

Hard Probes 2004: International Conference on
Hard and Electromagnetic Probes of High Energy Nuclear Collisions
Ericeira, Portugal, November 8th 2004

Flow effects on jet profiles and multiplicities

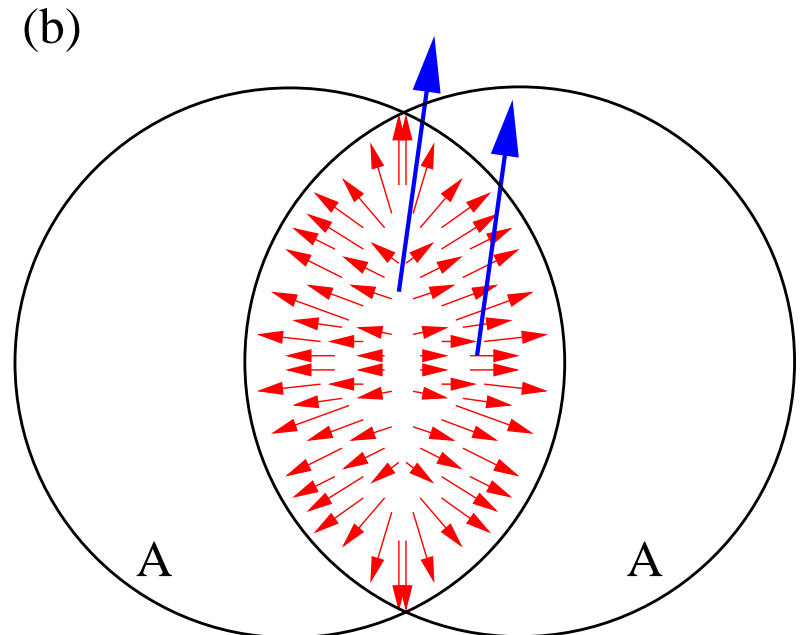
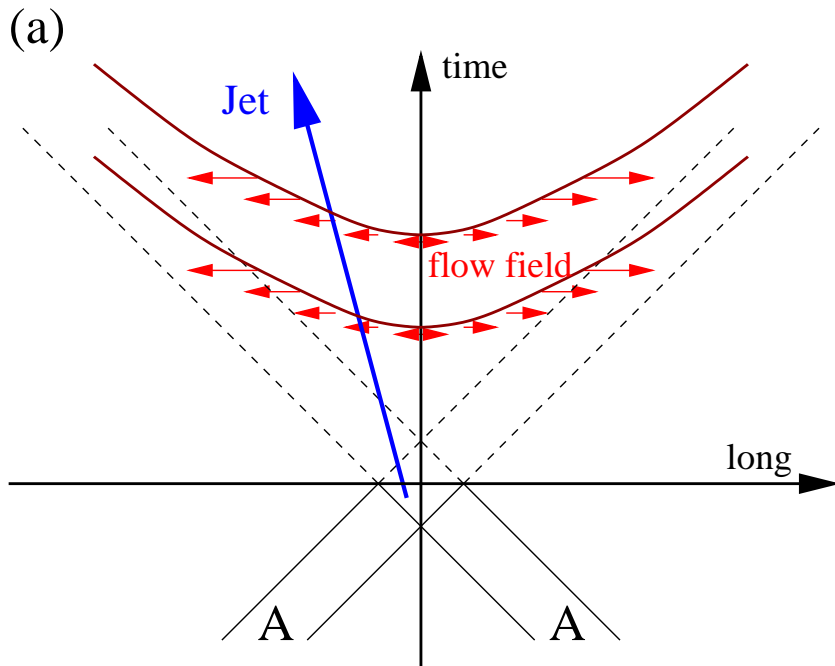
Néstor Armesto

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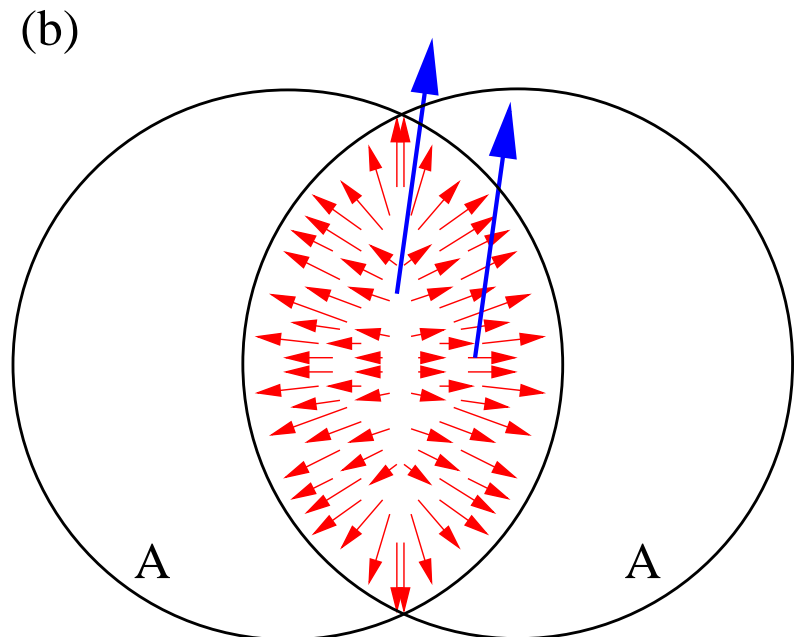
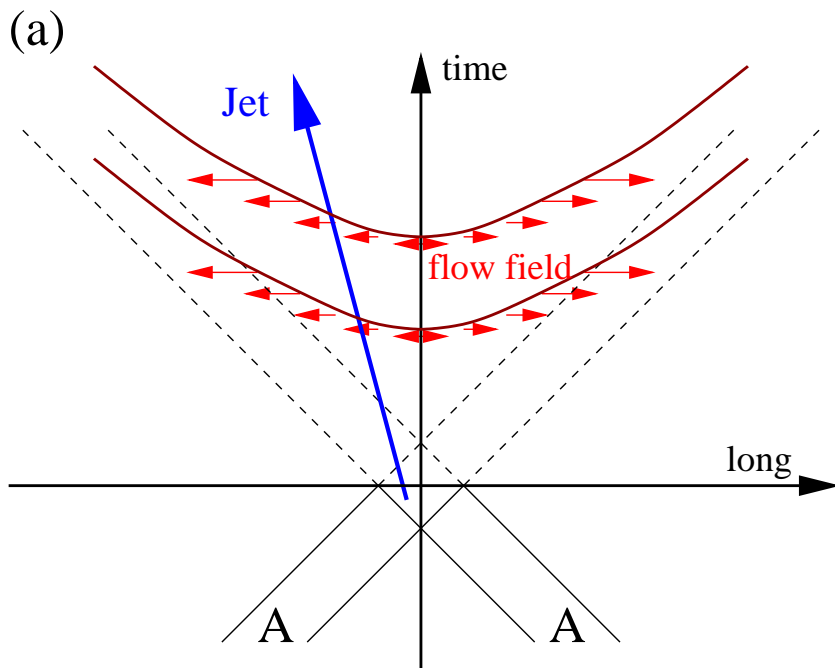
1. Motivation and formalism.
2. Exercises:
 - LHC: jet shapes.
 - RHIC: widths of particle distributions.
 - RHIC: elliptic flow.
3. Summary.

