

# Photon testbeam

## Data/G4 comparison



- ① Motivation
- ② Testbeam setup & simulation
- ③ Analysis & results
- ④ Conclusion

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## Motivation

Photon identification is crucial for physics analysis in **ATLAS**

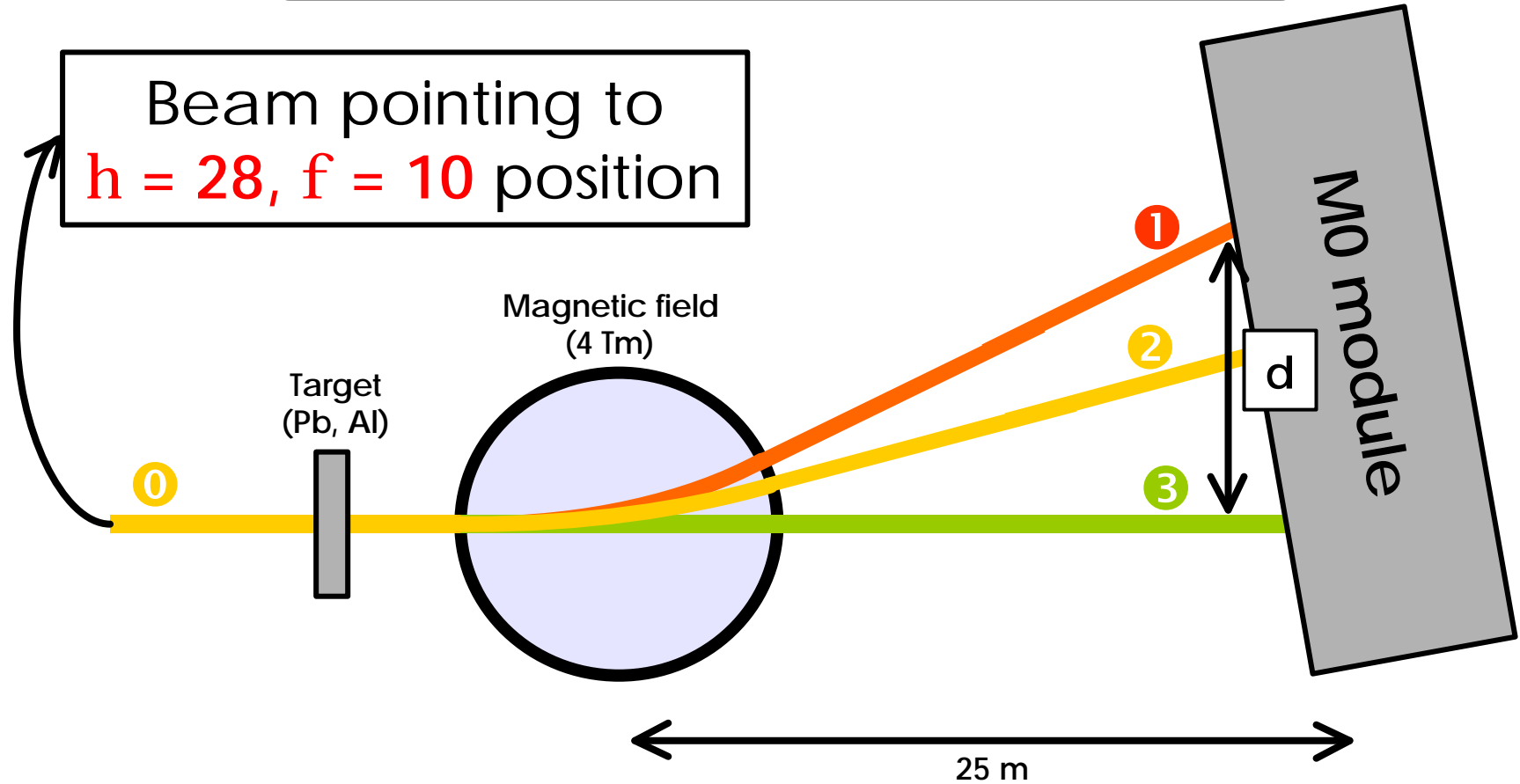
Numerous complex effects are involved (cross-talk, electronic noise,...)

Is **G4** able to reproduce correctly photon in the Ecal?

Compare data (taken on Ecal barrel in **August 2000**) to a G4 simulation taking into account those effects

The Photon TestBeam experimental setup  
1. Technical description

Beam pointing to  
 $h = 28, f = 10$  position



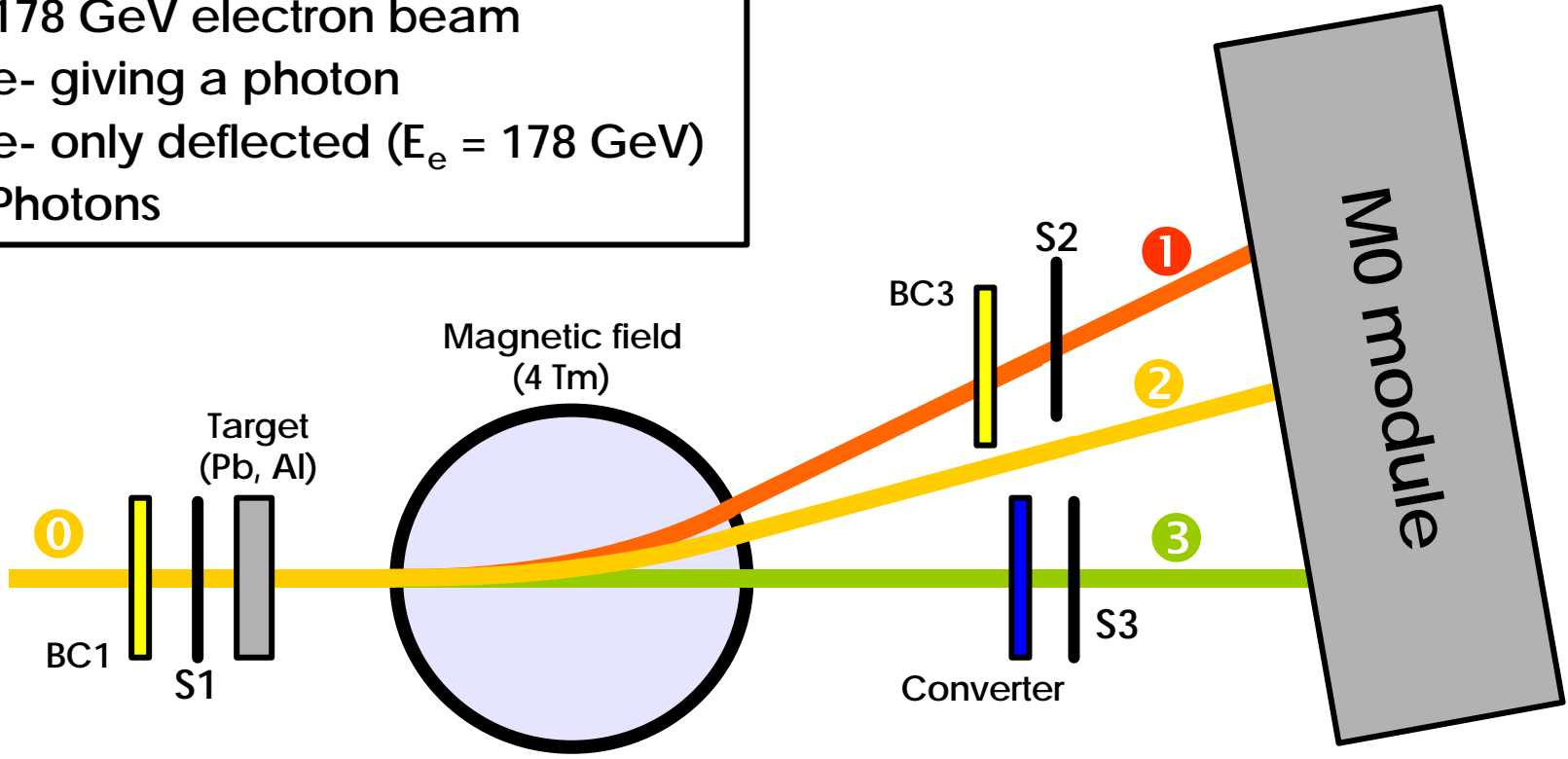
- 0 178 GeV electron beam
- 1 e- giving a photon
- 2 e- only deflected ( $E_e = 178$  GeV)
- 3 Photons

$$d = \left( \frac{3060}{E_e} \right) \text{ cm}$$

# The Photon TestBeam experimental setup

## 2. Triggering & event filtering

- 0 178 GeV electron beam
- 1 e- giving a photon
- 2 e- only deflected ( $E_e = 178$  GeV)
- 3 Photons

