

$$K^+ \rightarrow \pi^+ \nu \bar{\nu}$$

at NNLO

Ulrich Haisch



in collaboration with A.J. Buras, M. Gorbahn & U. Nierste,
hep-ph/0508165

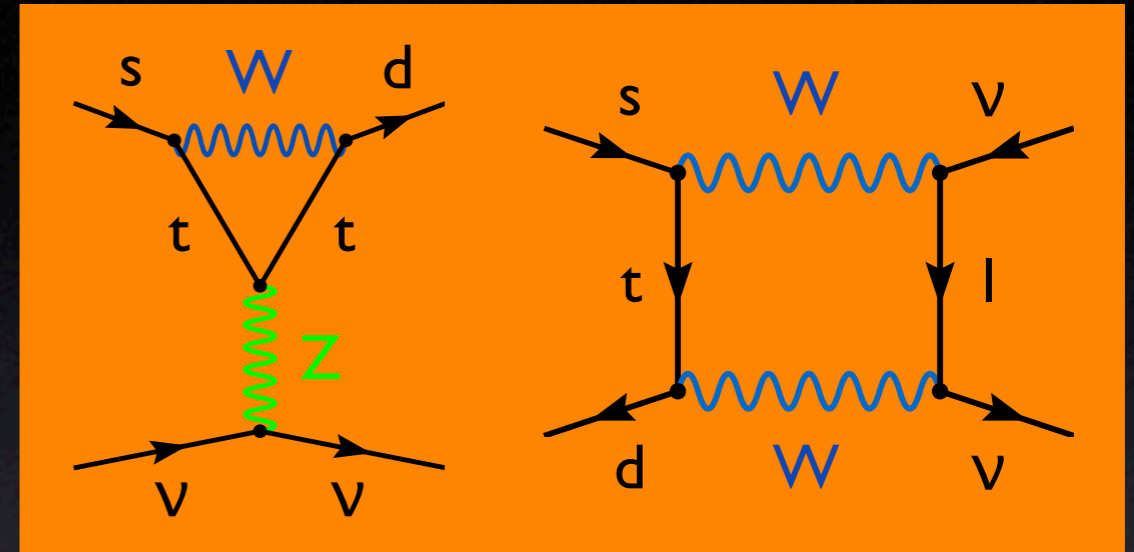
LoopFest IV, August 19, 2005, Snowmass, CO

Next 29 minutes before

- Introduction
- Theoretical status of $K \rightarrow \pi \nu \bar{\nu}$
- NNLO calculation of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$
- Determination of unitarity triangle
- Conclusions

Introduction

- FCNC processes strongly suppressed in SM by loop and CKM factors
- SD effects are significant and calculable with high precision
- LD hadronic effects are small and under good theoretical control



precise
determination of flavor
structure of SM

Introduction

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precise
determination of flavor
structure of SM

$$\sigma(\mathcal{B}(K_L \rightarrow \pi^0 \nu \bar{\nu})) = \pm(1 - 2)\%$$

$$\sigma(\mathcal{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu})) = \pm(3 - 4)\%$$

$$\sigma(\mathcal{A}_{\text{FB}}(B \rightarrow X_s l^+ l^-)) = \pm(6 - 9)\%$$

$$\sigma(\mathcal{B}(B \rightarrow X_s \gamma)) = \pm(8 - 14)\%$$

$$\sigma(\mathcal{B}(B \rightarrow X_s l^+ l^-)) = \pm(10 - 17)\%$$

$$\sigma(\mathcal{B}(B_s \rightarrow \mu^+ \mu^-)) = \pm 15\%$$

$$\sigma(\mathcal{A}_{\text{FB}}(B \rightarrow K^{(*)} l^+ l^-)) = \pm 15\%$$

$$\sigma(\mathcal{B}(B \rightarrow (K^*, \rho, \omega) \gamma)) = \pm(15 - 30)\%$$

$$\sigma(\mathcal{B}(B \rightarrow K^* l^+ l^-)) = \pm(30 - 35)\%$$

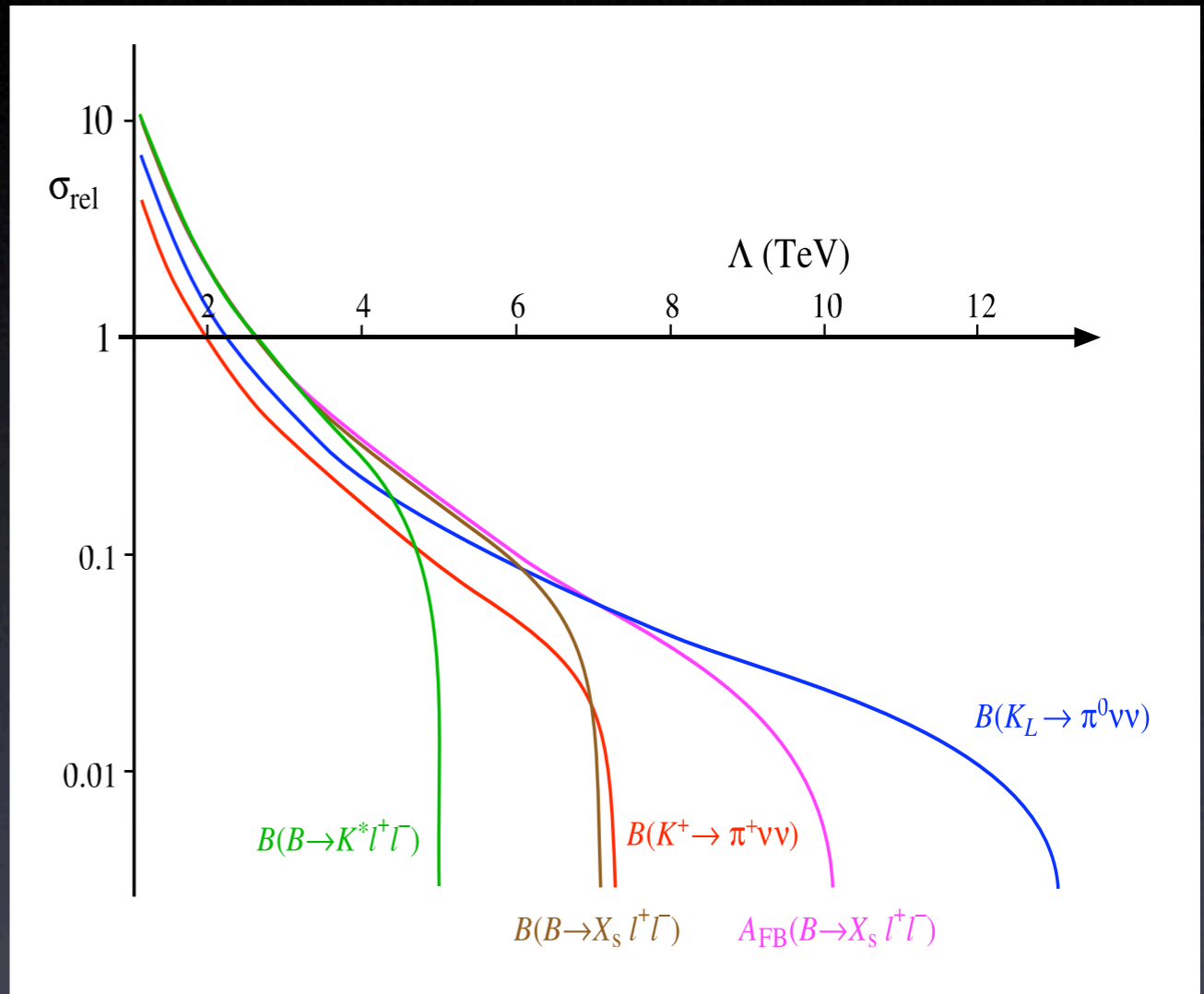
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Introduction

- FCNC processes strongly suppressed in SM by loop and CKM factors
- SD effects are significant and calculable with high precision
- LD hadronic effects are small and under good theoretical control



enhanced
sensitivity to flavor
dynamics of NP



Bryman, Buras, Isidori & Littenberg '05

General properties of $K \rightarrow \pi \nu \bar{\nu}$

$$\mathcal{L}_{\text{eff}} = \frac{4G_F}{\sqrt{2}} \frac{\alpha}{2 \sin^2 \theta_W} \sum_{i=u,c,t} C^i(\mu) Q_\nu$$

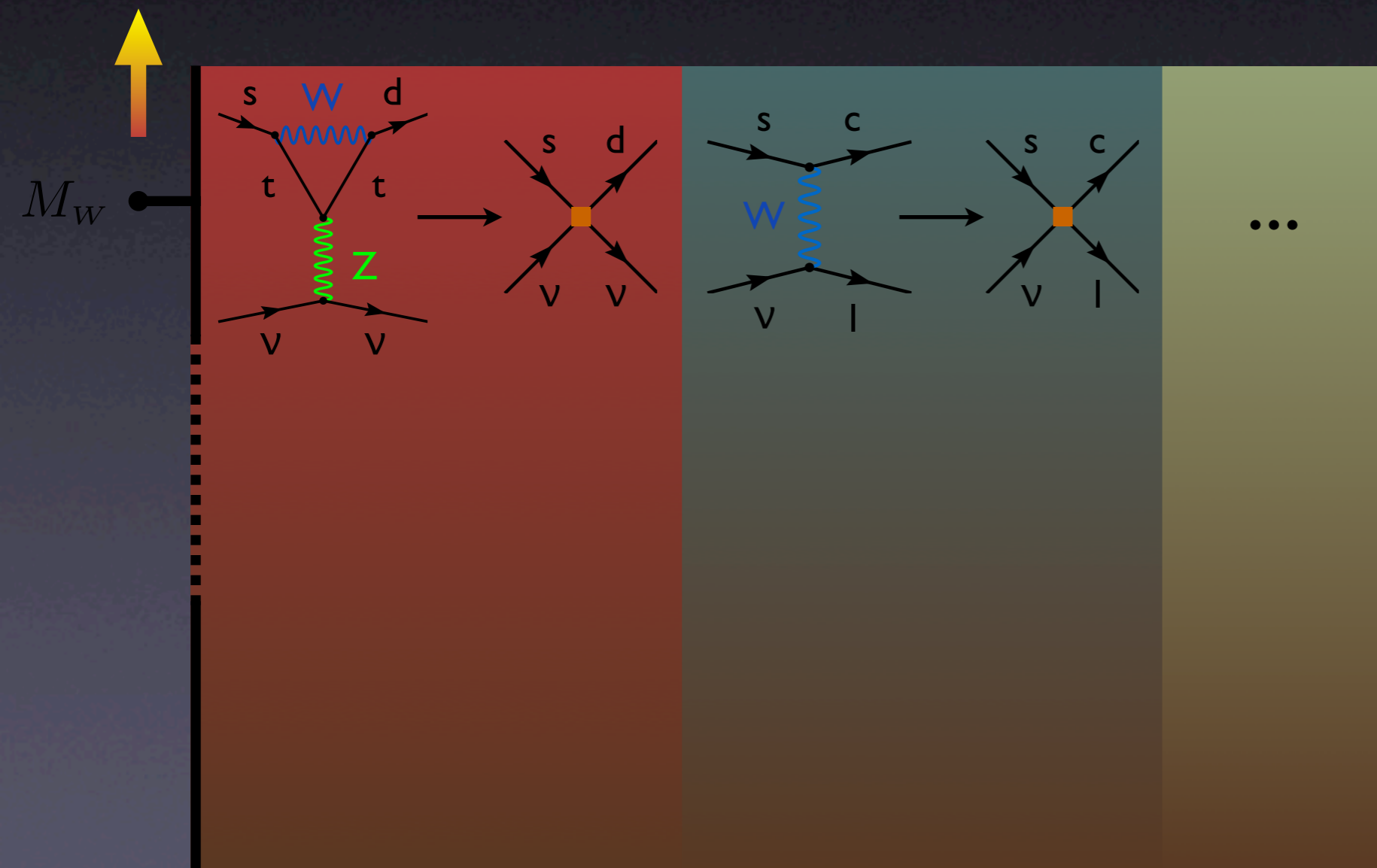
$$Q_\nu = \sum_{l=e,\mu,\tau} (\bar{s}_L \gamma_\mu d_L) (\bar{\nu}_{lL} \gamma^\mu \nu_{lL})$$

$$C^i(M_W) \propto m_i^2 V_{is}^* V_{id} \propto \begin{cases} \Lambda_{\text{QCD}}^2 \lambda & \text{u} \\ m_c^2 (\lambda + i\lambda^5) & \text{c} \\ m_t^2 (\lambda^5 + i\lambda^5) & \text{t} \end{cases}$$

u

c

t



- power-like GIM mechanism
- top quark contributions dominates
- QCD corrections are small
- large CP phase

General properties of $K \rightarrow \pi \nu \bar{\nu}$

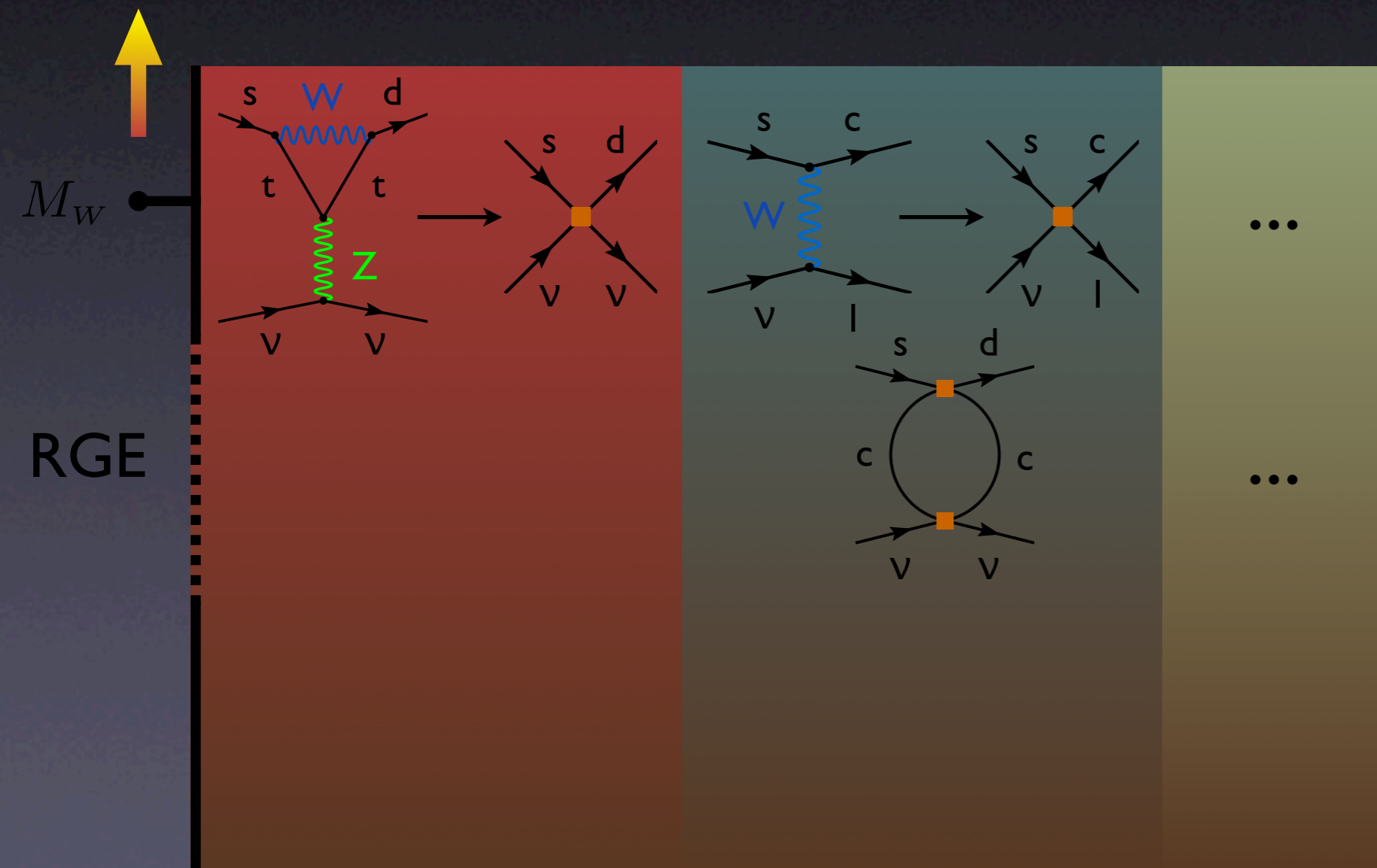
$$\mathcal{L}_{\text{eff}} = \frac{4G_F}{\sqrt{2}} \frac{\alpha}{2 \sin^2 \theta_W} \sum_{i=u,c,t} C^i(\mu) Q_\nu$$

$$Q_\nu = \sum_{l=e,\mu,\tau} (\bar{s}_L \gamma_\mu d_L) (\bar{\nu}_{lL} \gamma^\mu \nu_{lL})$$

$$\mu \frac{d}{d\mu} C_i(\mu) = \sum_j \underbrace{\gamma_{ji}(\mu)}_{\text{ADM}} C_j(\mu)$$

u

c



- top quark contribution does not evolve
- charm effects moderate for K^+ while negligible for K_L

