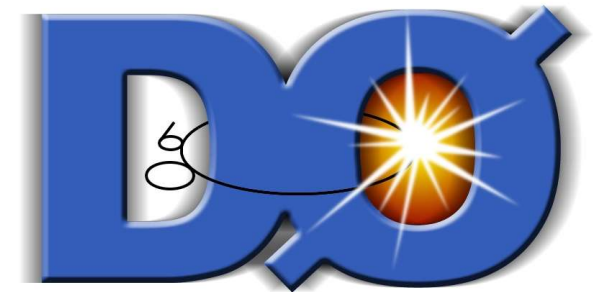


# $\tau$ -ID at D0 and ATLAS

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Michael Heldmann,  
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University of Freiburg

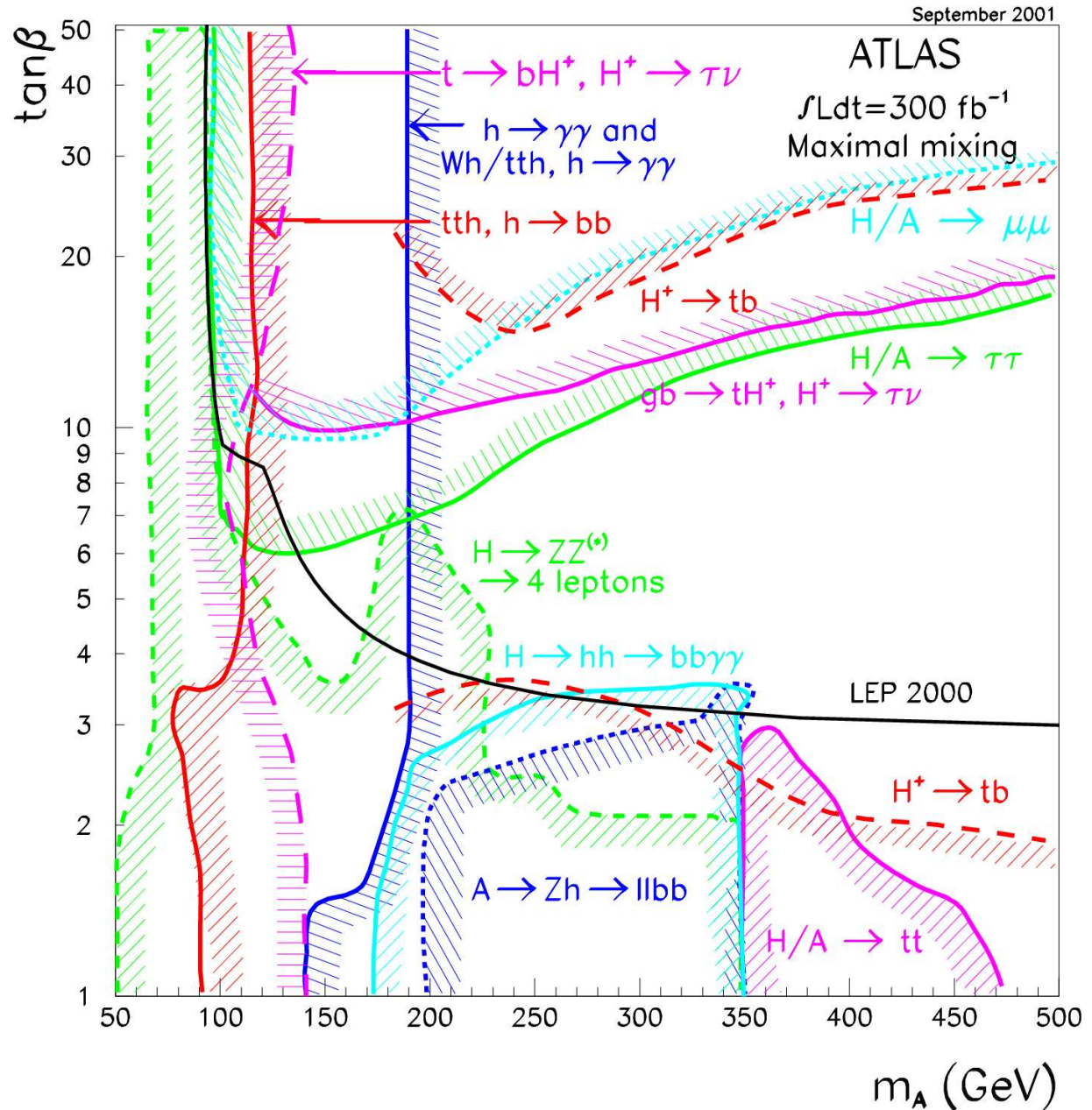


- Motivation
- D0
- ATLAS
- preliminary comparison
- summary



# Motivation

- At ATLAS we expect a big number of final states involving taus
- Channels using taus
  - $A^0/H^0 \rightarrow \tau\tau$
  - $H^\pm \rightarrow \tau\nu$
  - SUSY with production of  $\tilde{\tau} \rightarrow \tau + \chi^0_1$
  - Standardmodell Higgs (VBF  $qq \rightarrow qq \tau\tau$ )
  - $Z \rightarrow \tau\tau$  (for commissioning)
  - $\tau$  are perhaps the only way to access the chiral structure of SUSY
- $\rightarrow\tau$ 's are an important signature



# Motivation

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- Since this is the TeV4LHC workshop, the questions are:
  - What can ATLAS learn from D0 about tau reconstruction and identification ?
  - How can we transfer this knowledge to ATLAS ?
- The steps we would like to follow are:
  - compare the D0 algorithm to what we use at ATLAS
  - look for input on how we can improve our algorithm
- many ATLAS Analysis rely on the understanding of tau identification
  - will we reach the performance we see on the MC at the moment ?
  - → learn from the D0 comparison between MC and data
- check if the description of MC-Generators of the low multiplicity jets is correct with D0 data
- get input on how to measure the performance using data

# Tau Identificaton

- How can one identify  $\tau$ -leptons ?

- most important decay modes

- Leptonical decay modes

- $\tau \rightarrow \nu_\tau + \nu_e + e$  (17.4%)

- $\tau \rightarrow \nu_\tau + \nu_\mu + \mu$  (17.8%)

- Hadronical decay modes

- 1 prong

- $\tau \rightarrow \nu_\tau + \pi^c$  (11.0%)

- $\tau \rightarrow \nu_\tau + \pi^c + \times \pi^0$  (36.2%)

- 3 prong

- $\tau \rightarrow \nu_\tau + 3 \cdot \pi^c + \times \pi^0$  (15.2%)

- 

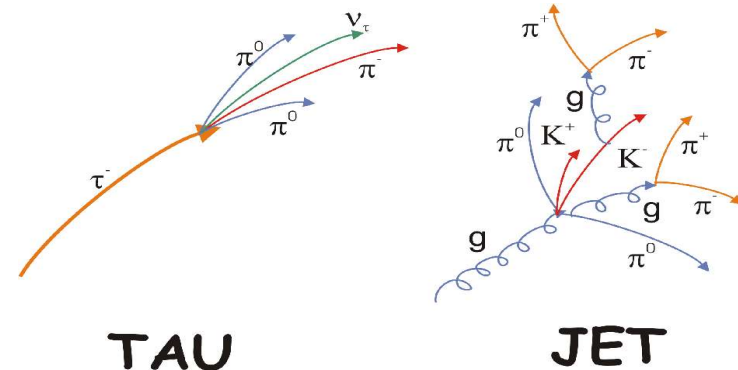
1 track

only difference  
from prompt leptons:  
impact parameter

1 track, impact parameter  
shower shape, energy sharing  
find the photon cluster

