

*New Frontiers in Subnuclear Physics*  
*September 12-17 2005, Milano*

# Test of CMS muon drift tube chambers with cosmic rays

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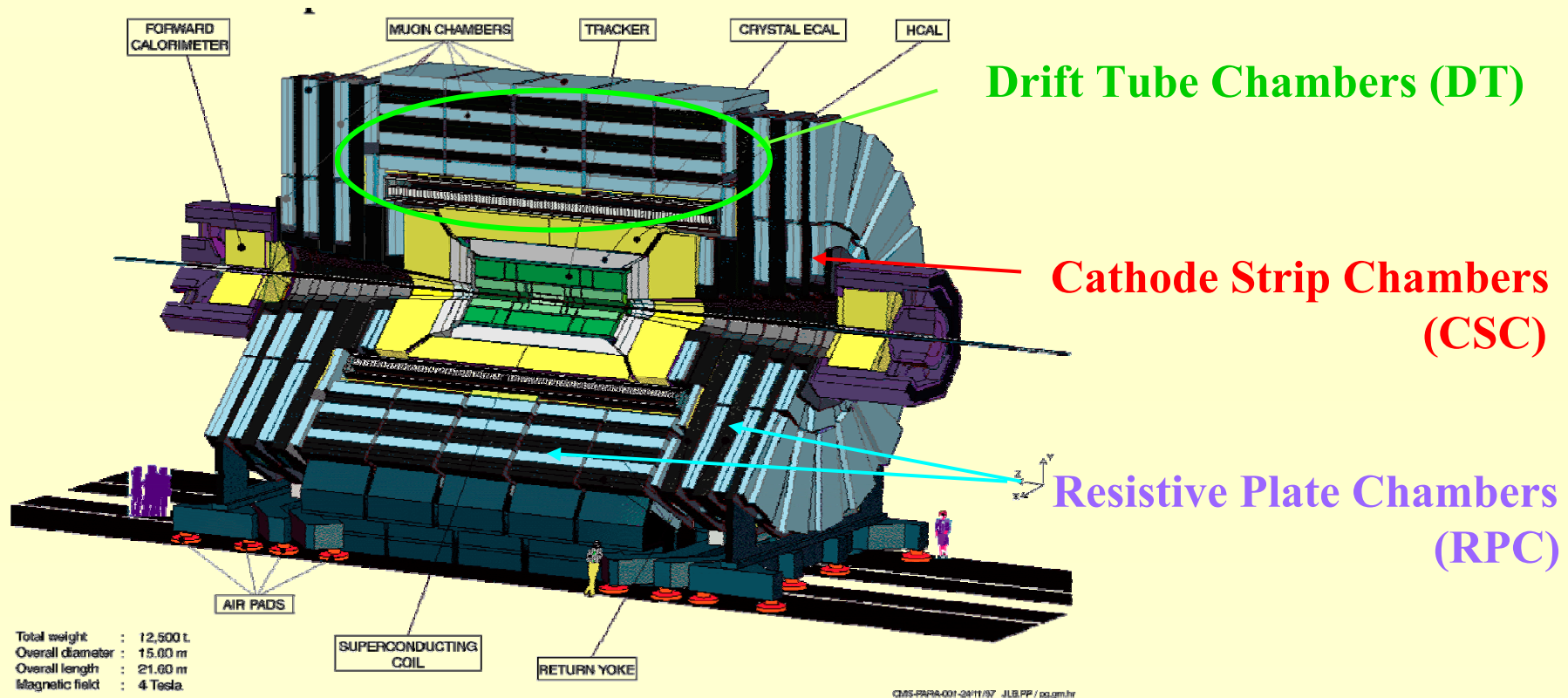
- Layout and properties of the detector;
- Quality control tests at the assembly sites;
- Measurements with cosmic muons and comparison with test beam results.



# Muon detectors in CMS



Robustness and redundancy ensured by 3 different muon sub-detectors:



4 layers of chambers. At least 3 track segments for a  $\mu$  track.

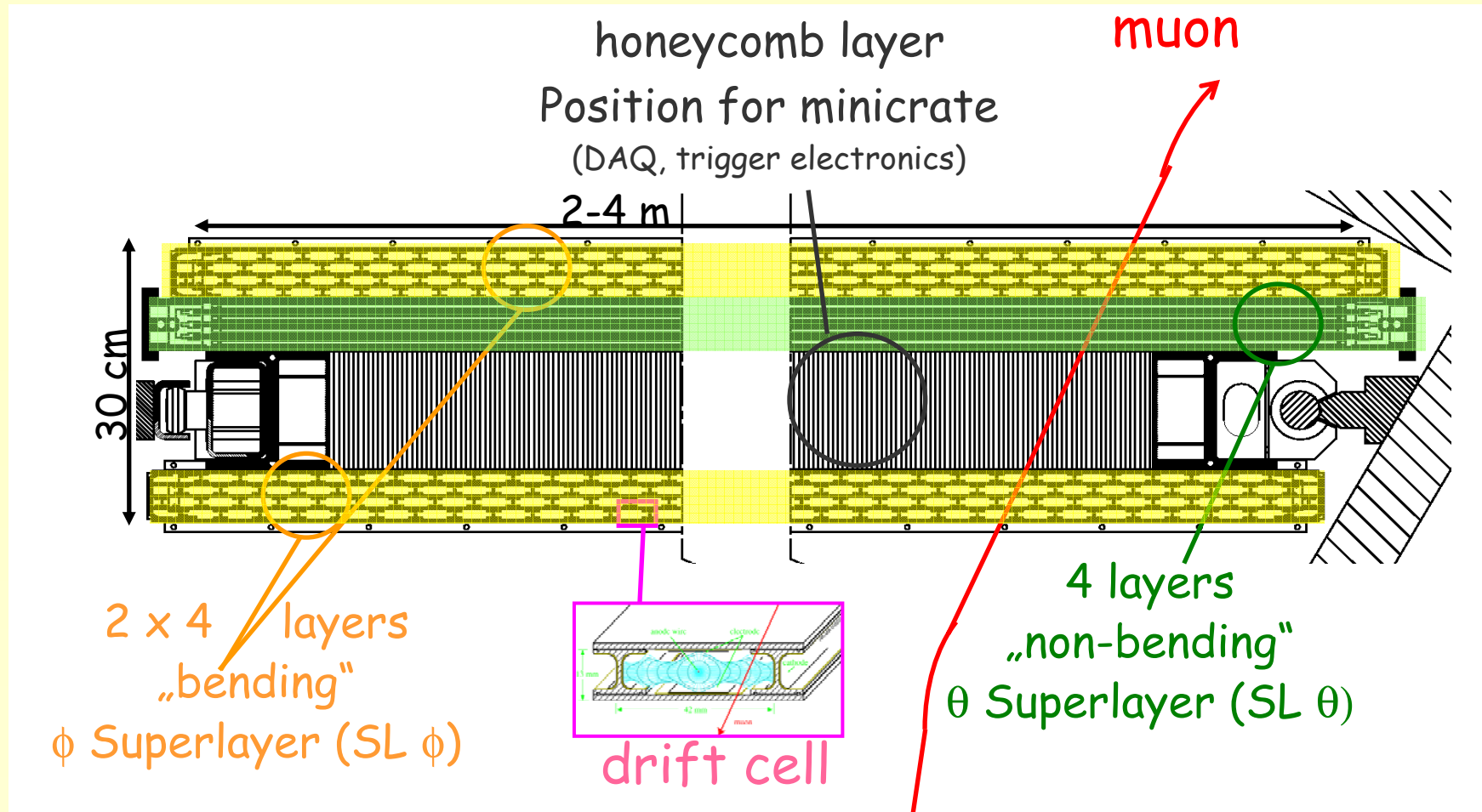




# Chamber structure

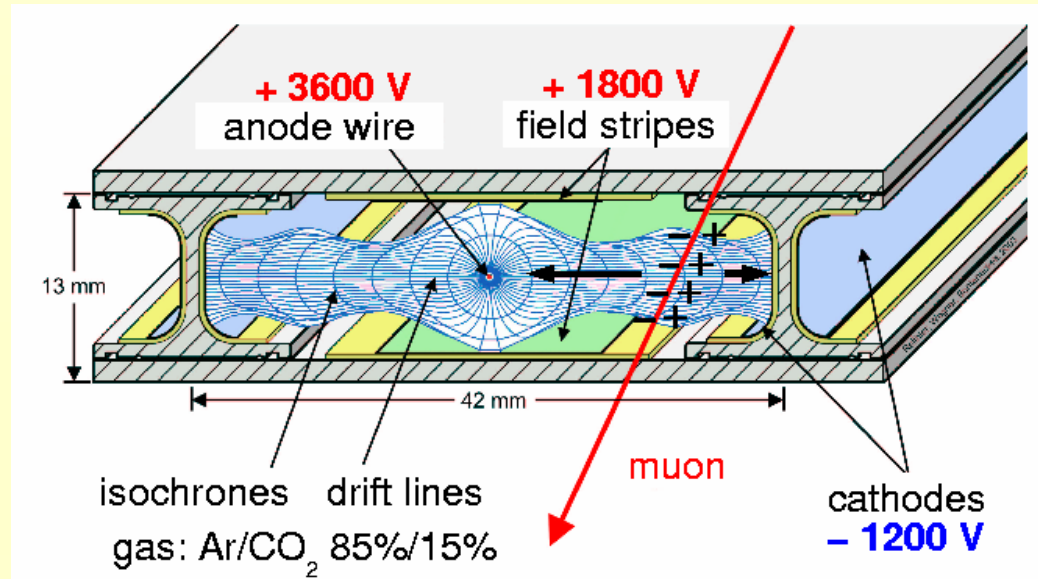


Each chamber is made of 3 or 2 (MB4) independent structures (**SuperLayers**)  
Each SL consists of 4 layers of drift tube cells, staggered by half a cell.





