

***Gravitational radiation reaction  
in compact binary systems:  
Quadrupole-monopole and  
magnetic dipole-magnetic dipole  
interaction***

***Zoltán Keresztes***

***László Árpád Gergely and Mátyás Vasúth***

# Dynamics:

## Lagrangian:

$$L = L_N + L_{QM} + L_{DD},$$

$$L_N = \frac{\mu v^2}{2} + \frac{Gm\mu}{r},$$

$$L_{QM} = \frac{G\mu m^3}{2r^5} \sum_{i=1}^2 p_i [3(\hat{S}_i \cdot \mathbf{r})^2 - r^2],$$

$$L_{DD} = \frac{1}{r^3} [3(\mathbf{n} \cdot \mathbf{d}_1)(\mathbf{n} \cdot \mathbf{d}_2) - \mathbf{d}_1 \cdot \mathbf{d}_2],$$

$p_i$  quadrupole scalar (axial symmetry)

$d_i$  magnetic dipole vector

## Identities:

$$v^2 = \frac{2E_N}{\mu} + \frac{2Gm}{r},$$

$$\dot{r}^2 = \frac{2E_N}{\mu} + \frac{2Gm}{r} - \frac{L_N^2}{\mu^2 r^2}.$$

## Solution of the radial equation

Definition of turning points:

$$\begin{aligned} r_{\min} &= r(0), & \dot{r}^2(0) &= 0, \\ r_{\max} &= r(\pi), & \dot{r}^2(\pi) &= 0. \end{aligned}$$

Definition of  $\chi$ :

$$\frac{2}{r} = \left( \frac{1}{r_{\min}} + \frac{1}{r_{\max}} \right) + \left( \frac{1}{r_{\min}} - \frac{1}{r_{\max}} \right) \cos \chi,$$

$$r = \frac{\bar{L}^2}{\mu(Gm\mu + \bar{A} \cos \chi)} + \frac{G\mu^2 m^3}{4\bar{A}\bar{L}^2(Gm\mu + \bar{A} \cos \chi)^2} \sum_{i=1}^2 p_i \Lambda^i + \frac{\mu d_1 d_2 \Lambda}{2\bar{A}\bar{L}^2(Gm\mu + \bar{A} \cos \chi)^2},$$

where:

$$\Lambda^i = \bar{A}[\bar{A}^2(\alpha_0^i + 4\beta_0^i) + (Gm\mu)^2(3\alpha_0^i + 10\beta_0^i)] + Gm\mu[\bar{A}^2(3\alpha_0^i + 11\beta_0^i) + (Gm\mu)^2(\alpha_0^i + 3\beta_0^i)] \cos \chi,$$

$$\Lambda = \bar{A}[(3G^2 m^2 \mu^2 + \bar{A}^2)A_0 + (G^2 m^2 \mu^2 + \bar{A}^2)B_0] + Gm\mu[(G^2 m^2 \mu^2 + 3\bar{A}^2)A_0 + 2\bar{A}^2 B_0] \cos \chi,$$

