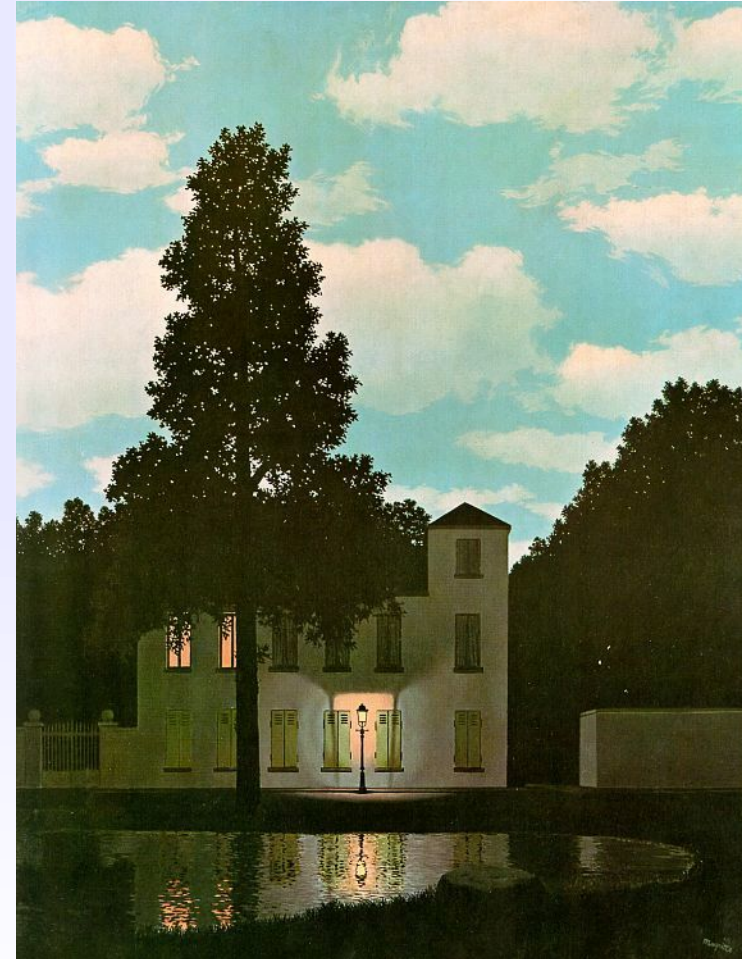


Particle Detectors – Principles and Techniques (3/5)

Lecture 3b – Photo-detection

Speaker: Thierry GYS (CERN-PH/DT2)



The Empire of Lights (René Magritte, Lessines 1898 – Brussels 1967)

(1954, Canvas, 146 x 114 cm, Brussels, Royal Museums of Fine Arts of Belgium, © SABAM 2001)



- **Lecture 1 - Introduction** C. Joram, L. Ropelewski
- **Lecture 2 - Tracking Detectors** L. Ropelewski, M. Moll
- **Lecture 3 - Scintillation and Photo-detection** C. D'Ambrosio, T. Gys

- **3a) Scintillation**

- **3b) Photo-detection**

Thierry Gys (CERN - PH/DT2)

- Photon detectors: purpose, basic principle and general requirements
- Vacuum photon detectors
- Solid-state photon detectors
- Hybrid photon detectors
- Literature

- **Lecture 4 - Calorimetry, Particle ID** C. Joram
- **Lecture 5 - Particle ID, Detector Systems** C. Joram, C. D'Ambrosio



Detailed outline

extra slide
not shown

3b Photo-detection

■ Photon detectors

- Purpose, basic principle and general requirements

■ Vacuum photon detectors

- The photoelectric effect, photo-cathodes and optical windows
- Photomultipliers:
 - Basic principle and gain fluctuations
 - Dynode configurations: traditional and position-sensitive
- Image intensifiers: principles, generations and Micro Channel Plates

■ Solid-state photon detectors

- Basic principle, PIN and avalanche diodes, light absorption
- A detailed example of CCD optimization for astronomy

■ Hybrid photon detectors

- Basic principle and gain fluctuations
- Description of various HPD types

■ Literature



Purpose:

- Convert light into detectable (electronic) signal

Principle:

- Use photoelectric effect to convert photons (γ) to photoelectrons (pe)

Standard requirements:

- High sensitivity, usually expressed as:

- quantum efficiency: $QE(\%) = \frac{N_{pe}}{N_{\gamma}}$

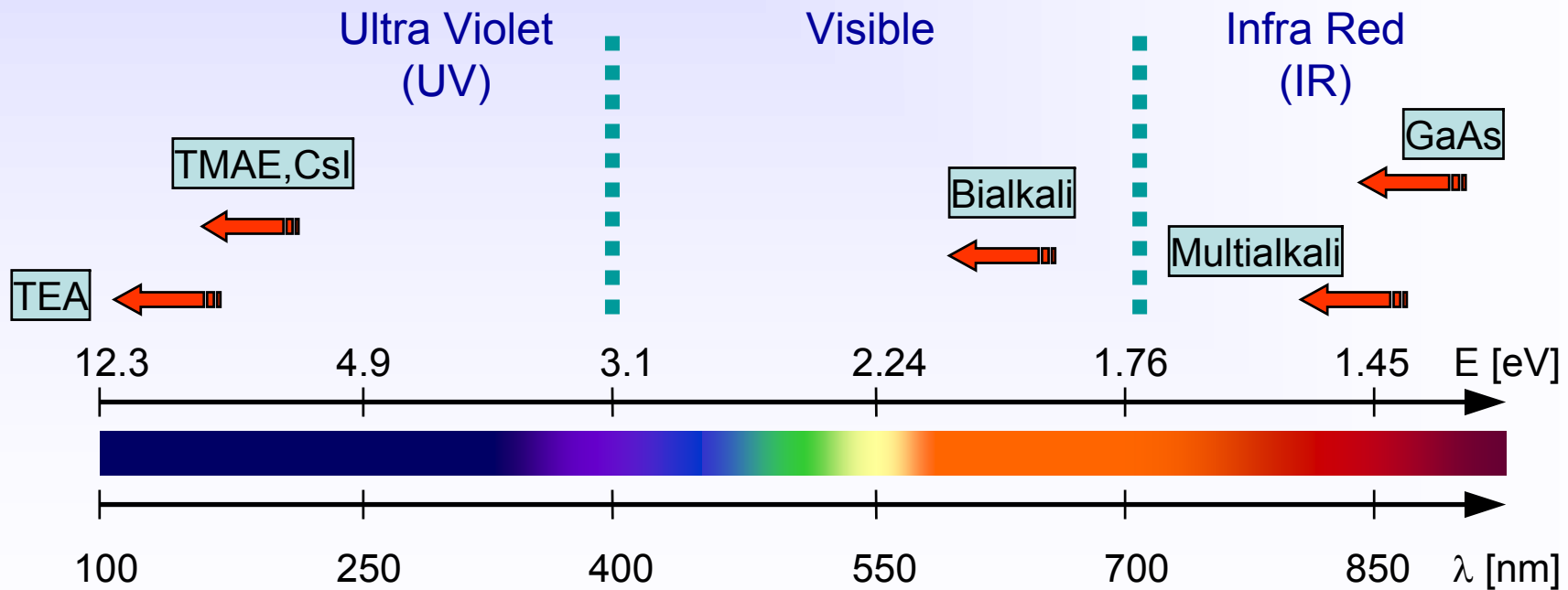
- radiant sensitivity $S(mA/W)$ with: $QE(\%) \approx 124 \cdot \frac{S(mA/W)}{\lambda(nm)}$

- Low intrinsic noise
- Low gain fluctuations
- High active area

Main types of photon detectors:

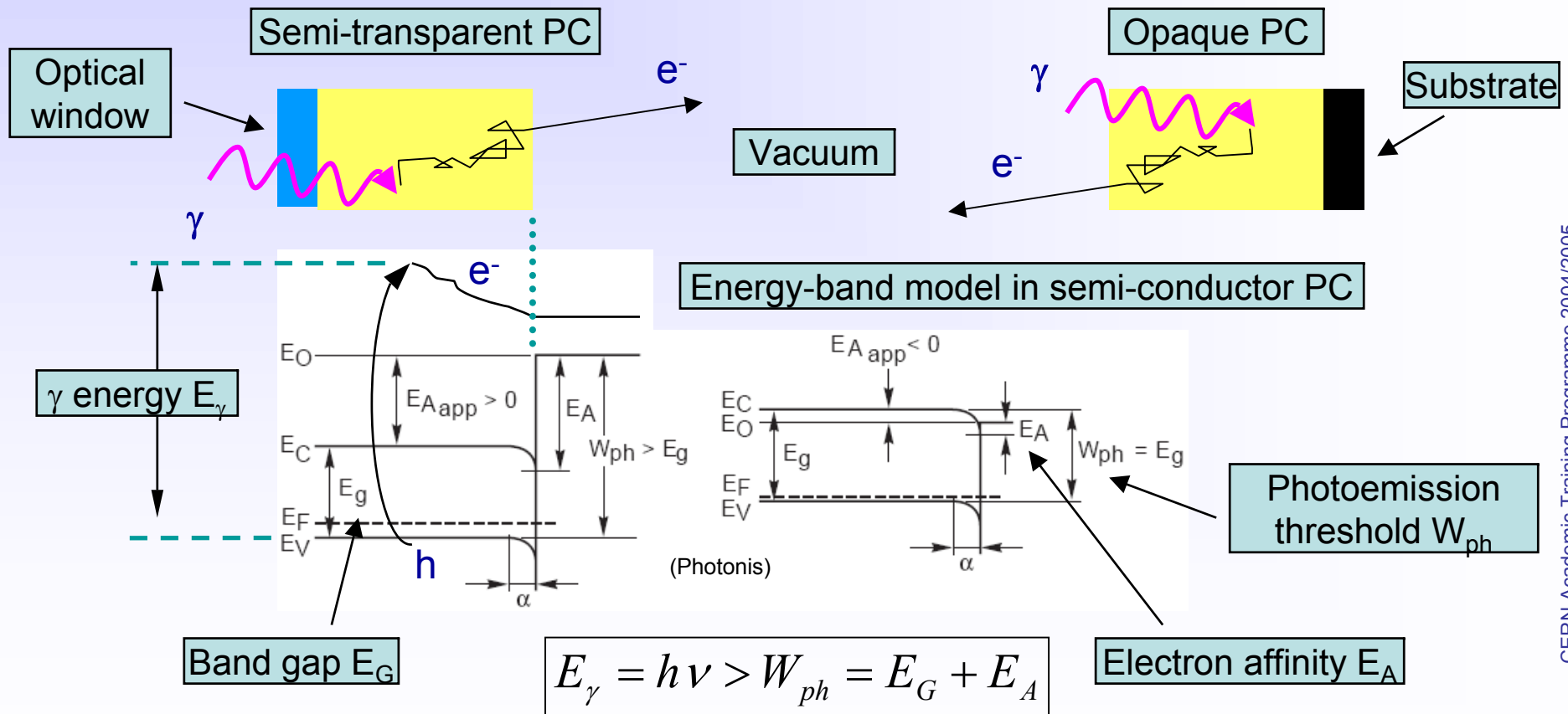
- gas-based (not covered in this lecture, see lecture 2a)
- vacuum-based
- solid-state (see also lecture 2b)
- hybrid