# Drift Cell structure



**Gas mixture:**  
**Ar(85%)+CO<sub>2</sub>(15%)**  
at atmospheric pressure

**Gain = 3÷5 10<sup>5</sup>**

## **Requirements:**

**Chamber spatial resolution ~ 100 μm**

**Single cell spatial resolution ~ 250 μm**

## **Trigger requirement:**

**disalignments between layers below few hundreds μm**



# DT chamber performance



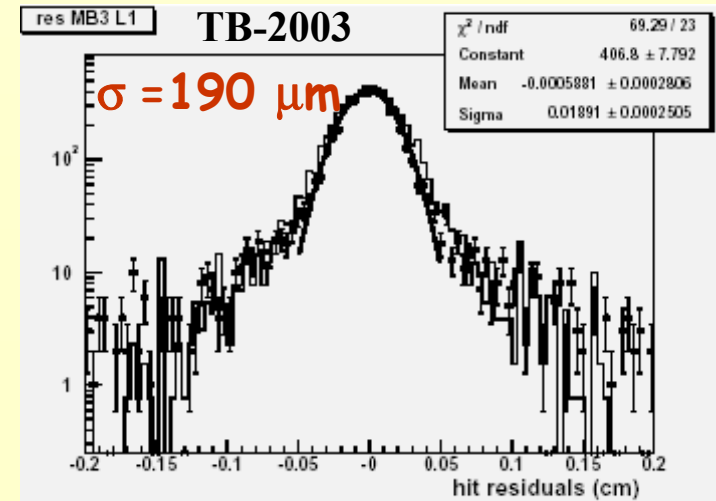
## Chamber performance from tests on beam

Single cell resolution  $< 200 \mu\text{m}$

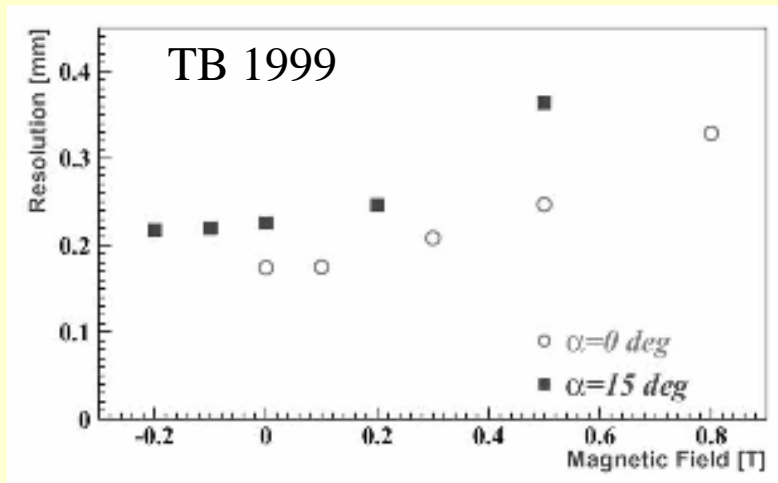
Cell efficiency  $> 99 \%$

Resolution and efficiency not affected by radiation noise

Deviations from linearity within  $100 \mu\text{m}$  in most of the drift length

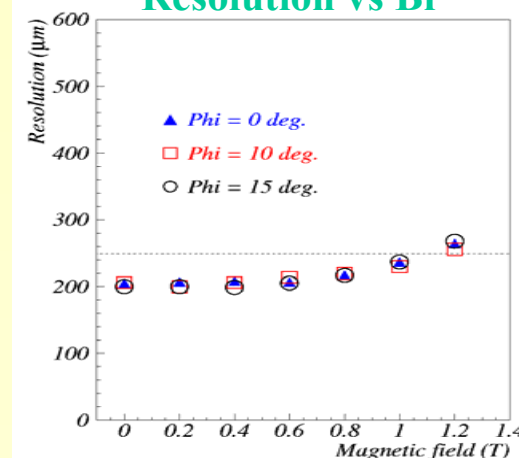


## $B_z$ – component along the $\phi$ wire

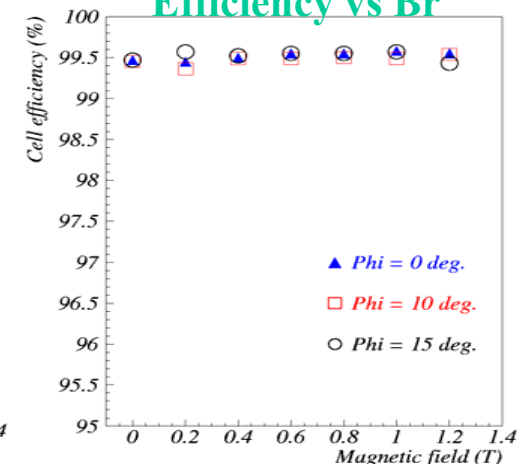


## $B_r$ – component orthogonal to the chamber

### Resolution vs $B_r$



### Efficiency vs $B_r$





# Quality control tests



**4 assembly sites:**

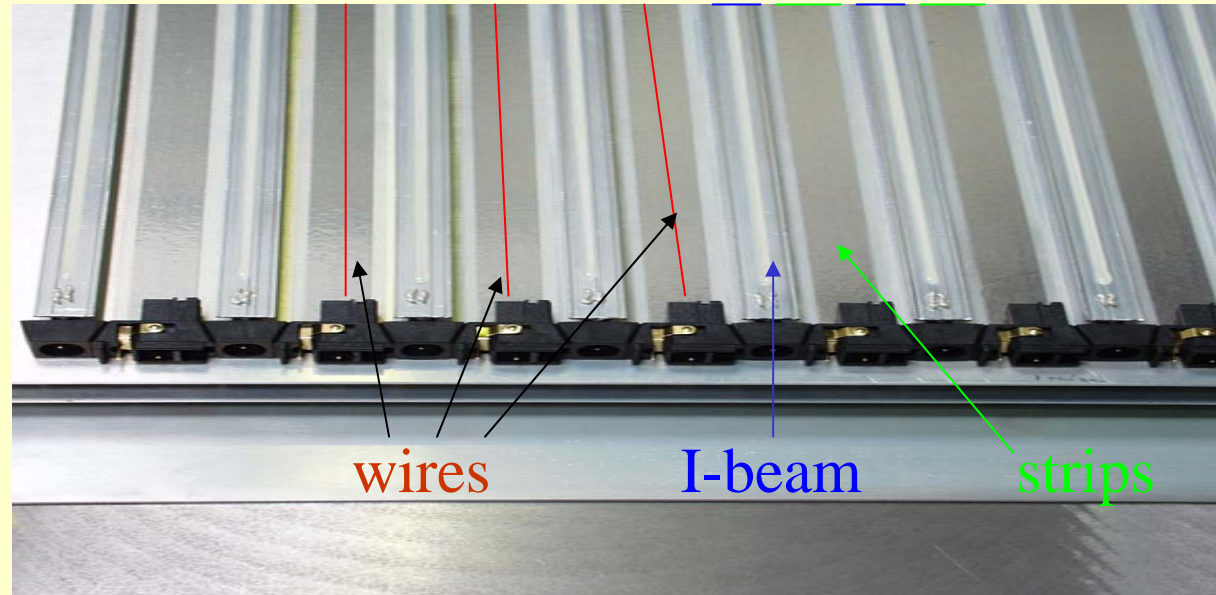
**Aachen** (MB1) 70+4+1

**Madrid** (MB2) 70+4+1

**Padova** (MB3) 70+4+1

**Torino** (MB4) 40+4

◆ The chamber production will be completed by April 06



**Quality control tests:**

**Assembly phase:**

- HV tests of strips and cathods
- Measurement of the wire tension
- Wire and cathode positions with CCD

