

# Highlights from the WG : Machine Detector (and Physics) Interface

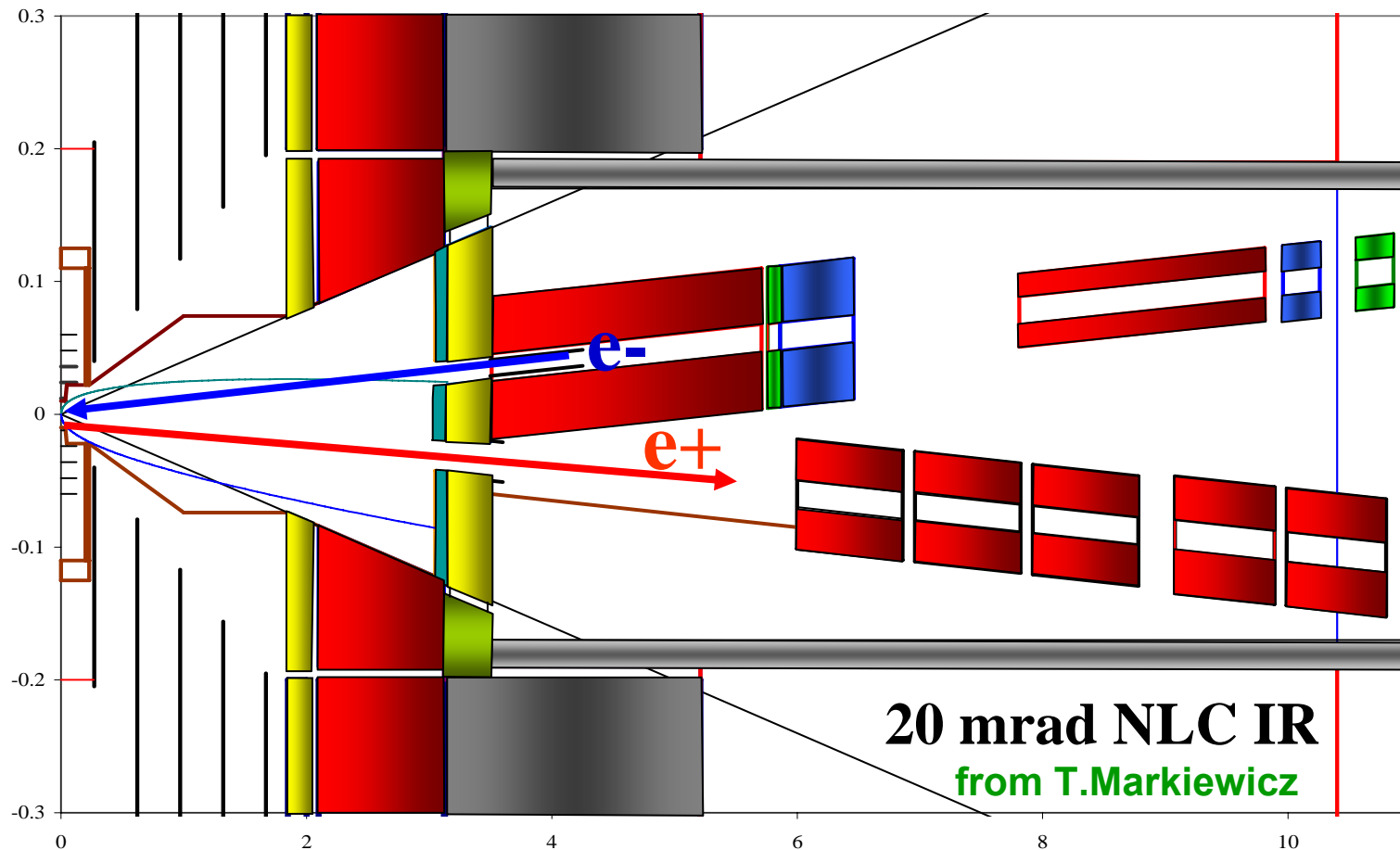
Four main themes :

1. IR design and crossing-angle choices
2. Very forward instrumentation
3. Machine backgrounds
4. Energy precision, stability, calibration

1. P. Bambade Summary of crossing-angle-or-not January 19 workshop
2. R. Appleby Alternative IR geometries for TESLA with small crossing angles
3. T. Aso Study of beam background at GLC including estimation by the BDS simulation from the exit of the LINAC to the beam dump
4. A. Stahl Beam-induced backgrounds in TESLA with  $l^*=4.1\text{m}$  optics and new masking scheme, with/without  $2^*10\text{mrad}$  crossing-angle.
5. T. Markiewicz IR design (collimation, backgrounds, crossing-angle)
6. K. Desch Physics impact of beam-beam hadron background
7. W. Lohmann Summary of Prague workshop on very forward instrumentation
8. V. Drugakov Detection of very forward Bhabha events and electron ID algorithm
9. J. Hauptman Detector technologies for forward calorimeters (quartz fiber-tungsten, Si-W, gas cherenkov-W, PbW04)
10. N. Delarue Beamstrahlung monitor
11. H. Yamamoto Pair (beam profile) monitor
12. M. Hildreth Energy spectrometers
13. T. Barklow Energy spread effects; energy precision
14. K. Kubo Energy spectrum measurement at the extraction line
15. S. Boogert Energy spectrum extraction from Bhabha events
16. All Discuss future inter-regional collaboration & working groups