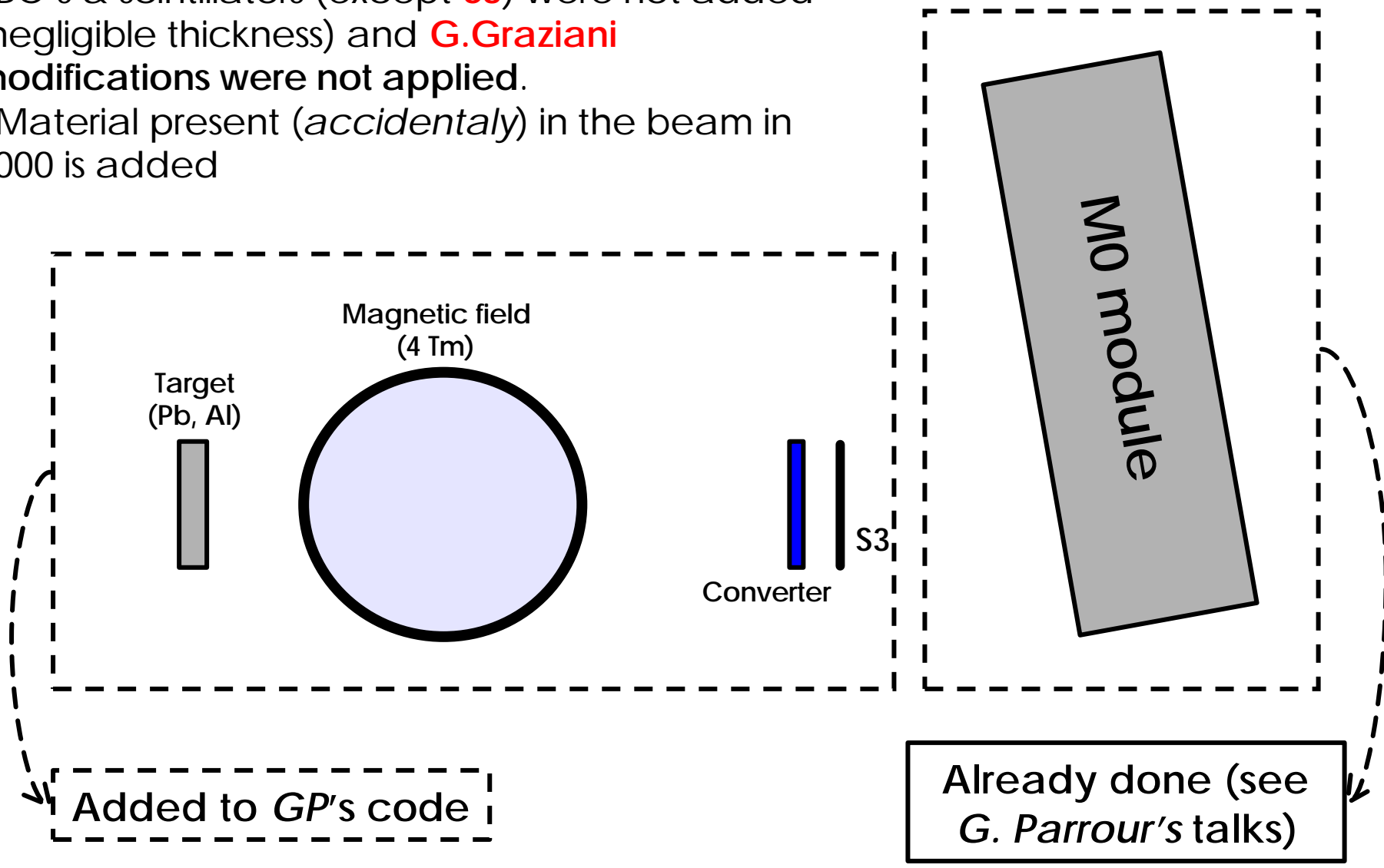


{ Trigger  $\hat{U}$  S1  $\hat{U}$  S2  
 Multiphoton rejection  $\hat{U}$  Converter + S3  
 Beam profile  $\hat{U}$  BC's

# The Photon TestBeam experimental setup

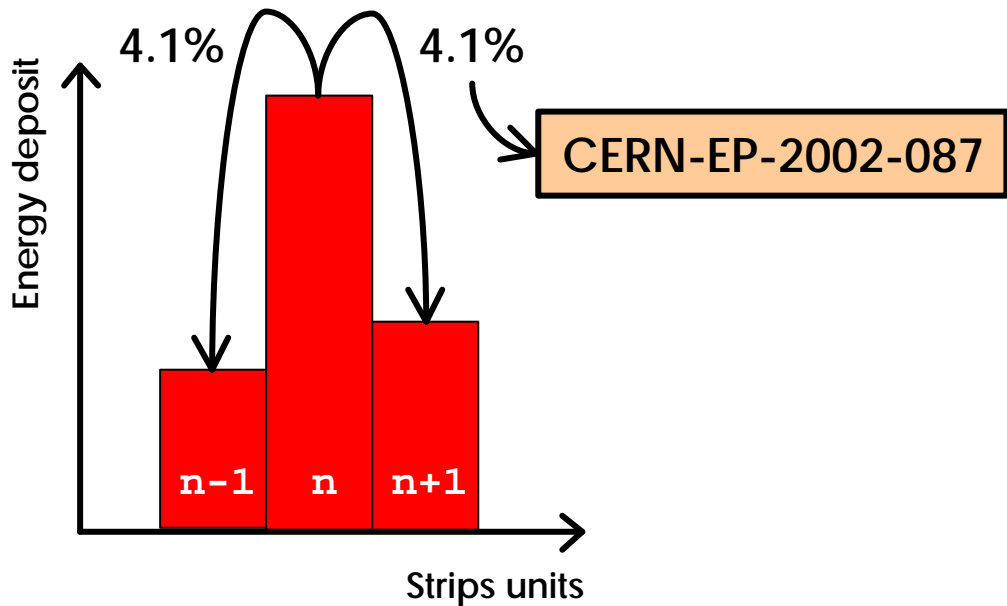
## 3. G4 Modelisation

- Simulation developed with G4 **release 4.4.1**
- BC's & scintillators (except **S3**) were not added (negligible thickness) and **G.Graziani modifications were not applied.**
- Material present (*accidentally*) in the beam in 2000 is added

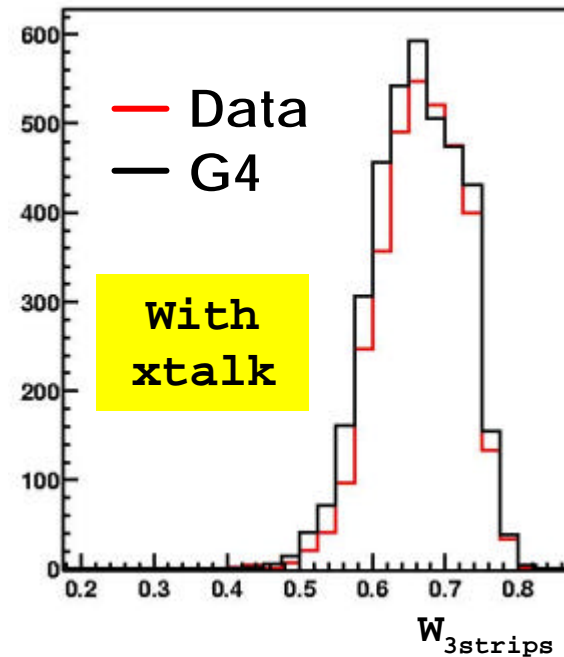
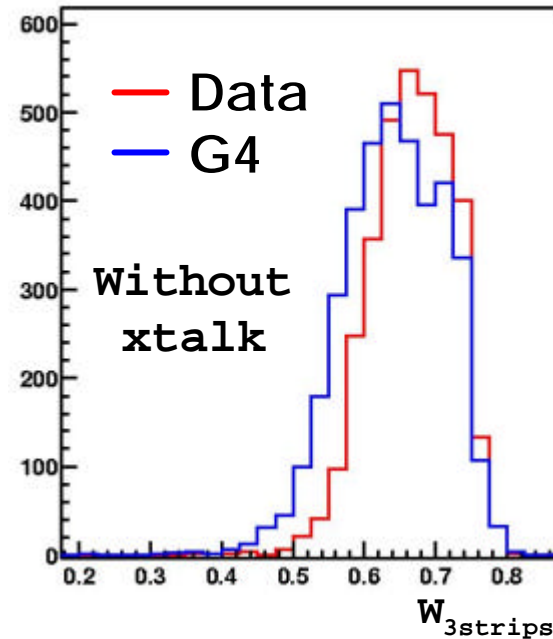


Calibration  
1. Crosstalk

Significant effect on precise parameters (Width on 3 strips,...)

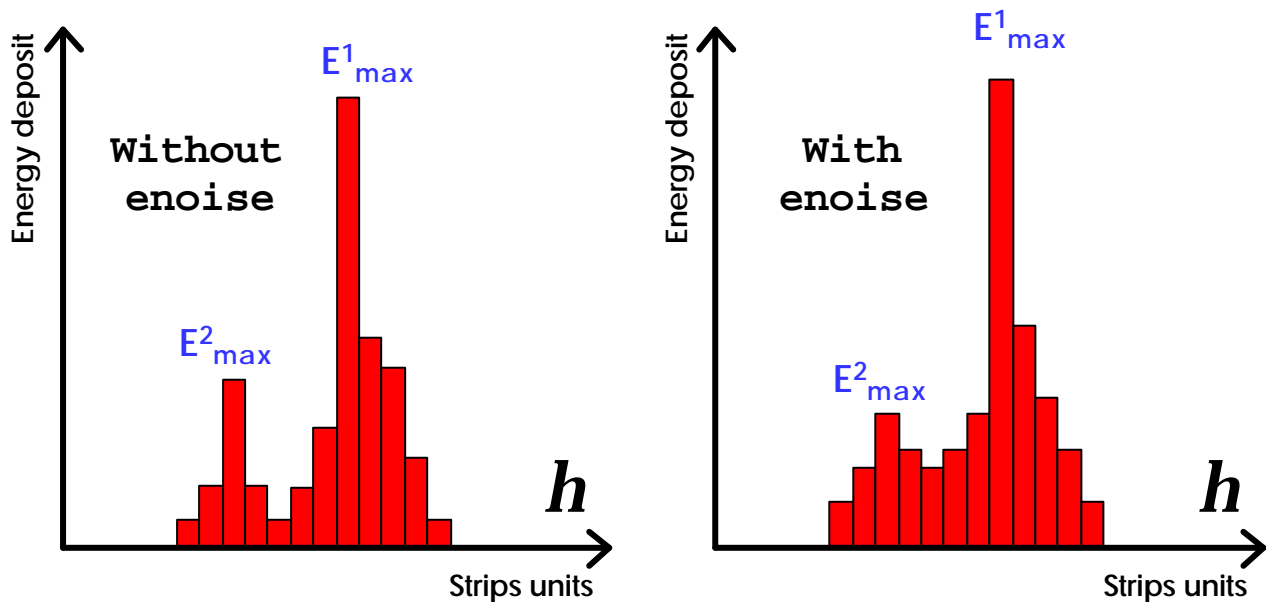


Only x-talk from closest neighbours in the strips is added

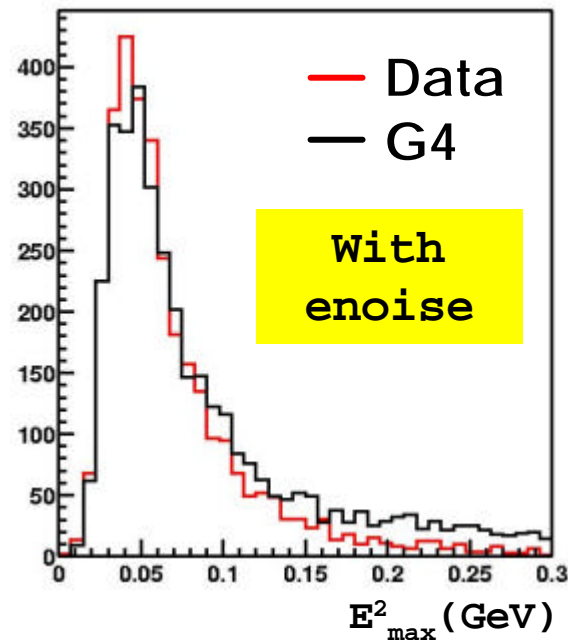
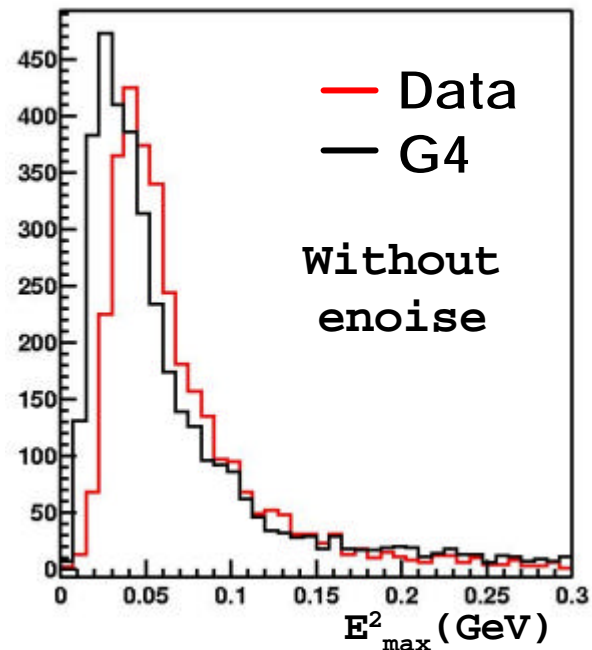