With C.A. Salgado and U.A. Wiedemann, hep-ph/0405301 (PRL) and in preparation.

- Strong momentum-position correlations in the expanding medium are suggested by the success of hydro at low p_T : **collective flow**.

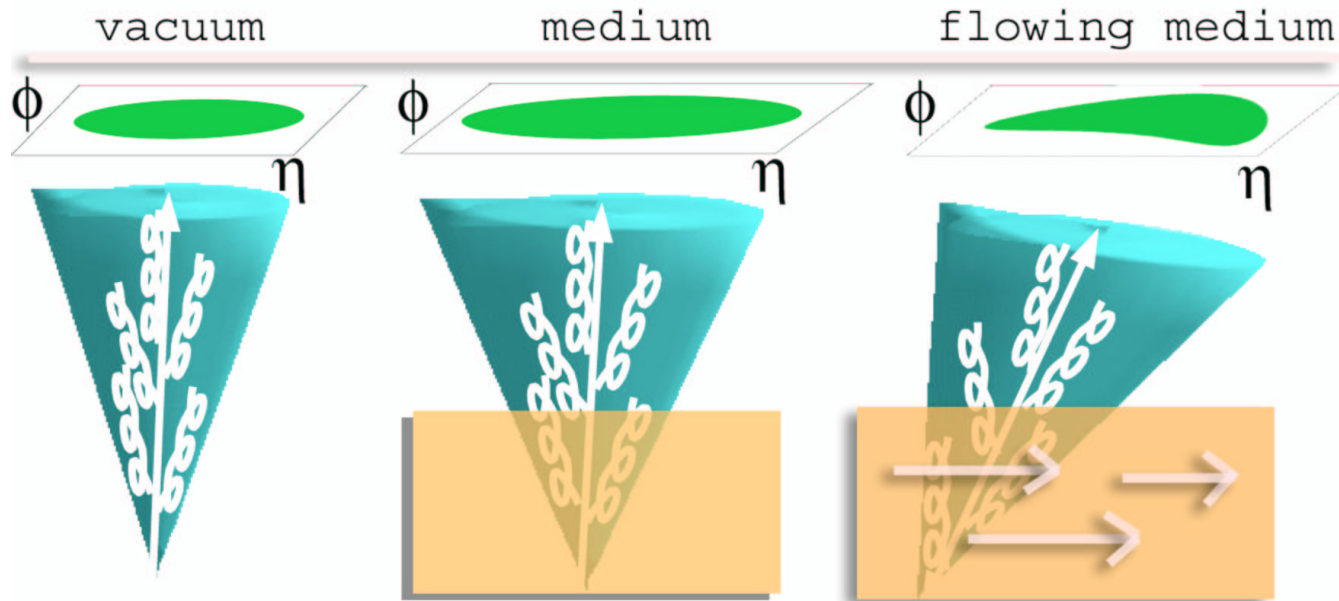


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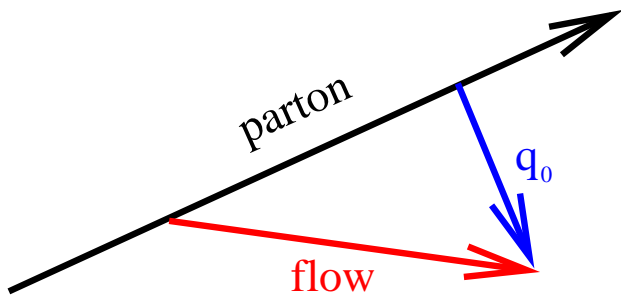
- Radiative energy loss is determined by momentum exchanges perpendicular to the trajectory of the parton.
- **Idea:** if the jet is produced in a frame not co-moving with the collective flow, **momentum exchanges become anisotropic and an additional contribution to energy loss comes from flow.**

$$T^{\mu\nu}(x) = (\epsilon + p) u^\mu u^\nu - p g^{\mu\nu}, \quad u^\mu = \gamma(1, \vec{\beta}).$$

- **Estimation:** $T^{ii*} = p \rightarrow T^{ii} = p + \Delta p$, where
 $\Delta p = (\epsilon + p)u^i u^i = 4p\gamma^2\beta^2$ (ideal EOS $\epsilon = 3p$) \rightarrow rapidity difference
 $\eta = 0.5, 1.0, 1.5$ between frames $\implies \Delta p/p \simeq 1, 5, 18$.

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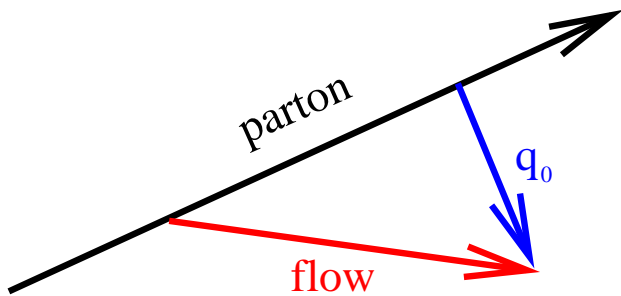
$$|a(\mathbf{q})|^2 = \frac{\mu^2}{\pi [(\mathbf{q} - \mathbf{q}_0)^2 + \mu^2]^2}.$$

- We modify the Yukawa-like scattering potential. We consider (Baier '02)

$$\hat{q} = \frac{\mu^2}{\lambda} \propto n\sigma, \quad \hat{q} [\text{GeV}^2/\text{fm}] = c\epsilon^{3/4} [(\text{GeV}/\text{fm}^3)^{3/4}] \Rightarrow q_0 \sim \mu.$$

