

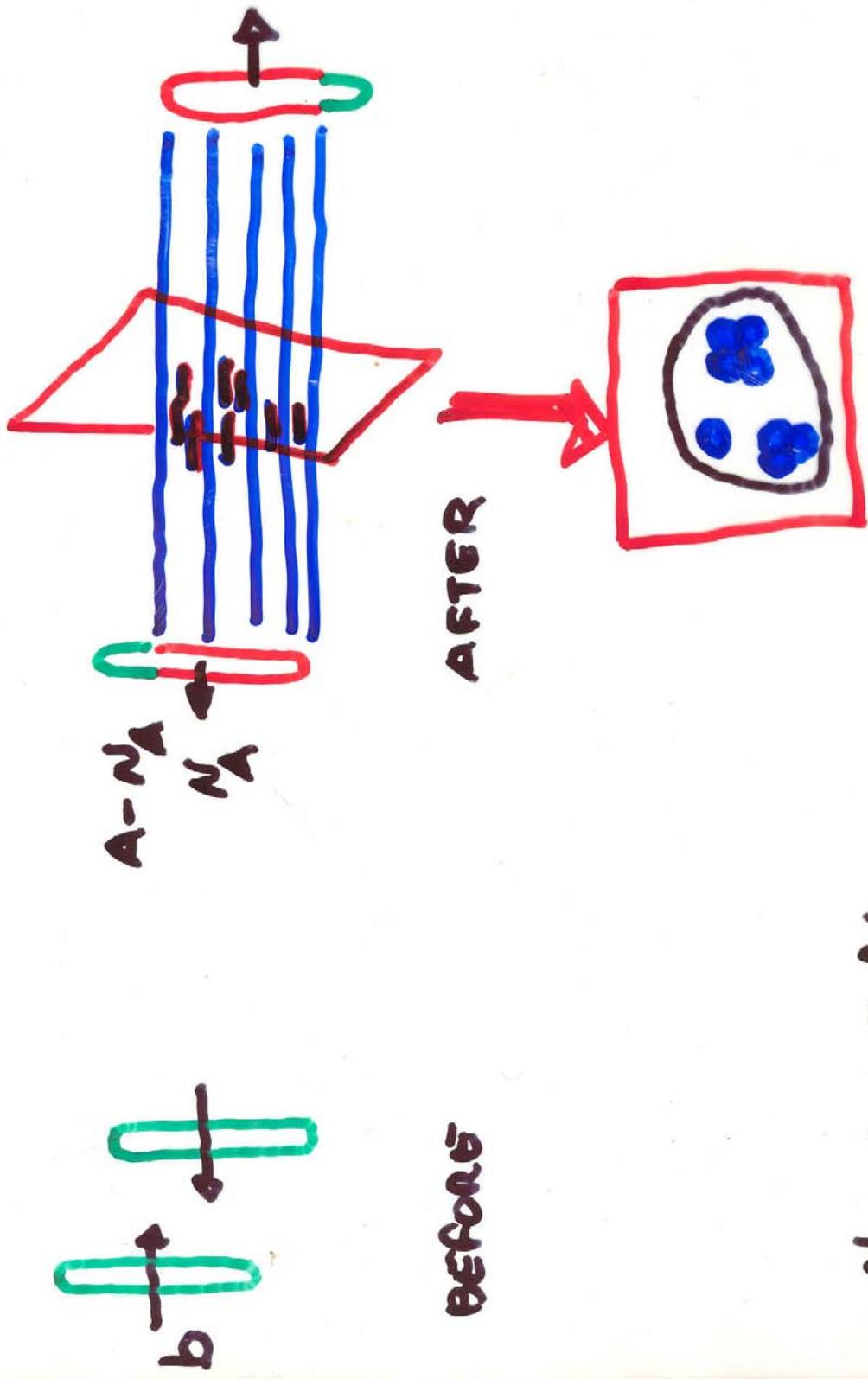
PERCOLATION AND P_T DISTRIBUTIONS

- 1 - DSM AND PERCOLATION
- 2 - SCHWINGER MODEL
- 3 - AA COLLISIONS AT RHIC AND LHC
- 4 - AWAY FROM $y=0$
- 5 - CONCLUSIONS

ROBERTO U.
CARLOS P.
ELENA F.
J. DD

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1 - DSM with PERCOLATION



² IMPACT PARAMETER PERCOLATION - 2 D

DENSITY : $\gamma \equiv \left(\frac{n}{R}\right)^2 \bar{N}_s$

n : STRINGS TRANSVERSE RADIUS
 $n \approx 0.2 \text{ fm}$

R : INTERACTION RADIUS

\bar{N}_s : AVERAGE NUMBER OF STRINGS

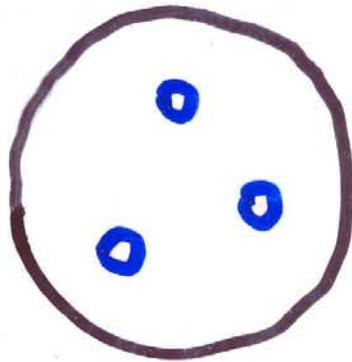
APPROXIMATIONS: $R \approx R^P n^{1/3}$ (Nuclear Phys.) $R \approx 1 \text{ fm}$

$$\bar{N}_s \approx N_P n^{4/3}$$
 (Multi-Hadron Physics)

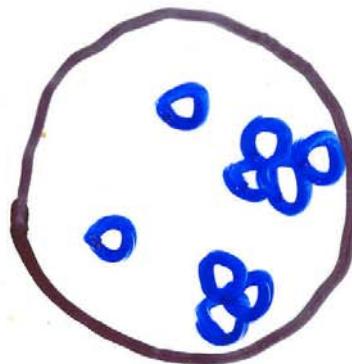
N_P : Production

$$\rho = \left(\frac{\pi}{R^3}\right)^2 N_A^{4/3}$$

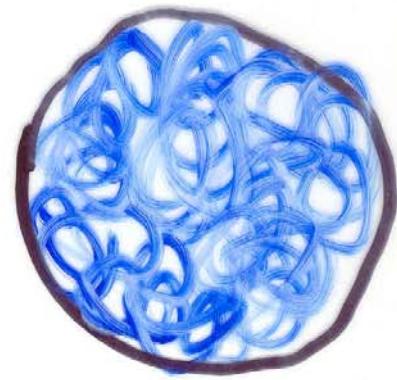
\uparrow Increases with Energy
 \uparrow Increases with Centrality



$\gamma \gg 1$



$\gamma \approx 1$



$\gamma \gg 1$

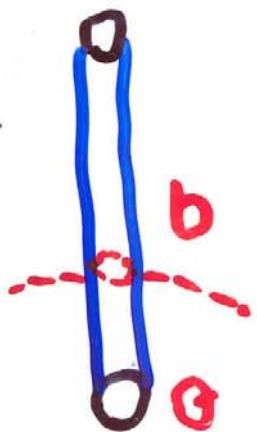
OVERRUNNING?