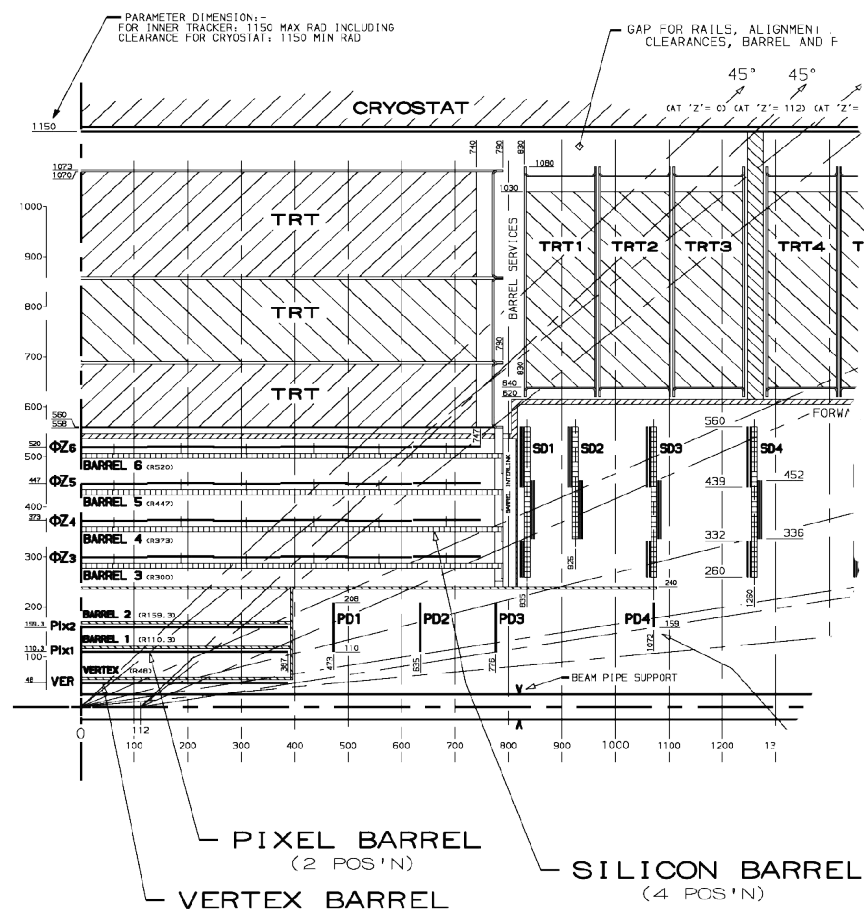
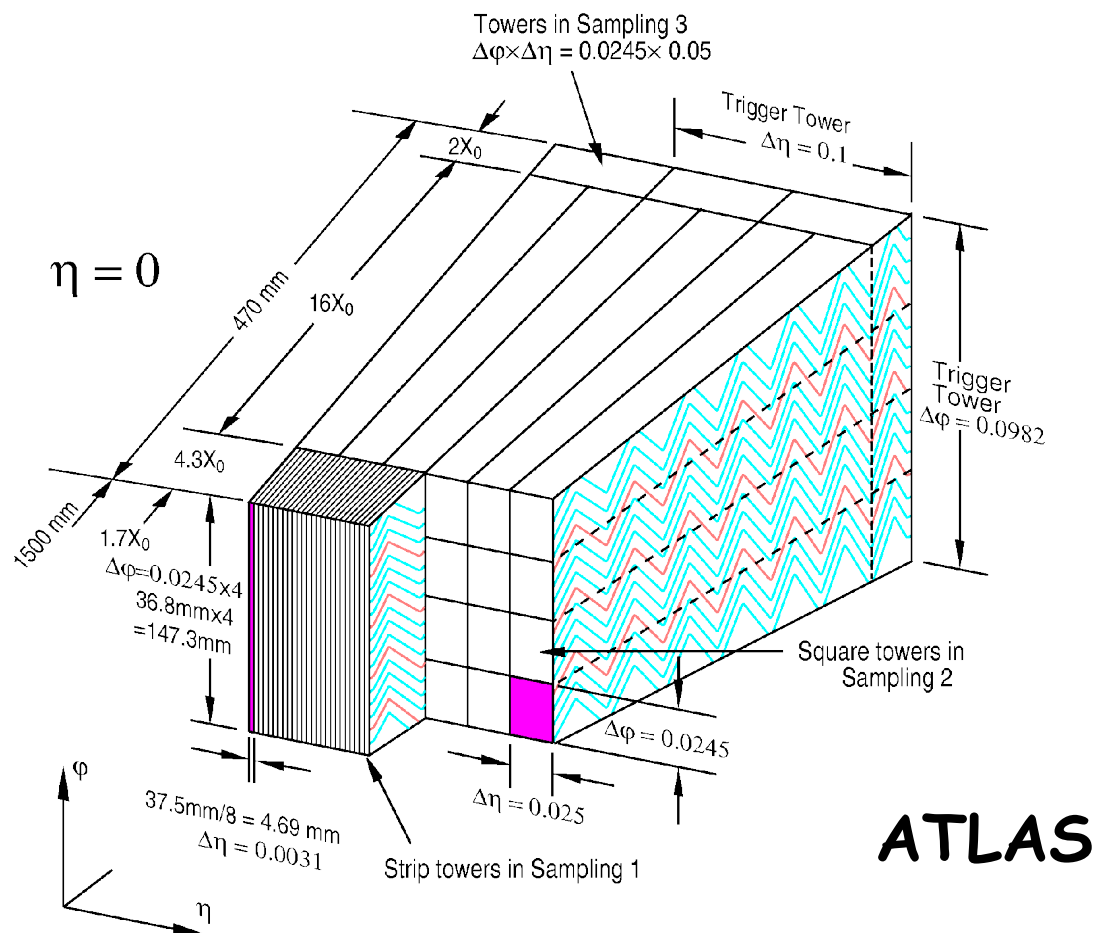
3 track, impact parameters,  
secondary vertex

- $\rightarrow \tau$ s are colimated calorimeter objects with one or three associated tracks



# ATLAS Calorimeter and ID

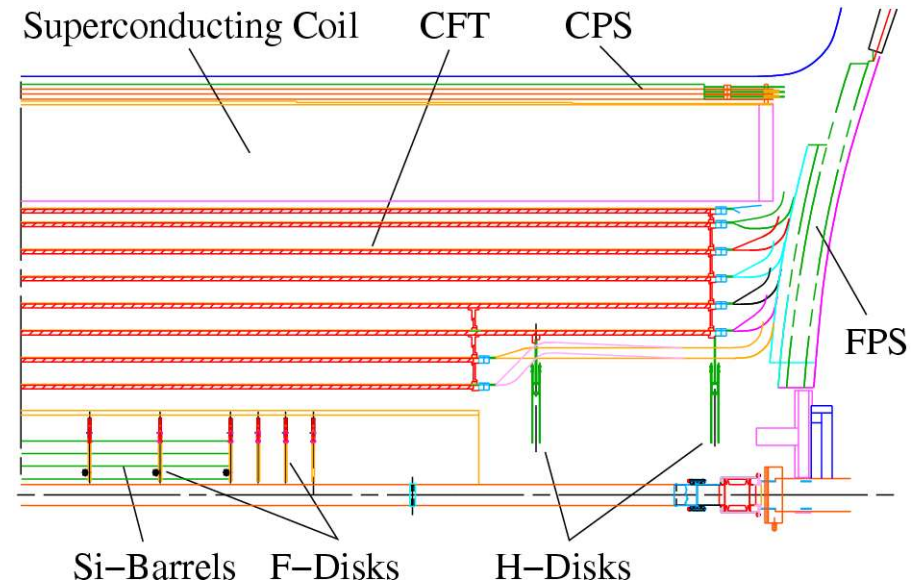
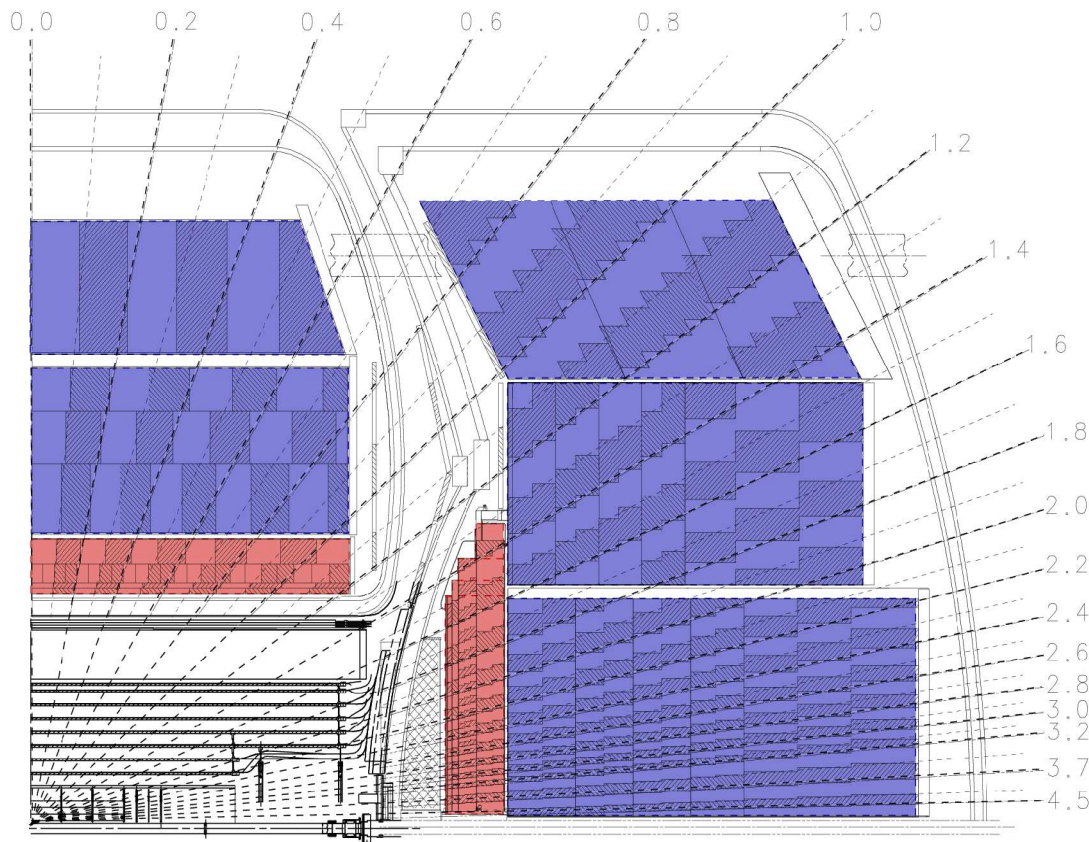
- The tau identification makes use of tracks and calorimeter objects
- Atlas has a presampler (0.025x0.1), eta-strip-layer (0.003x0.1), middle (0.025x0.025), back layer (0.05x0.025) and three hadronical layers with 0.1x0.1 and 0.2x0.1
- The ID has three pixel layers, four stereo microstrip layers and a straw tube tracker



# D0 Calorimeter and ID

- D0 has an EM with four layers of  $0.1 \times 0.1$ ,  $0.1 \times 0.1$ ,  $0.05 \times 0.05$ ,  $0.1 \times 0.1$  and four hadronical layers  $0.1 \times 0.1$
- The ID consists of a silicon tracker with four stereo layers and eight stereo layers for the fiber tracker
- the ID of D0 covers  $|\eta| < 3.0$  !