Calibration  
2. e-noise

Significant effect on  $g/p^0$  rejection parameters (Energy of 2<sup>nd</sup> maximum,...)



e-noise softens the difference between  $E^2_{max}$  the valley, particularly at low energy



Analysis  
1.Cuts applied

Photon candidate if :

$$5 \text{ GeV} < E_{\text{ph}} < 55 \text{ GeV}$$

*Too much background under 5 GeV*

$$0.93 E_{\text{calo}} < E_{\text{ph}} + E_{\text{el}} < 0.96 E_{\text{calo}}$$

~~*Multiphoton*~~

$$226. < \eta_{\text{strips}}(\text{Photon}) < 230.$$

~~*Electron position dispersion*~~

$$S3 \text{ signal } \hat{U} \text{ pedestal}$$

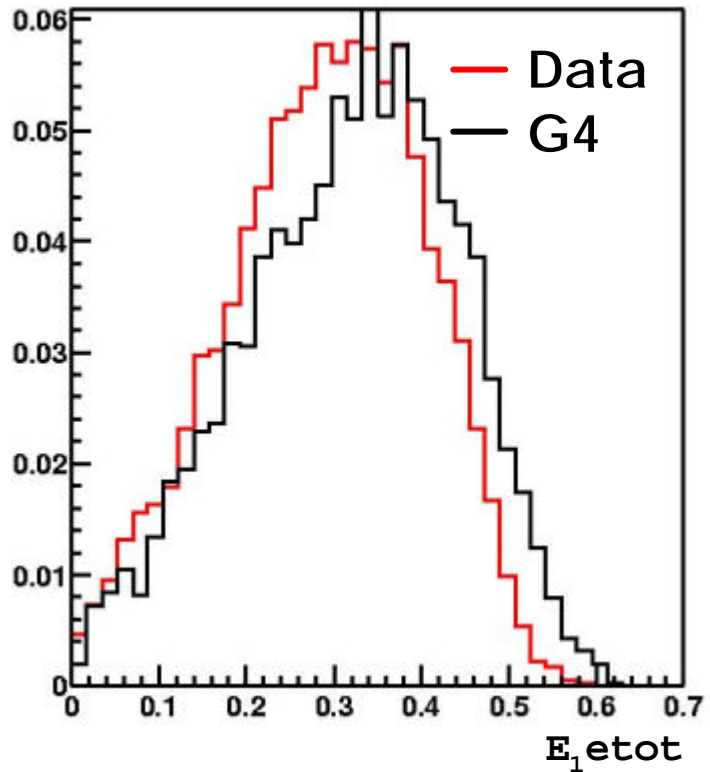
~~*Multiphoton*~~



**Results**  
 1. Energy in the strips & leakage proportion

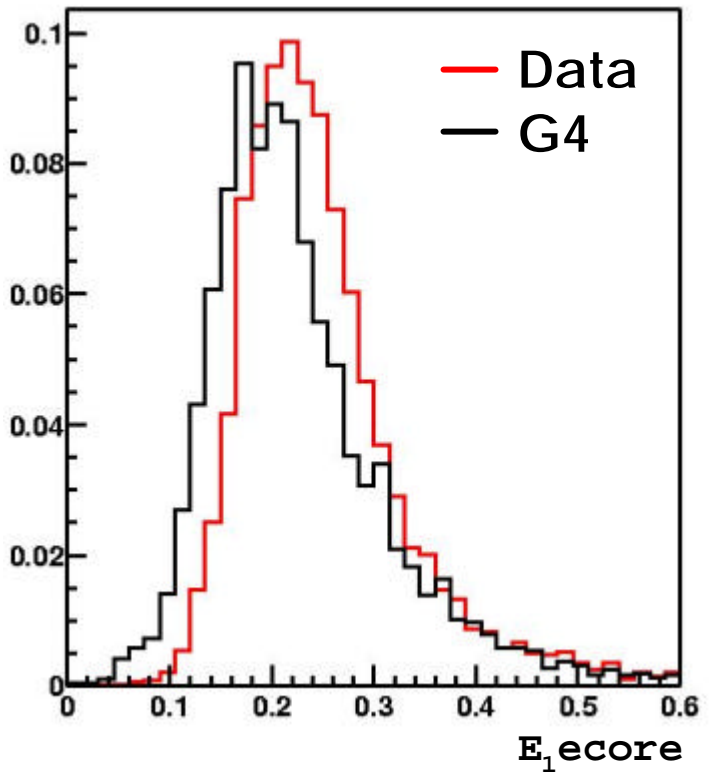
Fraction of energy in the strips

$$E_{1etot} = \frac{E_{\text{photon}}(\text{strips})}{E_{\text{photon}}}$$



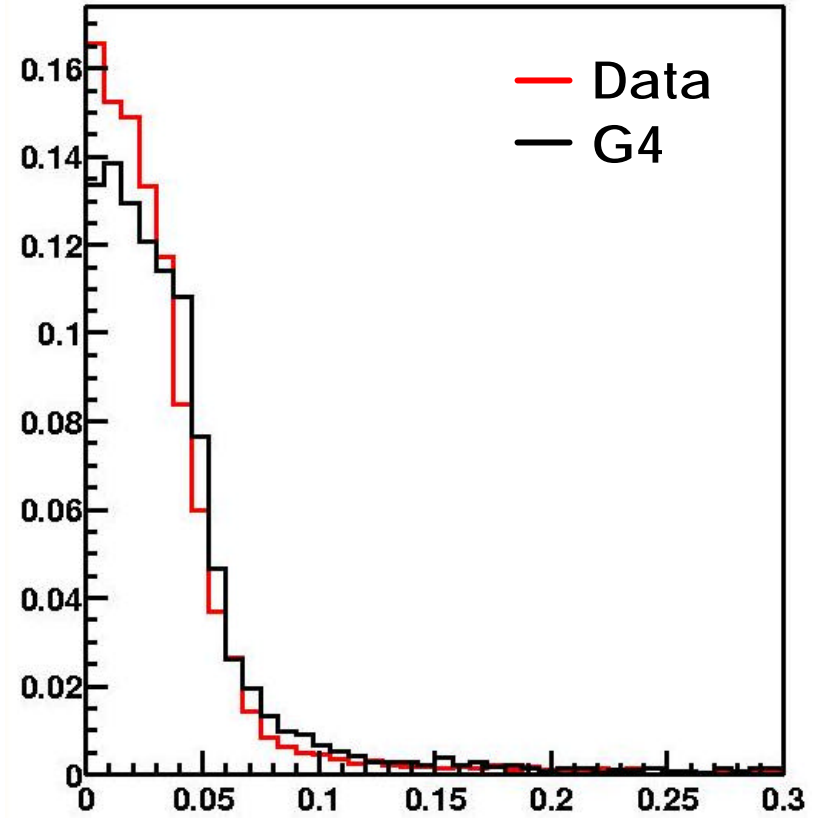
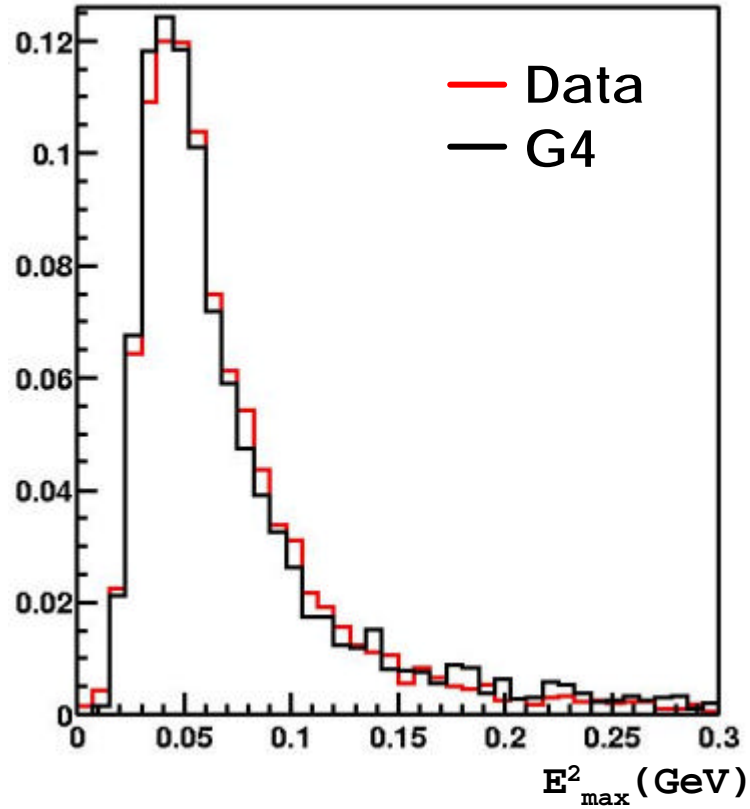
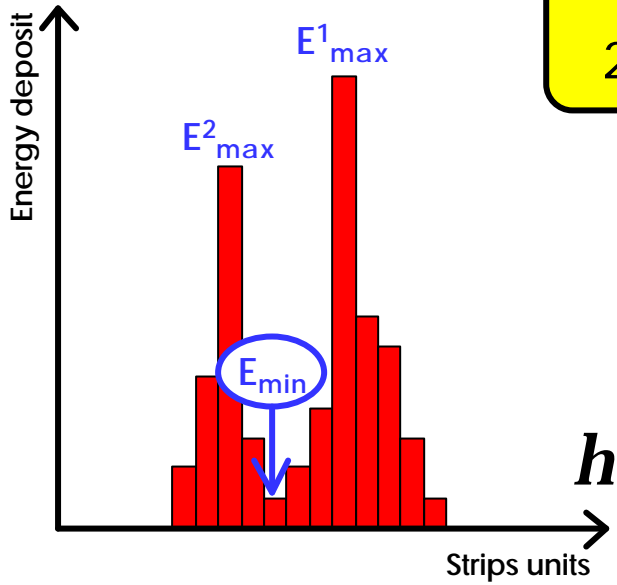
Leakage fraction in the strips

$$E_{1ecore} = \frac{E_{\text{photon}}(7\text{strips}) - E_{\text{photon}}(3\text{strips})}{E_{\text{photon}}(3\text{strips})}$$



➔ Problem with upstream material modelisation and calibration with xtalk in the strips (see G. Graziani 's 01/10/03 talk)