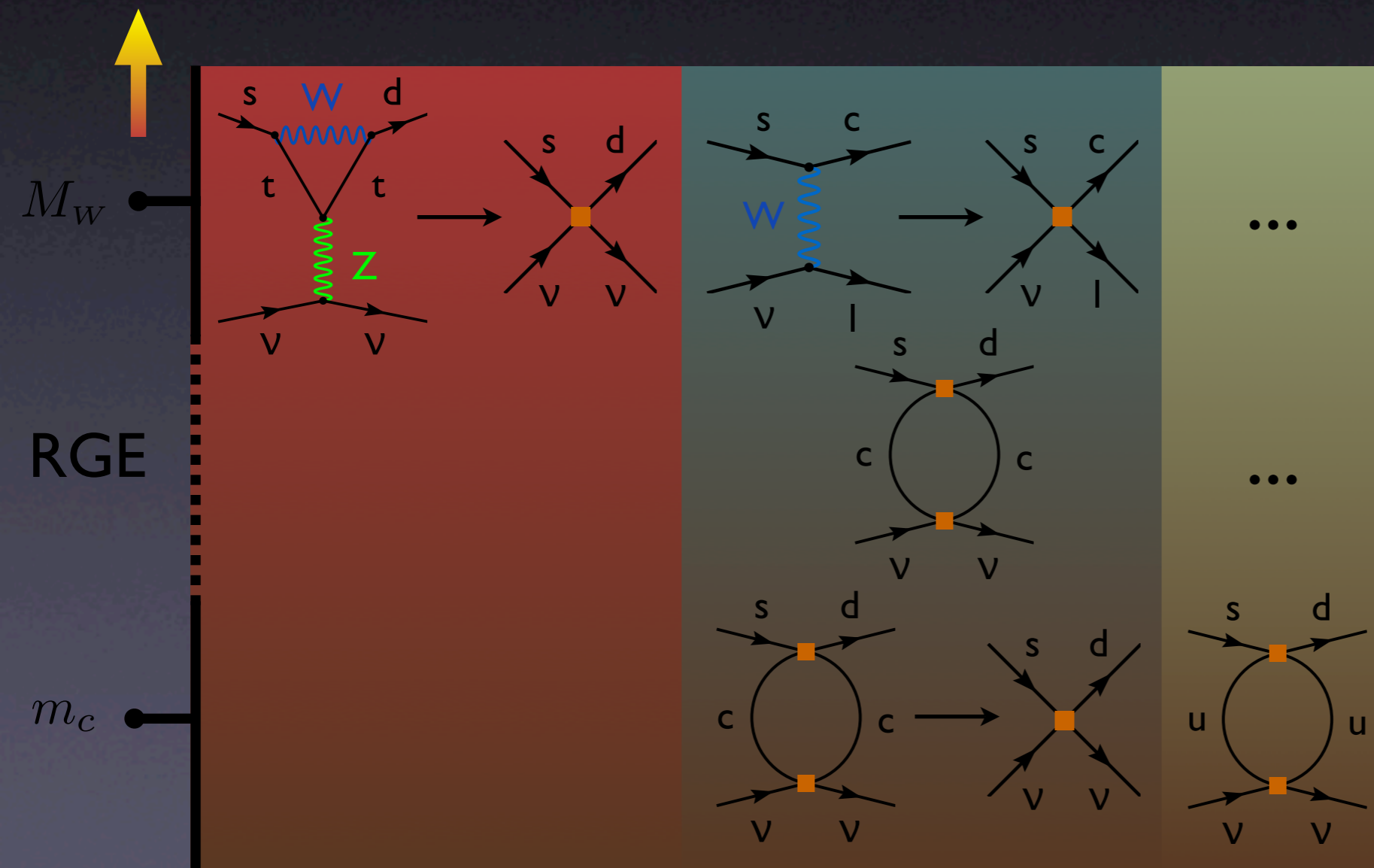
General properties of $K \rightarrow \pi \nu \bar{\nu}$

$$\mathcal{L}_{\text{eff}} = \frac{4G_F}{\sqrt{2}} \frac{\alpha}{2 \sin^2 \theta_W} \sum_{i=u,c,t} C^i(\mu) Q_\nu$$

u

$$Q_\nu = \sum_{l=e,\mu,\tau} (\bar{s}_L \gamma_\mu d_L) (\bar{\nu}_{lL} \gamma^\mu \nu_{lL})$$

$$\langle \pi | \bar{s}_L \gamma_\mu d_L | K \rangle \propto \langle \pi^0 | \bar{s}_L \gamma_\mu u_L | K^+ \rangle$$



- hadronic matrix elements precisely known from K_{e3}^+
- neutrino pair in CP eigenstate
- K_L decay purely CP

NLO SM prediction of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$

$$\mathcal{B}(K^+) = \kappa_+ \left[\left(\frac{\text{Im}\lambda_t}{\lambda^5} X \right)^2 + \left(\frac{\text{Re}\lambda_t}{\lambda^5} X + \frac{\text{Re}\lambda_c}{\lambda} (P_c + \delta P_c) \right)^2 \right]$$

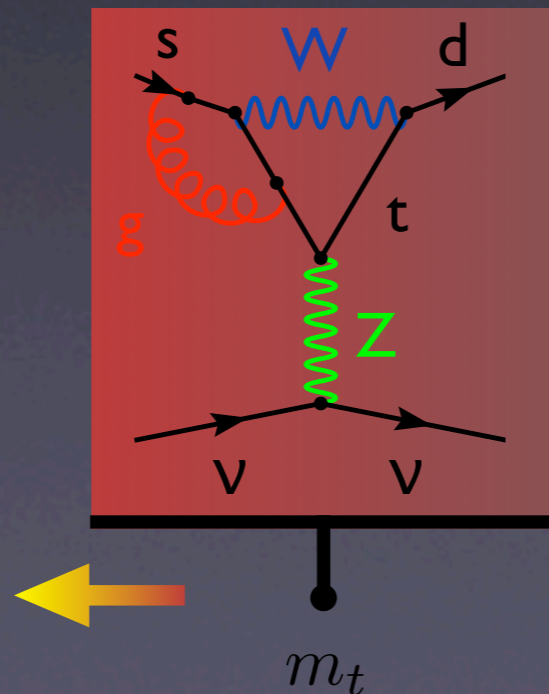
$$\kappa_+ = r_{K^+} \frac{3\alpha^2 \mathcal{B}(K_{e3}^+)}{2\pi^2 \sin^4 \theta_w}$$

$$X = 1.46 \pm 0.04 \text{ (NLO)}$$

Buchalla & Buras '93, '99;
Misiak & Urban '99

$$\lambda_i = V_{is}^* V_{id}$$

$$\lambda = |V_{us}| \approx 0.22$$



t

NLO SM prediction of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$

$$\mathcal{B}(K^+) = \kappa_+ \left[\left(\frac{\text{Im}\lambda_t}{\lambda^5} X \right)^2 + \left(\frac{\text{Re}\lambda_t}{\lambda^5} X + \frac{\text{Re}\lambda_c}{\lambda} (P_c + \delta P_c) \right)^2 \right]$$

c

t

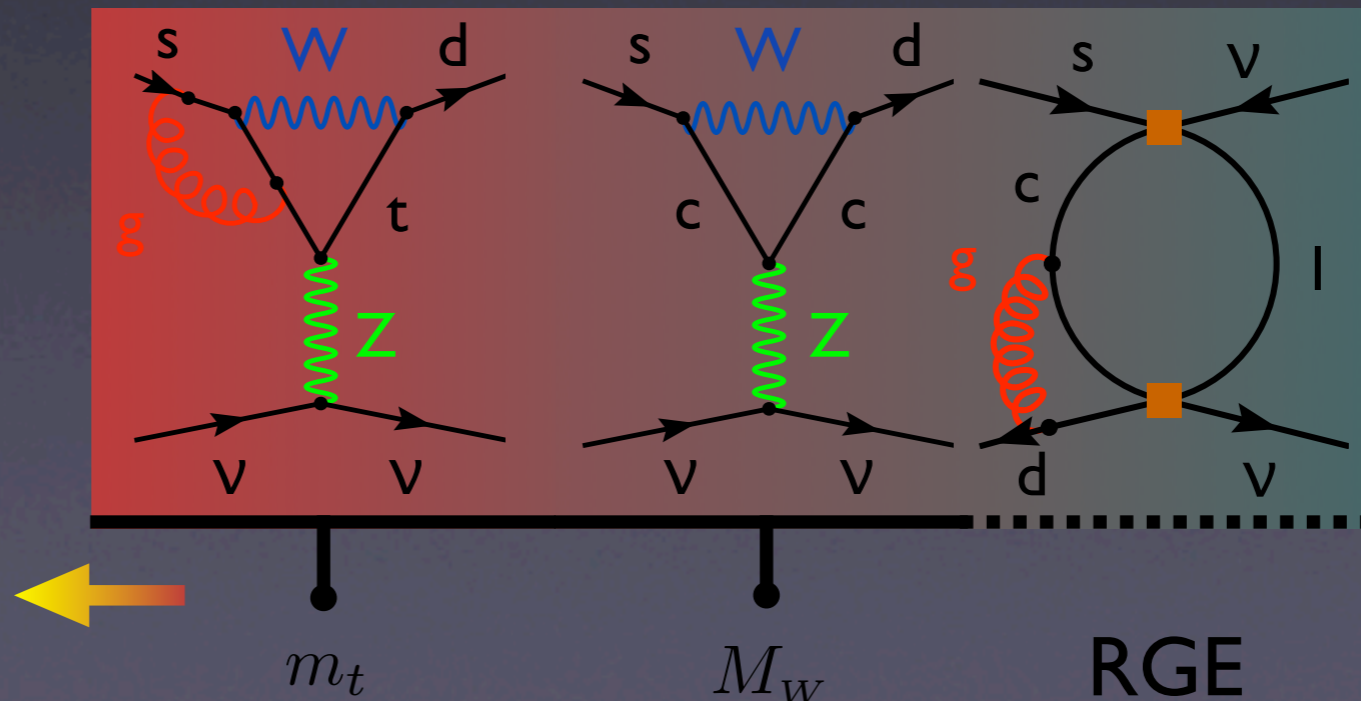
$$\kappa_+ = r_{K^+} \frac{3\alpha^2 \mathcal{B}(K_{e3}^+)}{2\pi^2 \sin^4 \theta_w}$$

$$P_c = 0.37 \pm 0.06 \quad (\text{NLO})$$

$$\lambda_i = V_{is}^* V_{id}$$

Buchalla & Buras '94, '99;
Buras, Gorbahn, Nierste & UH '05

$$\lambda = |V_{us}| \approx 0.22$$



NLO SM prediction of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$

u
c
t

$$\mathcal{B}(K^+) = \kappa_+ \left[\left(\frac{\text{Im}\lambda_t}{\lambda^5} X \right)^2 + \left(\frac{\text{Re}\lambda_t}{\lambda^5} X + \frac{\text{Re}\lambda_c}{\lambda} (P_c + \delta P_c) \right)^2 \right]$$

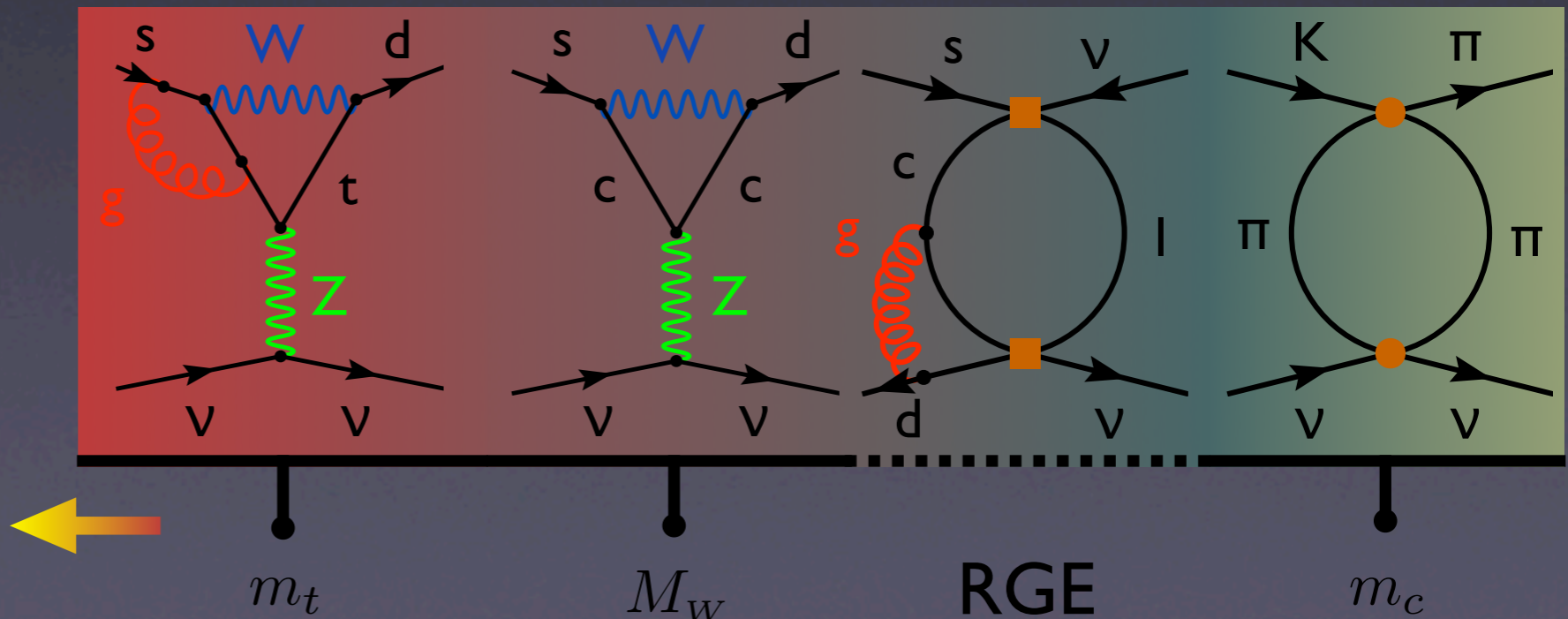
$$\kappa_+ = r_{K^+} \frac{3\alpha^2 \mathcal{B}(K_{e3}^+)}{2\pi^2 \sin^4 \theta_w}$$

$$\delta P_c = 0.04 \pm 0.02 \quad (\text{CHPT})$$

Isidori, Mescia & Smith '05

$$\lambda_i = V_{is}^* V_{id}$$

$$\lambda = |V_{us}| \approx 0.22$$



NLO SM prediction of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$

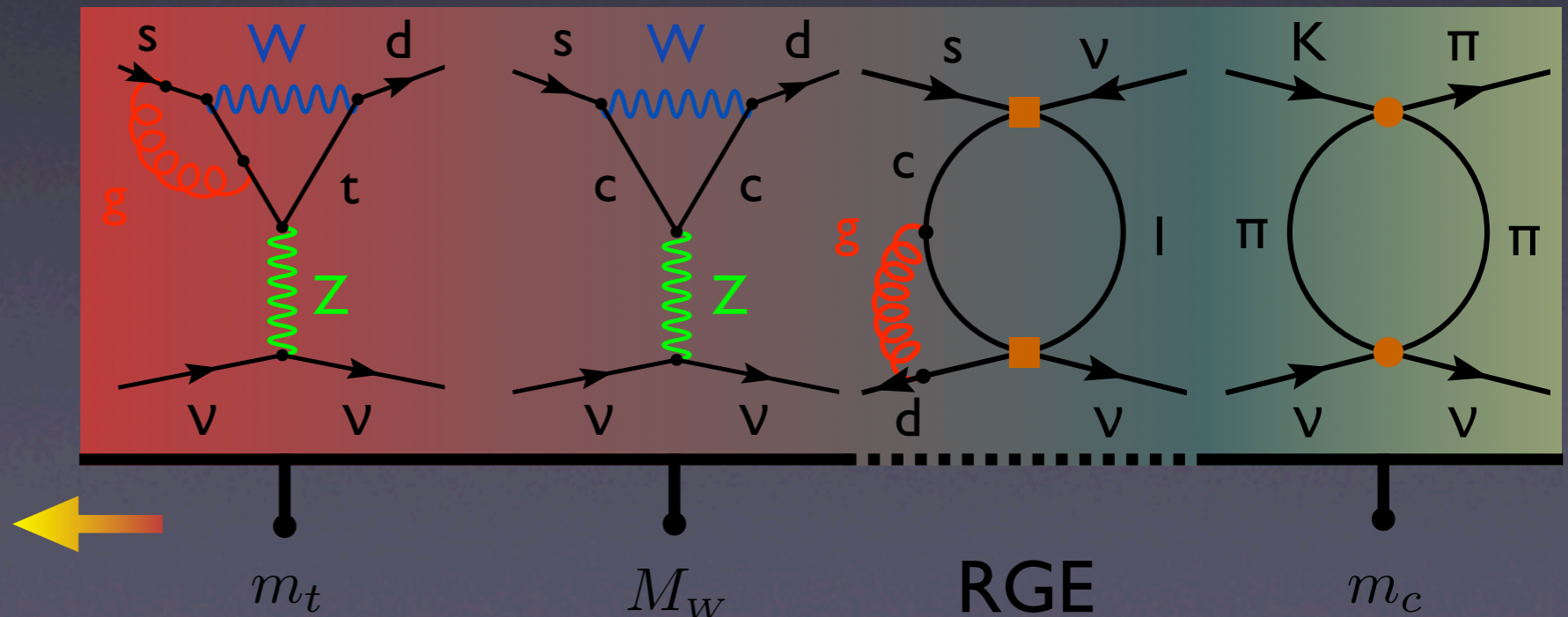


$$\mathcal{B}(K^+) = \kappa_+ \left[\left(\frac{\text{Im}\lambda_t}{\lambda^5} X \right)^2 + \left(\frac{\text{Re}\lambda_t}{\lambda^5} X + \frac{\text{Re}\lambda_c}{\lambda} (P_c + \delta P_c) \right)^2 \right]$$

$$\kappa_+ = r_{K^+} \frac{3\alpha^2 \mathcal{B}(K_{e3}^+)}{2\pi^2 \sin^4 \theta_w}$$

