



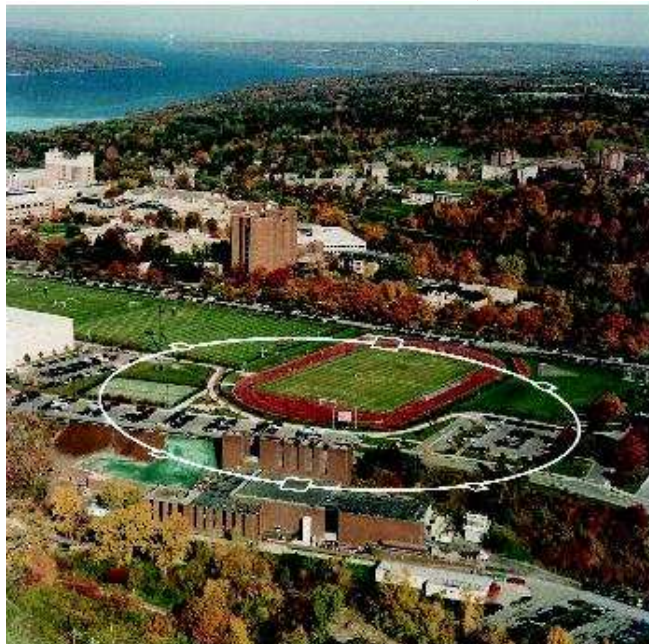
Cornell University
Laboratory for Elementary-Particle Physics

CesrTA Program and Recent Results

M. Palmer

for the CesrTA Collaboration

November 21, 2008





- CEsrTA Program
- CESR Reconfiguration & Startup
- EC Data from Commissioning Run
- EC Data-Simulation Comparisons
- Summary and Comments



- **CesrTA is an ILC Damping Rings R&D project**
 - Joint NSF/DOE funding
 - Funding spans FY08-FY10
 - Funding levels consistent with a 2 year experimental program
- **Project Goals**
 - Characterize electron cloud build-up
 - Develop electron cloud suppression techniques
 - Develop modeling tools for electron cloud instabilities
 - Determine electron cloud instability thresholds
 - Conduct experiments in an ultra-low emittance positron ring
 - Provide significant amounts of dedicated experimental time (~240 running days over course of program)



CesrTA Program II

R&D Targets:

- 2008 through mid-2009
 - Machine reconfiguration for ultra low emittance operation
 - Deploy and commission instrumentation needed to characterize low emittance beams
 - Deployment of EC diagnostics and studies of EC growth and suppression in wigglers, dipoles, quadrupoles and drift regions
 - Implement ultra low emittance lattice and begin low emittance tuning program
- Mid-2009 through April 2010
 - Work to achieve ultra low emittance operation
 - Ongoing EC growth and mitigation studies
 - EC beam dynamics studies at the lowest achievable emittances
 - Focus shifts much more heavily to experiment versus machine modifications
- Provide evaluations for the ILC Technical Design Phase in mid-2010

Low Emittance Parameters

Parameter	Value
E^\dagger	2.0 GeV
N_{wiggler}	12
B_{max}	1.9 T
ϵ_x (geometric)	2.5 nm
ϵ_y (geometric) Target	5-10 pm → 20pm
$\tau_{x,y}$	56 ms
σ_E/E	8.1×10^{-4}
Q_z	0.070
Total RF Voltage	7.6 MV
σ_z	8.9 mm
α_p	6.2×10^{-3}
$N_{\text{particles/bunch}}$	2×10^{10}
τ_{Touschek}	10s of minutes
Bunch Spacing	Multiples of 4ns and 14ns