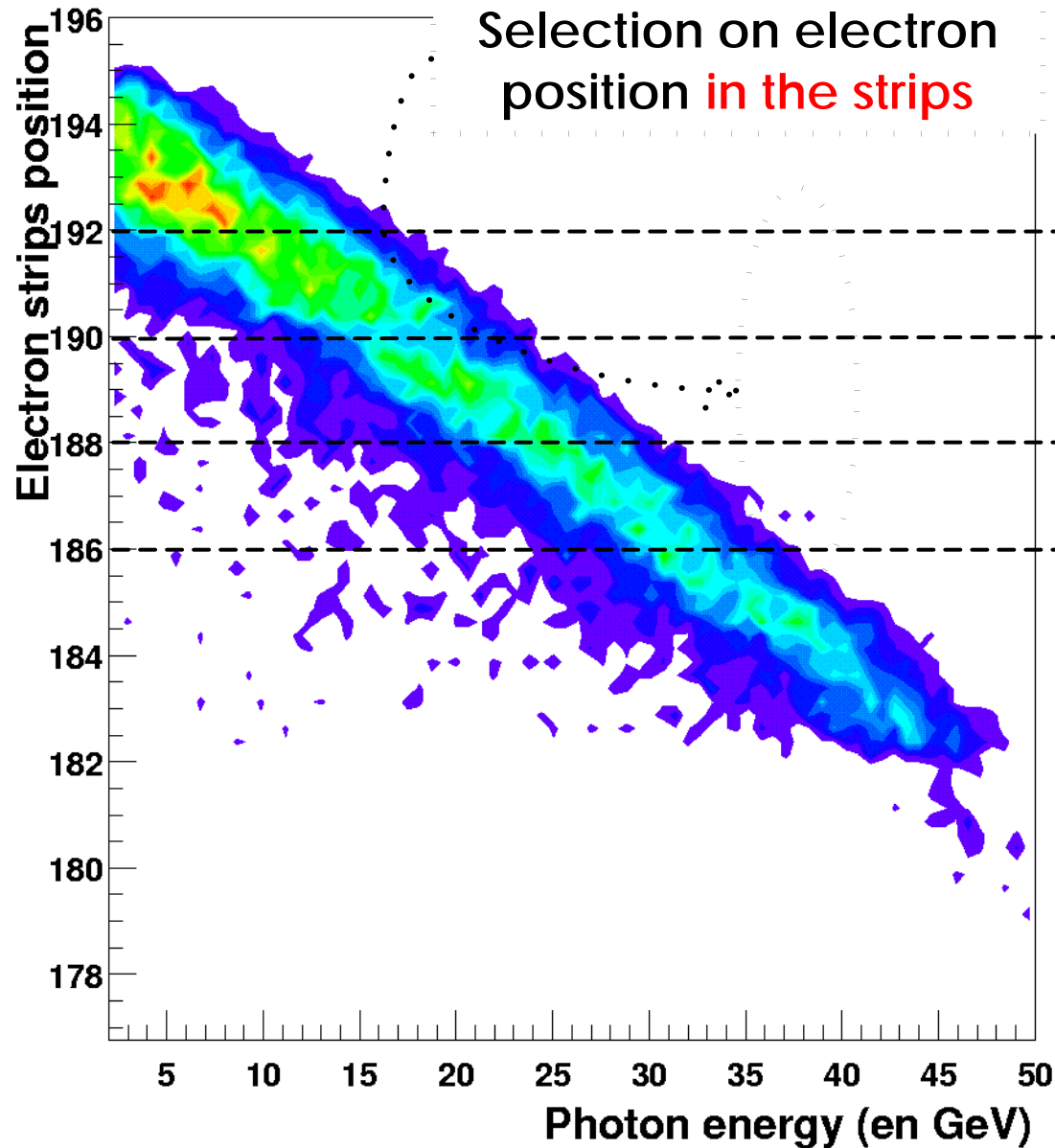
Results  
2.  $\gamma/\pi^0$  rejection parameters



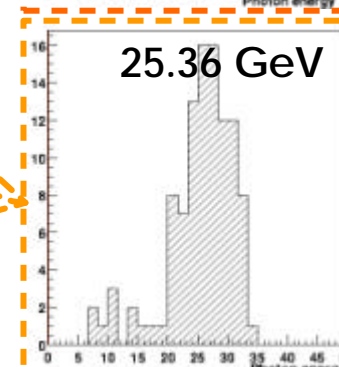
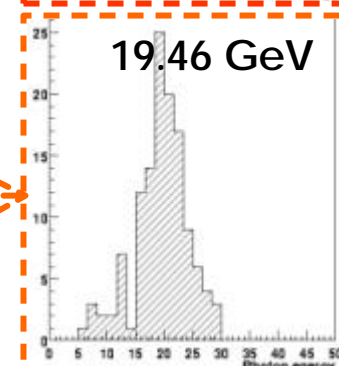
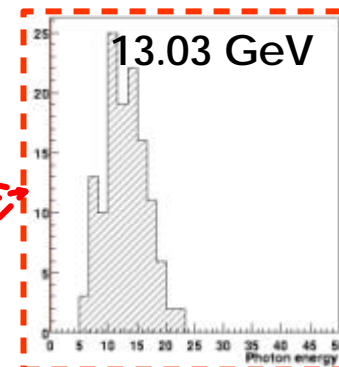
$$E_{\max}^{d2} = E_{\max}^2 - E_{\min}$$

➔ Good agreement G4/Data

**Analysis**  
 2. Using correlation between electron deviation & photon energy



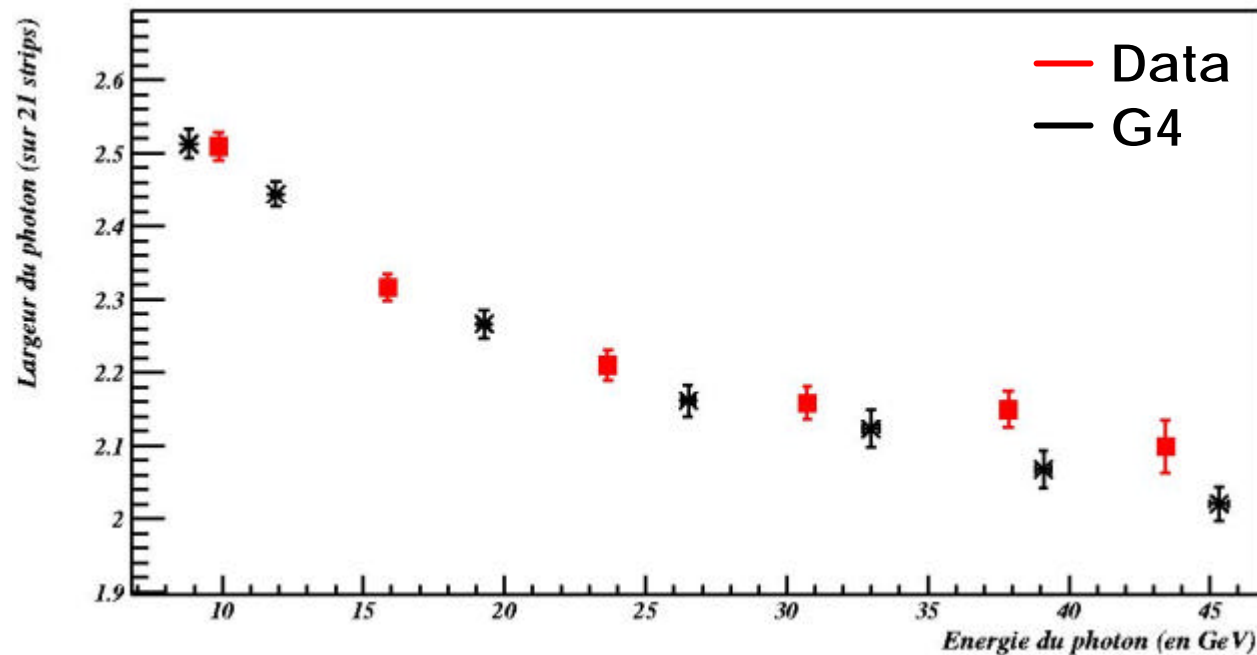
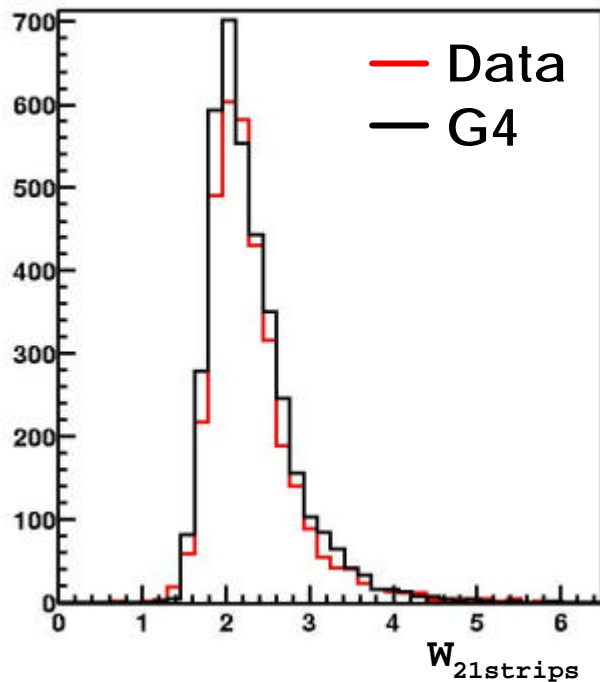
Selection on electron position **in the strips**