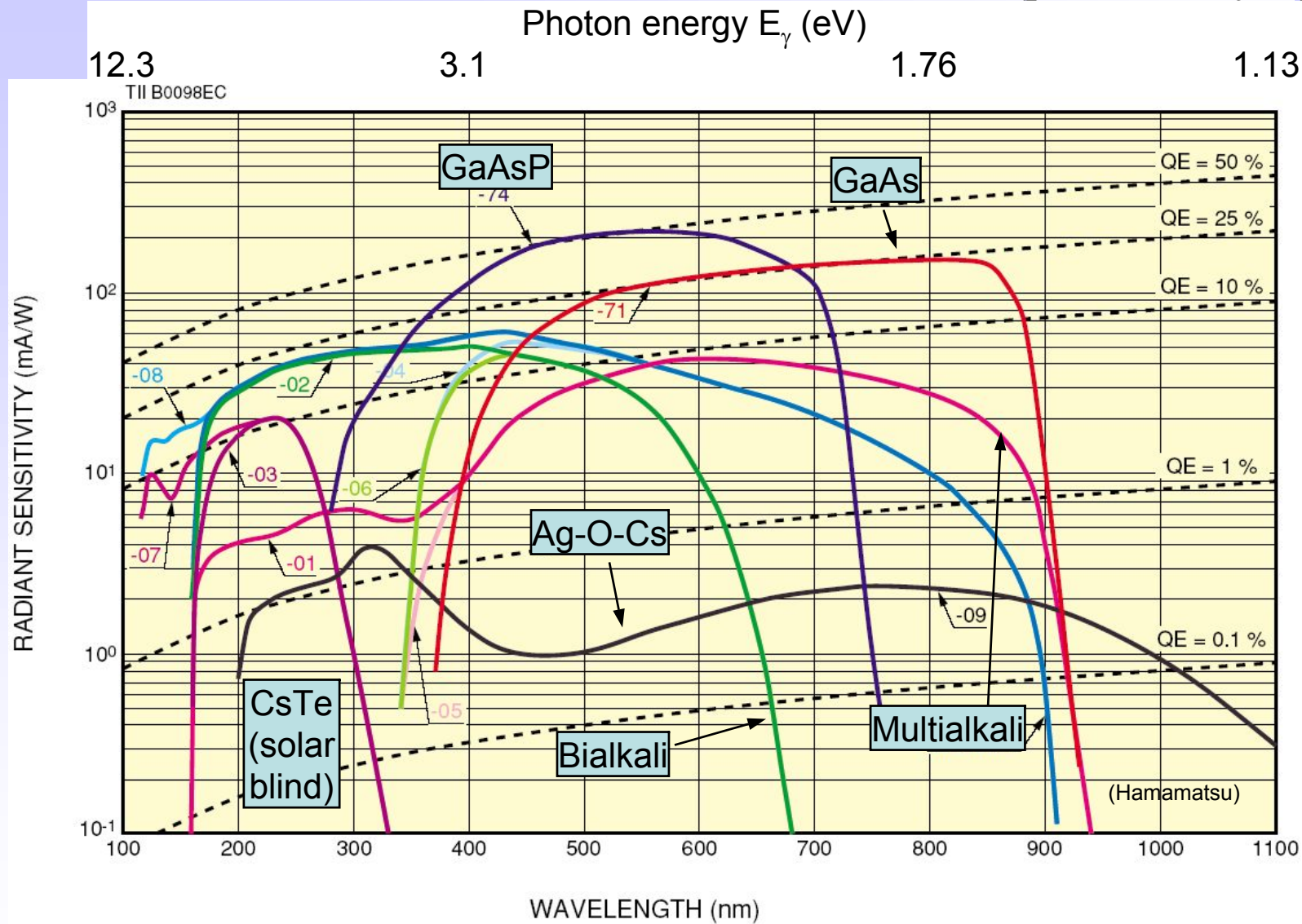
Photoemission threshold W_{ph} of various materials



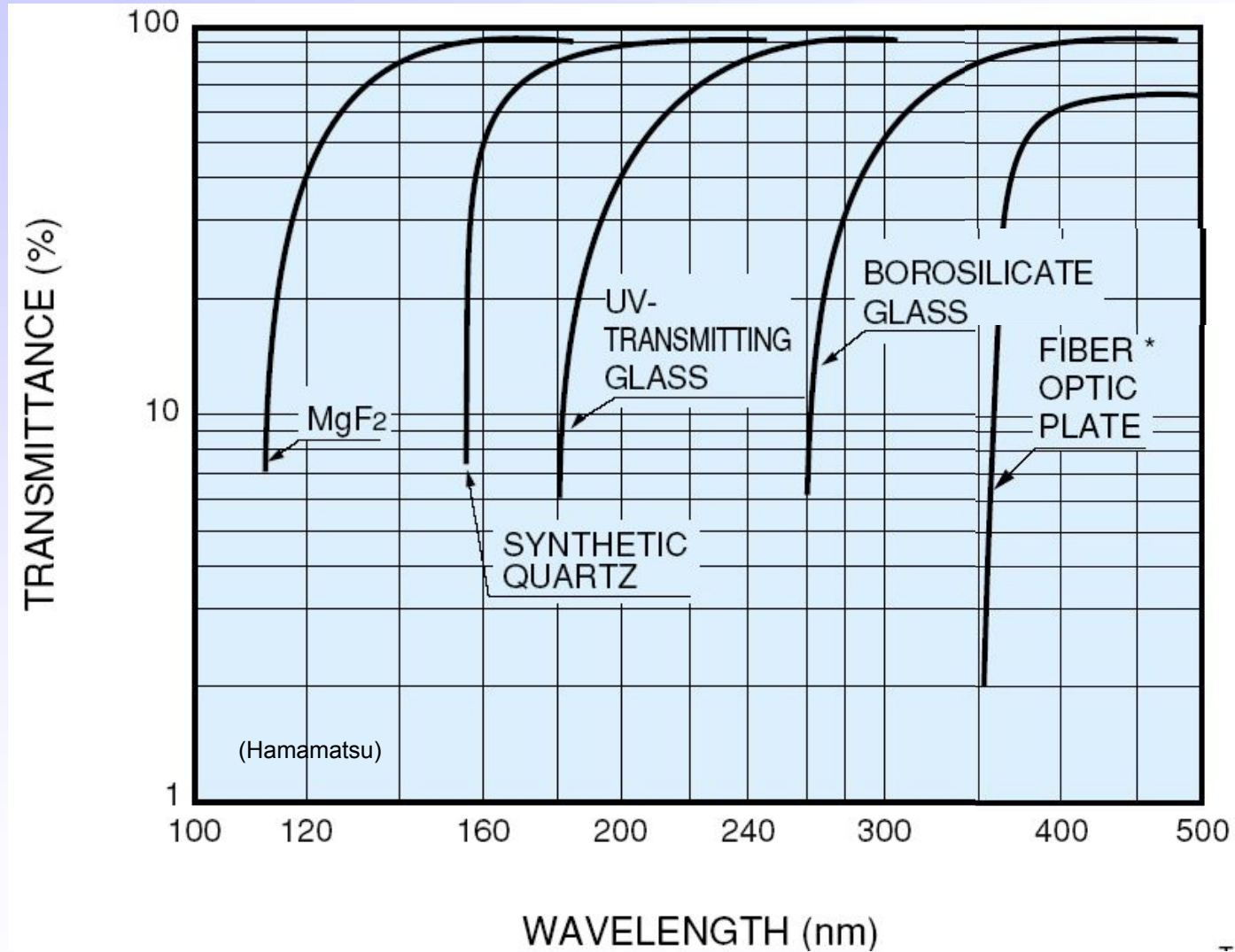
3-step process:

- absorbed γ 's impart energy to electrons (e) in the material;
 - energized e's diffuse through the material, losing part of their energy;
 - e's reaching the surface with sufficient excess energy escape from it;
- ⇒ ideal photo-cathode (PC) must absorb all γ 's and emit all created e's





Bialkali: SbKCs, SbRbCs **Multialkali:** SbNa₂KCs (alkali metals have low work function)

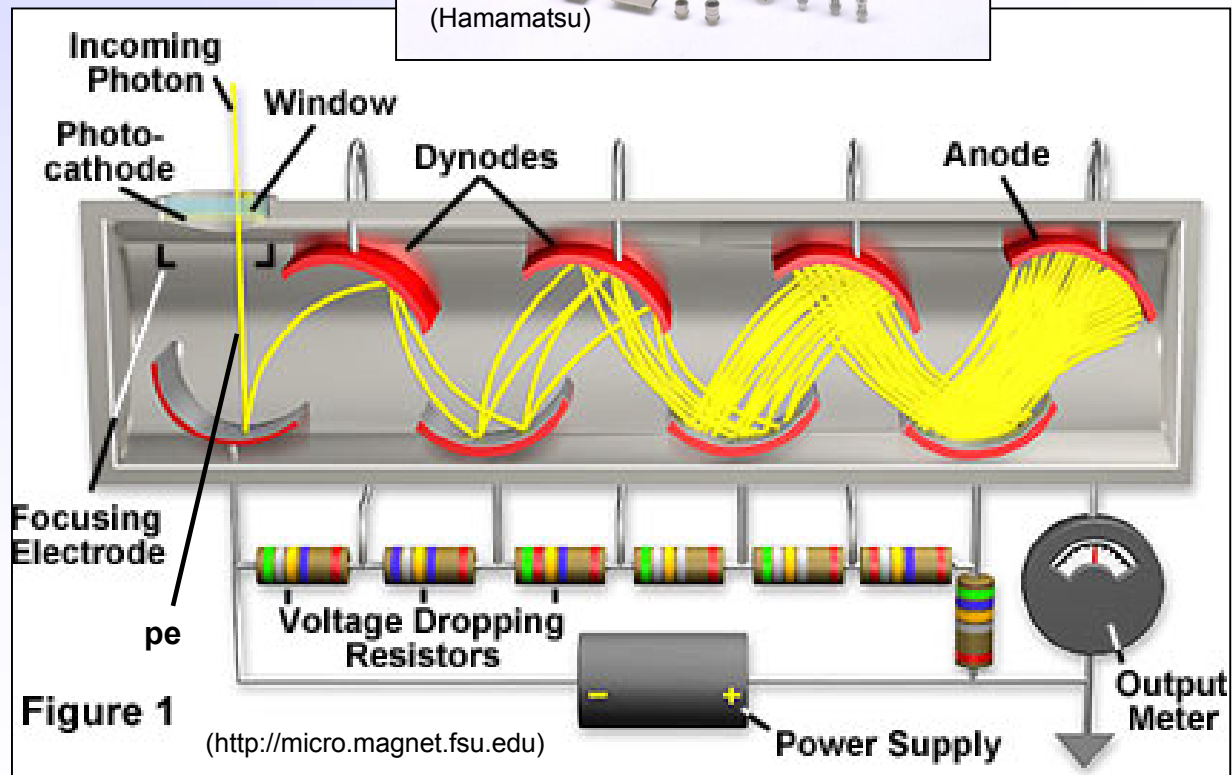


Basic principle:

- Photo-emission from photo-cathode
- Secondary emission (SE) from N dynodes:
 - dynode gain $g \approx 3-50$ (function of incoming electron energy E);
 - total gain M :

$$M = \prod_{i=1}^N g_i$$

- Example:
 - 10 dynodes with $g=4$
 - $M = 4^{10} \approx 10^6$





Gain fluctuations of PMT's

- Mainly determined by the fluctuations of the number $m(\delta)$ of secondary e's emitted from the dynodes;

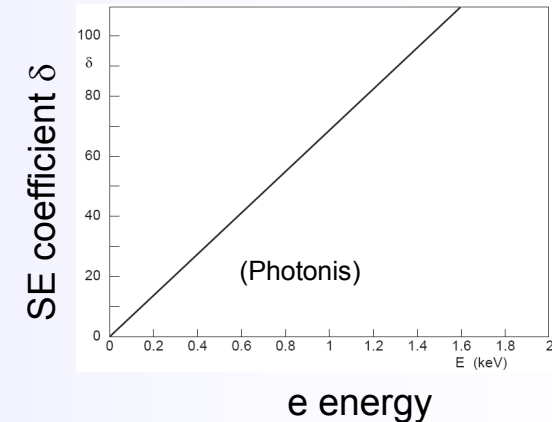
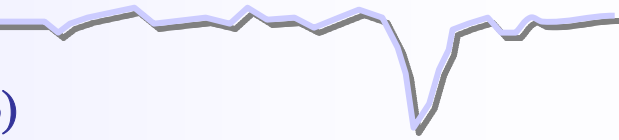
- Poisson distribution:

$$P_{\delta}(m) = \frac{\delta^m e^{-\delta}}{m!}$$

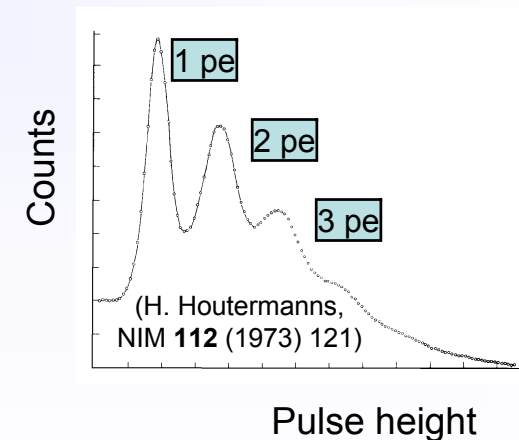
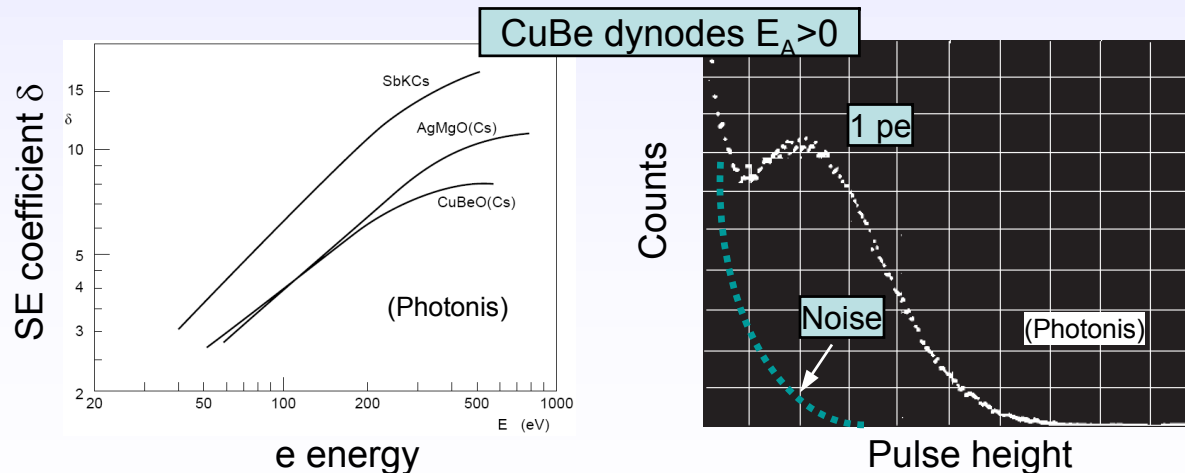
- Standard deviation:

$$\frac{\sigma_m}{\delta} = \frac{\sqrt{\delta}}{\delta} = \frac{1}{\sqrt{\delta}}$$

