

Horizontal Crossing Angle

Necessary only at warm machines

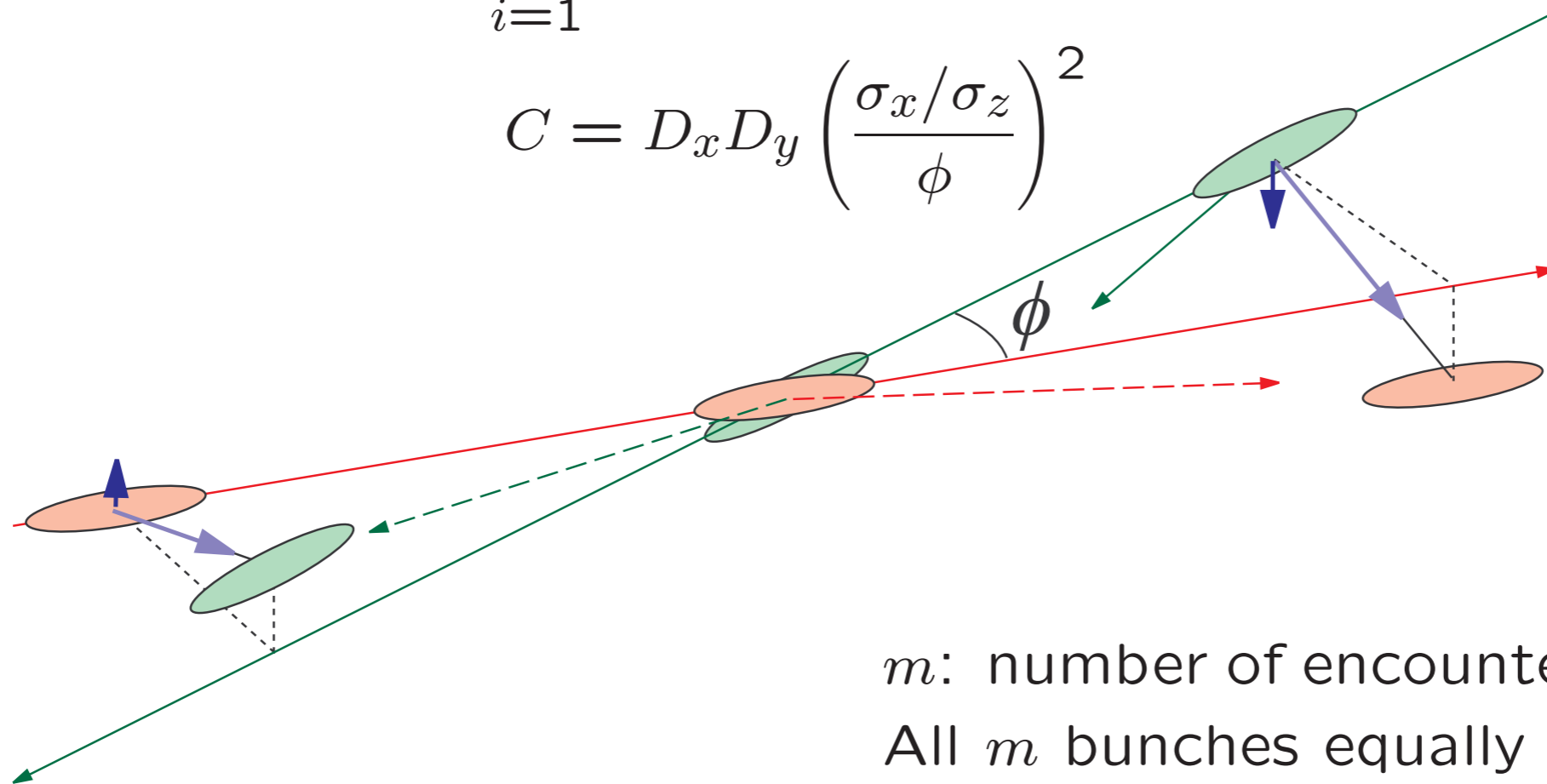
Multi-bunch Crossing Instability

by Yokoya

- $\Delta_k^\pm \equiv$ vertical offset of k -th e^\pm bunch at IP
- Evolution of $\Delta_k = \Delta_k^+ - \Delta_k^-$ is