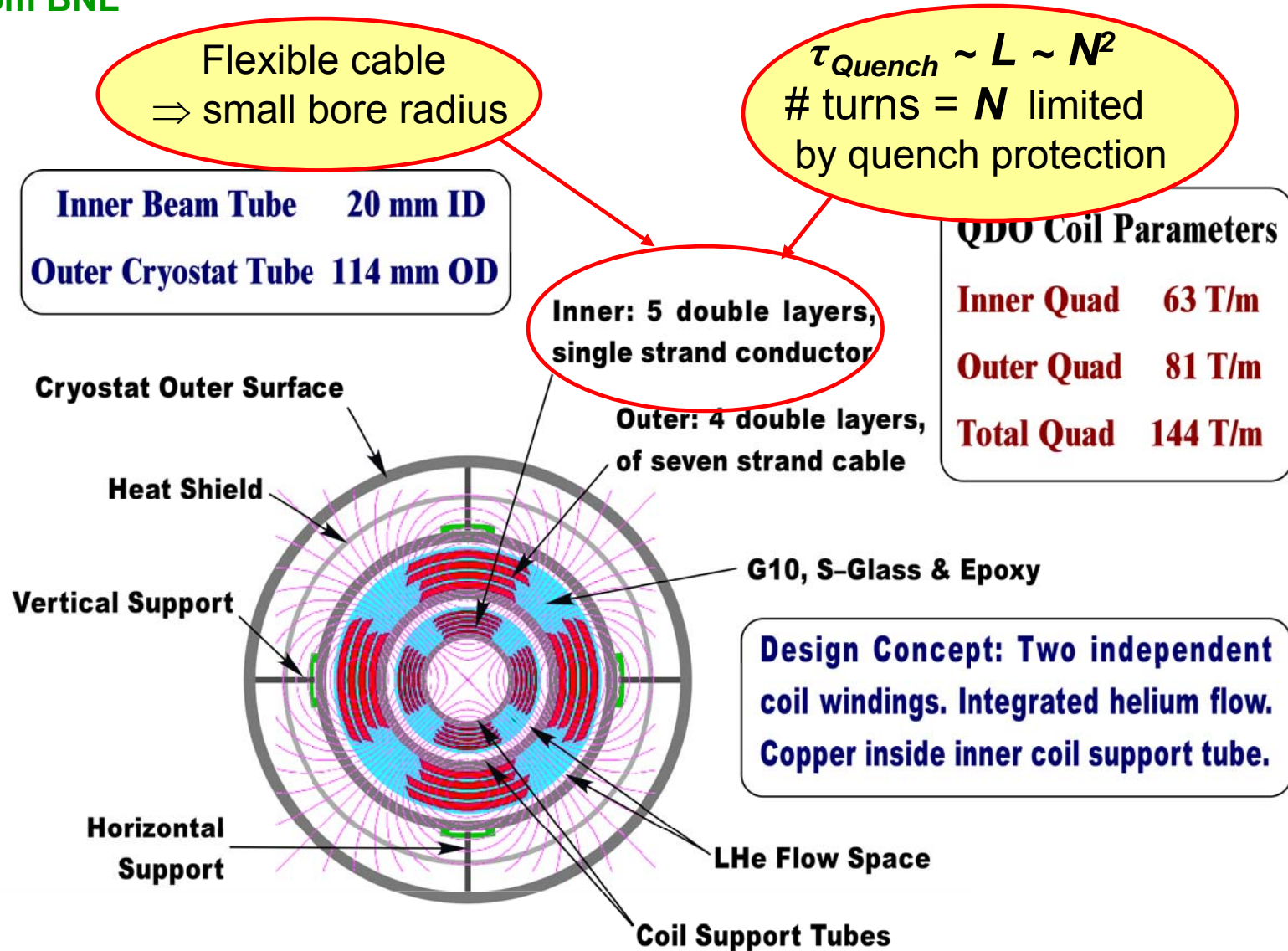
# 1 - IR design and X-ing angle

- **NLC-GLC** : X-ing angle mandatory to avoid **parasitic crossings** at every 20 cm (1.4 ns) or 40 cm (2.8 ns)



