

From HERA to the LHC

John Ellis,

DESY,

March 23rd, 2005

Preview

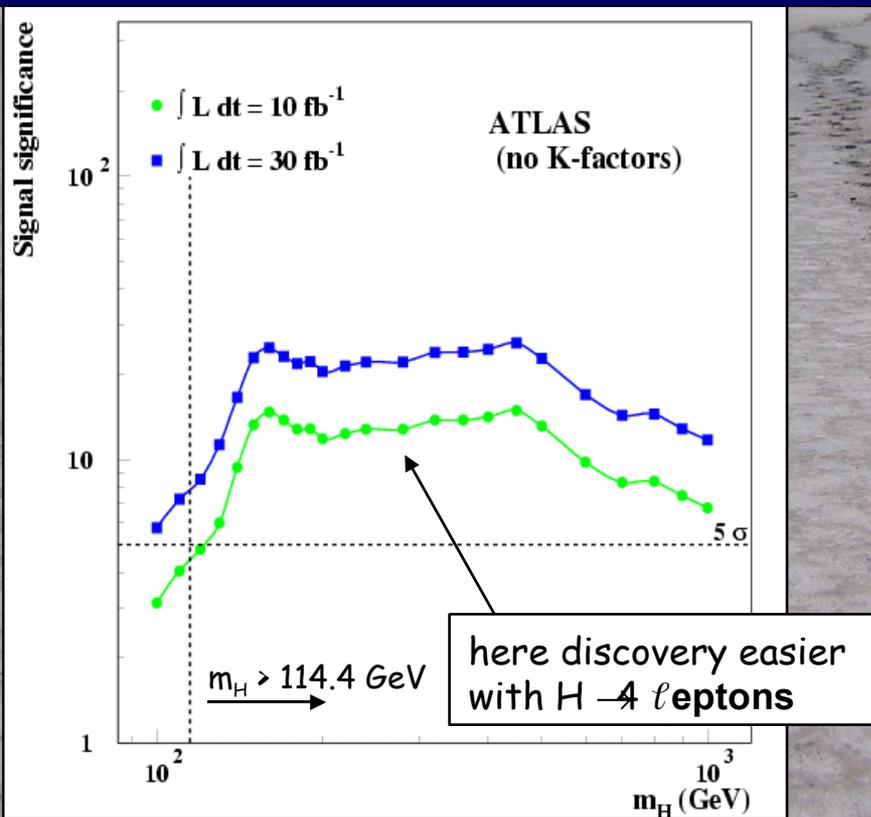
- LHC's 'core business': Higgs & Susy
- Importance of understanding QCD
- Parton saturation, RHIC and the CGC
- UHECRs & Forward physics @ LHC
- Diffractive Higgs production @ LHC



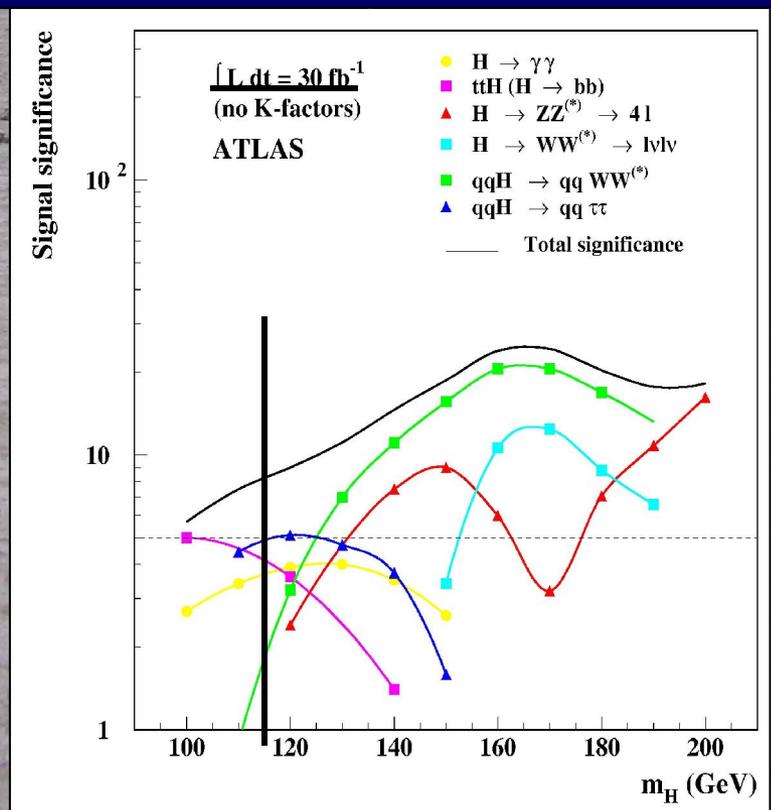
Prospects in Higgs Physics

Higgs Detection at the LHC

The Higgs may be found quite quickly ...



... in several different channels

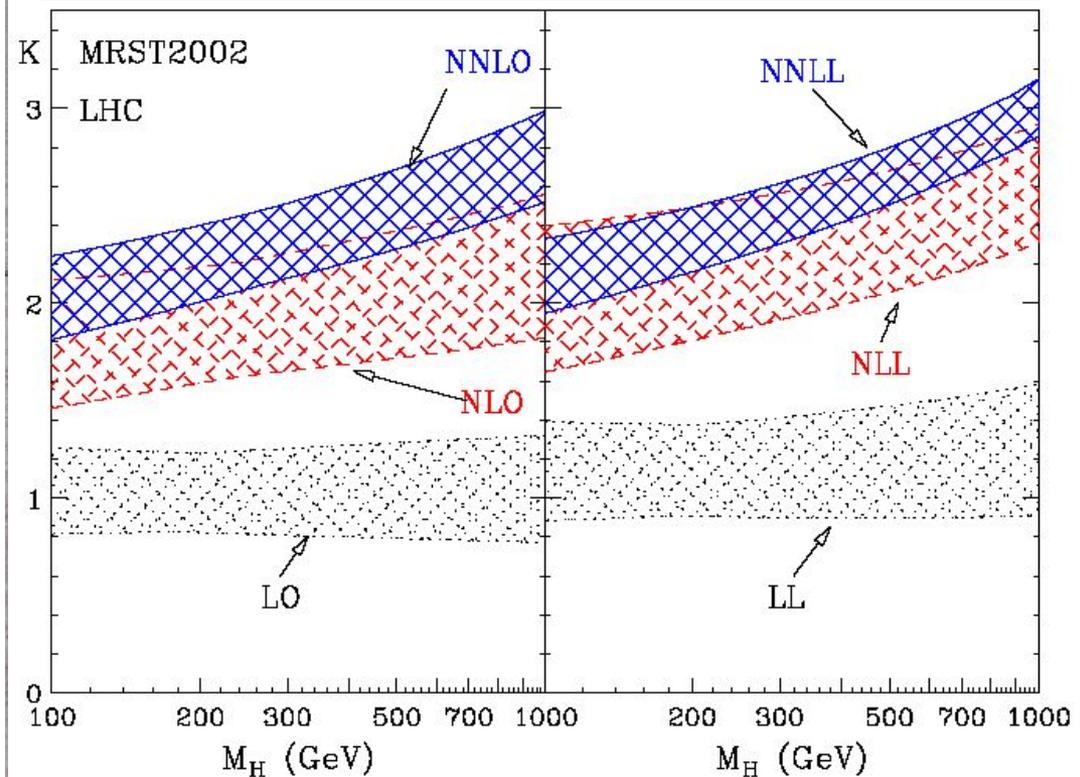
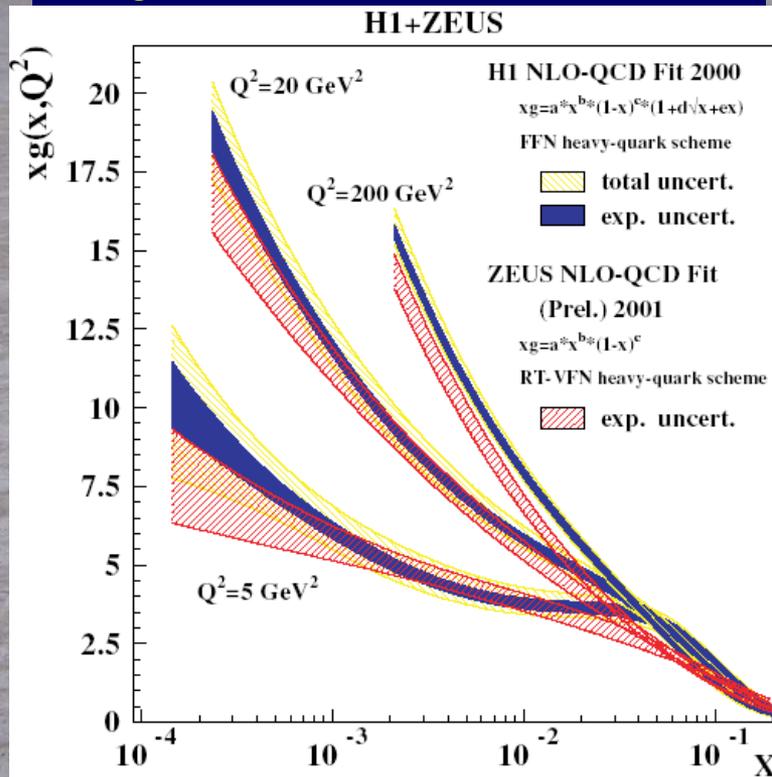


How Accurately can the Higgs Cross Section be Calculated?

Glover

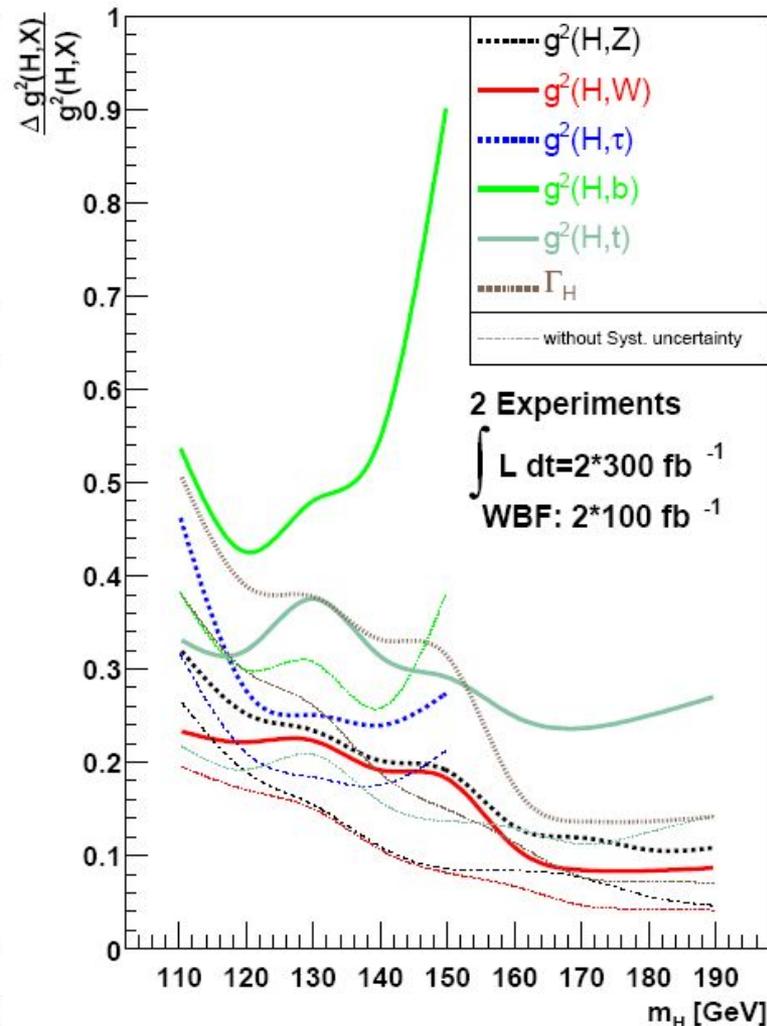
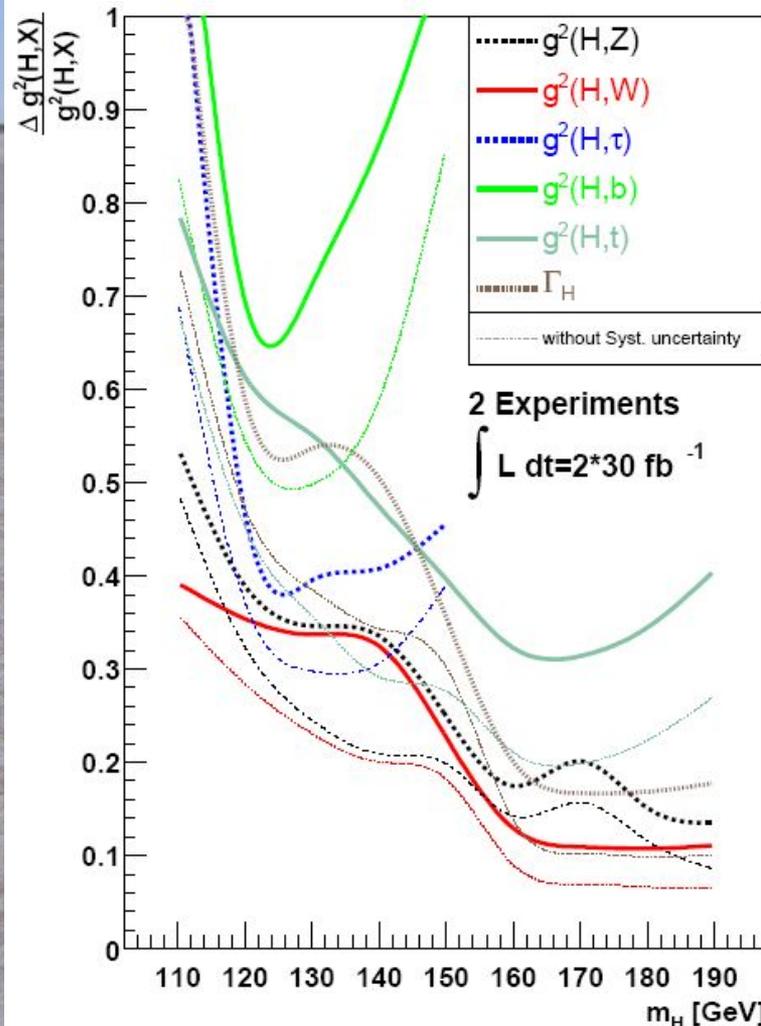
Need gluon distribution @
 $x_g \sim 10^{-2}$, $Q^2 \sim 10^4 \text{ GeV}^2$

Resumming next-to-next-to-leading
 logarithms: 10% uncertainty



Catani + de Florian + Grazzini + Nason

Accuracy in LHC Determinations of Higgs Couplings



Theorists getting Cold Feet

- Composite Higgs model?
conflicts with precision electroweak data
- Interpretation of EW data?
consistency of measurements? Discard some?
- Higgs + higher-dimensional operators?
corridors to higher Higgs masses?
- Little Higgs models?
extra 'Top', gauge bosons, 'Higgses'
- Higgsless models?
strong WW scattering, extra D?

Little Higgs Models

- Embed SM in larger gauge group
- Higgs as pseudo-Goldstone boson
- Cancel top loop

$$\delta m_{H,top}^2(SM) \sim (115\text{GeV})^2 \left(\frac{\Lambda}{400\text{GeV}}\right)^2$$

with new heavy T quark

$$m_T > 2\lambda_t f \sim 2f \quad f > 1 \text{ TeV}$$

$$\delta m_{H,top}^2(LH) \sim \frac{6G_F m_t^2}{\sqrt{2}\pi^2} m_T^2 \log \frac{\Lambda}{m_T} \gtrsim 1.2 f^2$$

- New gauge bosons, Higgses
- Higgs light, other new physics heavy

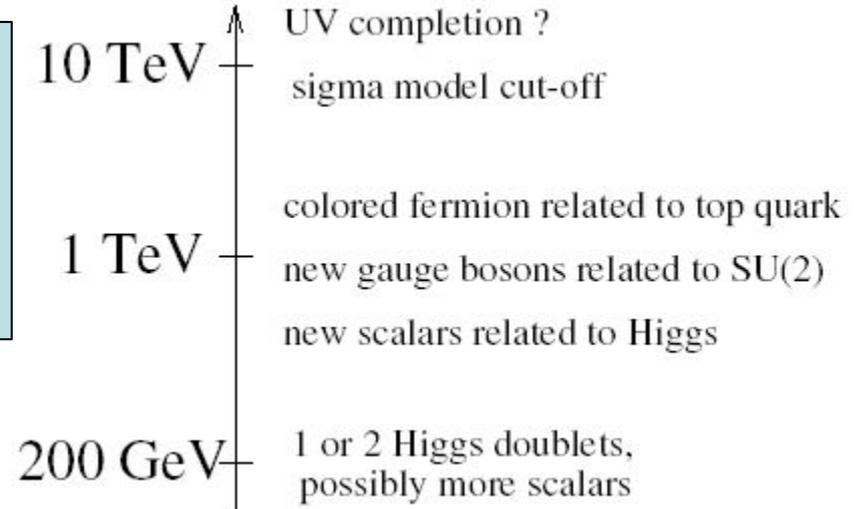
$$M_T < 2 \text{ TeV} \quad (m_h / 200 \text{ GeV})^2$$

$$M_{W'} < 6 \text{ TeV} \quad (m_h / 200 \text{ GeV})^2$$

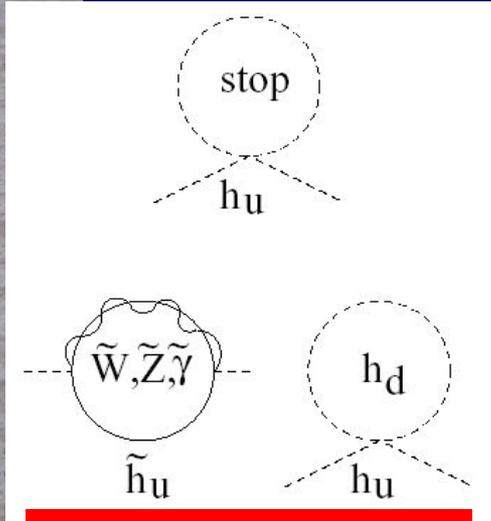
$$M_{H^{++}} < 10 \text{ TeV}$$

Not as complete as susy: more physics > 10 TeV

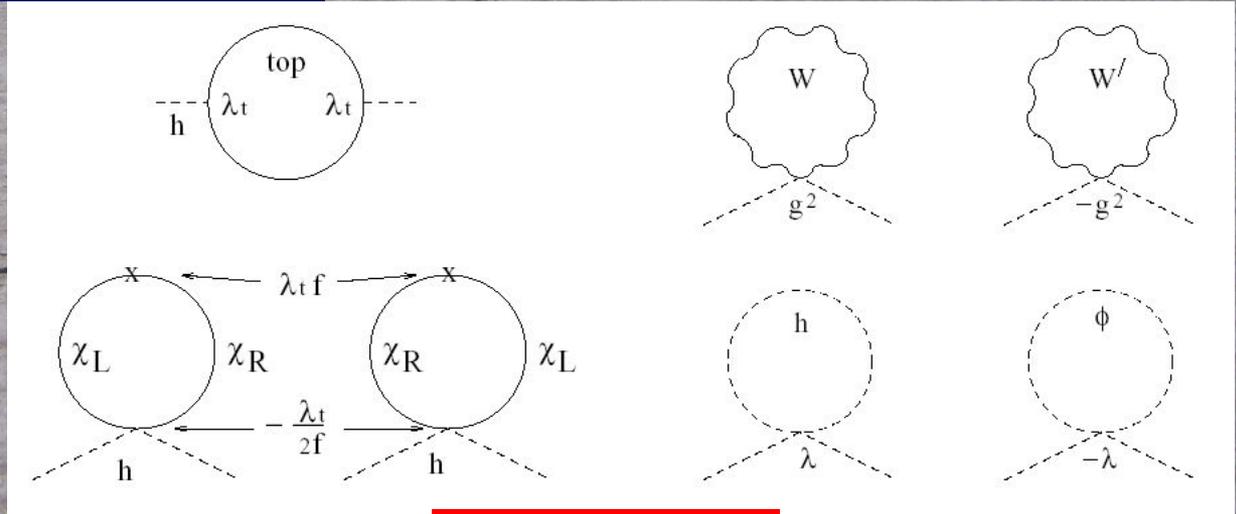
Generic Little Higgs Spectrum



Loop cancellation mechanisms

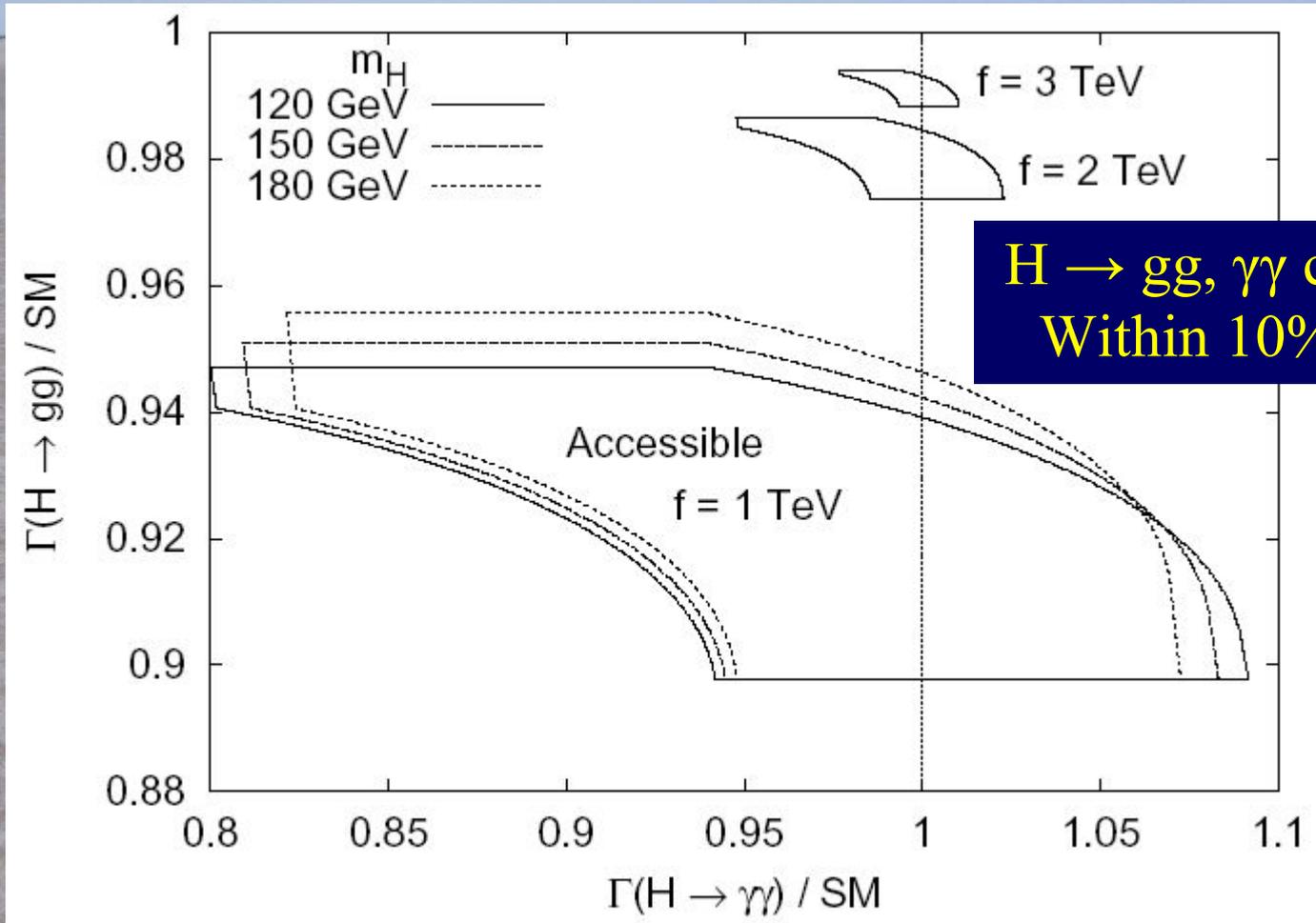


Supersymmetry



Little Higgs

Measuring Little Higgs Couplings



A wide, flat, sandy beach with a path of footprints leading towards the horizon under a clear blue sky. The beach is vast and flat, with a path of footprints leading from the foreground towards the horizon. The sky is a clear, pale blue. In the distance, there are some low mountains or hills. The overall scene is bright and open.