$$r_{K^+} = 0.90 \pm 0.03 \quad (\text{SU}(2))$$

Marciano & Parsa '96



NLO SM prediction of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$

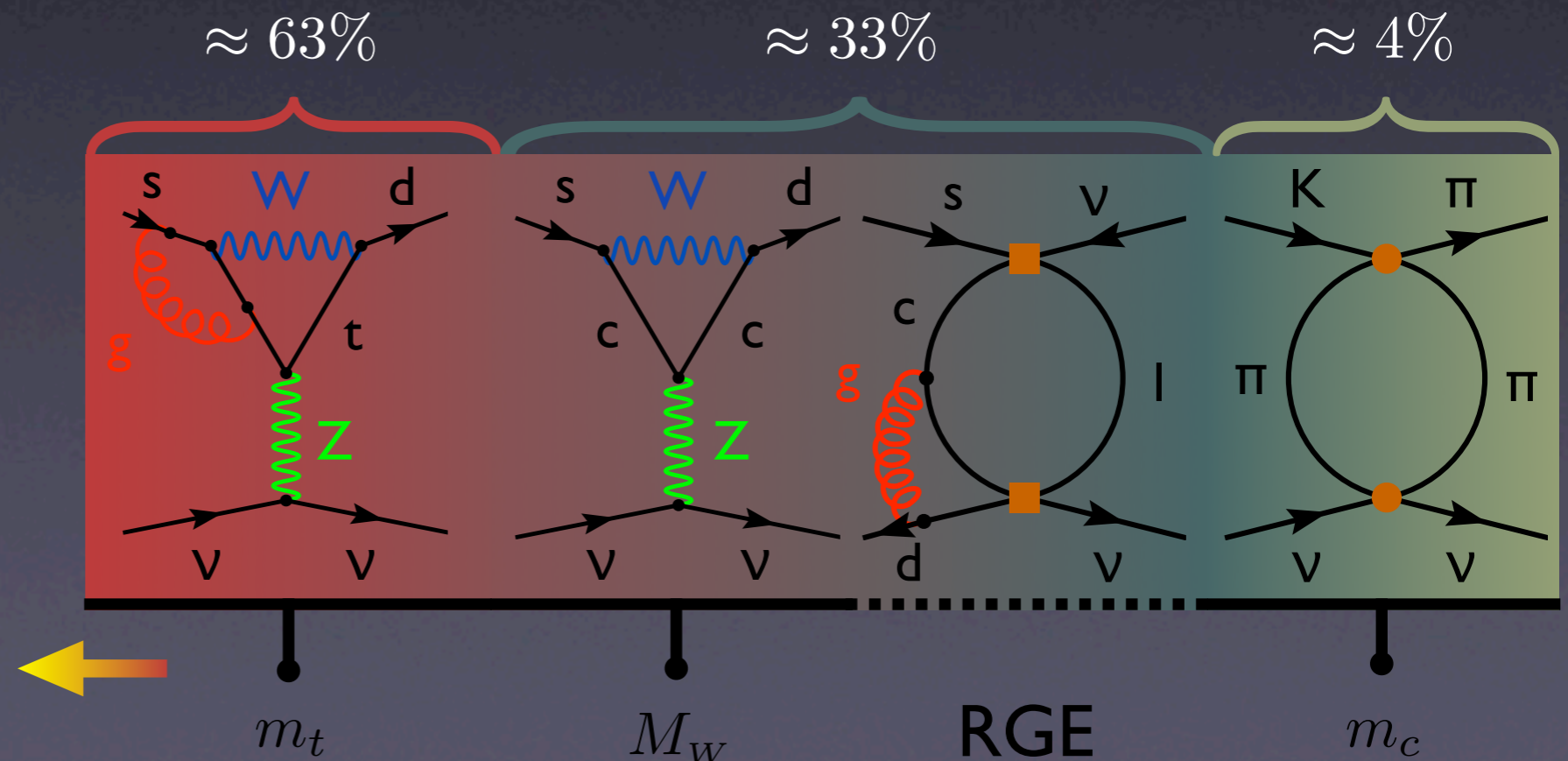
$$\mathcal{B}(K^+) = \kappa_+ \left[\left(\frac{\text{Im}\lambda_t}{\lambda^5} X \right)^2 + \left(\frac{\text{Re}\lambda_t}{\lambda^5} X + \frac{\text{Re}\lambda_c}{\lambda} (P_c + \delta P_c) \right)^2 \right] = (7.9 \pm 1.3) \times 10^{-11}$$



$$\kappa_+ = r_{K^+} \frac{3\alpha^2 \mathcal{B}(K_{e3}^+)}{2\pi^2 \sin^4 \theta_w}$$

Buras, Gorbahn, Nierste & UH '05

- dominant theoretical error of $\approx 6\%$ due to perturbative charm contribution P_c
- $\approx 50\%$ of total error of $\approx 16\%$ due to m_c and CKM elements



NLO SM prediction of $K_L \rightarrow \pi^0 \nu \bar{\nu}$

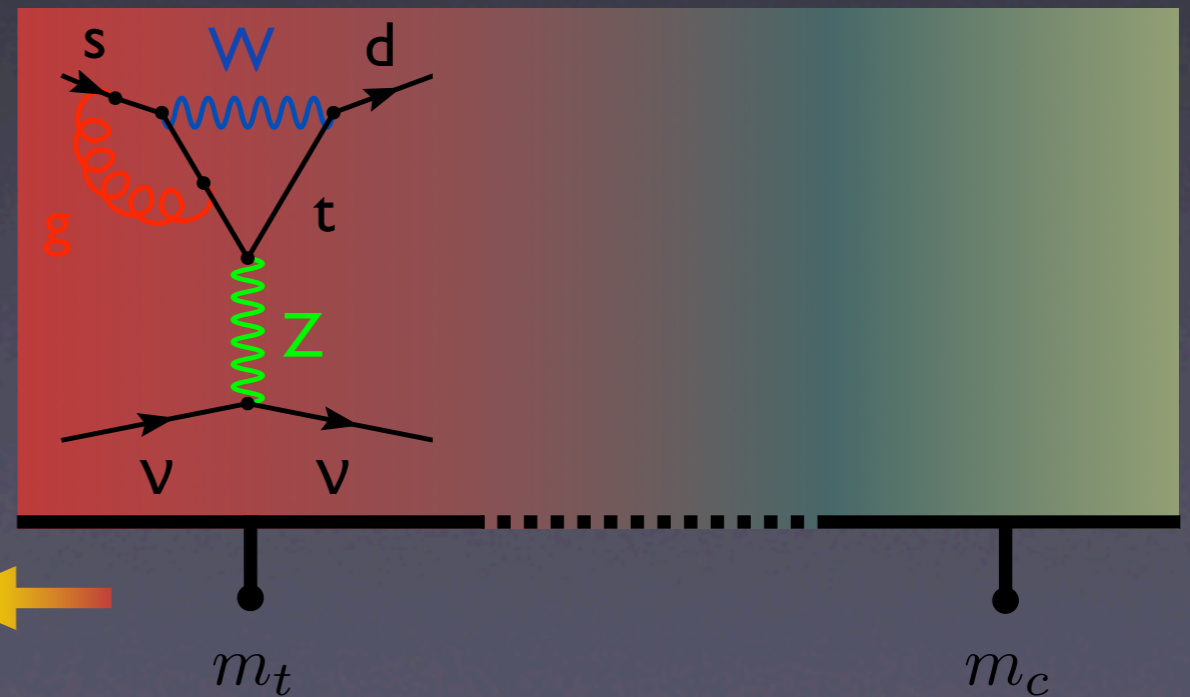
$$\mathcal{B}(K_L) = \kappa_L \left(\frac{\text{Im} \lambda_t}{\lambda^5} X \right)^2$$

t

$$\kappa_L = \kappa_+ \frac{r_{K_L} \tau(K_L)}{r_{K^+} \tau(K^+)}$$

$$r_{K_L} = 0.94 \pm 0.03 \quad (\text{SU}(2))$$

Marciano & Parsa '96



NLO SM prediction of $K_L \rightarrow \pi^0 \nu \bar{\nu}$

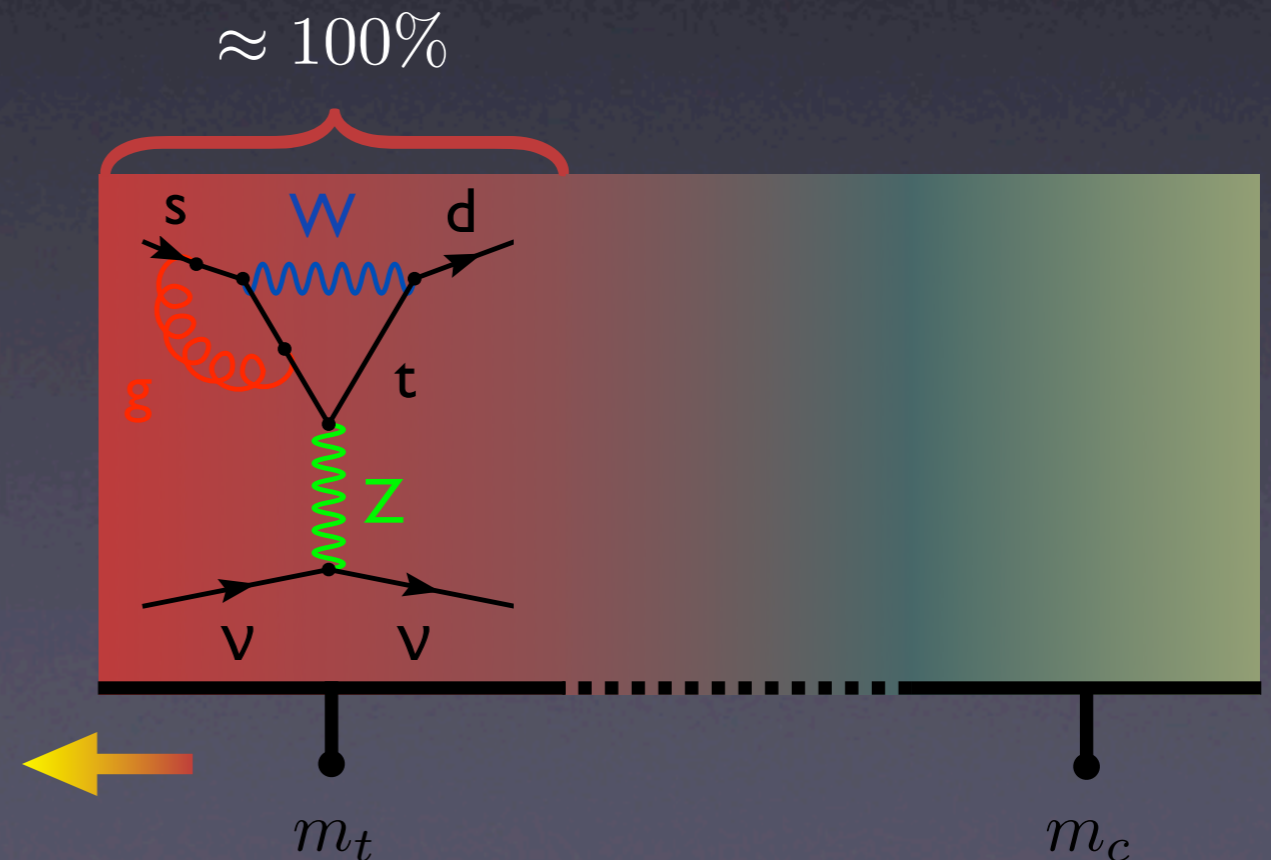
$$\mathcal{B}(K_L) = \kappa_L \left(\frac{\text{Im} \lambda_t}{\lambda^5} X \right)^2 = (2.9 \pm 0.4) \times 10^{-11}$$

t

$$\kappa_L = \kappa_+ \frac{r_{K_L} \tau(K_L)}{r_{K^+} \tau(K^+)}$$

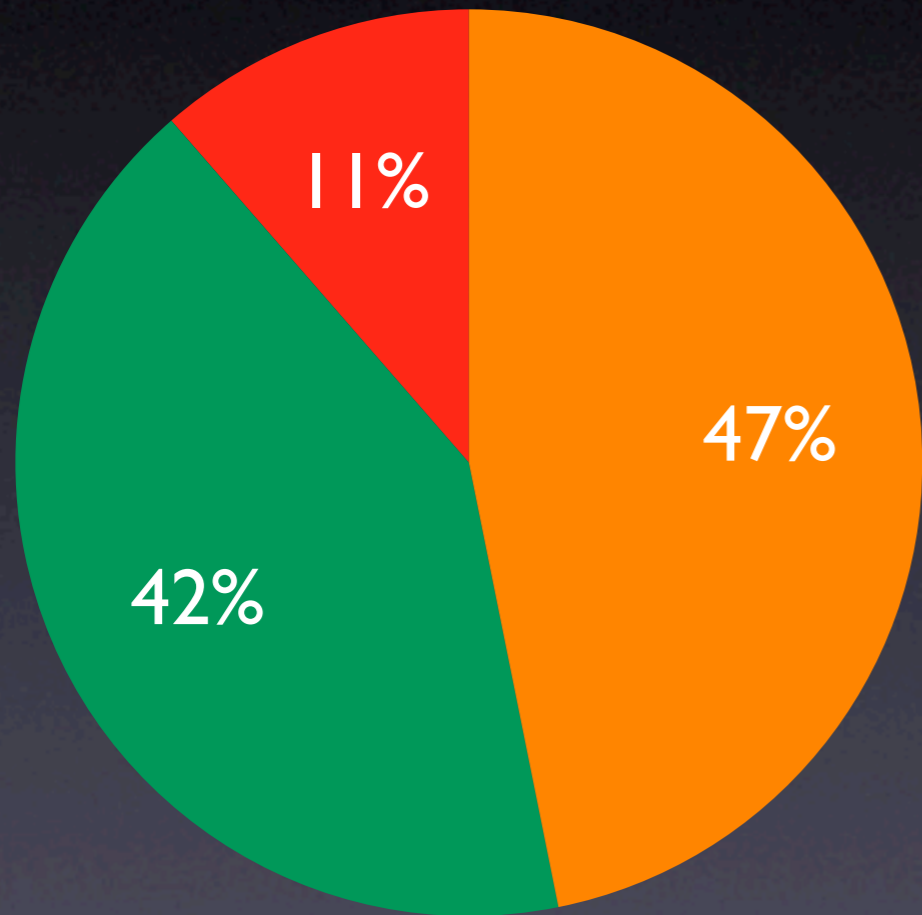
Littenberg '89;
 Buchalla & Buras '96;
 Buchalla & Isidori '98;
 Buras, Gorbahn, Nierste & UH '05

- very small theoretical error of only $\approx 1\%$
- $\approx 90\%$ of total error of $\approx 15\%$ due to CKM parameters
- within SM amount of \mathcal{CP} can in principle be determined with unmatched precision