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- Dilution of the medium (Baier, Dokshitzer, Mueller, Schiff, '98; Gyulassy, Vitev, Wang, '00) can be taken into account by rescaling \hat{q} (Salgado, Wiedemann, '02):

$$\langle \hat{q} \rangle = \frac{2}{L^2} \int d\tau \tau \hat{q}(\tau).$$

- In the single hard scattering approximation (Wiedemann, '00; Gyulassy, Levai, Vitev, '00),

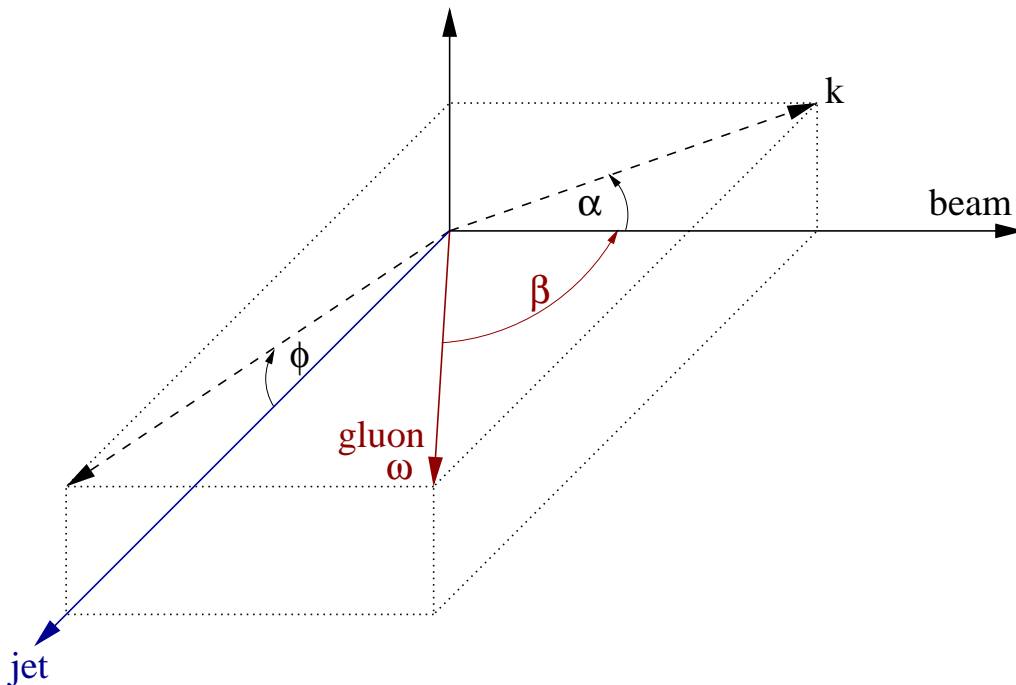
$$\omega \frac{dI^{\text{med}}}{d\omega d\mathbf{k}} = \frac{\alpha_s}{(2\pi)^2} \frac{4C_R n_0}{\omega} \int d\mathbf{q} |a(\mathbf{q})|^2 \frac{\mathbf{k} \cdot \mathbf{q}}{\mathbf{k}^2} \frac{-L \frac{(\mathbf{k}+\mathbf{q})^2}{2\omega} + \sin\left(L \frac{(\mathbf{k}+\mathbf{q})^2}{2\omega}\right)}{[(\mathbf{k} + \mathbf{q})^2 / 2\omega]^2}.$$

Similar results (Salgado, Wiedemann, '03) in BDMPS (Baier, Dokshitzer, Mueller, Peigné, Schiff, '96), $\sigma(\mathbf{r}) = 2 \int d\mathbf{q} |a(\mathbf{q})|^2 (1 - e^{-i\mathbf{q}\cdot\mathbf{r}})$, $n(\tau) \sigma(\mathbf{r}) \simeq \frac{1}{2} \hat{q}(\tau) \mathbf{r}^2$.

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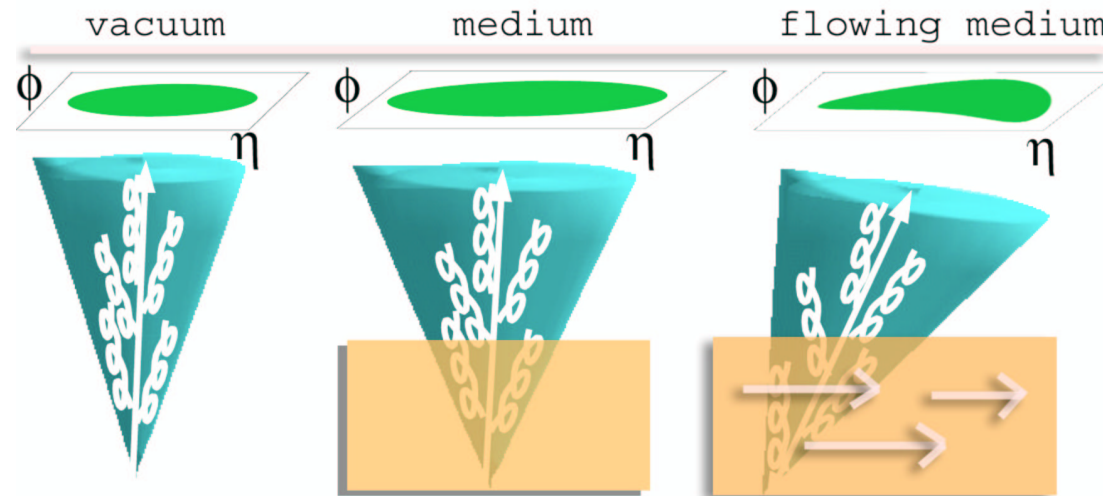


$$k dk d\alpha = \omega^2 \frac{\cos \phi}{\cosh^3 \eta} d\eta d\phi$$

(it already induces an $\eta - \phi$ -asymmetry).