Supersymmetry

Why Supersymmetry (Susy)?

- Hierarchy problem: why is $m_W \ll m_P$?

($m_P \sim 10^{19}$ GeV is scale of gravity)

- Alternatively, why is

$$G_F = 1/m_W^2 \gg G_N = 1/m_P^2 ?$$

- Or, why is

$$V_{\text{Coulomb}} \gg V_{\text{Newton}} ? \quad e^2 \gg G m^2 = m^2 / m_P^2$$

- Set by hand? What about loop corrections?

$$\delta m_{H,W}^2 = O(\alpha/\pi) \Lambda^2$$

- Cancel boson loops fermions

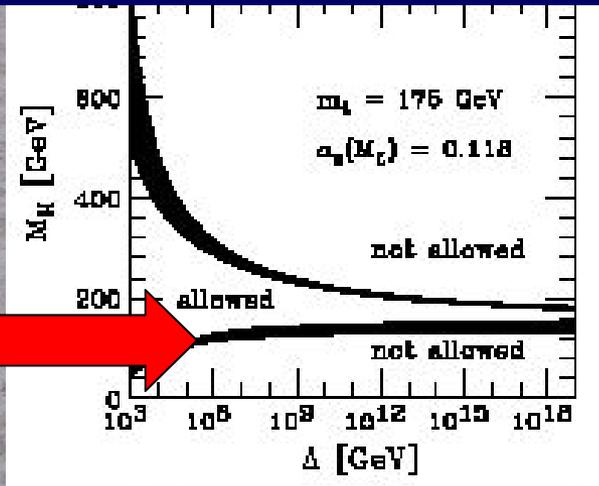
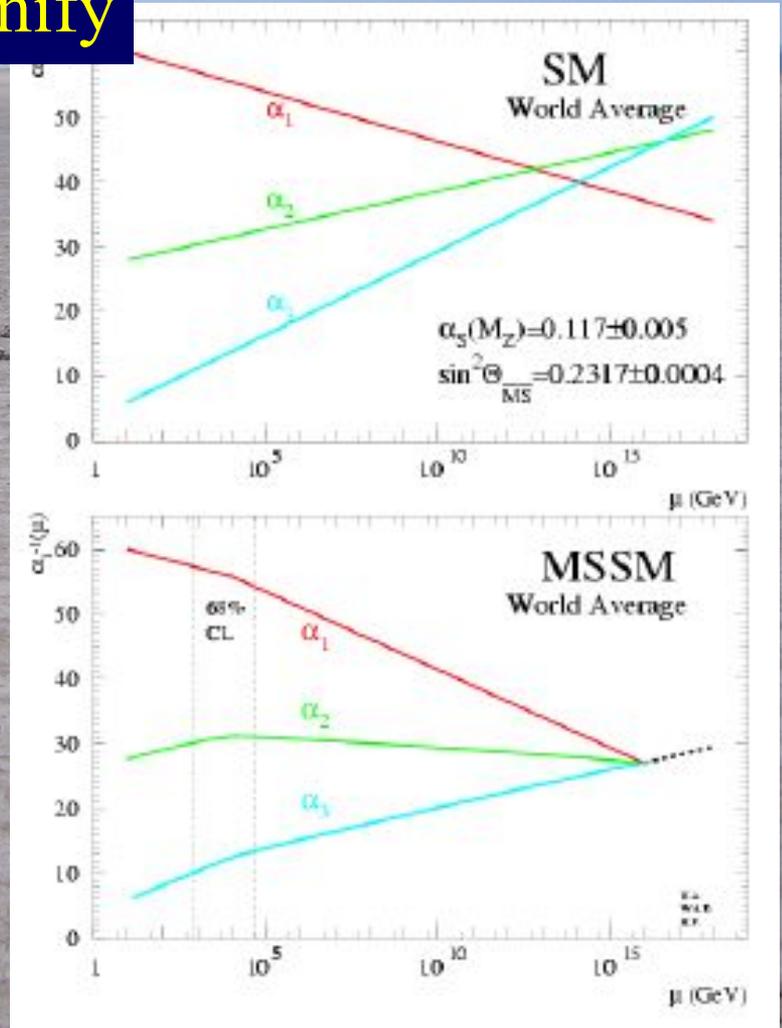
- Need $|m_B^2 - m_F^2| < 1 \text{ TeV}^2$

Other Reasons to like Susy

It enables the gauge couplings to unify

It predicts $m_H < 150$ GeV

It stabilizes the Higgs potential for low masses



Approved by Fabiola Gianotti

Dark Matter in the Universe

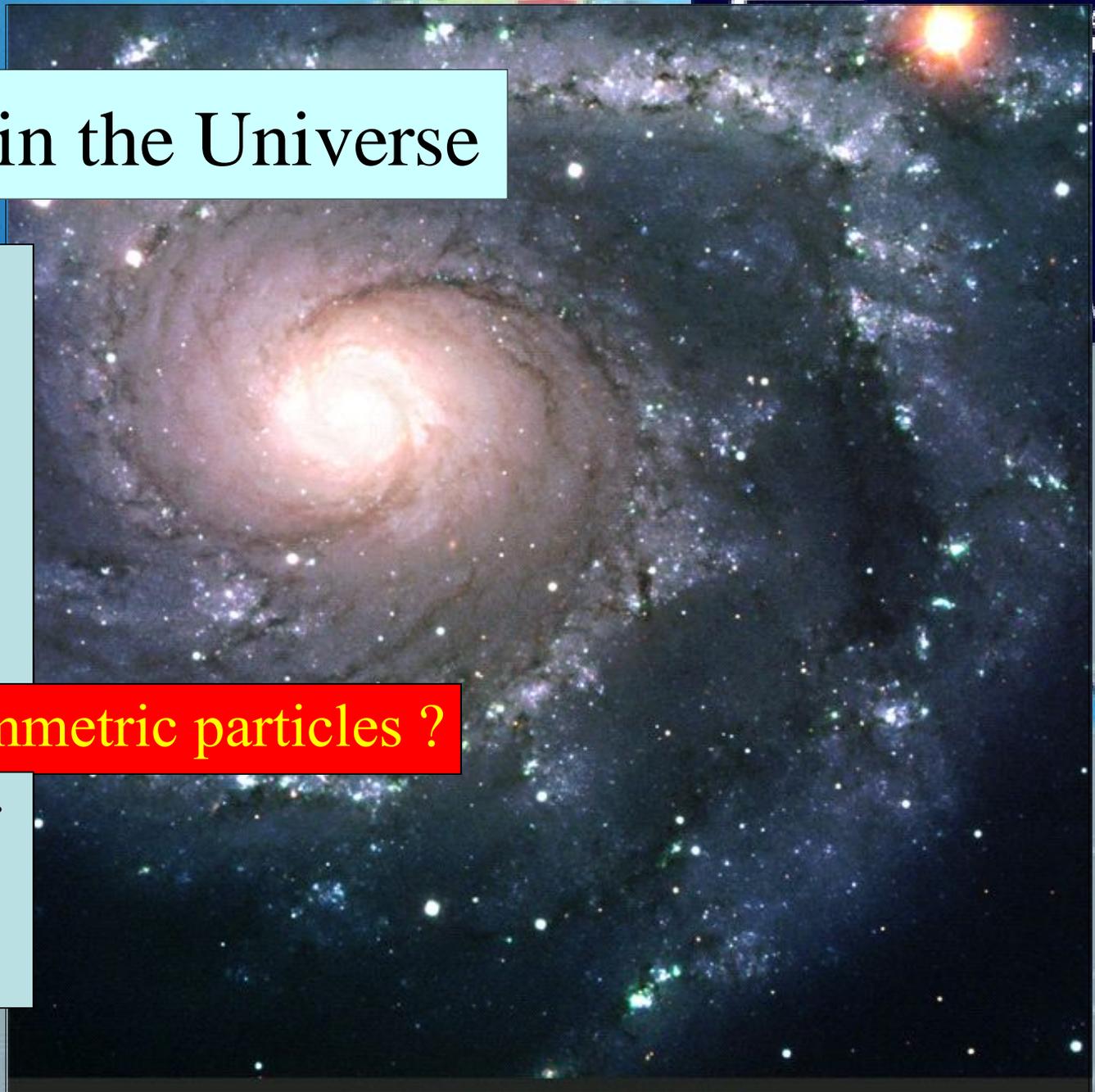
Astronomers say
that most of the
matter in the
Universe is
invisible

Dark Matter

Lightest Supersymmetric particles ?

We shall look for
them with the

LHC



Constraints on Supersymmetry

- Absence of sparticles at LEP, Tevatron

selectron, chargino > 100 GeV

squarks, gluino > 250 GeV

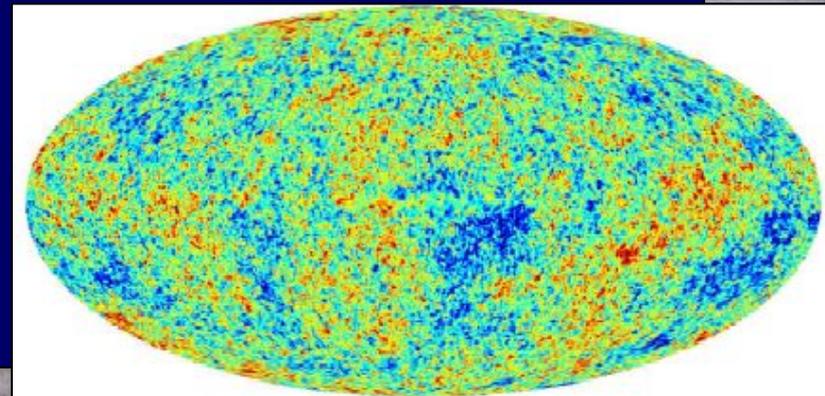
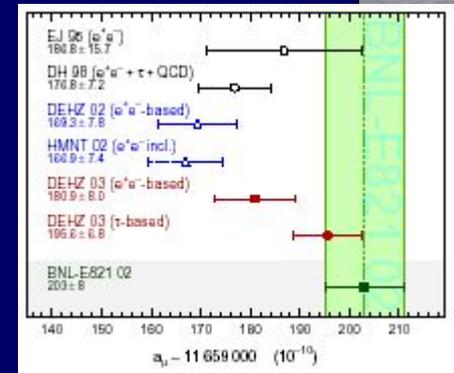
- Indirect constraints

Higgs > 114 GeV, $b \rightarrow s \gamma$ $g_{\mu} - 2$

- Density of dark matter

lightest sparticle χ :

WMAP: $0.094 < \Omega_{\chi} h^2 < 0.124$



Current Constraints on CMSSM

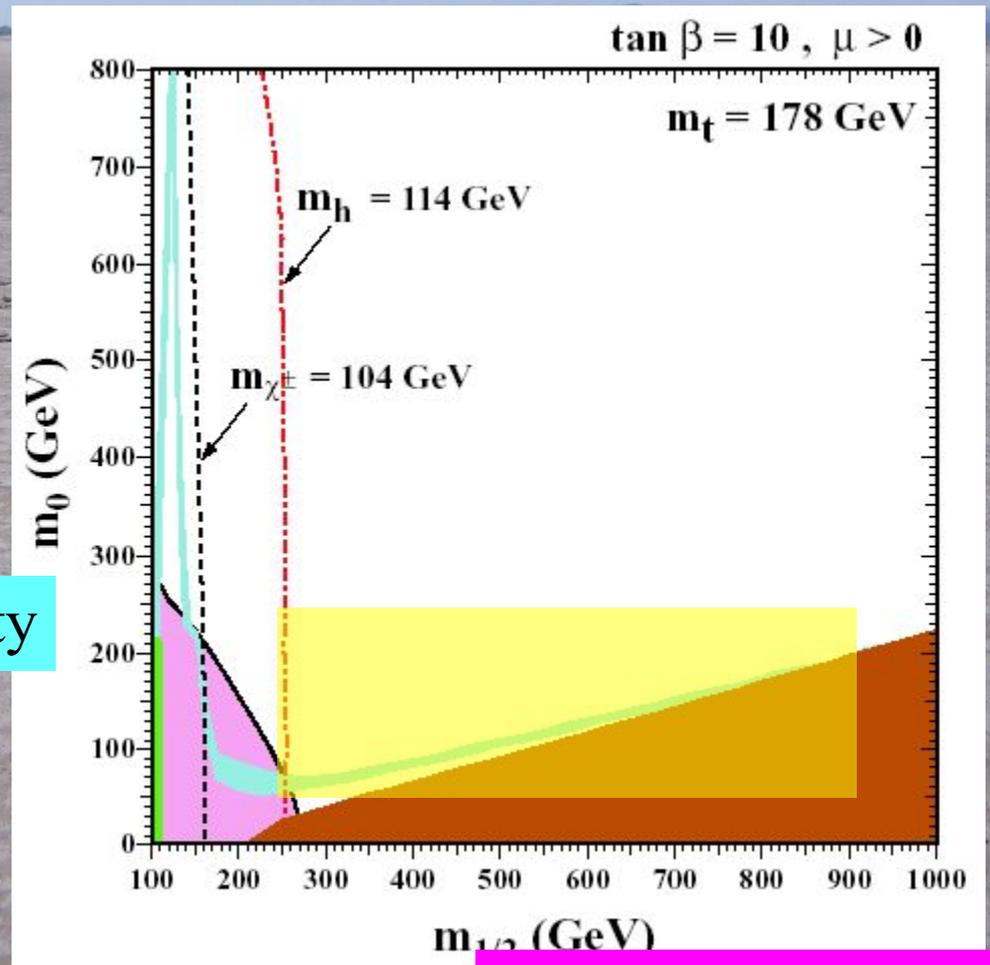
Assuming the lightest sparticle is a neutralino

Excluded because stau LSP

Excluded by $b \rightarrow s$ gamma

WMAP constraint on relic density

Excluded (?) by latest $g - 2$

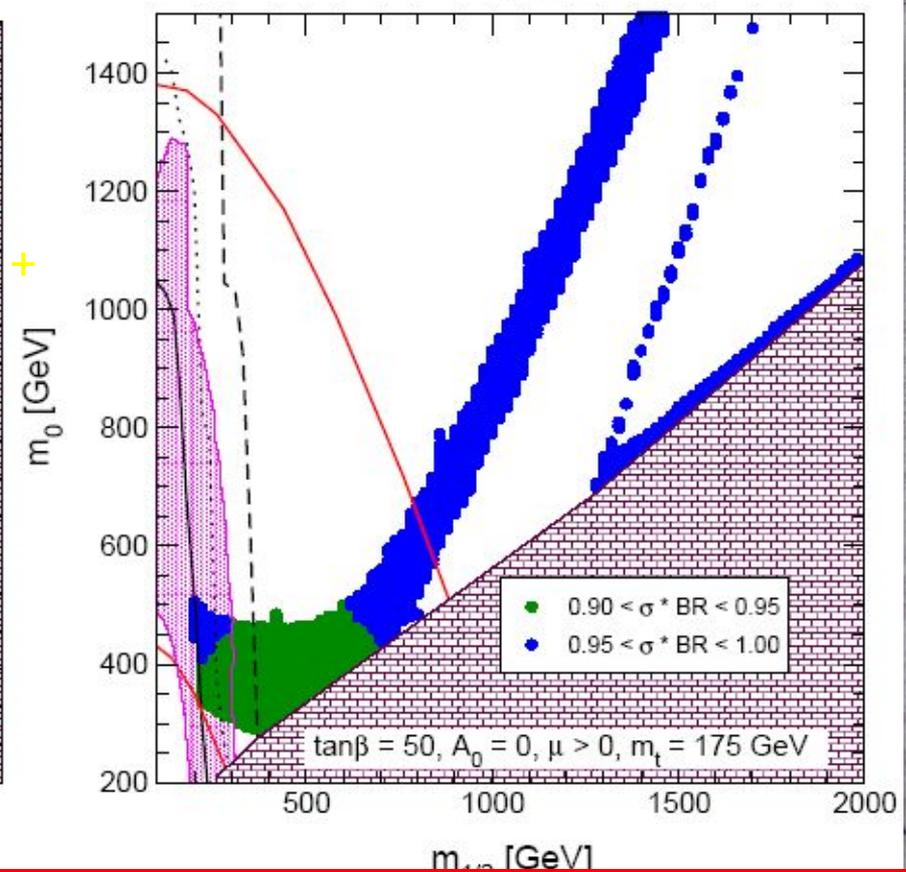
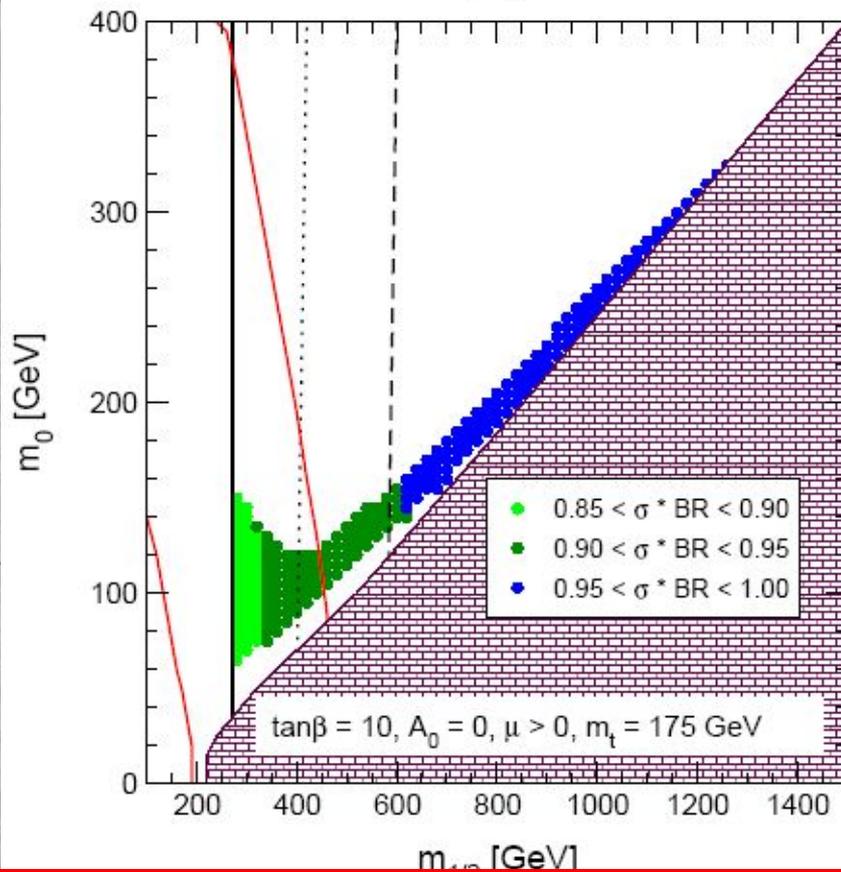


Higgs Production: CMSSM vs SM

$$\frac{[\sigma(gg \rightarrow h) \times \mathcal{B}(h \rightarrow \gamma\gamma)]_{\text{CMSSM}}}{[\sigma(gg \rightarrow h) \times \mathcal{B}(h \rightarrow \gamma\gamma)]_{\text{SM}}}$$

(a)

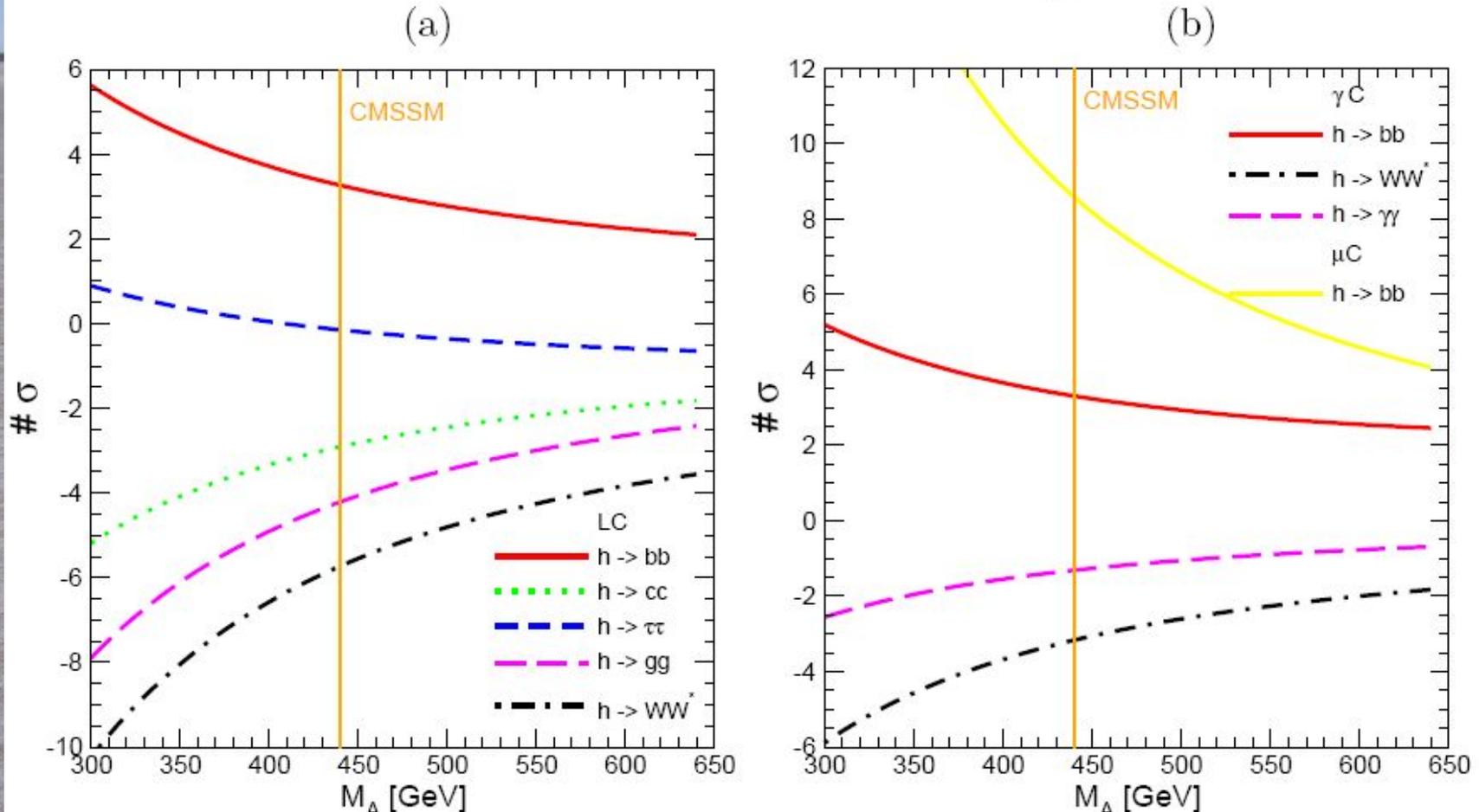
(b)