## D0

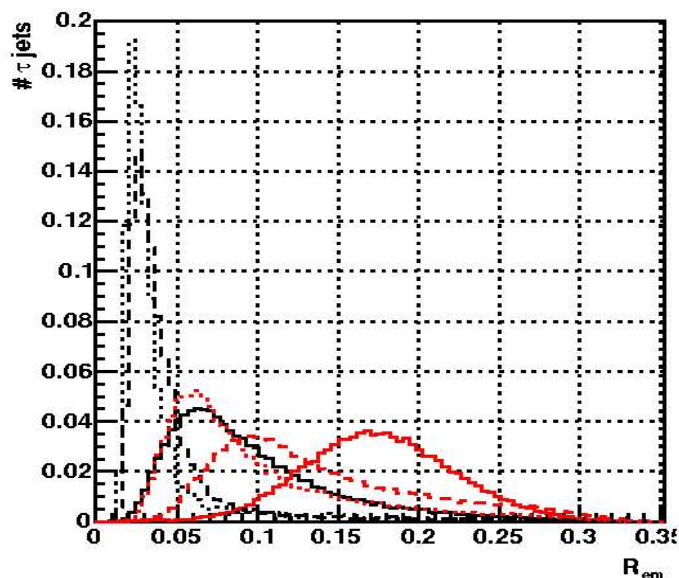


# TauID at ATLAS

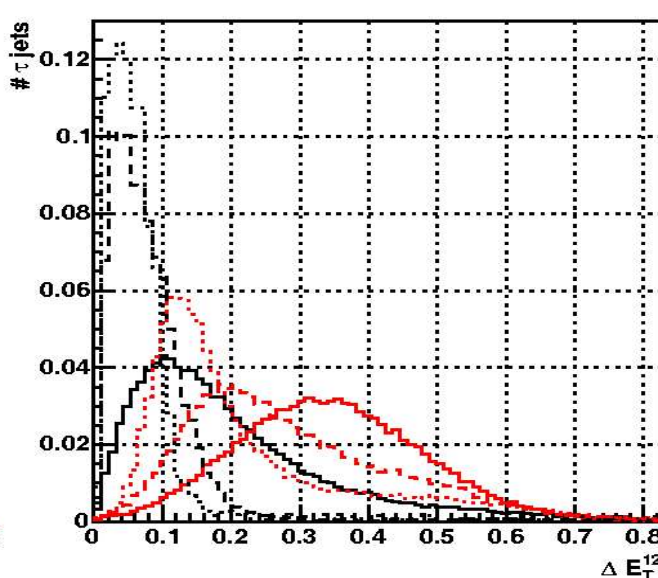
- Half a year ago a tau group has formed for ATLAS -> there is very much work in progress
- I will show only the “standard” way at the moment because we are more interested on D0 side anyway but there exist other algorithms as well
- Our  $\tau$ -reconstruction package tauRec starts from clusters found by a sliding window algorithm
- We use the following quantities to discriminate taus against jets
  - $R_{em}$  = Radius of the cluster in the em-calorimeter  $\Delta R=0.4$
  - $\Delta E_T^{12}$  = Fraction of the transverse Energy between  $\Delta R=0.2$  and  $\Delta R=0.1$  around the center of the cluster
  - $N_{tr}$  = Number of Tracks within 0.3,  $p_T > 2\text{GeV}$
  - $N_{em} / N_{strip}$  = Number of Hits in EM calo/ $\eta$ -Strip,  $E_T > 200\text{MeV}$
  - $E_{T,width,strip}$  = Width in the  $\eta$ -Strip
  - $E_{T,em}/E_T$ , Charge,  $E_{T,had}/\Delta p_{T(tracks)}$
  - Lifetime Signed Impact Parameter
  - for 3 prong decays: secondary vertex

# TauID at ATLAS

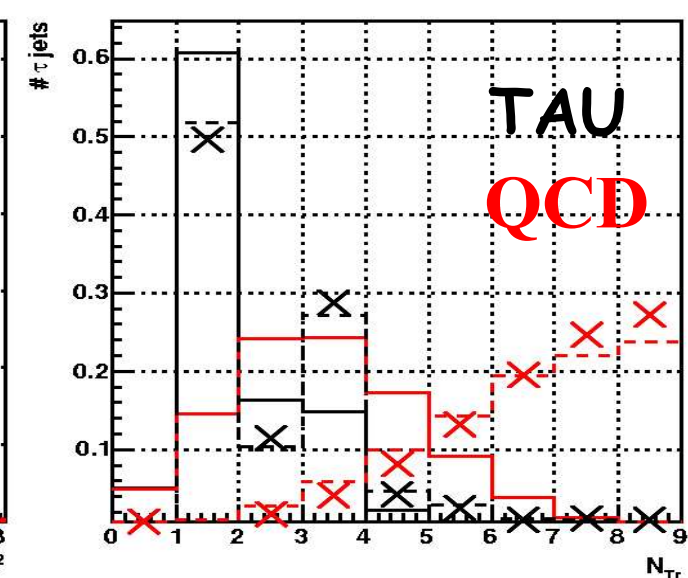
$R_{em}$  for  $\tau$  jets and qcd jets



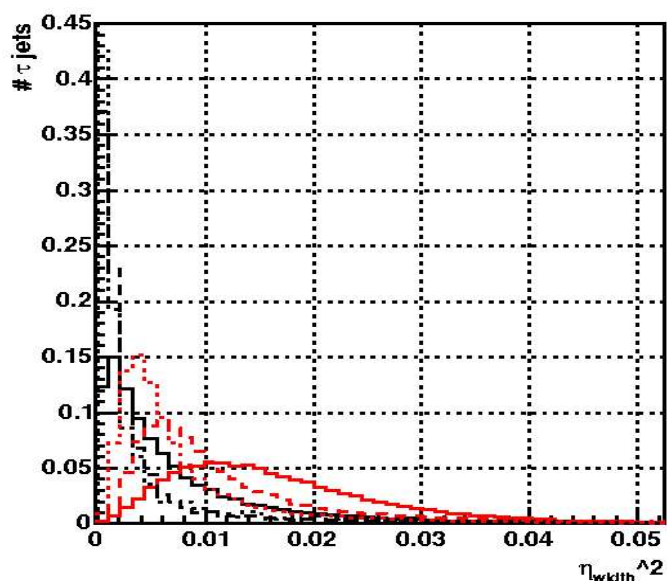
$\Delta E_T^{12}$  for  $\tau$  jets and qcd jets



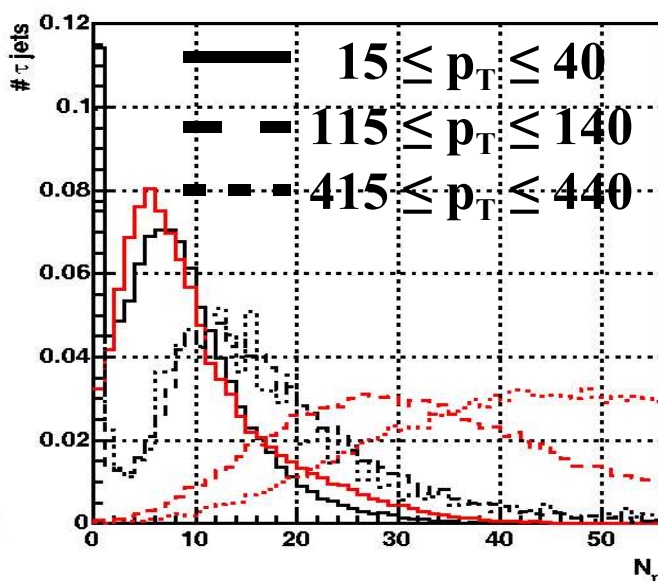
$N_{Tr}$  for  $\tau$  jets and qcd jets



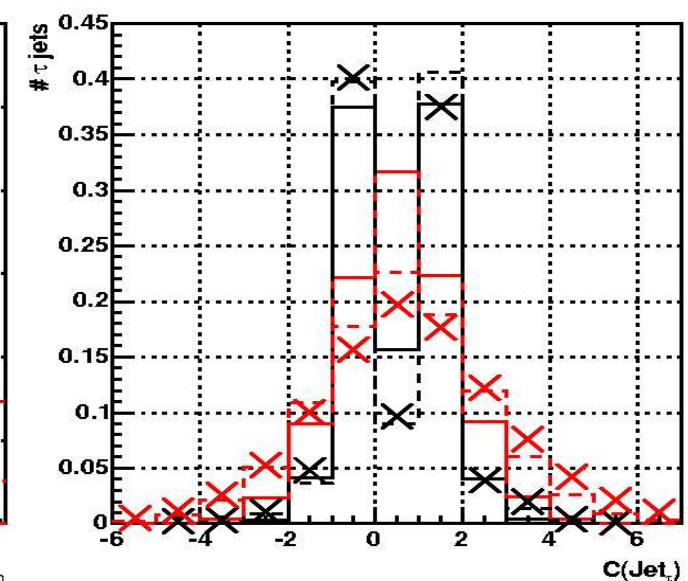
$\eta_{width}^2$  for  $\tau$  jets and qcd jets



