{rn}^* \lambda'_{inm}$$

# Very often in the literature one neglects the difference between $\lambda'$ and $\widetilde{\lambda}'$ based on

$$V_{jn} = \delta_{jn} + O(\lambda)$$
 so  $\widetilde{\lambda}_{ijk} \approx \lambda'_{ijk} + O(\lambda)$ 

This is not always a good approximation. Example: if for a given i, j, k, n,  $\lambda'_{ijk} \neq 0$ ,  $\lambda'_{ink} \neq 0$  and all other  $\lambda'$  couplings vanish,

$$\widetilde{\lambda}_{ijk}'\widetilde{\lambda}_{ink}'^{\dagger} \approx \lambda_{ijk}'\lambda_{ink}'^{\dagger}$$
 if  $\lambda_{ijk}' \sim \lambda_{ink}'$ 

Otherwise, if  $\lambda'_{ijk} >> \lambda'_{imk}$ 

$$\widetilde{\lambda}_{ijk}^{\prime} \, \widetilde{\lambda}_{ink}^{\prime *} \, \approx \, \lambda_{ijk}^{\prime} \, \lambda_{ink}^{\prime *} \, + \, V_{nj}^{\phantom{\prime}} \, \big| \, \lambda_{ijk}^{\prime} \, \big|^2 \, \neq \, \lambda_{ijk}^{\prime} \, \lambda_{ink}^{\prime *}$$

#### Or even

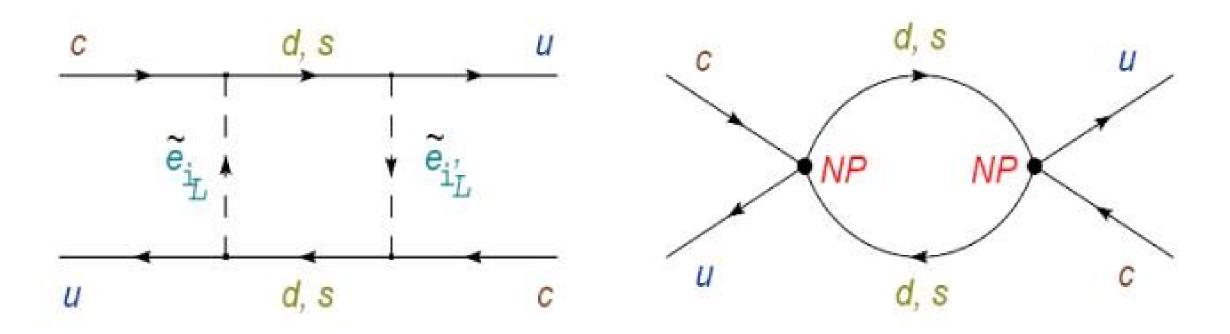
$$\widetilde{\lambda}_{ijk}^{\,\prime}\,\widetilde{\lambda}_{ink}^{\,\prime^{\star}} \;\approx\; V_{nj}^{\phantom{\dagger}}\,\left|\,\lambda_{ijk}^{\,\prime}\,\right|^2 >>\; \lambda_{ijk}^{\,\prime}\,\lambda_{ink}^{\,\prime^{\star}}$$

If one neglects the difference between  $\lambda'$  and  $\widetilde{\lambda}'$  one may misuse the existing constraints on RPV couplings.

S. L. Chen, X. G. He, A. Hovhannissyan, H.S. Tsai JHEP 09, 044 (2007)-RPV SUSY contribution to  $y_D$  is rather small – true if only there is no hierarchy in the bounds (and hence the values) of the relevant RPV couplings

More generally, if the hierarchy is allowed, then....

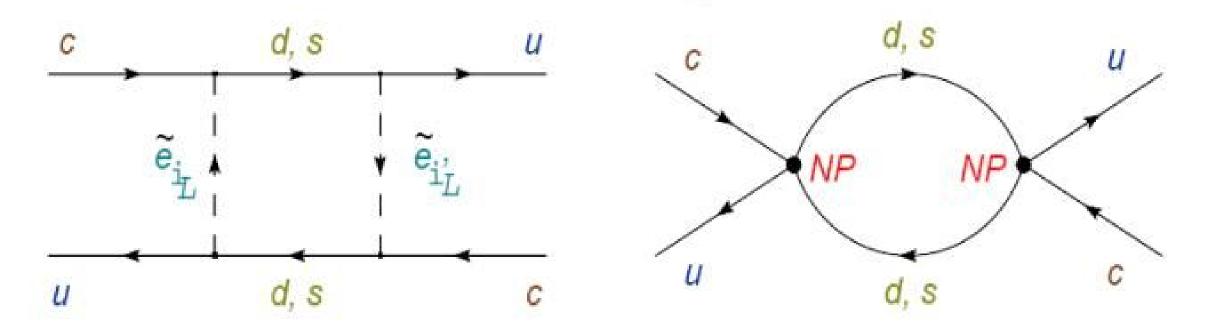
#### The dominant diagrams.