**Tests on the assembled SL:**

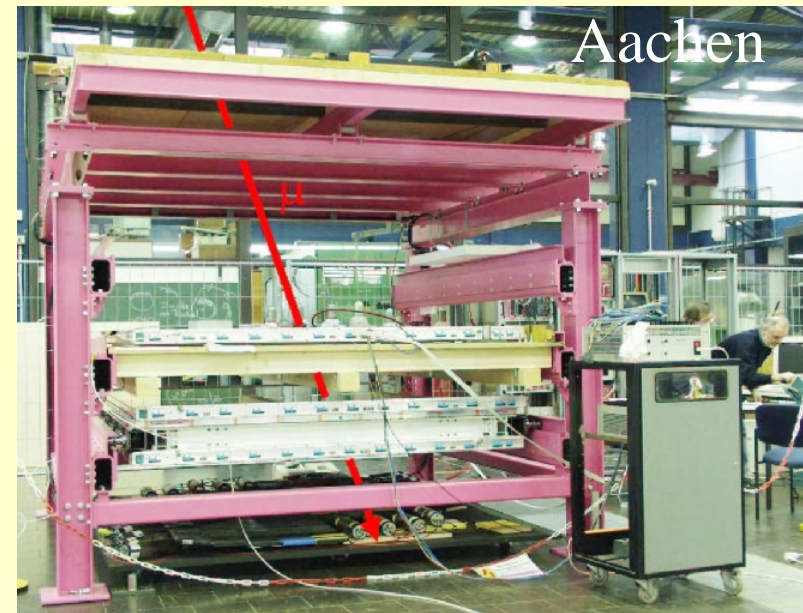
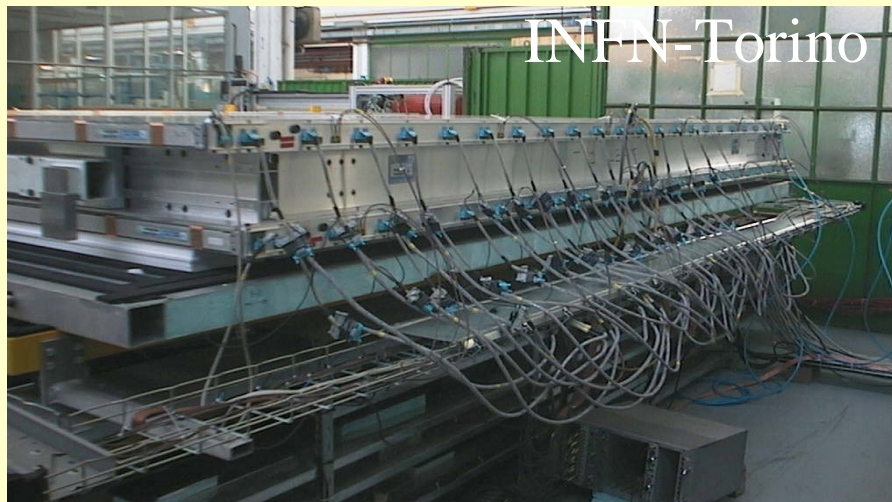
- Planarity
- Gas tightness
- HV behaviour
- noise



# Tests with cosmics



Measurements with cosmic muons are performed in each production site to certificate the chamber performance.



Cosmic-ray events are triggered by an external scintillator system.  
Drift times are measured with external TDCs and custom DAQ.  
Typical trigger rates  $50\div 100$  Hz  $\rightarrow 10^6$  events in few hours.



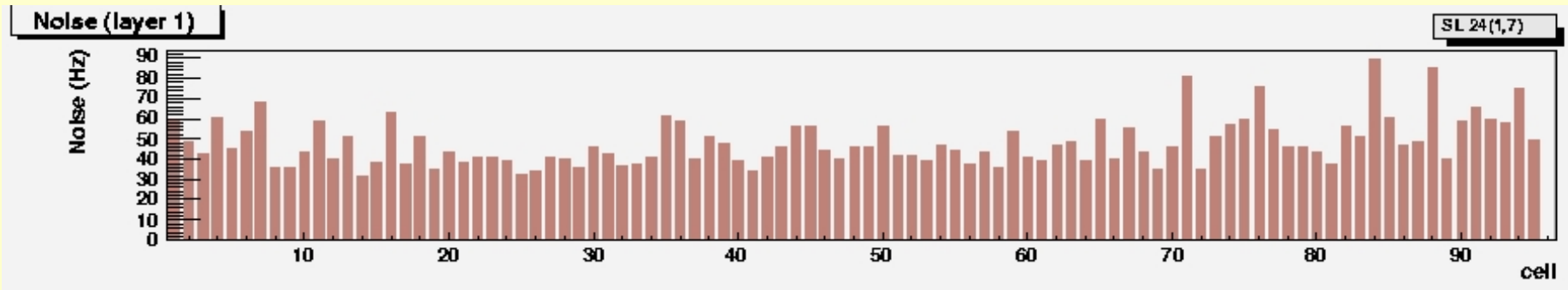


# Hit rate

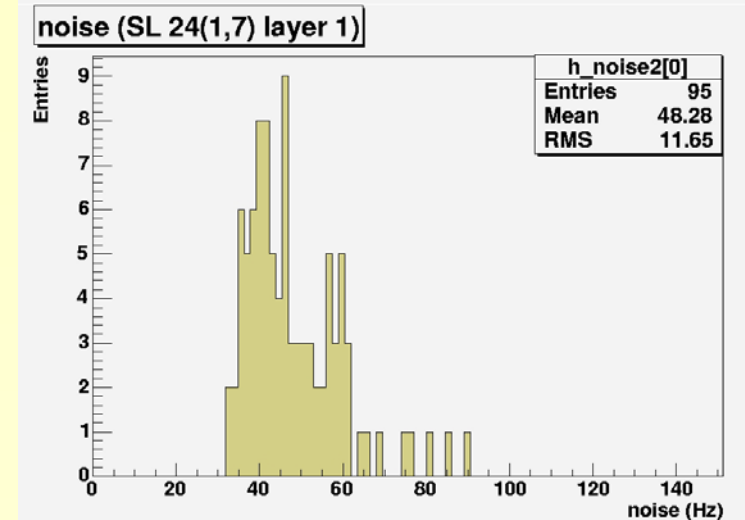


Measured with TDC (random triggers) or scalers.

Typical hit rate distribution in one layer:



High noise levels in one or more channel could indicate FE problems, wrong HV behaviour, problems in the capacitive coupling, etc.

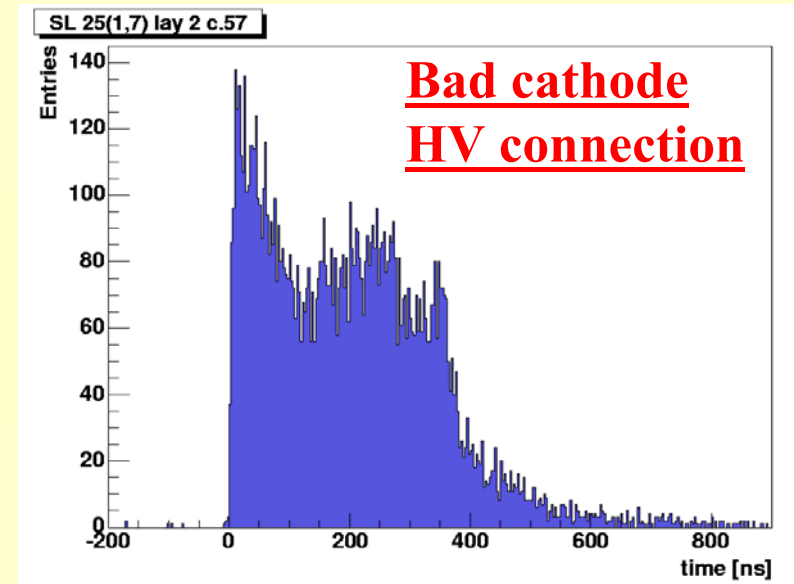
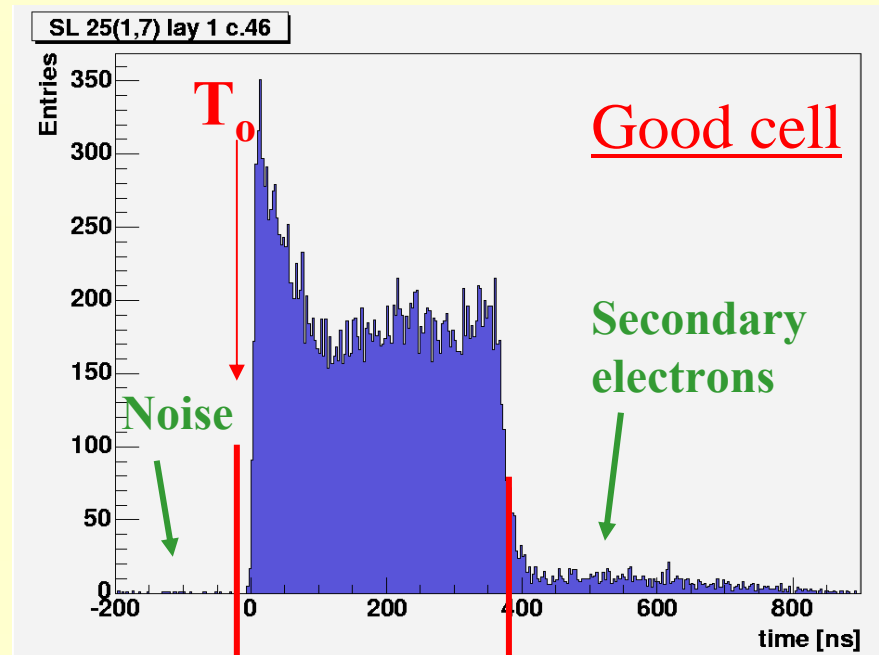




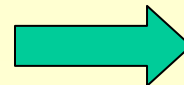
# Drift time distribution



Unconnected cathodes can be easily identified from the drift time distribution.