$N_\eta$  for  $\tau$  jets and qcd jets



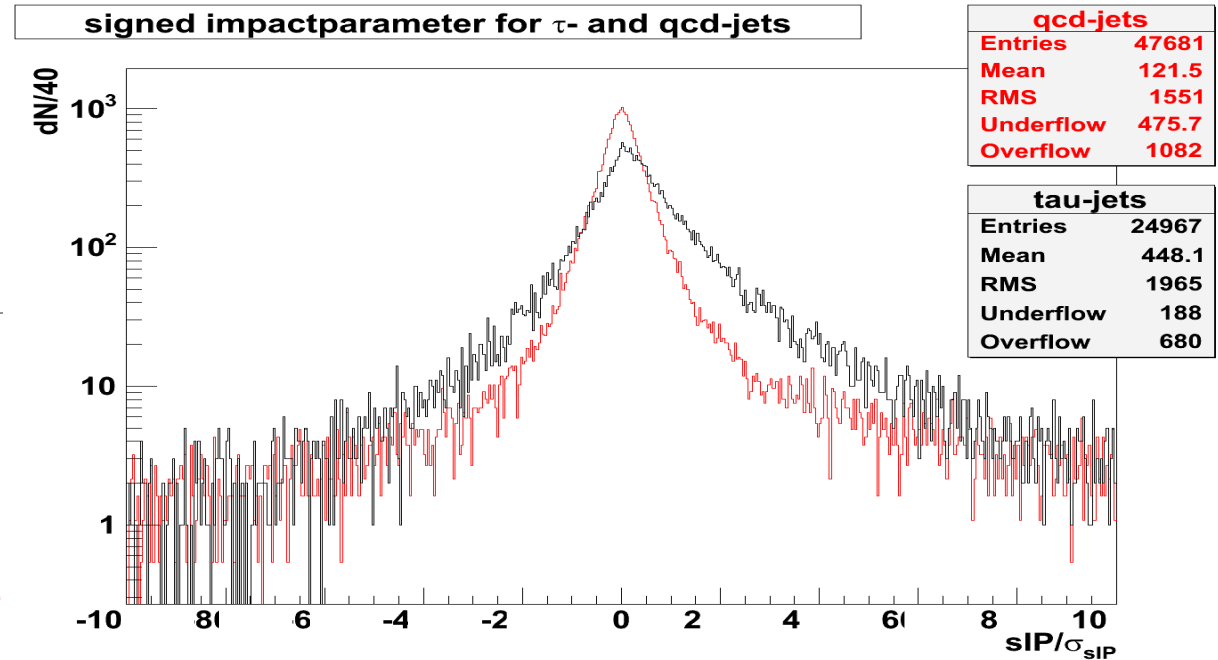
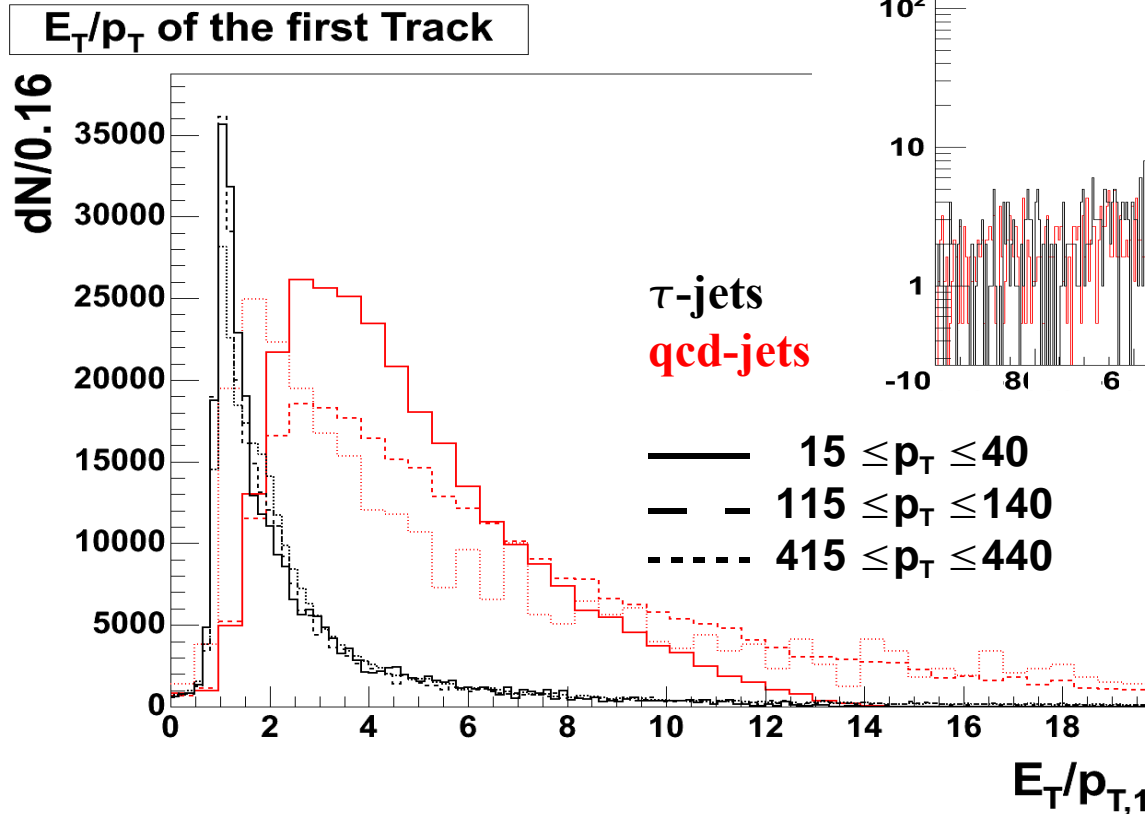
$C(Jet_i)$  for  $\tau$  jets and qcd jets





# TauID at ATLAS

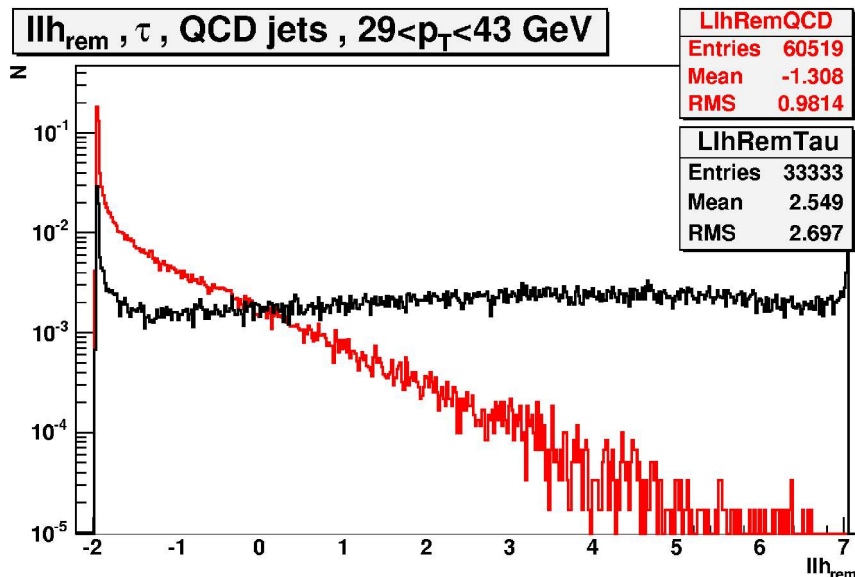
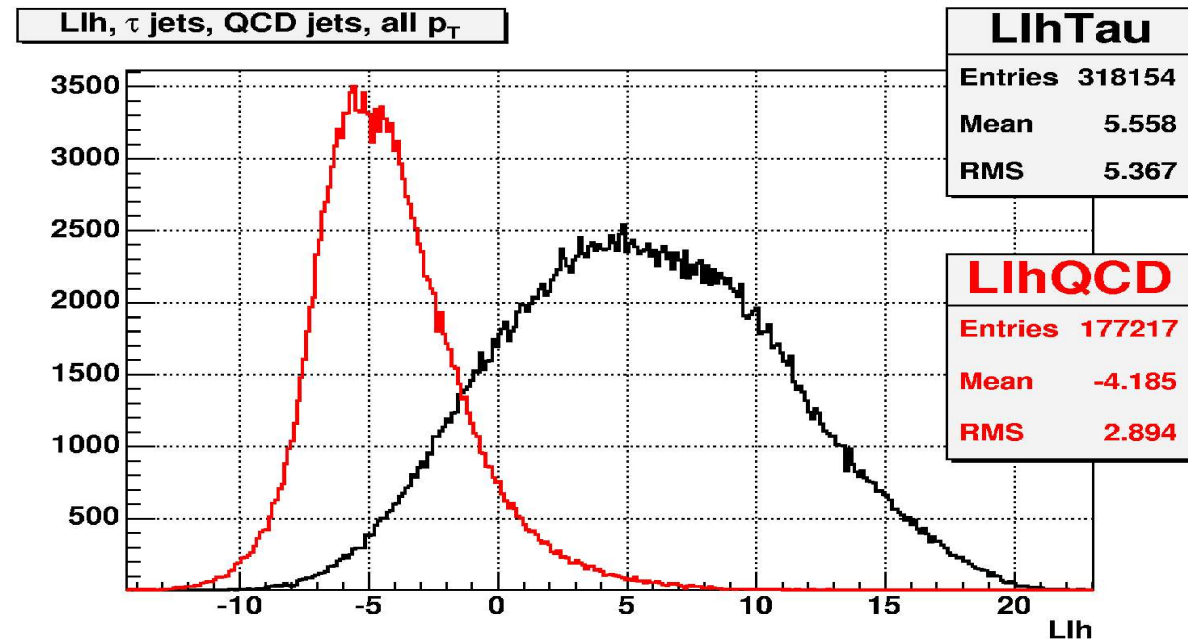
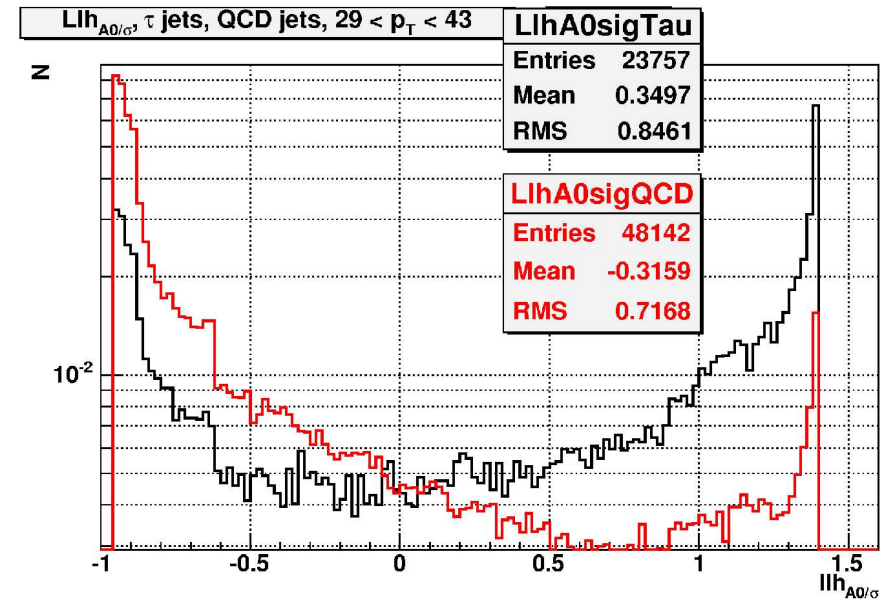
- Impact parameter significance  $A_0/\sigma_{ip}$  : only 2d information, no reconstructed primary vertex (soon to come)
- Sign is defined as  $\text{sign}(\sin(\varphi_{\text{track}} - \varphi_{\text{cluster}}))$