Good news: no suppression – Bad news: cannot distinguish CMSSM

CMSSM vs SM @ LC, $\gamma\gamma$ Collider

Variation of the $\sigma \times \mathcal{B}$ with M_A



Good news: can hope to distinguish CMSSM from Standard Model

Possible CP-Violating Asymmetries

in CP-violating
scenario with
three-way mixing:

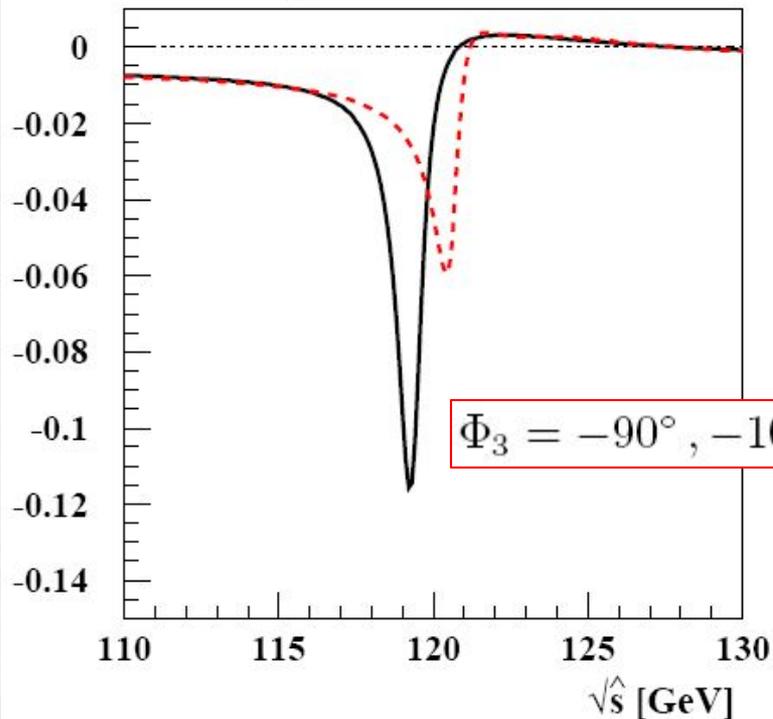
$$\tan \beta = 50, \quad M_{H^\pm}^{\text{pole}} = 155 \text{ GeV},$$

$$M_{\tilde{Q}_3} = M_{\tilde{U}_3} = M_{\tilde{D}_3} = M_{\tilde{L}_3} = M_{\tilde{E}_3} = M_{\text{SUSY}} = 0.5 \text{ TeV},$$

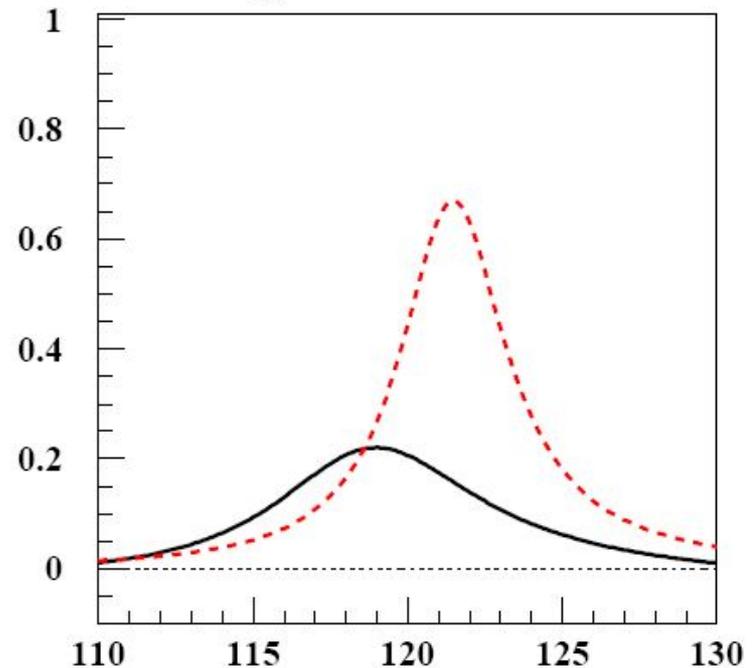
$$|\mu| = 0.5 \text{ TeV}, \quad |A_{t,b,\tau}| = 1 \text{ TeV}, \quad |M_2| = |M_1| = 0.3 \text{ TeV}, \quad |M_3| = 1 \text{ TeV},$$

$$\Phi_\mu = 0^\circ, \quad \Phi_A = \Phi_{A_t} = \Phi_{A_b} = \Phi_{A_\tau} = 90^\circ, \quad \Phi_1 = \Phi_2 = 0^\circ,$$

$a_{\text{CP}}(pp(gg,bb) \rightarrow \tau^+ \tau^- X)$

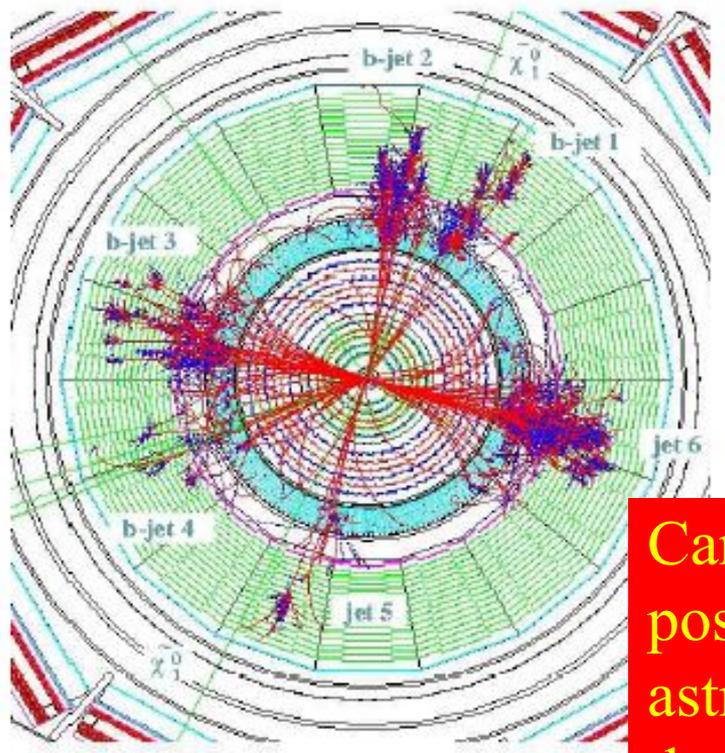


$a_{\text{CP}}(pp(W^+W) \rightarrow \tau^+ \tau^- X)$

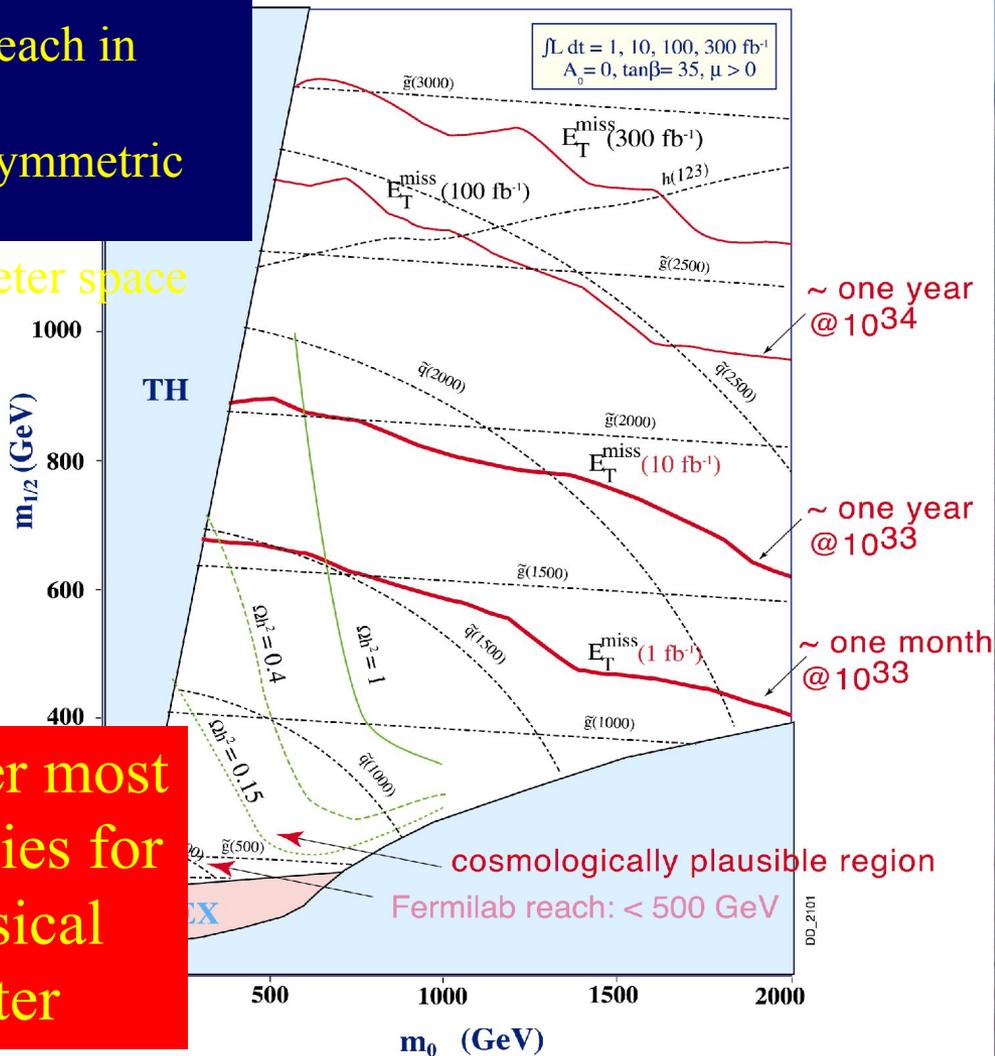


Supersymmetry Searches at LHC

'Typical' supersymmetric
Event at the LHC



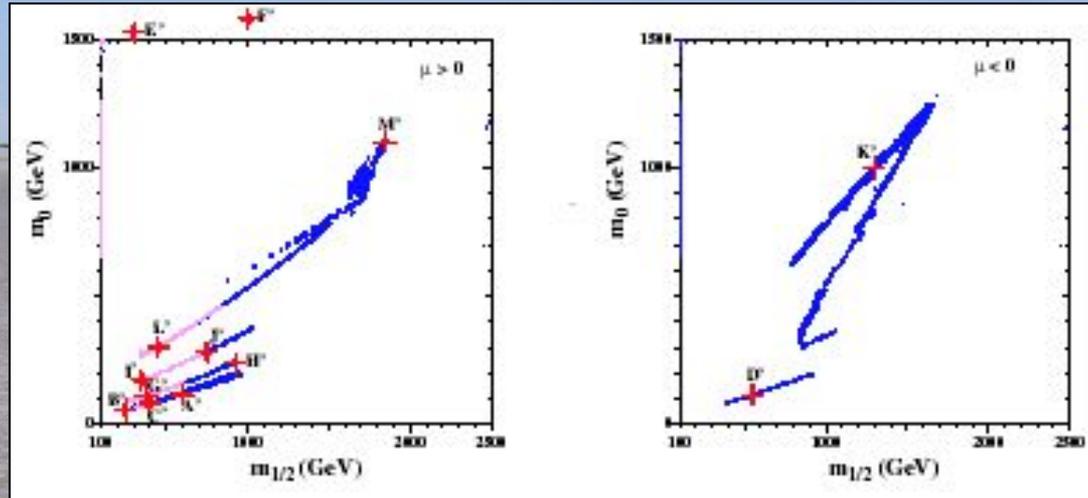
LHC reach in
supersymmetric
parameter space



Can cover most
possibilities for
astrophysical
dark matter

Supersymmetric Benchmark Studies

Lines in
susy space
allowed by
accelerators,



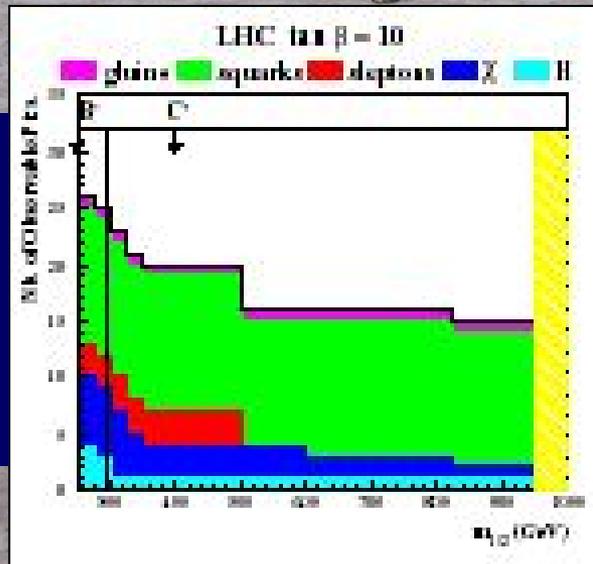
Specific
benchmark
Points along

WMAP lines

WMAP data

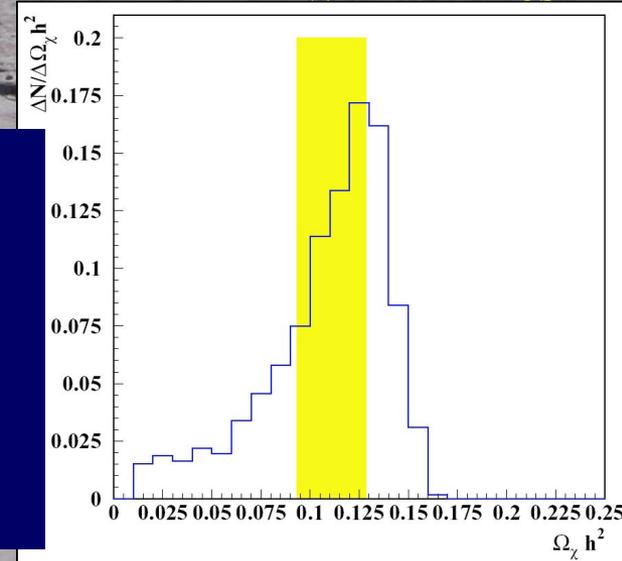
Sparticle
detectability
Along one

WMAP line



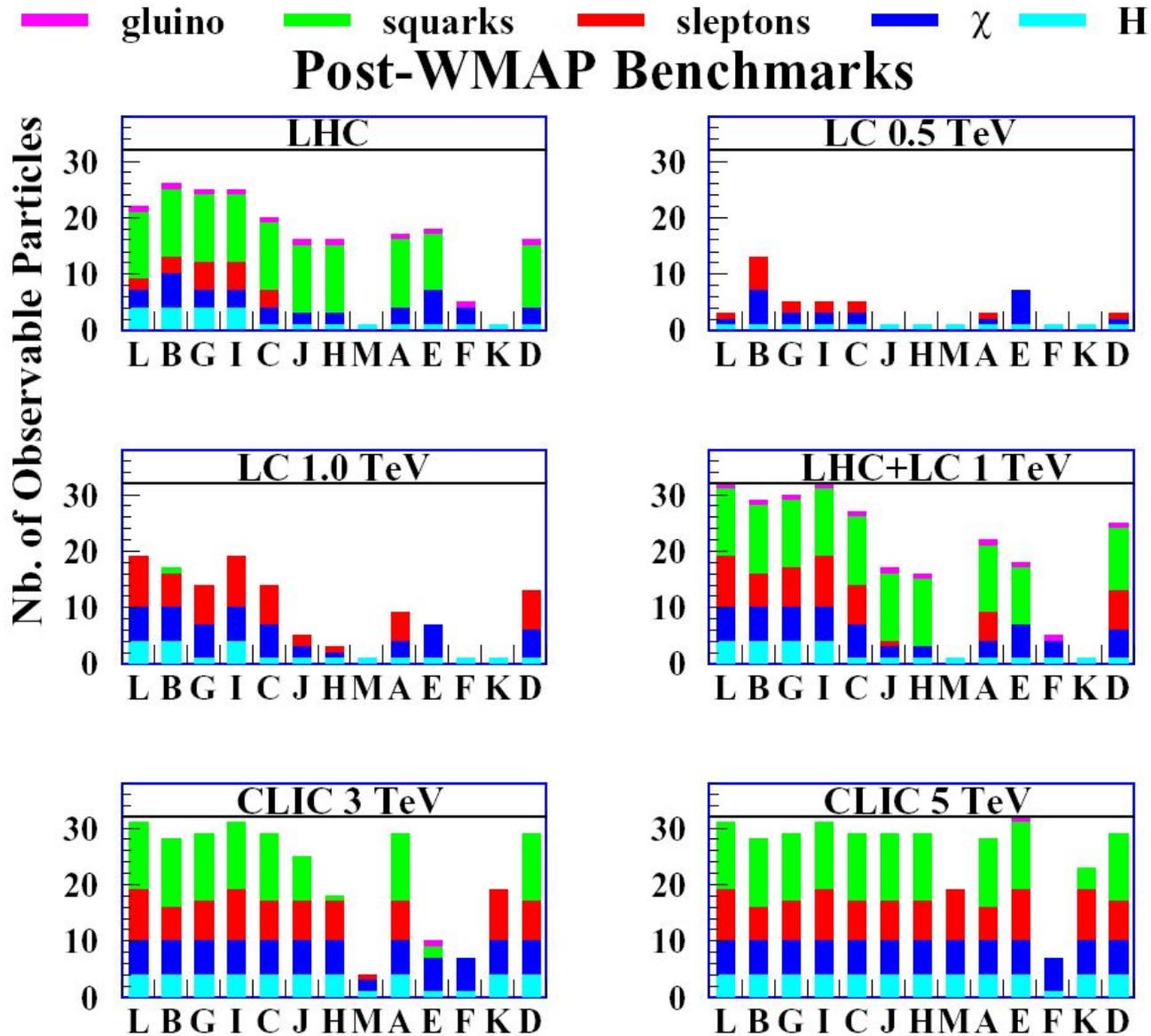
Calculation
of relic
density at a

benchmark



Summary of LHC Scapabilities ... and Other Accelerators

LHC almost
'guaranteed'
to discover
supersymmetry
if it is relevant

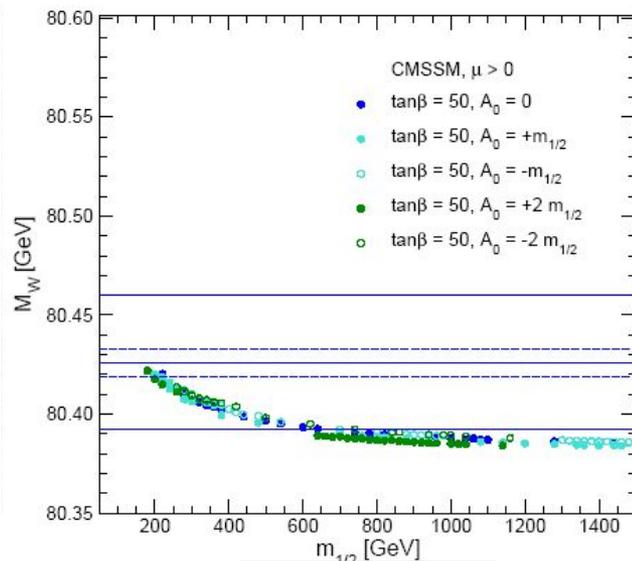
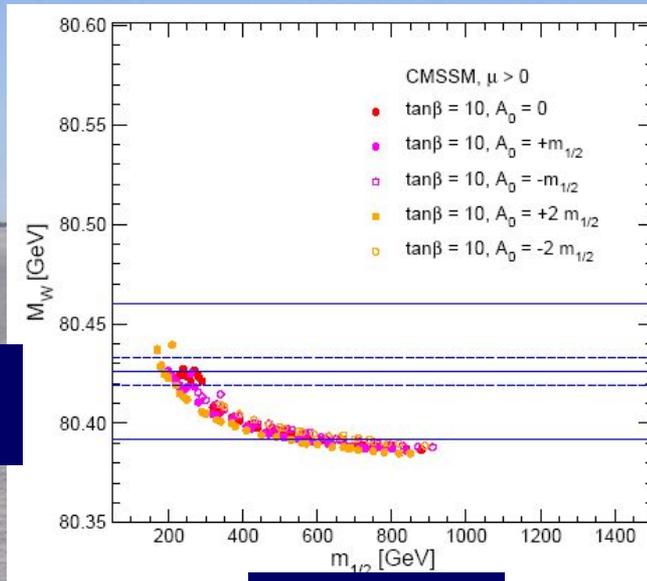


Precision Observables in Susy

Sensitivity to $m_{1/2}$
in CMSSM
along WMAP lines
for different A

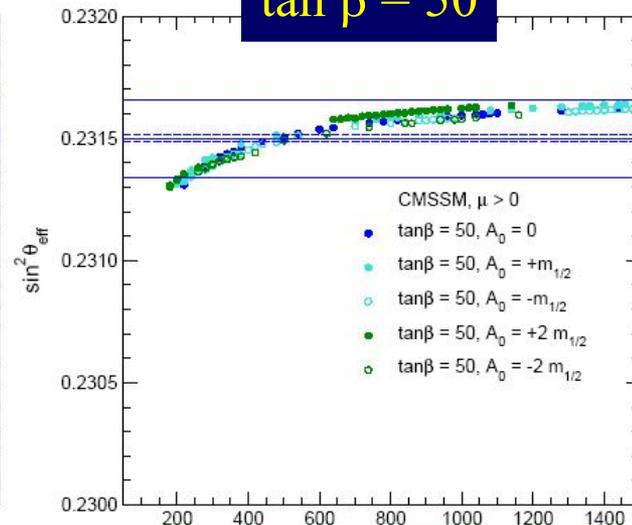
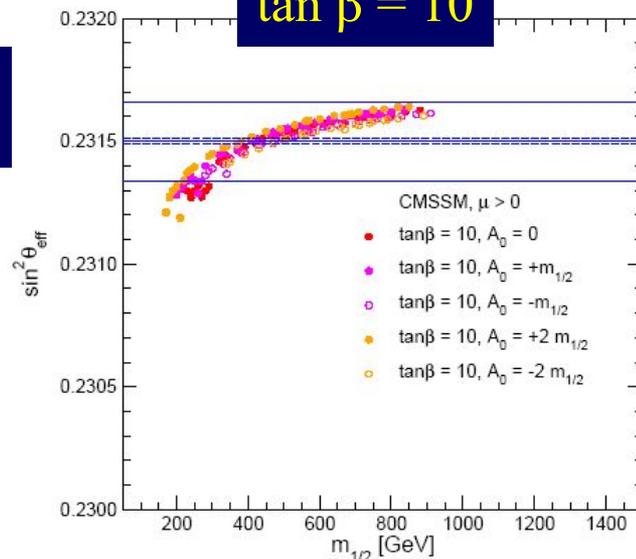
m_W