Max.drift time  $\sim 380$  ns

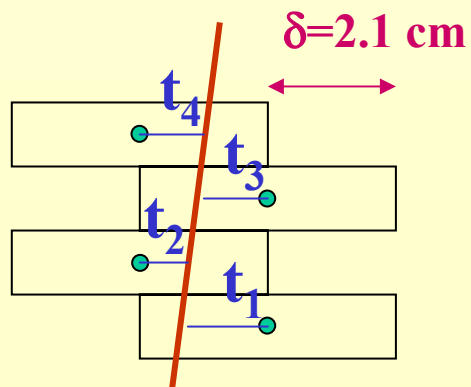


$v_{\text{drift}} \sim 55$   $\mu\text{m/ns}$

Fraction of disconnected cells (mainly for HV problems):  $\sim 0,08$  %



# Mean Time

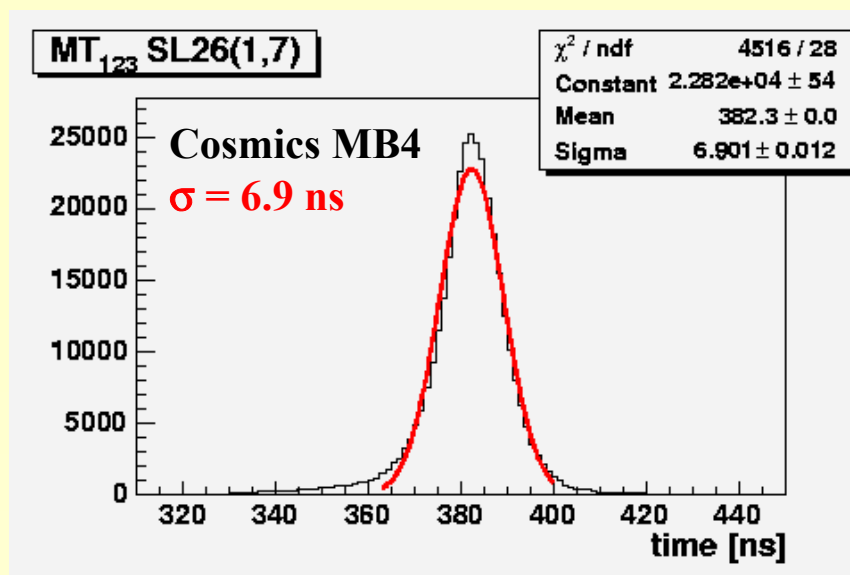
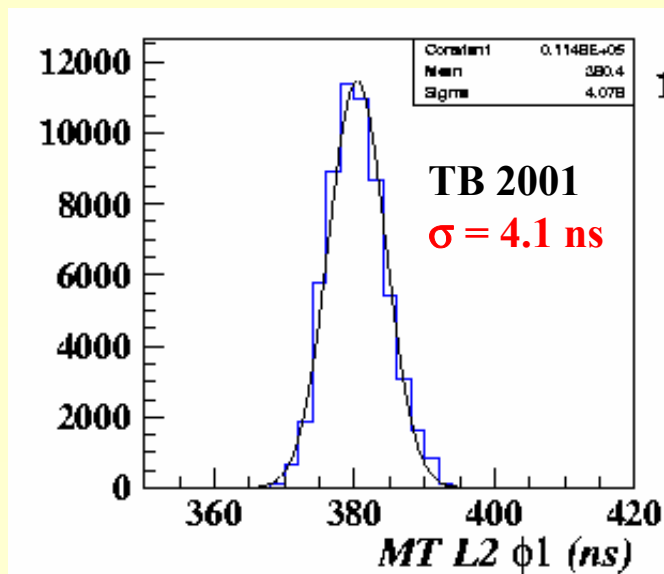