Jet shapes with flow (I):

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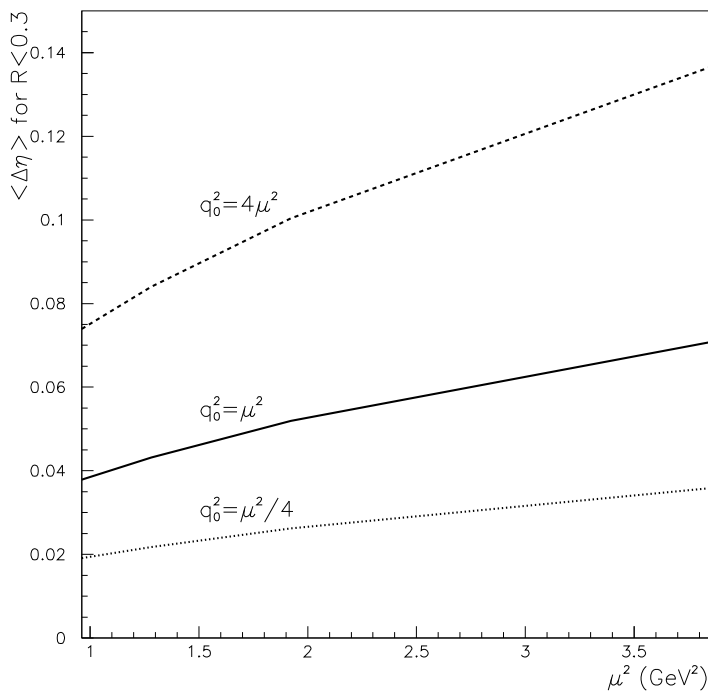
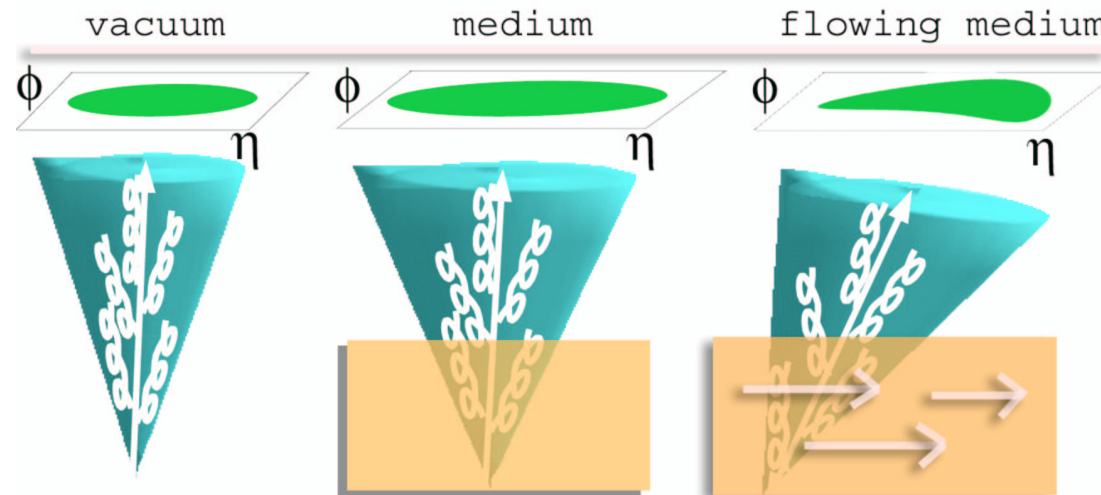
- For the vacuum, $R = \sqrt{\eta^2 + \phi^2}$,

$$\rho_{\text{vac}}(R) \equiv \frac{1}{N_{\text{jets}}} \sum_{\text{jets}} \frac{E_T(R)}{E_T(R=1)}$$

taken from D0 (LHC: pp and pA references needed).

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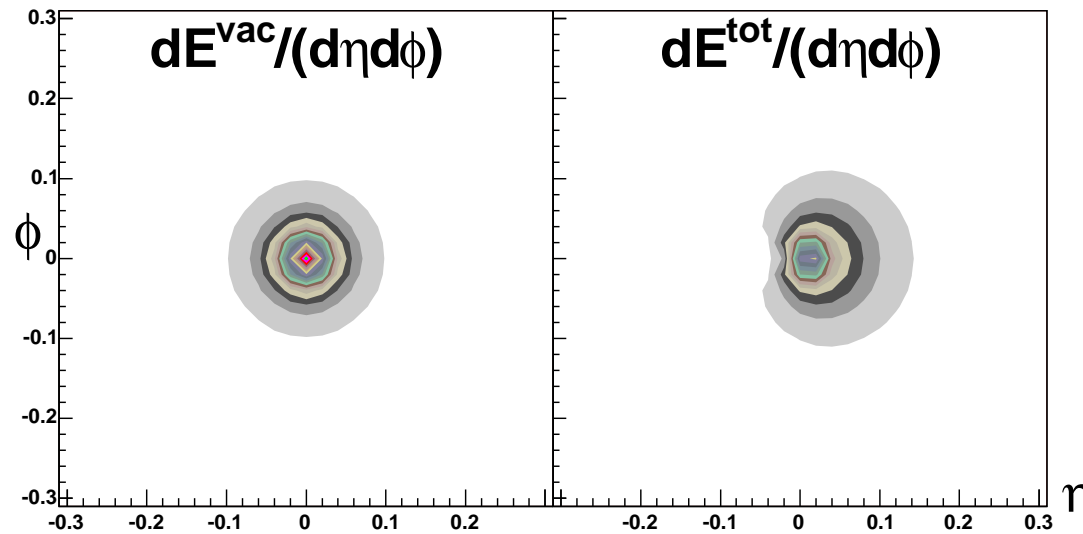
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- $n_0 L \alpha_s C_R = 1$, $L = 6 \text{ fm} \implies$