# New Development : Compact SC Quad Design

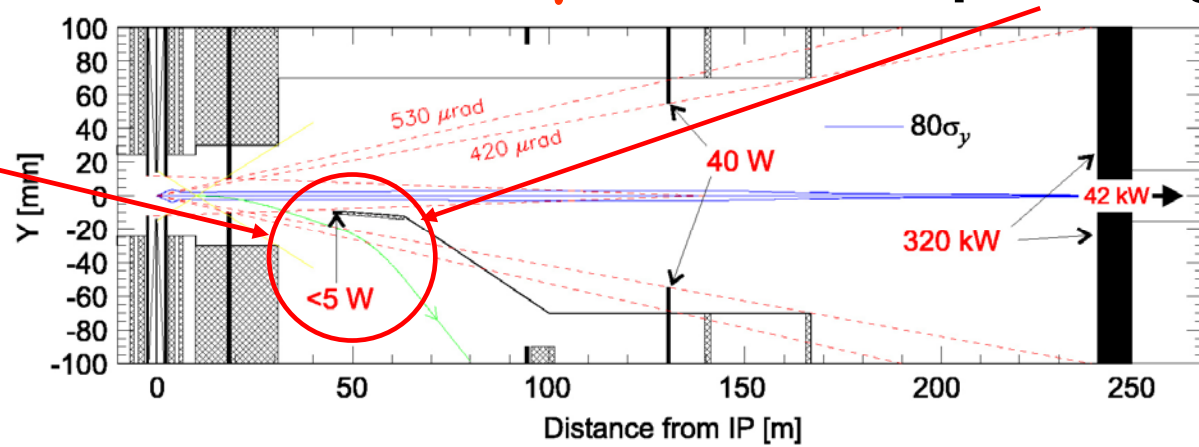
from BNL



# Crossing Angle Choices for TESLA

- **TESLA** : problem with **beam** and  $\gamma$  losses in septum region

~ 15 kW of beam and beamstrahlung loss in realistic conditions

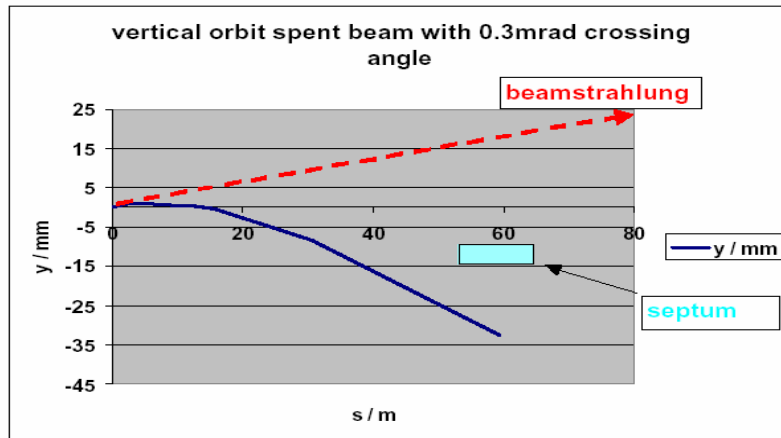


→ **3 options**

1. **300 μrad collision + quadruplet to reduce beam losses**

Suggestion: vertical crossing angle ~0.3mrad at IP

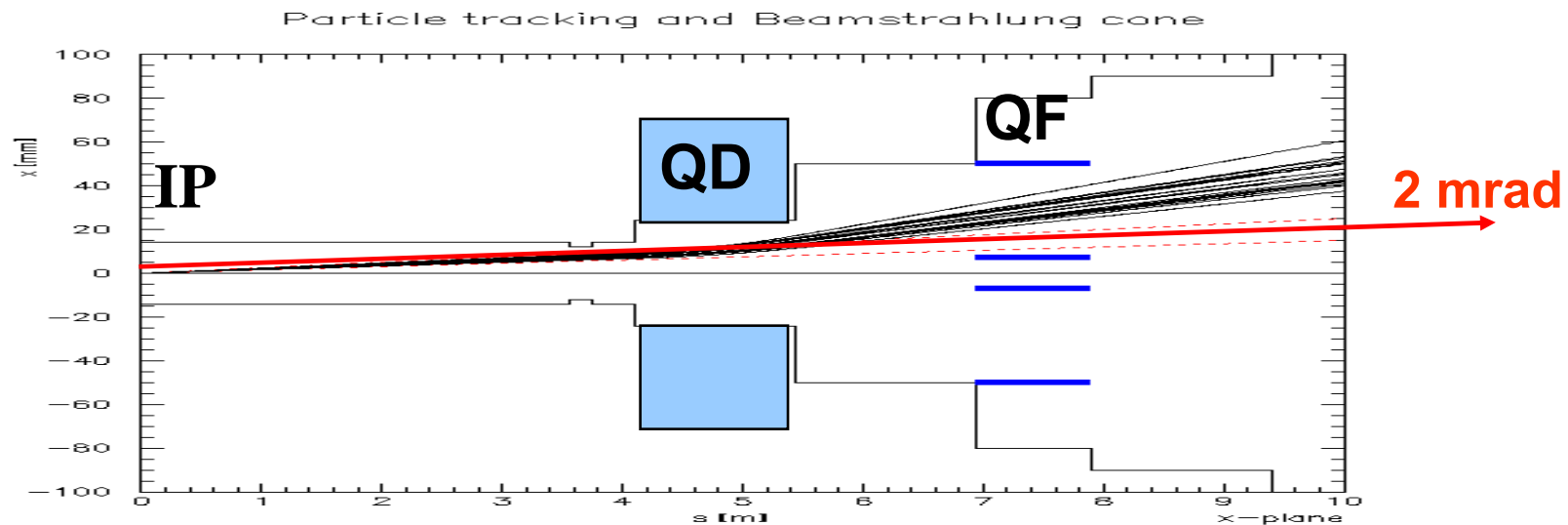
**D. Angal-Kalinin, R. Appleby, R. Brinkmann**