Compare **photon widths** at different  $\langle E_{ph} \rangle$

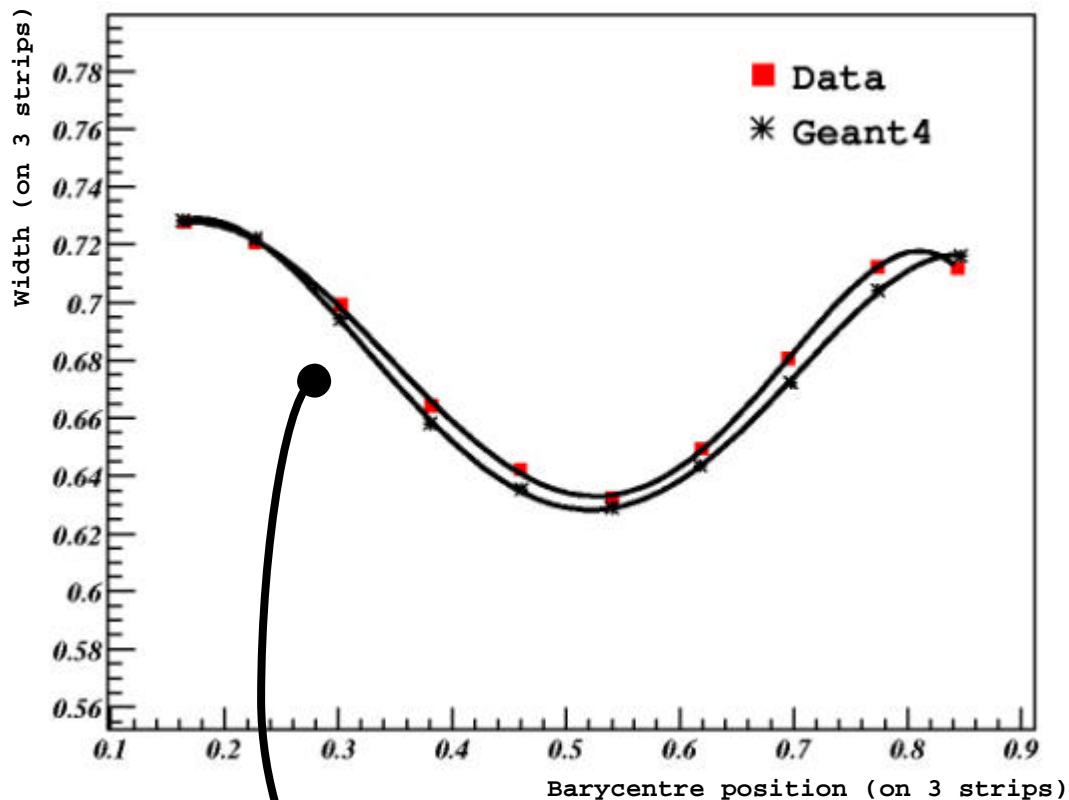
## Results

### 3. Width on 21 strips

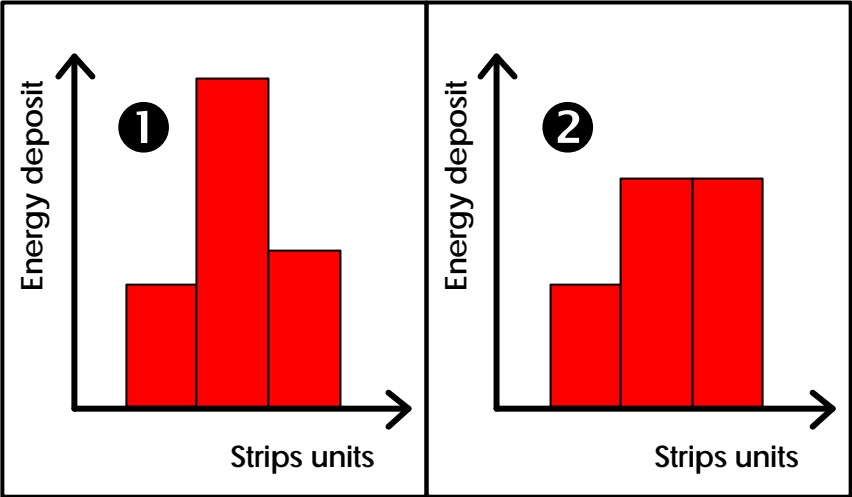


- Good agreement (the **width reduction with energy** is correctly reproduced with G4)
- xtalk effect **small**
- enoise effect **significant** at **low** energy

Results  
4. Width on 3 strips (1/2)



Strip granularity effect



$$E^{\textcircled{1}}_{\text{strips}} = E^{\textcircled{2}}_{\text{strips}}$$

but

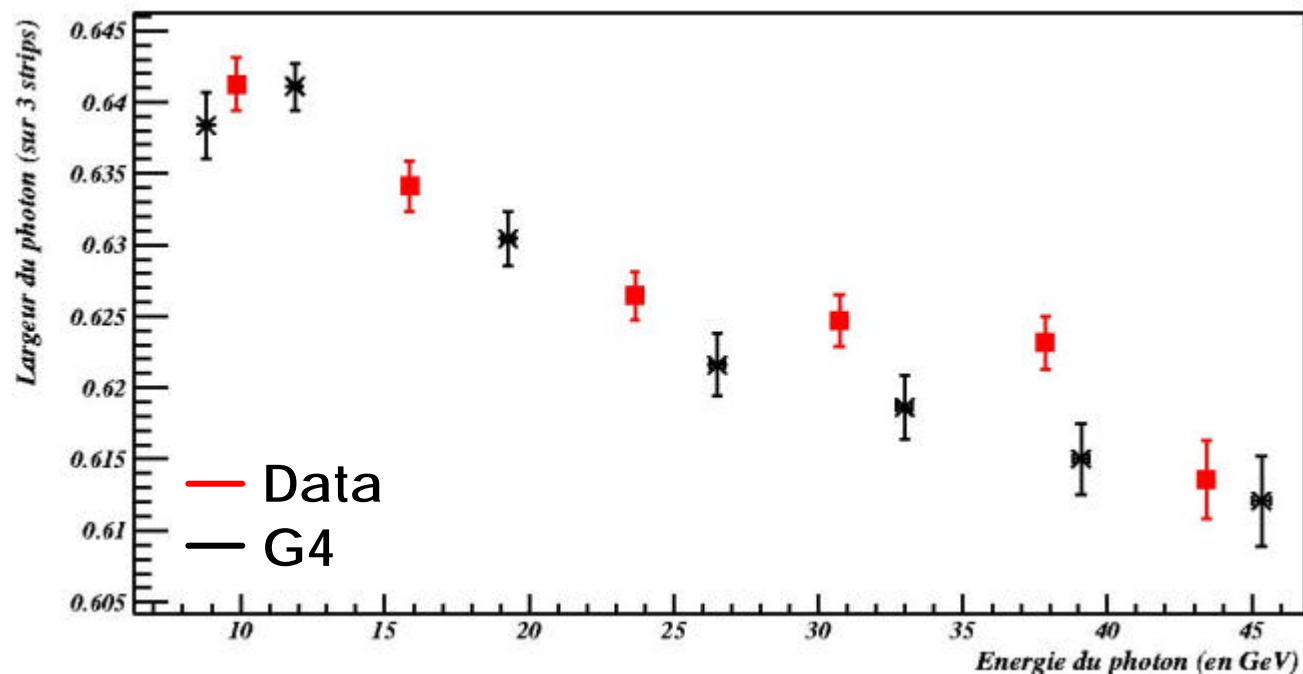
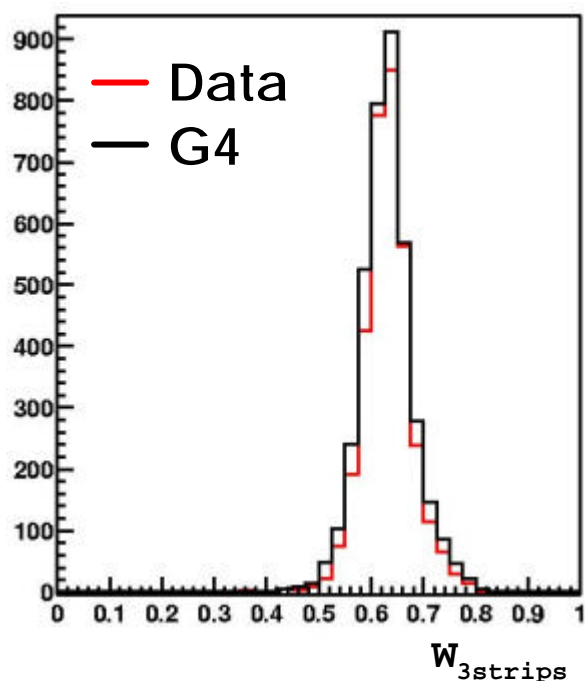
$$W^{\textcircled{1}}_3 < W^{\textcircled{2}}_3$$

**Geometrical effect** well reproduced by G4 simulation **including xtalk**

## Results

### 5. Width on 3 strips (2/2)

Ⓜ After correction of granularity effect (polynomial fit)



- Good agreement
- xtalk effect **significant** (see slide 9)
- enoise effect **small**

## Conclusions

- A comparison G4/Data of photon shape parameters in the ATLAS Ecal was performed.
- Crosstalk & e-noise were included
- The real experimental setup (with mag. field, converter, and target) was simulated (in order to compare width evolution with photon energy).

Energy independent measurements were compatible with G3 and data (except  $E_1^{\text{ecore}}$  and  $E_1^{\text{etot}}$ ).

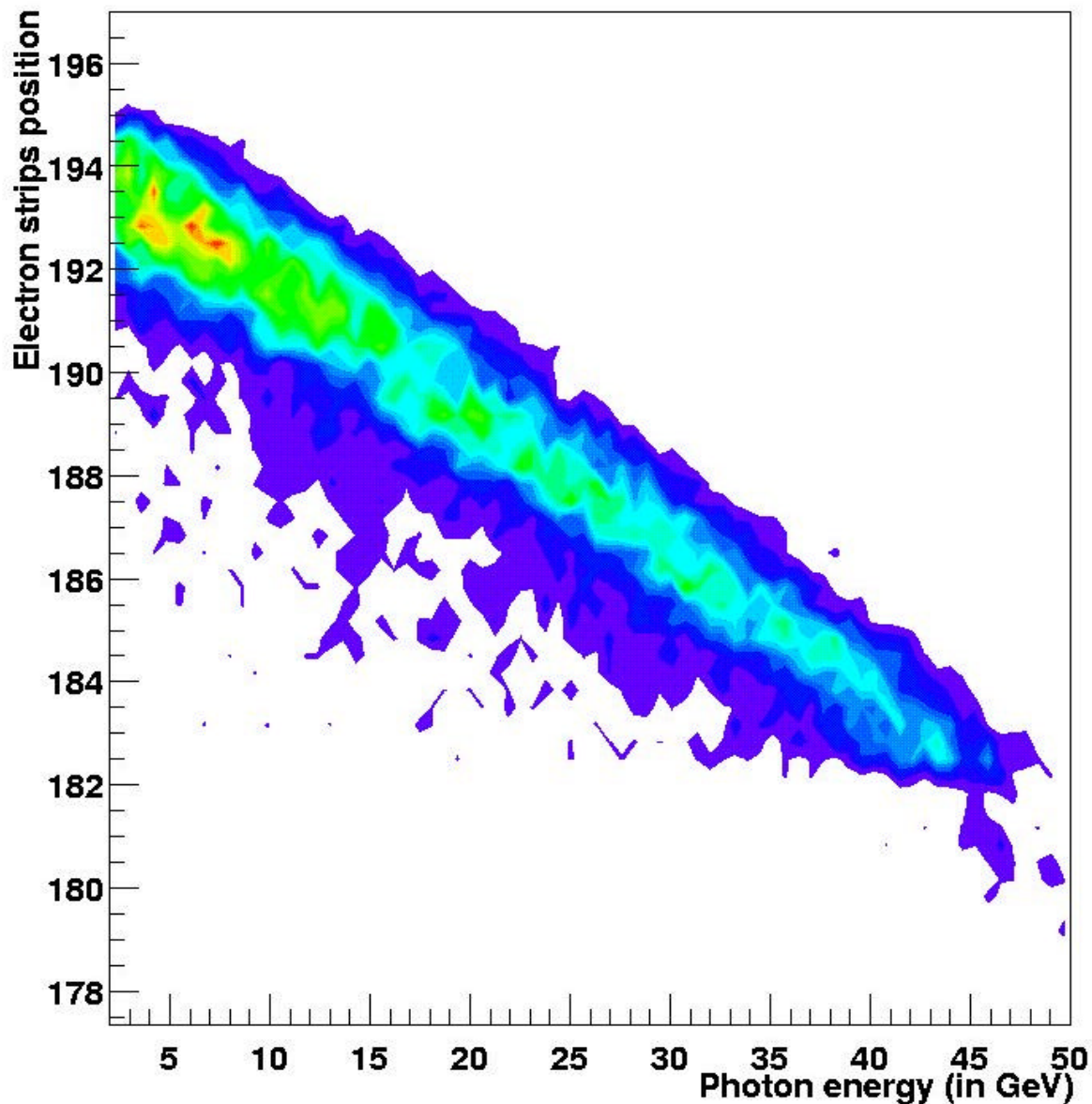
Energy dependent measurements (*not done with G3*) were in good agreement with data. The width reduction is well reproduced by Geant4

- The description of photon shower by G4 is equivalent to G3's.
- But some points still need to be understood (Is G. Graziani's electron's description valid for photons ?, ...).