\dagger Operating range of 1.5 to 5.3 GeV



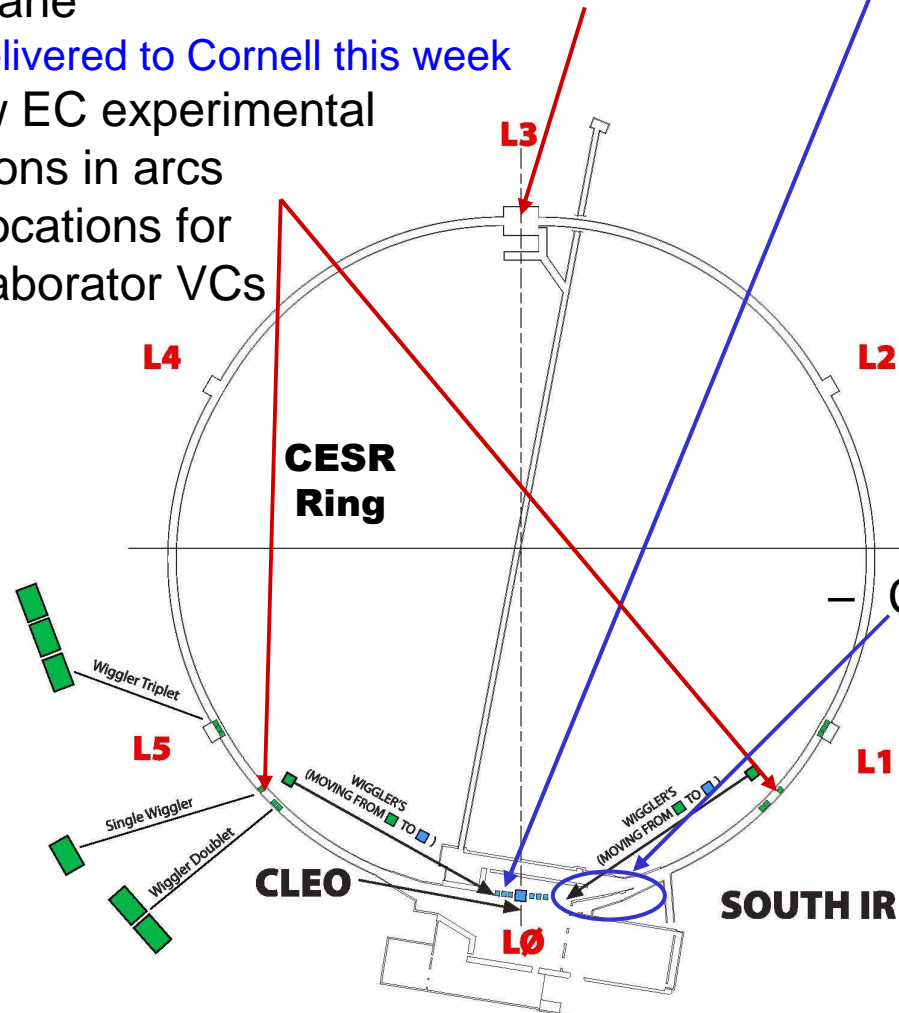
CESR Reconfiguration

- **Electron Cloud Diagnostics**

- L3 region prepped for arrival of PEP-II EC hardware including diagnostic chicane

Delivered to Cornell this week

- New EC experimental regions in arcs w/ locations for collaborator VCs



- L0 region reconfigured as a wiggler straight

Instrumented with EC diagnostics
Wiggler chambers with retarding field analyzers (fabricated at LBNL) - scheduled for installation ~Oct 23rd
Chambers with EC mitigation (TiN coatings by SLAC)

- CHES D-line

Windowless (all vacuum) x-ray line upgrade for x-ray detector capable of single pass measurements
Addition of x-ray optics box
Detector setup in CHES hutch
Option for longer optical lever arm by extending line to detector located in L0