**Not a choice** if no progress on 50 kV/cm reliable electro-static separators (20-30 m long)

## Crossing Angle Choices for TESLA (cont.<sup>ed</sup>)

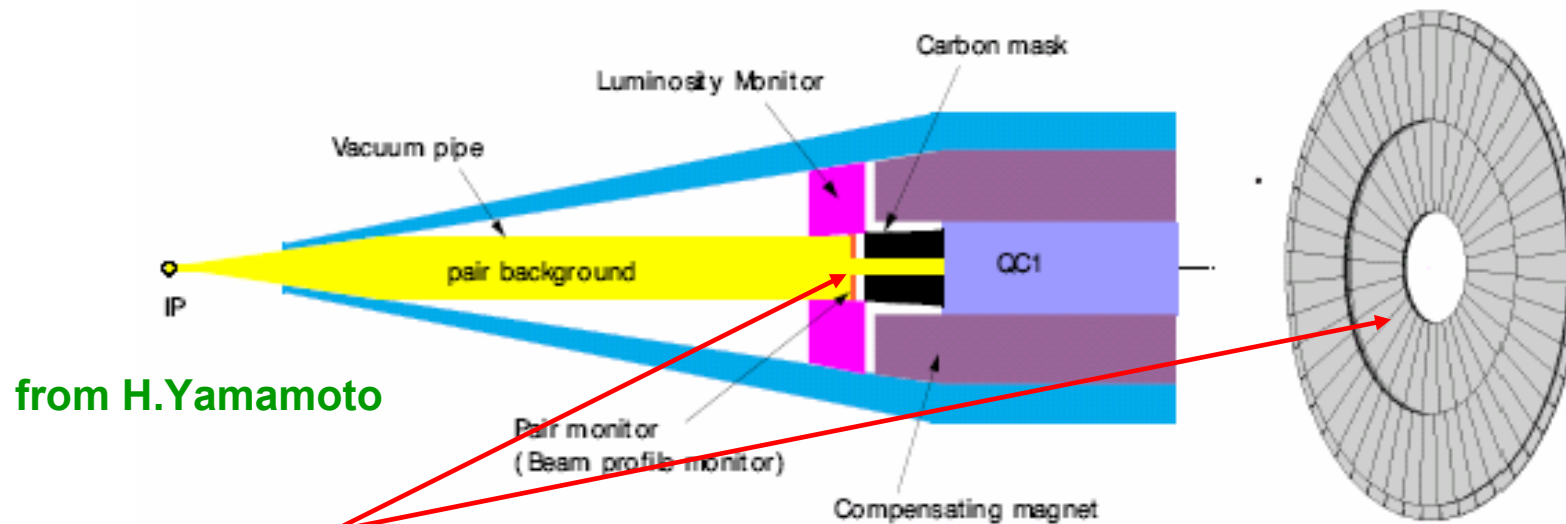
2. 2 mrad Xing angle : no electrostatic separators, 15% Lumi loss compensated by angular dispersion @ IP (~ crab-crossing)



3. 20 mrad Xing angle : NLC like final focus (cf. US-LC study), implies RF cavity crab crossing

⇒ Detector and Physics Implications (later)

## 2 - Very Forward Instrumentation

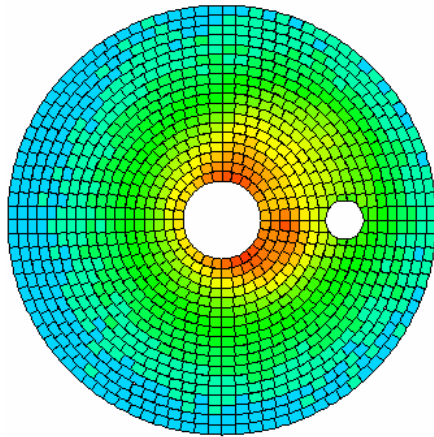


**Forward Calorimeter** has a dual function :

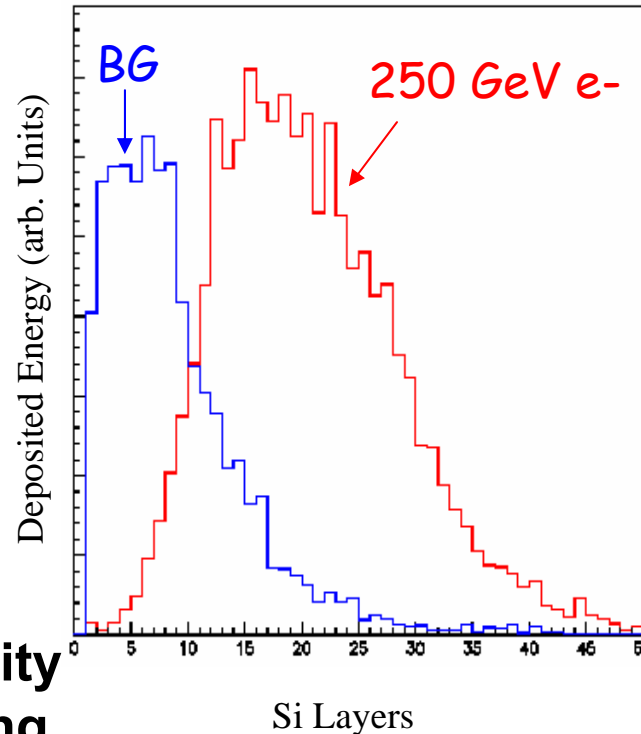
- **MACHINE** : Fast luminosity monitor, **keystone** of the collider (cf. SLC 1997 run). Can be used as a beam profile monitor
- **DETECTOR** : Low angle  $e\pm$  tagging : removes photon-photon background events

# Highly Segmented Detector

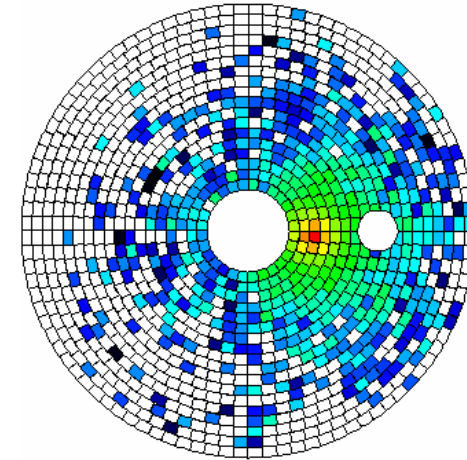
Pair Background



from N. Graf, T. Maruyama



250 GeV Electron



$N_{\text{tot}}$  or  $E_{\text{tot}} \propto \text{Luminosity}$   
 $\Rightarrow$  fast machine tuning

$\sigma_y$	$\Delta\sigma_y$
3.5 %	11 %

from A Stahl

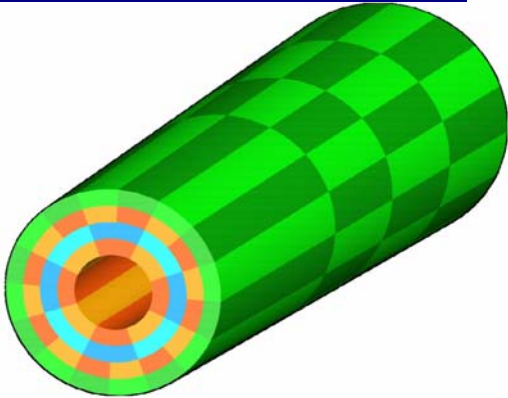
250 GeV Electron  
 detection efficiency  $\sim 1$