**Backup  
slides**



# The correlation between electron position and Photon energy

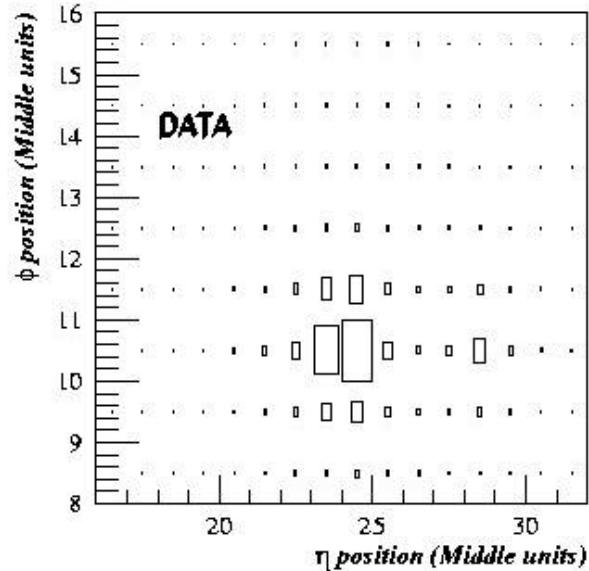
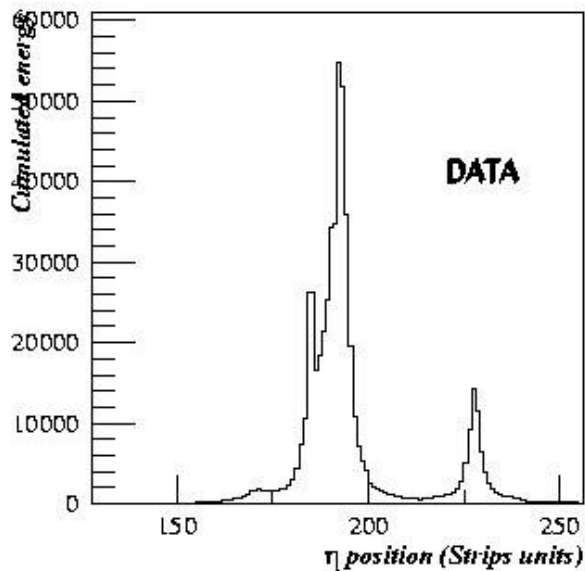
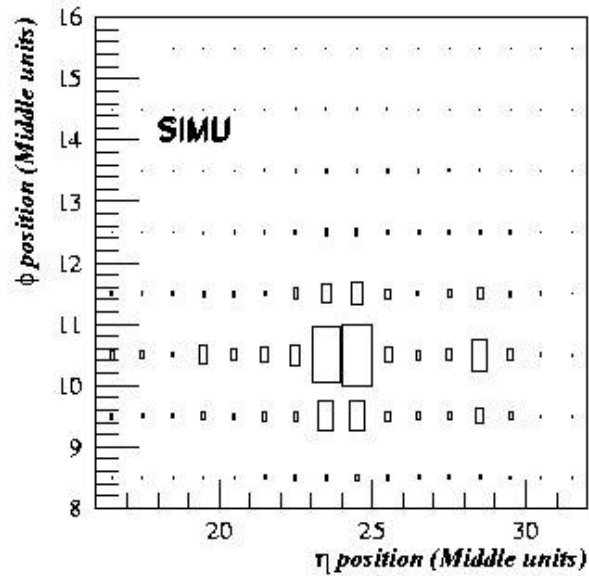
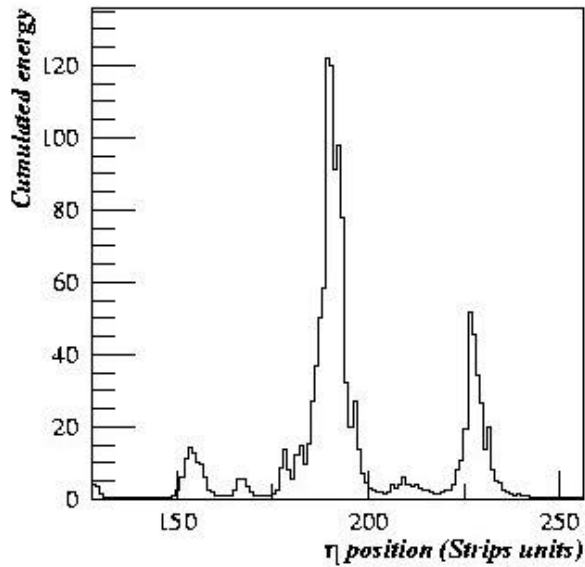