Error budget of P_c at NLO

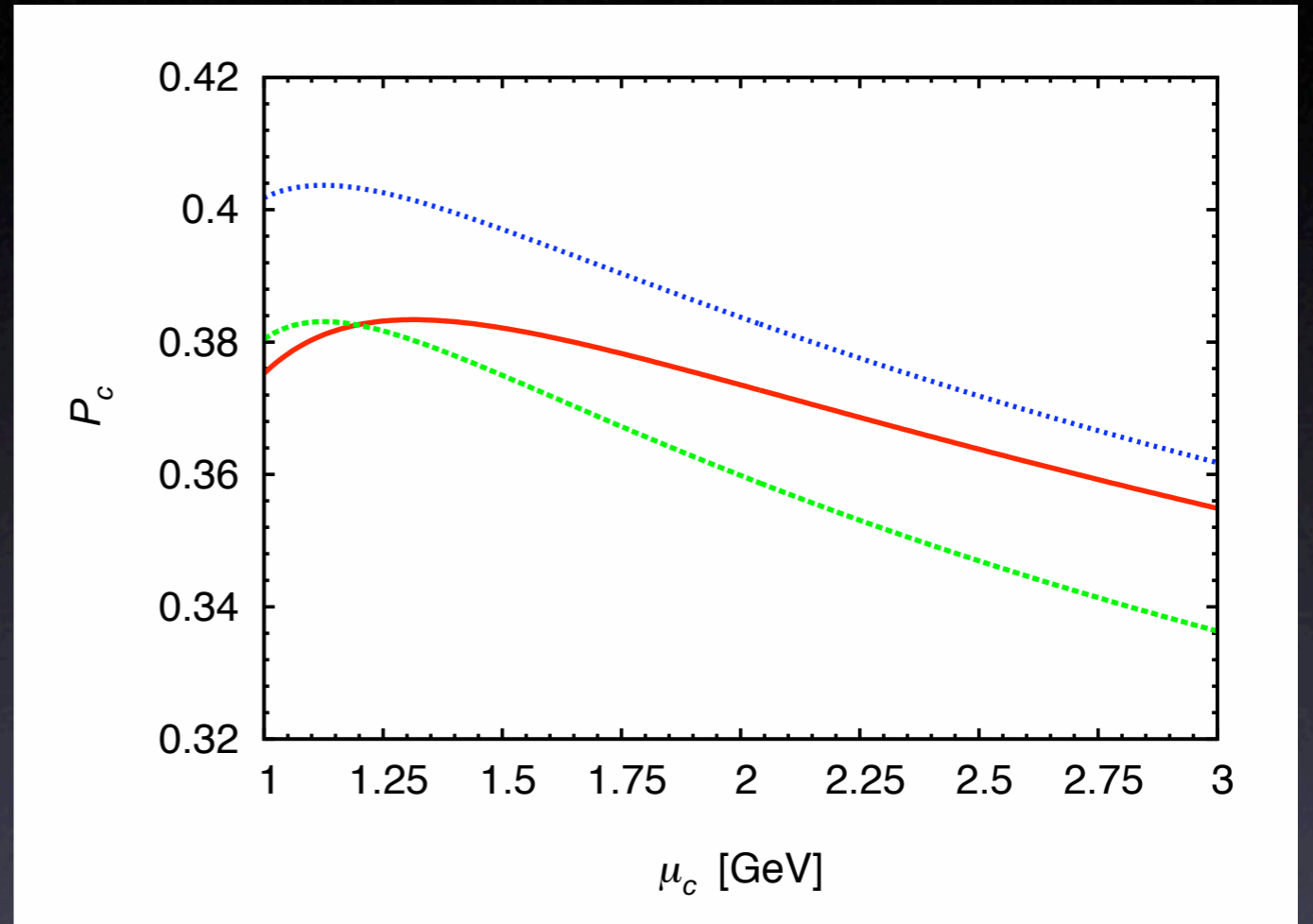
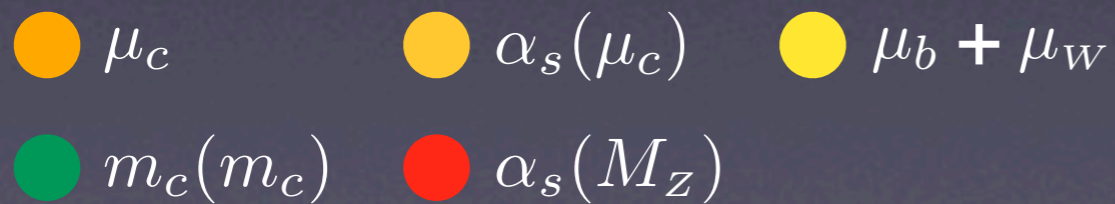
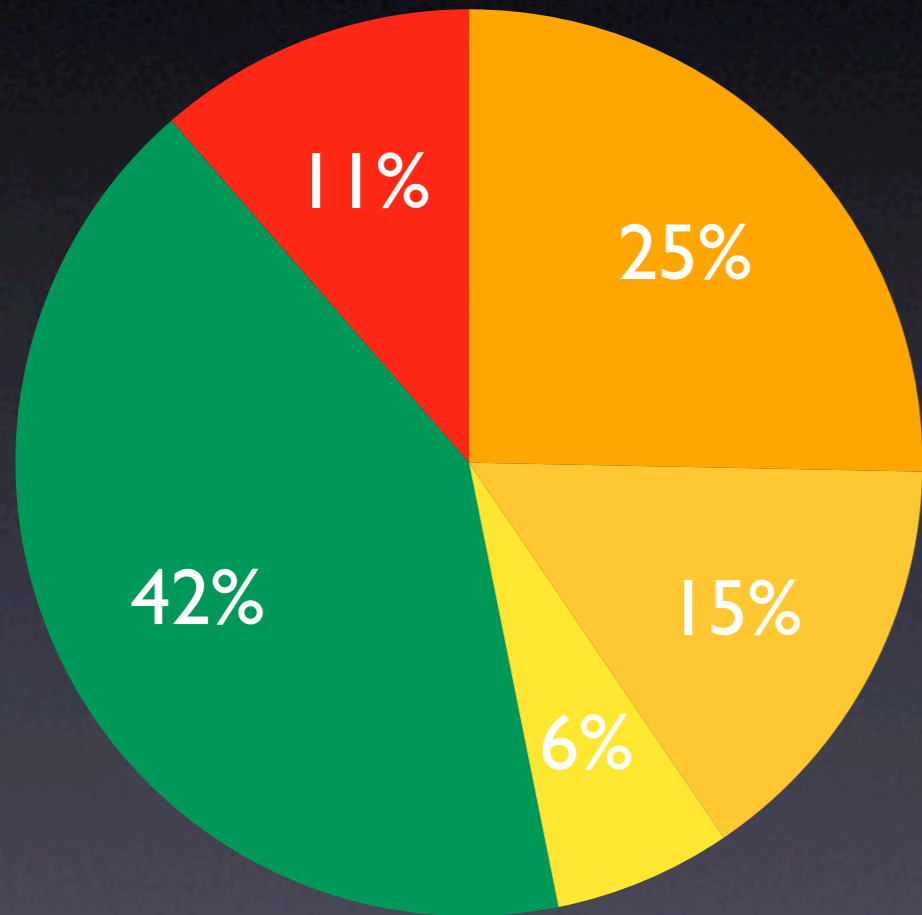
$$P_c = 0.367 \pm 0.037 \pm 0.033 \pm 0.009$$



● theory ● $m_c(m_c)$ ● $\alpha_s(M_Z)$

Error budget of P_c at NLO

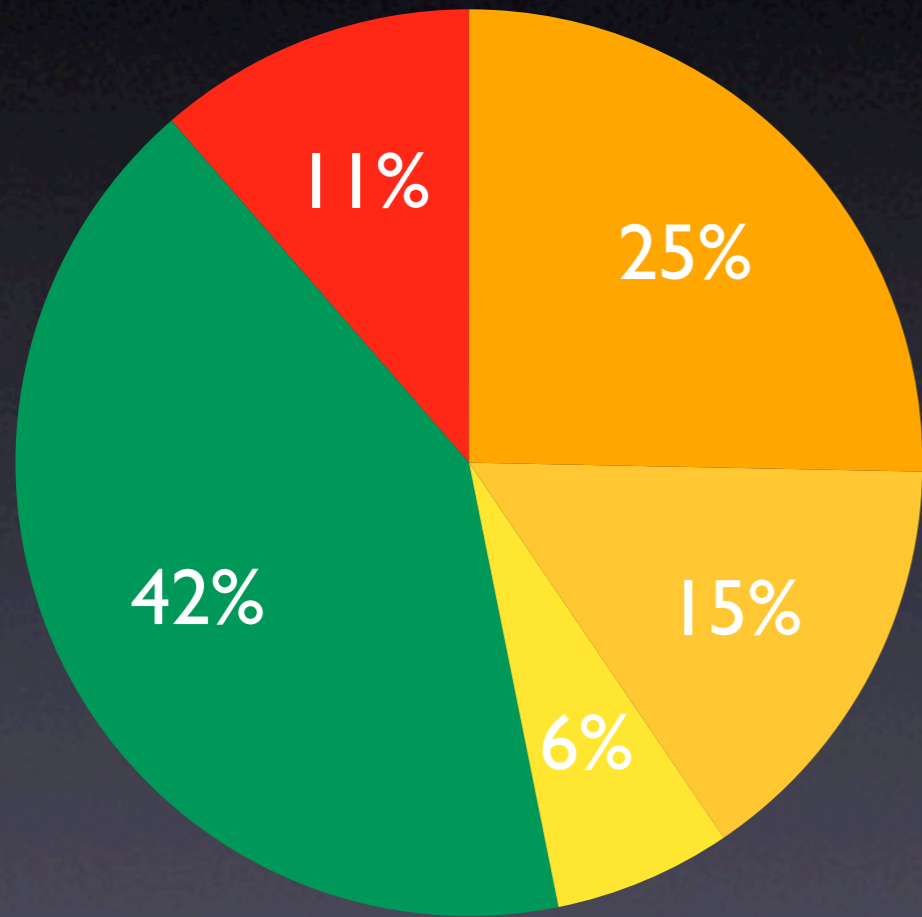
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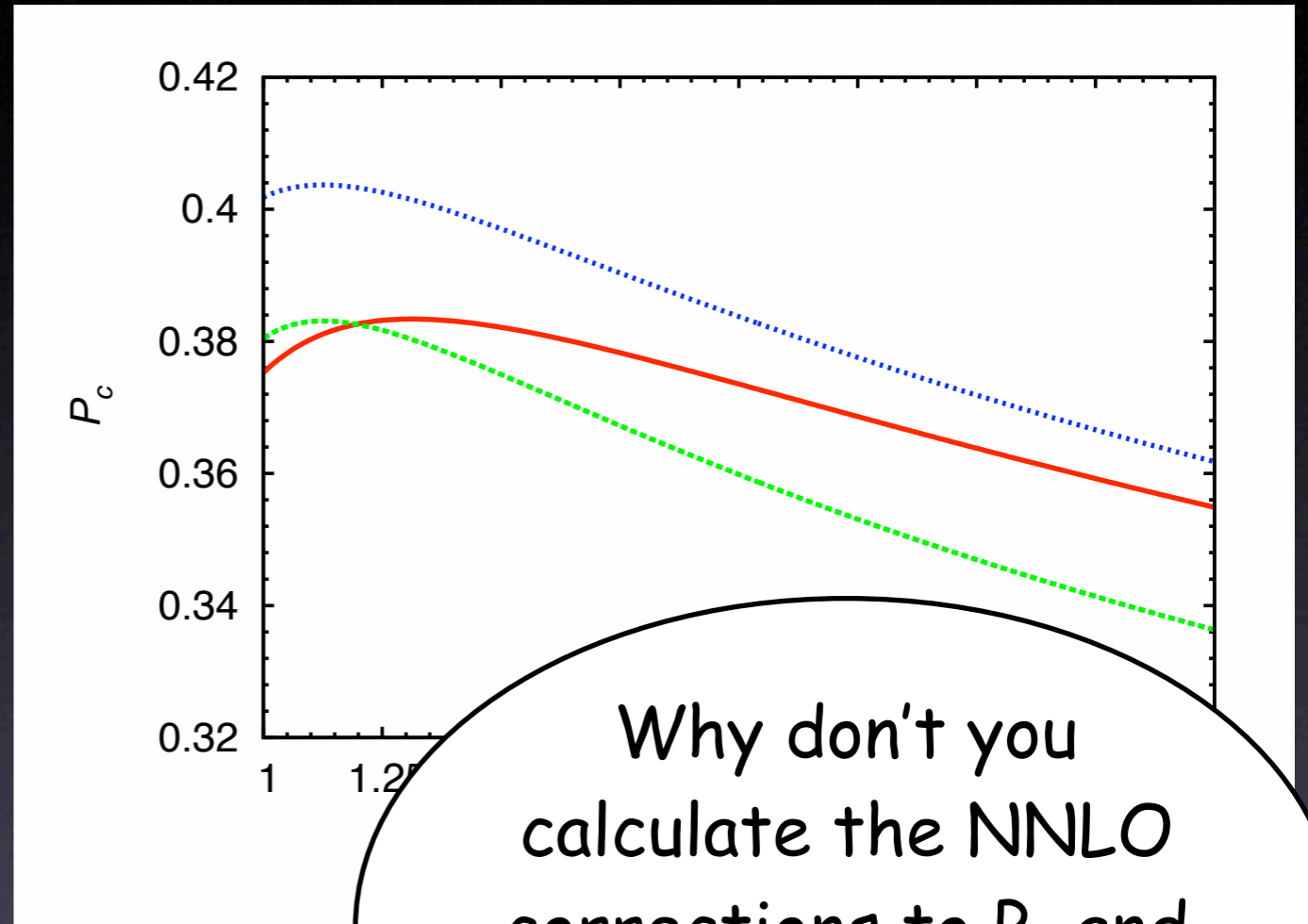
Buras, Gorbahn, Nierste & UH '05

Error budget of P_c at NLO

$$P_c = 0.367 \pm 0.037 \pm 0.033 \pm 0.009$$



- μ_c
- $\alpha_s(\mu_c)$
- $\mu_b + \mu_W$
- $m_c(m_c)$
- $\alpha_s(M_Z)$

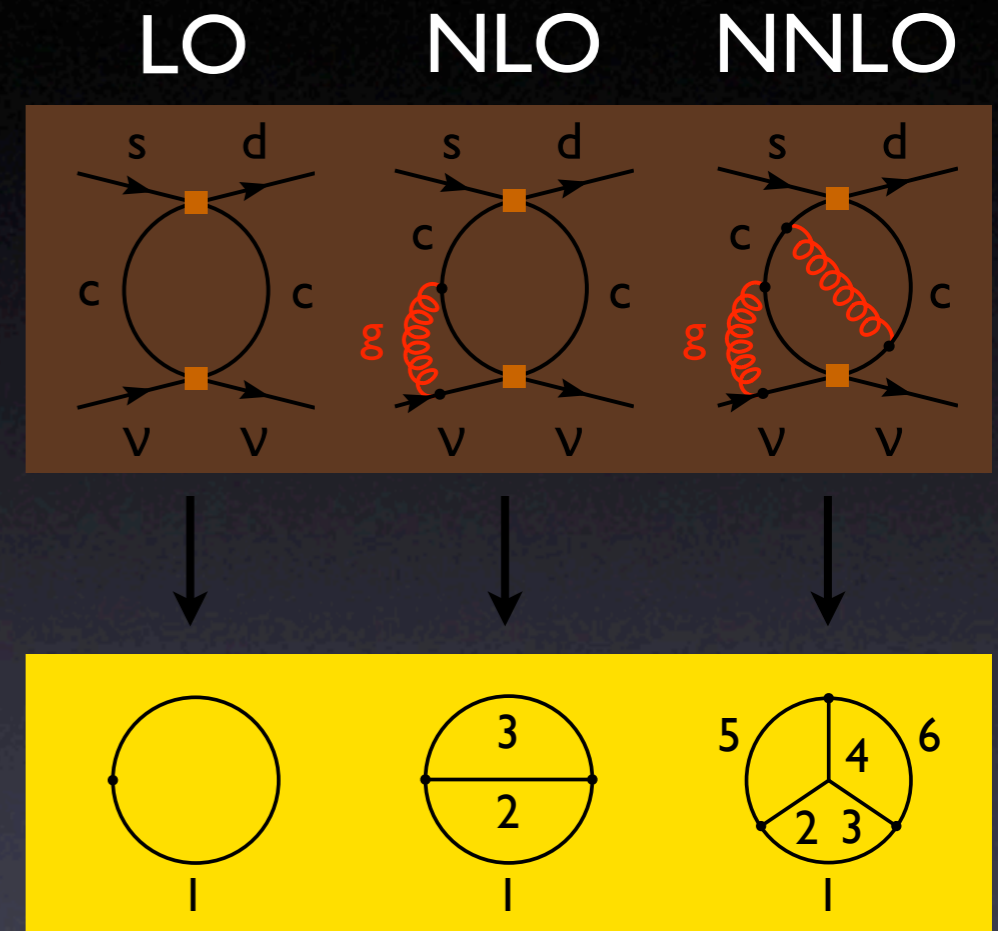


Why don't you calculate the NNLO corrections to P_c and reduce the theoretical error?



NNLO ADM calculation of P_c

- ADM is determined from $1/\epsilon_{UV}$ of 1-, 2- and 3-loop diagrams
- integrals have $1/\epsilon_{IR}$ and $1/\epsilon_{UV}$ that are indistinguishable in DR
- in \overline{MS} scheme $1/\epsilon_{UV}$ are polynomial in masses and momenta after subtraction of subdivergences
- calculation of counterterms reduces to computation of massive tadpoles



$$\frac{1}{(k+p)^2 - m^2} = \frac{1}{k^2 - M^2} - \frac{p^2 + 2k \cdot p - m^2 + M^2}{k^2 - M^2} \frac{1}{(k+p)^2 - m^2}$$

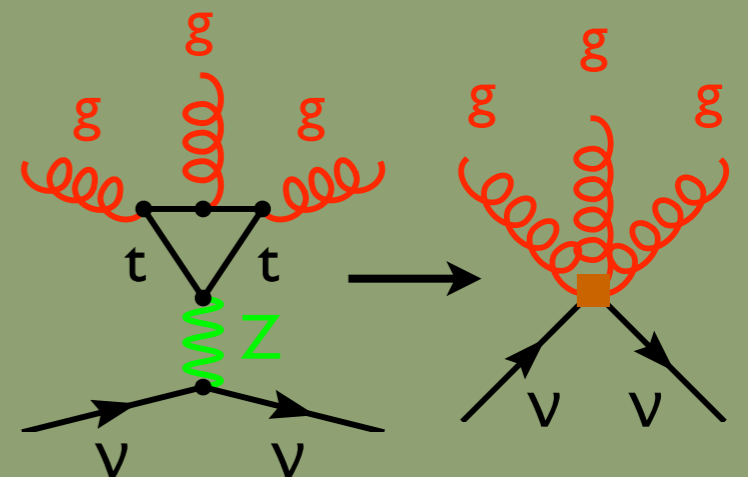
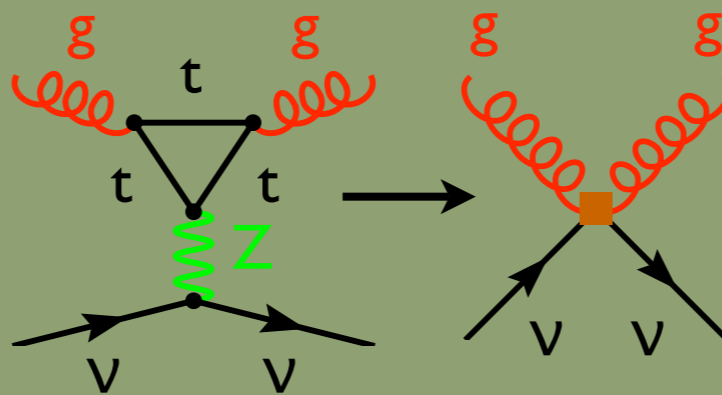
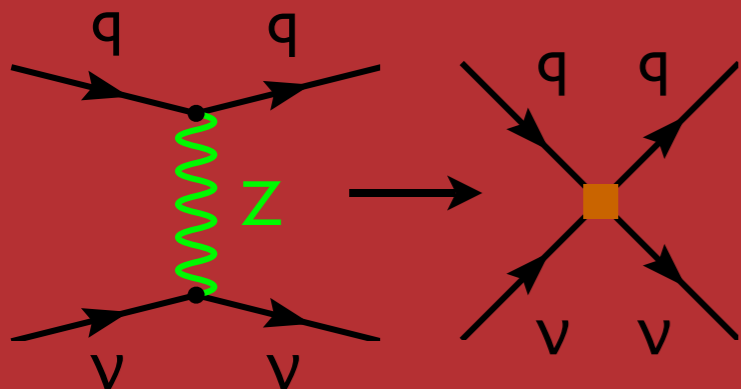
New features in NNLO calculation of P_c

$$Q_Z = \sum_q \left((I_q^3 - 2e_q \sin^2 \theta_w) Q_V^q - I_q^3 (Q_A^q + Q_{CS}) \right) \quad a_{CS} = \begin{cases} 2 & \text{HV} \\ \frac{2}{3} & \text{DRED} \end{cases}$$

$$Q_V^q = \sum_{l=e,\mu,\tau} (\bar{q} \gamma_\mu q) (\bar{\nu}_{lL} \gamma^\mu \nu_{lL})$$

$$Q_A^q = \sum_{l=e,\mu,\tau} (\bar{q} \gamma_\mu \gamma_5 q) (\bar{\nu}_{lL} \gamma^\mu \nu_{lL})$$

$$Q_{CS} = \frac{g^2}{16\pi^2} a_{CS} \epsilon^{\mu\nu\lambda\kappa} \left(G_\mu^a \partial_\nu G_\lambda^a + \frac{1}{3} g f^{abc} G_\mu^a G_\nu^b G_\lambda^c \right) \sum_{l=e,\mu,\tau} (\bar{\nu}_{lL} \gamma_\kappa \nu_{lL})$$