$\sin^2\theta_W$



$\tan\beta = 10$

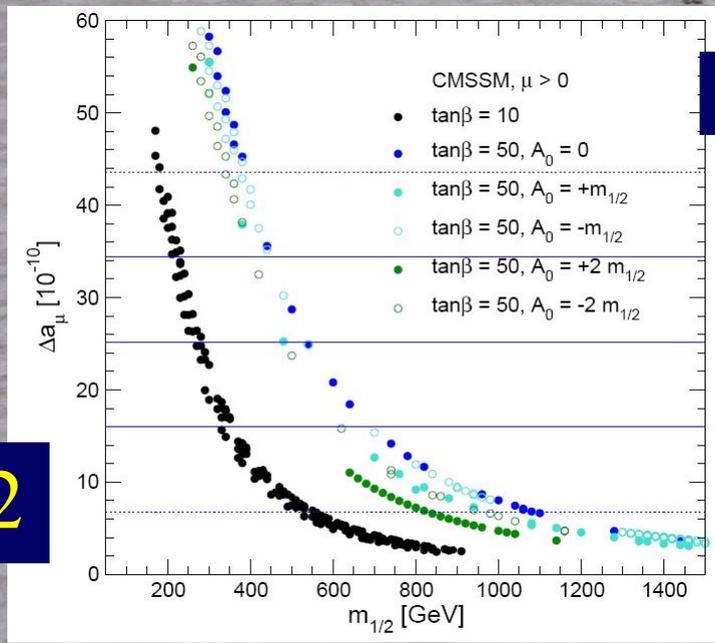
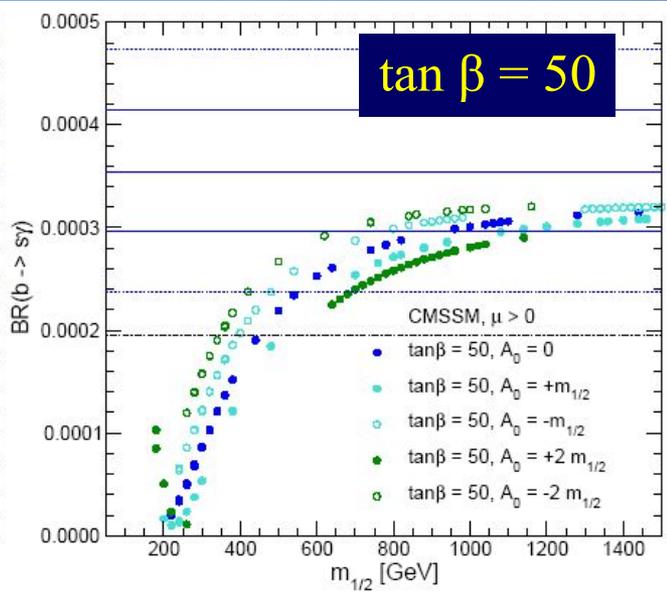
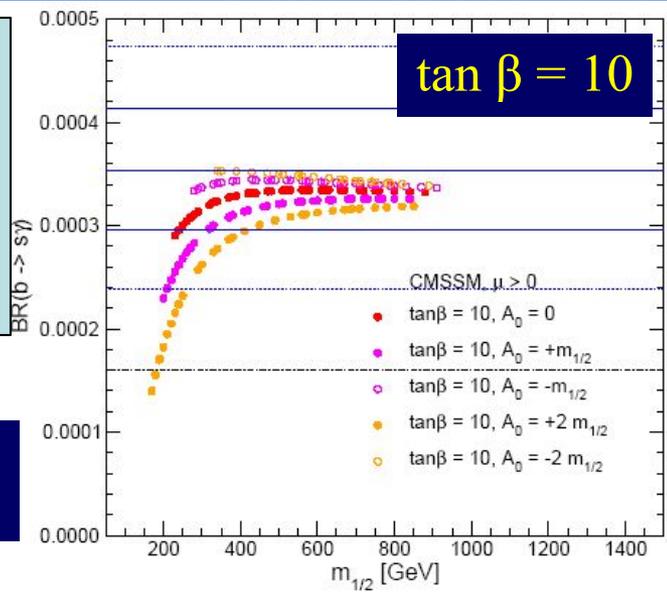
$\tan\beta = 50$



Present & possible
future errors

More Observables

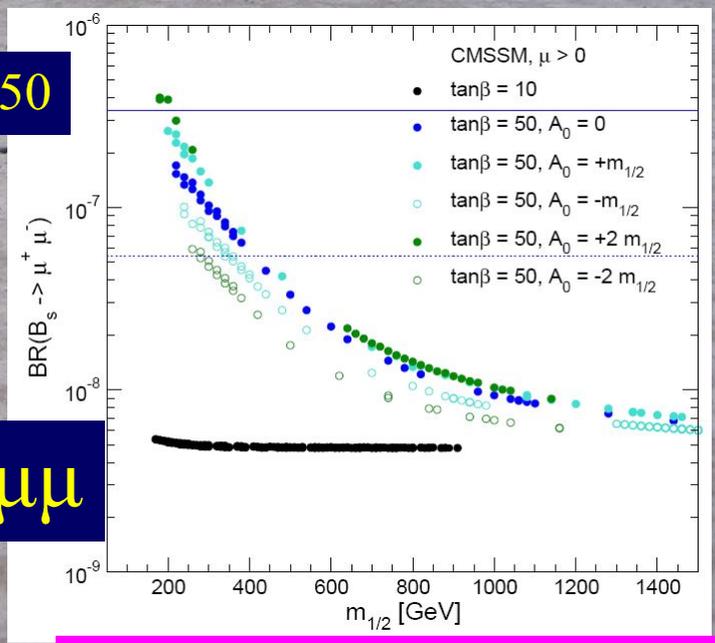
$b \rightarrow s\gamma$



$g_\mu - 2$

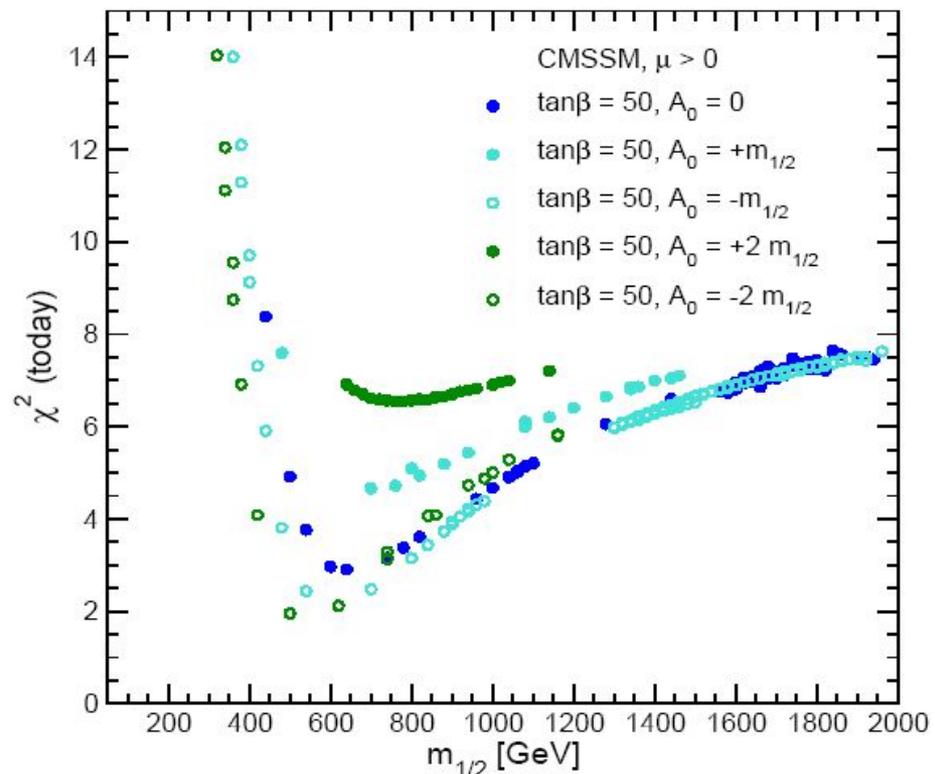
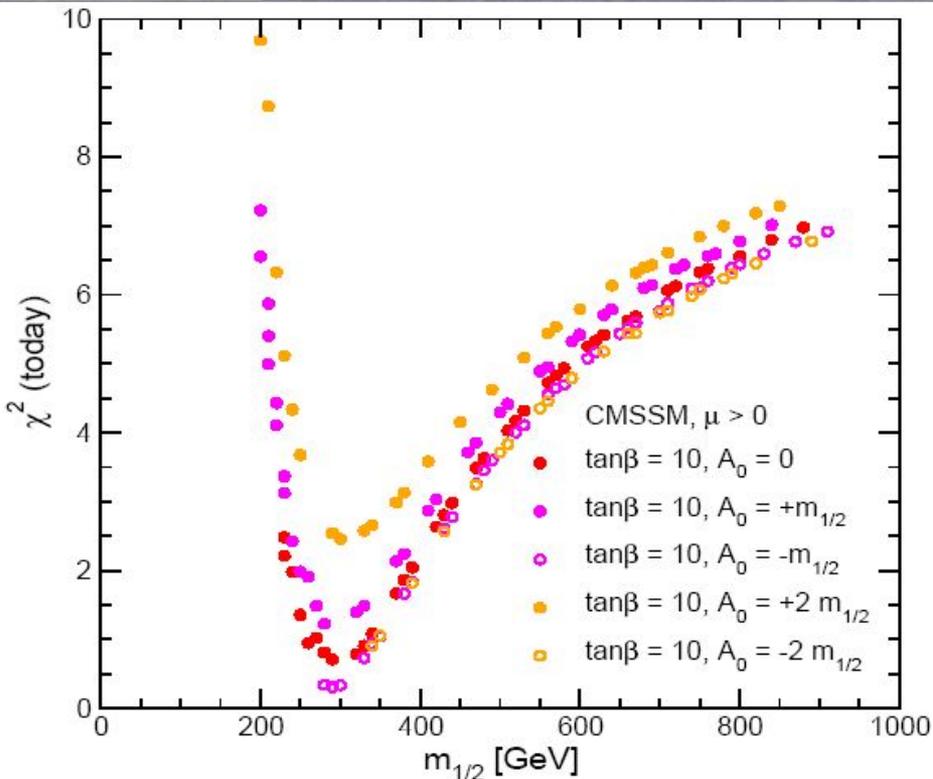
$\tan \beta = 10, 50$

$B_s \rightarrow \mu\mu$



Global Fits to Present Data

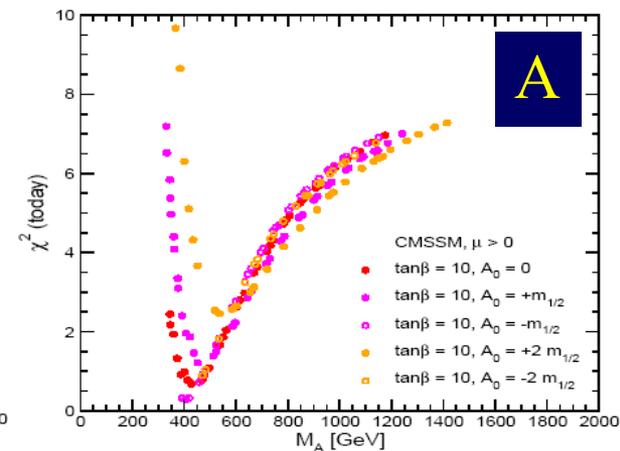
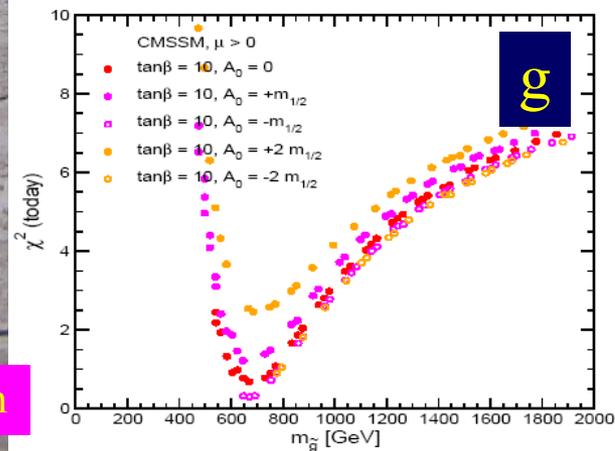
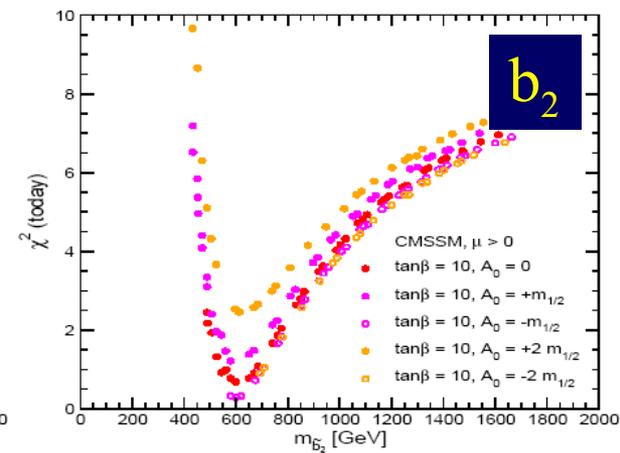
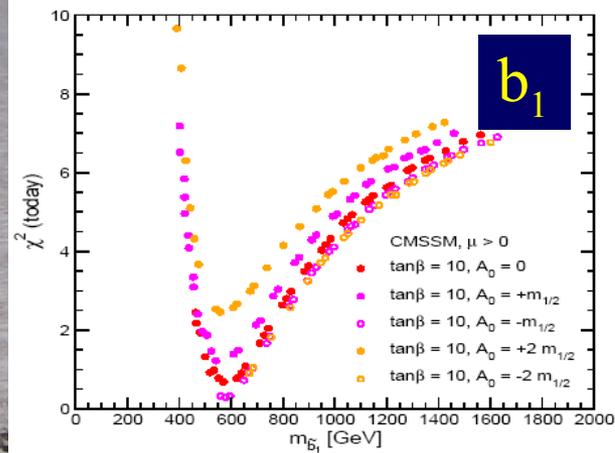
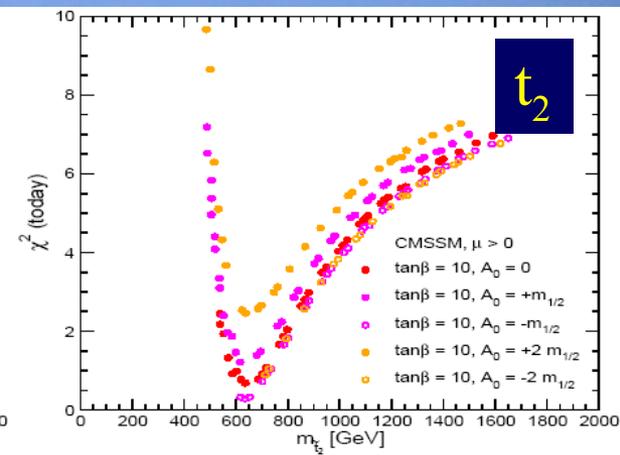
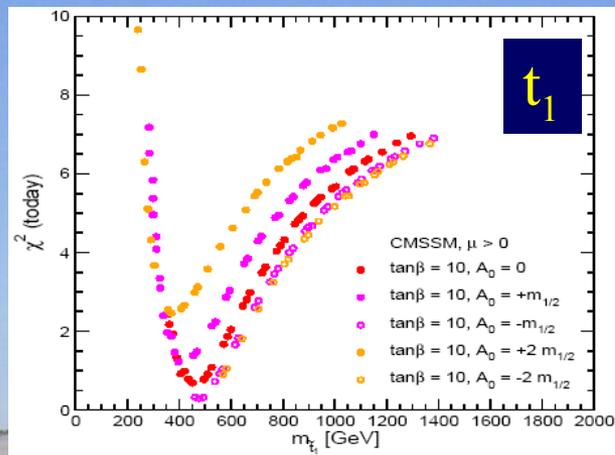
As functions of $m_{1/2}$ in CMSSM for $\tan\beta = 10, 50$



Global Fits to Present Data

Preferred
sparticle
masses for
 $\tan \beta = 10$

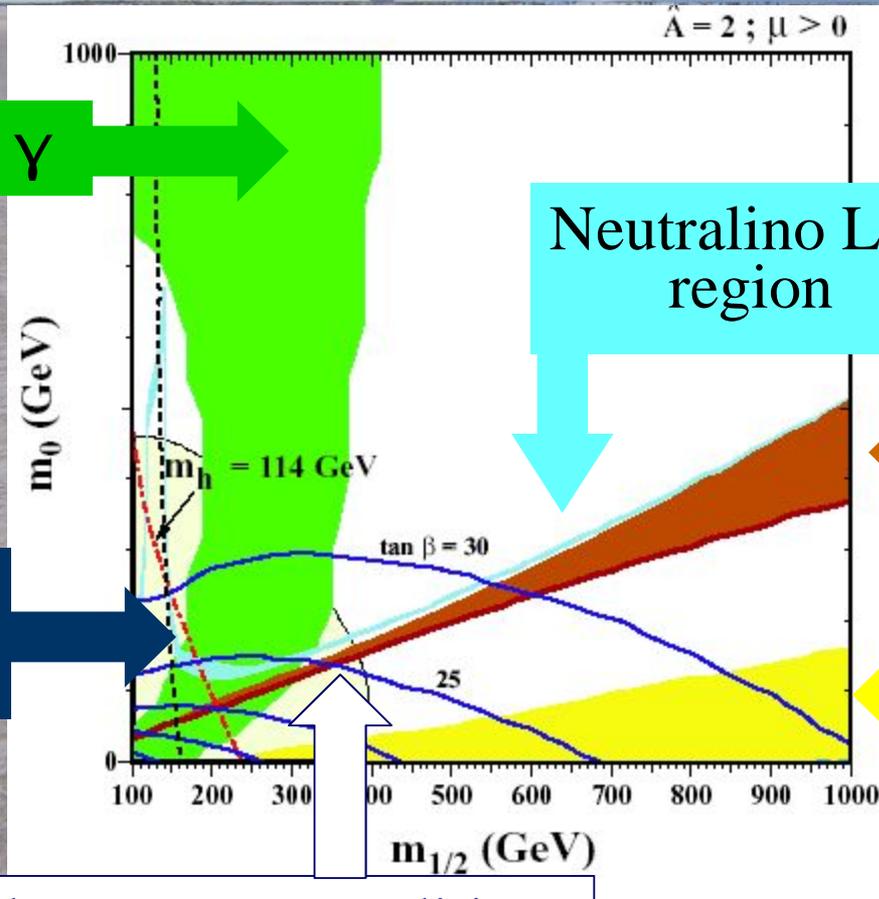
JE + Heinemeyer + Olive + Weiglein



Minimal Supergravity Model

More constrained than CMSSM: $m_{3/2} = m_0$, $B_\lambda = A_\lambda - 1$

Excluded by $b \rightarrow s \gamma$



Neutralino LSP region

stau LSP (excluded)

Gravitino LSP region

LEP constraints
On m_h , charging

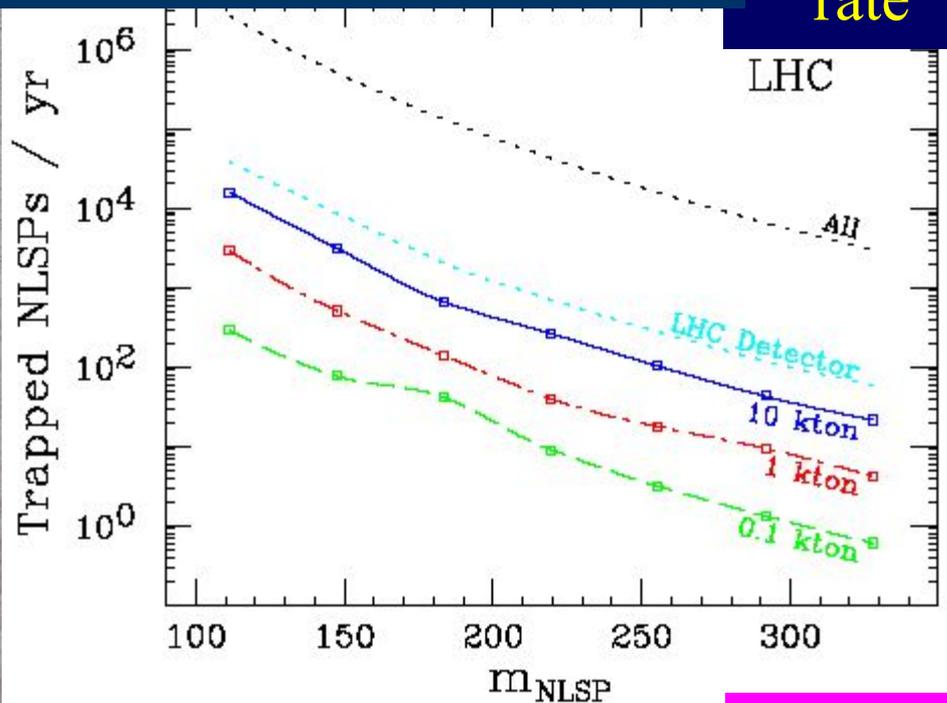
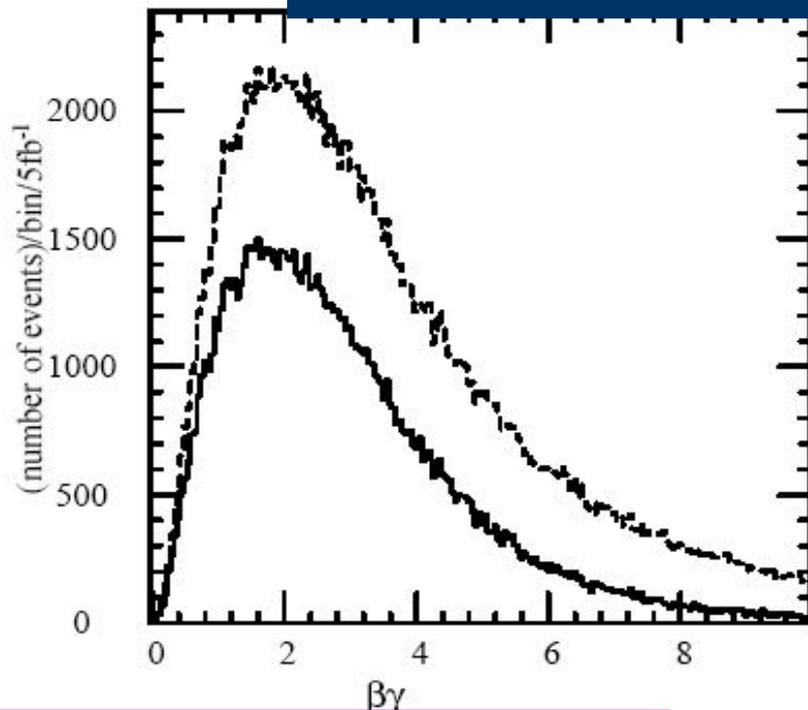
$\tan \beta$ fixed by vacuum conditions

Slepton Trapping at the LHC?

If stau next-to-lightest sparticle (NLSP)
may be metastable
may be stopped in detector/water tank
wait for them to decay: days, weeks, ...?