$$E_{ph} + E_e = 178 \text{ GeV}$$

$$d = \left( \frac{3060}{178 - E_{ph}} \right) \text{ cm}$$

Applying cuts on  $d$ , we could study photons at **different** energies

# Geometry & magnetic field verification

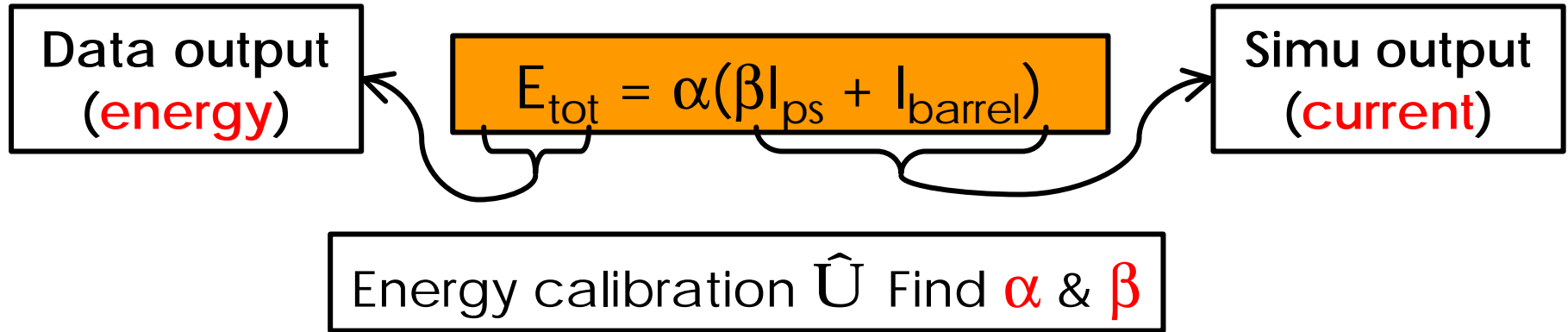


*f* & *h* position in agreement with data

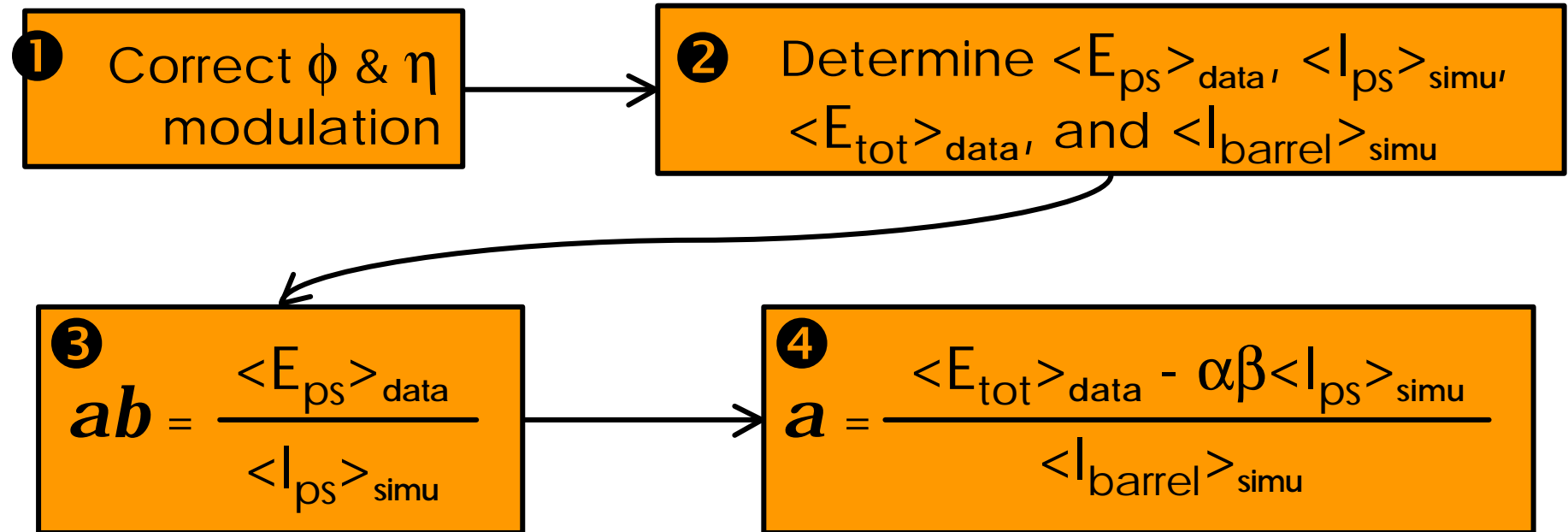
Geometry & Magnetic field are OK

Next step: Physics calibration

Energy calibration  
1. The method

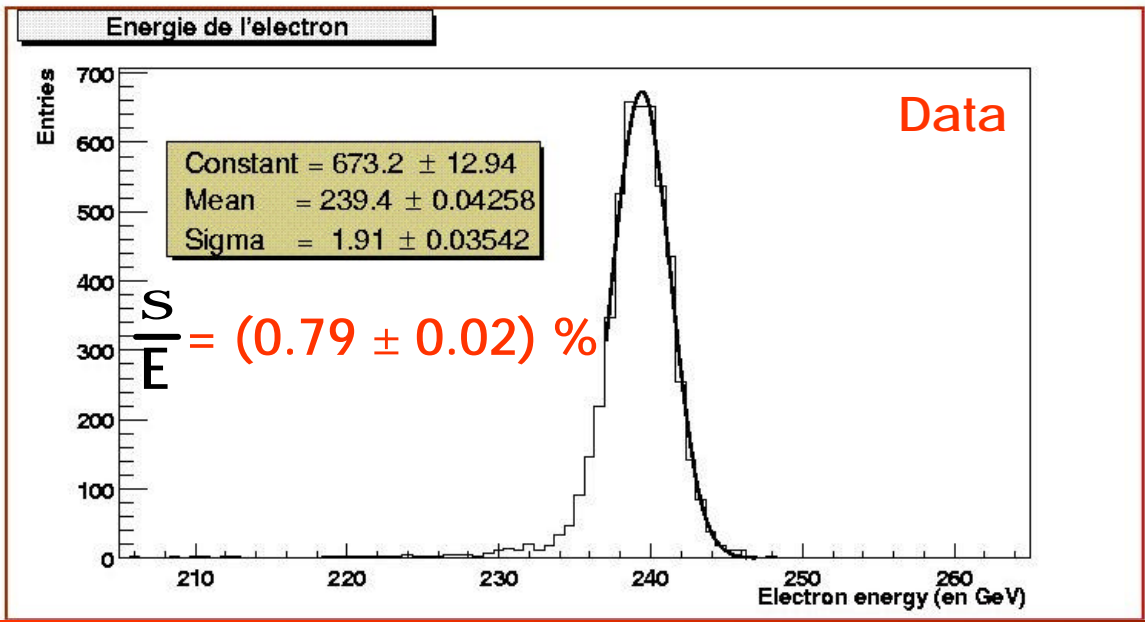
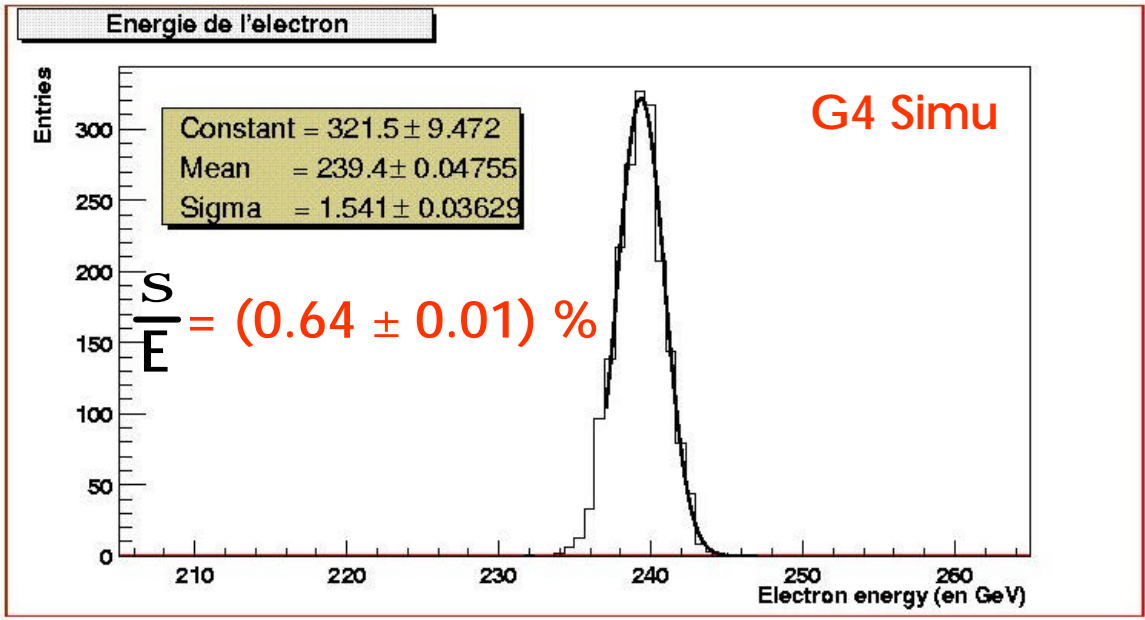


Basic method (only one cell (28,10)):



# Energy calibration

## 2. The problem



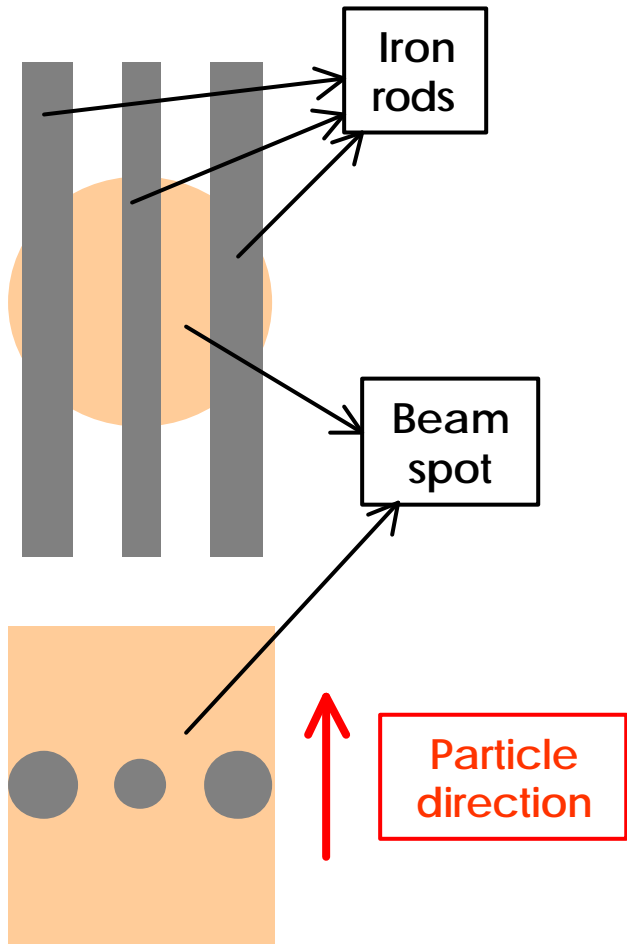
245 GeV electrons were simulated with the same program

MC resolution was **better** than expected

Something was missing in the simulation

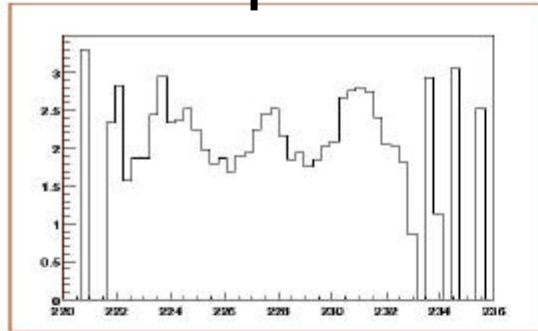
Material found in the beam in 2001 could be an answer to this problem

# Energy calibration 3. Adding the material

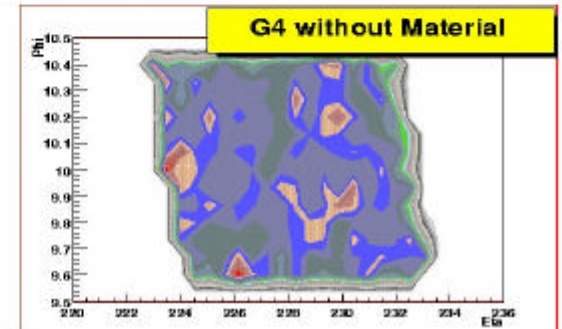
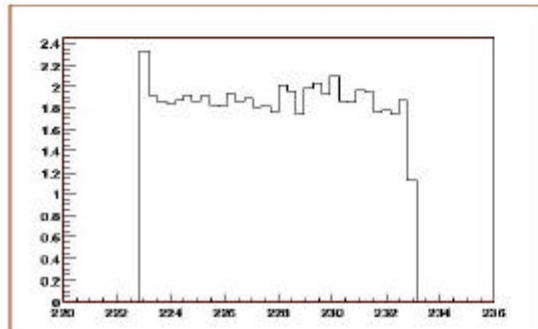
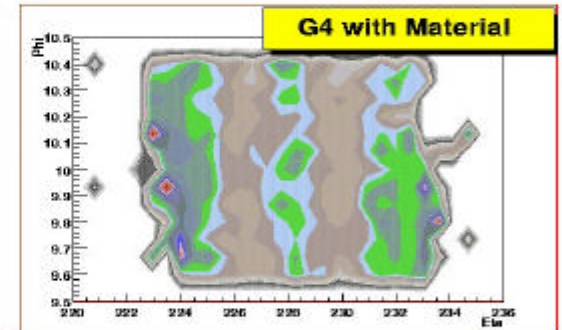
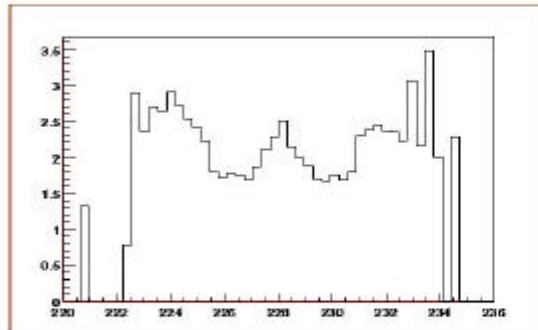
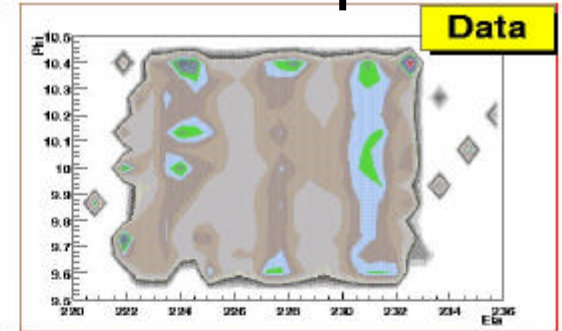