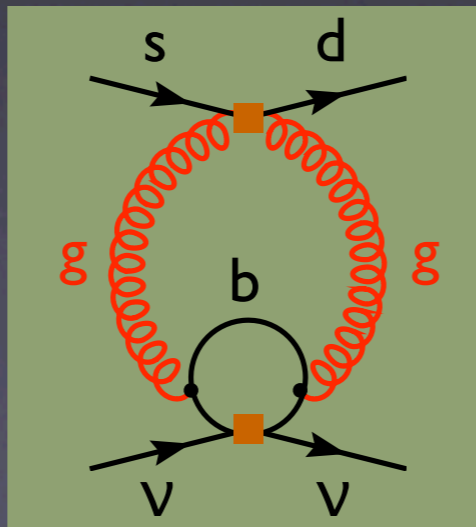
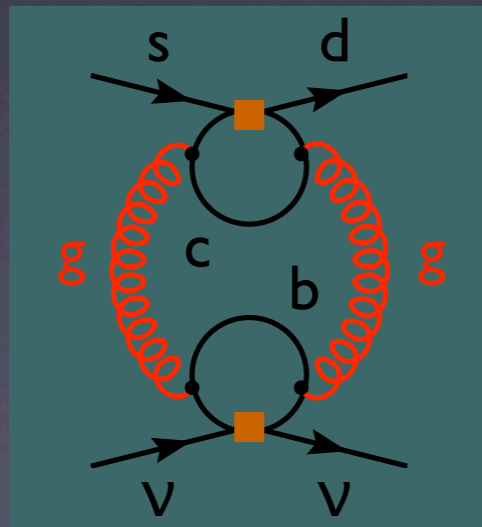
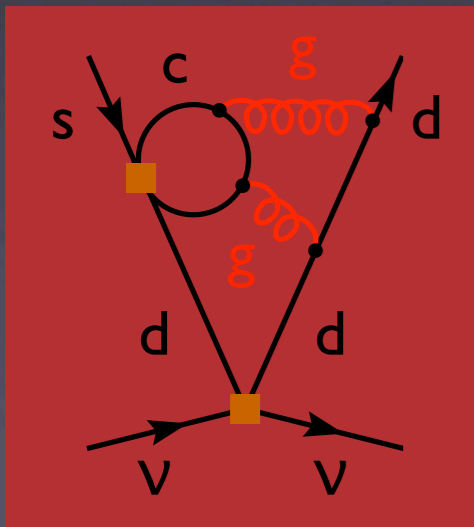
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$$Q_V^q = \sum_{l=e,\mu,\tau} (\bar{q} \gamma_\mu q) (\bar{\nu}_{lL} \gamma^\mu \nu_{lL})$$

$$Q_A^q = \sum_{l=e,\mu,\tau} (\bar{q} \gamma_\mu \gamma_5 q) (\bar{\nu}_{lL} \gamma^\mu \nu_{lL})$$

$$Q_{CS} = \frac{g^2}{16\pi^2} a_{CS} \epsilon^{\mu\nu\lambda\kappa} \left(G_\mu^a \partial_\nu G_\lambda^a + \frac{1}{3} g f^{abc} G_\mu^a G_\nu^b G_\lambda^c \right) \sum_{l=e,\mu,\tau} (\bar{\nu}_{lL} \gamma_\kappa \nu_{lL})$$



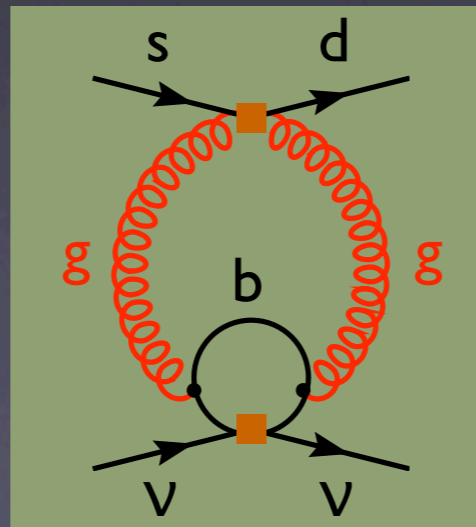
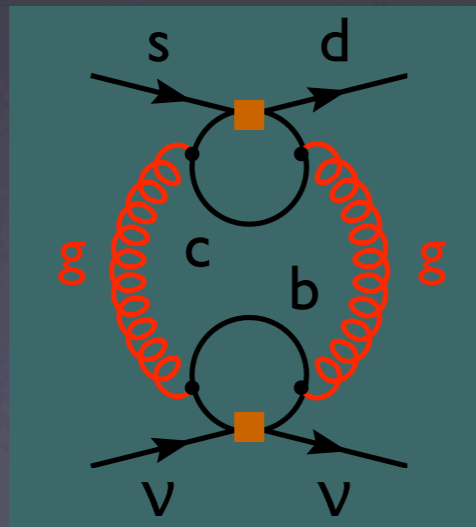
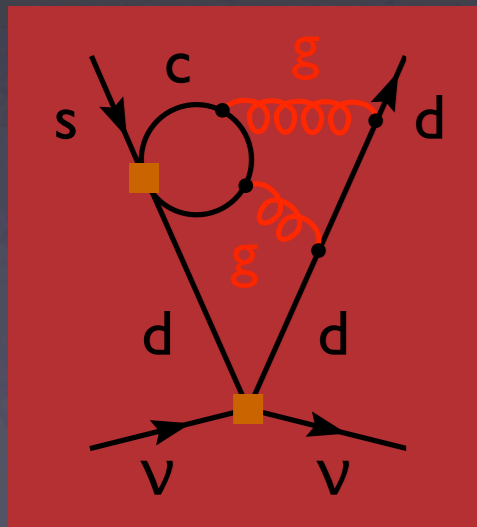
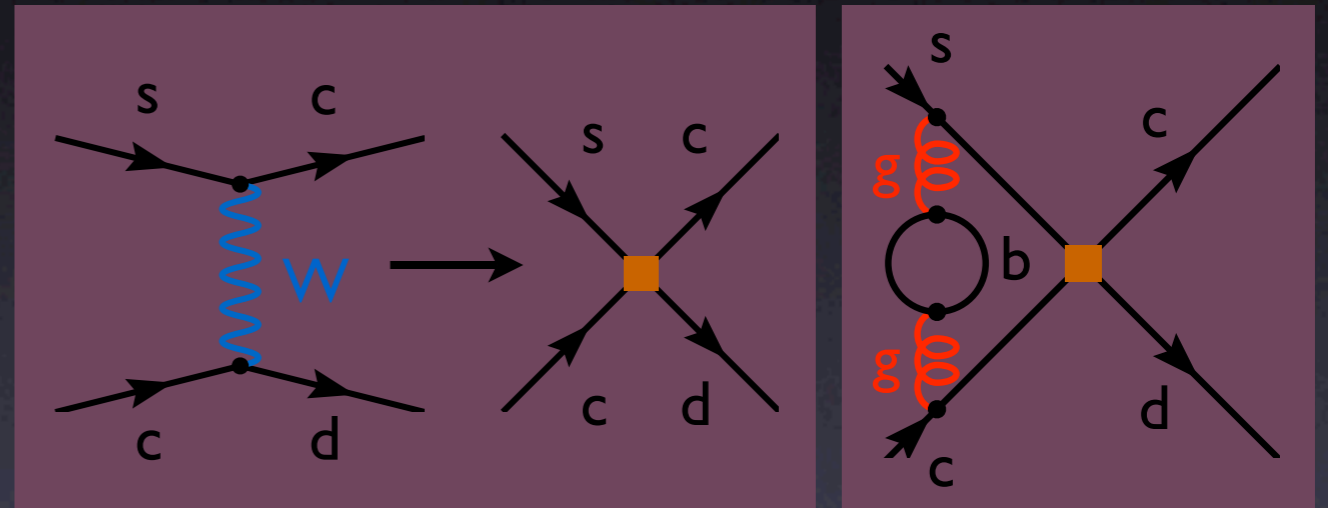
- vector part of neutral-current coupling starts to contribute
- contributions from axial-vector and Chern-Simons operator cancel exactly

New features in NNLO calculation of P_c

$$Q_Z = \sum_q ((I_q^3 - 2e_q \sin^2 \theta_w) Q_V^q - I_q^3 (Q_A^q + Q_{CS}))$$

$$Q_1 = (\bar{s}_L \gamma_\mu c_L)(\bar{c}_L \gamma^\mu d_L)$$

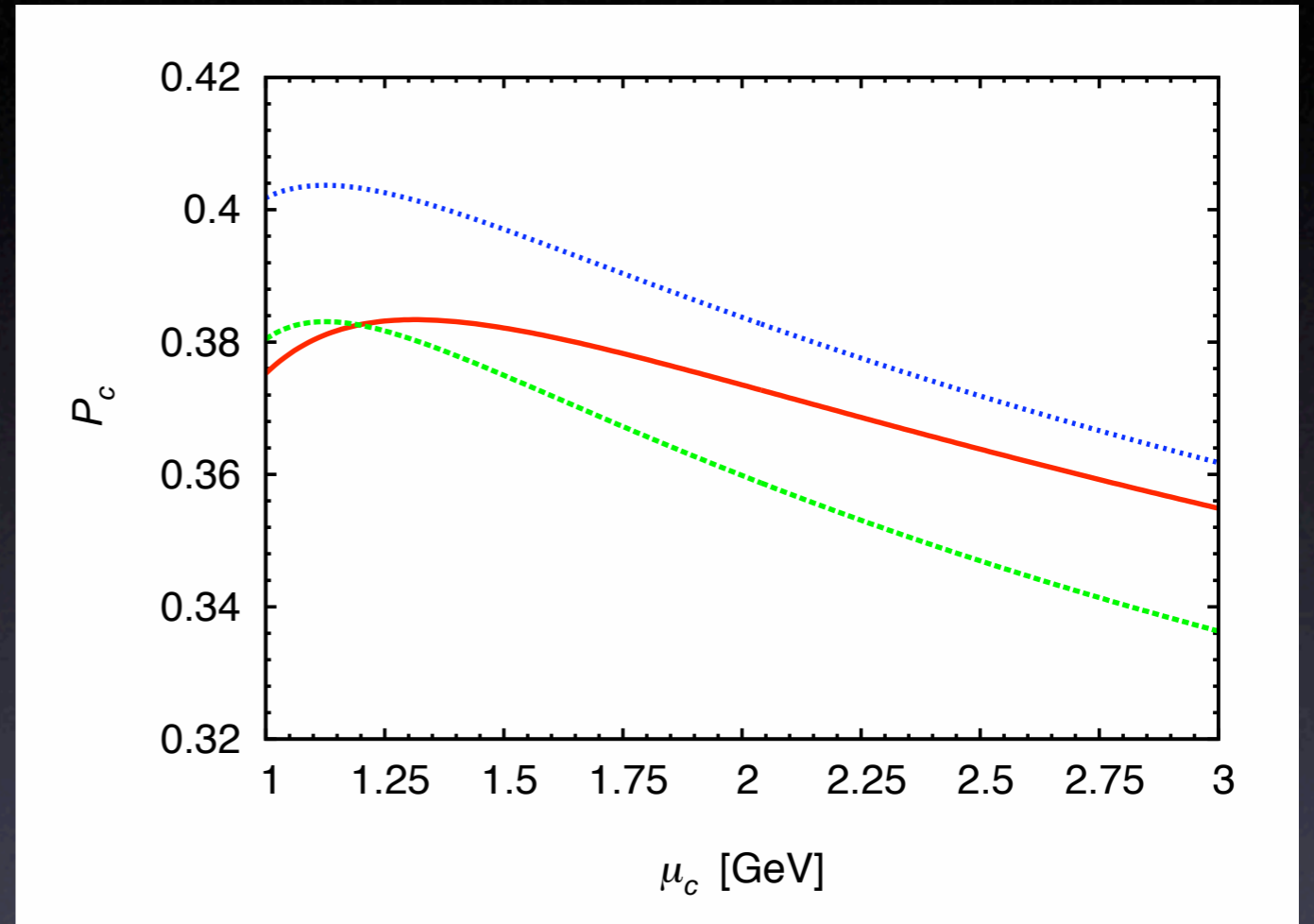
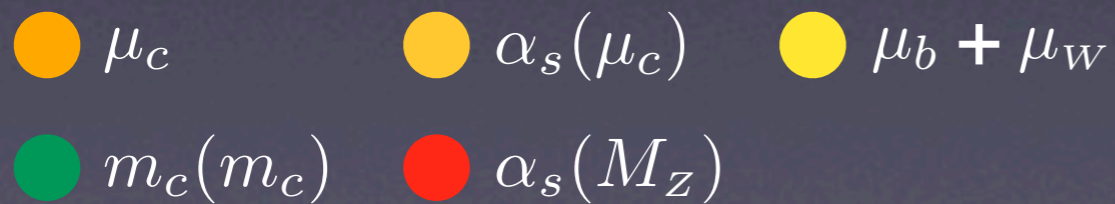
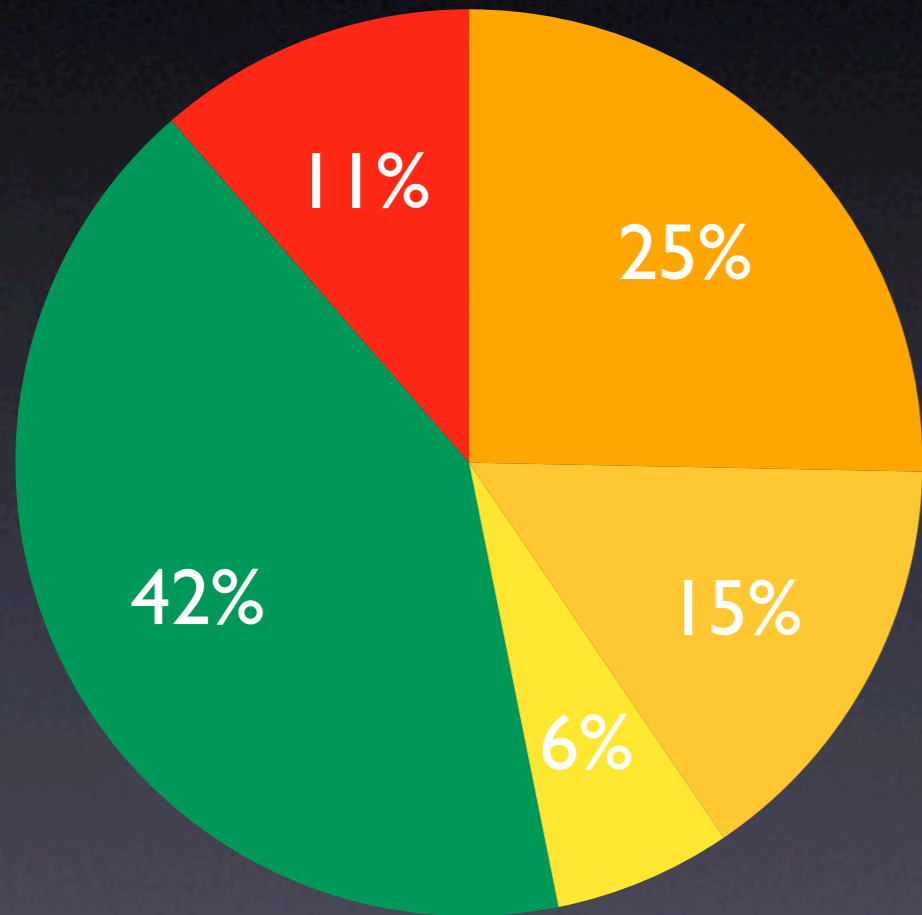
$$Q_2 = (\bar{s}_L \gamma_\mu T^a c_L)(\bar{c}_L \gamma^\mu T^a d_L)$$



- non-trivial matching corrections at bottom quark threshold arise for current-current operators