Kinematics

Trapping rate

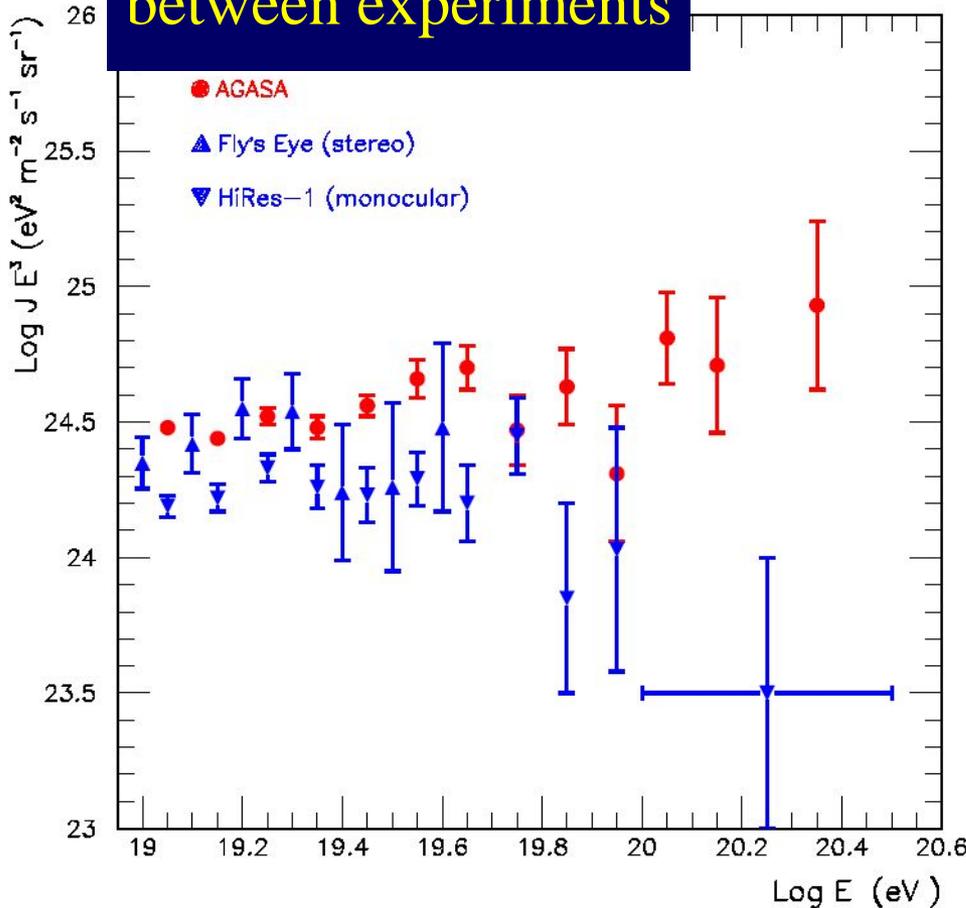


A wide, flat, sandy landscape under a clear blue sky. A path of footprints leads from the foreground towards the horizon. In the distance, there are low mountains and a small cluster of buildings. The overall scene is desolate and open.

Something completely different

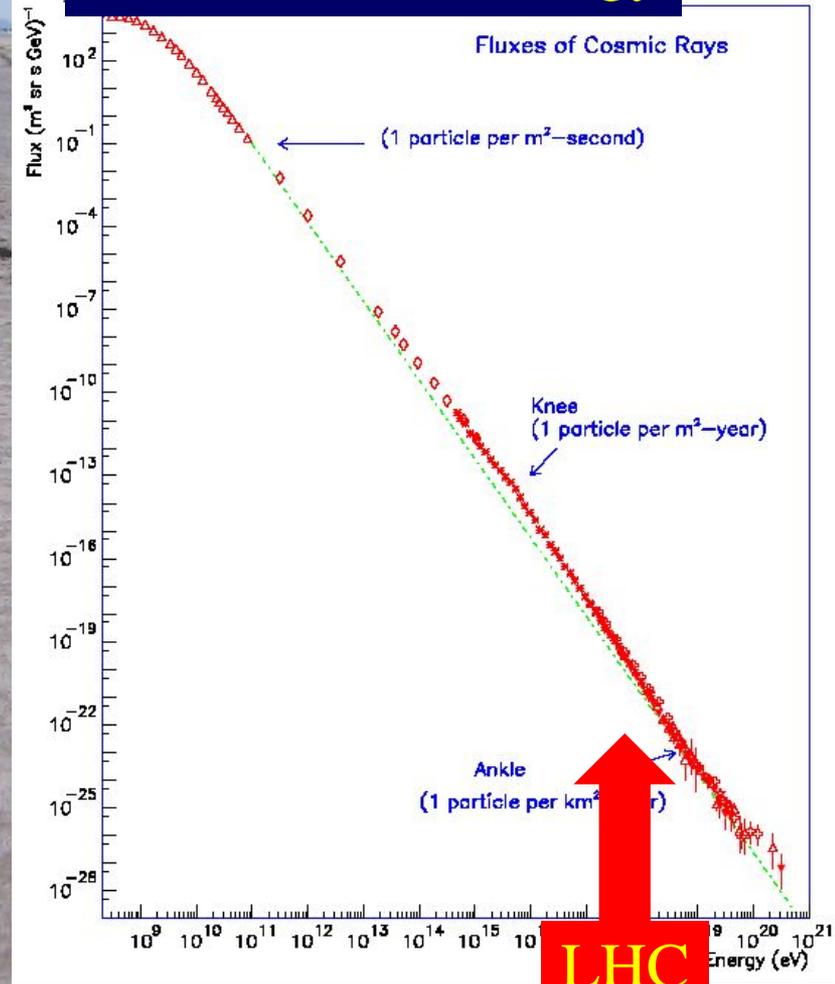
Ultra-High-Energy Cosmic Rays

Discrepancies between experiments



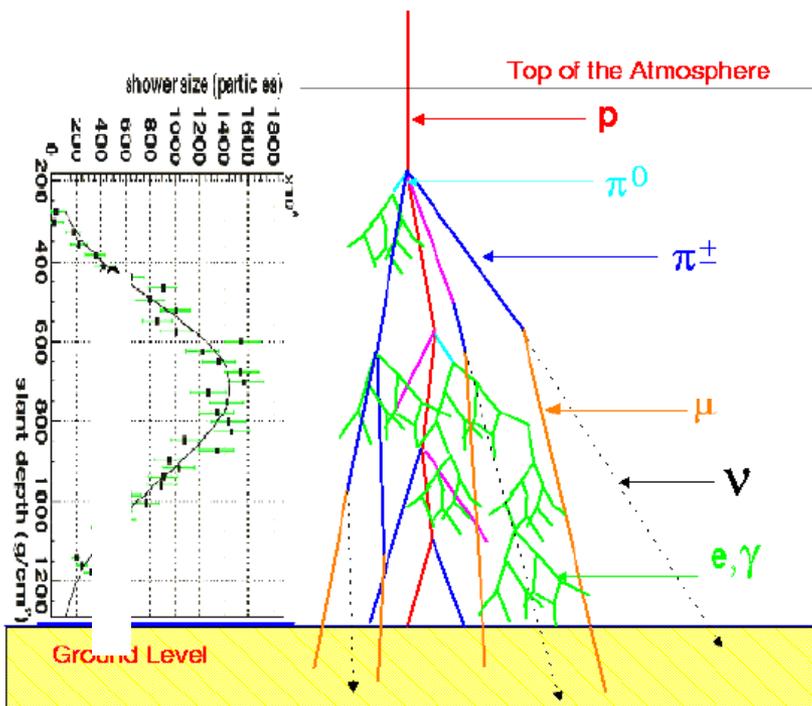
To be resolved by Auger laboratory

LHC closest in energy



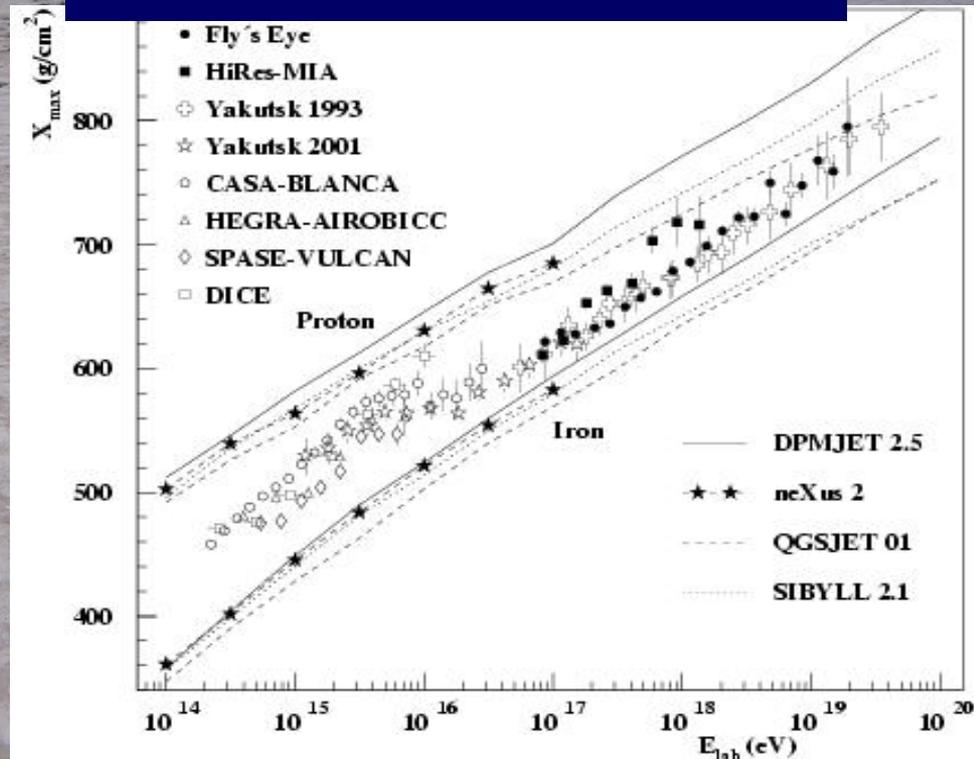
Understanding High-Energy Cosmic Rays

Development of extensive air shower

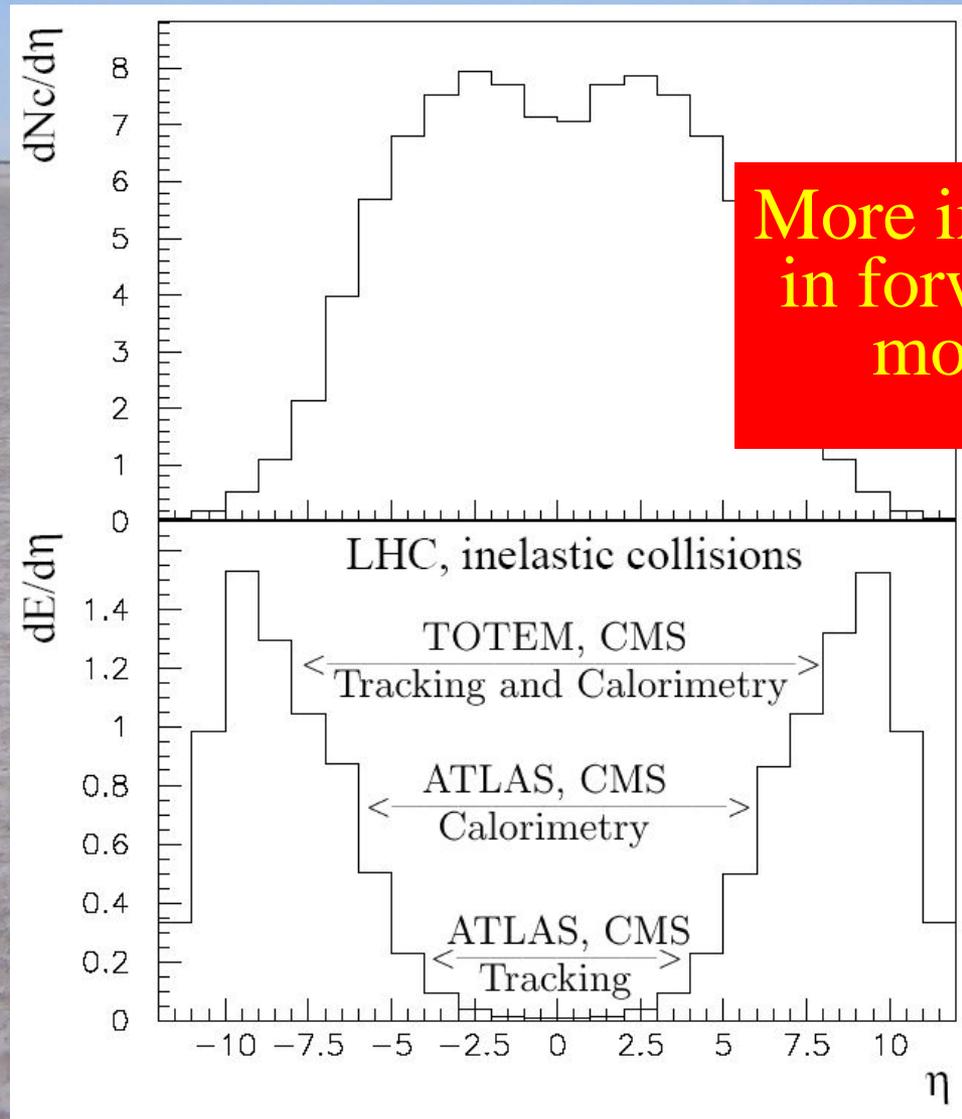


Extensive Air Showers

Uncertainties in composition due to model dependence



(Lack of) Coverage by LHC Detectors



More instrumentation
in forward direction
most welcome

A wide, flat, sandy landscape under a clear blue sky. A path of footprints leads from the foreground towards the horizon. In the distance, there are low mountains and a small cluster of buildings. The overall scene is bright and open.

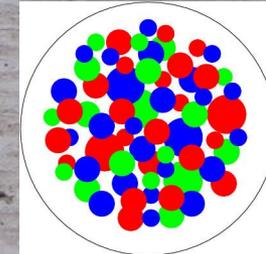
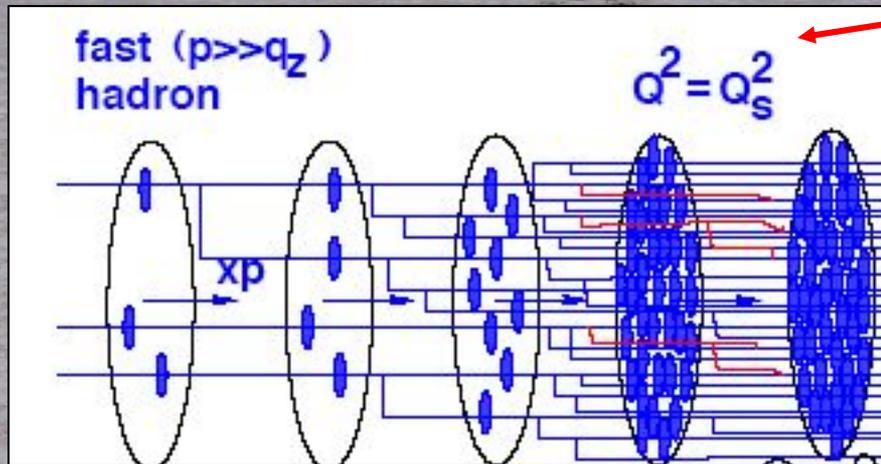
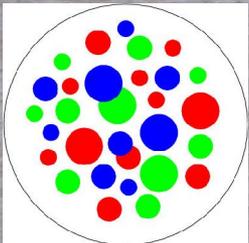
Back to forward QCD

Parton Saturation Effects

High energies s open access to small $x = Q^2/s$

Boundary of non-linear regime:
partons of size $1/Q > 1/Q_s$
overlap

large x



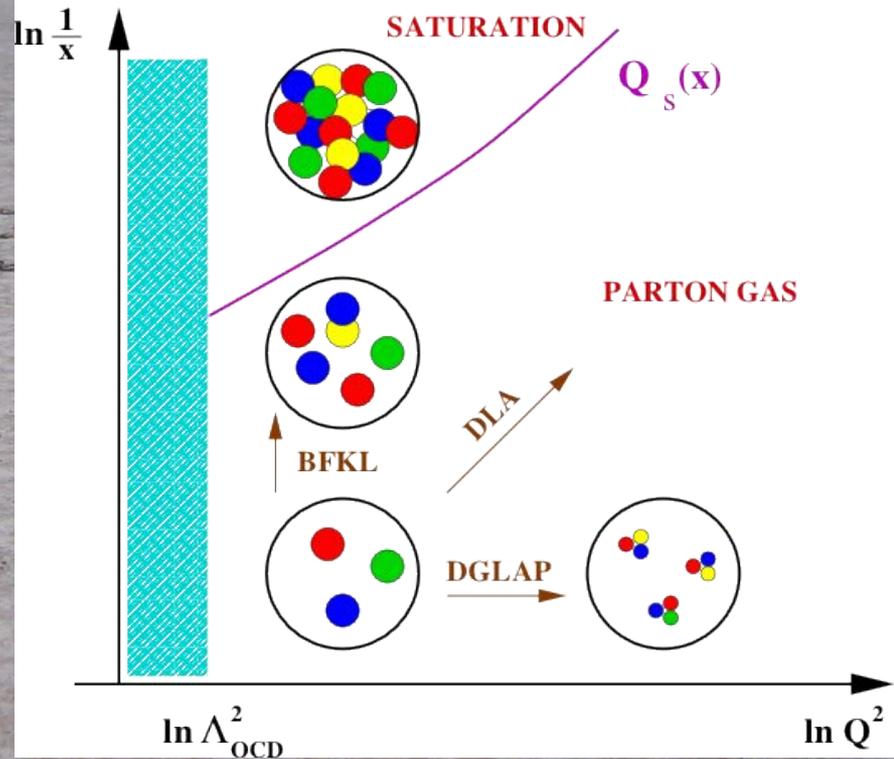
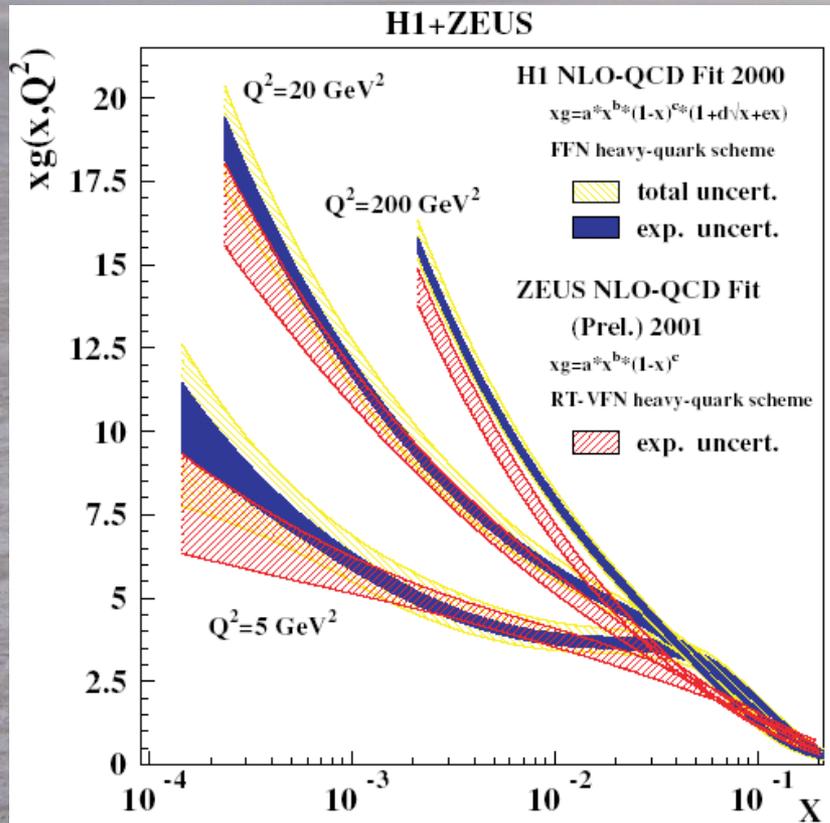
small x

Large probability to emit an extra gluon $\sim \alpha \ln(1/x) \sim 1$:
number of gluons at small x grows, transverse area limited
Transverse density becomes large