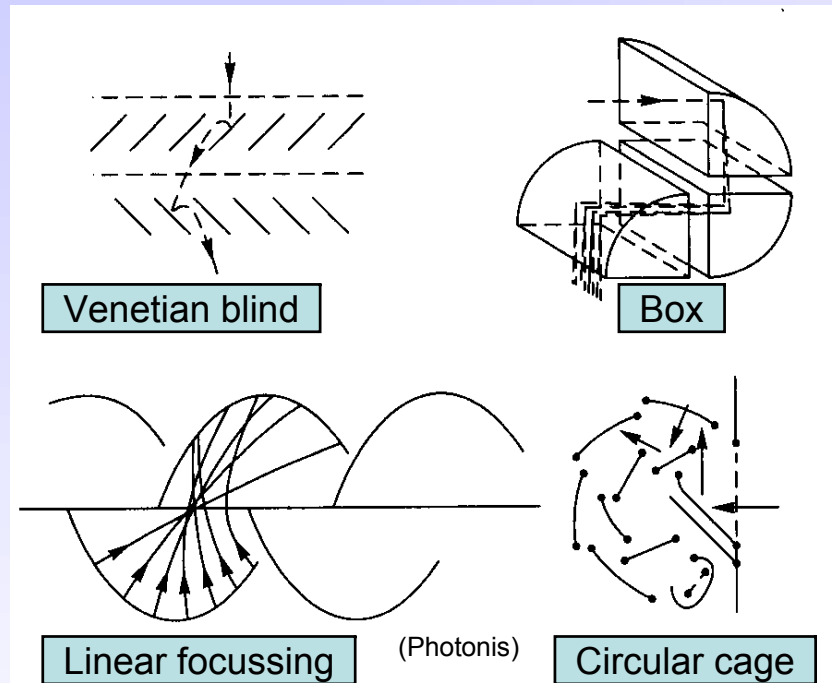
⇒ fluctuations dominated by 1st dynode gain;



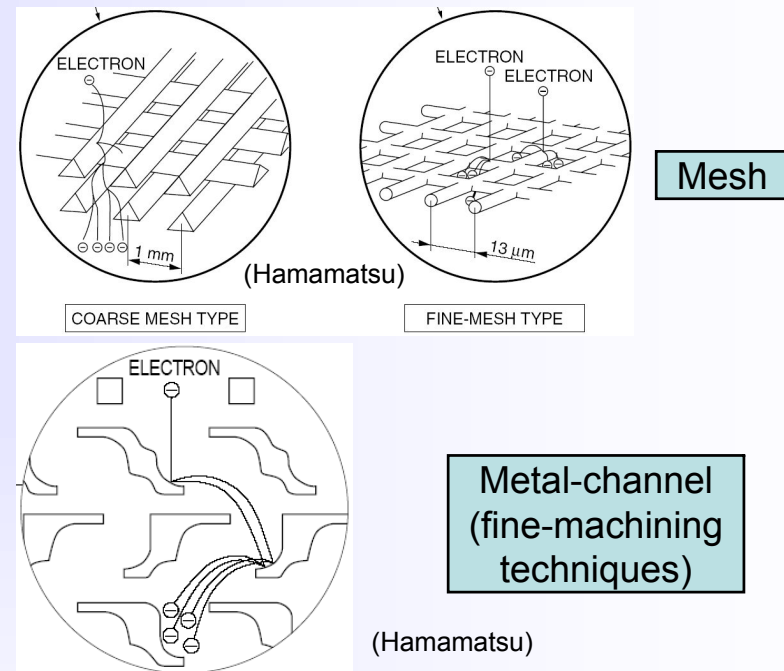
GaP(Cs) dynodes $E_A < 0$



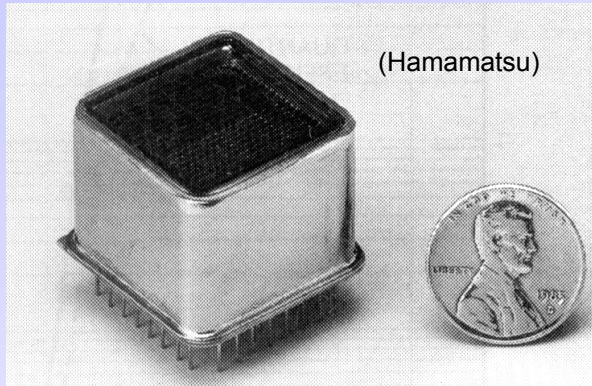
Traditional



Position-sensitive



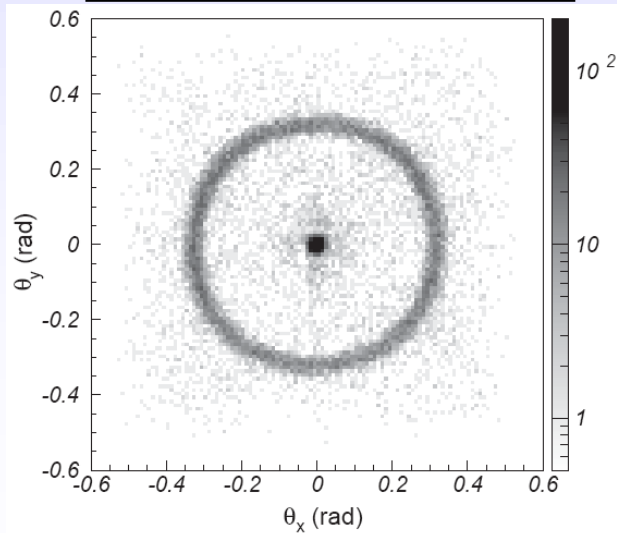
- “Fast” PMT’s require well-designed input electron optics to limit (e) chromatic and geometric aberrations \rightarrow transit time spread < 200 ps;
- PMT’s are in general very sensitive to magnetic fields, even to earth field (30-60 μT). Magnetic shielding required.



Multi-anode (Hamamatsu H7546)

- Up to 8×8 channels ($2 \times 2 \text{ mm}^2$ each);
- Size: $28 \times 28 \text{ mm}^2$;
- Active area $18.1 \times 18.1 \text{ mm}^2$ (41%);
- Bialkali PC: $QE \approx 20\% @ \lambda_{\text{max}} = 400 \text{ nm}$;
- Gain $\approx 3 \cdot 10^5$;
- Gain uniformity typ. 1 : 2.5;
- Cross-talk typ. 2%

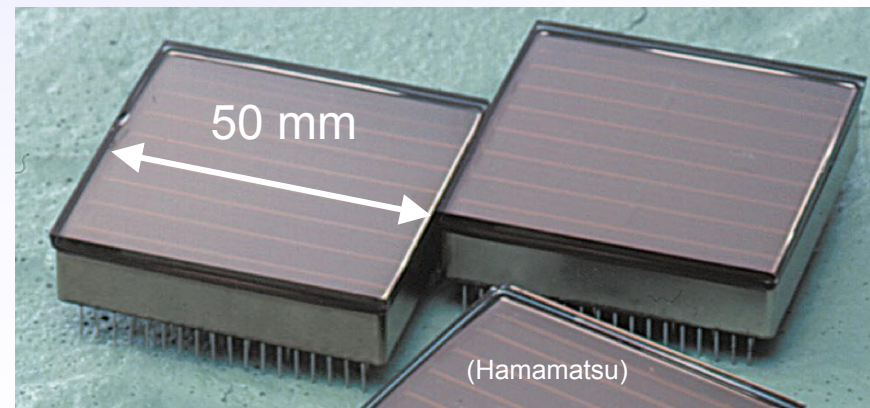
Cherenkov rings from
3 GeV/c π^- through aerogel



(T. Matsumoto et al., NIMA **521** (2004) 367)

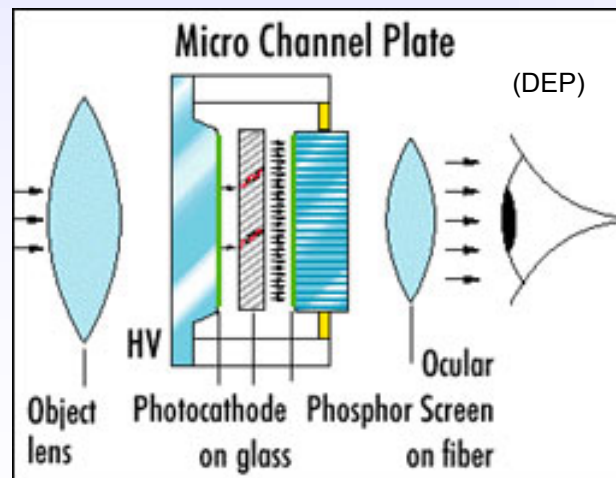
Flat-panel (Hamamatsu H8500):

- 8 x 8 channels ($5.8 \times 5.8 \text{ mm}^2$ each);
- Excellent surface coverage (89%)



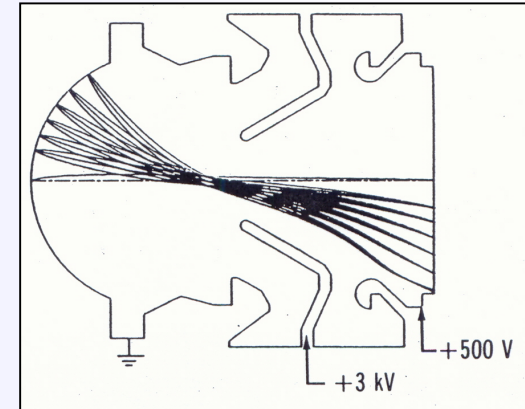
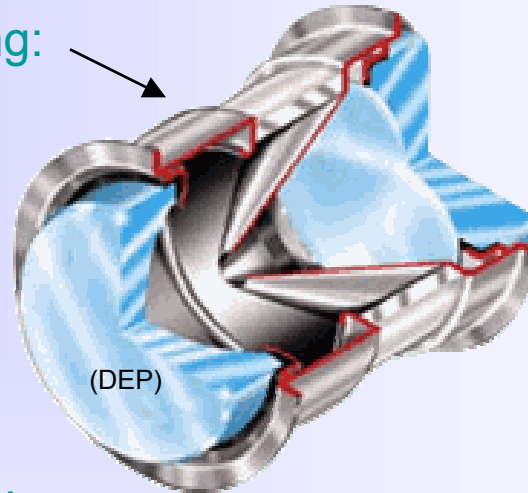
Basic principle:

- Vacuum photon detectors amplifying low light-level *image* to observable levels;
- Input: collection lens, optical window, photo-cathode;
- Gain: achieved by high voltage and possibly by additional imaging electron multiplier;
- Output: phosphor on optical window, ocular, observer (eye, CCD)



Gen. I - electrostatic focussing:

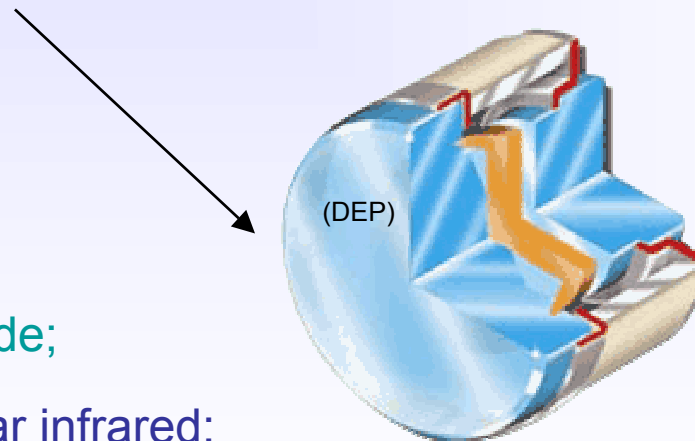
- high image resolution;
- wide dynamic range;
- low noise;