$$MT_{123} = \frac{t_1 + t_3}{2} + t_2$$

$$MT_{234} = \frac{t_2 + t_4}{2} + t_3$$

$$MT_{123} = MT_{234} = T_{\max} = \frac{\delta}{v_d}$$

$$\sigma_{MT} = \sqrt{\frac{3}{2}} \cdot \frac{\sigma_x}{v_d}$$



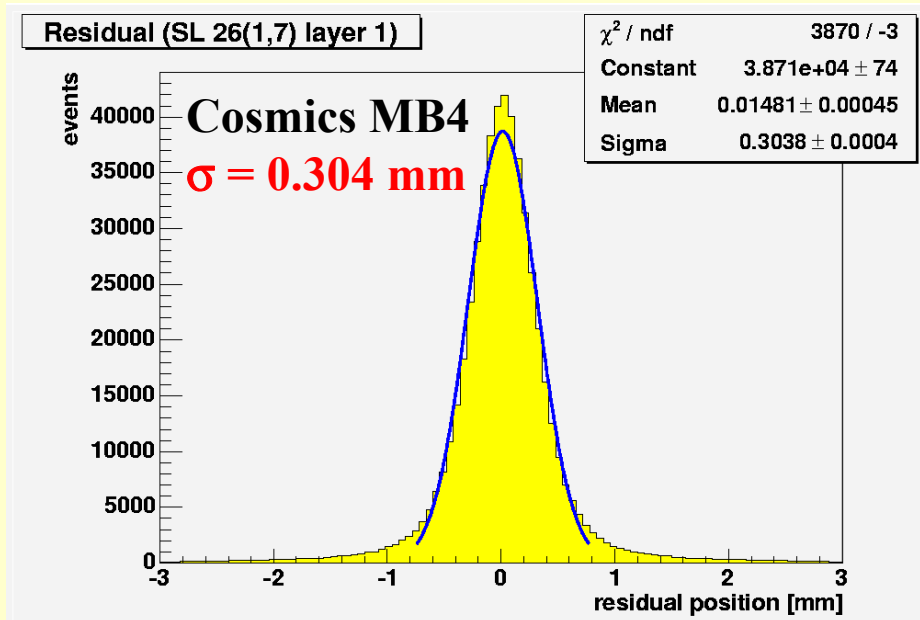
$\sigma_{MT} \sim 7.0 \text{ ns}$



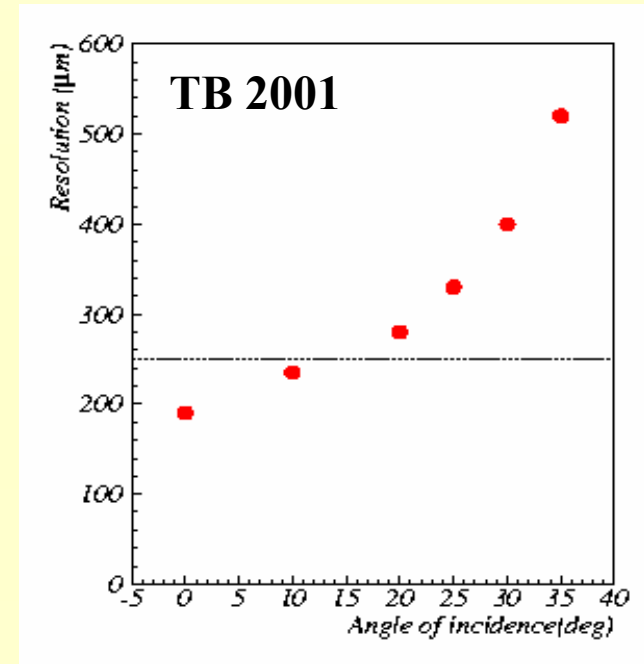
$\sigma_x \sim 310 \mu\text{m}$



# Cell resolution



## Resolution vs angle of incidence



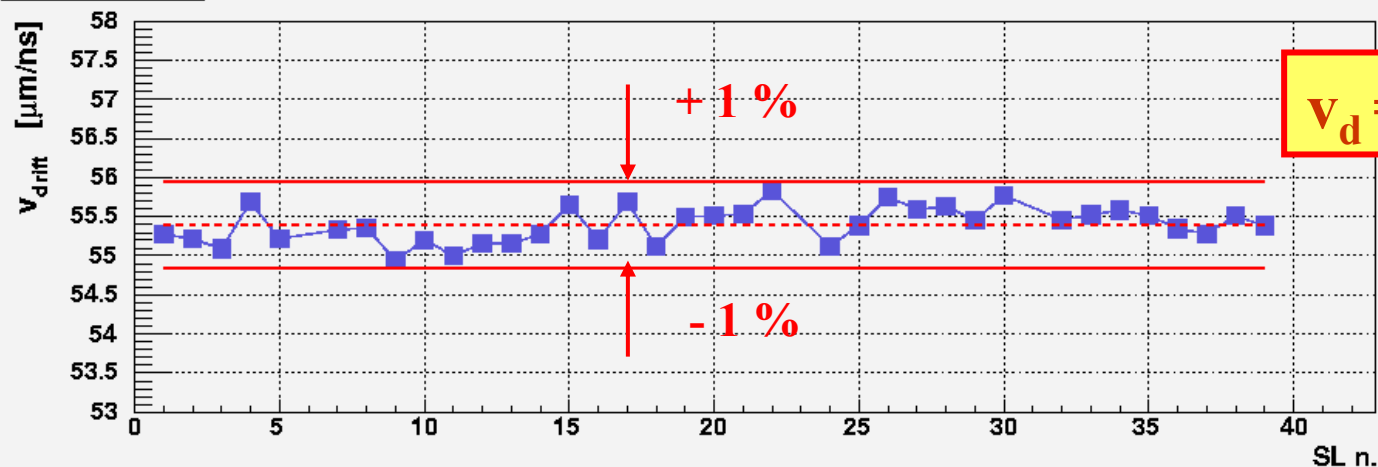
The cell resolution measured with cosmics is higher than that measured from test beam data (wider angular distributions, multiple scattering from low energy muons, trigger timing resolution)



# Calibration stability



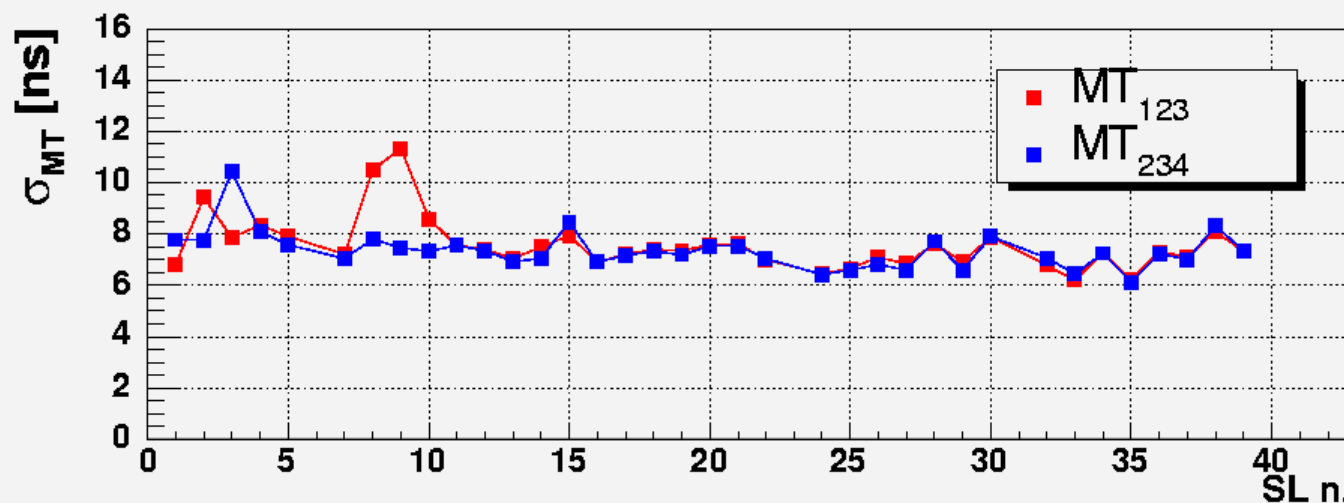
cosmics MB4



$v_d = 55.4 \mu\text{m/ns}$

stable within  
 $\pm 1 \%$

cosmics MB4



High  $\sigma_{MT}$  in  
the first MB4  
SLs !

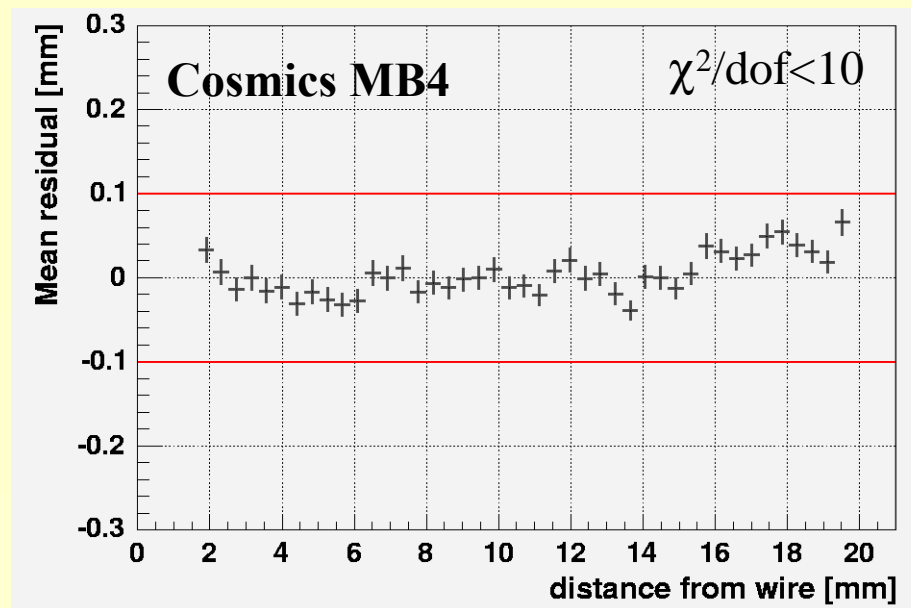
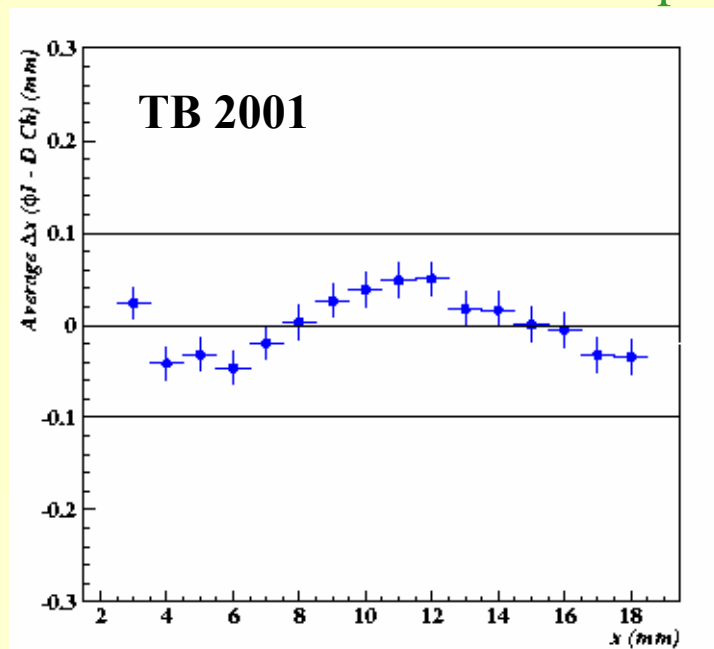


# Linearity



SL local linear fit assuming linear drift parameterisation.  
2D segment reconstruction.

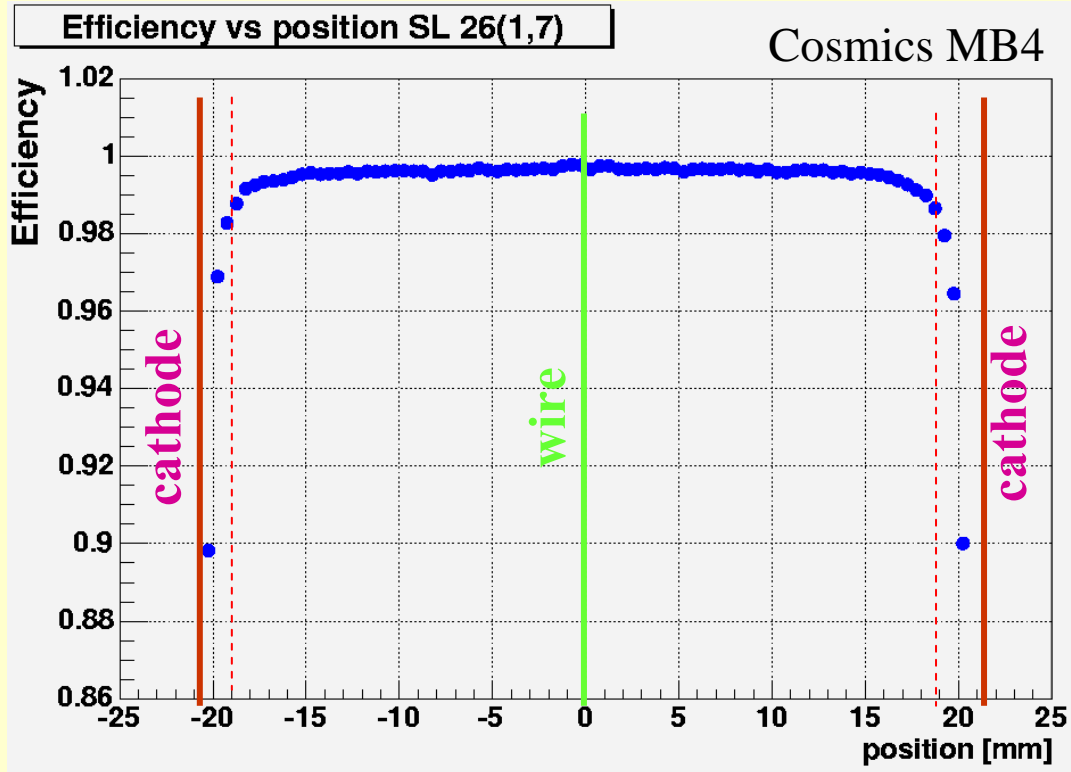
Residual vs position inside the cell



Deviations from linearity  $< 100 \mu\text{m}$  excluding few mm close to the wire and the cathode