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$$\text{OBJECTS} \sim \sqrt{s}$$
$$\text{AREA OF INTERACTION} \sim (\text{mass})^2$$
$$\text{OBJECTS / AREA} \rightarrow \infty$$

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2 - SCHWINGER MODEL



$$\frac{d\eta}{dy} \sim Q$$

$\left\{ \right.$

$$\langle p_y^+ \rangle \sim \sigma$$

$$\text{GAUSS: } Q = S \sigma$$

\downarrow

$$\frac{d\eta}{dy}$$

$\left. \right\} \langle p_y^+ \rangle$

CLUSTER OF N -STRINGS

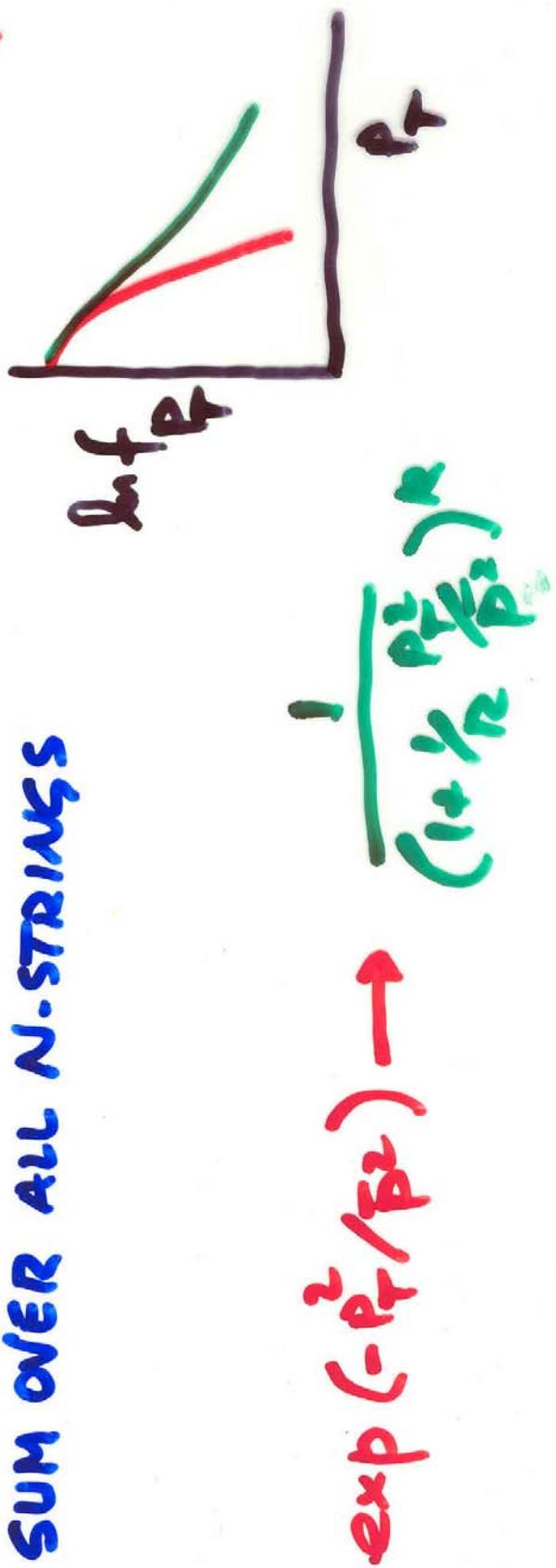
$$Q_N = \sqrt{\frac{s_N}{N s_1}} N Q_1$$

$$Ns_1 \gg s_N \gg s_1$$

$$\left\{ \begin{array}{l} \frac{d\eta}{dy}_N = \sqrt{\frac{S_N}{N}} \propto \sqrt{\eta}, \\ \langle P_T^2 \rangle_N = \sqrt{N} \sqrt{\frac{S_N}{N}} \propto P_T^2 \end{array} \right.$$

($\bar{\eta}$, \bar{P}_T : single string)

P_T -DISTRIBUTION OF A N-STRINGS
SUM OVER ALL N-STRINGS

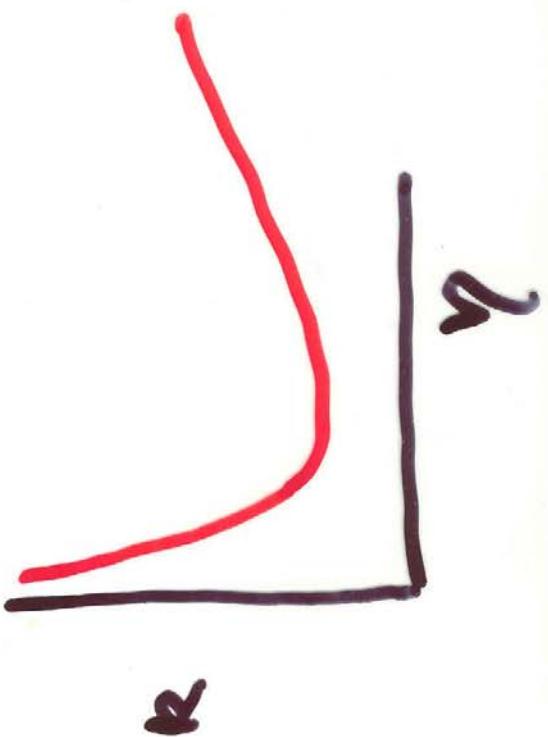


γ_k : NORMALIZED WIDTH OF N-CLUSTERS DISTRIBUT.

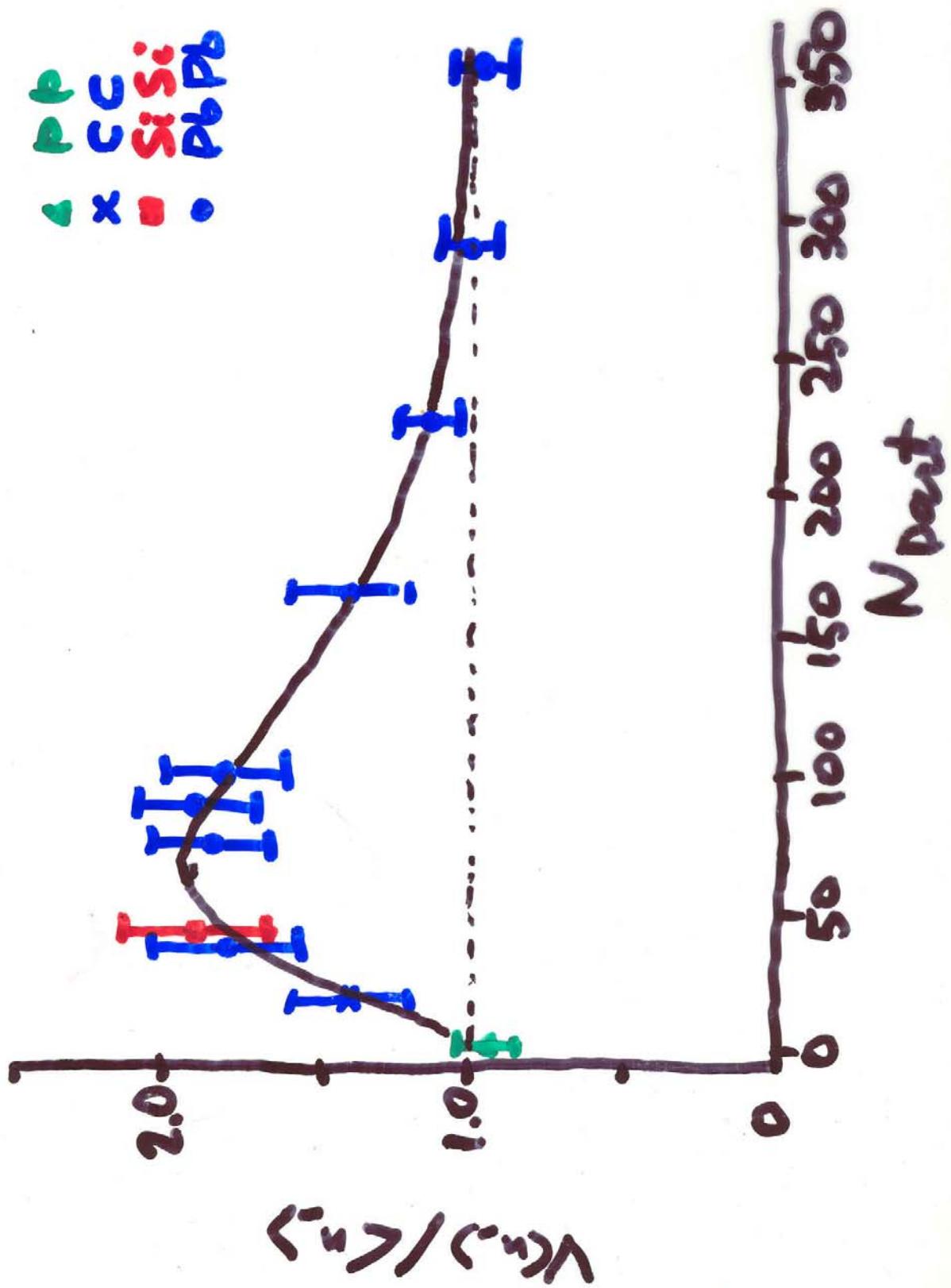
$$\gamma_k = \frac{\langle N^2 \rangle - \langle N \rangle^2}{\langle N \rangle^2}$$