(I. P. Csorba, Image Tubes, Sams (1985))

Gen. II - Micro Channel Plate:

- worse resolution;
- much higher gain;



Gen. III – GaAs photo-cathode;

- enhanced sensitivity in near infrared;

extra slide
not shown

Principle:

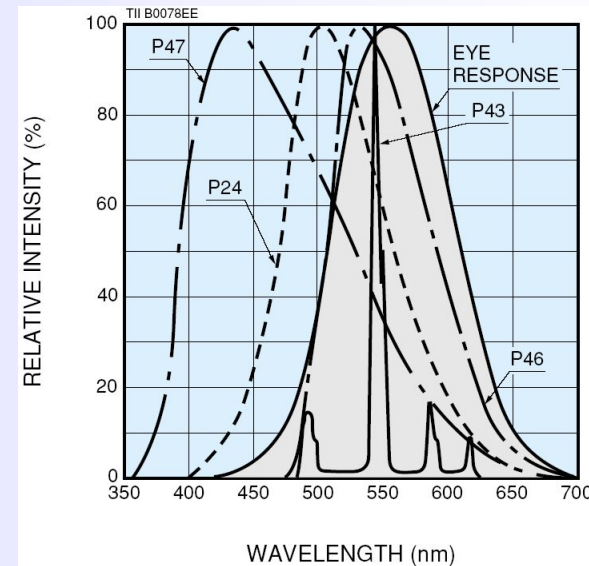
- absorb electrons;
- emit light on a characteristic λ of their material;

Spectral response:

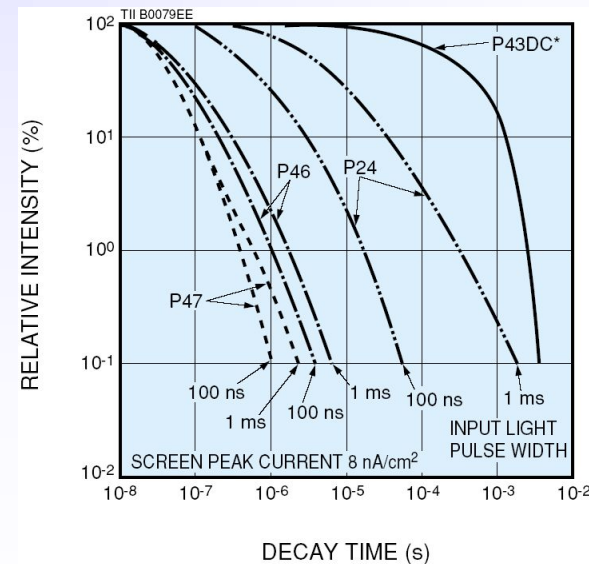
- originally adapted to human eye response;
- must now match solid-state sensor response (e.g. CCD's);

Decay time:

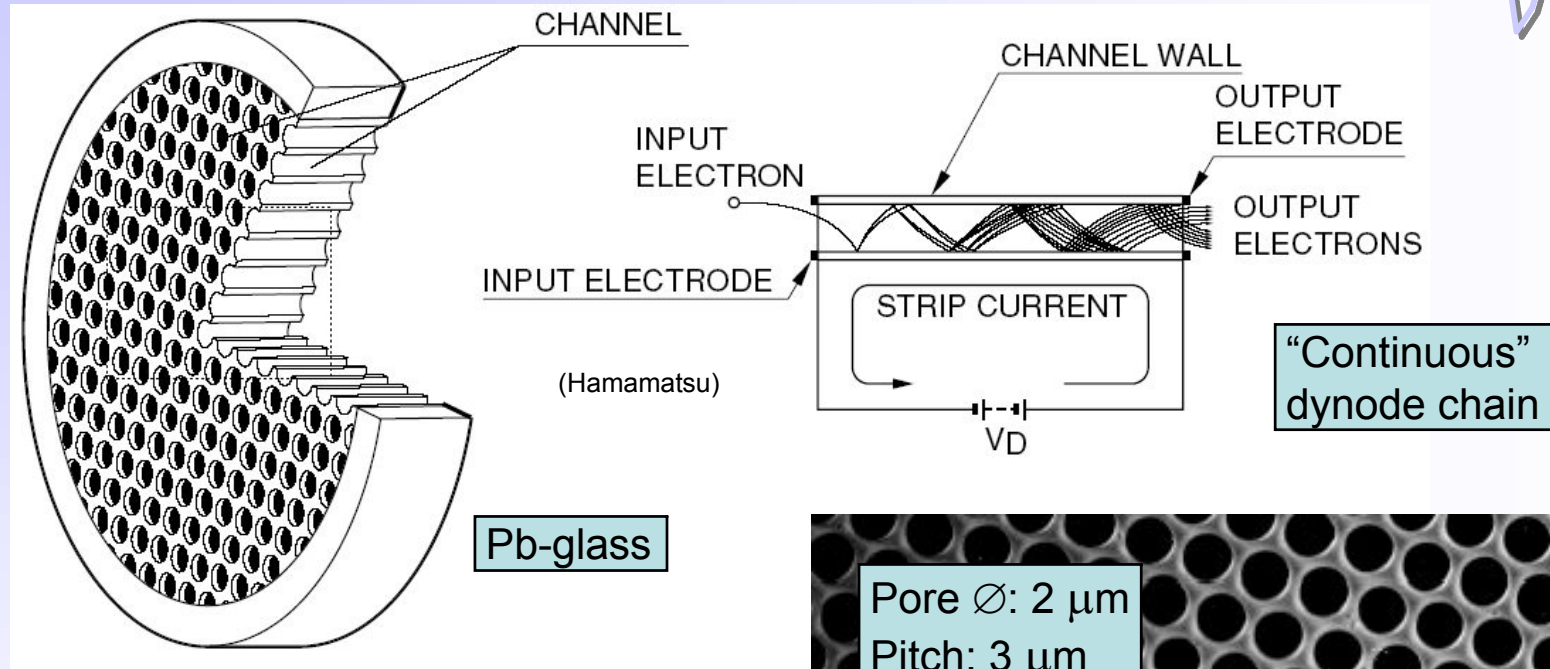
- short (<100ns) for e.g. high-speed CCD's to minimize afterglow;
- long (~1ms) for night-vision and surveillance to minimize flicker;



(Hamamatsu)



(Hamamatsu)

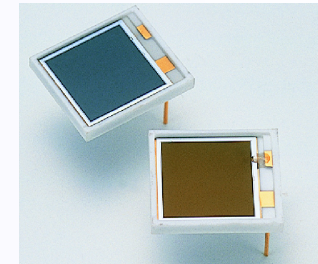
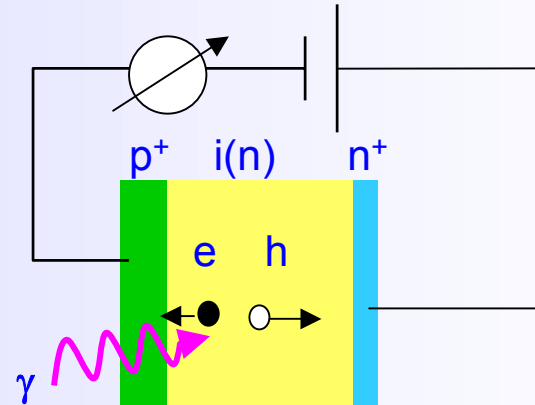


Kind of 2D PMT:

- + high gain up to $5 \cdot 10^4$;
- + fast signal (transit time spread ~ 50 ps);
- + less sensitive to B-field (0.1 T);
- limited lifetime (0.5 C/cm²);
- limited rate capability (μ A/cm²);

Photodiodes:

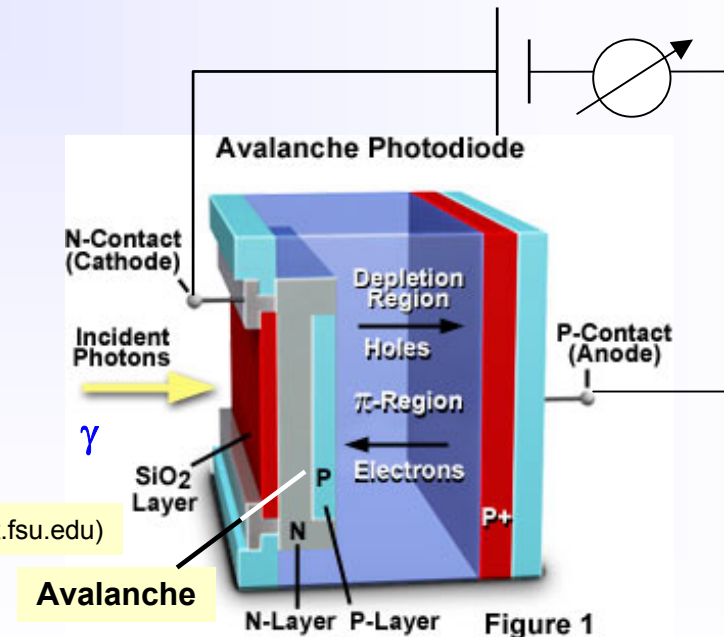
- P(I)N type (see lecture 2b);
- p layer very thin ($<1 \mu\text{m}$), as visible light is rapidly absorbed by silicon (see next slide);
- High QE (80% @ $\lambda \approx 700\text{nm}$);
- No gain: cannot be used for single photon detection;



Avalanche photodiode:

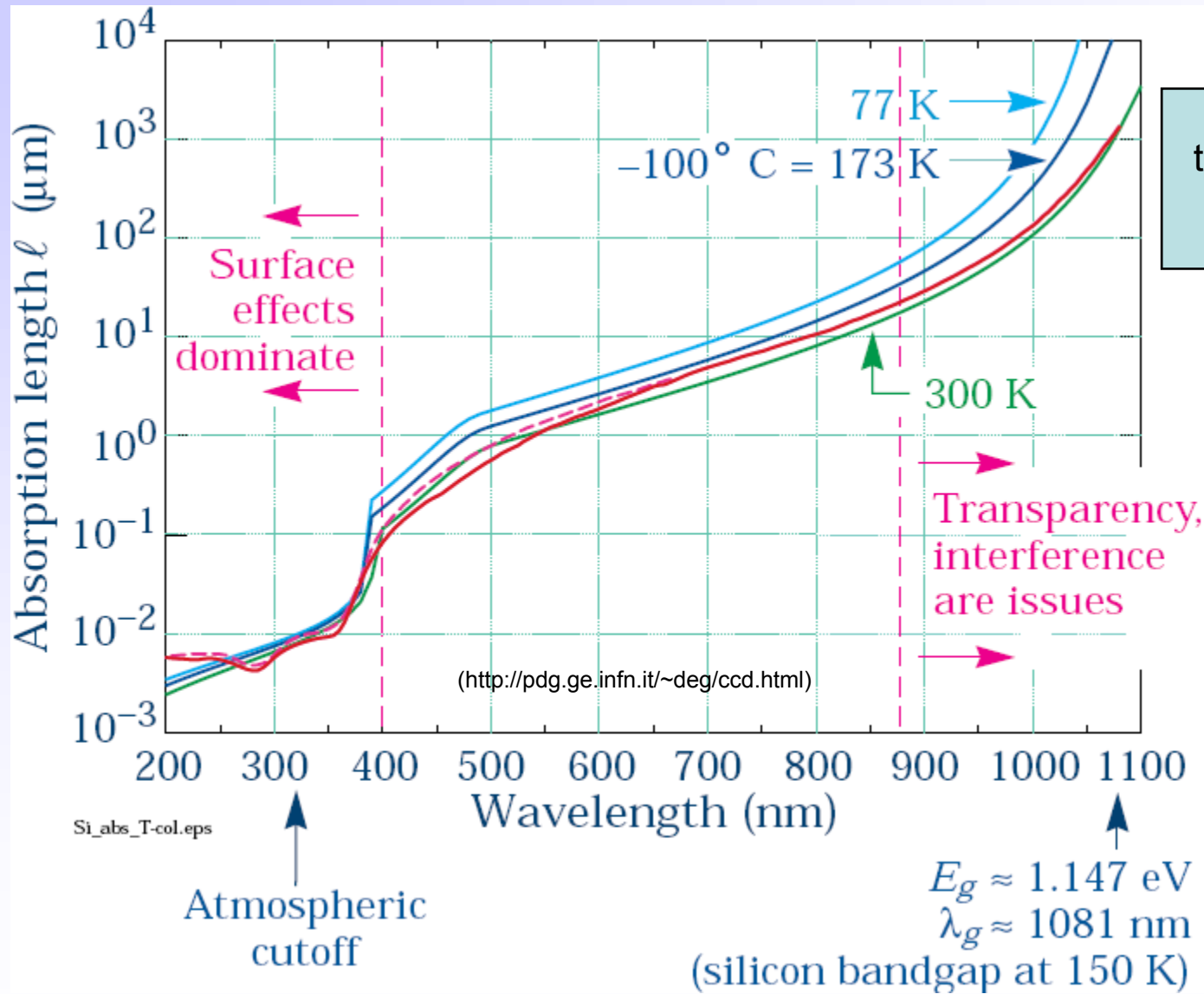
- High reverse bias voltage: typ. 100-200 V
- ⇒ due to doping profile, high internal field and avalanche multiplication;
- High gain: typ. 100-1000;
 - Used in CMS ECAL;

(<http://micro.magnet.fsu.edu>)





Light absorption in Silicon





Many more types exist ...