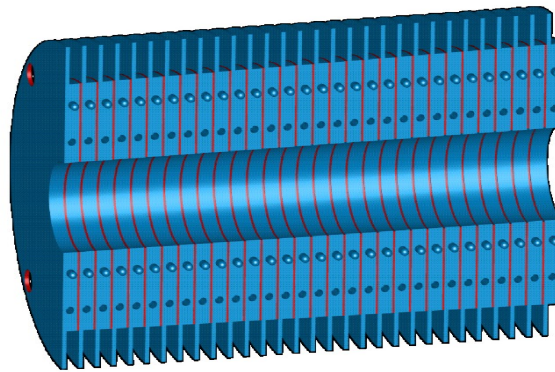
# Detector Design

from W. Lohmann

Heavy crystals

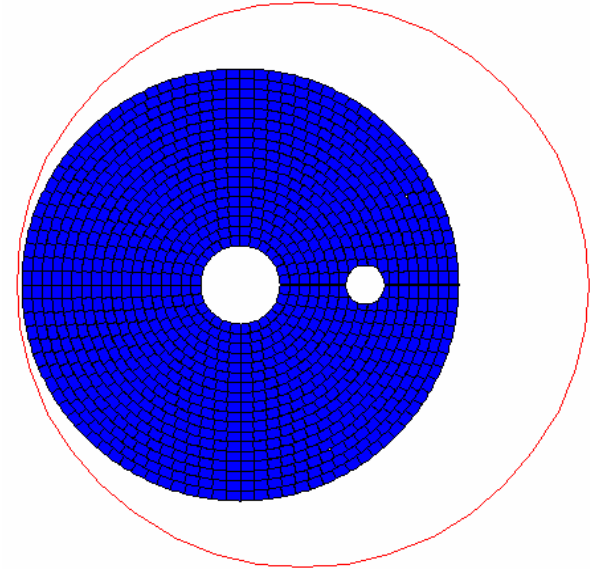


Diamond-W Sandwich

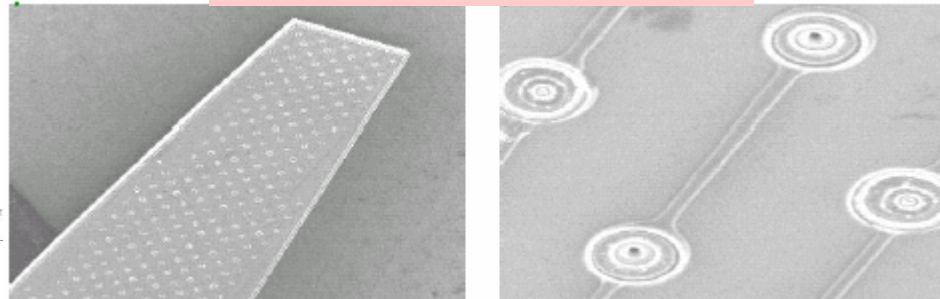
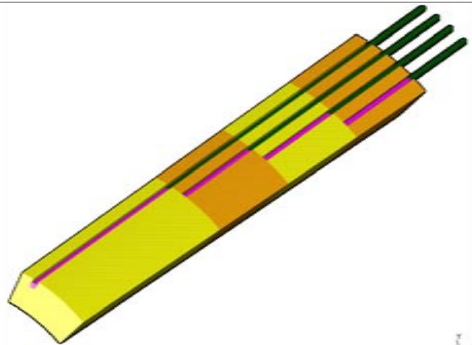


from N. Graf, T. Maruyama

Si W



3D Pixel Sensor



Technology  
Comparison  
made by  
J. Hauptman

from H. Yamamoto

# 3 - Machine and Beam-Beam Backgrounds

Impact on warm vs. cold technology choice ?  
via

- Crossing angle :  $\sim 0$  mrad vs. 7 mrad vs. 20 mrad
- Bunch Crossing : 1.4 ns vs. 2.8 ns vs 337 ns

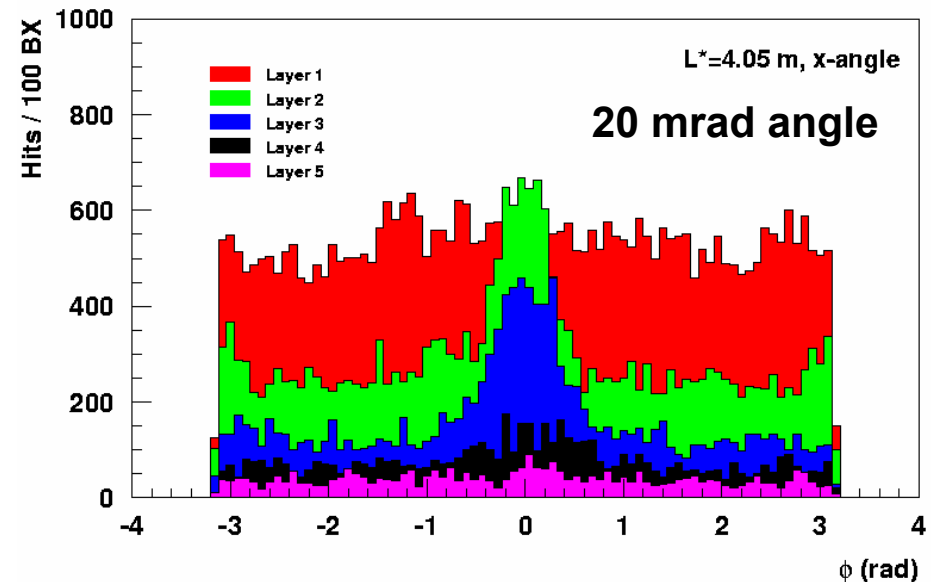
# Hit density on Vertex Detector from e<sup>+</sup>e<sup>-</sup> pairs : contradicting studies ?