Error budget of P_c and $B(K^+)$ at NLO

$$P_c = 0.367 \pm 0.037 \pm 0.033 \pm 0.009$$

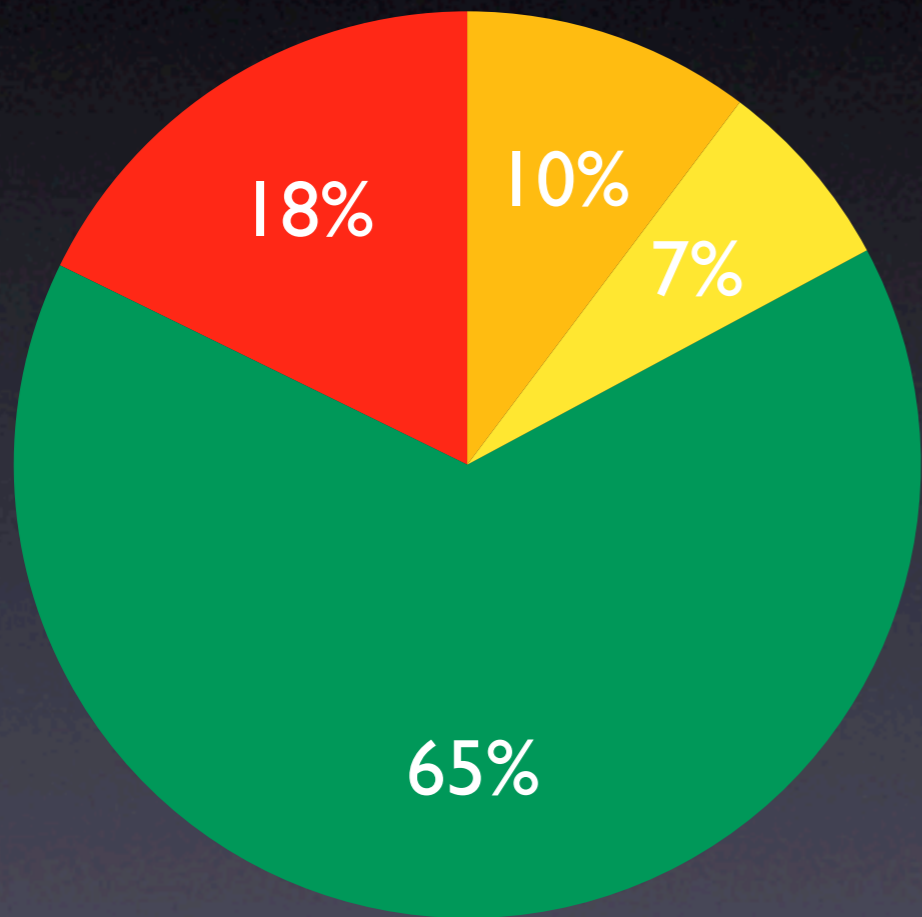


$$\mathcal{B}(K^+) = (7.93 \pm 0.77 P_c \pm 0.84_{\text{other}}) \times 10^{-11}$$

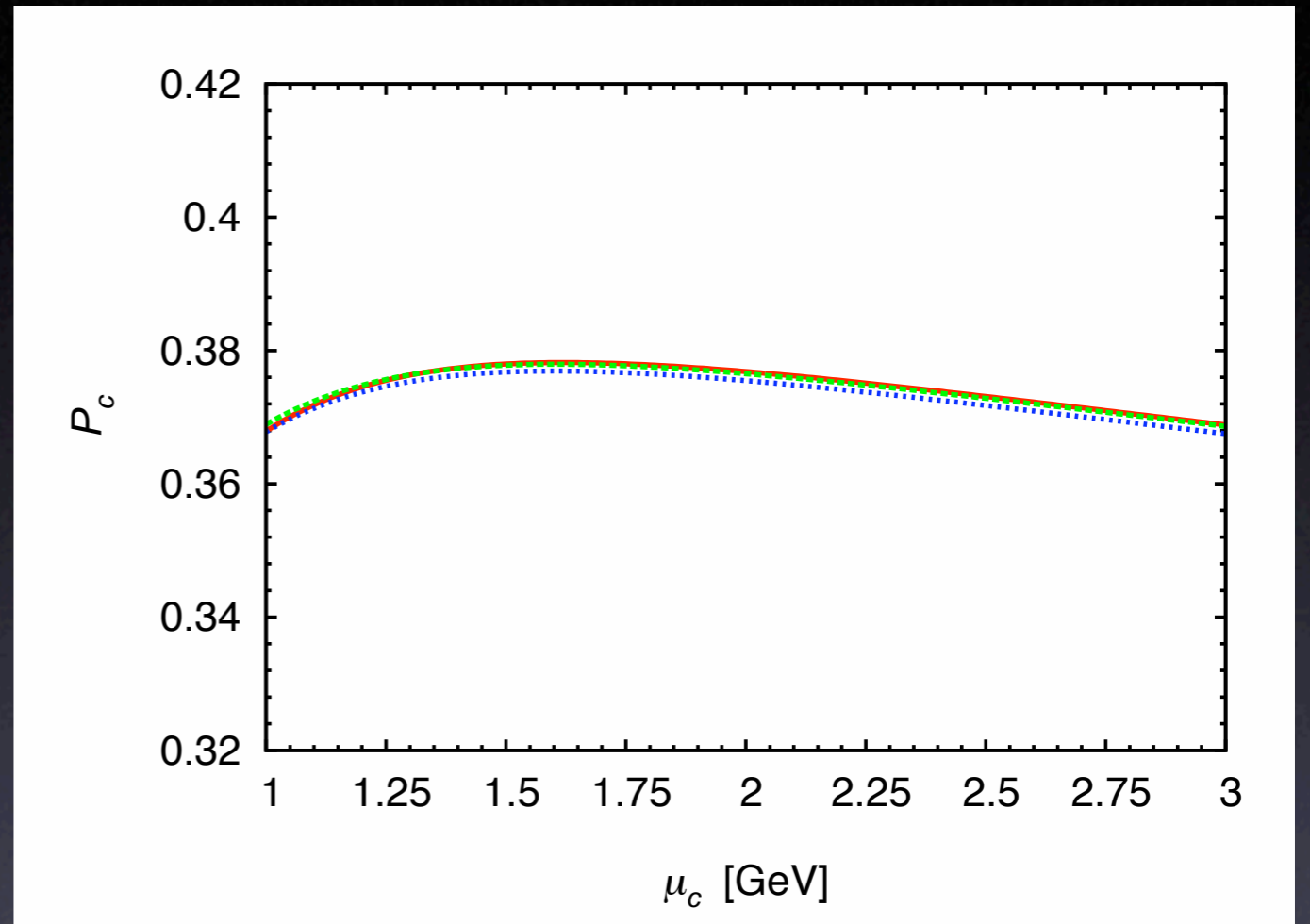
Buras, Gorbahn, Nierste & UH '05

Error budget of P_c and $B(K^+)$ at NNLO

$$P_c = 0.371 \pm 0.009 \pm 0.031 \pm 0.009$$



- $\mu_c + \alpha_s(\mu_c)$ ● $\mu_b + \mu_W$
- $m_c(m_c)$ ● $\alpha_s(M_Z)$

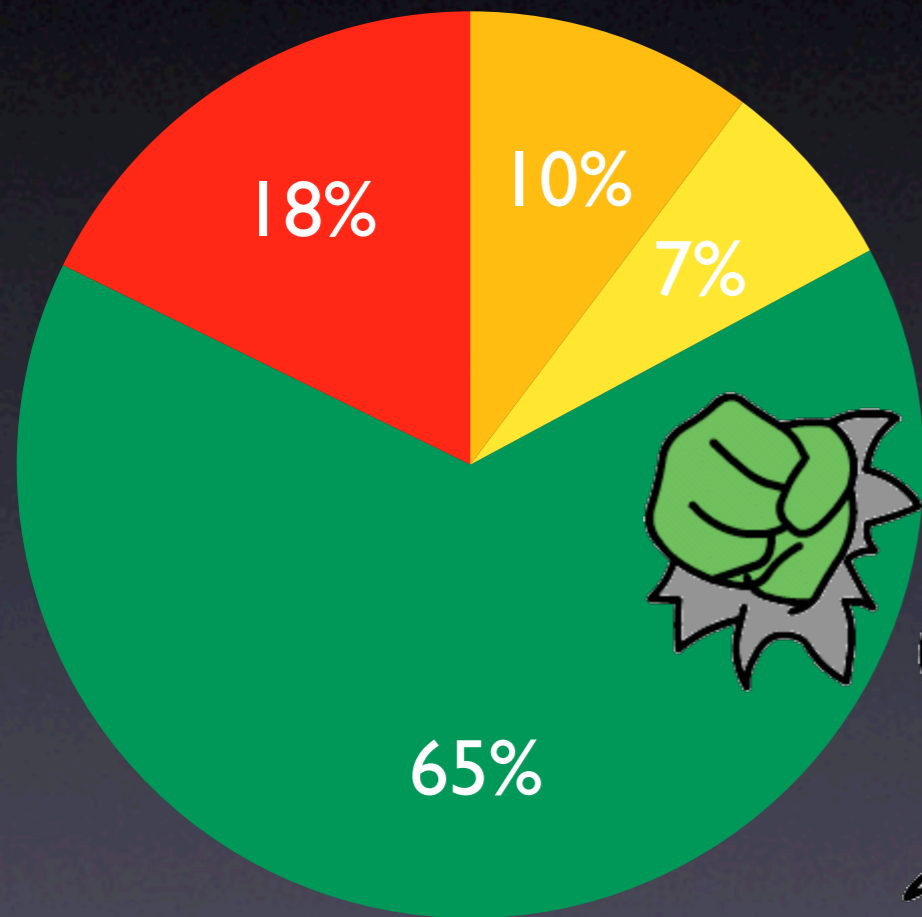


$$\mathcal{B}(K^+) = (7.96 \pm 0.49_{P_c} \pm 0.84_{\text{other}}) \times 10^{-11}$$

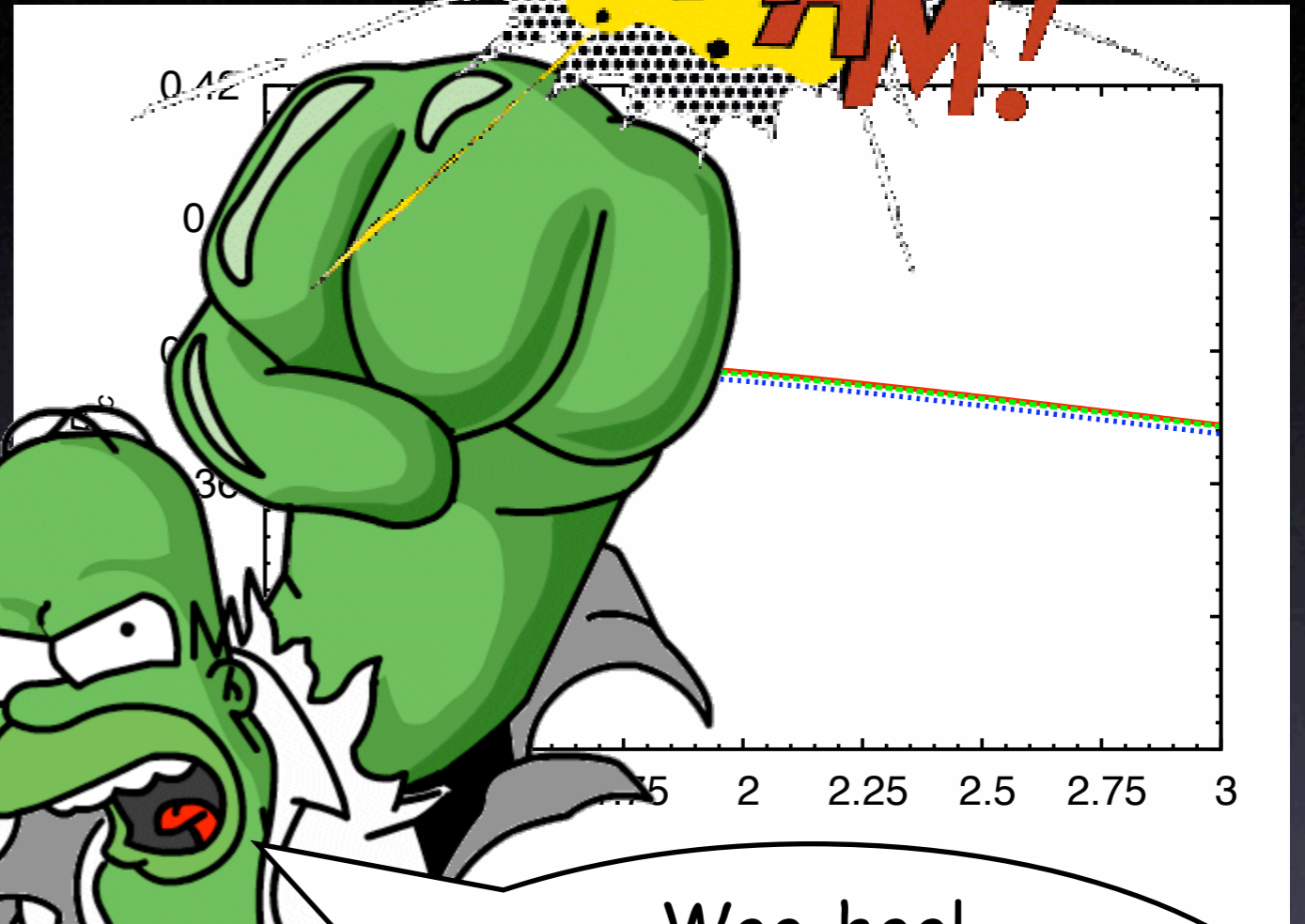
Buras, Gorbahn, Nierste & UH '05

Error budget of P_c and $B(K^+)$

$$P_c = 0.371 \pm 0.009 \pm 0.031 \pm 0.009$$



- $\mu_c + \alpha_s(\mu_c)$
- $\mu_b + \mu_w$
- $m_c(m_c)$
- $\alpha_s(M_Z)$



Woo-hoo!
NNLO calculation reduces
theoretical error of P_c and
 $B(K^+)$ by factor 4!

Unitarity triangle from $K \rightarrow \pi\nu\bar{\nu}$

d

s

b

$1 - \frac{\lambda^2}{2}$	λ	V_{ub}
$-\lambda$	$1 - \frac{\lambda^2}{2}$	V_{cb}
V_{td}	V_{ts}	1

u

c

t

$$A \approx 0.82$$

$$\bar{\rho} = \rho \left(1 - \frac{\lambda^2}{2} \right)$$

$$\bar{\eta} = \eta \left(1 - \frac{\lambda^2}{2} \right)$$

$$V_{cb} = -V_{ts} = A\lambda^2$$

$$V_{ub} = A\lambda^3(\rho - i\eta)$$

$$V_{td} = A\lambda^3(1 - \bar{\rho} - i\bar{\eta})$$

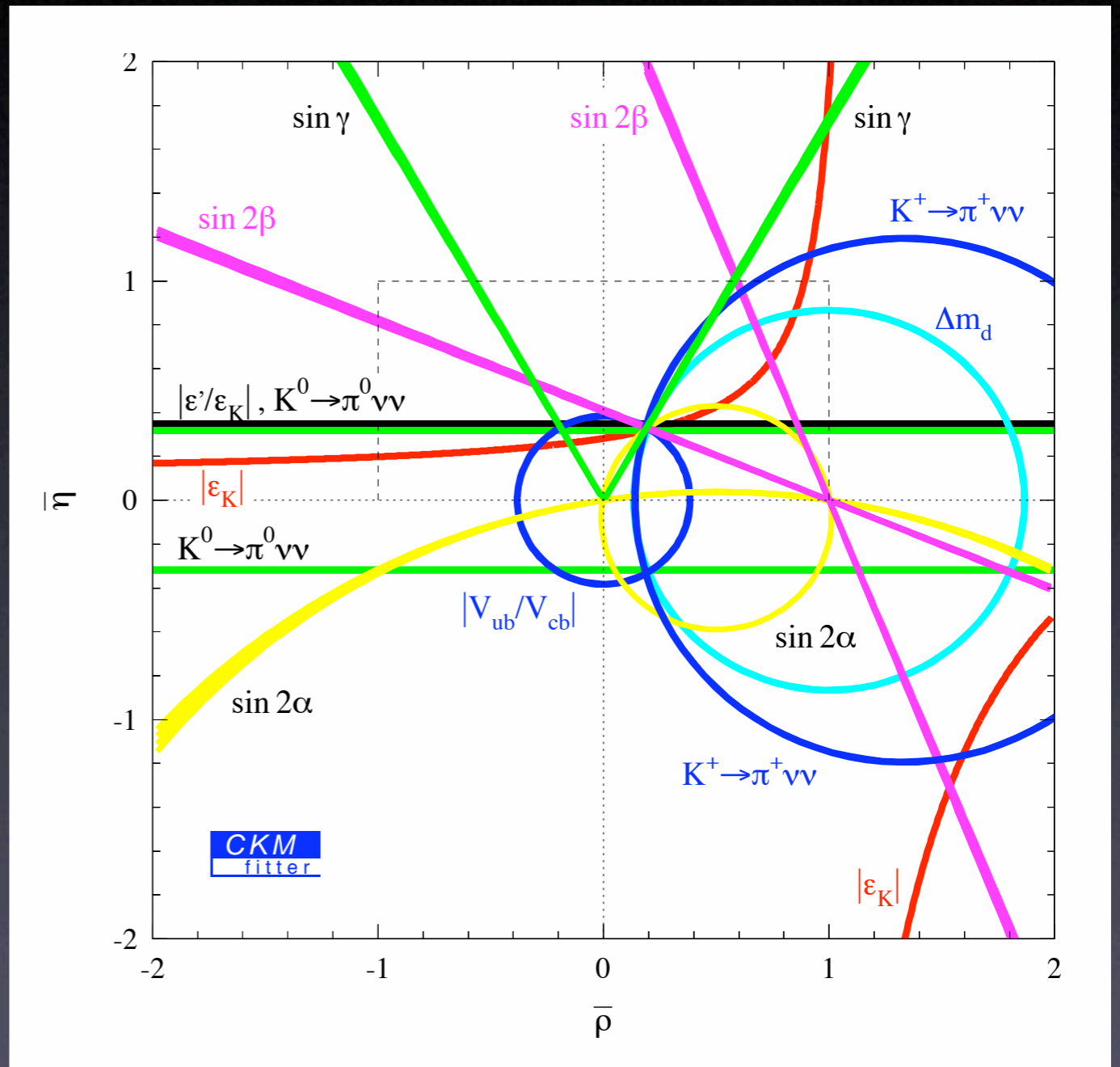
Unitarity triangle from $K \rightarrow \pi \nu \bar{\nu}$