3b Photo-detection

Non-exhaustive list:

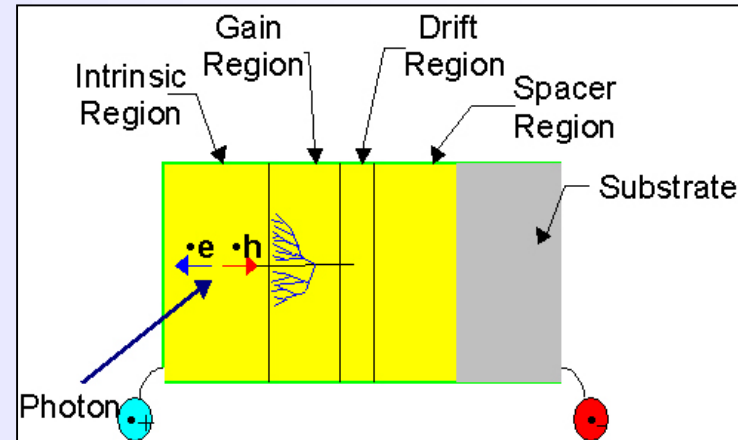
- Visible Light Photon Counter (VLPC);
- Silicon Photo-Multiplier (Si-PMT);
- Strip, pad and pixel arrays;
- CCD's:
 - conventional, front-illuminated;
 - thinned, back-illuminated;
 - fully-depleted, back-illuminated;

(see a detailed example of the latter 2 for astronomical applications in the next slides)

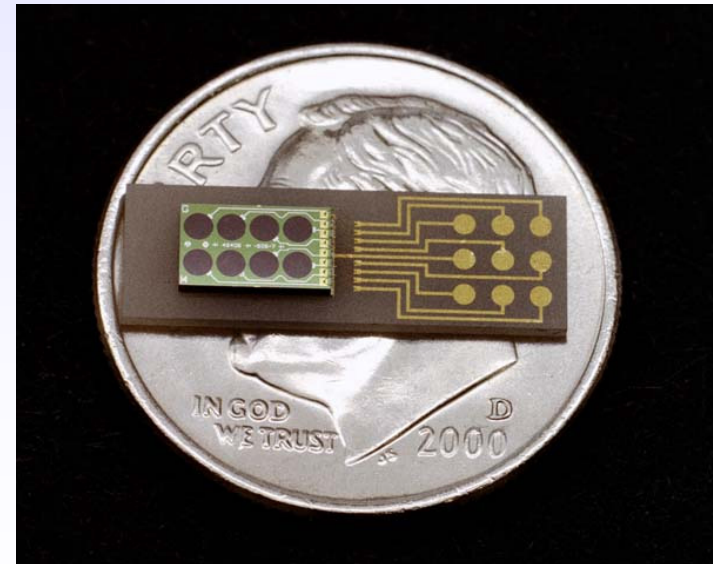
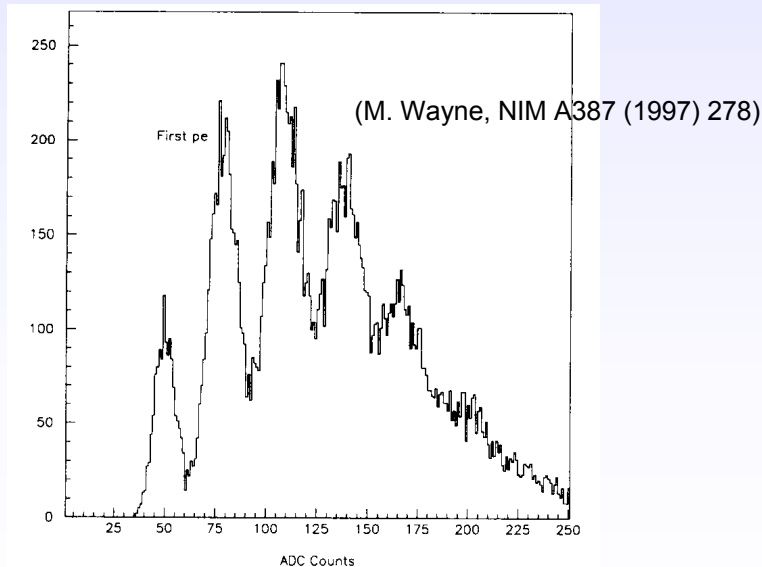
extra slide
not shown

Visible Light Photon Counter (VLPC):

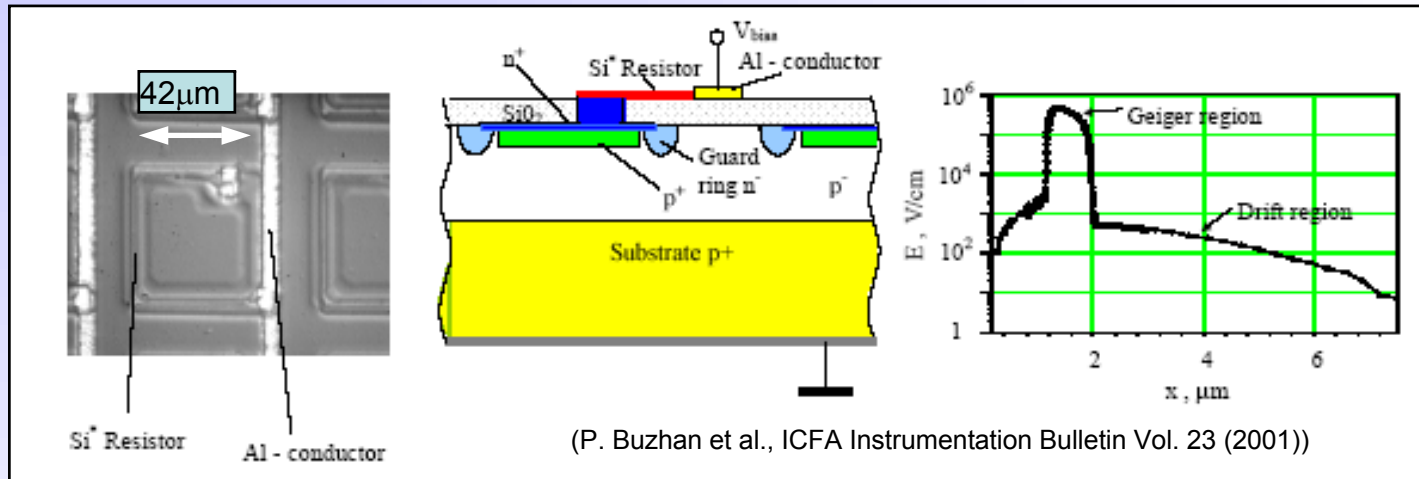
- Originally developed by Rockwell;
- Operation at low bias voltage (7V);
- High IR sensitivity:
⇒ *requires cooling at liquid He T° (7K)!*
- Q.E. \approx 70% around 500 nm;
- Gain up to 50.000 !
- used in the D0 Central Scintillating Fibre Tracker



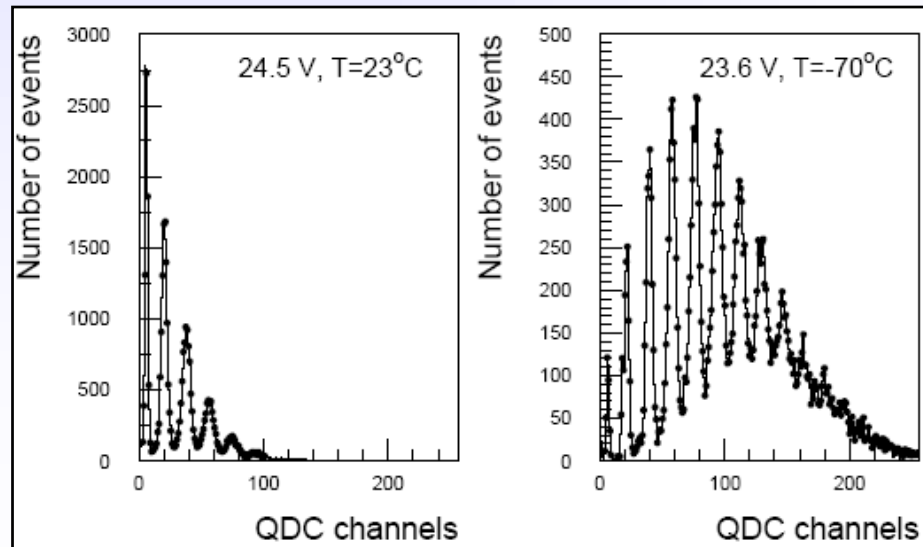
(http://d0server1.fnal.gov/projects/scifi/pictures/vlpc_related.html)



extra slide
not shown



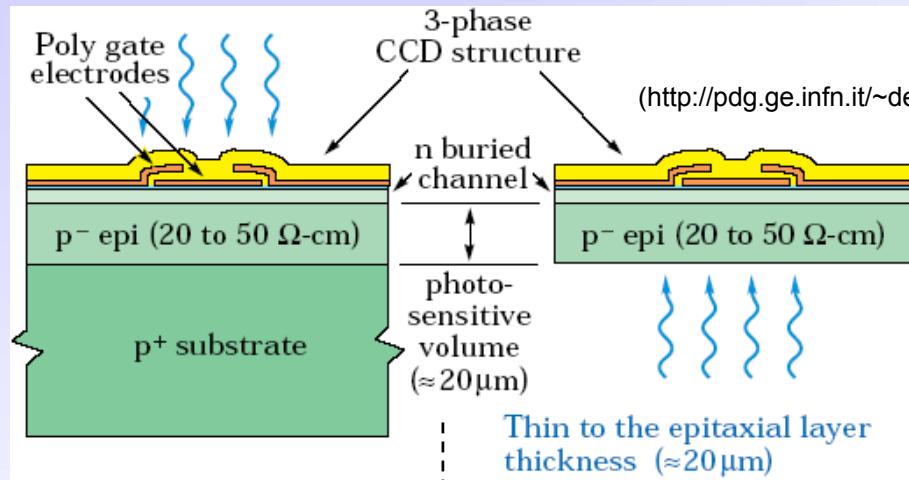
Kind of APD array
operating in Geiger
mode