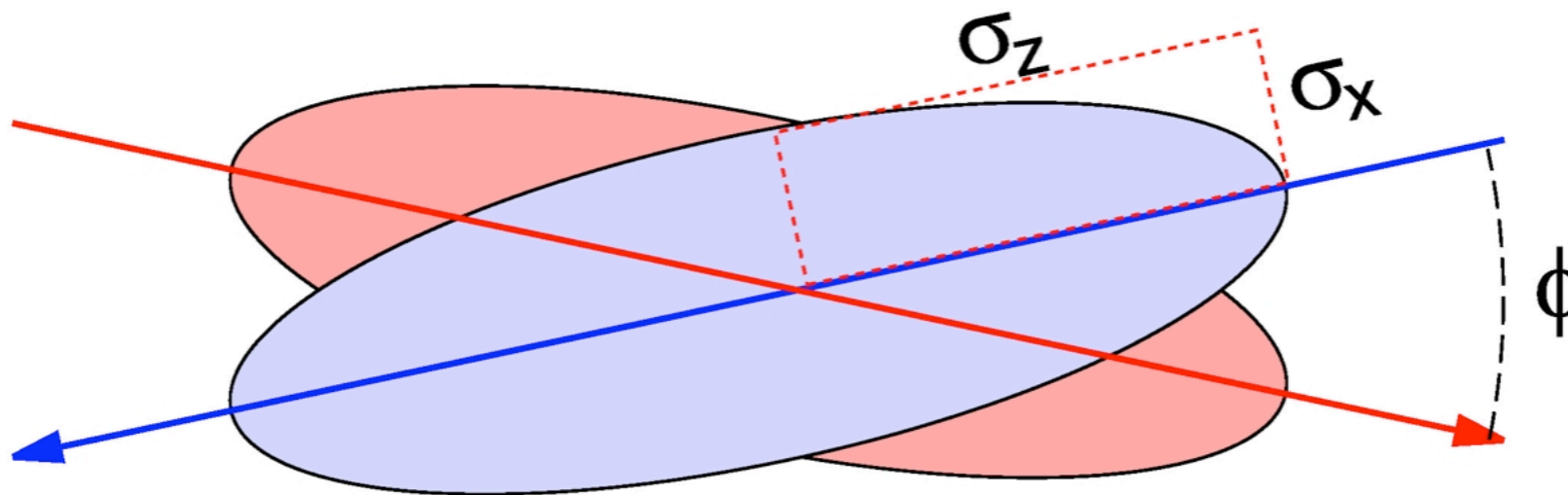
$$\Delta_k = C \sum_{i=1}^m F(\Delta_{k-i}) + \Delta_{k,0},$$

$$C = D_x D_y \left(\frac{\sigma_x / \sigma_z}{\phi} \right)^2$$



IR: Crossing Angle Issue

K.Yokoya



$$\frac{2\sigma_x}{\sigma_z} = 4.4 \times 10^{-3}$$

Small angle : $\phi \lesssim \frac{2\sigma_x}{\sigma_z}$,

Large angle : $\phi \gtrsim \frac{2\sigma_x}{\sigma_z}$

Why Small Crossing Angle ?

7 mrad

vs

Why Large Crossing Angle ?

20 mrad

- Detector $\cos\theta$ coverage
- Timing of crab cavity
- Radiation in the solenoid magnet

$$\sigma(\delta y) \propto \phi^{5/2} = 0.074\text{nm with } \phi = 20\text{mrad}$$