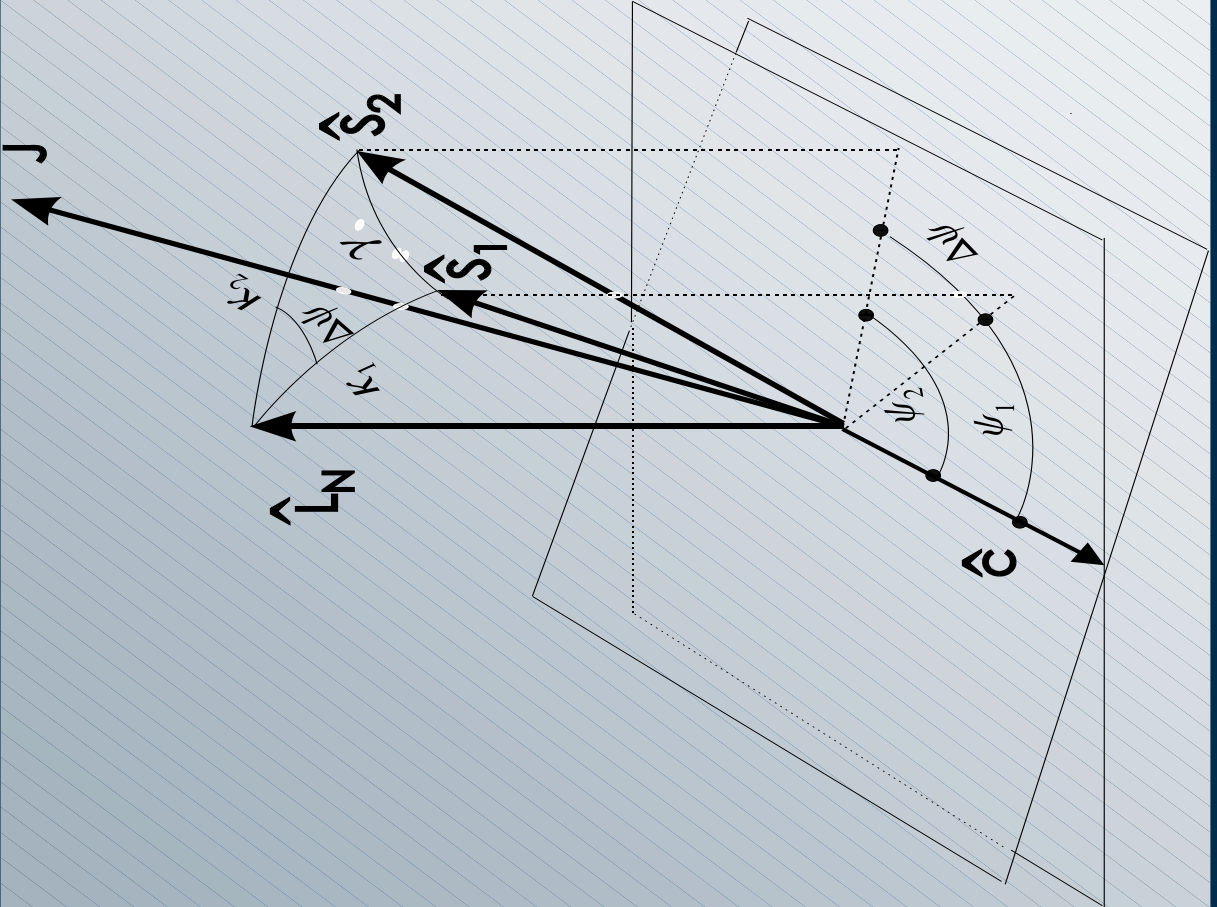
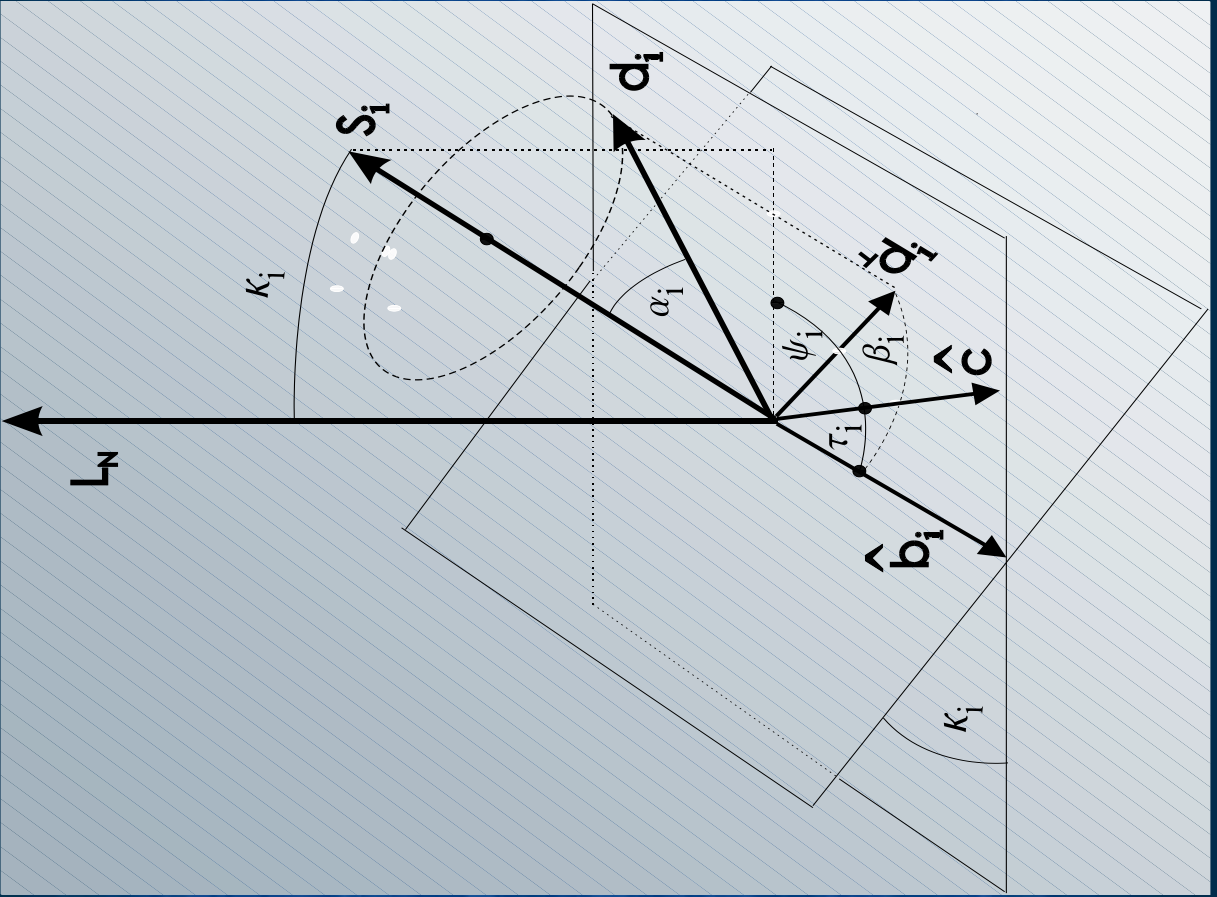
Definition of  $\xi$ :