Front-illuminated CCD

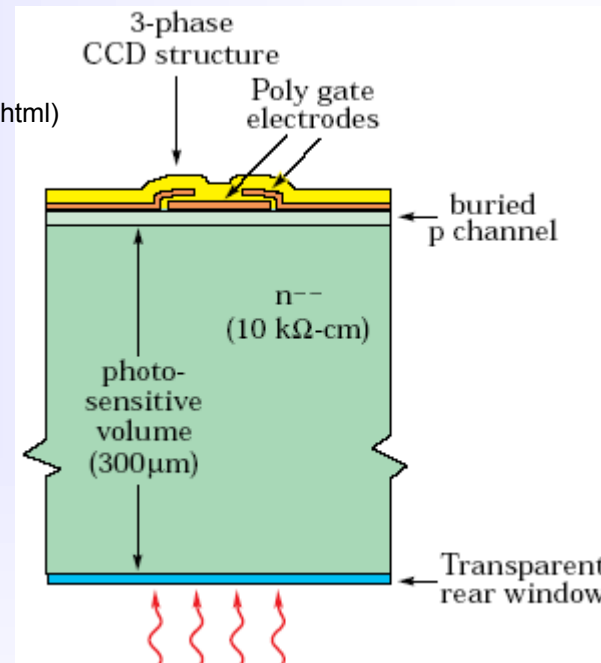
Back-illuminated thinned CCD

Back-illuminated fully depleted CCD

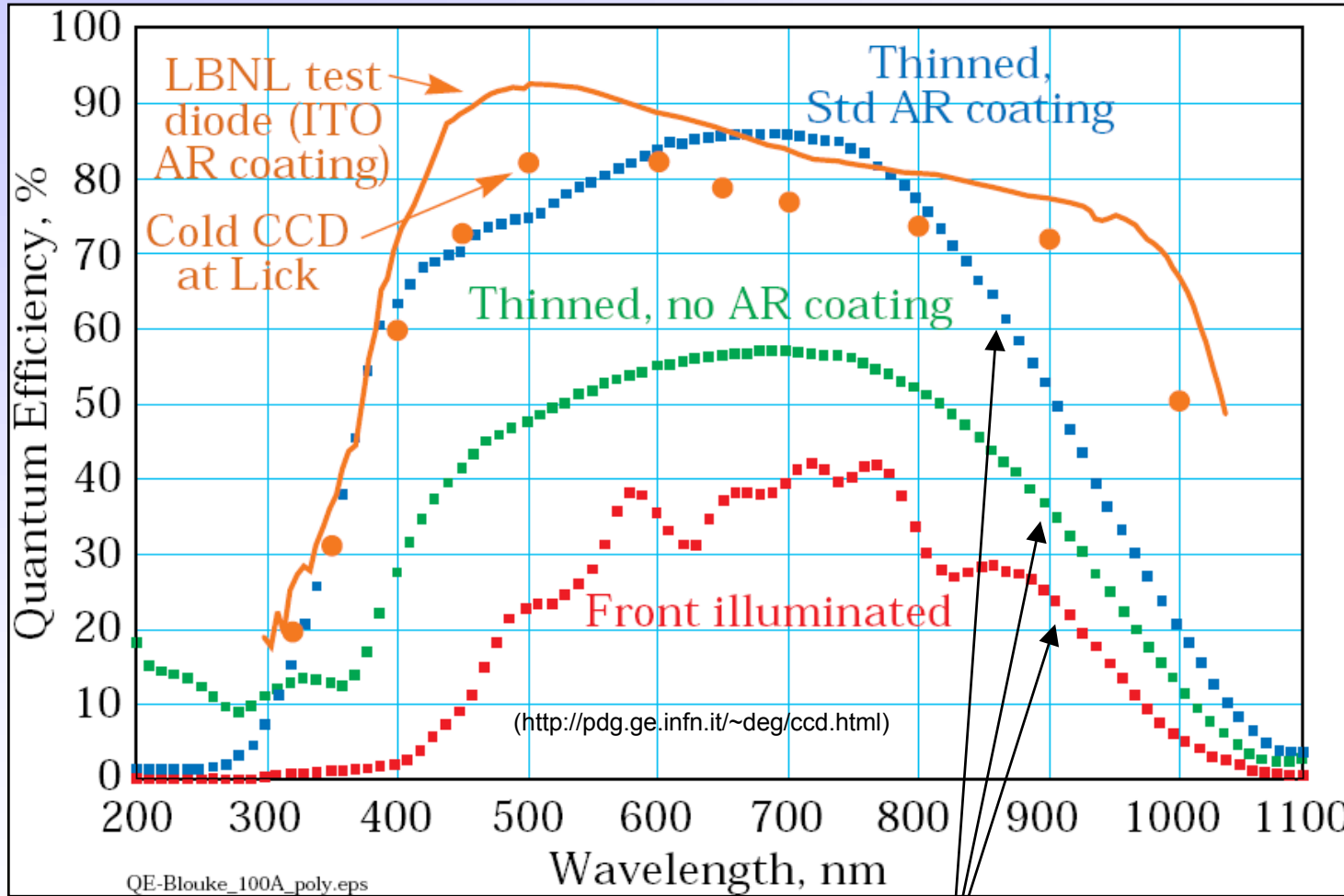


- poor response in blue (poly-Si) and IR (thin epitaxial layer);
- interference (gate);

- thinning difficult, expensive and not flat;
- poor IR response;
- fringing;
- lateral diffusion ⇒ degraded PSF;
- charge build-up at rear surface;



- +conventional MOS process;
- +full QE up to $\lambda=1\mu\text{m}$, (no fringing);
- +good blue response;
- enhanced sensitivity to radiation



(M. Blouke and M. Nelson, SPIE **1900** (1993), 228-240)



And the result is ...

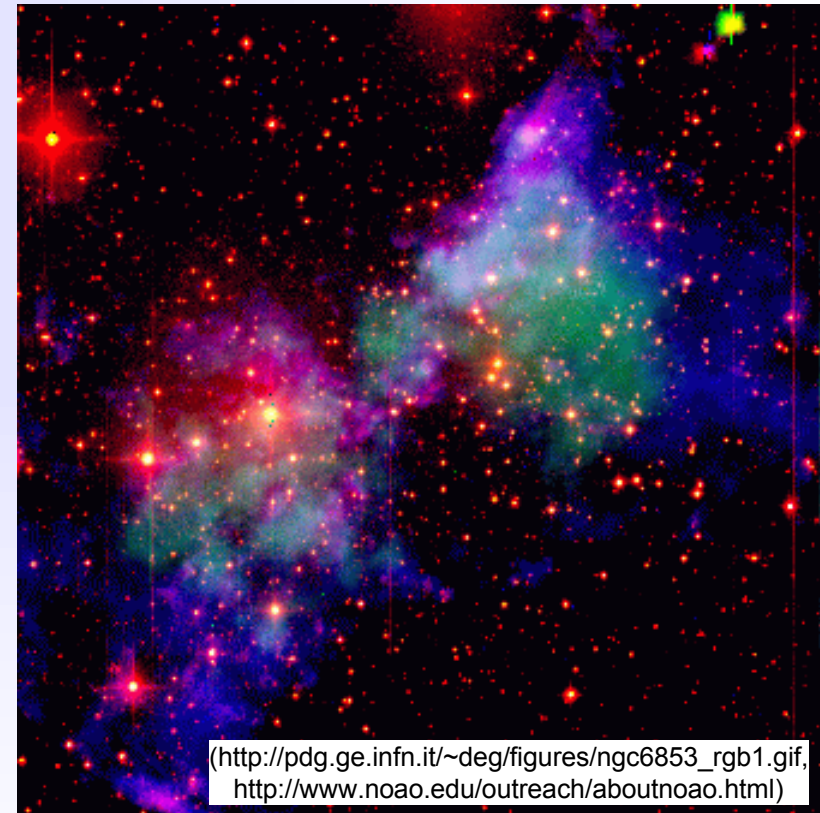
3b Photo-detection

Dumbbell Nebula in Vulpecula (M27, NGC 6853)



(<http://antwrp.gsfc.nasa.gov/apod/ap981009.html>)

FORS false color image using a Tektronix back-illuminated $2k \times 2k$ CCD with $24\mu m$ pixels thinned and anti-reflection coated. This image was obtained on ESO 8.2-m VLT Unit Telescope (UT) 1 on September 28, 1998.

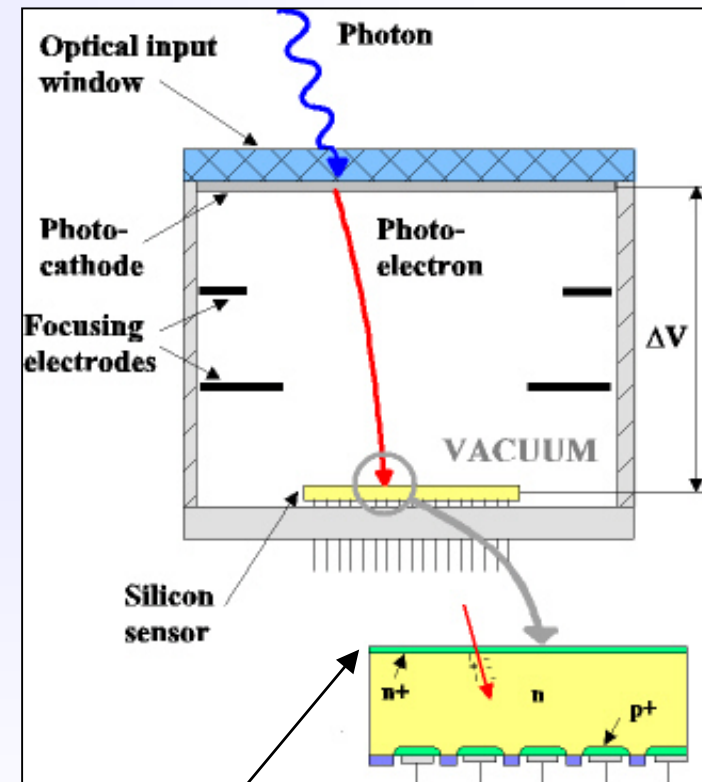


(http://pdg.ge.infn.it/~deg/figures/ngc6853_rgb1.gif,
<http://www.noao.edu/outreach/aboutnoao.html>)

NOAO false color image using a back-illuminated fully depleted $2k \times 2k$ CCD with $15\mu m$ pixel. This image was obtained on WIYN 3.5-m Telescope on June 7, 2001.

Basic principle:

- Combination of vacuum photon detectors and solid-state technology;
- Input: collection lens, (active) optical window, photo-cathode;
- Gain: achieved *in one step* by energy dissipation of keV pe's in solid-state detector anode; this results in low gain fluctuations;
- Output: direct electronic signal;
- Encapsulation in the tube implies:
 - compatibility with high vacuum technology (low outgassing, high T° bake-out cycles);
 - internal (for speed and fine segmentation) or external connectivity to read-out electronics;
 - heat dissipation issues;



Basic properties:

- Photo-emission from photo-cathode;
- Photo-electron acceleration to $\Delta V \approx 10\text{-}20\text{kV}$;
- Energy dissipation through ionization and phonons ($W_{\text{Si}}=3.6\text{eV}$ to generate 1 e-h pair in Si) with low fluctuations (Fano factor $F \approx 0.12$ in Si);

- Gain M :
$$M = \frac{e(\Delta V - V_{th})}{W_{\text{Si}}}$$

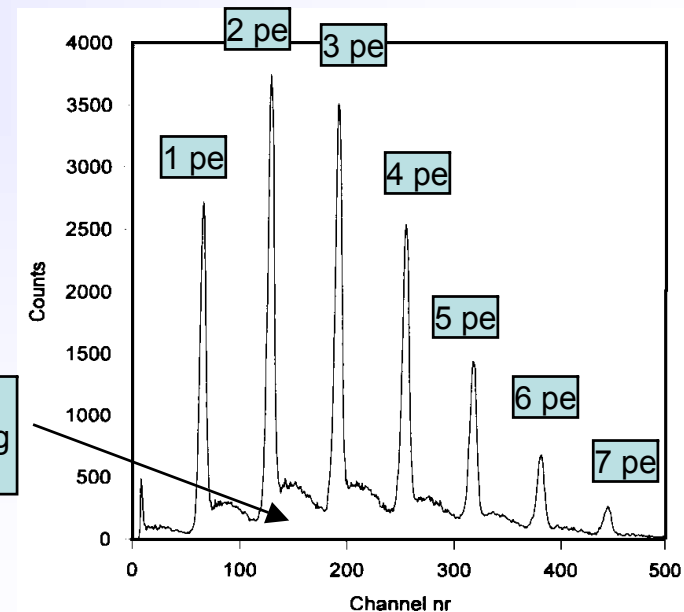
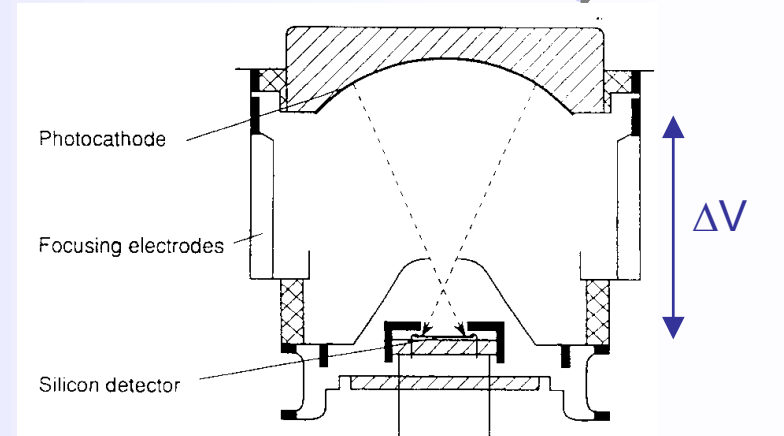
- Gain fluctuations σ_M :
$$\sigma_M = \sqrt{F \times M}$$

⇒ dominated by electronics

- Example: $\Delta V = 20\text{kV}$

⇒ $M \approx 5000$ and $\sigma_M \approx 25$

- suited for single photon detection with high resolution;



(C.P. Datema et al., NIM A 387(1997) 100)