NAIVE X:

$$\begin{aligned} \gamma &\ll 1, \quad N=1, \quad \gamma_k \rightarrow 0 \\ \gamma &> 1, \quad N=N_s, \quad \gamma_k \rightarrow 0 \end{aligned}$$



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RANDOM DISTRIBUTION

$$\langle Q \rangle = \sqrt{\frac{1}{N} \sum_i \langle \bar{N}_s(Q_i) \rangle}$$

$\langle A \rangle$: Average cluster Area
(in units of f^2_1)

$\langle N \rangle$: Average Number of strings
per clusters

BUT : $\left\{ \begin{array}{l} \bar{N}_c \langle N \rangle = \bar{N}_s \\ \bar{N}_c \langle A \rangle = \left(\frac{R}{N}\right)^2 (1 - \bar{a}^2) \end{array} \right.$

$$\rightarrow \frac{\langle A \rangle}{\langle N \rangle} = \frac{(1 - \bar{a}^2)}{\bar{N}}$$

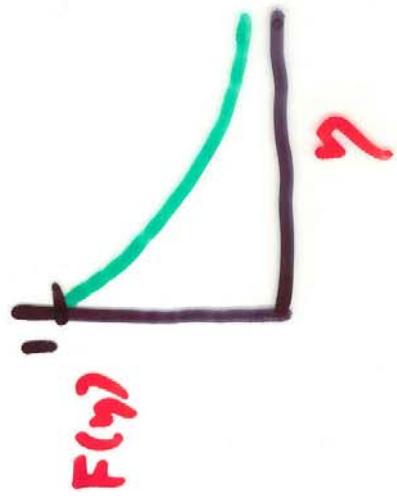
FINALLY:

$$\left\{ \begin{array}{l} \frac{dn}{dy} = F(\eta) \bar{N}_s = \\ \langle \bar{P}_T^2 \rangle = \end{array} \right.$$

$$\frac{R(\eta)}{k(\eta) - 2} \frac{1}{F(\eta)} \bar{P}^2$$

WITH

$$F(\eta) \equiv \sqrt{\frac{1 - \alpha^2}{\eta}}$$



CONSEQUENCES:

SLOW INCREASE OF $\alpha \gamma dy/dy$ (with $\sqrt{\eta_1}$ and η_2)
INCREASE OF $\langle \bar{P}_T^2 \rangle$ (with $\sqrt{\eta_1}$ and η_2)

UNIVERSALITY :

γ IS UNIVERSAL

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$\langle P_T^2 \rangle$: UNIVERSAL FUNCTIONS OF γ

$d\eta/dy$: NOT UNIVERSAL !

$$[\eta = \left(\frac{R}{n}\right)^2 \bar{\eta}_s]$$

But: $\bar{N}_s = \left(\frac{R}{n}\right)^2 \gamma \Rightarrow \frac{d\eta}{dy} = F(\eta) \gamma \left(\frac{R}{n}\right)^2 \bar{\eta}$

$$(R = R_p n_A^{1/3})$$

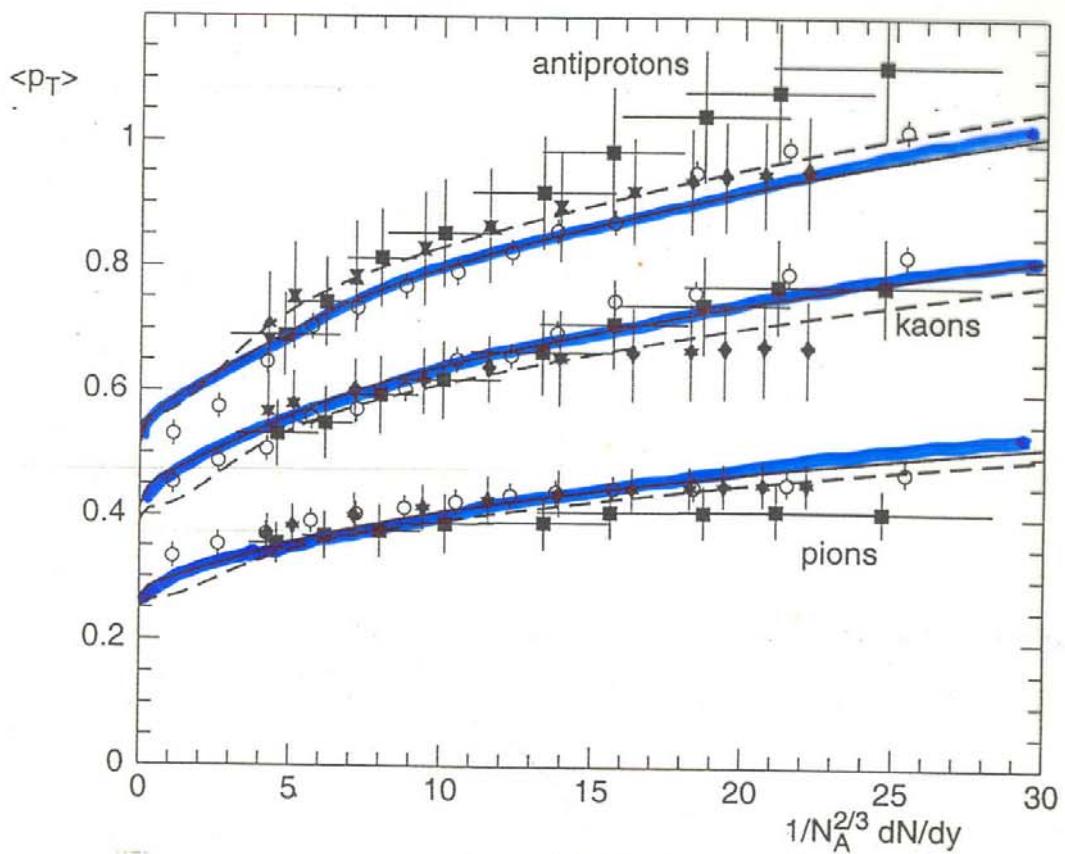
$$\Rightarrow \frac{1}{n_A} \eta_s \frac{d\eta}{dy} = F(\eta) \gamma \left(\frac{R_p}{n}\right)^2 \bar{\eta}$$