- Background to the detector
- Multi-bunch crossing instability
- Design of the final quadrupole magnet
- Layout of the beam dump

$$\frac{\Delta y}{\sigma_y} =$$

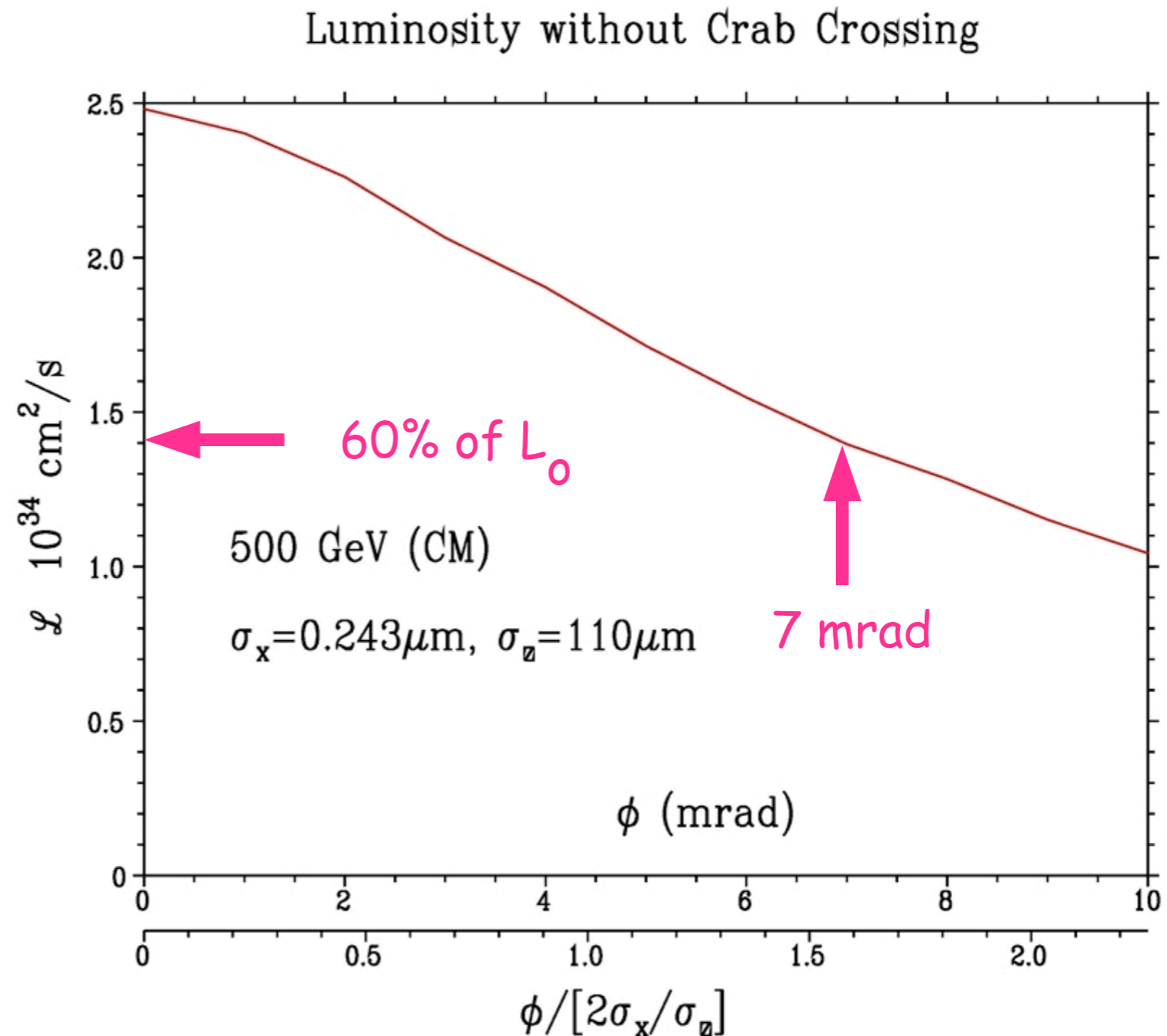
1.8 vs 0.6
at $L^*=3.5\text{m}$
($\Delta y^0=0.5\sigma_y$)

IR: 7 mrad Crossing Angle

K.Yokoya

- We decided $\phi \approx 7\text{mrad}$ many years ago
- Background tolerable
- Luminosity loss not too significant (σ_x was larger than today's value, σ_z was smaller)