- $E_T/p_T$  : Ratio of total to charged transverse Energy of the Jet
- Shows  $E_T$  dependance for QCD-Jets but none for  $\tau$ -Jets

# TauID at ATLAS

- All variables are then combined into a LikelihoodRatio
- preselection cut before the Lh:  $1 \leq N_{Tr} \leq 3$
- 3 discrete variables,  $N_{Tr}, N_{strip}, Charge$ , Lh directly from histograms
- 4 continuous variables,  $R_{em}, \Delta E_T^{12}, E_{T,width,strip}, A0/\sigma$ , fitted with arbitrary functions (normally gaus\*polynom)
- 10  $p_T$  bins from 15 to 600 GeV, for Noise and NoNoise



# TauID at D0

- D0 makes use of sometimes similar variables but all are defined in slightly different way
- D0 defines a  $\tau$  – **Type** as follows
  - Type 1 : 1-prong – no em subcluster ( $\tau \rightarrow \pi + \nu$ )
  - Type 2 : 1-prong – with em subcluster ( $\tau \rightarrow \pi + \nu + x \pi^0$ )
  - Type 3 : 3-prongs (more than one  $\tau$  track)
- The em subcluster is found by the following algorithm
  - Find the leading cells in the em layer with finest granularity
  - collect all neighbour cells
  - around the leading neighbour cell, in turn collect all its neighbour cells
  - collect all em cells (from other layers) which have an overlap with any of the so far collected cells
  - if their energy > 800 MeV they are called the em subcluster
- $\tau$ -Tracks are defined in the following way
  - only tracks with  $\Delta R < 0.5$ ,  $p_T > 1.5$  GeV are considered, and sorted in  $p_T$
  - the first track is always a  $\tau$  track, the second/third tracks are  $\tau$  tracks if their invariant mass together < 1.1 / 1.7 GeV

# TauID at D0

- all variables act within a cone of 0.5 around the calo center
  - definition  $\text{variablename}(x)$  means variable calculated using objects within  $dR < x$  around the calo center
- Profile =  $(E_T(\text{Tower1}) + E_T(\text{Tower2}))/E_T(0.5)$ , Towers defined on  $\text{deta} \times \text{dphi} = 0.1 \times 0.1$  granularity
- Isolation =  $(E_T(0.5) + E_T(0.3)) / E_T(0.3)$
- $M(\text{Track1, em subcluster})$
- $p_{T1} / E_T = p_T$  of the leading tracks divided by the calorimeter energy
- EM12Frac =  $(E_T(\text{EM1}) + E_T(\text{EM2})) / E_T$ , where  $E_T(\text{EM1})$  means trans.energy in the first em layer
- $\text{trkiso} = p_T(\tau \text{ tracks}) / p_T(\text{all tracks})$
- $e1e2/E_T = \text{sqrt}(\text{sum}(\tau \text{ tracks } p_T) * E_T(\text{EM})) / E_T(0.3)$
- $\text{em3iso} = E_T(\text{em subcluster}) / E_T(\text{EM3})$
- ntr1030 = number of tracks within  $10^\circ - 30^\circ$  around the calo center

# Preliminary comparison

- The goal is to understand what difference we can expect from the results in MC to performance with real data
  - We think we can accomplish establishing a chain of understanding
  - ATLAS Algorithm on ATLAS MC → D0 Algorithm on ATLAS MC →
    - → D0 Algorithm on D0 MC → D0 Algorithm on D0 data
  - Therefor we need to implement all variables D0 uses at ATLAS
- Of course because of differences in the detector design the “translation” of variables is not unambiguous
- our convention:
  - D0 EM3 (finely granulated layer in the EM) → ATLAS EM2
  - D0 EM1, EM2 → ATLAS eta-strip layer
  - tower granularity in both cases 0.1x0.1
  - energy thresholds have to be adjusted

# Preliminary comparison

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- because of time constraints the samples are not as similar as they could be, so some differences (beside detector design) are to be expected
- D0: signal sample is  $Z \rightarrow \tau \tau$ , background is  $W \rightarrow \mu$  MC,  $W \rightarrow \mu$  data, QCD data (with muon trigger  $\rightarrow$  mainly  $b\bar{b}$ )
- ATLAS: signal sample is  $Z \rightarrow \tau \tau$ , background is QCD – dijet events (no trigger)
- with time and more statistics at ATLAS we will try to implement all cuts + triggers

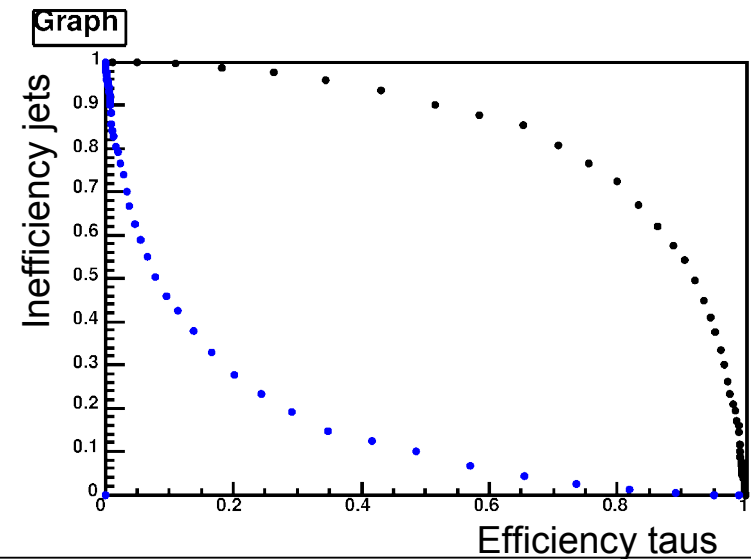
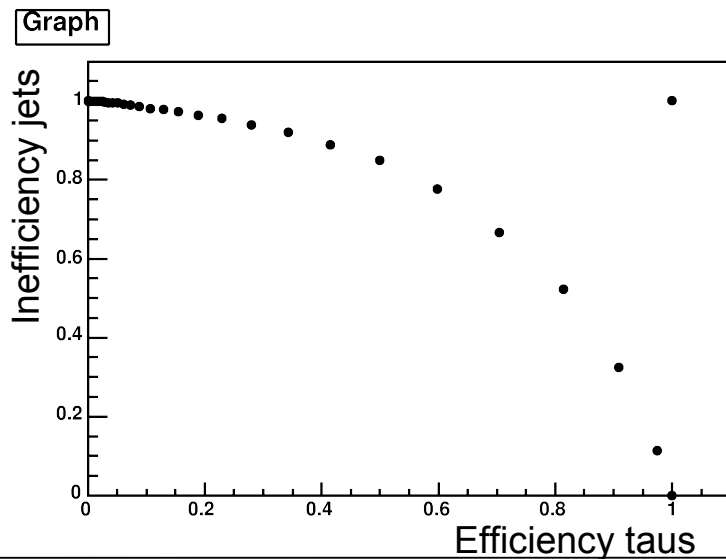
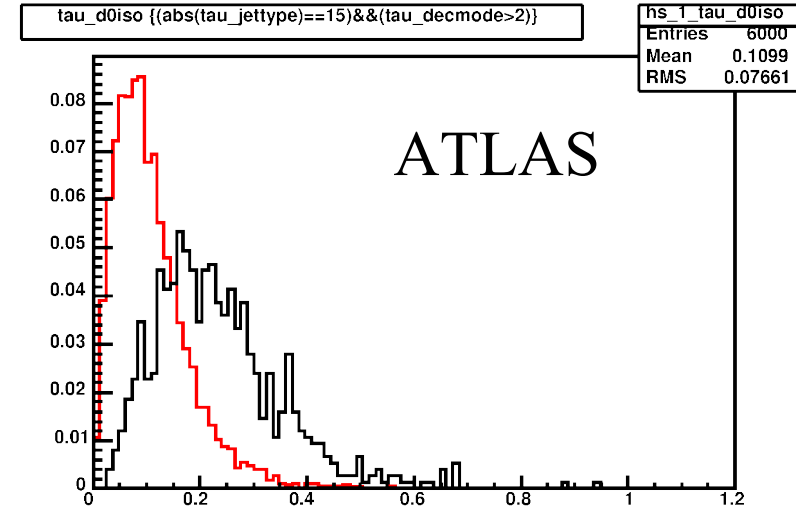
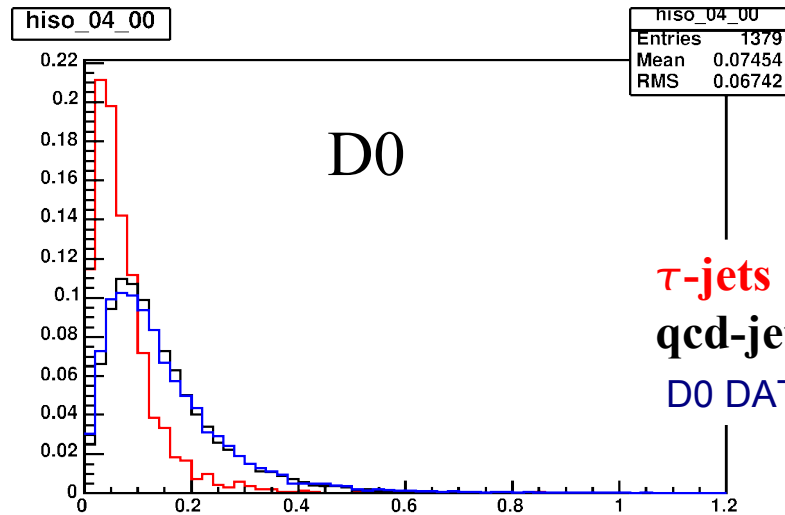
# Comparison - Isolation

- These comparisons are meant as consistency checks and not as real performance comparisons because of the stated difficulties !