UNIVERSAL FUNCTIONS OF γ

$i = \pi, \kappa, \bar{P}$

$$\boxed{\langle P_T^2 \rangle_i = \bar{P}_i \phi(x)}$$

$$x = \frac{1}{n_A} \frac{d\eta}{dy}$$

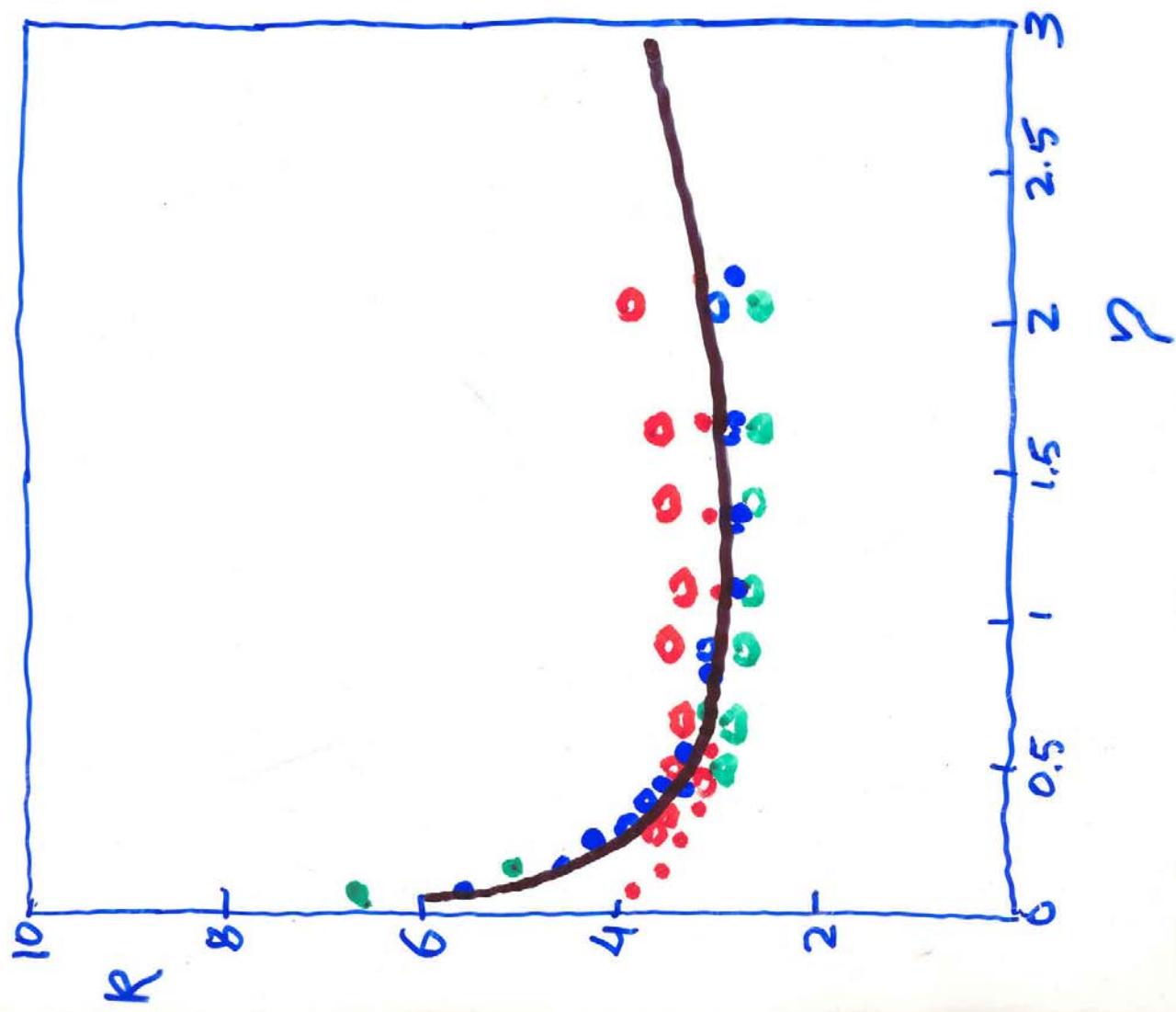


$p\bar{p} : \sqrt{s} = 1.8 \text{ TeV}$

PHOBOS + PANDA_{AuAu}: $\sqrt{s} = 200 \text{ GeV}$

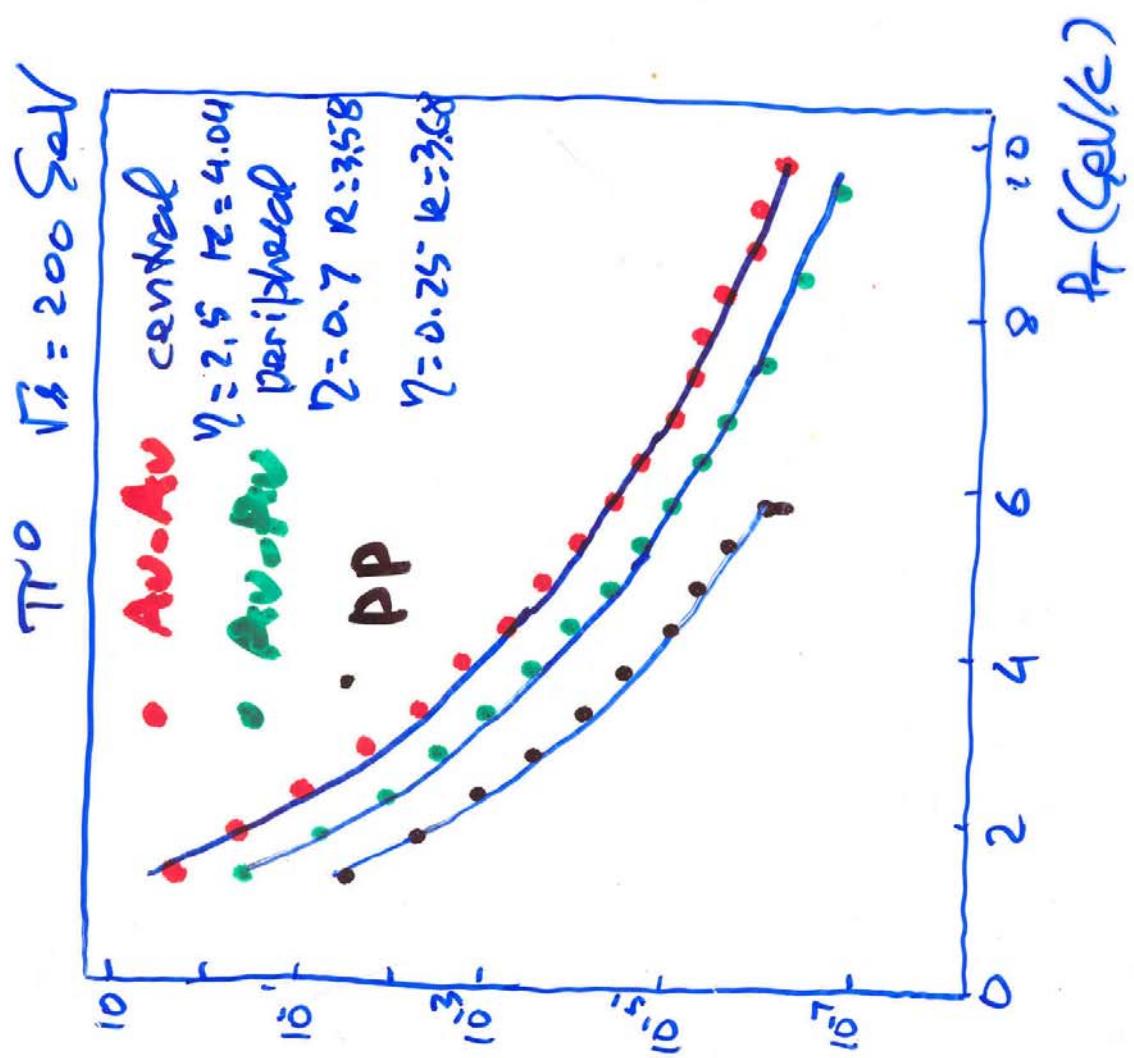
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- π/PP
- π/AuAu
- κ/PP
- κ/AuAu
- $\bar{\rho}/\text{PP}$
- $\bar{\rho}/\text{AuAu}$