# Neglecting numerically subdominant terms,

$$y_{\tilde{\ell}\tilde{\ell}} \approx \frac{-m_c^2 f_D^2 B_D m_D}{288\pi\Gamma_D m_{\tilde{\ell}}^4} \left[ \frac{1}{2} + \frac{5}{8} \frac{\bar{B}_D^S}{B_D} \right] \left[ \lambda_{ss}^2 + \lambda_{dd}^2 \right]$$
$$\lambda_{ss} \equiv \sum_i \tilde{\lambda}_{i12}^{\prime*} \tilde{\lambda}_{i22}^{\prime} \gg \sum_i \lambda_{i12}^{\prime*} \lambda_{i22}^{\prime}$$
$$\lambda_{dd} \equiv \sum_i \tilde{\lambda}_{i11}^{\prime*} \tilde{\lambda}_{i21}^{\prime} \gg \sum_i \lambda_{i11}^{\prime*} \lambda_{i21}^{\prime}$$

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It is non-vanishing in the exact flavor SU(3) limit. Else, 
$$|\lambda_{ss}| \le 0.29$$
,  $|\lambda_{dd}| \le 0.29$  or  $|\lambda_{ss}| \le 0.0841$  and  $|\lambda_{dd}| \le 0.0841$ , M<sub>NP</sub>~100GeV, contribution of this type of diagrams is large.

#### Numerically

$$-0.12 \left(\frac{100 GeV}{m_{\tilde{\ell}}}\right)^4 \le y_{\tilde{\ell}\tilde{\ell}} < 0$$

#### Compare to

$$y_D^{exp} = (7.3 \pm 1.8) \cdot 10^{-3}$$

Radical approach: place severe constraint on  $\lambda_{ss}$  and  $\lambda_{dd}$  to make NP contribution to  $y_D$  negligible

Moderate approach: demand a large positive contribution from the SM (to have a destructive interference of two contributions).

A. Falk et al., PRD 65, 054034 (2002): due do the longdistance effects, y<sub>SM</sub> may be up to ~1%.

Thus, RPV SUSY contribution to y<sub>D</sub> should be ~1% or less as well.

**Impose** 

$$-0.01 \le y_{new} \approx y_{\tilde{\ell}\tilde{\ell}}$$

then

either

$$m_{\tilde{\ell}} > 185 {\rm GeV}$$

or if

$$m_{\tilde{\ell}} \leq 185 \text{GeV}$$

then

$$|\lambda_{ss}| \le 0.082 \left(\frac{m_{\tilde{\ell}}}{100 GeV}\right)^2$$

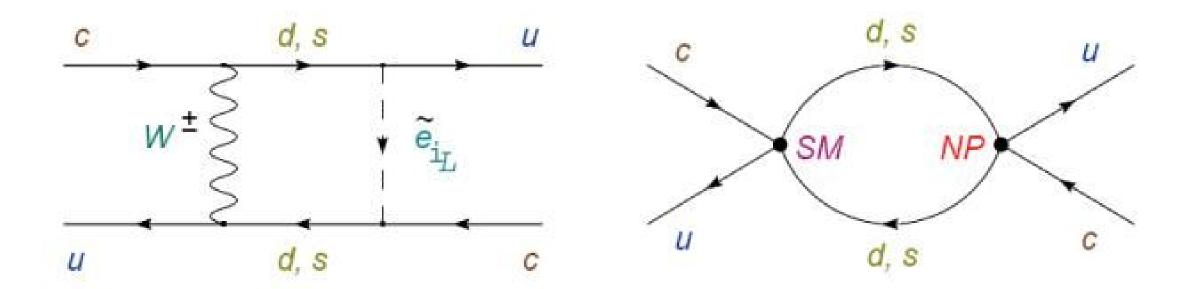
$$|\lambda_{dd}| \le 0.082 \left(\frac{m_{\tilde{\ell}}}{100 GeV}\right)^2$$

Our bounds are insensitive or weakly sensitive on the assumptions on R-conserving SUSY sector

### **Summary and conclusions**

- Lifetime difference in D<sup>0</sup> D<sup>0</sup> mixing has been revisited within RPV SUSY models.
- NP contribution to  $y_D$  may be large in absolute value: it may exceed  $y_D^{exp} = (7.3 \pm 1.8) \cdot 10^{-3}$  by an order of magnitude.
- When being large y<sub>NP</sub> is negative in sign. The existing experimental data may be the result of destructive interference of the SM and RPV SUSY contributions.
- We derive new bounds on the RPV coupling pair products and/or supersymmetric particle masses.
   These bounds are insensitive or weakly sensitive to assumptions on the R-conserving sector of the theory.