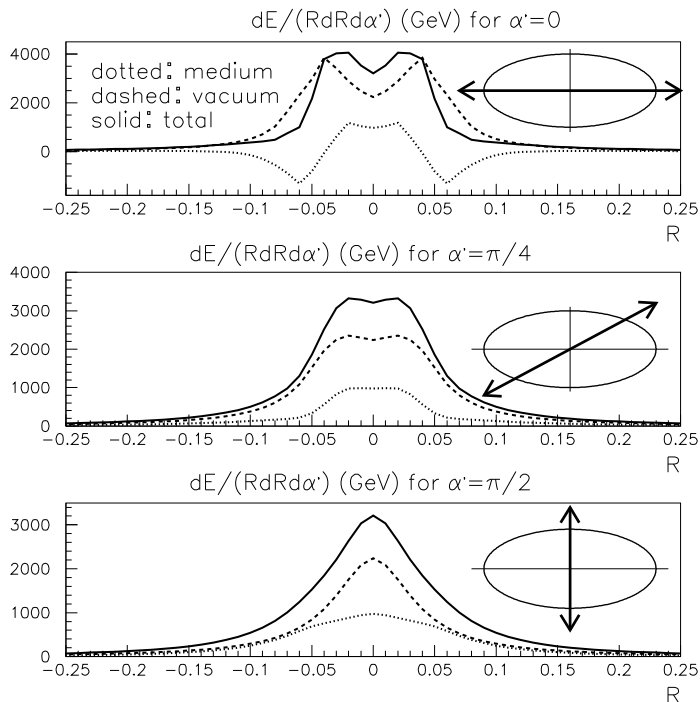
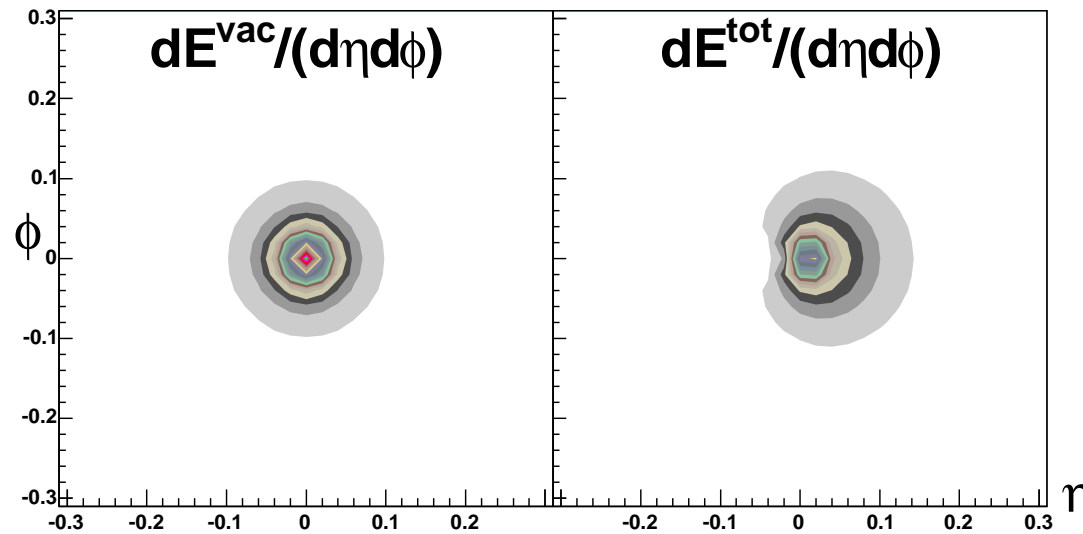
the jet center is not too displaced.

$\rightarrow \bar{p}p$: $R = 0.7 \div 1$ ($E_{\text{within}} \simeq 100\%$).

$\rightarrow \text{PbPb at LHC}$: $R = 0.3 \div 0.4$ ($E_{\text{within}} \simeq 75 \div 80\%$).



- 100 GeV jet; $\mu = 1$ GeV, $q_0 = \mu \implies \langle \Delta\eta \rangle_{R<0.3} \simeq 0.04$, $\Delta E_T = \int d\omega \omega \frac{dI^{\text{med}}}{d\omega} = 23$ GeV redistributed by medium-induced gluon radiation.



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- **Clear asymmetry.** (If) flow comes equally from both $+\eta$ and $-\eta$, **width in η becomes larger than in ϕ** ($d\eta d\phi = R dR d\alpha'$).

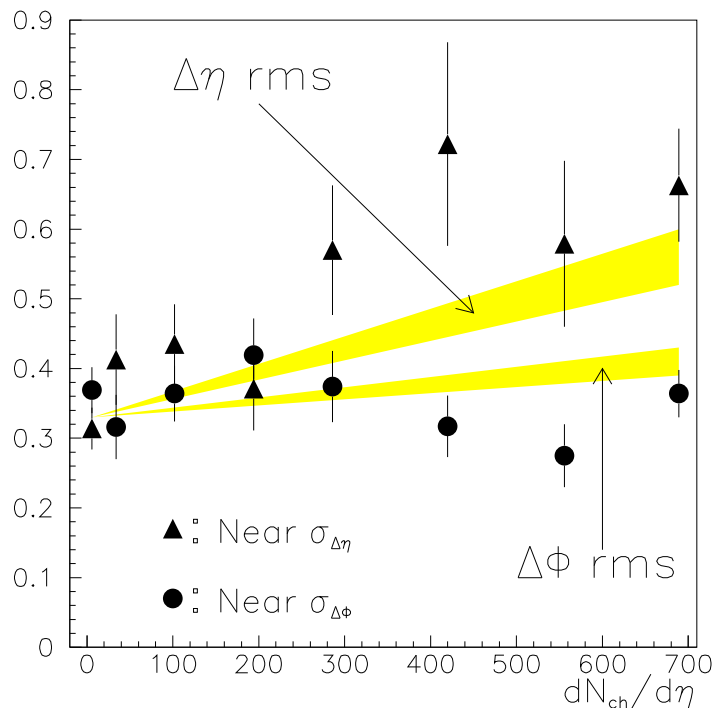
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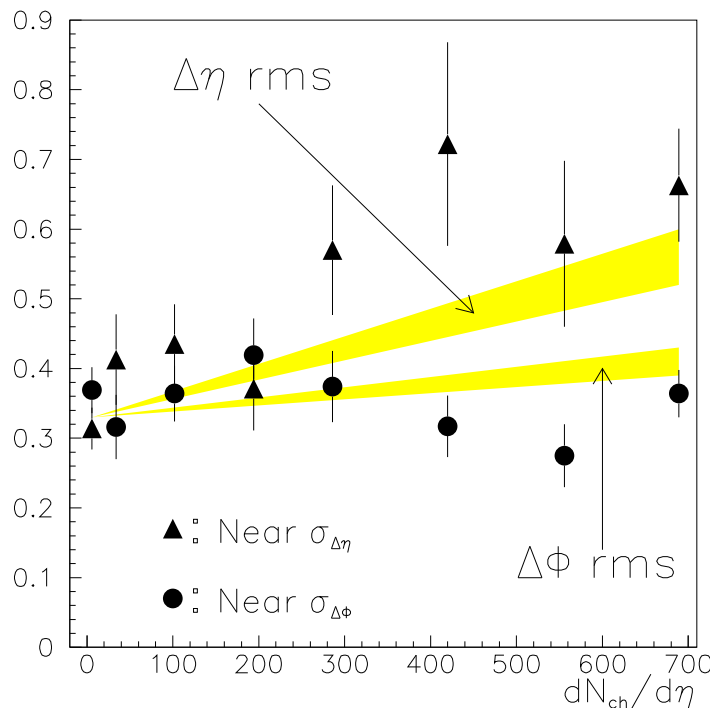


- We take a 10 GeV parton, $L = 2$ fm (near side jet), $\mu = 0.7 \div 1.4$ GeV, and $q_0/\mu = 4 \div 2$.
- Preliminary STAR data: F.Wang at QM04, near-side charged distribution associated to trigger with $4 \text{ GeV} < p_T^{\text{trigger}} < 6 \text{ GeV}$ in AuAu@200 GeV.