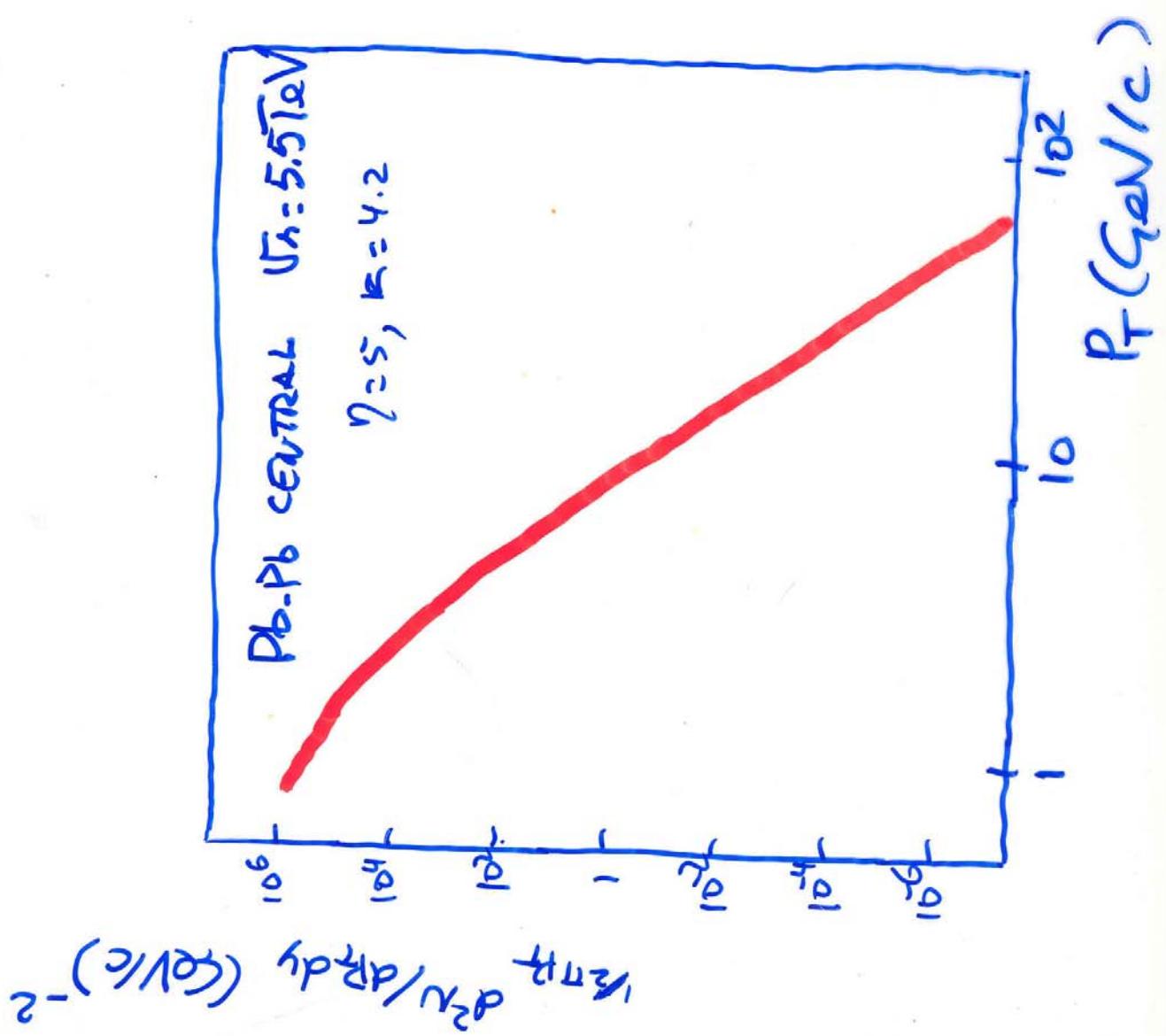
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phy/0304068



$h(p_T)/h(p_T^0)$

12-b



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3 - AA Collisions at RHIC (PHENIX) AND LHC.

$$1) \frac{1}{N_A} \bar{y}_3 \frac{dn}{dy} = F(\eta) \bar{\eta} \left(\frac{P_p}{R_p} \right)^2 \bar{n}$$

$$2) \langle p_T \rangle_i = \frac{nc(\eta)}{n(\eta)-2} \frac{\bar{P}_i}{F(\eta)}$$

$$3) \langle \eta \rangle_i = \bar{P}_i \left[\frac{1}{N_A^2} \frac{dn}{dy} \right]$$

$i = \pi, \kappa, \rho$

$$\gamma = \left(\frac{2\pi}{R_p} \right)^2 N_s^{\rho} N_A^2 \bar{y}_3$$

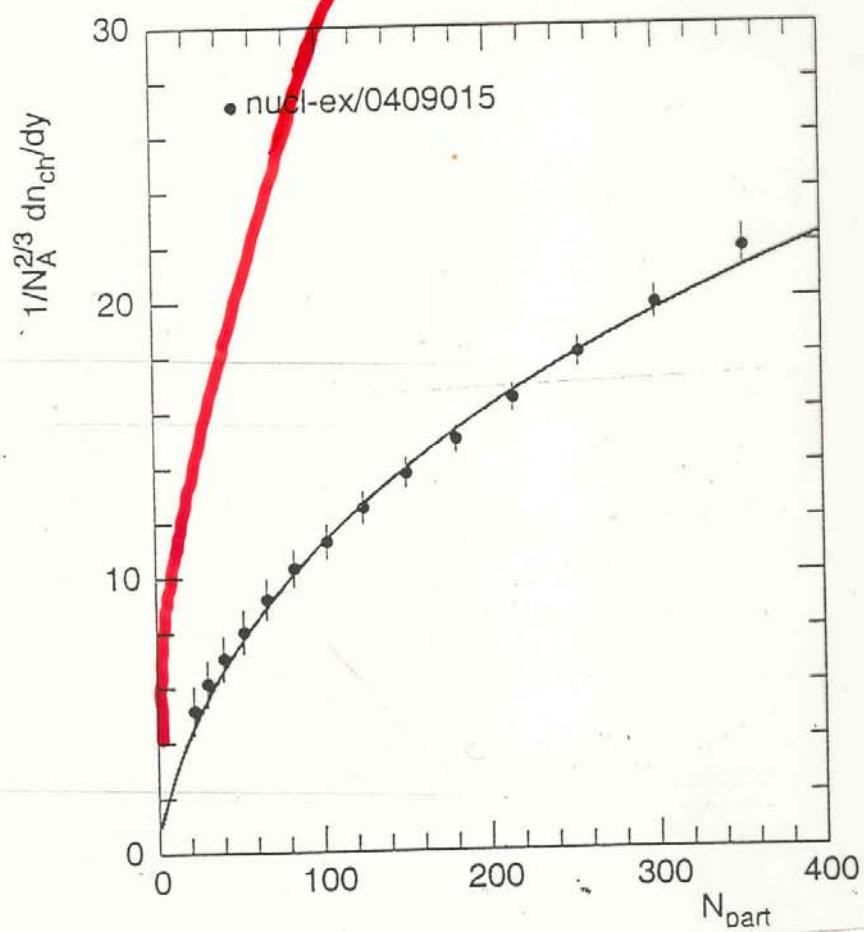
PARAMETERS : $n/R_p \approx 0.2$, $\bar{n} = 0.12$, $N_s^{\rho} (N_A = 200) = 8.0$,
 \bar{P}_i : $\bar{P}_{\pi} = 0.21 \text{ GeV/c}$, $\bar{P}_{\kappa} = 0.33$, $\bar{P}_{\rho} = 0.45$