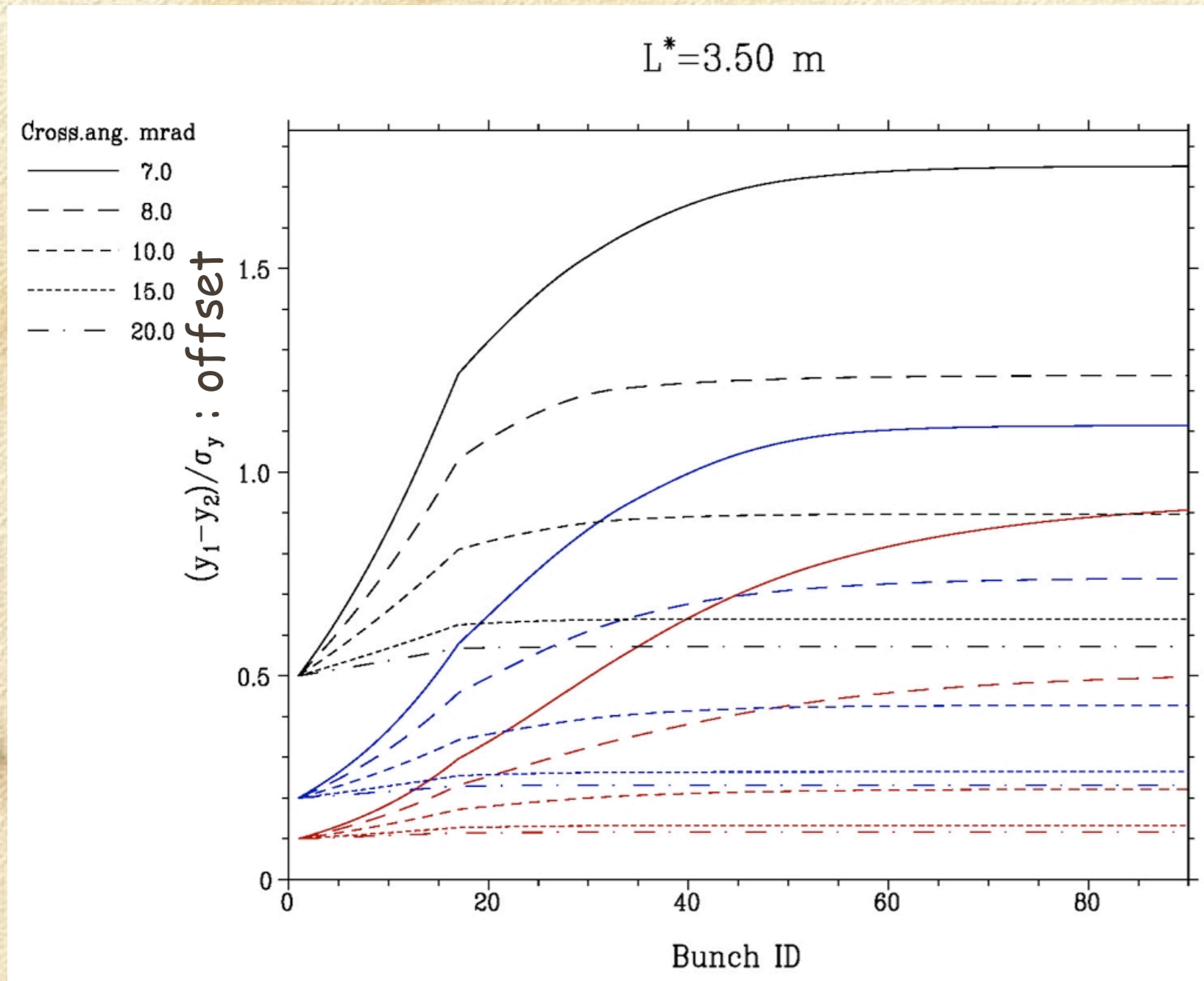
Luminosity vs. $\phi \Rightarrow$ by today's params.