Material found in the beam in 2001

Eta profile



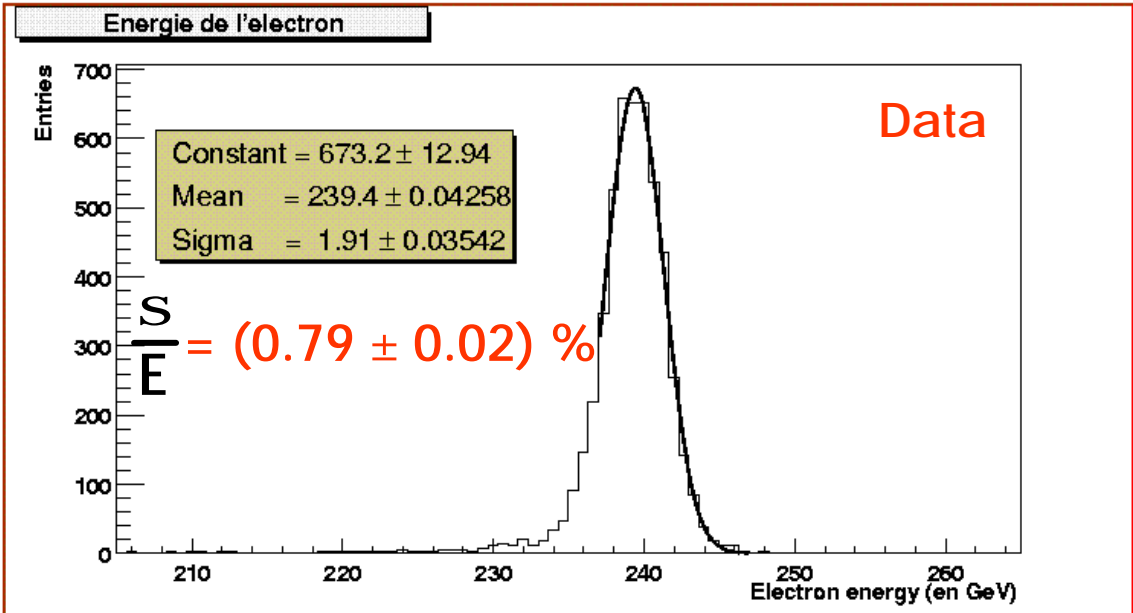
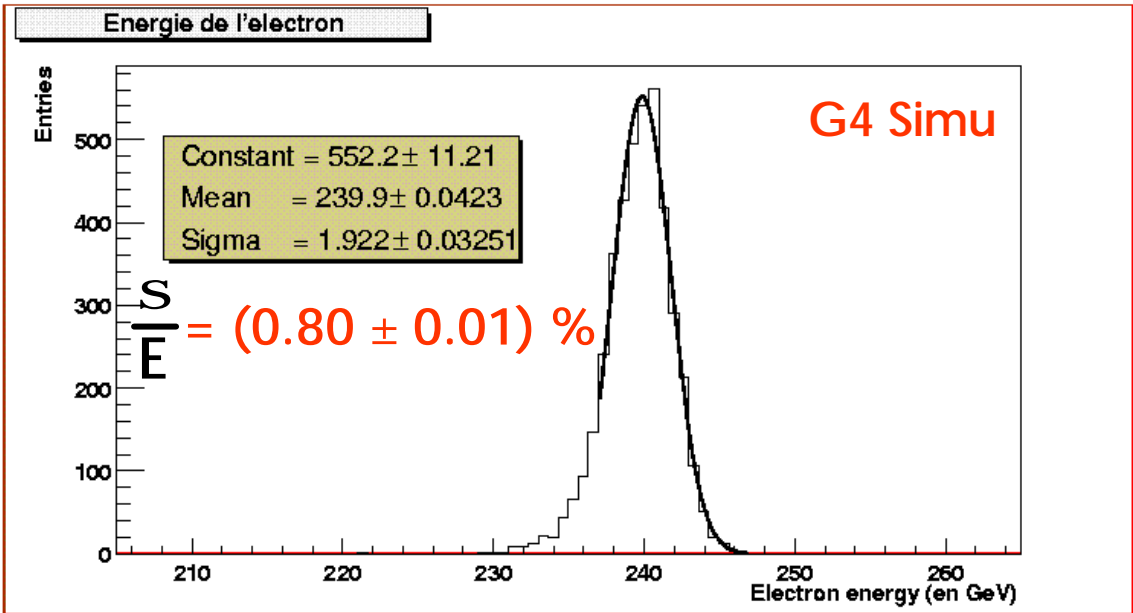
Eta vs. phi



Energy deposit in PS: material seems to be well reproduced

# Energy calibration

## 4. It looks better...



$$\alpha = 2.614$$
$$\beta = 5.787$$

MC & Data resolution  
in agreement

Energy calibration OK  
(for (28,10) cell)

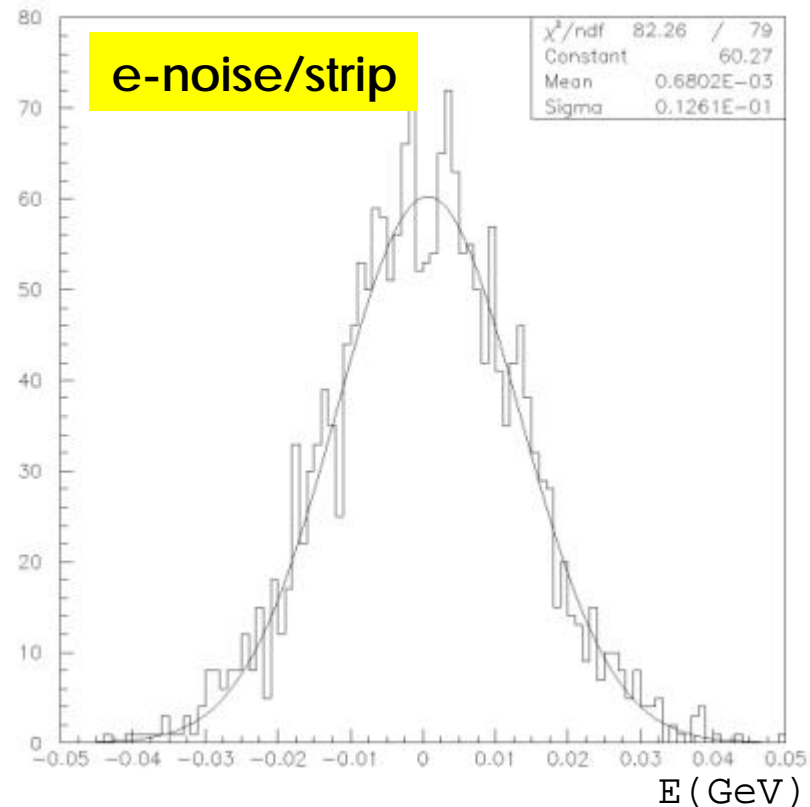
Next step: photons

## e-noise determination

e-noise could be deduced using  
**random** events (no signal)

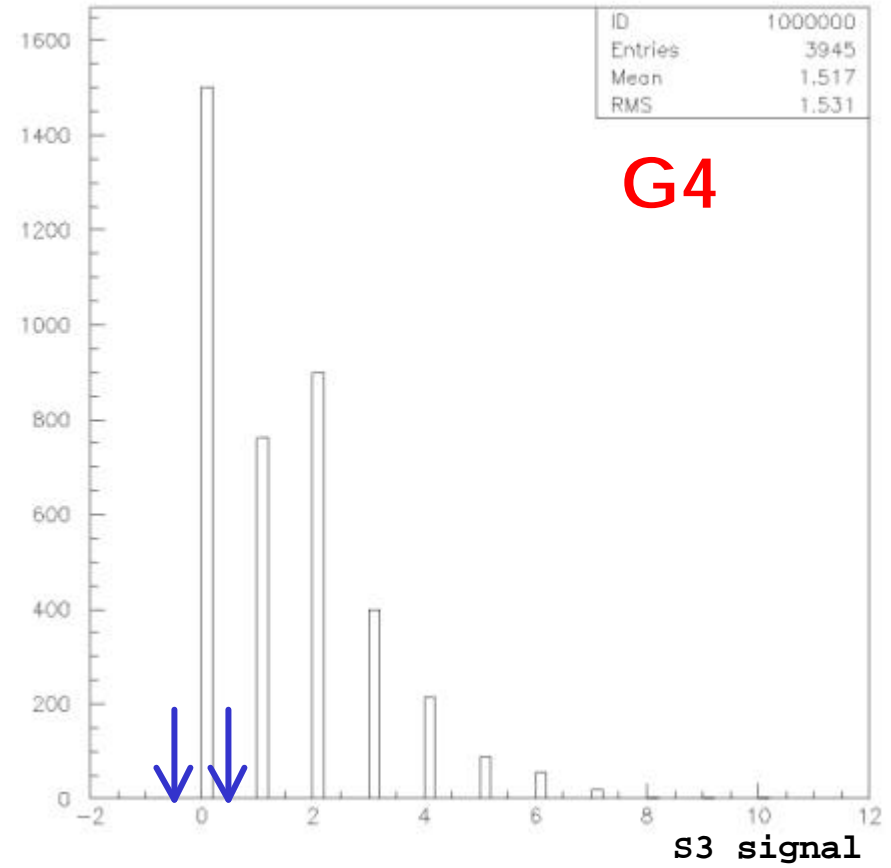
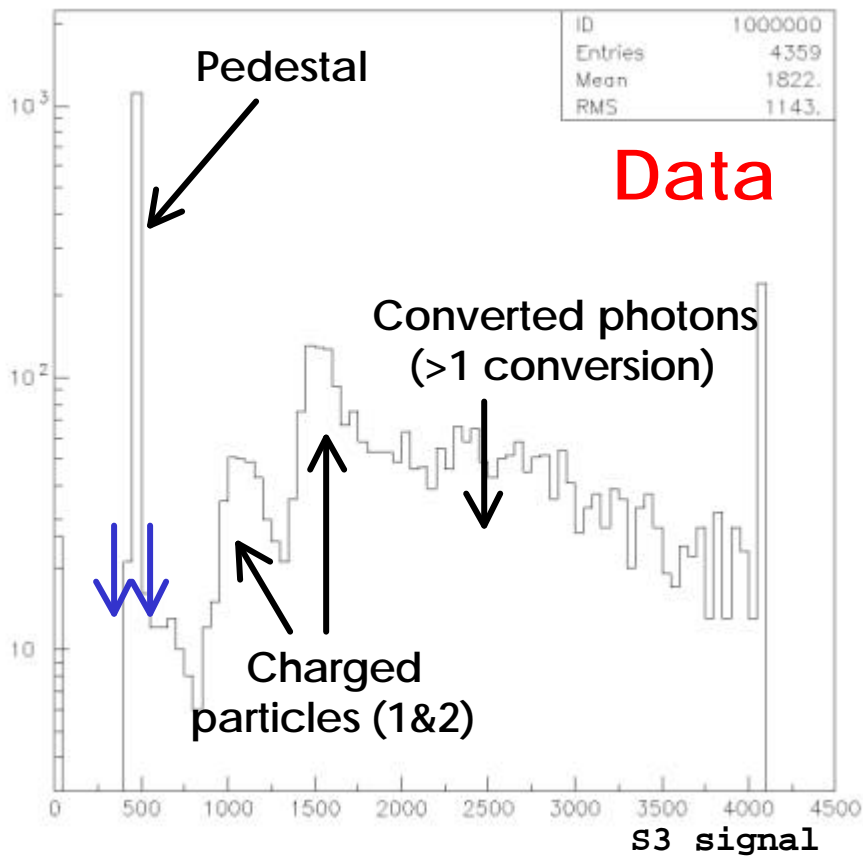


Fitted with a gaussian centered  
on **0** and with  **$s = 13\text{MeV}$**   
(compatible with TDR value)





# S3 cut



↓ ® cut values

Peak heights are not similar because data's values are smeared