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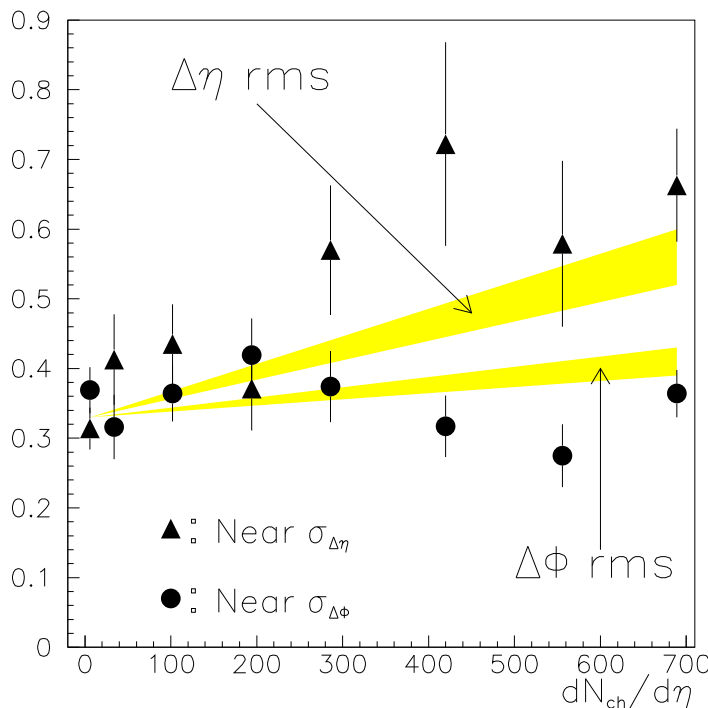
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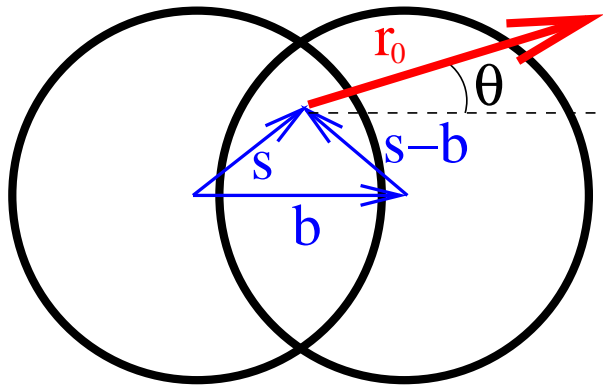
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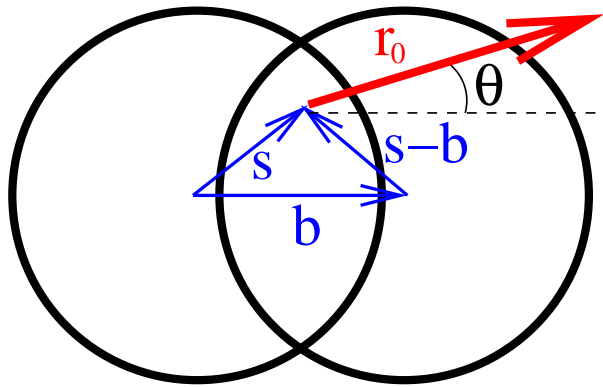


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- With radiative E -loss, the widths are expected to increase with decreasing $p_T = \omega$: $\sin^2 \theta = k_T^2/\omega^2 \simeq \sqrt{\hat{q}\omega}/\omega^2$, and radiation is harder than in vacuum.

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- Parton produced at (x_0, y_0) according to $T_A(\mathbf{s}) T_B(\mathbf{b} - \mathbf{s})$, with trajectory $\mathbf{r}_0(\xi) = (x_0 + \xi \cos \theta, y_0 + \xi \sin \theta)$ uniform in θ (Gyulassy, Vitev, Wang, '00).



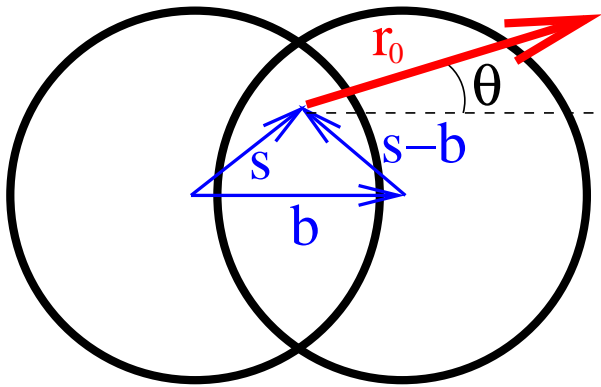
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with $\Omega(\mathbf{r})$ a time-dependent density distribution of produced matter.