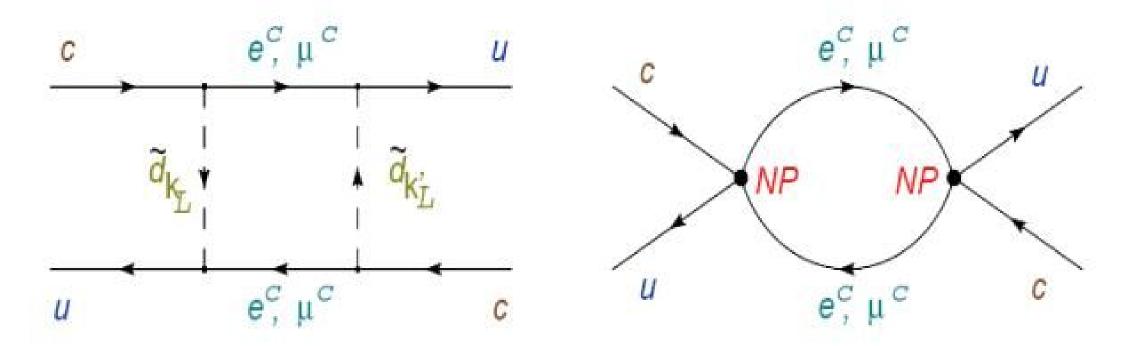
#### Supplementary slide N1



$$y_{SM,NP} = \frac{-G_F}{\sqrt{2}} \frac{f_D^2 B_D m_D}{6\pi \Gamma_D} \left(\frac{m_c^2}{m_{\tilde{\ell}}^2}\right) \left[C_1(m_c) + C_2(m_c)\right] \left[\lambda_{sd} \sqrt{x_s x_d} + \lambda \left(\lambda_{ss} x_s - \lambda_{dd} x_d\right) - \lambda^2 \lambda_{ds} \sqrt{x_s x_d}\right]$$

Contribution of this type of diagrams vanishes in the exact flavor SU(3) symmetry limit. As flavor SU(3) is broken  $y_{SM,NP}$  is suppressed as  $x_s = m_s^2/m_c^2$  or  $x_d = m_d^2/m_c^2$ . It is not hard to show that contribution of this type of diagrams is rather small.

#### Supplementary Slide N2



$$y_{\tilde{q}\tilde{q}} = \frac{m_c^2 \ f_D^2 B_D m_D}{288 \pi \Gamma_D \ m_{\tilde{q}}^4} \left[ \ \frac{5 \ \bar{B}_D^S}{8 \ B_D} - 1 \ \right] \left[ \ \lambda_{\mu\mu}^2 + \lambda_{ee}^2 + 2 \ \lambda_{\mu e} \lambda_{e\mu} \ \right]$$

Contribution of this type of diagrams does not vanish in the exact flavor SU(3) limit, however it is numerically subdominant:

$$y_{\tilde{q}\tilde{q}} \le 5.34 \cdot 10^{-5}$$

because of the stringent bounds on the RPV coupling products

#### Supplementary slide N3

$$\lambda_{ss} \equiv \sum_{i} \tilde{\lambda}_{i12}^{\prime*} \tilde{\lambda}_{i22}^{\prime} = \sum_{i} \lambda_{i12}^{\prime*} \lambda_{i22}^{\prime} + \lambda \left[ \sum_{i} |\lambda_{i22}^{\prime}|^{2} - \sum_{i} |\lambda_{i12}^{\prime}|^{2} \right] + A\lambda^{2} \sum_{i} \lambda_{i12}^{\prime*} \lambda_{i32}^{\prime} + A\lambda^{3} (1 + \rho - i\eta) \sum_{i} \lambda_{i32}^{\prime*} \lambda_{i22}^{\prime} + A^{2} \lambda^{5} (\rho - i\eta) \sum_{i} |\lambda_{i32}^{\prime}|^{2}$$

$$|\lambda_{i12}^{\prime*} \ \lambda_{i22}^{\prime}| \leq 6.3 \cdot 10^{-5} \left(\frac{m_{\tilde{q}}}{300 GeV}\right)^{2} \qquad |\sum_{i} \lambda_{i12}^{\prime*} \ \lambda_{i22}^{\prime}| \leq 2.7 \times 10^{-3} \left(\frac{m_{\tilde{\ell}}}{100 GeV}\right)^{2}$$

For 
$$m_{\tilde{q}} \ge 300 \text{GeV}$$
,  $|\lambda_{322}| \le 1.12$   $|\lambda'_{312}| \le 0.33 (m_{\tilde{q}}/300 \text{GeV})$ ,

$$|\lambda_{312}| \le 1.12$$

$$-0.025 \left(\frac{m_{\tilde{q}}}{300 GeV}\right)^2 \le \lambda_{ss} \le 0.29$$
, if  $m_{\tilde{q}} \le 1 TeV$ ,  $-0.29 \le \lambda_{ss} \le 0.29$ , if  $m_{\tilde{q}} \ge 1 TeV$