	d	s	b	
u	$1 - \frac{\lambda^2}{2}$	λ	V_{ub}	
c	$-\lambda$	$1 - \frac{\lambda^2}{2}$	V_{cb}	
t	V_{td}	V_{ts}	1	

$$V_{cb} = -V_{ts} = A\lambda^2$$

$$V_{ub} = A\lambda^3(\rho - i\eta)$$

$$V_{td} = A\lambda^3(1 - \bar{\rho} - i\bar{\eta})$$



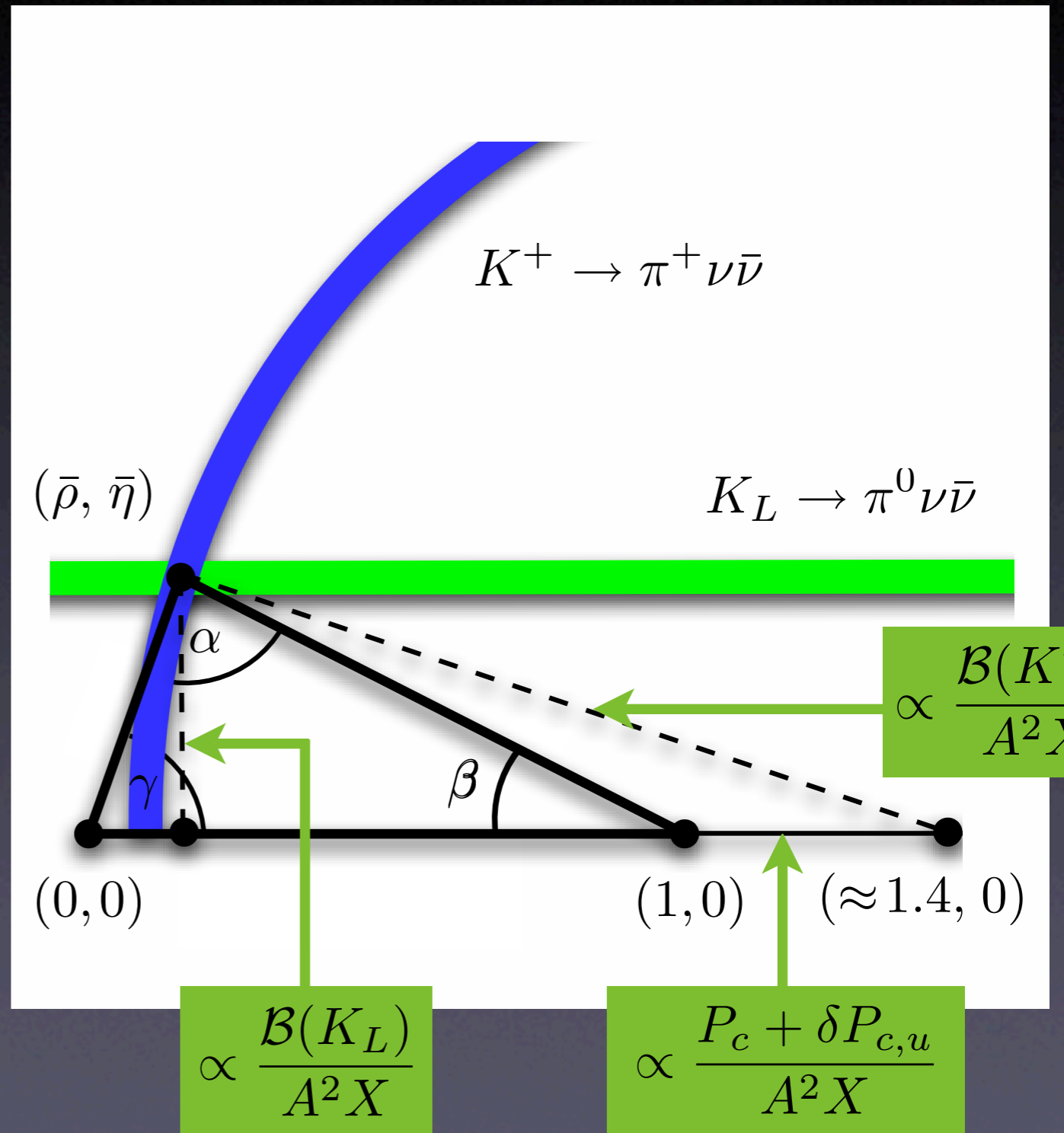
Unitarity triangle from $K \rightarrow \pi \nu \bar{\nu}$

	d	s	b	
u	$1 - \frac{\lambda^2}{2}$	λ	V_{ub}	
c	$-\lambda$	$1 - \frac{\lambda^2}{2}$	V_{cb}	
t	V_{td}	V_{ts}	1	

$$V_{cb} = -V_{ts} = A\lambda^2$$

$$V_{ub} = A\lambda^3(\rho - i\eta)$$

$$V_{td} = A\lambda^3(1 - \bar{\rho} - i\bar{\eta})$$



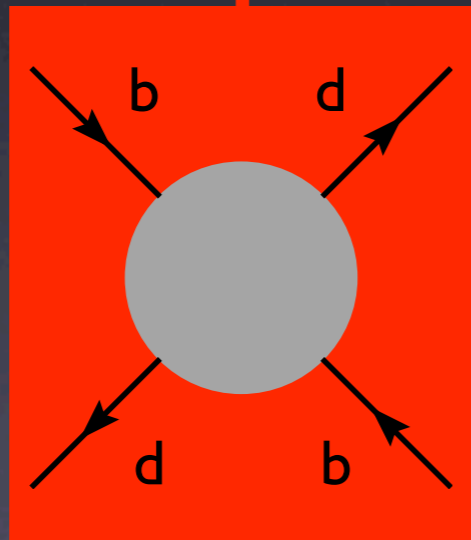
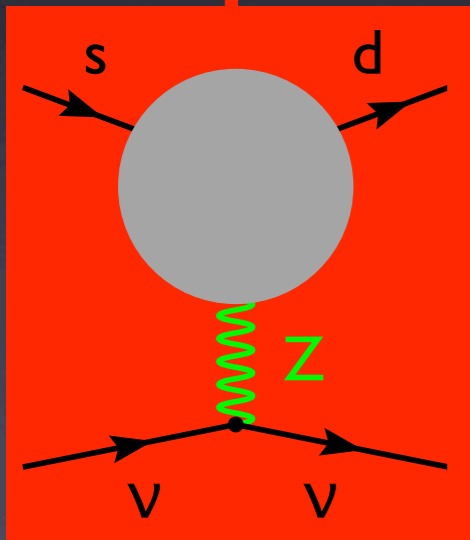
Unitarity triangle from $K \rightarrow \pi \nu \bar{\nu}$

$$(\sin 2\beta)_{K \rightarrow \pi \nu \bar{\nu}} = (\sin 2\beta)_{B \rightarrow \psi K_s}$$

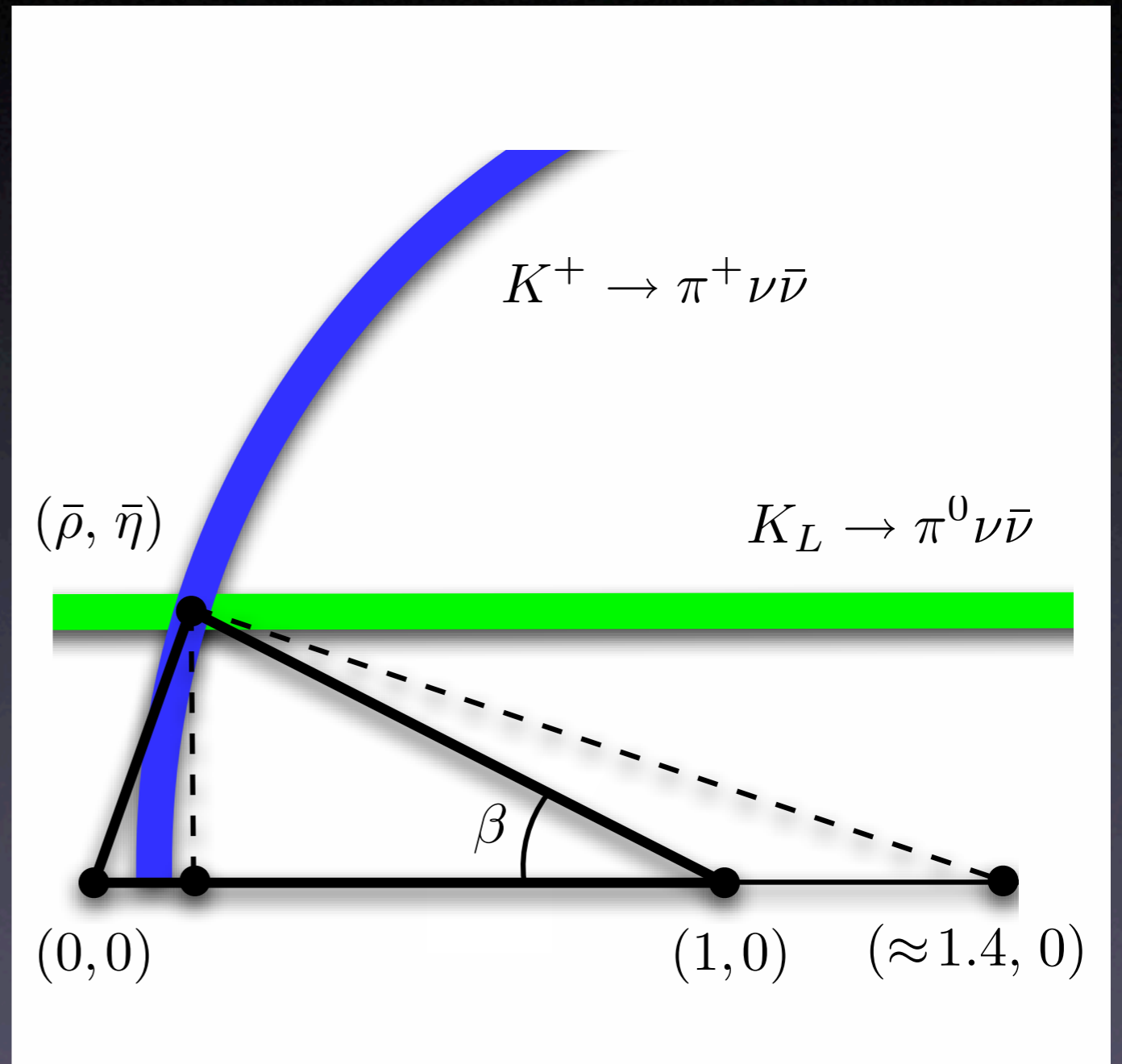
β

SM, MFV

β



K physics \longleftrightarrow B physics



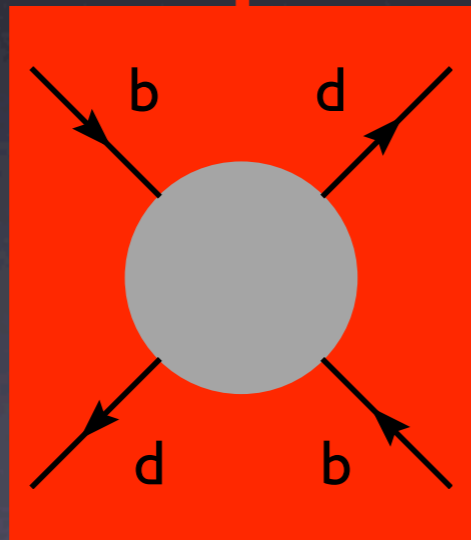
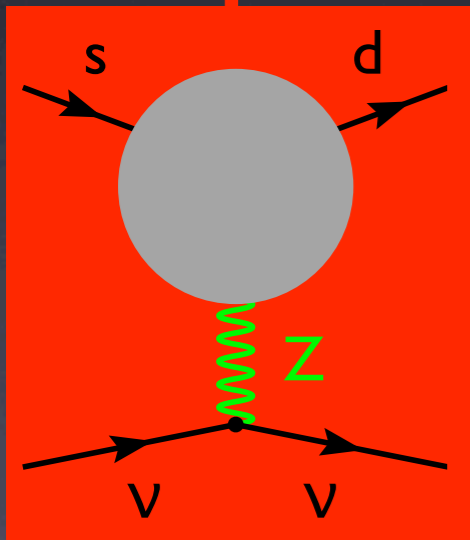
Unitarity triangle from $K \rightarrow \pi \nu \bar{\nu}$

$$(\sin 2\beta_X)_{K \rightarrow \pi \nu \bar{\nu}} \neq (\sin 2\beta_S)_{B \rightarrow \psi K_S}$$