Non-linear QCD evolution and population growth

Linear evolution:

maximum population density

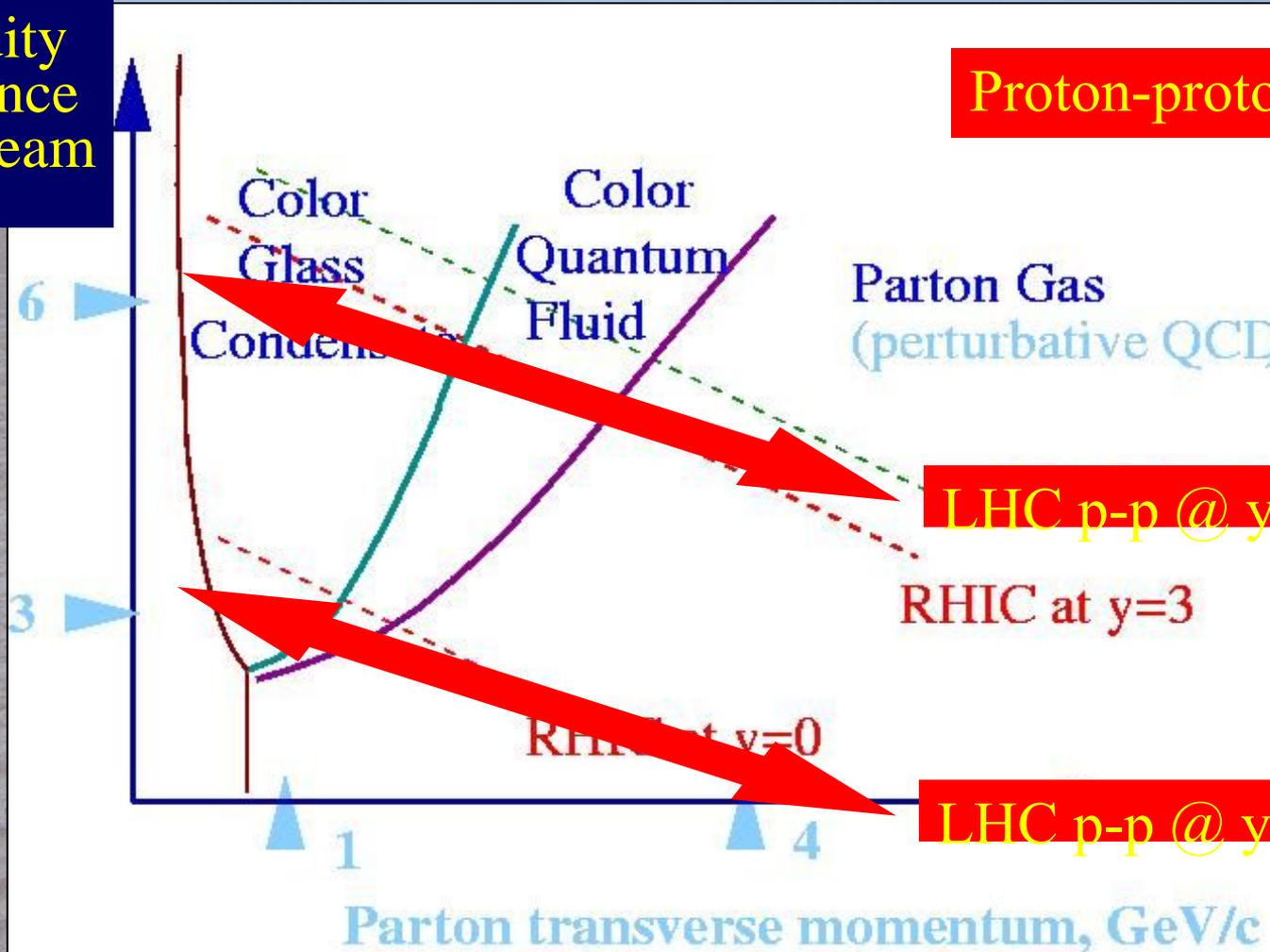


Malthus

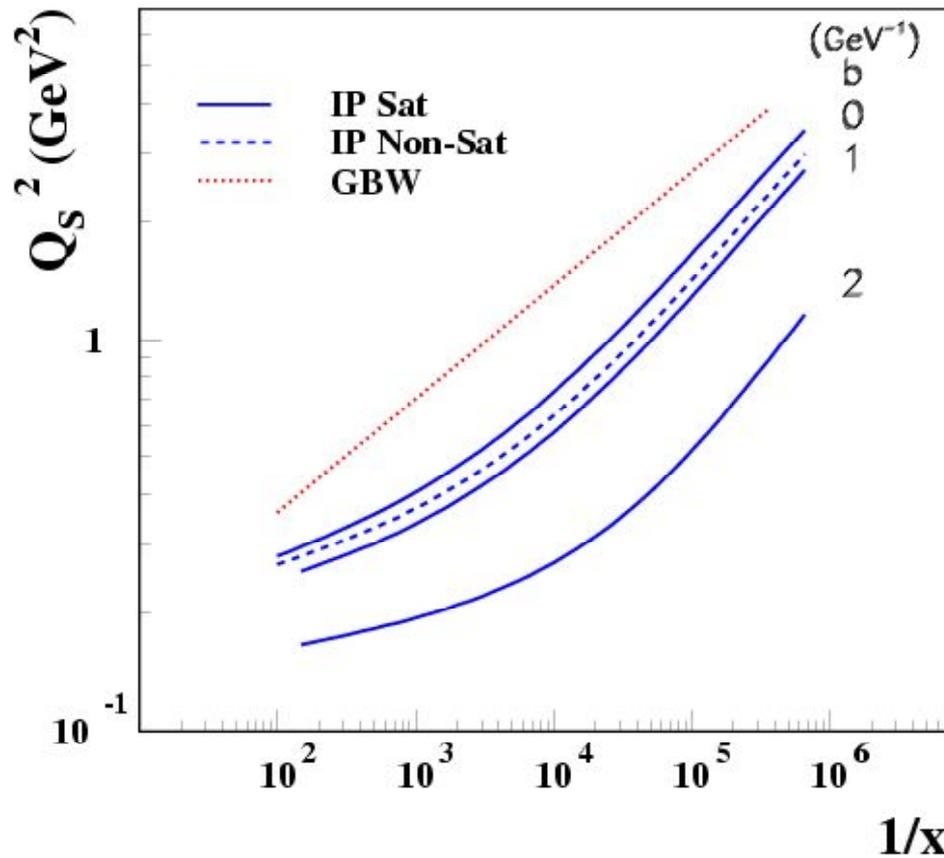
Color glass condensate

Phase diagram of high energy QCD

Rapidity difference from beam

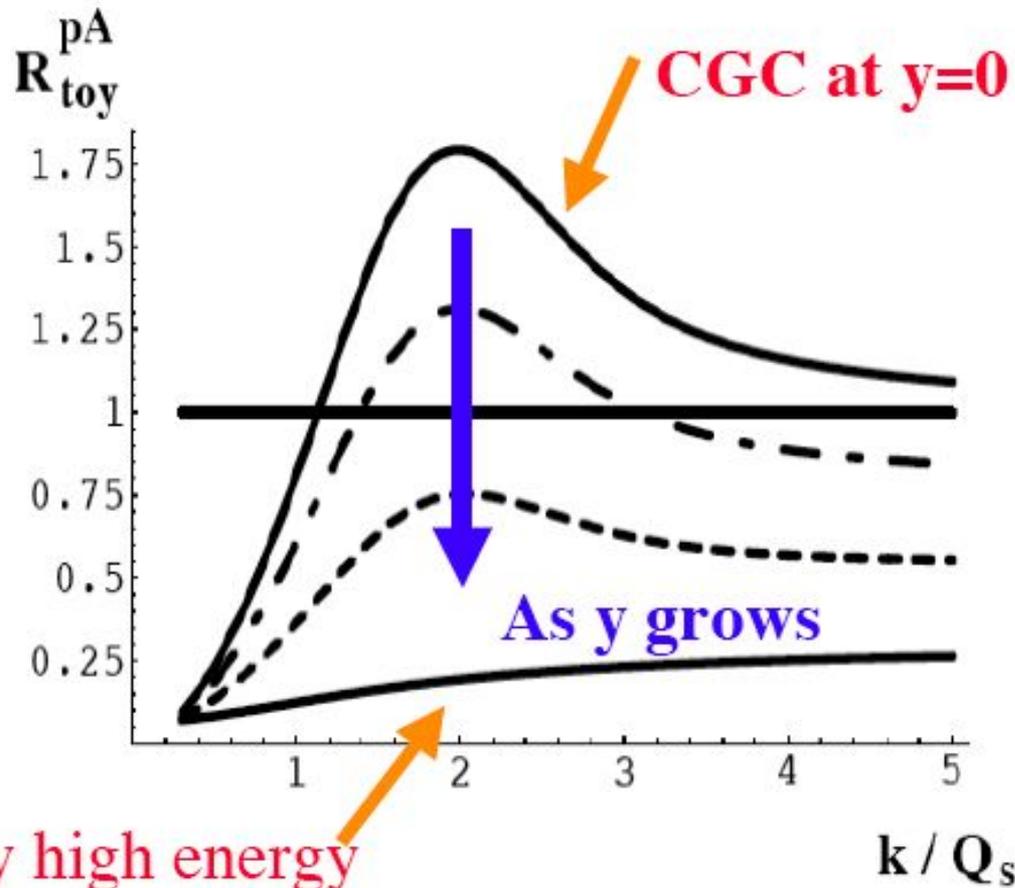


Extraction of Saturation Scale from HERA Data



Gluon multiplication in a limited (nuclear) environment

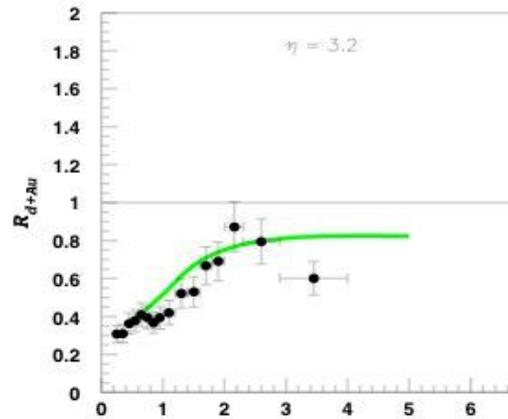
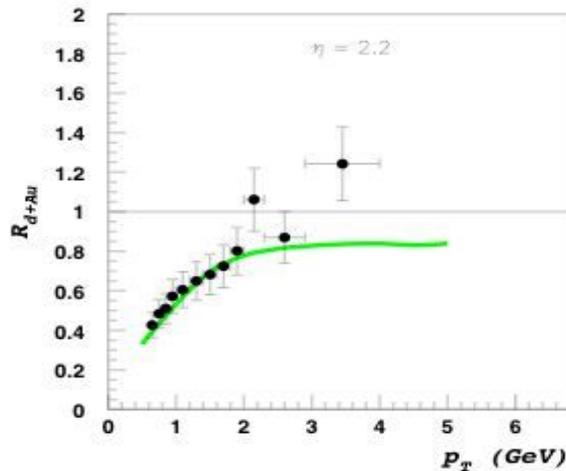
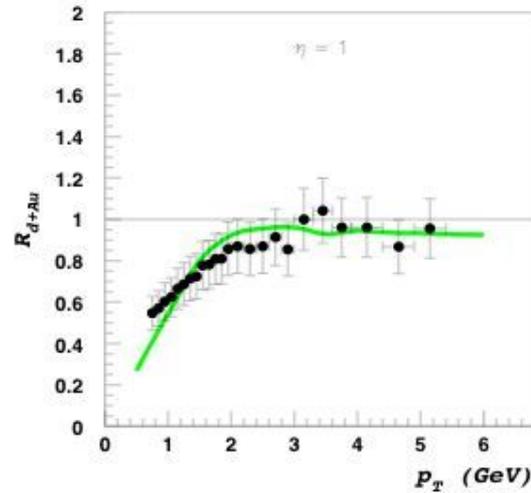
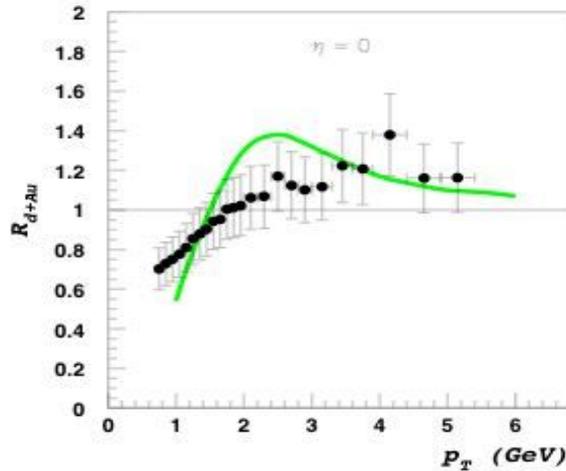
Ratio of pA and pp Cross sections



transverse momentum

At large rapidity y (small angle) expect suppression of hard particles

Color Glass Condensate: confronting the data



BRAHMS data:
 $\eta = 0, 1, 2.2, 3.2$



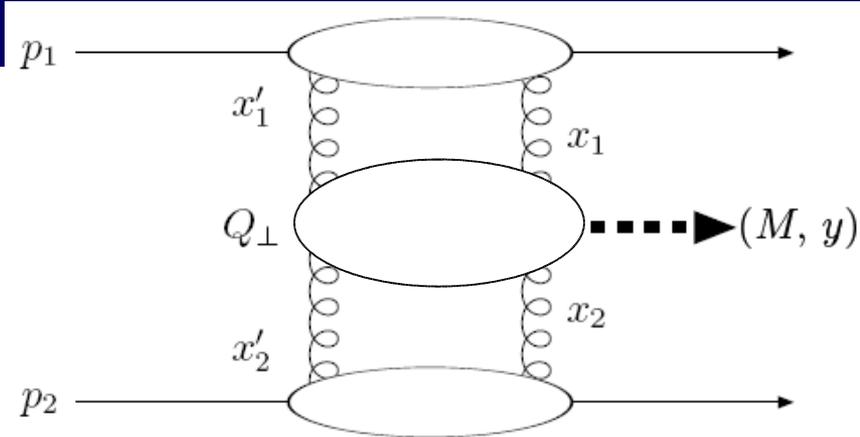
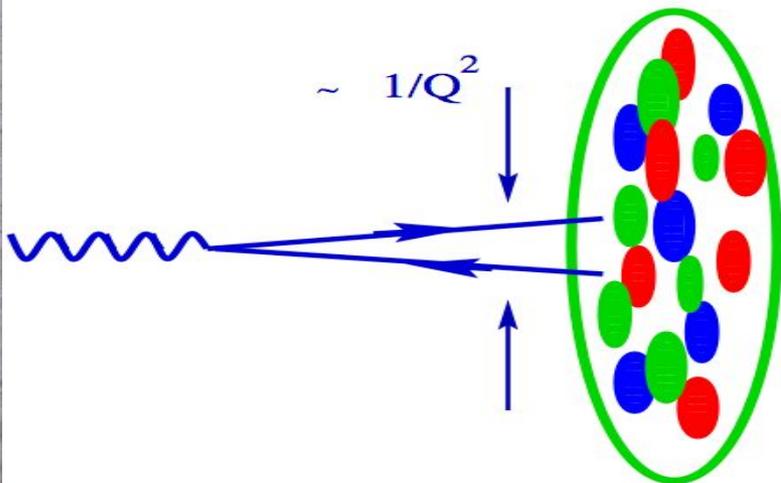
New Physics in Diffraction?

The Diffractive Menagerie

- Soft diffraction dissociation:
Peripheral proton-proton collision
dissociate proton \rightarrow low-mass system

- Hard diffraction:
Small colour dipole penetrates proton
produces very high-mass system

- Soft double diffraction:
Peripheral proton-proton collision
produces low-mass central system



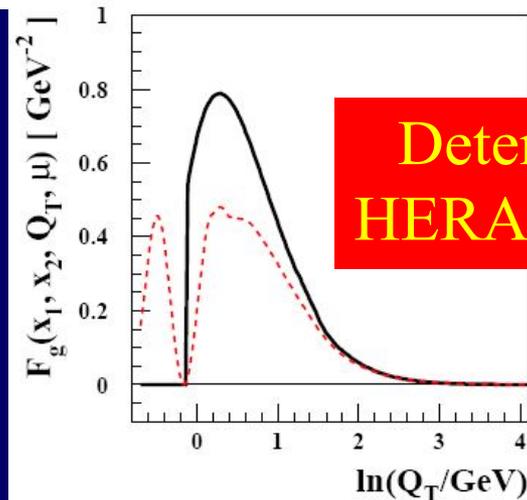
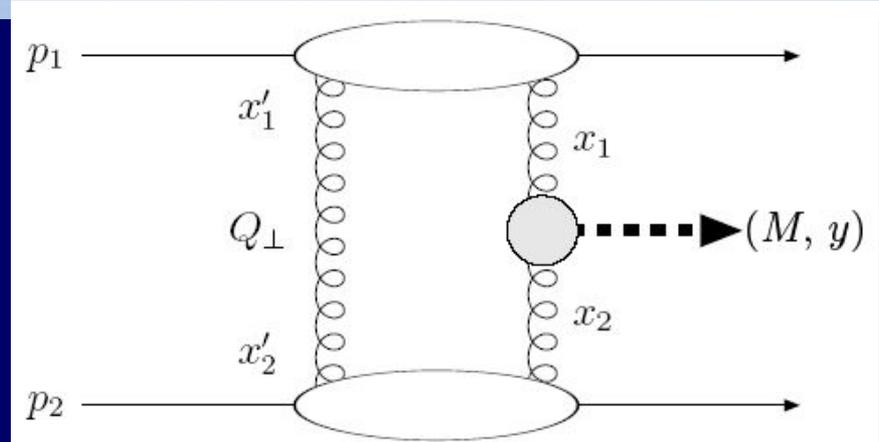
Diffractive Higgs Production

- Double-diffractive Higgs production mechanism:
- Effective luminosity:

$$M^2 \frac{\partial^2 \mathcal{L}}{\partial y \partial M^2} = 4.0 \times 10^{-4} \left[\frac{\int_{\ln Q_{\min}}^{\ln \mu} F_g(x_1, x_2, Q_T, \mu) d \ln Q_T}{\text{GeV}^{-2}} \right]^2 \left(\frac{\hat{S}^2}{0.02} \right) \left(\frac{4}{b \text{ GeV}^2} \right)^2 \left(\frac{R_g}{1.2} \right)^4$$

for nominal values
of other inputs

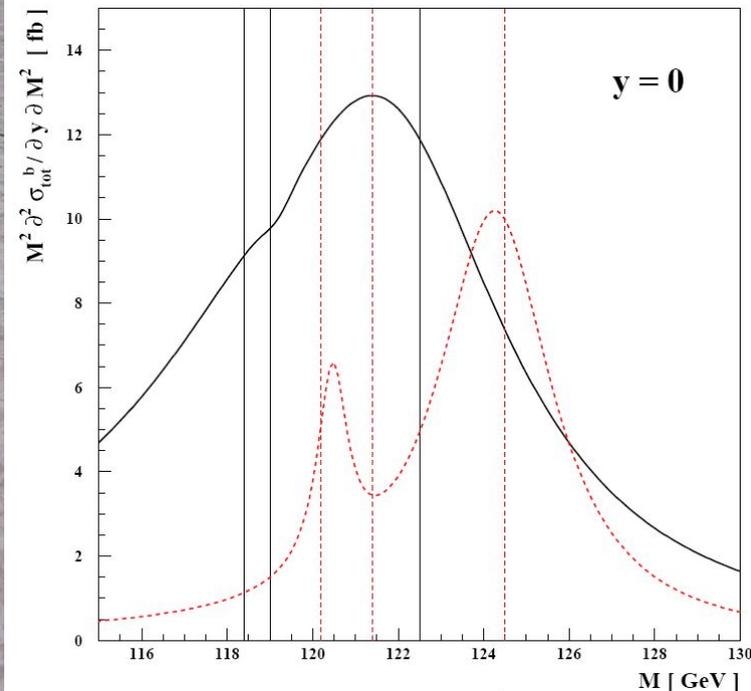
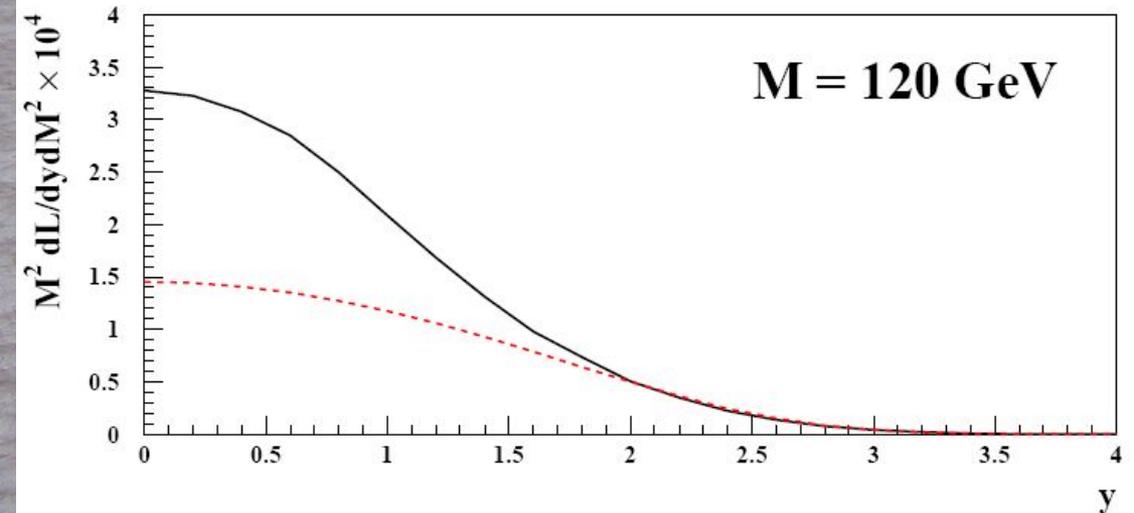
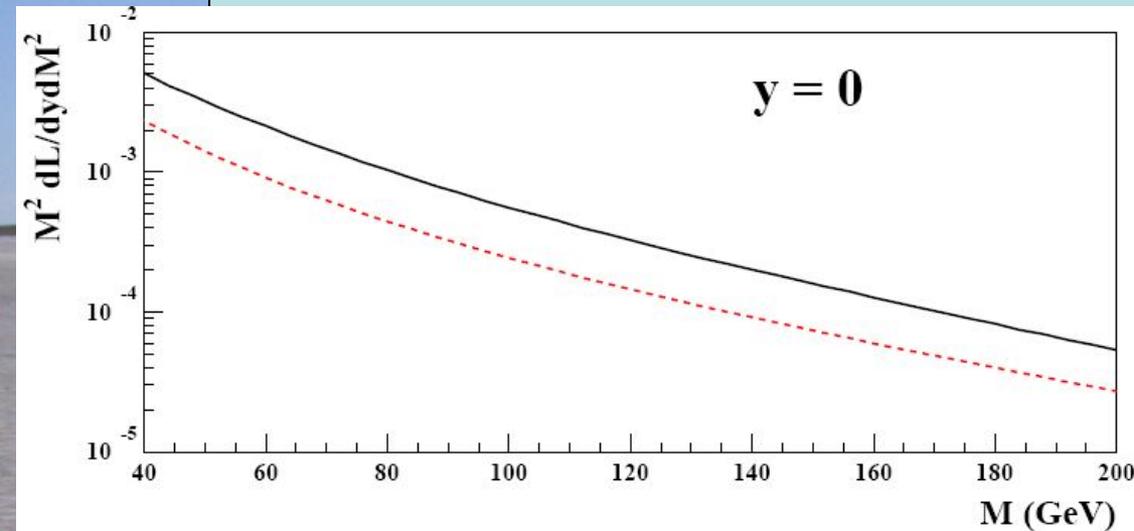
- Gluon collision factor for different PDFs:



Determine using
HERA & LHC data!

Effective Luminosity: Double-Diffractive Higgs production

Cross section in CP-violating scenario: three-way mixing
 $\tan \beta = 50$, $m_{H^\pm} = 155$ GeV

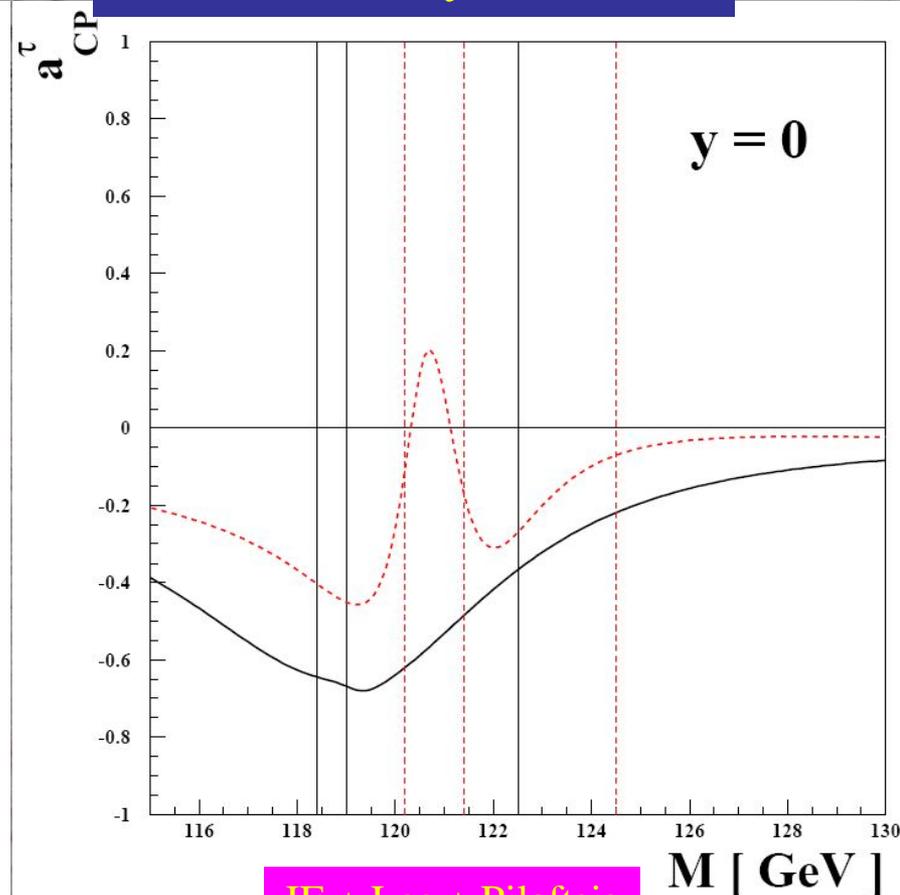
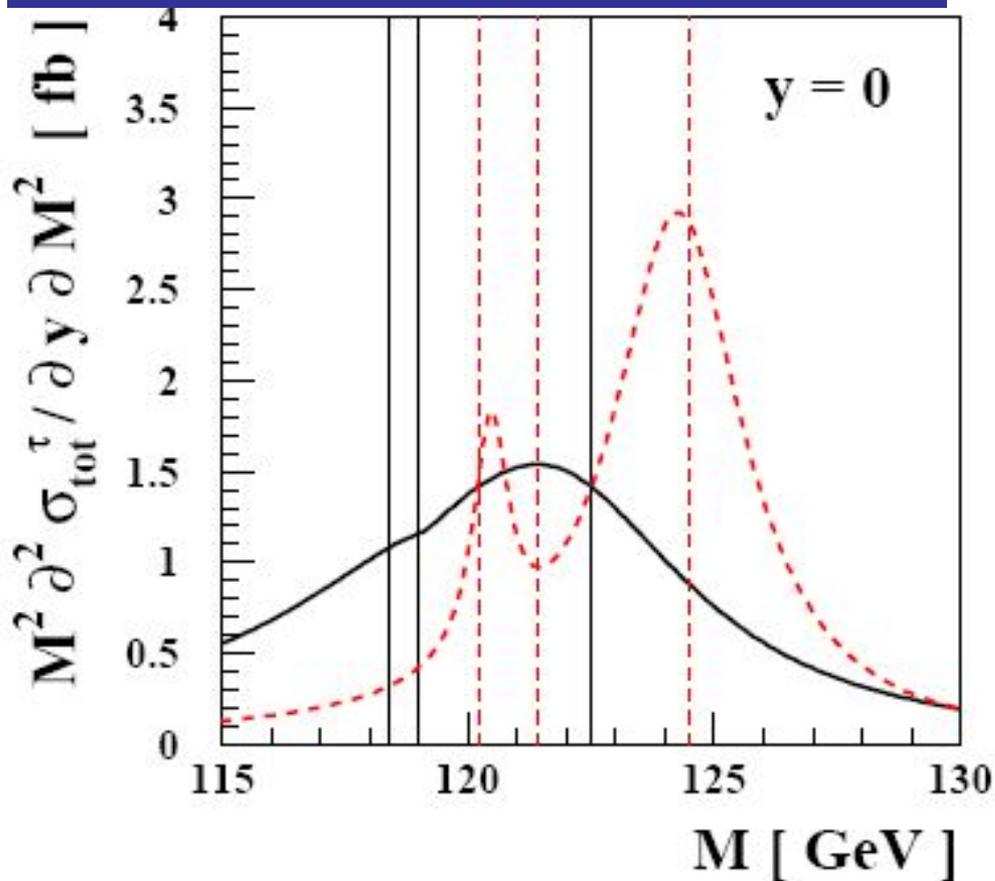


Can hope to measure line-shape using forward proton measurements?