$$2r = (r_{\max} + r_{\min}) - (r_{\max} - r_{\min}) \cos \xi,$$

$$r = \frac{Gm\mu - \bar{A} \cos \xi}{-2E} + \frac{G\mu^2 m^3}{4\bar{A}\bar{L}^2} \sum_{i=1}^2 p_i \Xi^i + \frac{\mu d_1 d_2}{2\bar{A}\bar{L}^2} \langle \bar{A}(A_0 + B_0) + Gm\mu A_0 \cos \xi \rangle,$$

where:

$$\Xi^i = \bar{A}(\alpha_0^i + 4\beta_0^i) + Gm\mu(\alpha_0^i + 3\beta_0^i) \cos \xi.$$





## Averaged radiative losses

$$\left\langle \frac{dE}{dt} \right\rangle = \left\langle \frac{dE}{dt} \right\rangle_N + \left\langle \frac{dE}{dt} \right\rangle_{QM} + \left\langle \frac{dE}{dt} \right\rangle_{DD},$$

$$\left\langle \frac{dE}{dt} \right\rangle_N = -\frac{G^2 m (-2E\mu)^{3/2}}{15c^5 \bar{L}^7} (148E^2 \bar{L}^4 + 732G^2 m^2 \mu^3 E \bar{L}^2 + 425G^4 m^4 \mu^6),$$

$$\left\langle \frac{dE}{dt} \right\rangle_{QM} = \frac{G^2 \mu m^3 (-2E\mu)^{3/2}}{30c^5 \bar{L}^{11}} \sum_{i=1}^2 p_i [C_1 \sin^2 \kappa_i \cos 2\delta_i + C_2 (2 - 3 \sin^2 \kappa_i)],$$

$$\left\langle \frac{dE}{dt} \right\rangle_{DD} = \frac{G^2 d_1 d_2 (-2E\mu)^{3/2}}{15c^5 \bar{L}^{11}} (C_1 B_0 + C_2 A_0),$$

$C_k$  coefficients:

$$C_1 = -\mu \bar{A}^2 (948E^2 \bar{L}^4 + 8936G^2 E \bar{L}^2 m^2 \mu^3 + 8335G^4 m^4 \mu^6),$$

$$C_2 = 708E^3 \bar{L}^6 + 10020G^2 E^2 \bar{L}^4 m^2 \mu^3 + 18865G^4 E \bar{L}^2 m^4 \mu^6 + 8316G^6 m^6 \mu^9.$$

$$\left\langle \frac{dL}{dt} \right\rangle = \left\langle \frac{dL}{dt} \right\rangle_N + \left\langle \frac{dL}{dt} \right\rangle_{QM} + \left\langle \frac{dL}{dt} \right\rangle_{DD},$$

$$\left\langle \frac{dL}{dt} \right\rangle_N = -\frac{4G^2 m (-2\mu E)^{3/2}}{5c^5 \bar{L}^4} (14E \bar{L}^2 + 15G^2 m^2 \mu^3),$$

$$\left\langle \frac{dL}{dt} \right\rangle_{QM} = \frac{G^2 m^3 \mu (-2\mu E)^{3/2}}{10c^5 \bar{L}^8} \sum_{i=1}^2 p_i [D_1 \sin^2 \kappa_i \cos 2\delta_i + D_2 (2 - 3 \sin^2 \kappa_i)],$$

$$\left\langle \frac{dL}{dt} \right\rangle_{DD} = \frac{G^2 d_1 d_2 (-2E\mu)^{3/2}}{5c^5 \bar{L}^8} [D_1 B_0 + D_2 A_0],$$