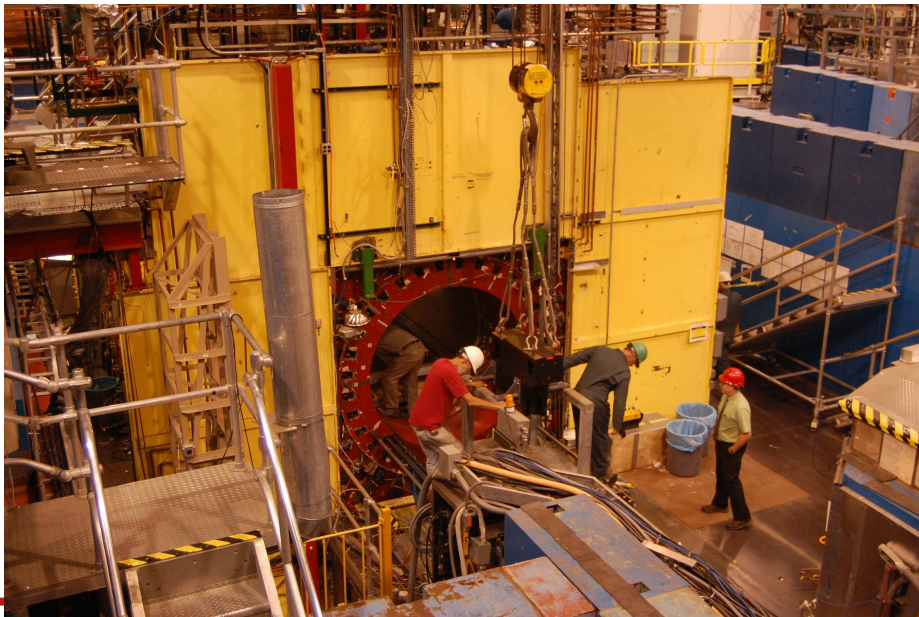
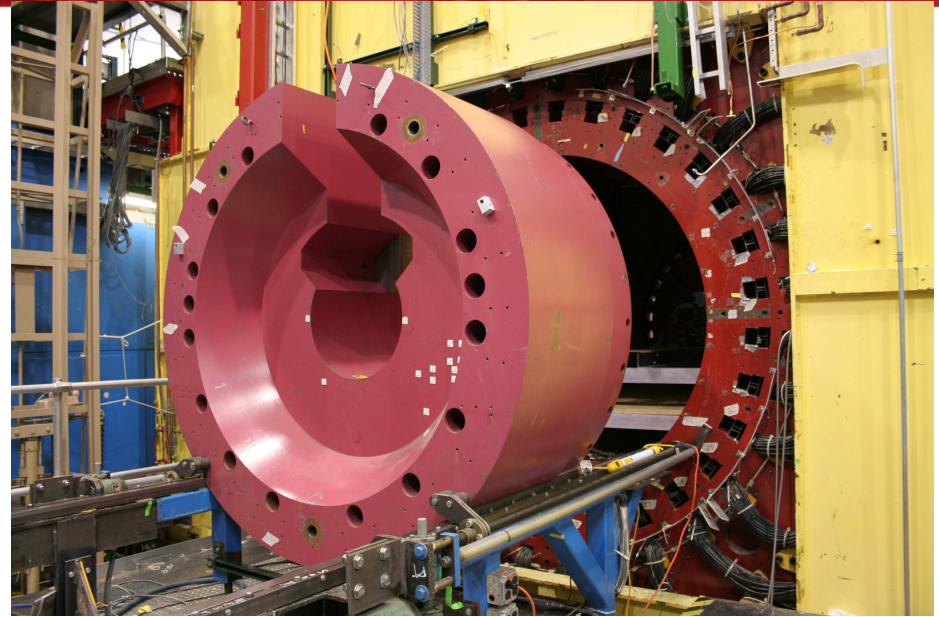
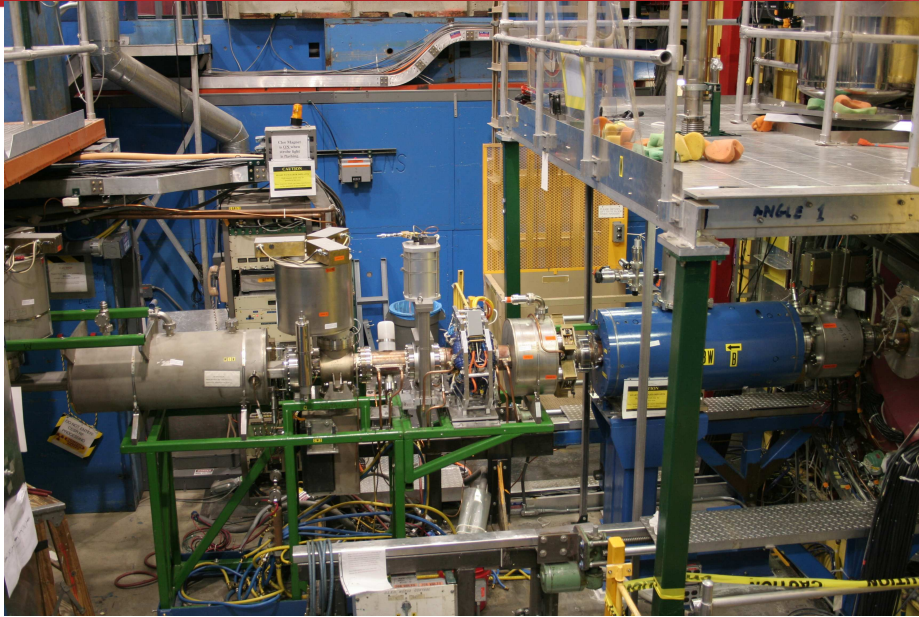