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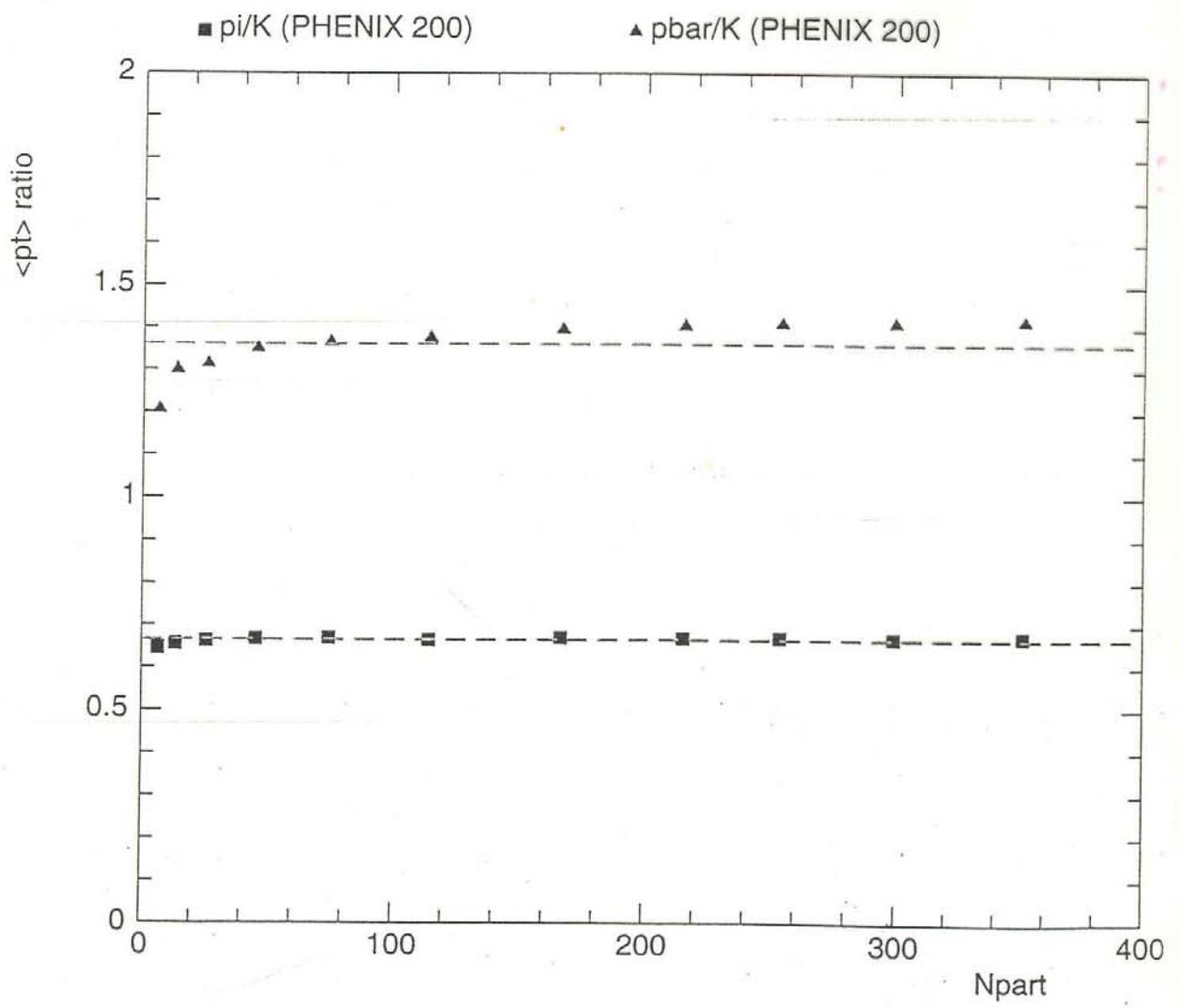
LHC