$D_k$  coefficients:

$$D_1 = -6\mu \bar{A}^2 (31E \bar{L}^2 + 90G^2 m^2 \mu^3),$$

$$D_2 = 252E^2 \bar{L}^4 + 1200G^2 E \bar{L}^2 m^2 \mu^3 + 805G^4 m^4 \mu^6.$$

$$\left\langle \frac{dy}{dt} \right\rangle = 0 ,$$

$$\left\langle \frac{dy}{dt} \right\rangle = 0 ,$$

$$\left\langle \frac{dx_i}{dt} \right\rangle = \left\langle \frac{dx_i}{dt} \right\rangle_{GM} + \left\langle \frac{dx_i}{dt} \right\rangle_{DD} ,$$

$$\left\langle \frac{dx_i}{dt} \right\rangle_{GM} = \frac{3G^2 m^3 \mu (-2\mu E)^{3/2}}{10c^5 \bar{I}^9} \sum_{j=1}^2 p_j \sin 2\kappa_j [V_1 \cos(\delta_i + \delta_j) + V_2 \cos(\delta_i - \delta_j)] ,$$

$$\left\langle \frac{dx_i}{dt} \right\rangle_{DD} = \frac{3Gd_1 d_2 (-2E\mu)^{3/2}}{5c^5 \bar{I}^9} \sum_{j=1}^2 \mu_{3-2j} [V_1 [\sigma_j \cos(\delta_i + \delta_j) - \rho_j \sin(\delta_i + \delta_j)] + V_2 [\sigma_j \cos(\delta_i - \delta_j) + \rho_j \sin(\delta_i - \delta_j)]] ,$$

$V_{1-3}$  coefficients:

$$V_1 = 40E^2 \bar{I}^4 + 90G^2 m^2 \mu^3 E \bar{I}^2 + 35G^4 m^4 \mu^6 ,$$

$$V_2 = 48E^2 \bar{I}^4 + 140G^2 m^2 \mu^3 E \bar{I}^2 + 70G^4 m^4 \mu^6 .$$

$$\left\langle \frac{d \cos \beta_i}{dt} \right\rangle_{DD} = \frac{Gd_1 d_2 (-2E\mu)^{3/2}}{10c^5 \bar{I}^9} \left\{ (6\rho_1 Z_1 - 2 \cos \beta_1 D_1) B_0 + (\rho_1 Z_2 - 2 \cos \beta_1 D_2) A_0 - \frac{6}{\sin \kappa_i} \sum_{j=1}^2 \zeta_{3-j} \times \right.$$

$$\left. \{ V_1 [(\rho_1 \sigma_j \cos \kappa_i - \sigma_j \cos \kappa_i \cos \beta_i + \rho_j \nu_i) \cos(\delta_i + \delta_j) + (-\rho_1 \rho_j \cos \kappa_i + \rho_j \cos \kappa_i \cos \beta_i + \sigma_j \nu_i) \sin(\delta_i + \delta_j)] + V_2 [(\rho_1 \sigma_j \cos \kappa_i - \sigma_j \cos \kappa_i \cos \beta_i - \rho_j \nu_i) \cos(\delta_i - \delta_j) + (\rho_1 \rho_j \cos \kappa_i - \rho_j \cos \kappa_i \cos \beta_i + \sigma_j \nu_i) \sin(\delta_i - \delta_j)] \} \right\} ,$$

where:

$$Z_1 = 96E^2 \bar{I}^4 + 314EG^2 \bar{I}^2 m^2 \mu^3 + 133G^4 m^4 \mu^6 ,$$

$$Z_2 = 396E^2 \bar{I}^4 + 1380EG^2 \bar{I}^2 m^2 \mu^3 + 735G^4 m^4 \mu^6 .$$

Algebraic relations:

$$S_1 \sin \kappa_1 \cos \psi_1 + S_2 \sin \kappa_2 \cos \psi_2 = 0 ,$$

$$\cos \gamma = \cos \kappa_1 \cos \kappa_2 + \cos(\psi_2 - \psi_1) \sin \kappa_1 \sin \kappa_2 ,$$

$$\cos \lambda = (\rho_2 \sigma_1 - \rho_1 \sigma_2) \sin(\delta_1 - \delta_2) + (\rho_1 \rho_2 + \sigma_1 \sigma_2) \cos(\delta_1 - \delta_2) + \zeta_1 \zeta_2 .$$

## REMARK:

### Circular orbit limit

of eccentric ENERGY LOSS agrees with earlier results:

(QM) Poisson, Phys Rev D **57**, 5287 (1998)

(Md - Md) Ioka & Taniguchi, Astrophys. J. **537**, 327 (2000)

## FUTURE PROSPECTS:

- 2.5PN order 1th correction to SO
- Extension to angular motion (non-conservation of  $L$  requires a generalized Damour-Deruelle formalism)
- Orbital and signal frequency evolution for eccentric orbits