Hit density on 1<sup>st</sup> Layer  
GLC study, from T. Aso

Crossing Angle	VTX Radius	Solenoid Field	Hit density /mm <sup>2</sup> /train
Head-on	15mm	4Tesla	0.99
7mrad	15mm	4Tesla	1.00
7mrad	24mm	3Tesla	0.38
20mrad	15mm	4Tesla	1.03
20mrad	15mm	3Tesla	1.71

Clear Dependence on VTX radius and B field.  
No dependence on crossing angle.

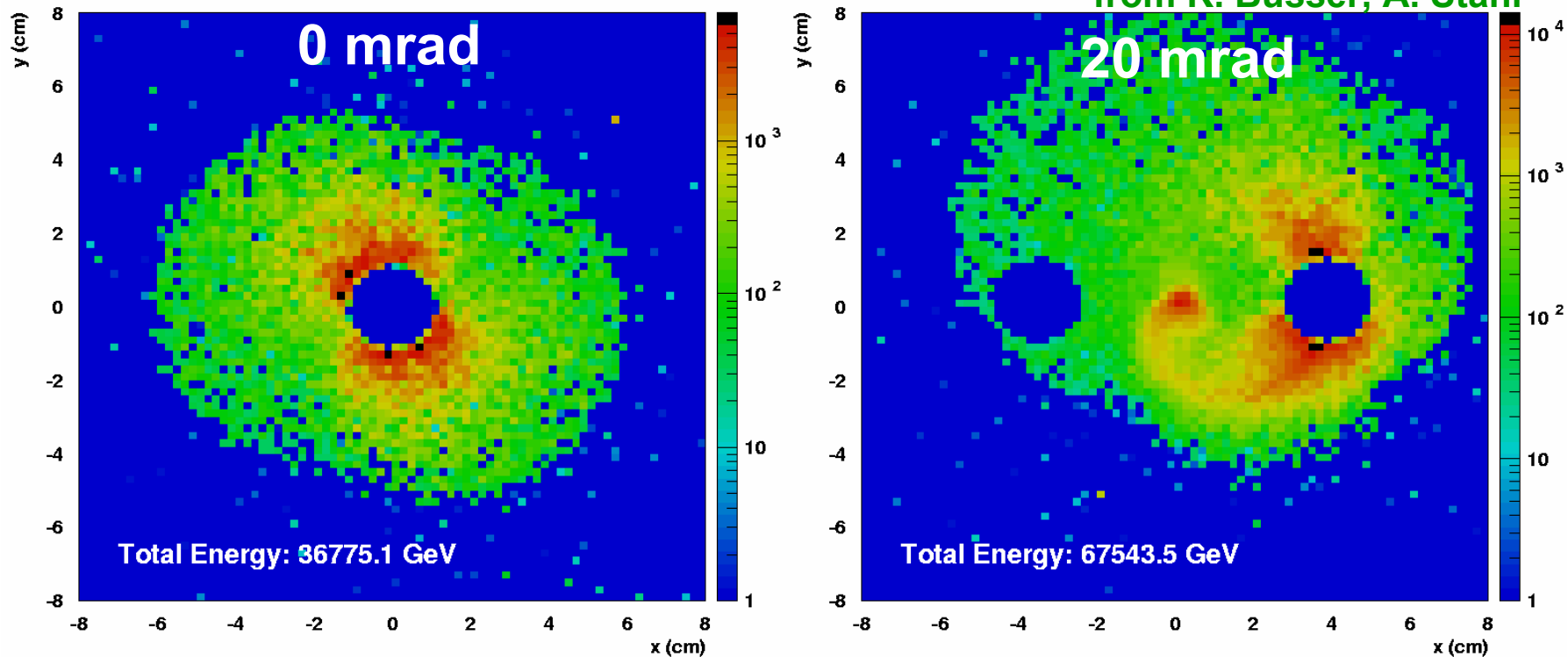
Azimuthal dependence  
TESLA study, from K. Büsser



TPC occupancy 2 times larger  
from 0 to 20 mrad crossing angle  
but  
All backgrounds so far studied are  
still on tolerable levels

# Forward Beam Calorimeter

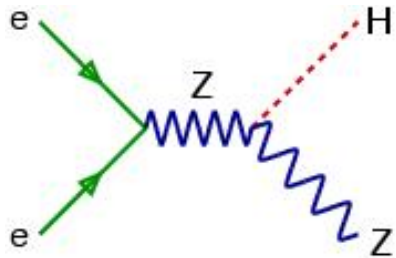
from K. Büsser, A. Stahl



- Energy deposition is twice as large in the 20 mrad x-angle case  
→ **impact on e – tagging efficiency**
- Origin of azimuthally asymmetric VTX pair hits