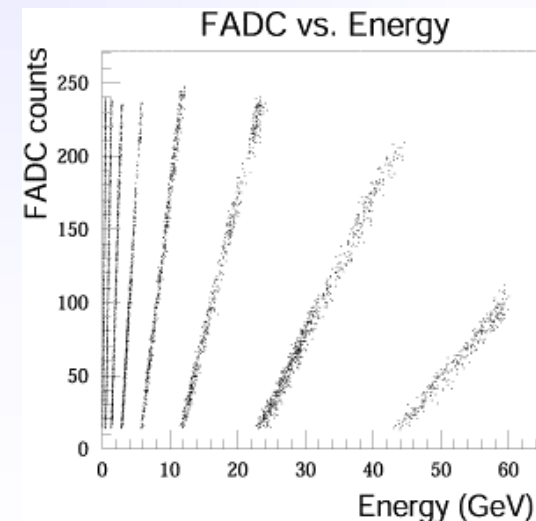
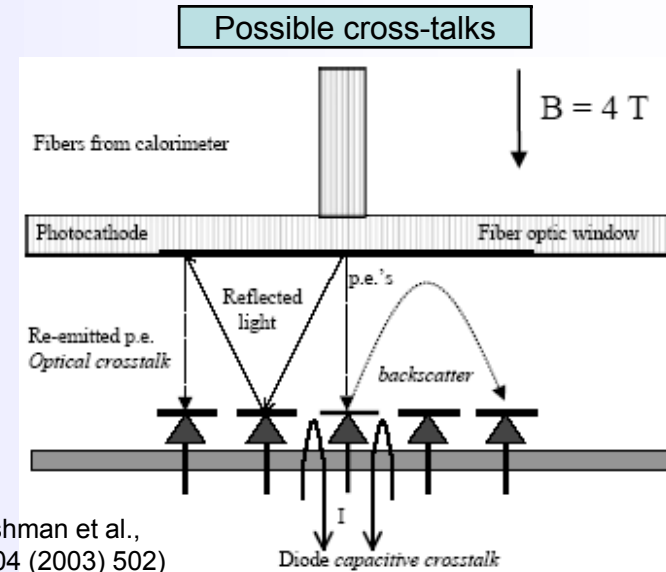


Multi-pixel proximity-focussed HPD

3b Photo-detection

DEP-CMS HCAL example:

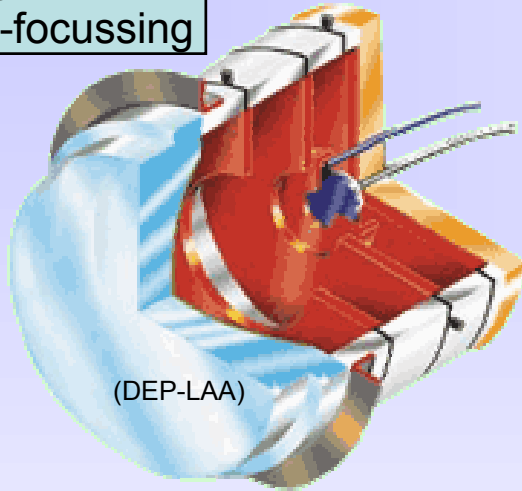
- $B=4T \Rightarrow$ proximity-focussing with 3.35mm gap and $HV=10kV$;
- Minimize cross-talks:
 - pe back-scattering: align with B ;
 - capacitive: Al layer coating;
 - internal light reflections: a-Si:H AR coating optimized @ $\lambda = 520nm$ (WLS fibres);
- Results in linear response over a large dynamic range from minimum ionizing particles (muons) up to 3 TeV hadron showers;



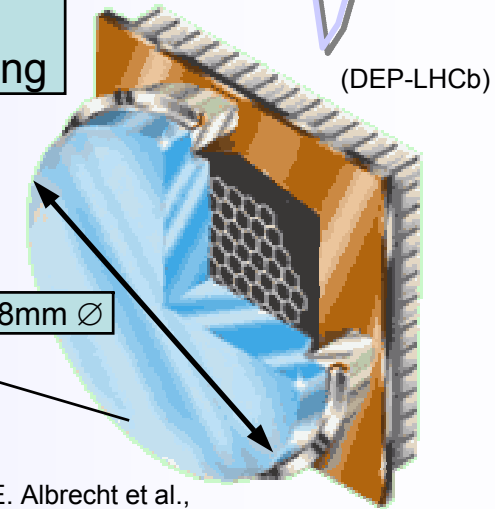
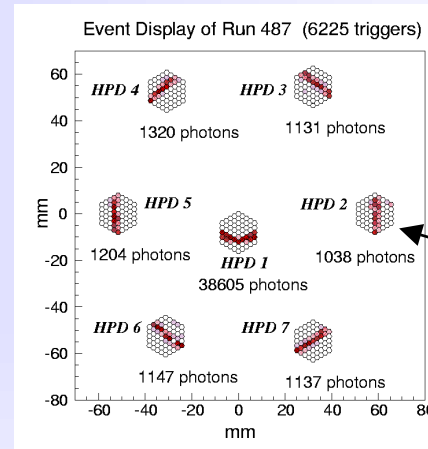
<http://cmsinfo.cern.ch/Welcome.html/CMSdetectorInfo/CMSHcal.html>

extra slide not shown

Single-diode cross-focussing

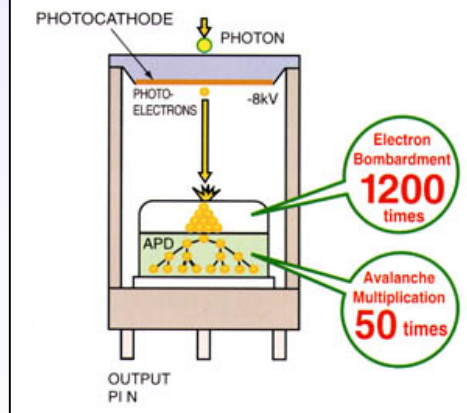


Multi-pixel proximity-focussing



(E. Albrecht et al., NIMA A 411 (1998) 249-264)

Single avalanche diode HPD



(Hamamatsu)

DEP-LHCb development:

- Multi-alkali photo-cathode;
- Commercial anode with 61 2mm-pixels; vacuum feed-throughs to external analog (VA2) readout electronics;
- Proximity-focussing electron optics;
- Poor intrinsic active area coverage (~50%);



Various kinds of commercial HPD's

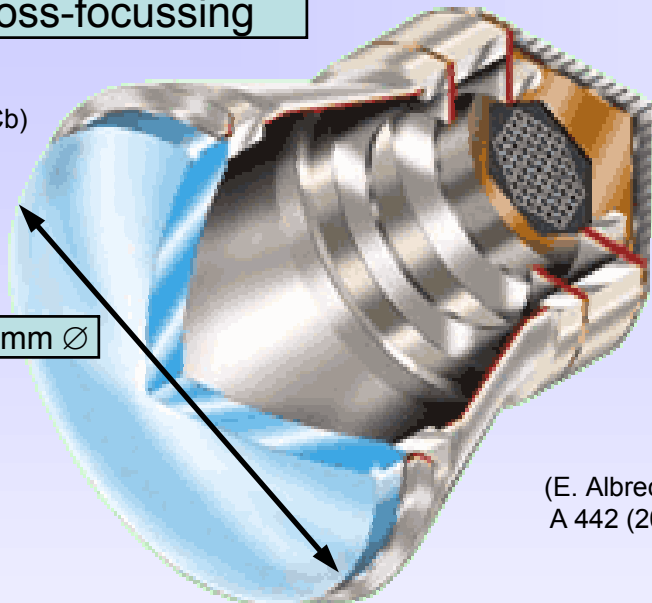
extra slide
not shown

3b Photo-detection

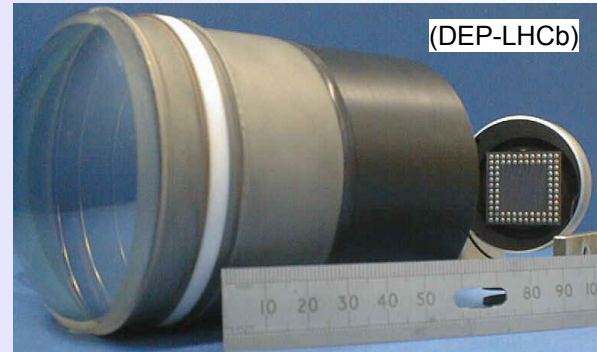
Multi-pixel,
cross-focussing

(DEP-LHCb)

72mm \varnothing



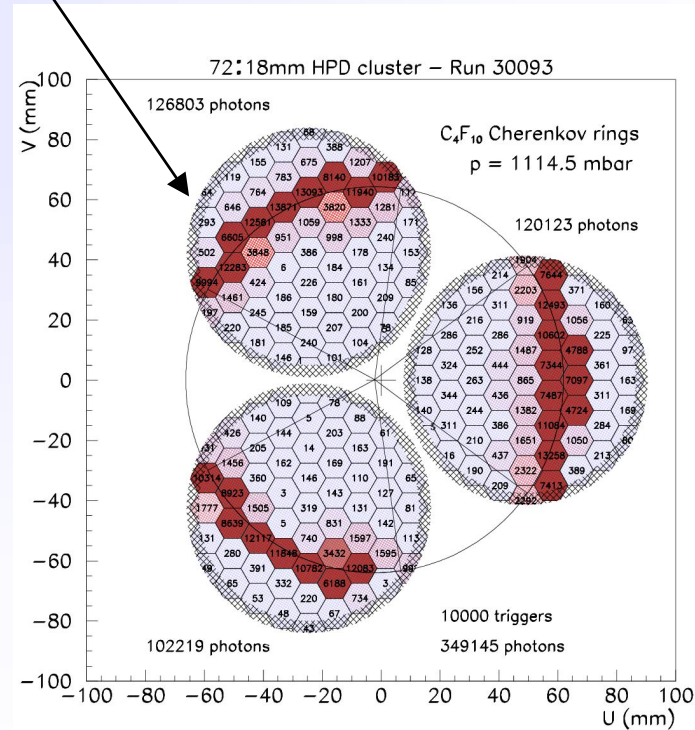
(E. Albrecht et al., NIMA A 442 (2000) 164-170)



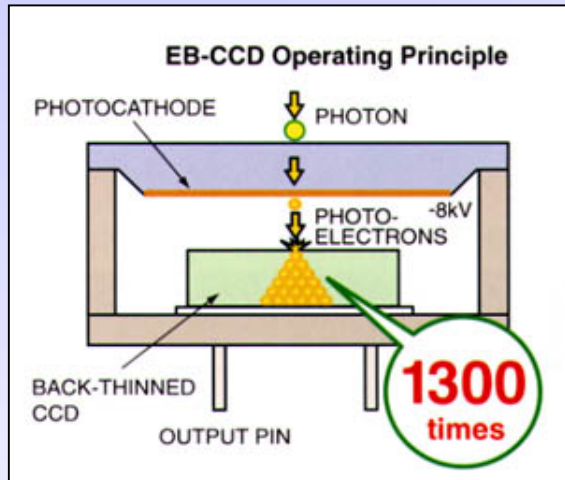
(DEP-LHCb)

DEP-LHCb development:

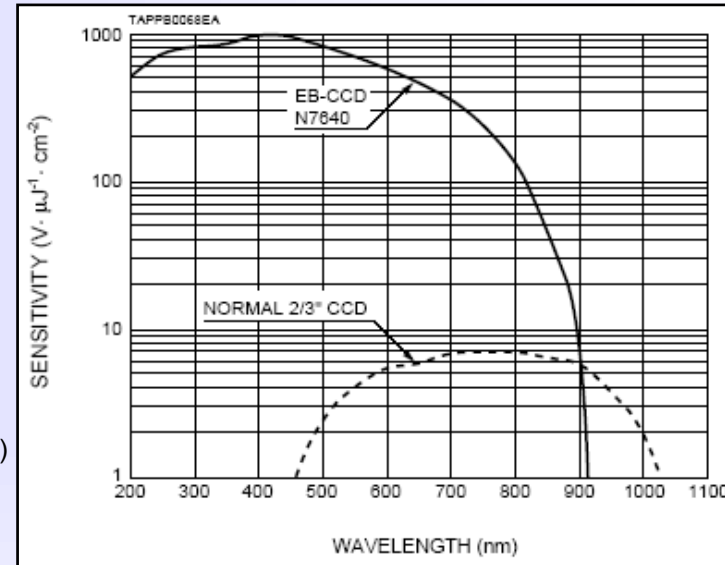
- Commercial anode;
- Cross-focussing electron optics (de-magnification by ~ 5);
- High intrinsic active area coverage (83%);



EBCCD
proximity-focussed



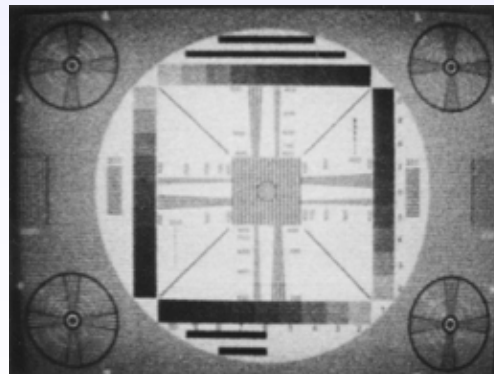
(Hamamatsu)