- Isolation is an important criteria both at D0 and at ATLAS

- the comparison shows that both D0 and ATLAS give similar distributions for taus and jets

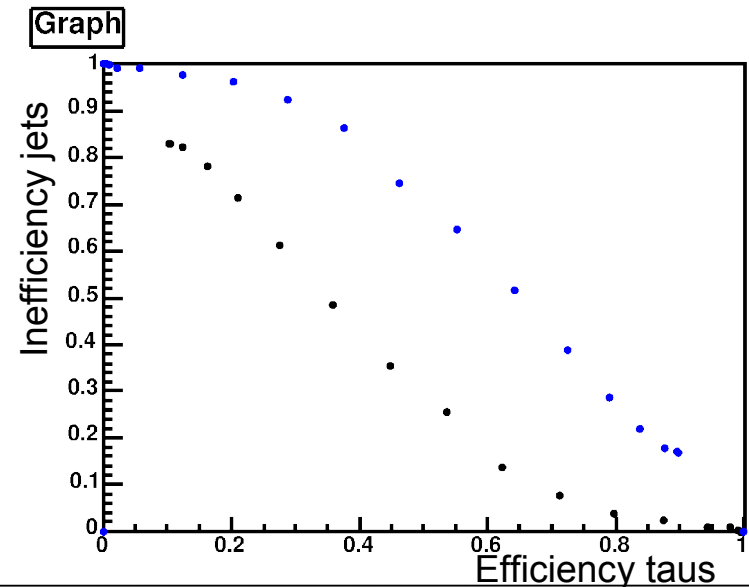
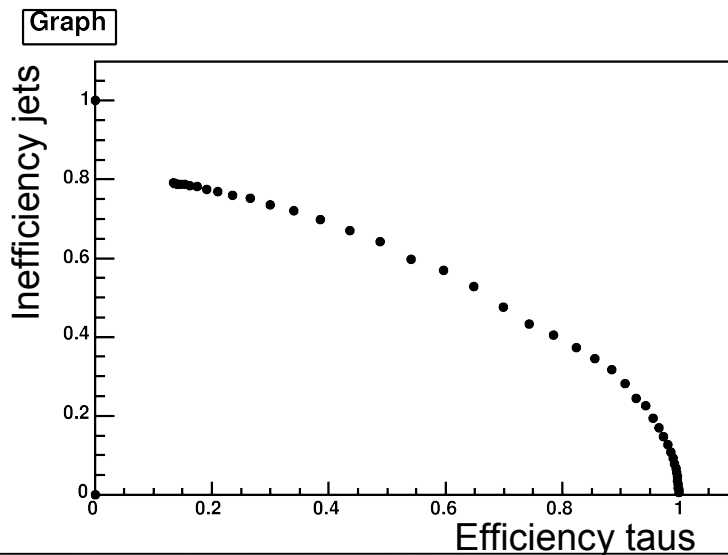
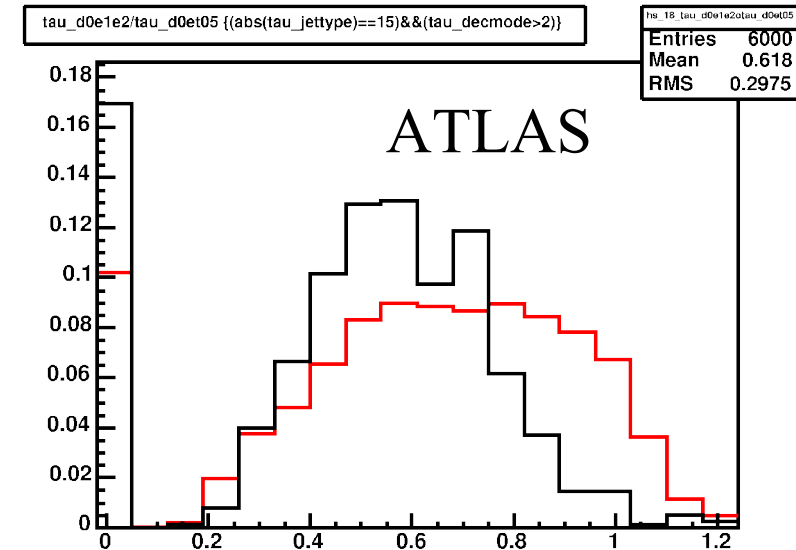
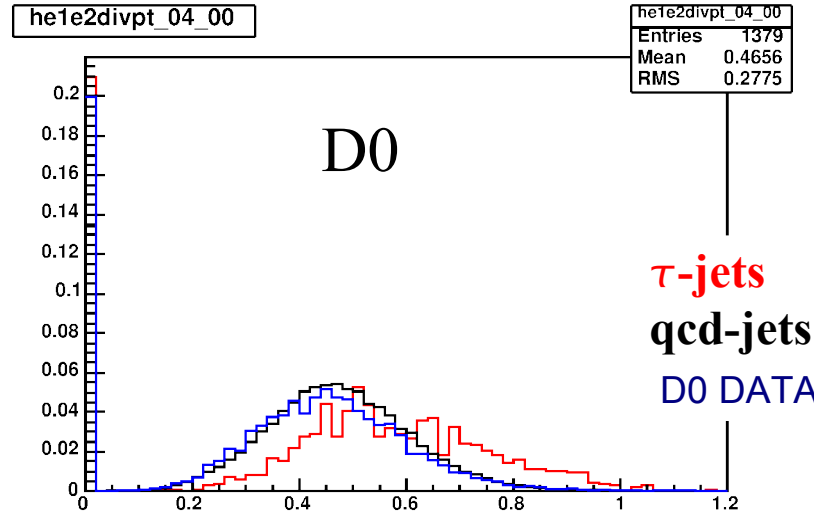
- also the performance of the variable is comparable



# Comparison $e1e2/ET(0.3)$

- $e1e2/ET$  is a measure of the difference in energy compared to the em calorimeter

- $e1e2/ET$  shows weaker discrimination for both
- distributions due not match perfectly, ATLAS show cases with very few EM energy
- the performance show different dependency on the efficiency
- but overall trend seems ok

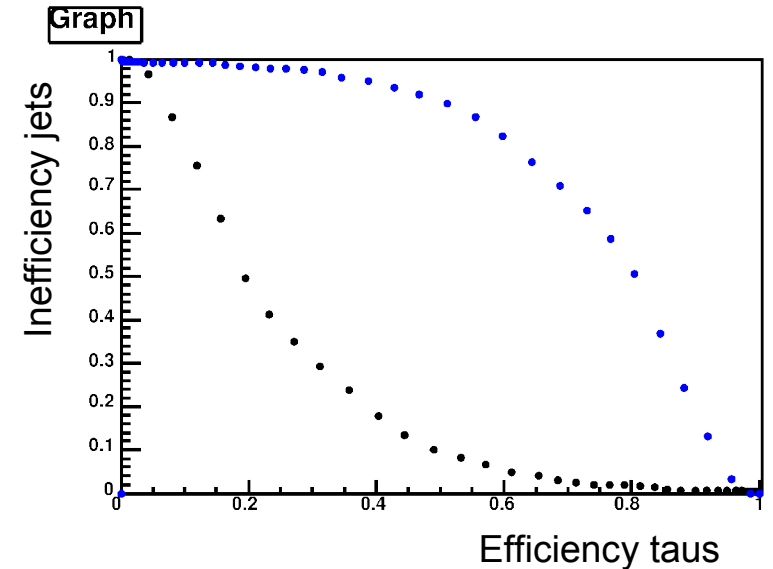
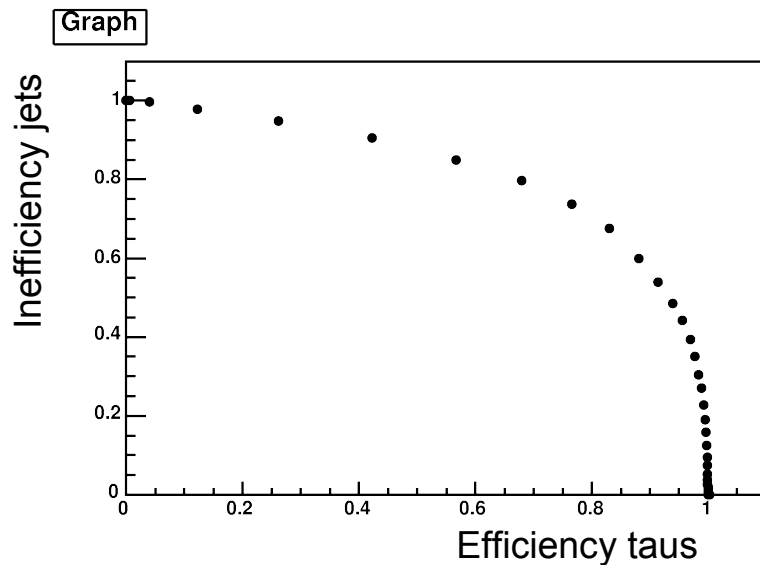
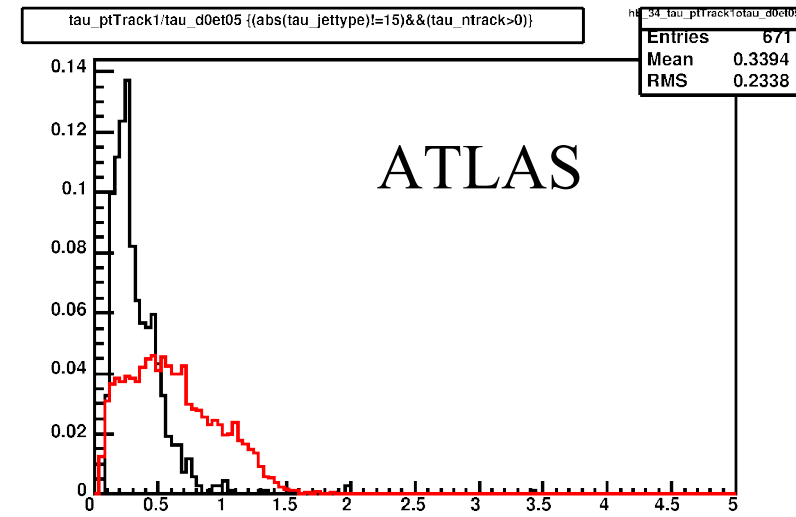
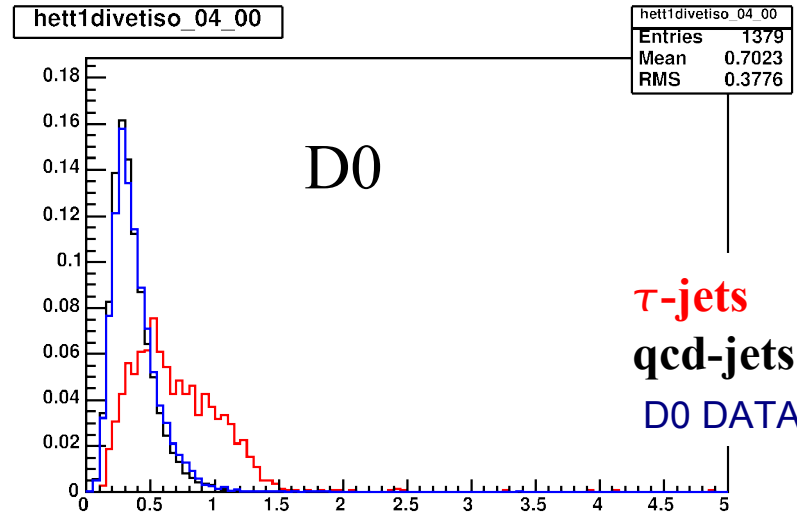