Cross Section, CP-Violating Asymmetry for $H_i \rightarrow \tau^+\tau^-$

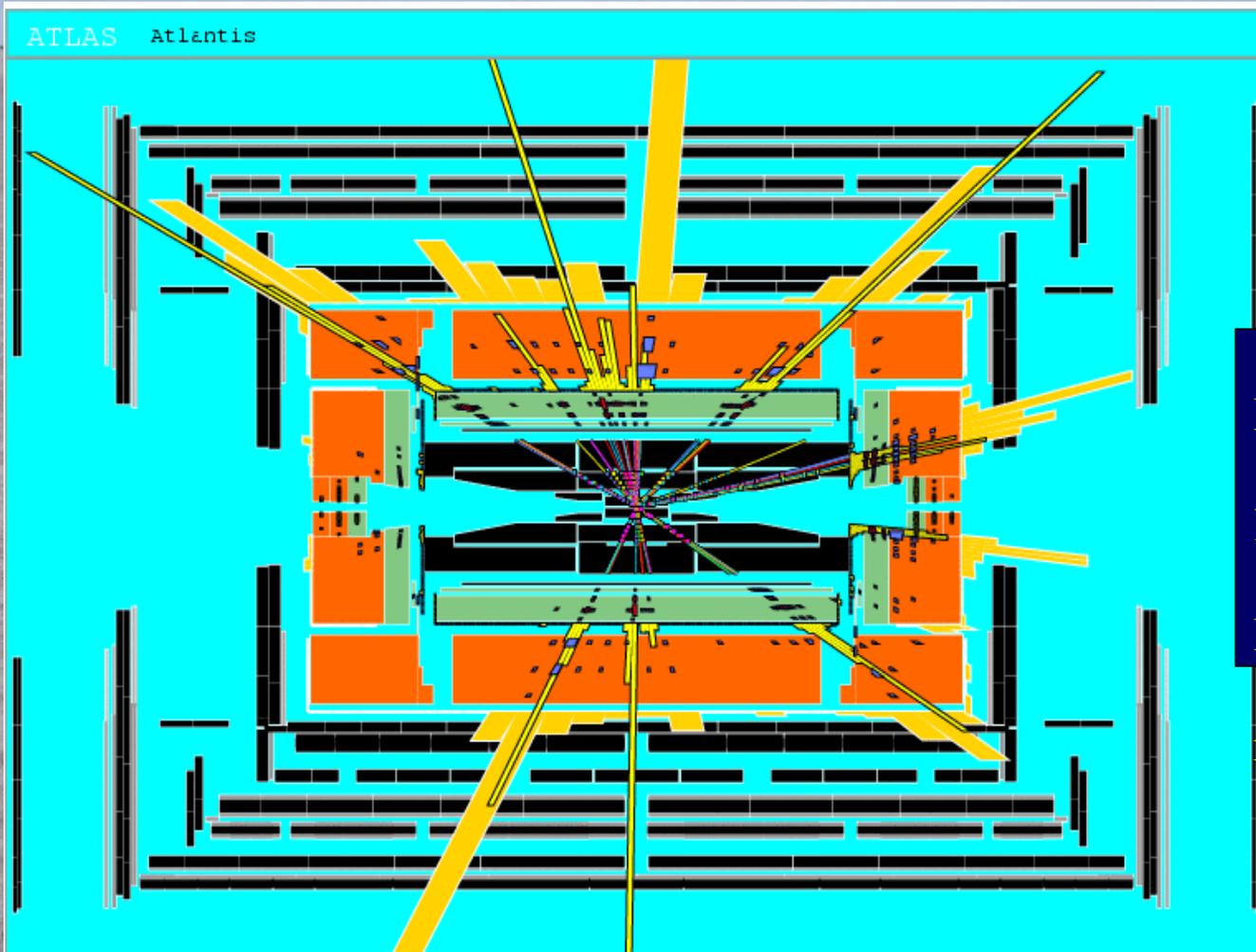
Cross section peaks not at poles ...

... nor are asymmetries



Perhaps none of the above?

Black Hole Production at LHC?



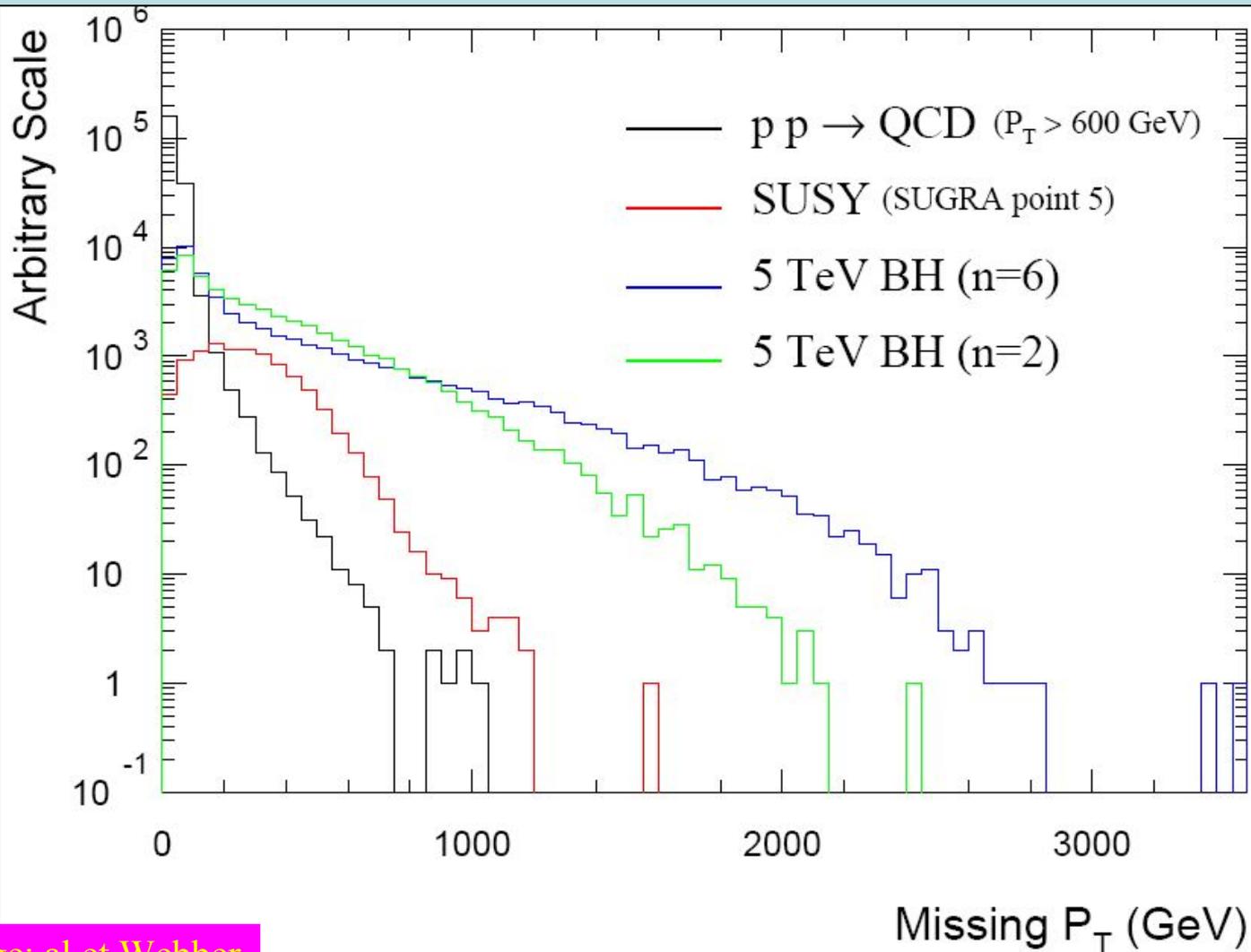
Multiple jets,
leptons from
Hawking

radiation

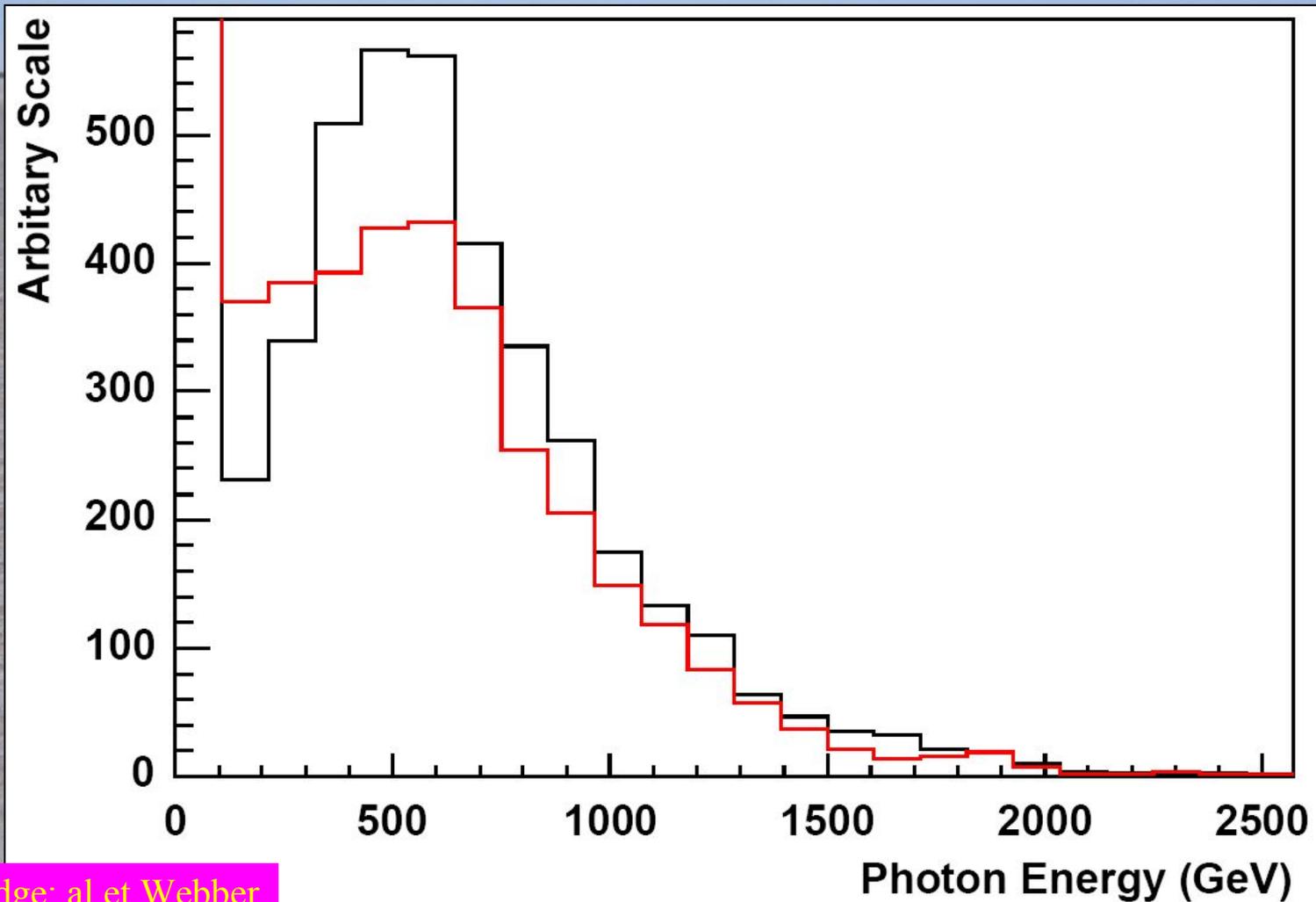
Summary

- We do not know what the LHC will find
- But HERA physics provides crucial inputs:
 - Parton distributions
 - Saturation effects
- Forward physics is potentially exciting area not covered by present detectors
 - Colour glass condensate
 - UHECRs
 - Diffraction Higgs production?

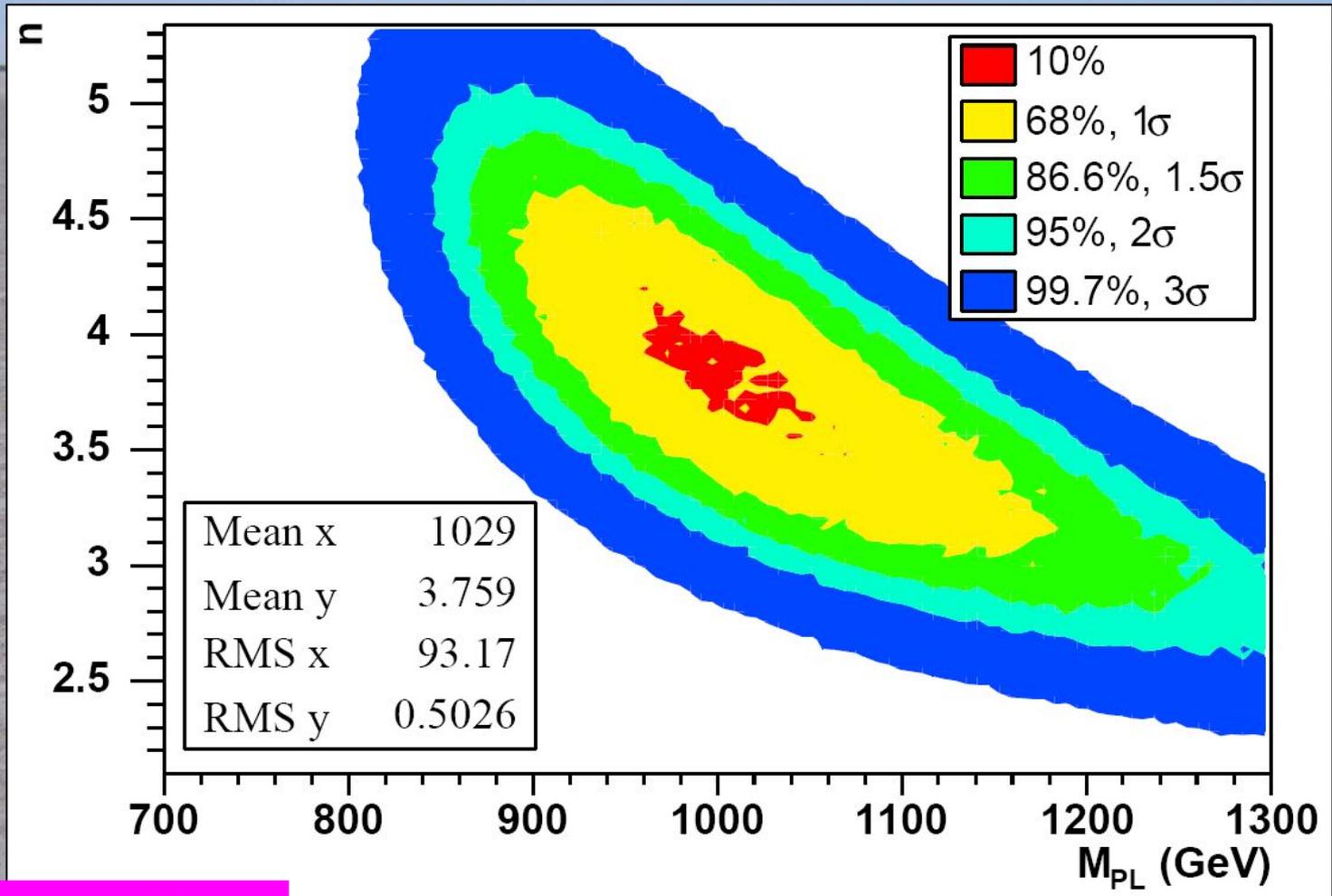
Black Hole Production



Black Hole Decay Spectrum



Measuring Extra Dimensions

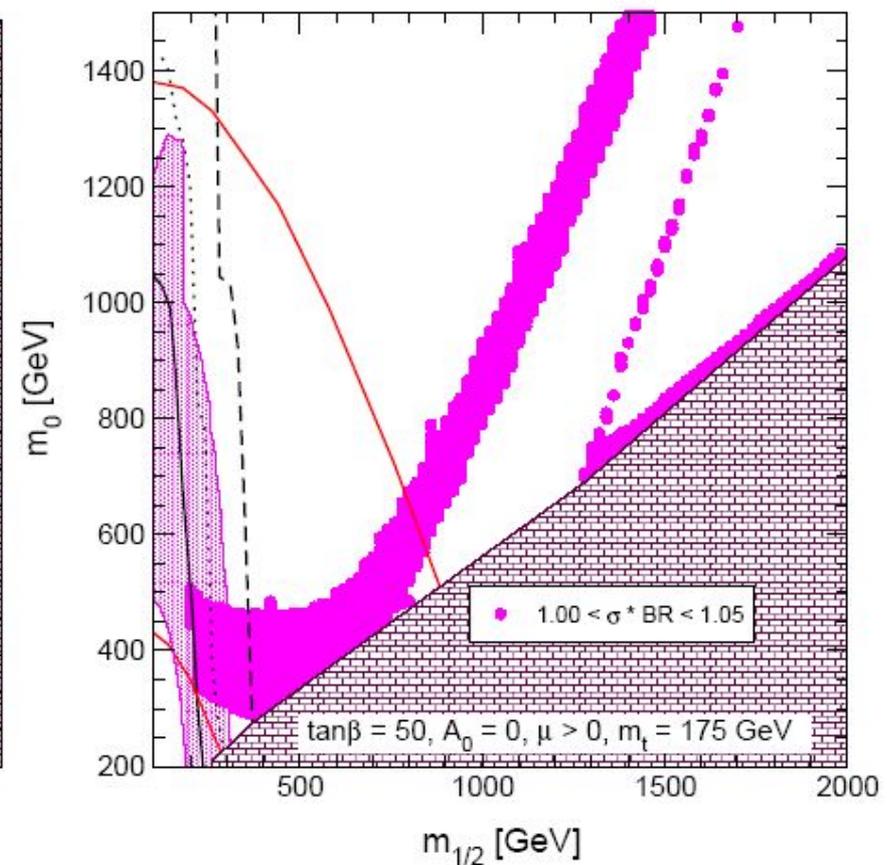
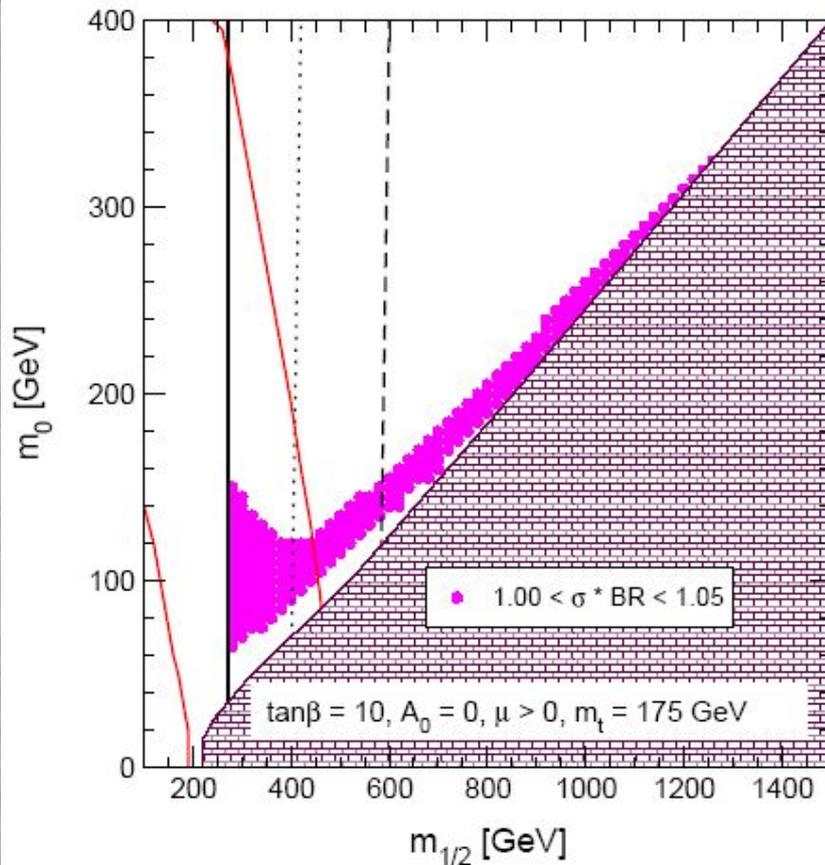


Higgs Production: CMSSM vs SM

$$\frac{[\sigma(W^\pm, Z^0/\bar{t}t + h) \times \mathcal{B}(h \rightarrow \bar{b}b)]_{\text{CMSSM}}}{[\sigma(W^\pm, Z^0/\bar{t}t + h) \times \mathcal{B}(h \rightarrow \bar{b}b)]_{\text{SM}}}$$

(a)

(b)