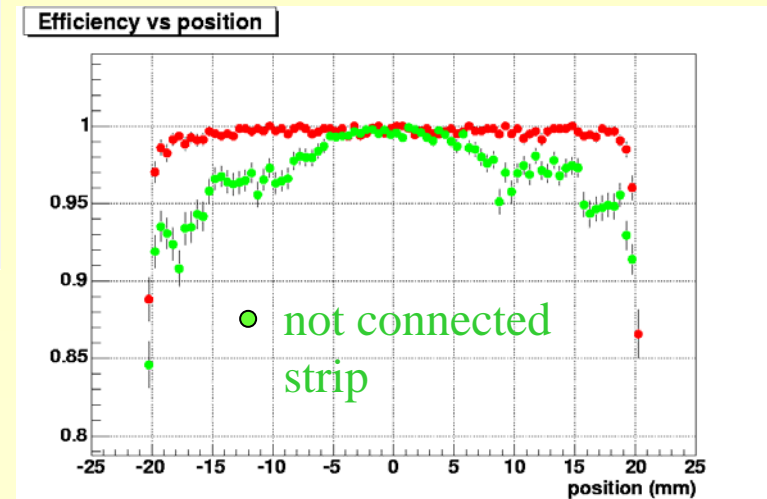
# Cell detection efficiency



$$\epsilon_{\text{cell}}(\mathbf{x}) = \frac{N_{\text{cell-found}}}{N_{\text{cell-expected}}(\mathbf{x})}$$

Detection efficiency in the active region > 99 %

Lower efficiency for cells with bad HV connections

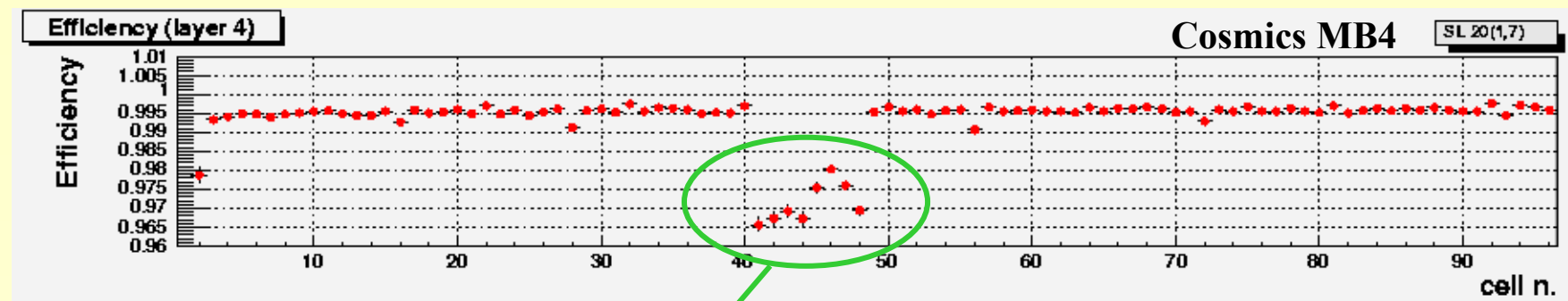
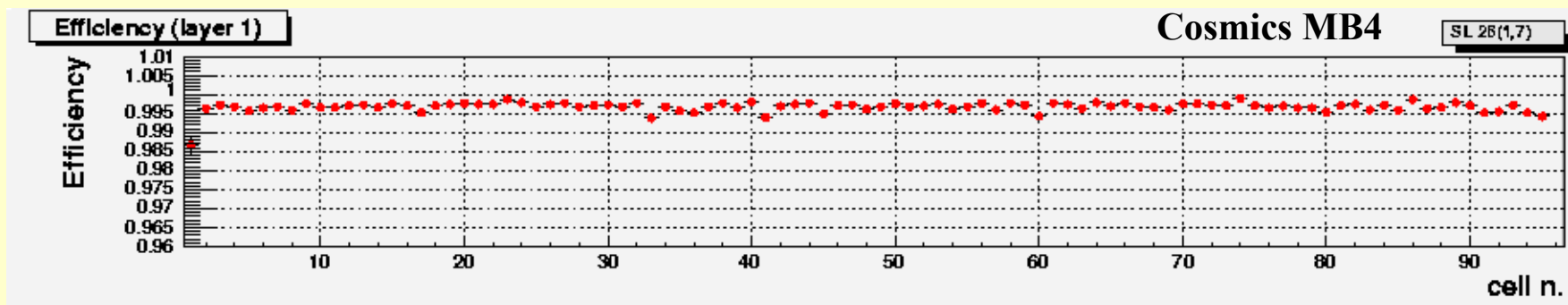




# Cell detection efficiency



Mean efficiency ~ 99,5 %

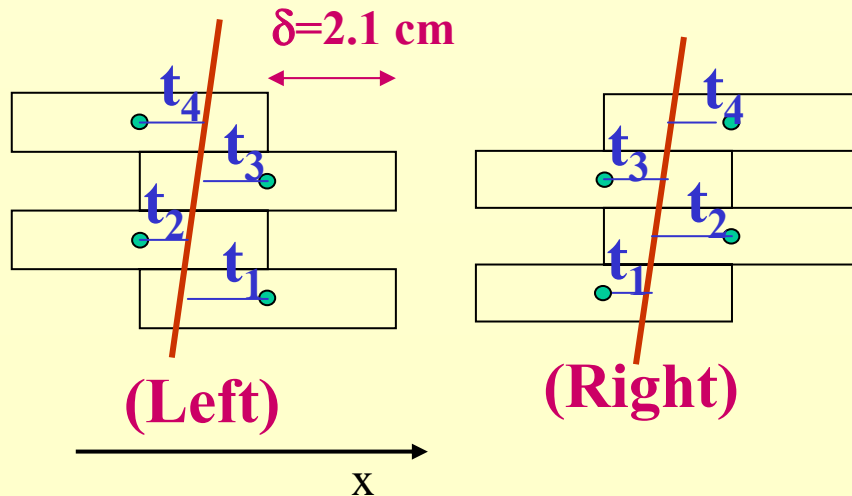


Cells with not connected strips





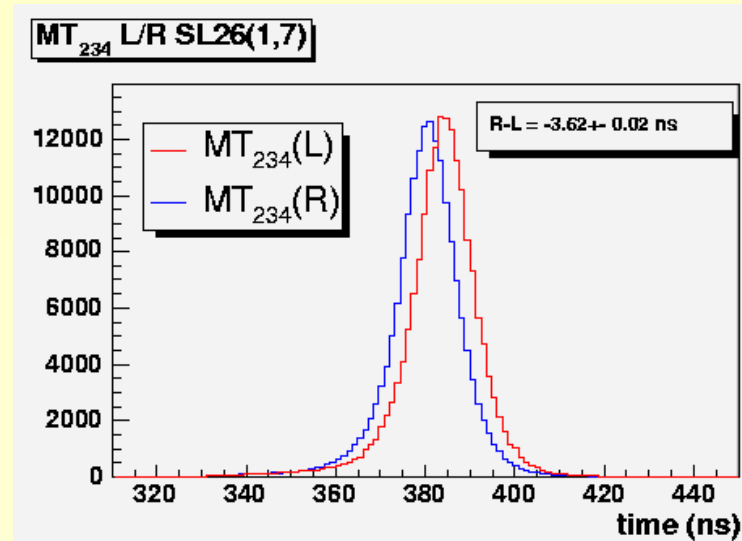
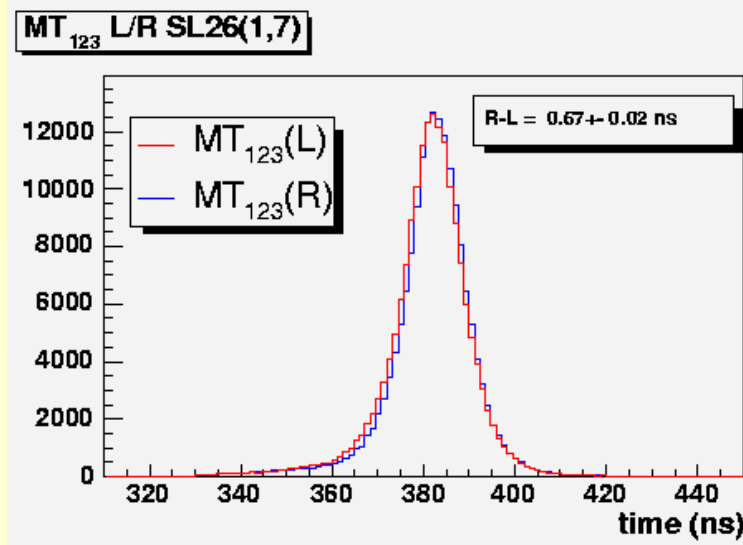
# Layer relative alignment