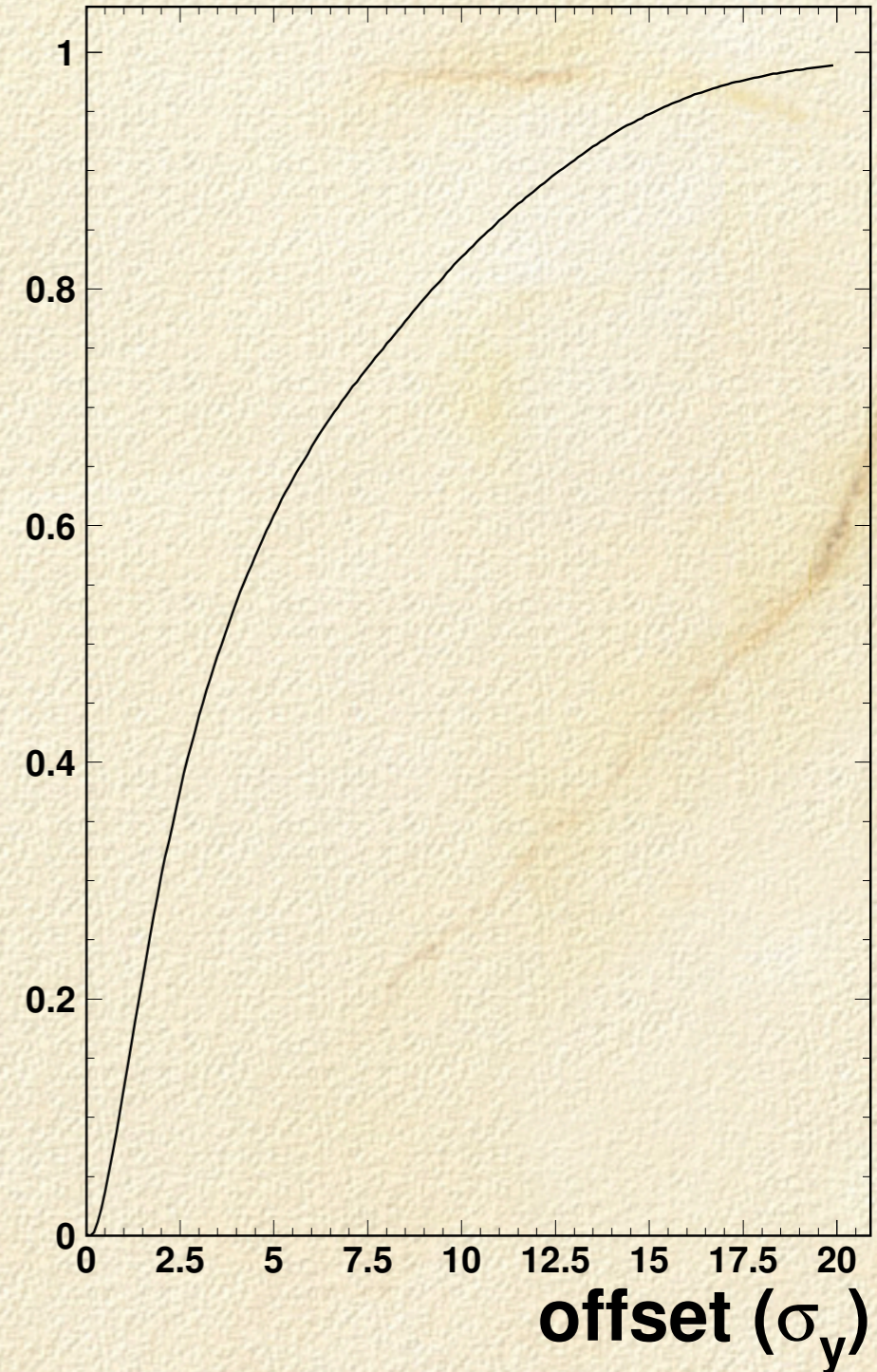
Multibunch Crossing Instability

Luminosity loss
due to offset

CAIN calculation by N.Delerue



Luminosity loss (L/L_0)



The offset must be corrected by Fast Feedback !