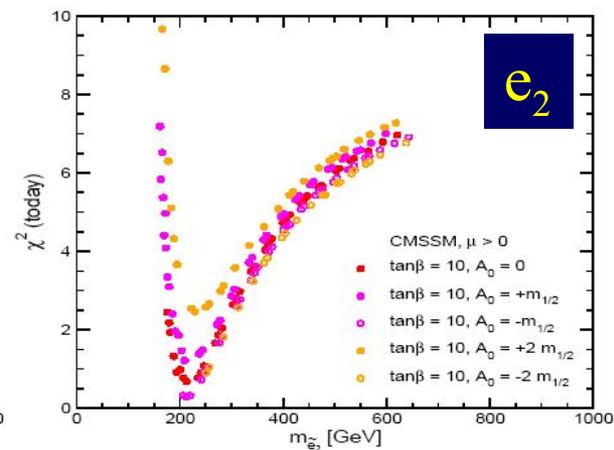
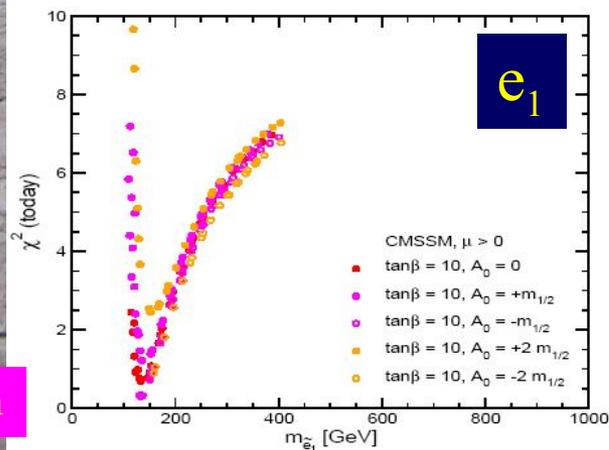
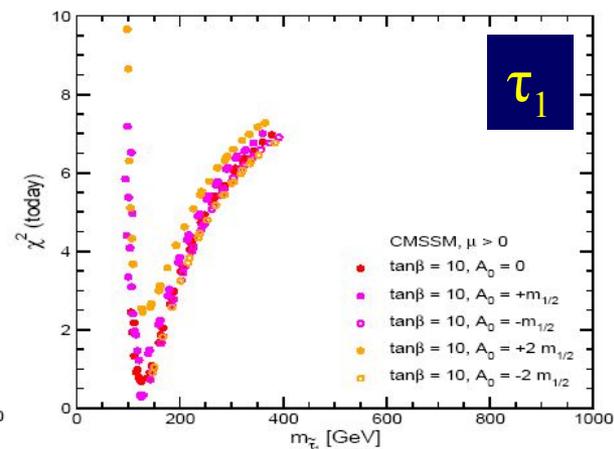
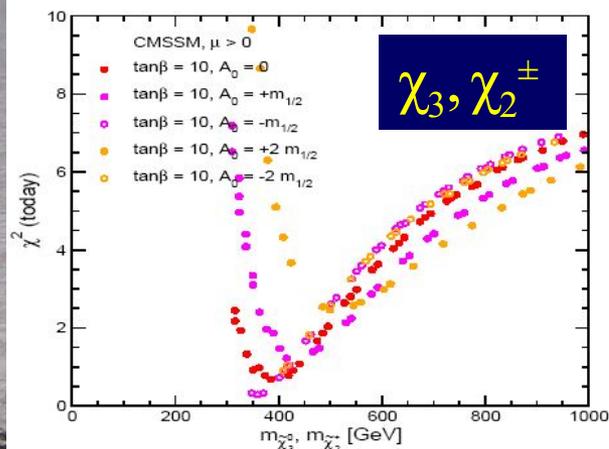
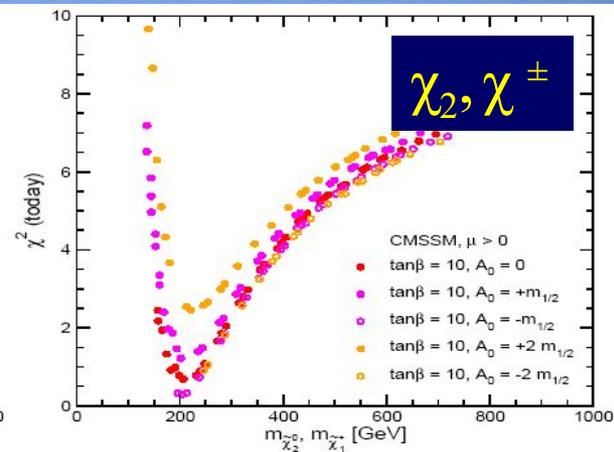
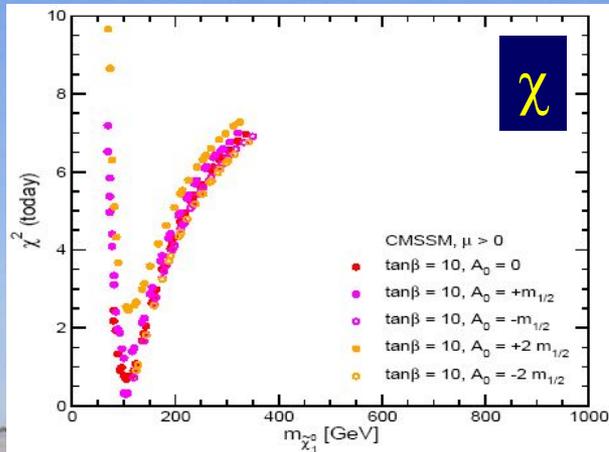


Good news: no suppression – Bad news: cannot distinguish CMSSM

Global Fits to Present Data

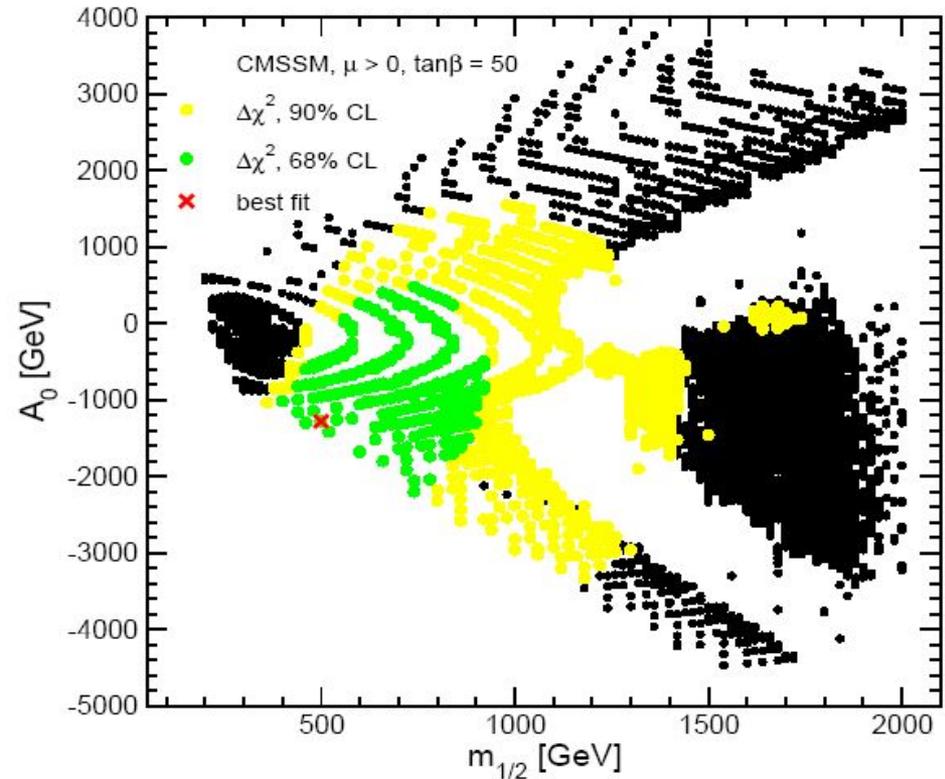
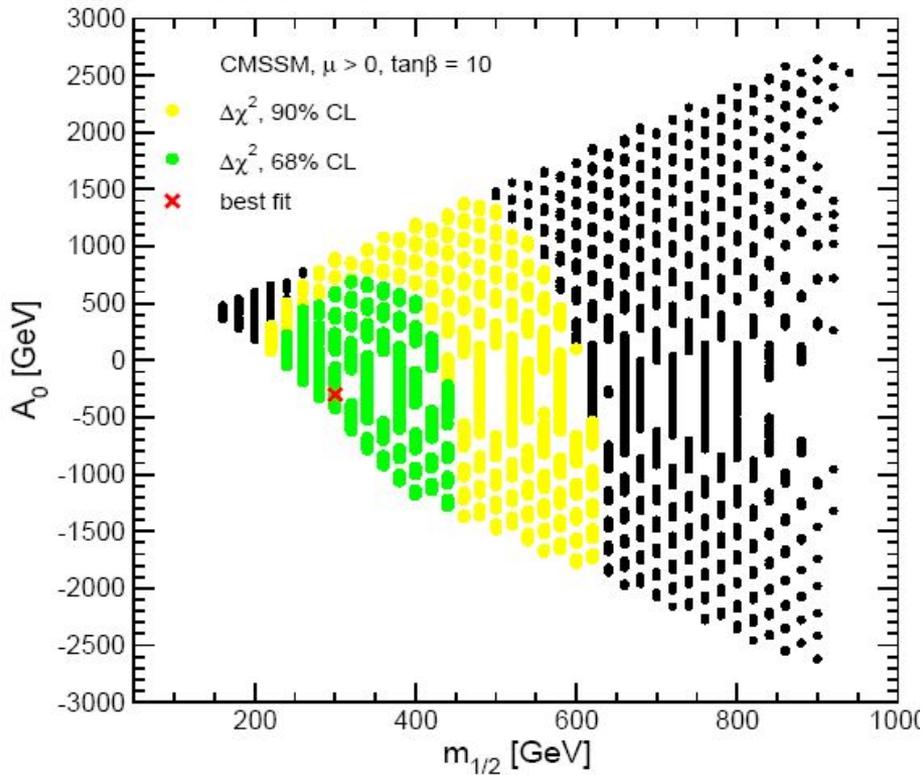
Preferred sparticle masses for $\tan \beta = 10$

JE + Heinemeyer + Olive + Weiglein



Global Fits to Present Data

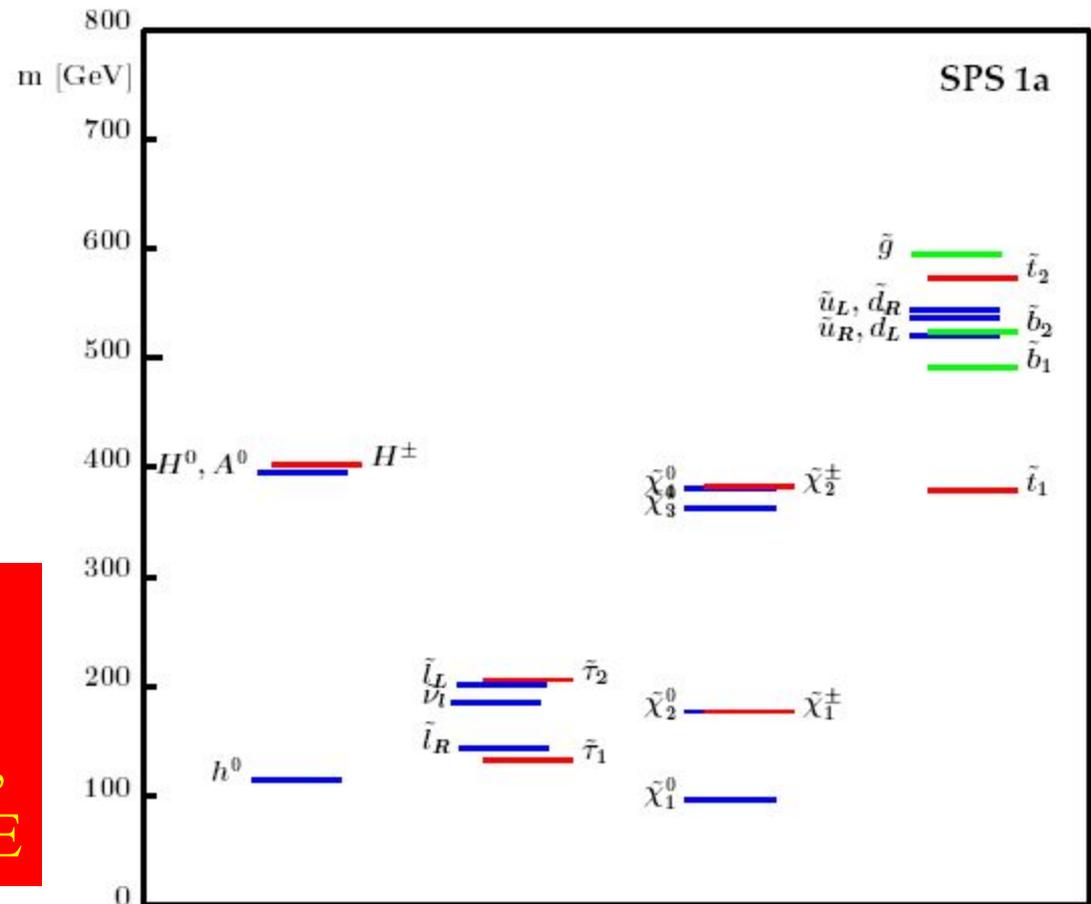
$(m_{1/2}, A_0)$ planes in CMSSM for $\tan\beta = 10, 50$



Example of Benchmark Point

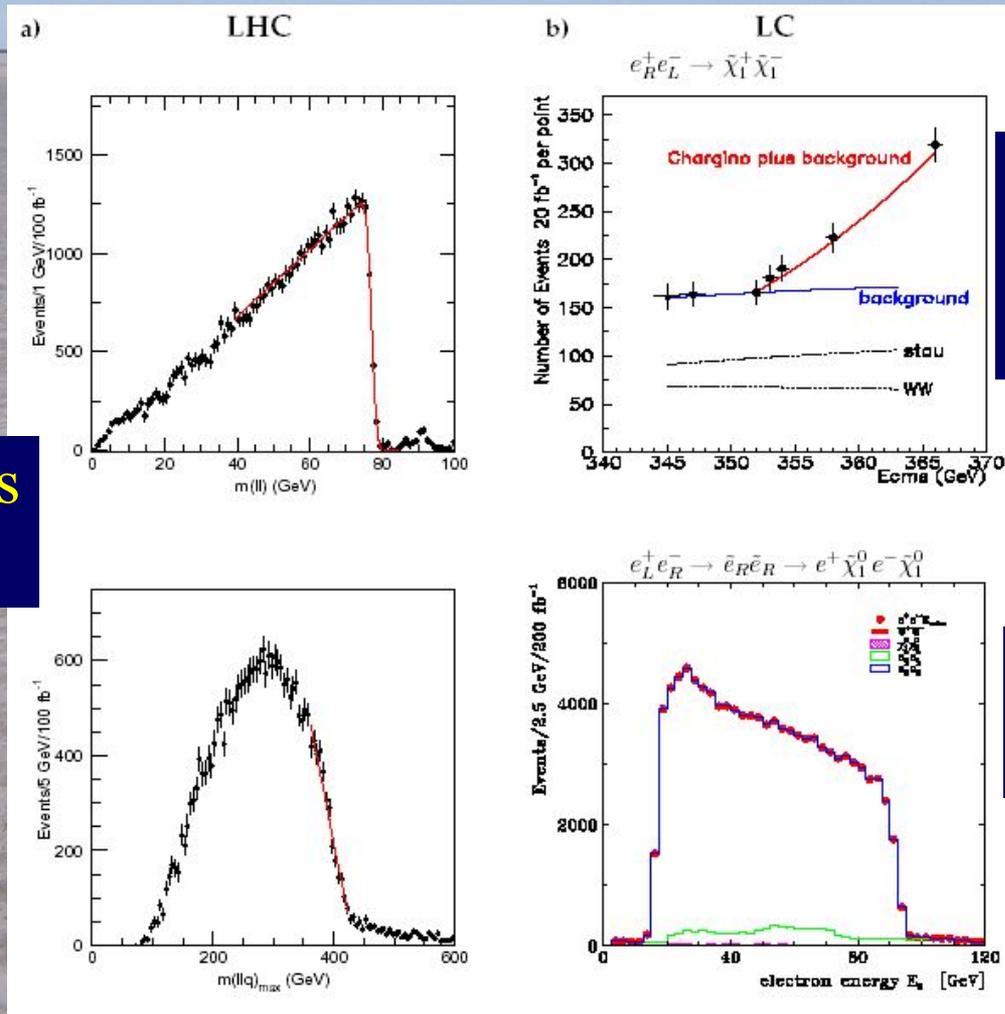
Spectrum of
Benchmark SPS1a
~ Point B of
Battaglia et al

Several sparticles
at 500 GeV LC,
more at 1000 GeV,
some need higher E



Examples of Sparticle Measurements

Spectrum edges
@ LHC



Threshold
excitation
@ LC

Spectra
@ LC

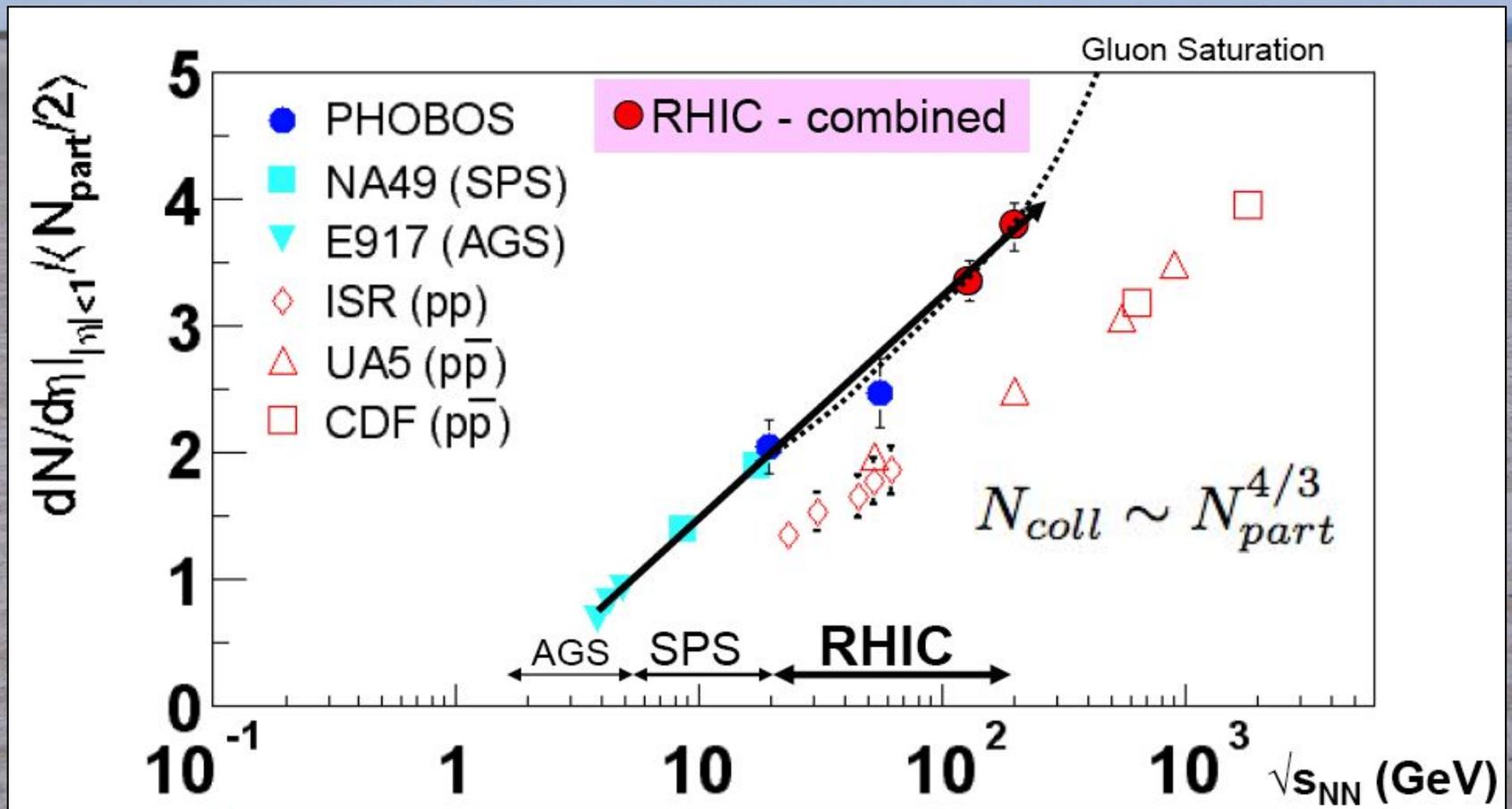
Added Value of LC Measurements

| | m_{SPS1a} | LHC | LC | LHC+LC | | m_{SPS1a} | LHC | LC | LHC+LC |
|---------------|--------------------|------|------|--------|---------------|--------------------|------|-----|--------|
| h | 111.6 | 0.25 | 0.05 | 0.05 | H | 399.6 | | 1.5 | 1.5 |
| A | 399.1 | | 1.5 | 1.5 | H_+ | 407.1 | | 1.5 | 1.5 |
| χ_1^0 | 97.03 | 4.8 | 0.05 | 0.05 | χ_2^0 | 182.9 | 4.7 | 1.2 | 0.08 |
| χ_3^0 | 349.2 | | 4.0 | 4.0 | χ_4^0 | 370.3 | 5.1 | 4.0 | 2.3 |
| χ_1^\pm | 182.3 | | 0.55 | 0.55 | χ_2^\pm | 370.6 | | 3.0 | 3.0 |
| \tilde{g} | 615.7 | 8.0 | | 6.5 | | | | | |
| \tilde{t}_1 | 411.8 | | 2.0 | 2.0 | | | | | |
| \tilde{b}_1 | 520.8 | 7.5 | | 5.7 | \tilde{b}_2 | 550.4 | 7.9 | | 6.2 |
| \tilde{u}_1 | 551.0 | 19.0 | | 16.0 | \tilde{u}_2 | 570.8 | 17.4 | | 9.8 |
| \tilde{d}_1 | 549.9 | 19.0 | | 16.0 | \tilde{d}_2 | 576.4 | 17.4 | | 9.8 |
| \tilde{s}_1 | | | | | | | 7.4 | | 9.8 |
| \tilde{c}_1 | | | | | | | 7.4 | | 9.8 |

Determination of CMSSM parameters

| | SPS1a | StartFit | LHC | Δ_{LHC} | LC | Δ_{LC} | LHC+LC | $\Delta_{\text{LHC+LC}}$ |
|------------------|-------|----------|--------|-----------------------|--------|----------------------|--------|--------------------------|
| \tilde{e}_1 | | | | | | | | |
| M_0 | 100 | 500 | 100.03 | 4.0 | 100.03 | 0.09 | 100.04 | 0.08 |
| $M_{1/2}$ | 250 | 500 | 249.95 | 1.8 | 250.02 | 0.13 | 250.01 | 0.11 |
| $\tilde{\tau}_1$ | 10 | 50 | 9.87 | 1.3 | 9.98 | 0.14 | 9.98 | 0.14 |
| $\tilde{\nu}_e$ | -100 | 0 | -99.29 | 31.8 | -98.26 | 4.43 | -98.25 | 4.13 |

Hadron multiplicities: the effect of parton coherence



Nuclear Modification of Hard Parton Scattering