# Comparison $p_T(\text{Tr1})/ET(0.5)$

- $p_T$  of the leading tracks divided by the total calorimeter energy

- shows a good performance
- distributions are comparable
- the performance shows different behaviour for high efficiency



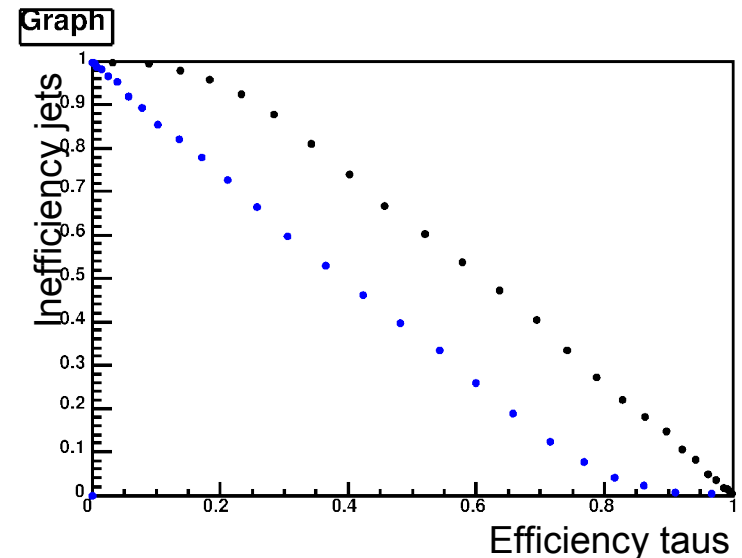
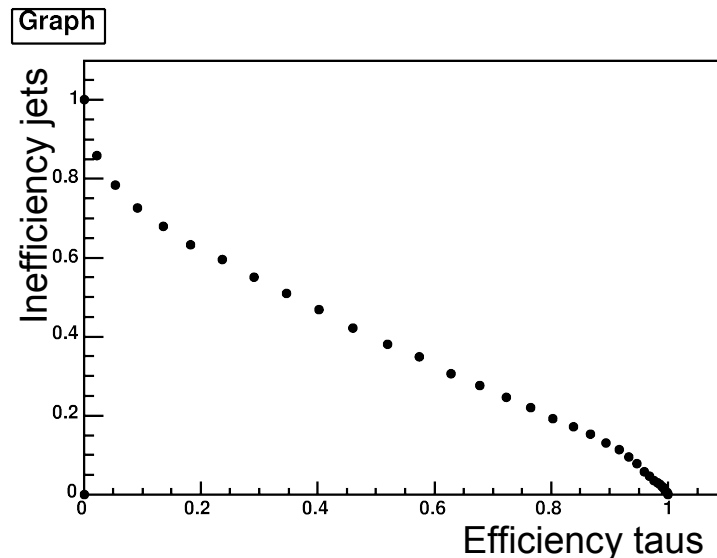
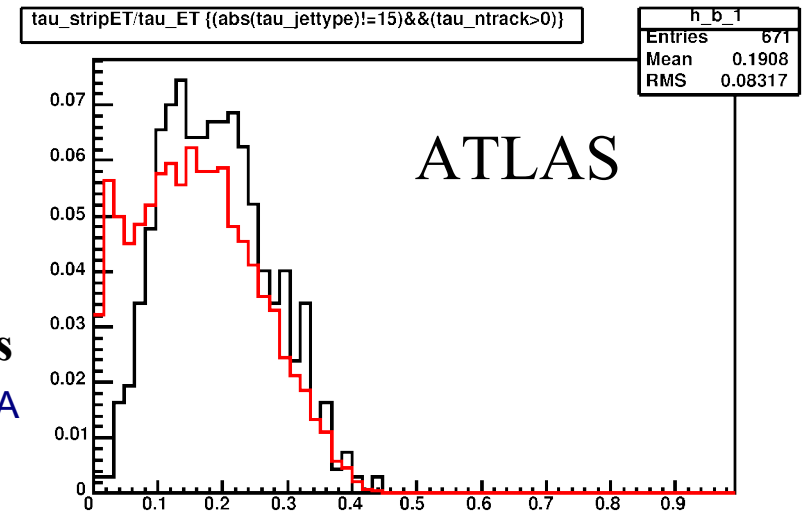
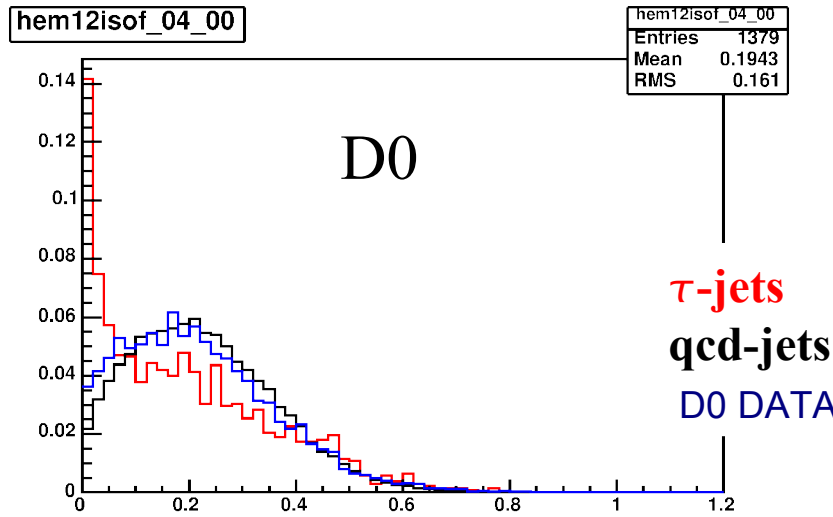
# Comparison EM12IsoF

- energy in the first two em layers divided by the total calorimeter energy

- for ATLAS I used only the eta-strip-layer, but this already has  $\sim 4.3X0 \rightarrow$  fewer cases were there is no energy