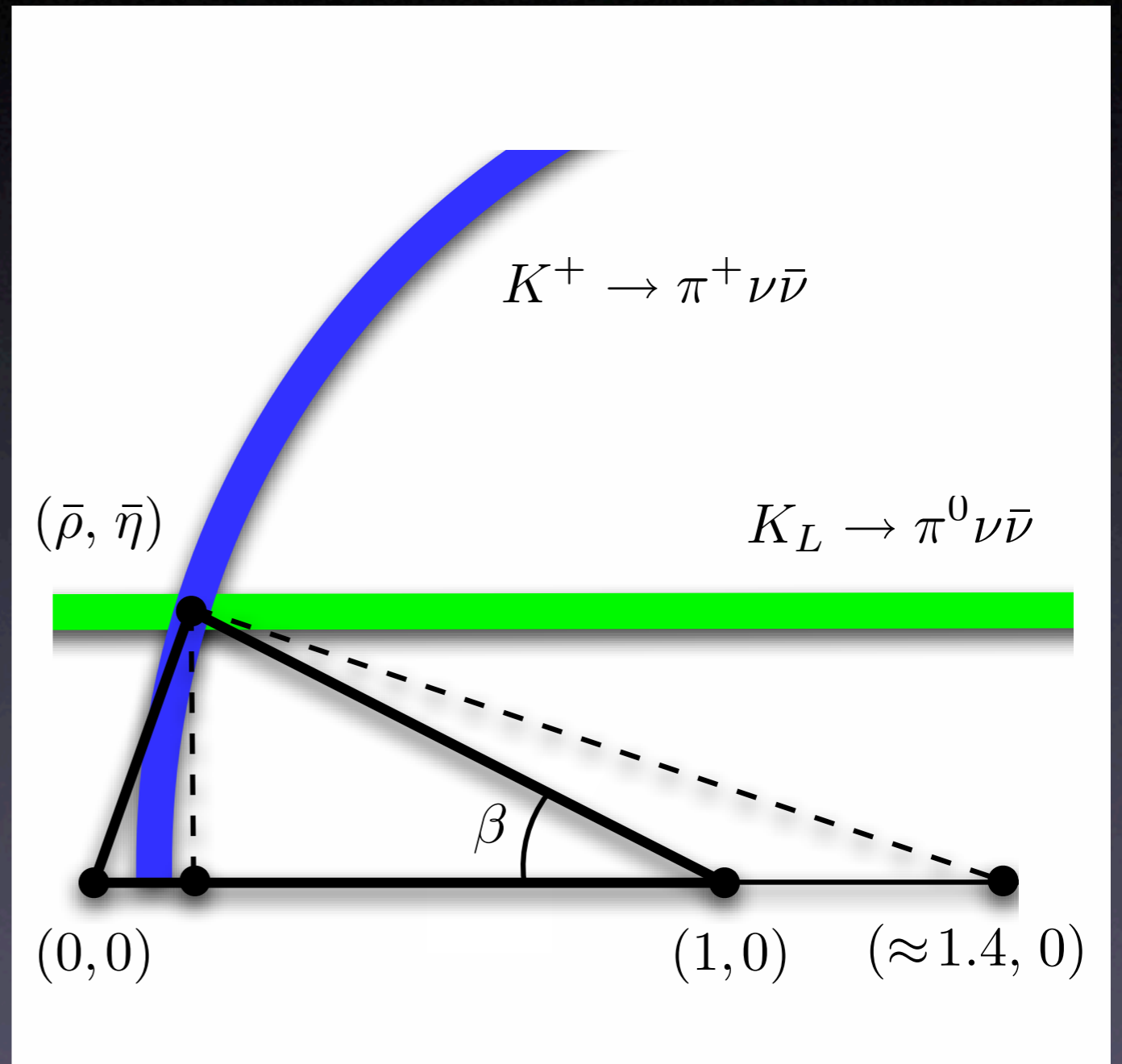
$\beta + \theta_X$

non-MFV

$\beta + \theta_S$



K physics \longleftrightarrow B physics



Unitarity triangle fit from $K \rightarrow \pi\nu\bar{\nu}$

$$\mathcal{B}(K^+) = (14.7_{-8.9}^{+13.0}) \times 10^{-11}$$

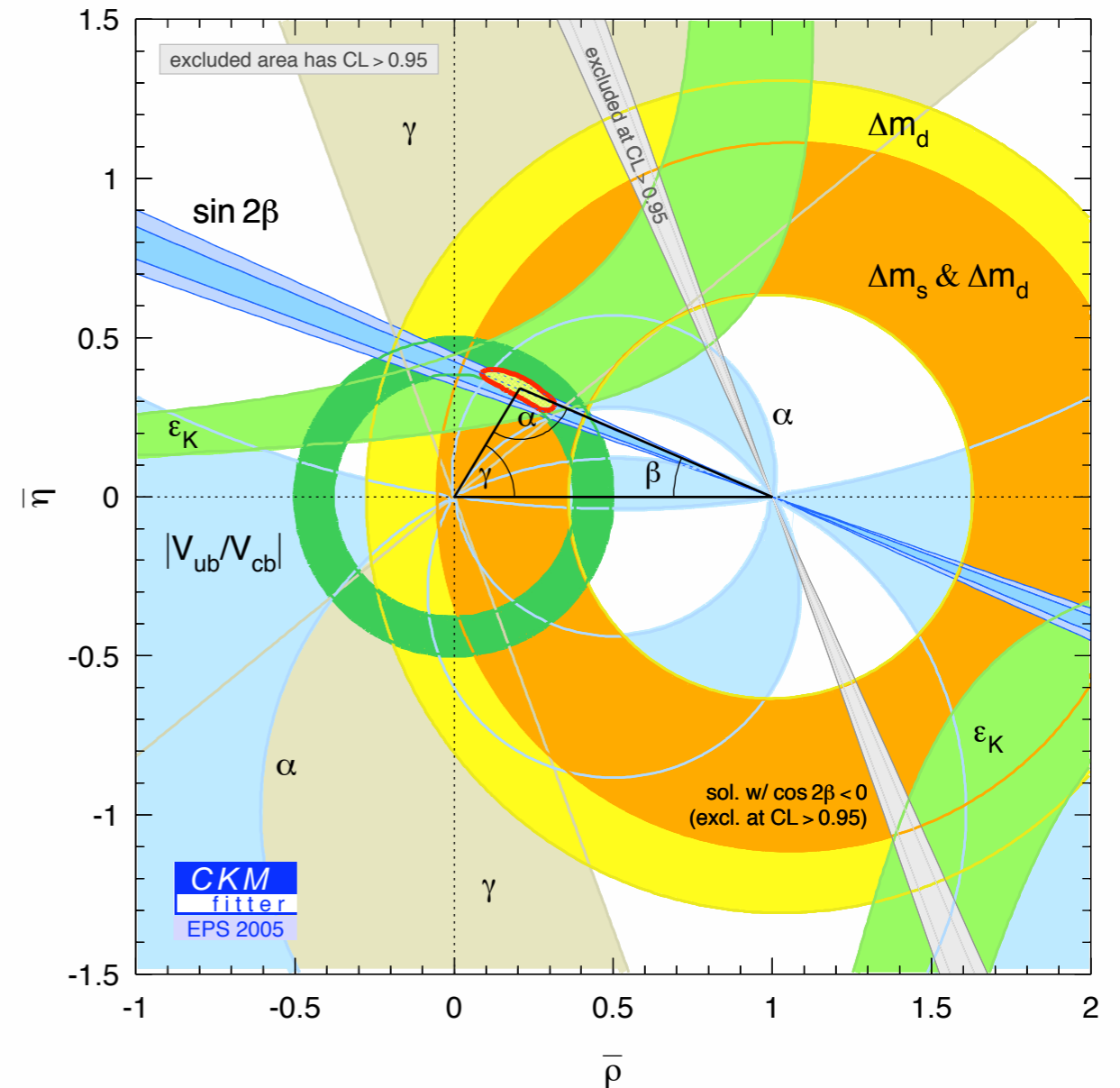
BNL AGS E787 & E949 '04

$$\mathcal{B}(K_L) < 5.9 \times 10^{-7} \quad (90\% \text{ CL})$$

FNAL KTeV E799-II '00

$$\mathcal{B}(K_L) < 2.86 \times 10^{-7} \quad (90\% \text{ CL})$$

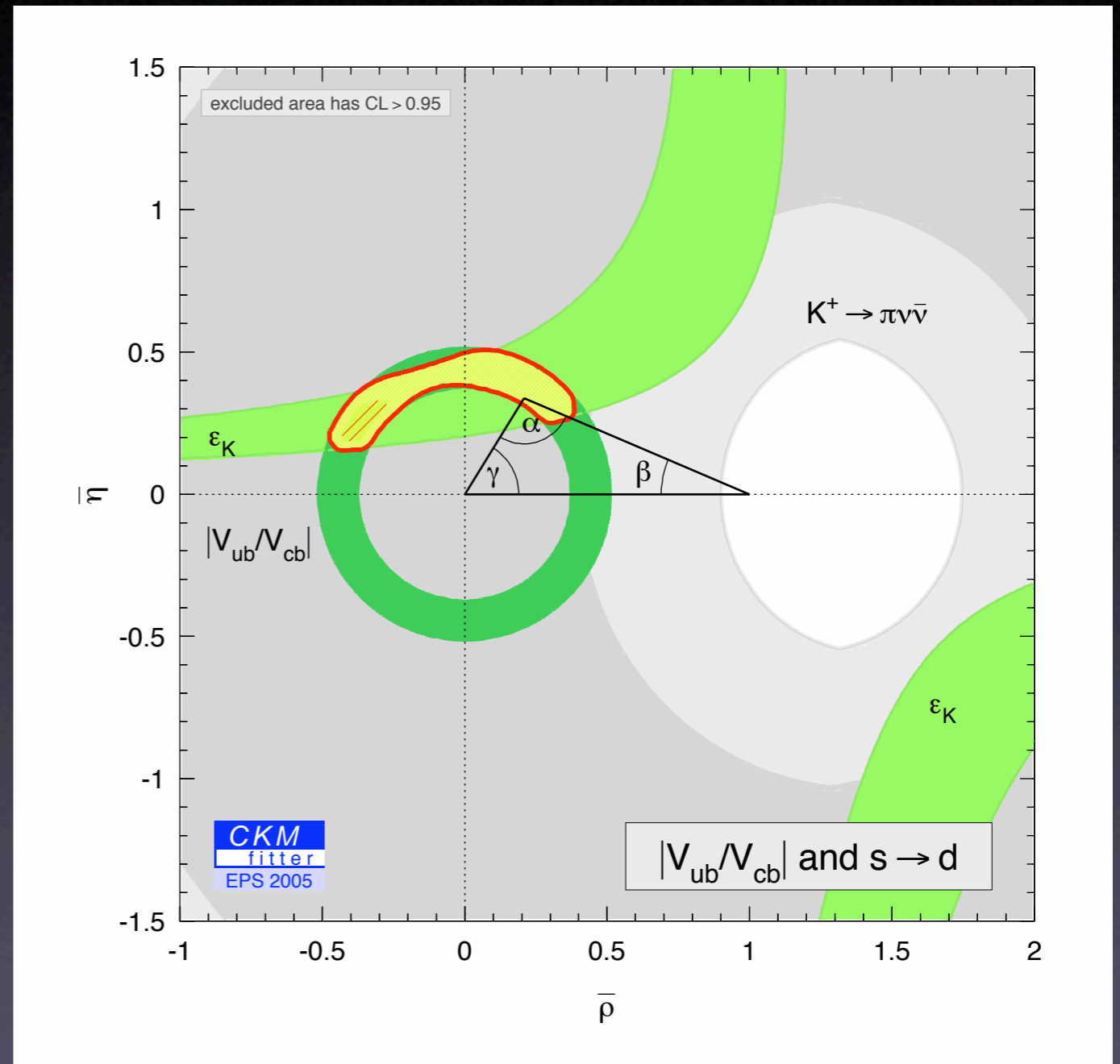
KEK PS E391a '05



Unitarity triangle fit from $K \rightarrow \pi \nu \bar{\nu}$

$$\mathcal{B}(K^+) = (14.7_{-8.9}^{+13.0}) \times 10^{-11}$$

BNL AGS E787 & E949 '04



Unitarity triangle fit from $K \rightarrow \pi\nu\bar{\nu}$

$$\mathcal{B}(K^+) = (7.50 \pm 0.75) \times 10^{-11}$$

$$\mathcal{B}(K_L) = (2.67 \pm 0.27) \times 10^{-11}$$

Future (?)

$$\frac{\sigma(|V_{td}|)}{|V_{td}|} = \pm 4.1\%$$

$$\sigma(\sin 2\beta) = \pm 0.025$$

$$\sigma(\gamma) = \pm 4.9^\circ$$

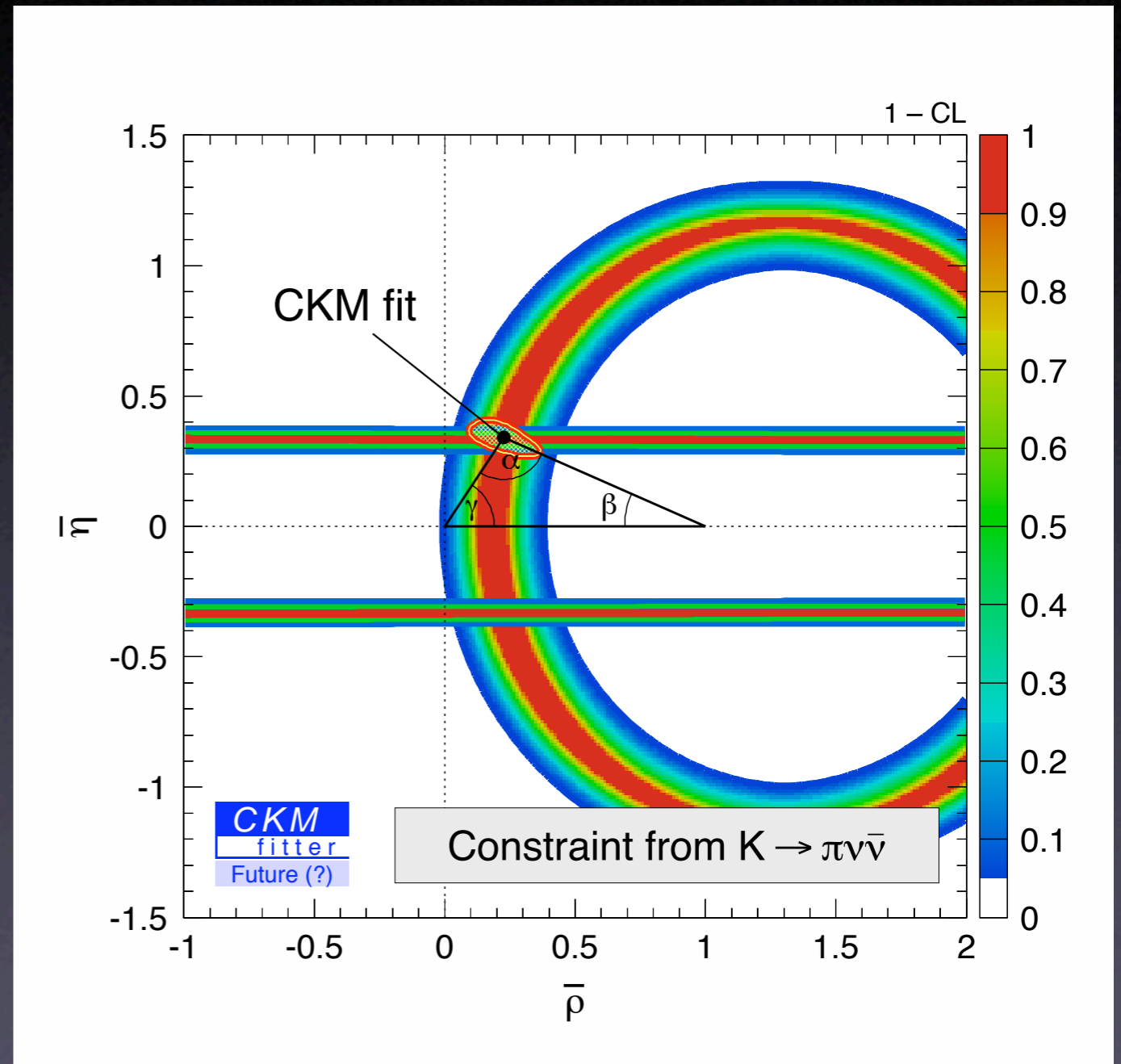
NLO
(theory error only)

$$\frac{\sigma(|V_{td}|)}{|V_{td}|} = \pm 1.0\%$$

$$\sigma(\sin 2\beta) = \pm 0.006$$

$$\sigma(\gamma) = \pm 1.2^\circ$$

NNLO
(theory error only)



Unitarity triangle fit from $K \rightarrow \pi\nu\bar{\nu}$

$$\mathcal{B}(K^+) = (7.50 \pm 0.75) \times 10^{-11}$$

$$\mathcal{B}(K_L) = (2.67 \pm 0.27) \times 10^{-11}$$

Future (?)

$$\frac{\sigma(|V_{td}|)}{|V_{td}|} = \pm 1.0\%$$

$$\sigma(\sin 2\beta) = \pm 0.006$$

$$\sigma(\gamma) = \pm 1.2^\circ$$

NNLO
(theory error only)

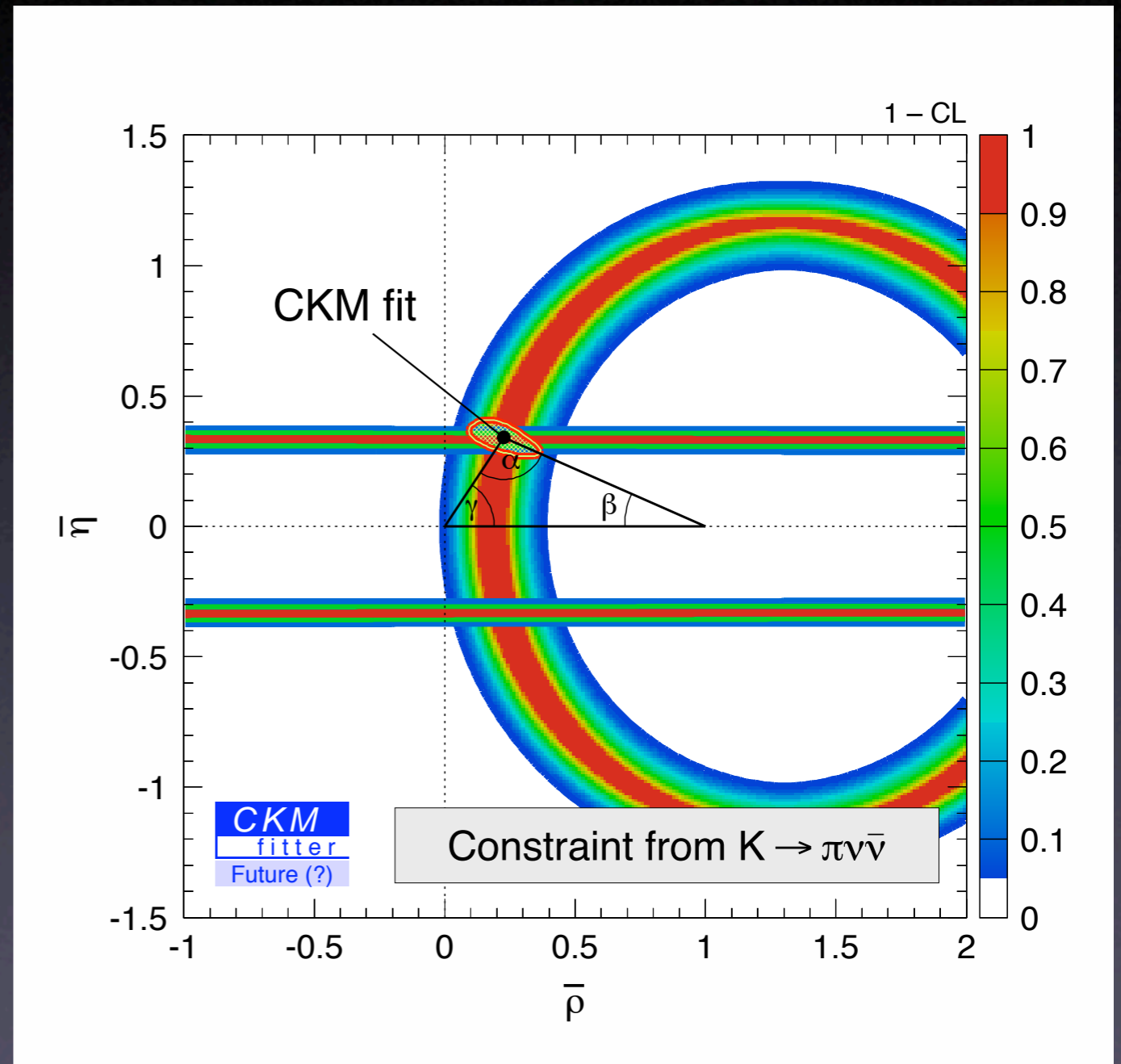
$$\frac{\sigma(|V_{td}|)}{|V_{td}|} = \pm 8.4\%$$

$$\sigma(\sin 2\beta) = \pm 0.056$$

$$\sigma(\gamma) = \pm 11^\circ$$

NNLO
(all uncertainties)

Future (?)



Conclusions

- SD dominated rare K^+ and K_L decays offer powerful and complementary test of flavor sector of SM
- NNLO calculation of charm contribution to K^+ is now available
- there is sizeable room for NP in this golden modes and their clean theoretical character remains valid in essential all extensions of SM
- measurements of branching ratios at $\approx 10\%$ would substantially improve our understanding of flavor dynamics at TeV scale



Credits

A.J. Buras, A. Höcker, M. Gorbahn & U. Nierste