$\Delta x_i$  = shift of the layer position with respect to the nominal position ( $\Delta x_1 = 0$ )

$$MT_{123}(R) - MT_{123}(L) = \frac{1}{v_d} \cdot (-\Delta x_3 + 2 \cdot \Delta x_2)$$

$$MT_{234}(R) - MT_{234}(L) = \frac{1}{v_d} \cdot (\Delta x_2 - 2 \cdot \Delta x_3 + \Delta x_4)$$

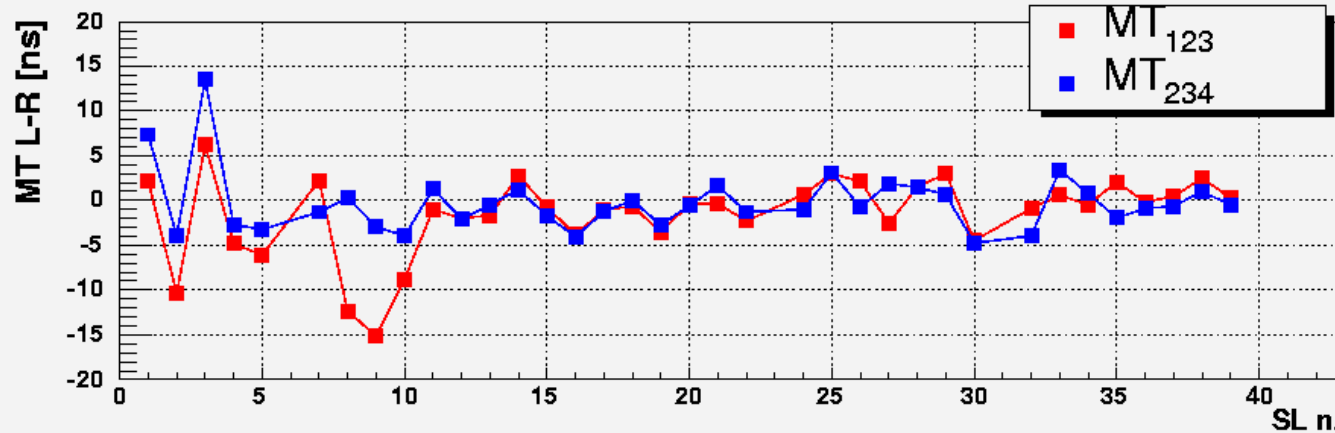




# Layer relative alignment



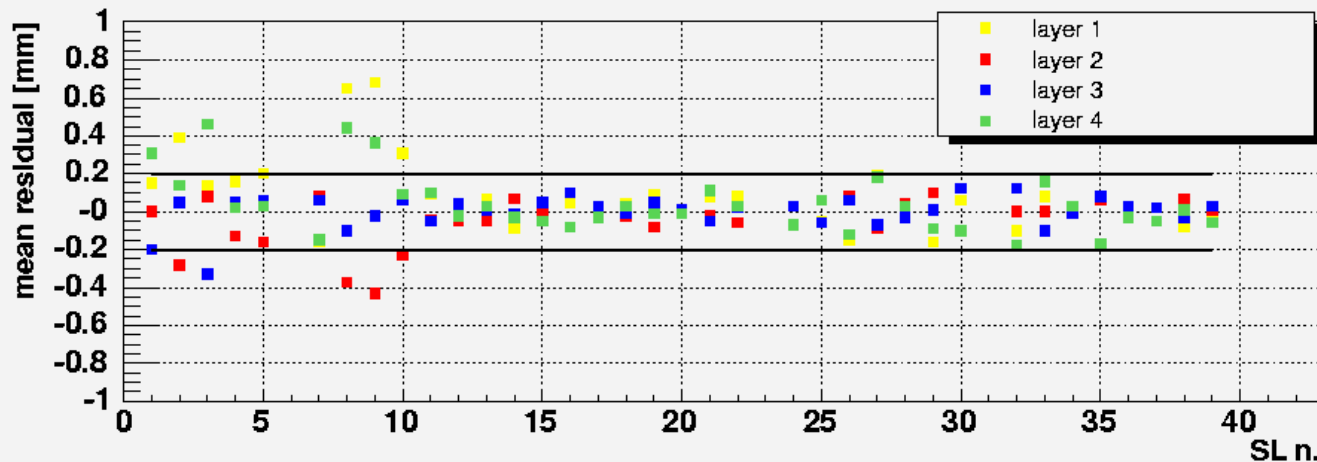
cosmlcs MB4



Indication of relatively large disalignment effects for SL n.  $\leq 10$

...

cosmlcs MB4



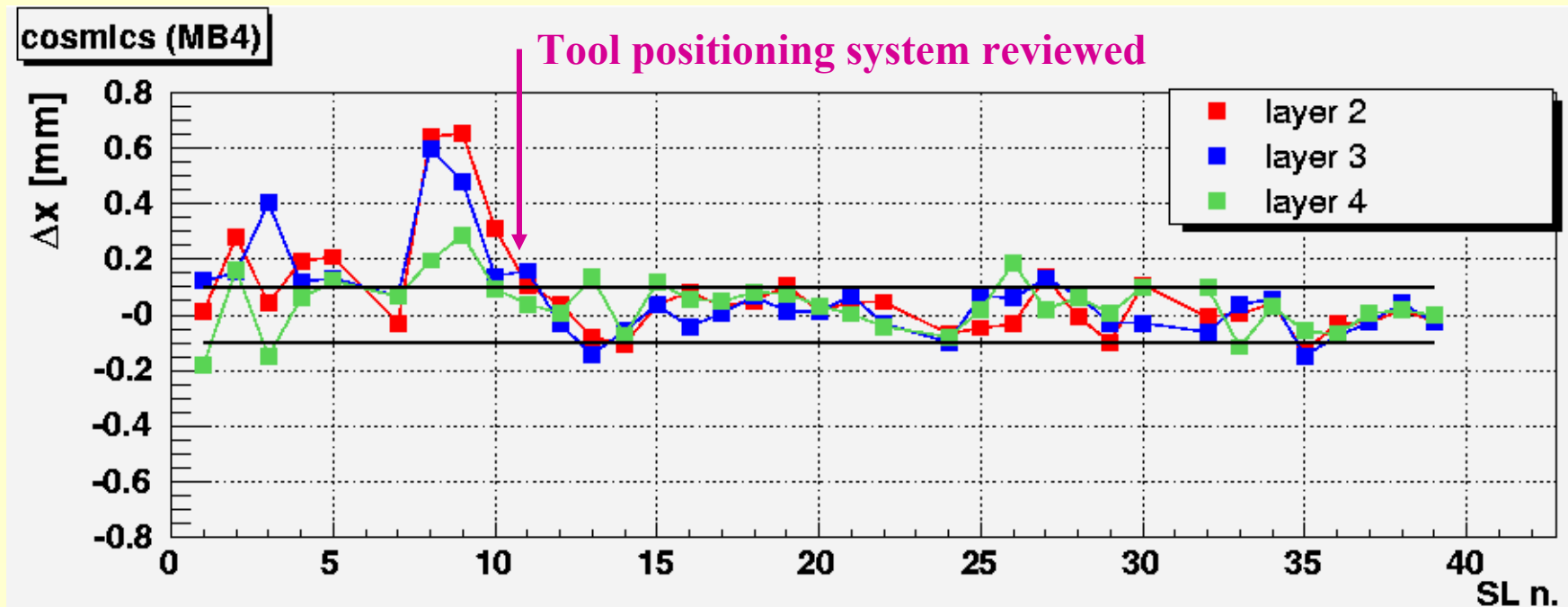
... confirmed by the mean layer residuals



# Layer relative alignment



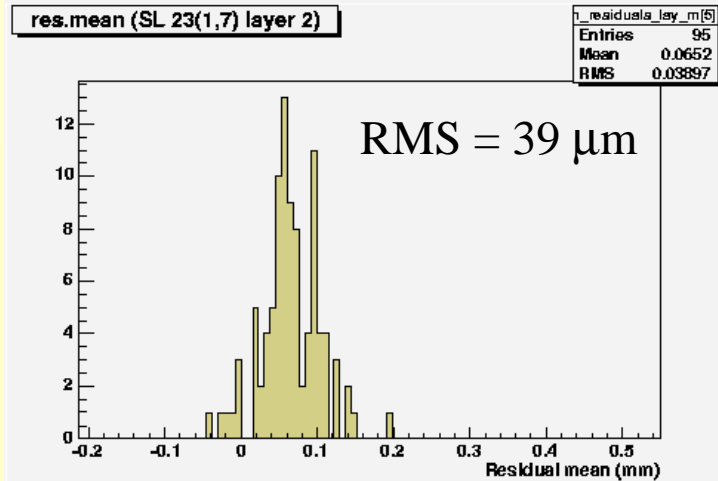
The plane disalignments  $\Delta x_i$  are determined assuming a cosmic angular distribution symmetric with respect to the vertical direction.