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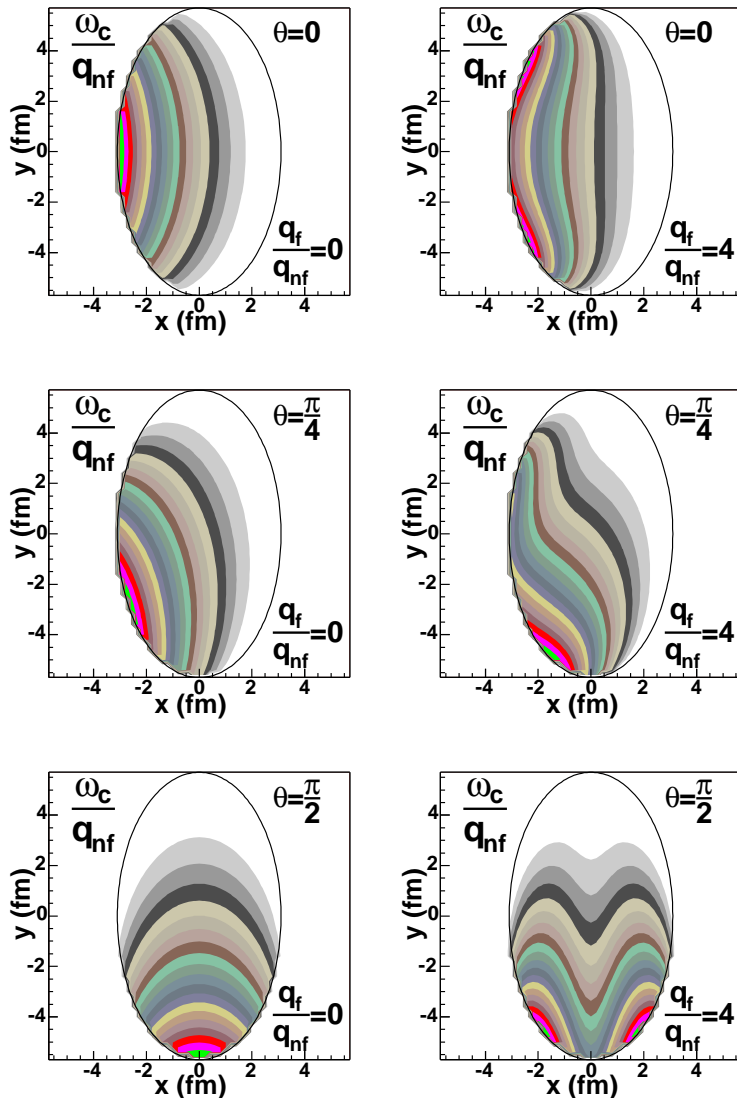
- For a given flow field (Lisa, Retière, '03), we take the **simple ansatz**

$$\hat{q}(\xi) = \bar{\hat{q}}_{nf} + \bar{\hat{q}}_{flow} |\mathbf{u}(\mathbf{r}_0(\xi)) \cdot \mathbf{n}_T|^2, \quad \mathbf{n}_T \cdot \mathbf{r}_0(\xi) = 0.$$

Contribution to elliptic flow (II):

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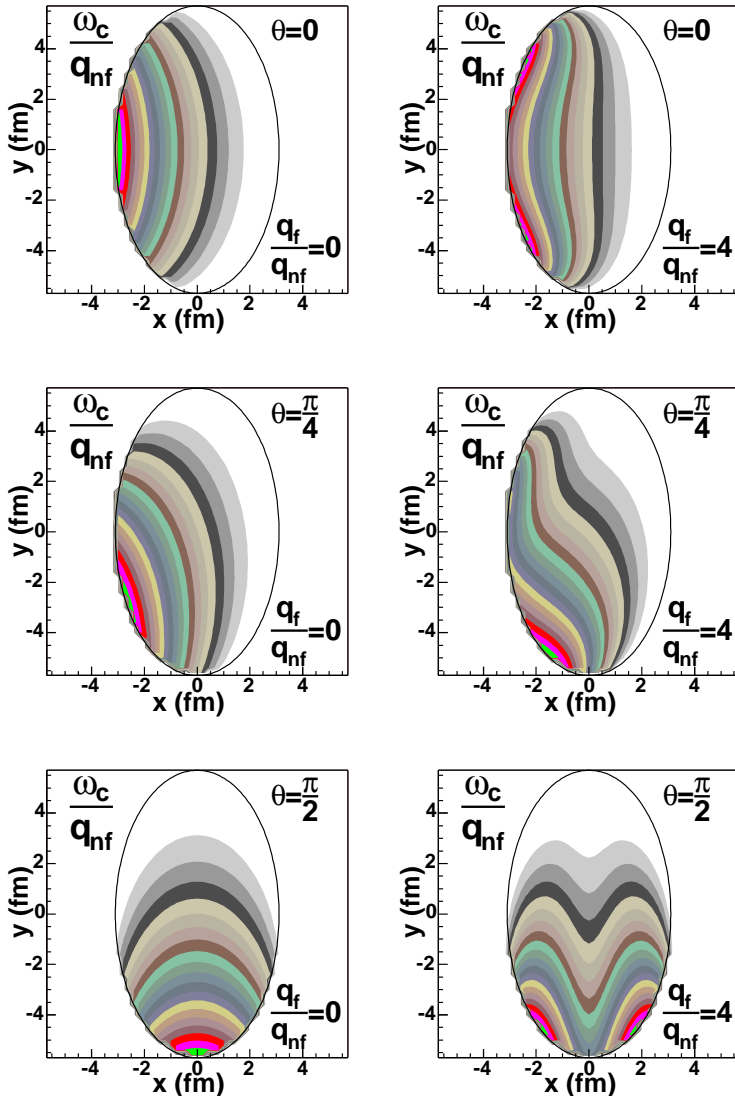
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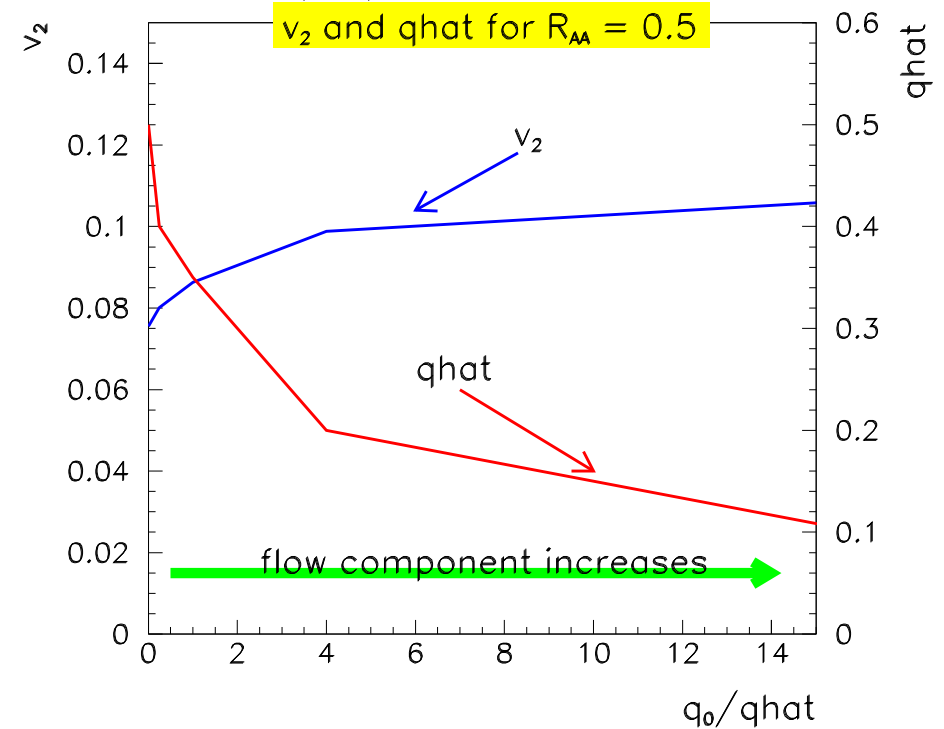