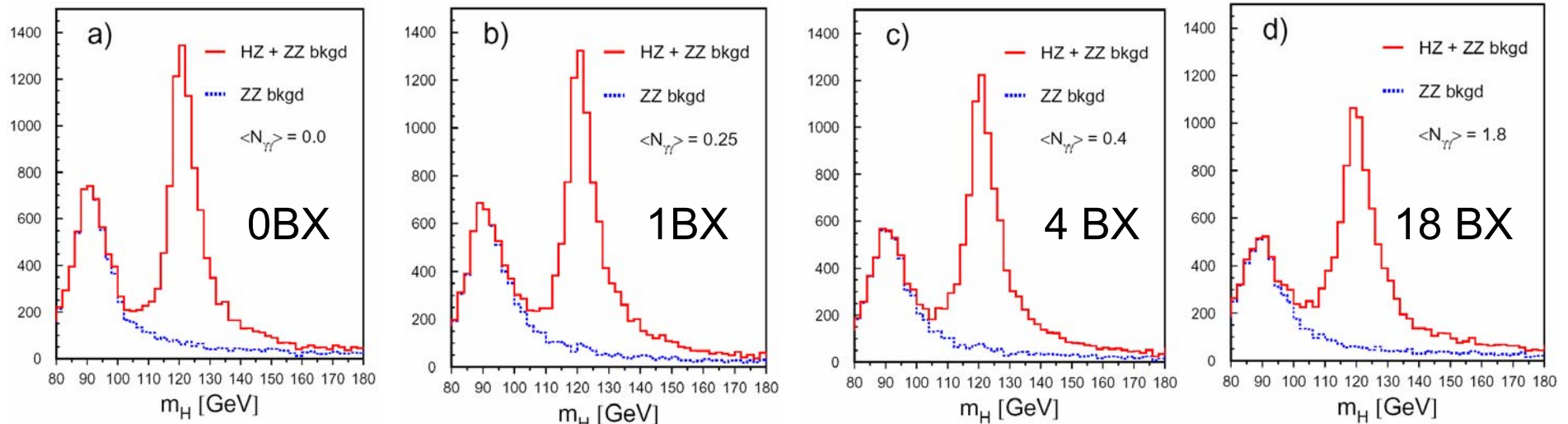
# Bunch Crossing : $\gamma\gamma \rightarrow$ hadrons background

from K. Desch



Mass measurement of light Higgs boson ( $m_H=120$  GeV)

$H \rightarrow bb, Z \rightarrow qq \Rightarrow$  4 jets reconstruction



*NLC, GLC and TESLA have about the same  $L / BX$*

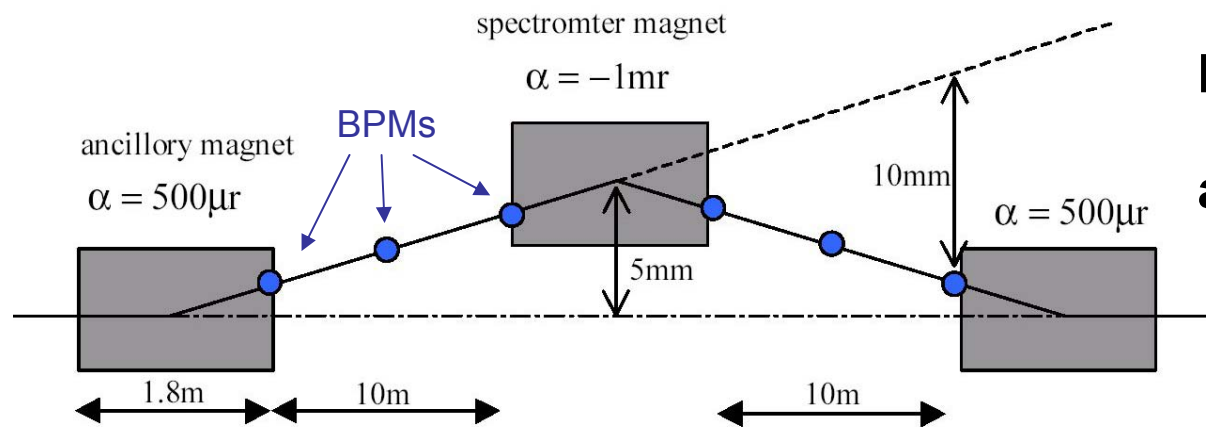
**Integrating over several BX hadronic backgrounds reduces the resolution on  $\Delta m_H$  from 75 MeV (1BX) to 92 MeV (18 BX)**

# 4 - Energy Precision

## Beam instrumentation goals

- Top mass: 200 ppm (35 MeV)
- Higgs mass: 200 ppm (25 MeV for 120 GeV Higgs)
- W mass: 50 ppm (4 MeV) ??
- ‘Giga’-Z  $A_{LR}$ : 200 ppm (20 MeV) (comparable to ~0.25% polarimetry)  
50 ppm (5 MeV) (for sub-0.1% polarimetry with e+ pol) ??