- the resulting distributions are therefore quite different

- the performance shows a different behaviour



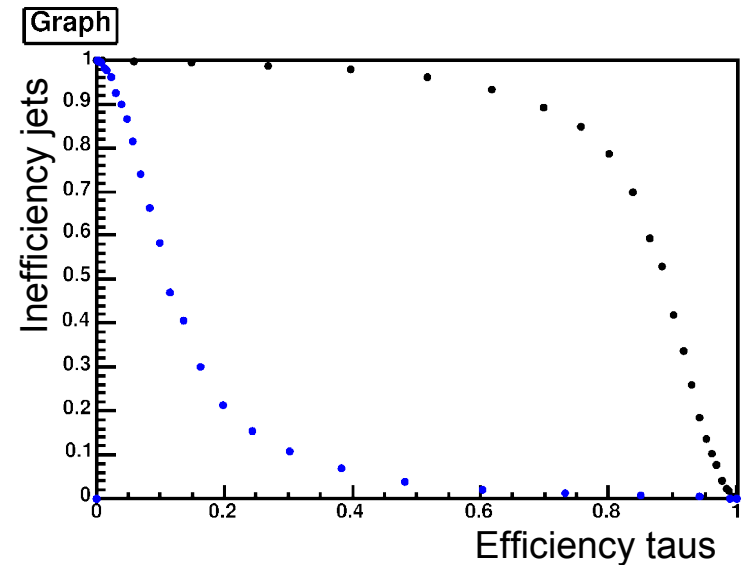
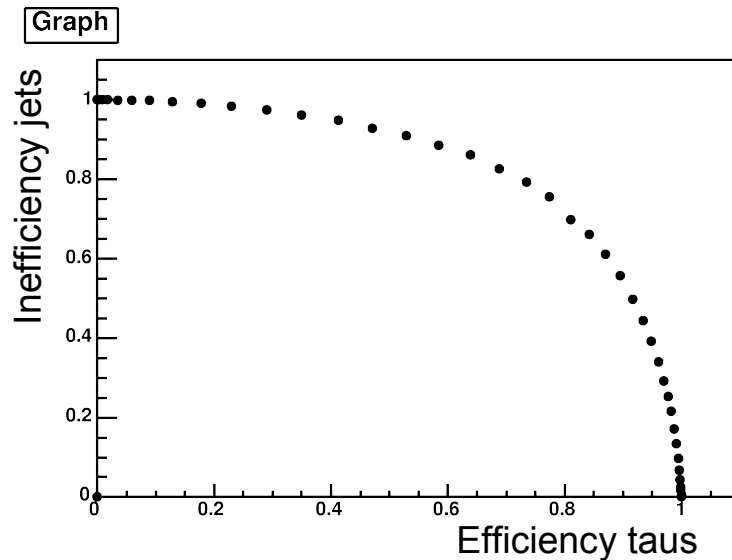
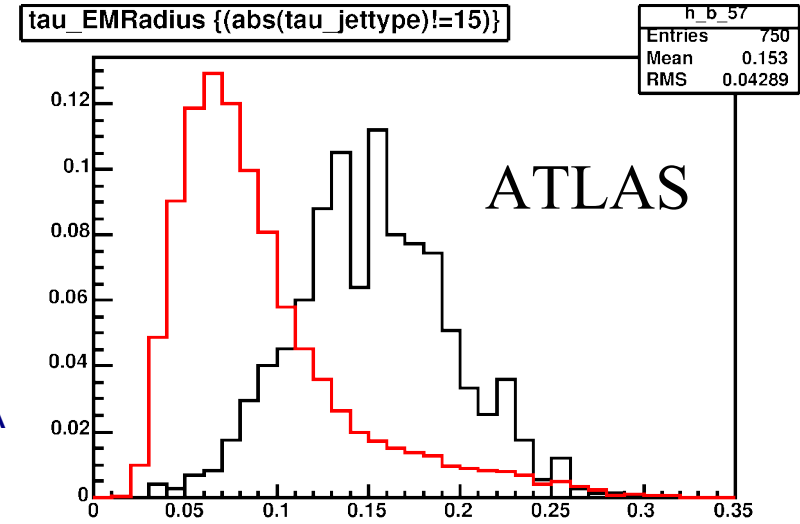
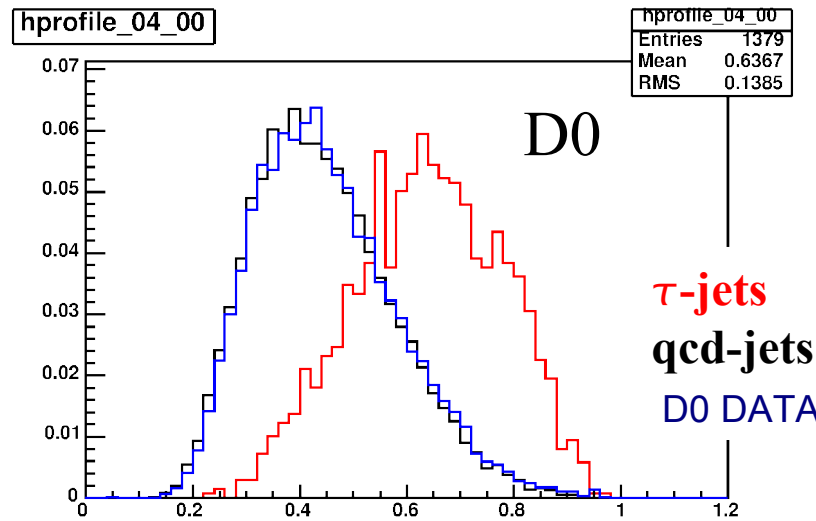
# Comparison Profile

- sorry, don't have the profile for ATLAS yet, the equivalent quantity is EMRadius

- for both ATLAS and D0 the “profile” is an important variable showing good discrimination

- only for the interest: the distributions are mirrored but show some similarity

- we will soon have the D0 style profile for ATLAS

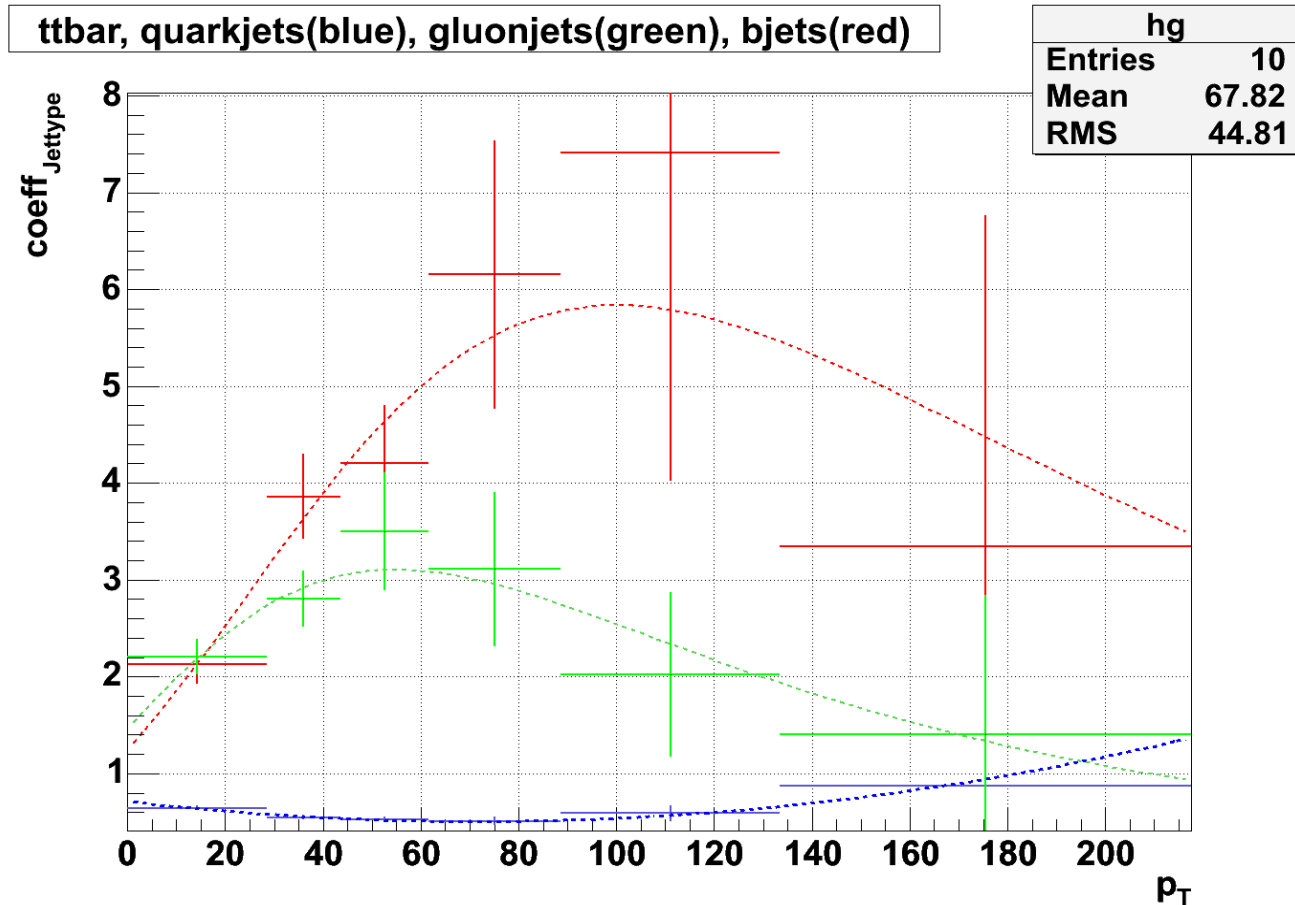


# Comparison Performance

- I don't show plots involving non-tau tracks, because D0 uses tracks down to 200MeV, and I have only  $> 1.5\text{GeV}$  at the moment
- As stated before the results are not really comparable, but just to give some idea
- we had a first look, cutting on the D0 variables for ATLAS
  - $\text{iso} < 0.2$
  - $e1e2/ETt > 0.4$
  - $pT(\text{Trk1})/ET > 0.25$
  - $EM12\text{isof} < 0.3$
  - $EM\text{Radius} < 0.15$
  - **eff = 56%, R=10.6**
- the D0 neural network gives
  - NN eff=57%, R=47                      or                      NN eff=36%, R=125
- for comparison the ATLAS LikelihoodRatio shows
  - LLH eff=57%, R=59                      or                      LLH eff=38%, R=334
- these numbers are normalized to tau-candidates, so they should be corrected for the reconstruction efficiency for the tau-candidates

# Plans

- this project has just begun
- a lot of details have still to be understood and differences made as small as possible (algorithm and samples)
- of course we need to finish the implementation of all variables
- the identification of tau types does not work at ATLAS → need to do something reasonably similar
- the samples have to be chosen carefully and enough statistics has to be available
  - at ATLAS we see that the rejection varies very much with the type of jet you are rejecting



# Plans

- → the samples we use for comparisons (also for the backgrounds) should be as similar as possible
- they should match in  $p_T$ ,  $\eta$ , and jettype
- → will probably use  $W+\text{jet}$  in the future for jets, sticking to  $Z \rightarrow \tau \tau$  for  $\tau$ s
- the preselection of D0 data should be imitated selecting the samples for ATLAS
- the  $p_T$  and  $\eta$  distribution of taus and jets should be as similar as possible
- the influence of the reconstruction has to be understood (perhaps normalizing to jets, but also the jets may show differences)
- → make a nice comparison between D0 MC, D0 data and ATLAS MC

# TeV for LHC(3)

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