Commercial 2/3" CCD



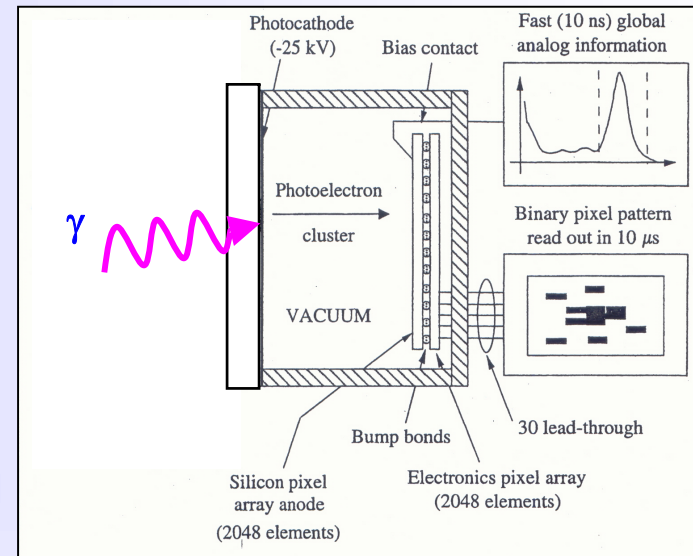
Hamamatsu N7640 EB-CCD



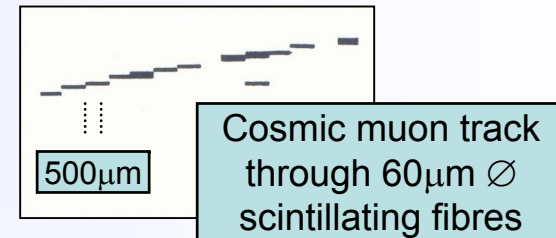
Object illuminance: 0.1lx

Imaging with Silicon Pixel Array:

- Pixel array sensor bump-bonded to binary electronic chip, developed for tracking (CERN-RD19);
- Flip-chip assembly encapsulated inside vacuum tube using standard parts, commercial ceramic carriers and packaging techniques;
- First ISPA prototype (1994) used to read small-diameter scintillating fibres developed for tracking (CERN-RD7);
- Spin-off applications for beta- and gamma-detection (quartz and YAP-crystal windows)

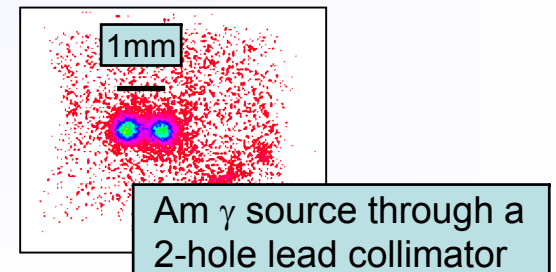


(T. Gys et al., NIMA 355 (1995) 386-389)



Cosmic muon track through 60 μm Ø scintillating fibres

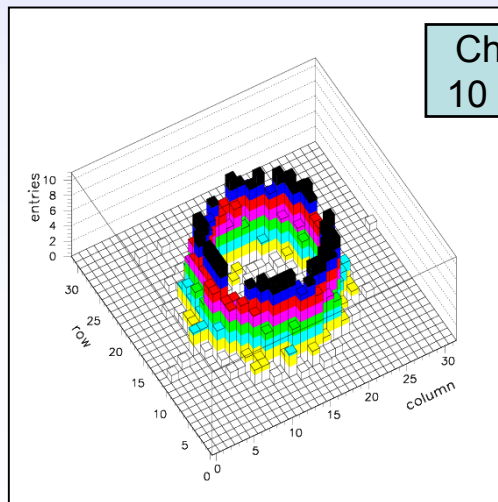
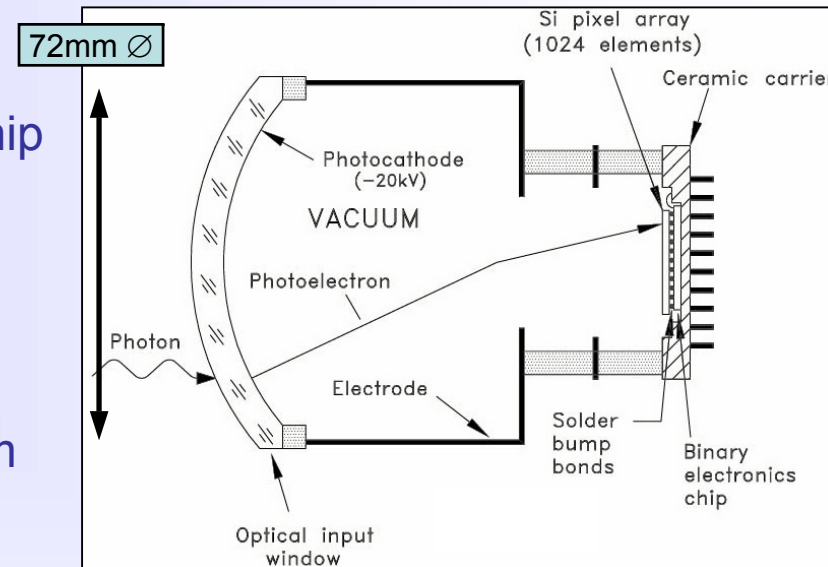
(F. Cindolo et al., IEEE TNS, Vol. 50, No. 1, February 2003, 126-132)



Am γ source through a 2-hole lead collimator

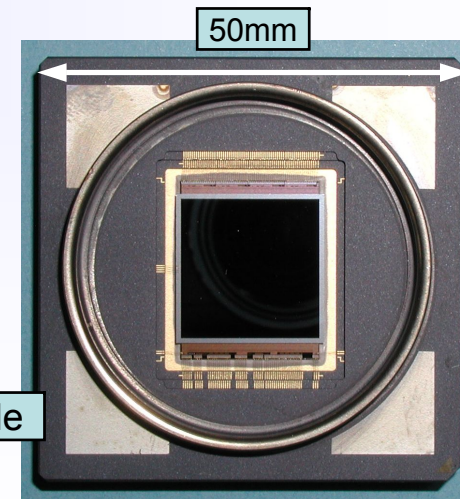
Industry-LHCb development:

- LHCb-dedicated pixel array sensor bump-bonded to binary electronic chip (in coll. w. ALICE-ITS), specially developed high T° bump-bonding;
- Flip-chip assembly encapsulated inside vacuum tube using full-custom ceramic carrier;



Cherenkov rings from 10 GeV/c π^- through air

(M. Moritz et al., IEEE TNS Vol. 51, No. 3,, June 2004, 1060-1066)





The pad HPD for RICH detectors

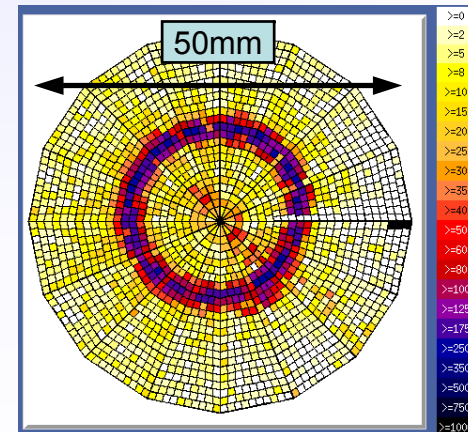
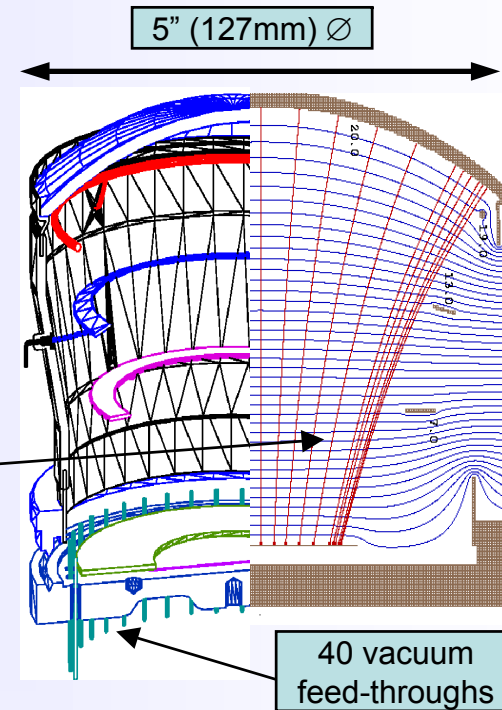
extra slide
not shown

3b Photo-detection

Full in-house (LHCb, CERN, Bologna, CdF) development:

- Aim for active area > 80%;
- Bi-alkali photo-cathode;
- “Fountain” focussing electron optics (de-magnification ~2.4);
- Si detector: $16 \times 128 = 2048$ pads ($\sim 1 \times 1 \text{ mm}^2$ each);
- Analogue electronics (16 VA3 chips) encapsulated inside vacuum tube;
- Standard Al wedge bonding;

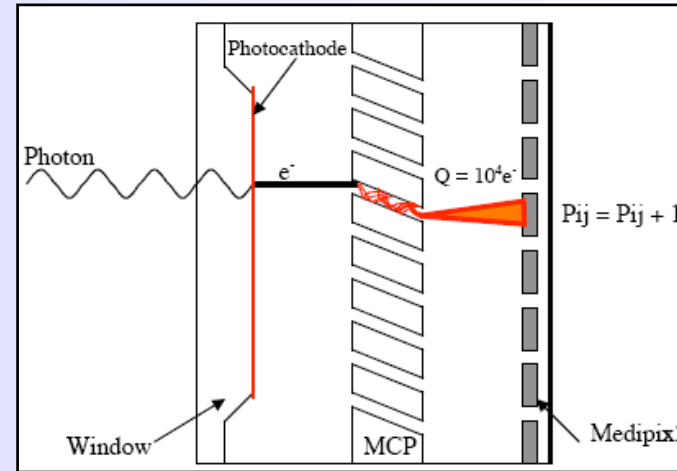
(LHCb 98-007, RICH)



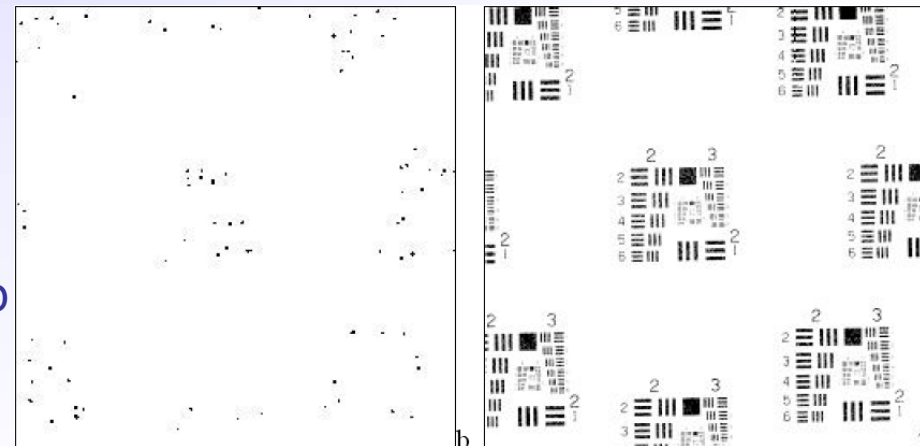
extra slide
not shown

Development of next-generation astronomical AO:

- Alternative to replace more conventional high-speed CCD's;
- Aim for IR response, ultra-low noise and several kHz frame-rates;
- GaAs photo-cathode;
- Proximity-focussing electron optics;
- High-gain wide dynamic range MCP;
- Anode: Medipix2 photon-counting chip used both as direct electron detector (55 μ m pixels) and FE readout electronics;



(J. Vallerga et al., Proc. SPIE, vol. 5490 (2004) 1256-1267)



Images of USAF test pattern, 100ms (left) and 100s (right) exposures, 50k MCP gain



Non-exhaustive list:

- www.photonis.com: “Photomultiplier tubes, principles and applications”;
- www.hamamatsu.com;
- www.dep.nl;
- A.H. Sommer, “Photoemissive materials”, J. Wiley & Sons (1968);
- H. Bruining, “Physics and Applications of Secondary Electron Emission”, Pergamon Press (1954);
- I. P. Csorba, “Image Tubes”, Sams (1985);
- Proceedings of the Beaune Conferences (1996-1999-2002) on “New Developments in Photo-detection”, published in NIMA;