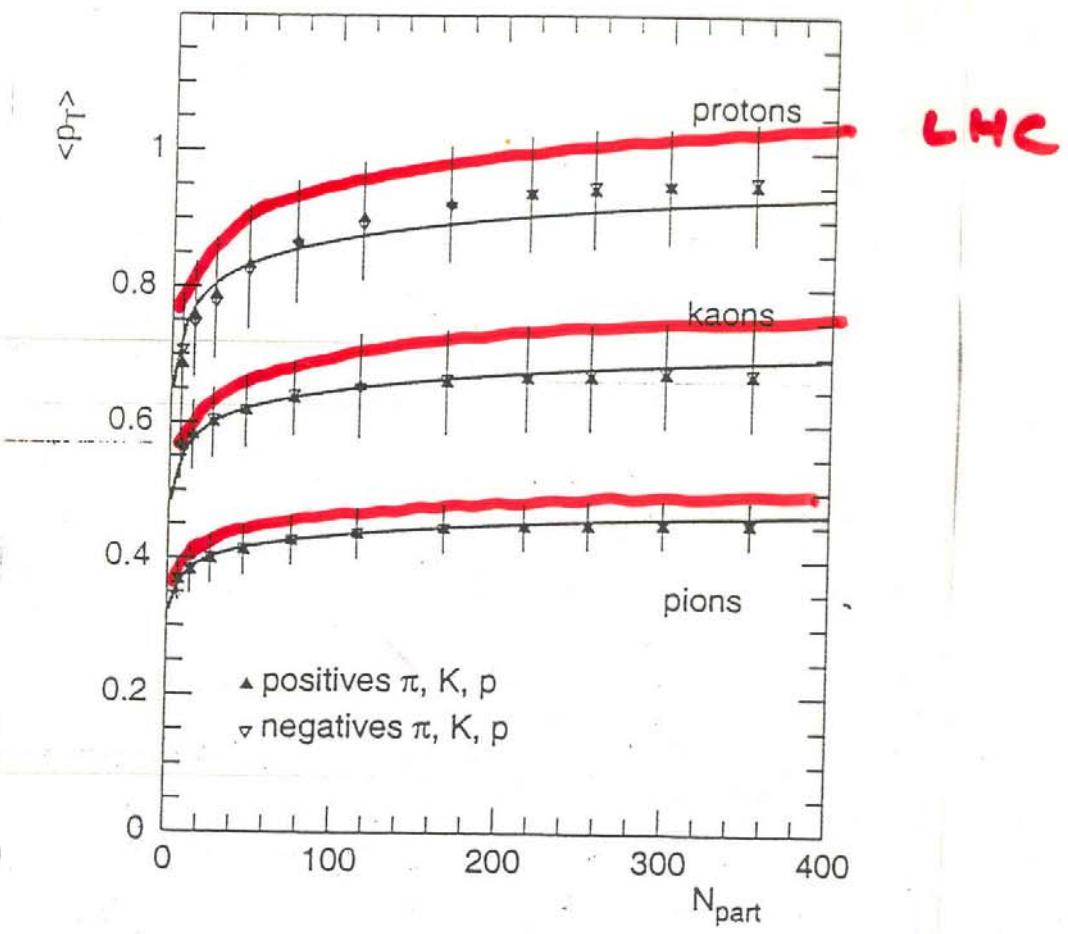
40 -

PHENIX 200 AGeV



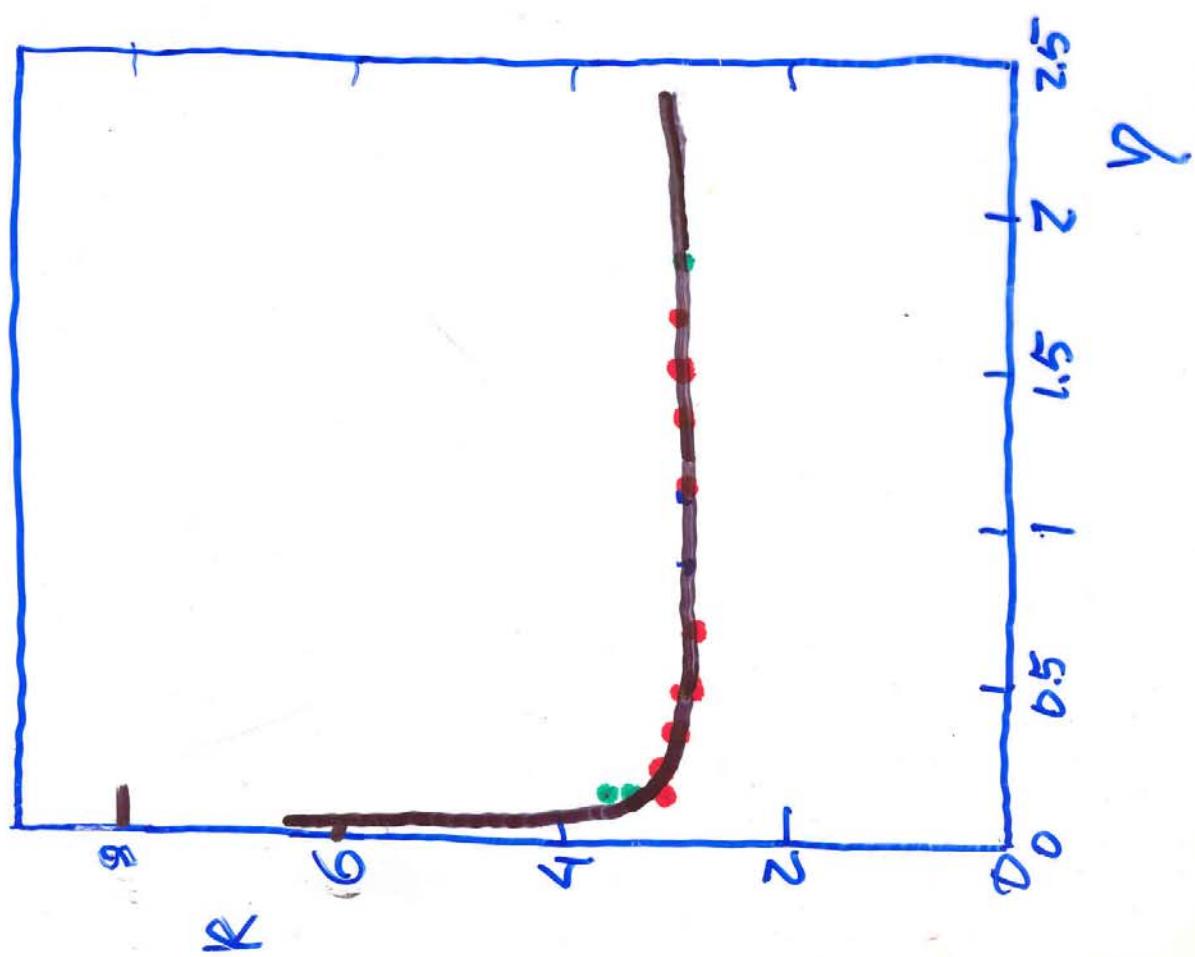
Au-Au





Au-Au

PHENIX



$$f(y_1, p_T) = \frac{dy}{dy_1} \frac{1}{\frac{\partial y}{\partial y_1}} F(y_1) \frac{R \cdot 1}{n} \frac{1}{(1 + \frac{F(y_1) P_T}{n})^R}$$

$$\frac{dy}{dy_1} = F(y_1) \bar{N}_s \bar{N}$$

$$y=0 \quad \left. \begin{array}{l} \bar{N}_s \sim N_A \\ \bar{N} \sim N_A \end{array} \right\} \text{Lyman } \alpha$$

$$R_{CP}(\gamma=0) = \frac{(\kappa'-1)/\kappa'}{\left(\kappa'-1\right)/\kappa} \left(\frac{F(\gamma')}{F(\gamma)} \right)^2 \left(\frac{\left(1 + \frac{F(\gamma)}{\kappa}\right) \frac{P_T^2}{P_T^2}}{\left(1 + \frac{F(\gamma')}{\kappa'}\right) \frac{P_T^2}{P_T^2}} \right)^{\kappa'/\kappa}$$

$$R_{CP} = \frac{(\kappa'-1)/\kappa'}{(\kappa'-1)/\kappa} \left(\frac{F(\gamma')}{F(\gamma)} \right)^2 < 1$$

$P_T = 0$

↑

R_{CP} small

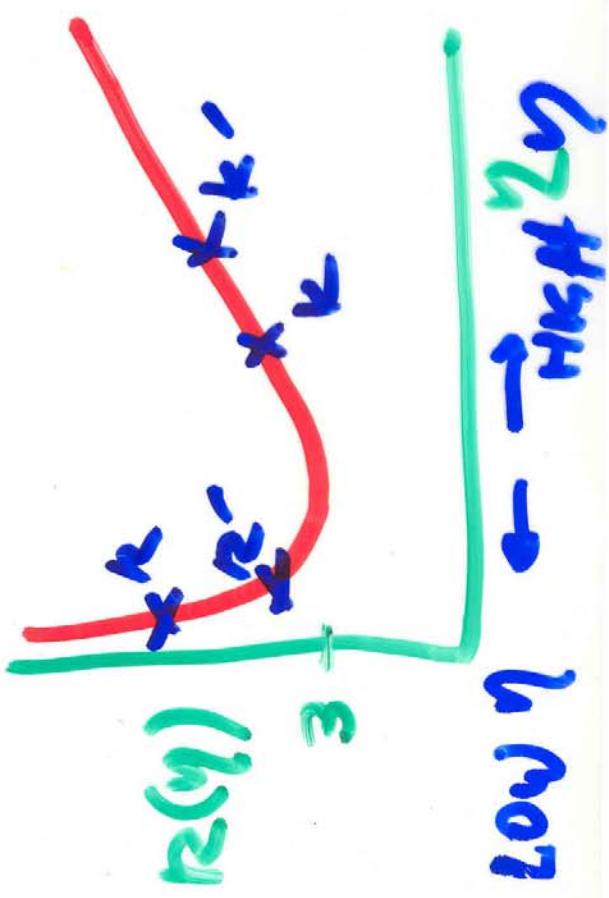
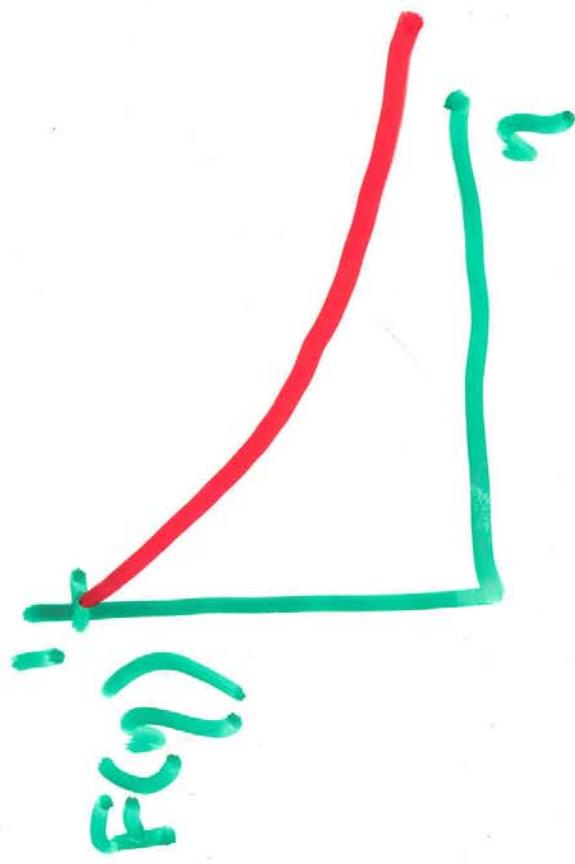
↑

$$R_{CP} \approx 1 + (F(\gamma) - F(\gamma')) \frac{P_T^2}{P_T^2}$$

1

η_{eq}

TO REMIND YOU ...