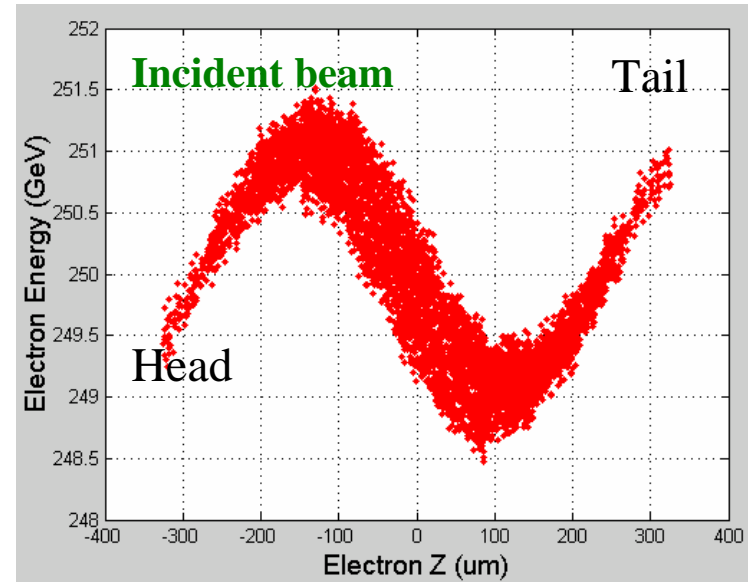
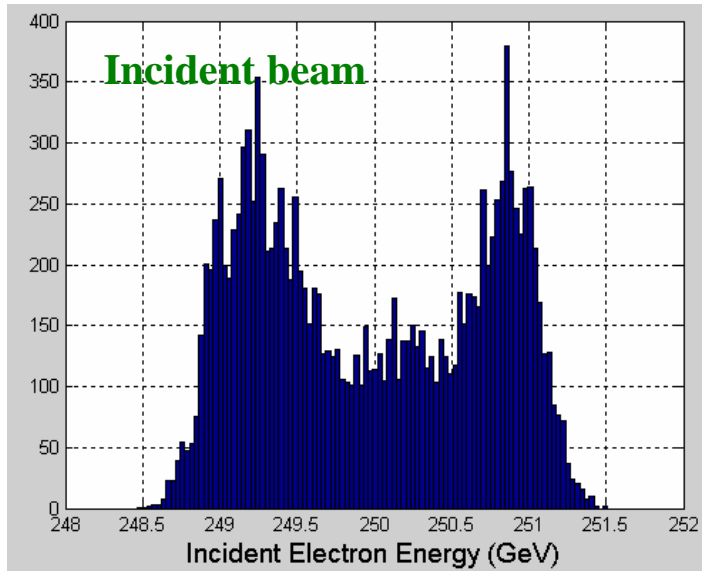
## Progress on Spectrometer measurement : $\langle E_{beam} \rangle$



**Progress at Zeuthen**  
 from H.J. Schreiber  
 and at SLAC + ...  
 from M. Hildreth

**BUT** : Machine bias to  $E_{CM}$  measurement :  $\langle \sqrt{s} \rangle_{lumi} \neq \langle E_{beam} \rangle$

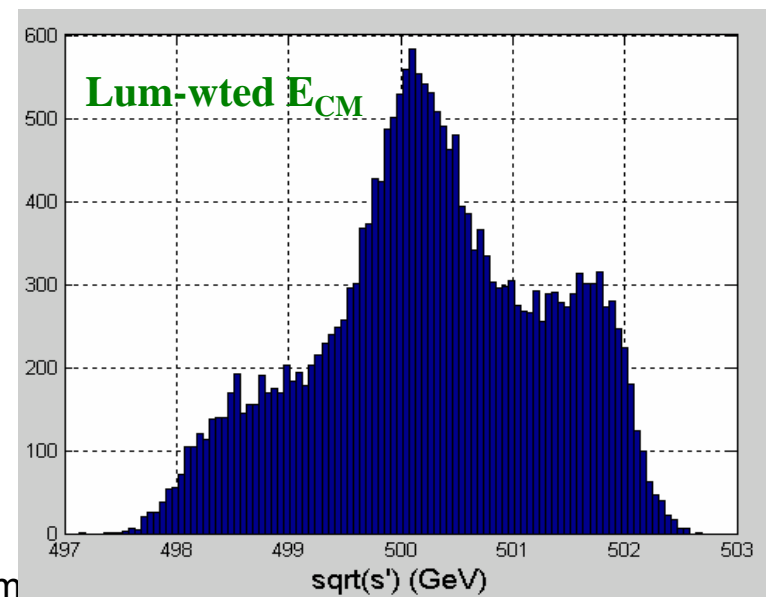
# Example of Lumi-weighted Energy Bias related to Beam Energy Spread at NLC-500 from T. Barklow, M. Woods



For energy bias study, **turn off beamstrahlung** and only consider beam energy spread.

$$\langle E_{CM}^{Bias} \rangle = \frac{\langle \sqrt{s'} \rangle - 500 \text{ GeV}}{500 \text{ GeV}} \approx 500 \text{ ppm}$$

Bhabha acolinearity analysis alone won't help resolve this bias.



# Kink instability and $E_{CM}$ Bias

from T. Barklow, M. Woods

(larger for NLC)                      (larger for TESLA)                      (comparable at NLC, TESLA)  
**Wakefields**                      +                      **Disruption**                       $\longrightarrow$                       **Kink instability**

(larger for NLC)                      (comparable at NLC, TESLA)                      (larger for NLC)  
**E-Spread + E-z correlation + Kink instability**                       $\longrightarrow$                        **$E_{CM}$  Bias**

$$E_{CM}^{Bias} = \frac{\langle E_1 \rangle + \langle E_2 \rangle - \langle E_{CM}^{lum-wt} \rangle}{\langle E_1 \rangle + \langle E_2 \rangle}, \quad E_1 \text{ and } E_2 \text{ are beam energies measured by the energy spectrometers}$$

## Summary of $E_{CM}^{bias}$

LC Machine Design	Collider Mode	$\langle E_{CM}^{bias} \rangle$ ( $\Delta y = 0$ )	$\sigma(E_{CM}^{bias})$ ( $\Delta y = 0$ )	Max( $E_{CM}^{bias}$ ) vary $\Delta y, \eta_y$
NLC-500	e <sup>+</sup> e <sup>-</sup>	+520 ppm	170 ppm	+1000 ppm
TESLA-500	e <sup>+</sup> e <sup>-</sup>	+50 ppm	30 ppm	+250 ppm



# Conclusions

- Many studies attempting to address the Cold / Warm technology comparison
- Trans-oceanic collaborations  
(even trans-Channel !!)
- Proposal to convene a specialized MDPI workshop in fall 2004