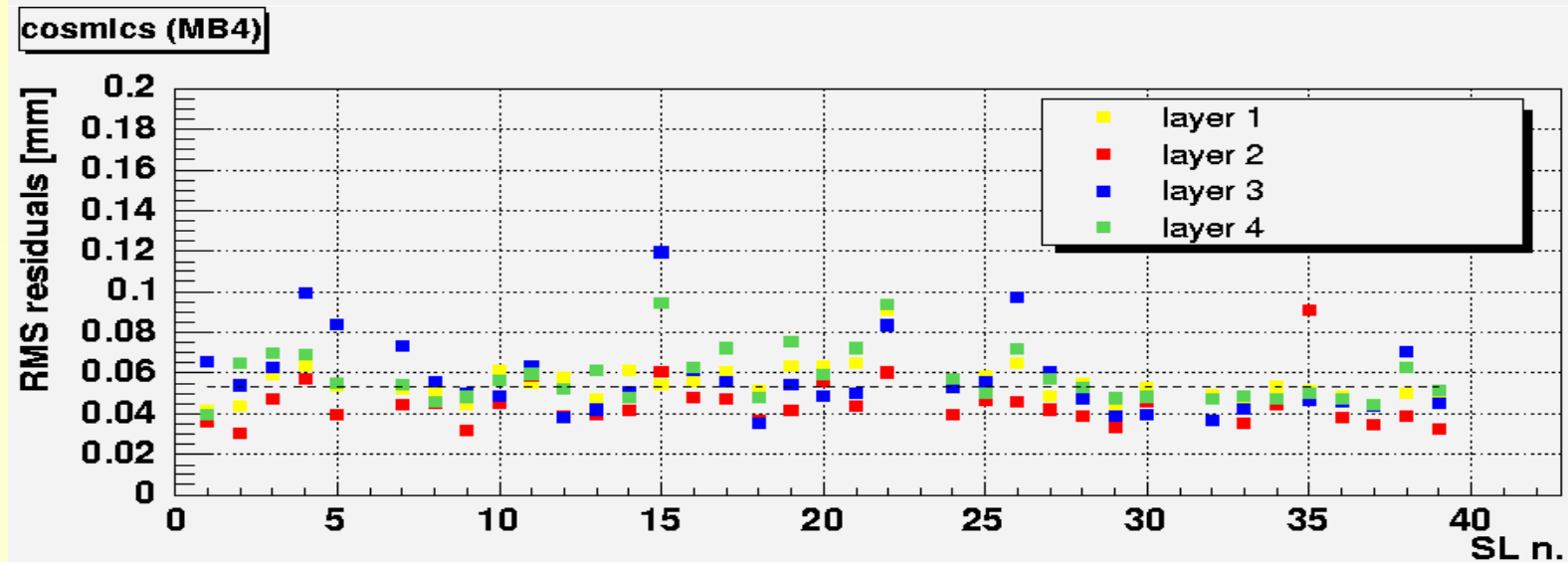
Better alignment in other production sites (smaller chambers)



# Wire positions



The RMS of the mean residual distribution for the cells in one layer is **< 100  $\mu\text{m}$**





# Conclusions



Data with cosmics provide a fast and efficient way to analyse the chamber performance on a cell-by-cell basis and to detect potential problems.

**All Drift Tube chambers are certified with cosmics:**

- **at the assembling sites**
- **at CERN before the installation**
- **on the CMS wheels after the installation**

At the moment, data taking with cosmic muons is the only way to commission and calibrate the detector.

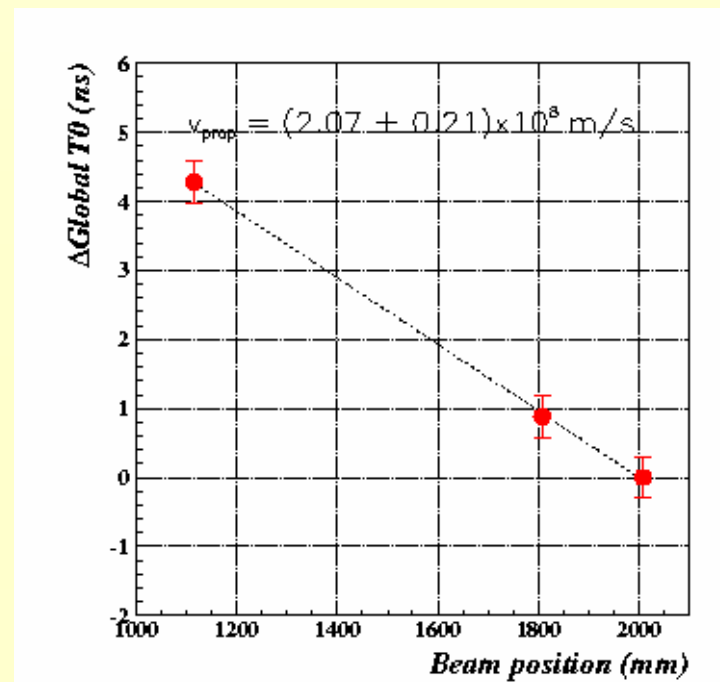
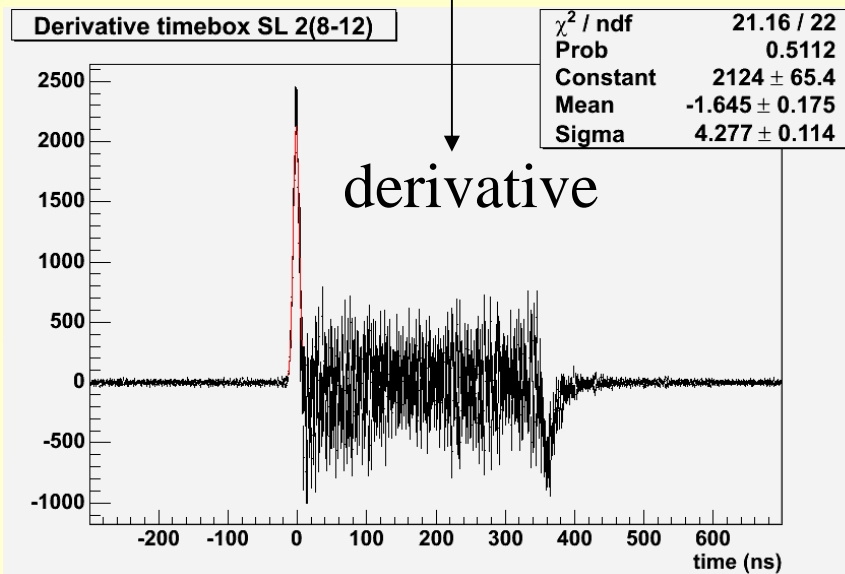
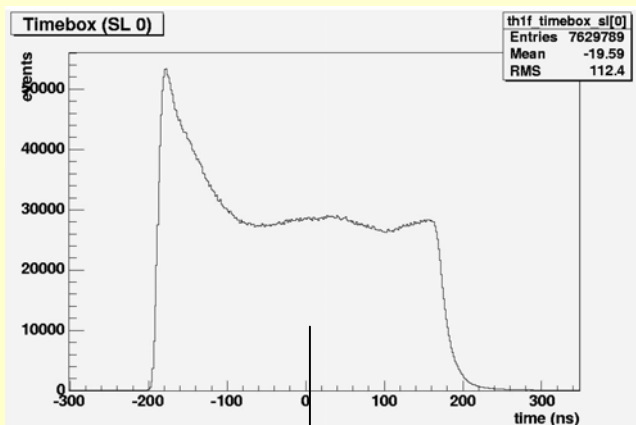


In 2006 a crucial combined test of all the CMS sub-detectors is foreseen.

Data collection with cosmics will be used for validation of the trigger/DAQ system and detector alignment studies.



# Timing calibration



Corrections due to the propagation time along the wire taken into account.

Wire-to-wire offsets measured with pulses

The peak of the derivative defines the **time** of particles crossing the wire.



# Construction phases (II)