→ Is there a Maximum?

$$\frac{d \text{Rep}}{d p_T^2} = 0 \rightarrow [F(\eta) \cdot F(\eta')] - F(\eta) \bar{F}(\eta') \left[\frac{1}{k^2} - \frac{1}{k'^2} \right] p_T^2 = 0$$

High η : $k' > k$, **YES** : Rep decreases
Low η : $k > k'$, **NO** : Rep increases (> 1)

$$\rightarrow p_T \text{ very large}$$

$$R_{\text{CP}} \sim (p_T^2)^{k-k'}$$

$$\left. \begin{array}{l} k' > k \\ k > k' \end{array} \right\} \quad \left. \begin{array}{l} k > k' \\ k' > k \end{array} \right\}$$

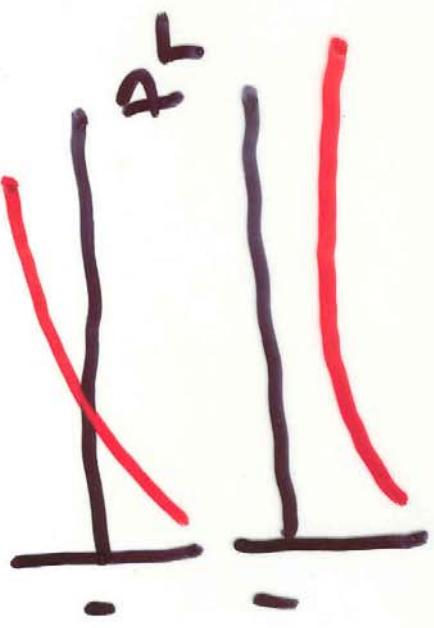
WHAT ABOUT pp collisions? ($N_{part} = 2$)

≥ 0

$R_{CP} \rightarrow R_{HL}$
(Central-peripheral) (High-low energy)

$$R_{HL} \equiv \frac{f(p_T, y, N_A, \sqrt{s}) / f(p_T, y, \sqrt{s})}{f(p_T, y, N_A, \sqrt{s}) / f(p_T, y, \sqrt{s})}$$

CROWN



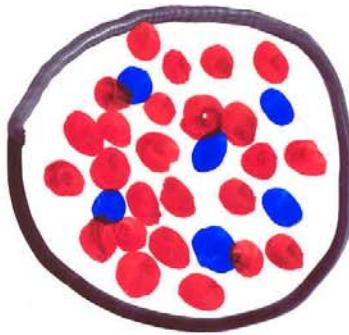
Low Energy
High Energy

LHC

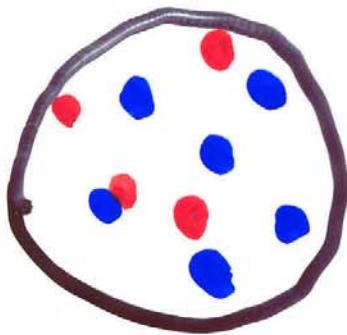
4) AWAY FROM $\gamma \approx 0$

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$$\gamma \approx 0$$



$$|\gamma| > 0$$



$$\gamma = \left(\frac{N}{R}\right)^2 N_s$$

$$\bar{N}_s \sim N_A$$

$$\bar{N}_s \sim \frac{N_A}{A}$$

$$R_{cap} \sim \frac{N_s}{N_{coll}} \sim 1$$

$$\sim \frac{1}{\lambda^3 A / 3}$$

$$\frac{R_{CP}(\gamma > \alpha)}{R_{CP}(\gamma = 0)} \sim \left(\frac{N_A}{N'_A} \right)^{\gamma_3}$$

IN GENERAL

EX:

d-Au

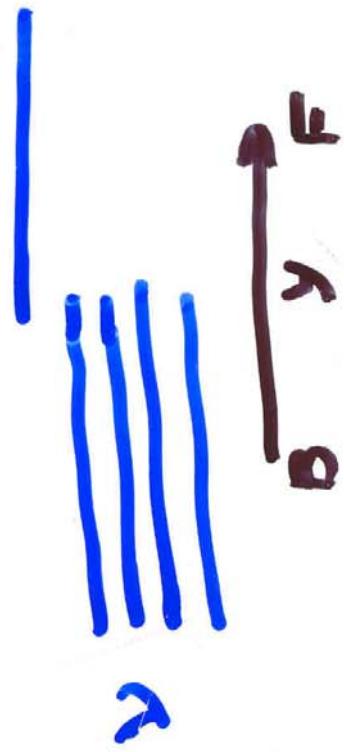
$$\frac{R_{CP}(P_T, \gamma = 3.2)}{R_{CP}(P_T, \gamma = 0)} \approx \frac{0.45}{1.25} = 0.36$$

$$\frac{R_{CP}(\gamma = 3.2)}{R_{CP}(\gamma = 0)} = \frac{(1.96 / 13.6)}{(1.39 / 3.3)} = 0.35$$

$$N'_d = 1.96, N'_{\text{coll.}} = 13.6$$

$$N_d = 1.39, N_{\text{coll.}} = 3.3$$

SIMPLIFIED VERSION FOR PAU (α_{Aee})



$N_{coll.} \approx 1$

$$R_{ep}(y=0) \approx \frac{6(1+\gamma)}{(1+\gamma)^2 + 4} = \frac{6}{5}, \quad R_{ep}(y=1) = \frac{6}{11}$$

$$R_{cp}(y=0) \approx \frac{(1-\gamma)}{(1+\gamma)} = \frac{1}{2}, \quad R_{cp}(0)$$

$$R_{cp}(y=1) = \frac{1}{11} > R_{cp}(0)$$

$$R_{PA} = f_{PA}^{(p_T)} / \langle \zeta_u \rangle f_{pp}^{(p_T)}$$

MB

$$\frac{f_{PA}^{(\gamma, \zeta_u)}}{f_{pp}^{(\gamma, \zeta_u)}} < R_{PA}^{(co)}$$

$$R_{PA}^{(\gamma, \zeta_u)} \sim \frac{1}{\zeta_u} < R_{PA}^{(co)}$$

$$R_{PA}^{(\gamma, \zeta_u)} \sim \frac{\zeta_u}{\zeta_u^2} < R_{PA}^{(co)}$$

$$\left\{ \frac{d\eta}{dy} = F(\eta) \right. \quad \bar{N}_s \bar{\eta}$$

$$\langle \hat{P}_T^2 \rangle = \bar{P}_T^2 \frac{\frac{d\eta}{dy}}{\eta(0) - \eta} \frac{1}{F(\eta)}$$

$$\bar{N}_s = \gamma \left(\frac{1}{\pi} \right) \eta^2 = \eta \left(\frac{R_p}{\pi} \right)^2 N_A^{1/3}$$

• REL. $\frac{d\eta}{dy}$ w. $\langle \hat{P}_T \rangle$ is the same

$$\eta = 0$$

$$|\eta| > 0$$

$$\sim N_A^{1/6}$$

$$\frac{1}{N_A^{2/3}} \frac{d\eta}{dy} \sim N_A^{1/3}$$

$$\langle \hat{P}_T \rangle \sim N_A^{1/6} \sim N_A^{1/2} \eta^{-1/2}$$

$$\sim N_A^{1/2}$$

JET QUENCHING

$\gamma > 0$

EFFECT



CONCLUSIONS

- IN THE TRANSITION TO GCP PERCOLATION EFFECTS SHOULD OCCUR
- GENERAL UNDERSTANDING OF $\langle n/dy \rangle, \langle p_T \rangle, f(p_T, y), P_{CP}, P_{PA}$
- WE EXPECT THE p_T DISTRIBUTION IN PP COLLISIONS TO BECOME WALKER AT HIGHER ENERGY (LHC^2)