- Exercise for semi-peripheral AuAu:

$$N(x_0, y_0, \theta, p_T) = \frac{d\sigma^{med}}{dp_T} \bigg/ \frac{d\sigma^{vac}}{dp_T}, \quad p_T =$$

5 GeV/c (Salgado, Wiedemann, '03),

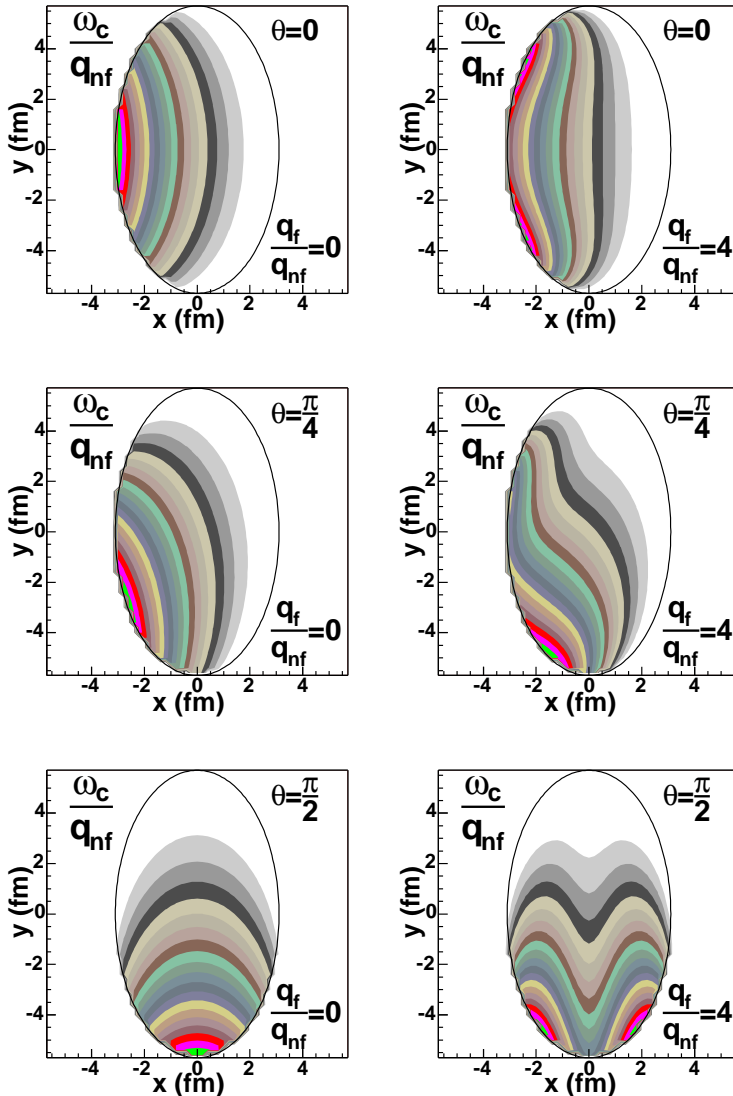
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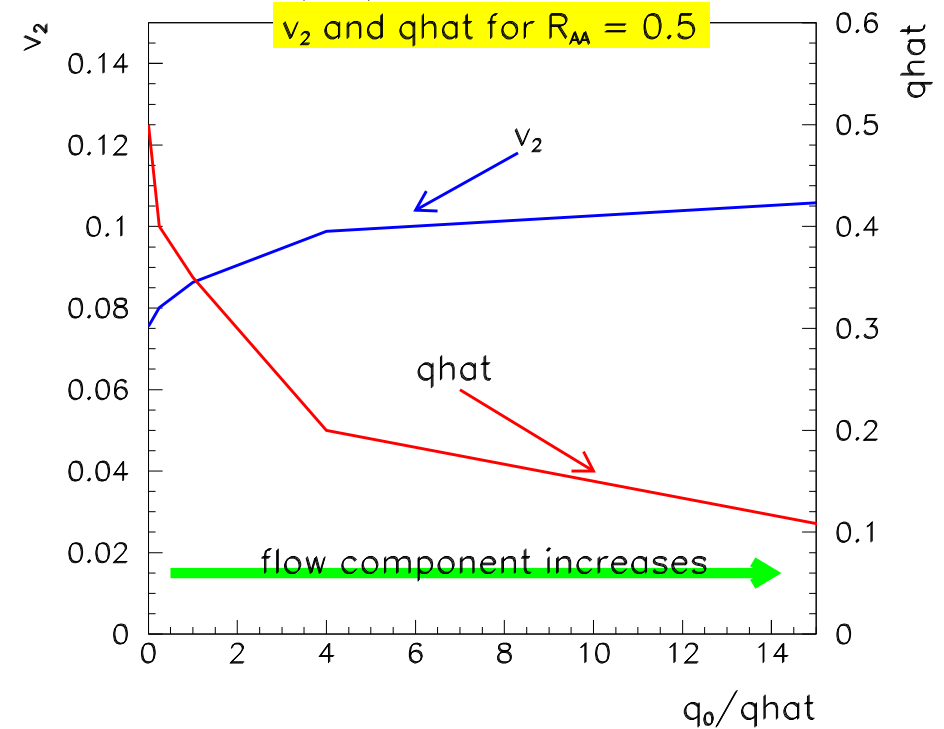


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- Effect on v_2 is not large, (Wang, '03; Drees, Feng, Jia, '03); flow effects may mimic a higher density.

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BUT

even a negative result provides information about the space-time evolution of the system (hard production coupled to the flow?) \longrightarrow **compute it within a full hydrodynamical simulation** (Hirano, Nara, '02; '03).