CESR Startup in DR Configuration

- Reconfiguration down started July 7 and concluded in early October. Highlights:
 - CESR vacuum closed week of Sept 29th
 - Linac/Synchrotron startup commenced Oct 2nd
 - First injection into CESR took place Oct 8th
 - Stored beam in CESR early Oct 13th
 - Magnet coil problem
 - Diagnostic wigglers installed Oct 23-24
 - Operation at 5.3 GeV through morning of Nov 3
 - Machine processing
 - CesrTA machine studies (particularly hardware checkout)
 - 2.0 GeV CesrTA Experiments : Nov 3-Nov 10
 - Electron Cloud Experiments
 - Low Emittance Optics
 - X-ray Beam Size Monitor
 - Instrumentation Checks

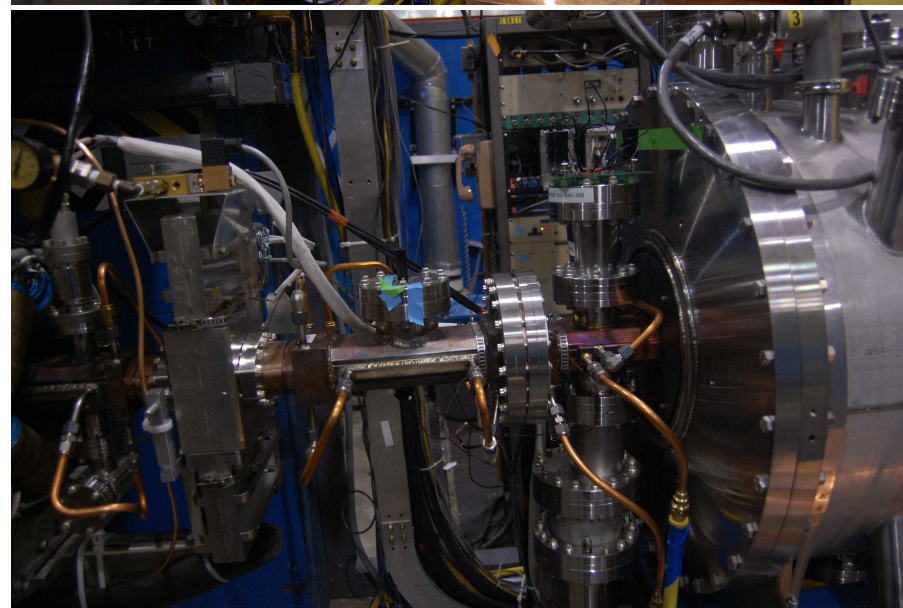
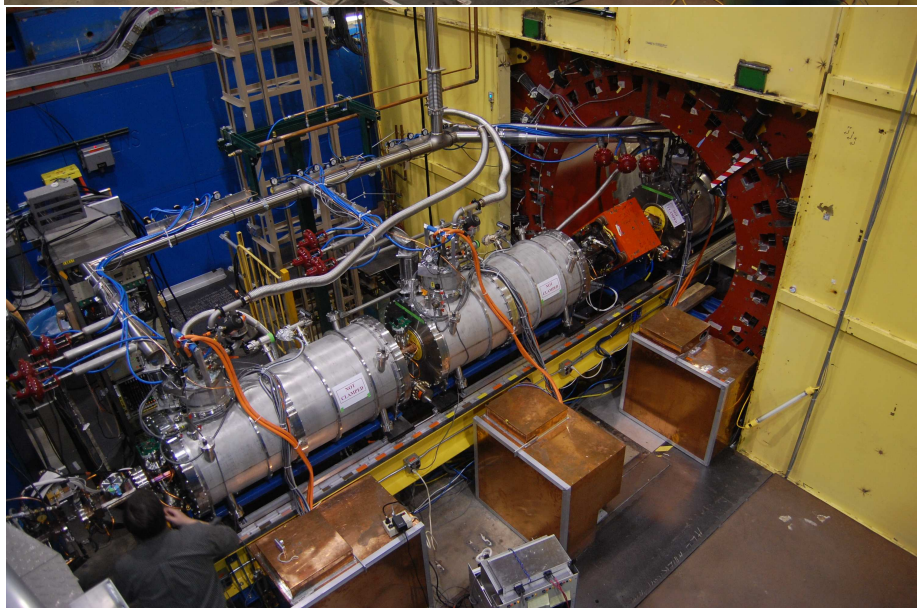
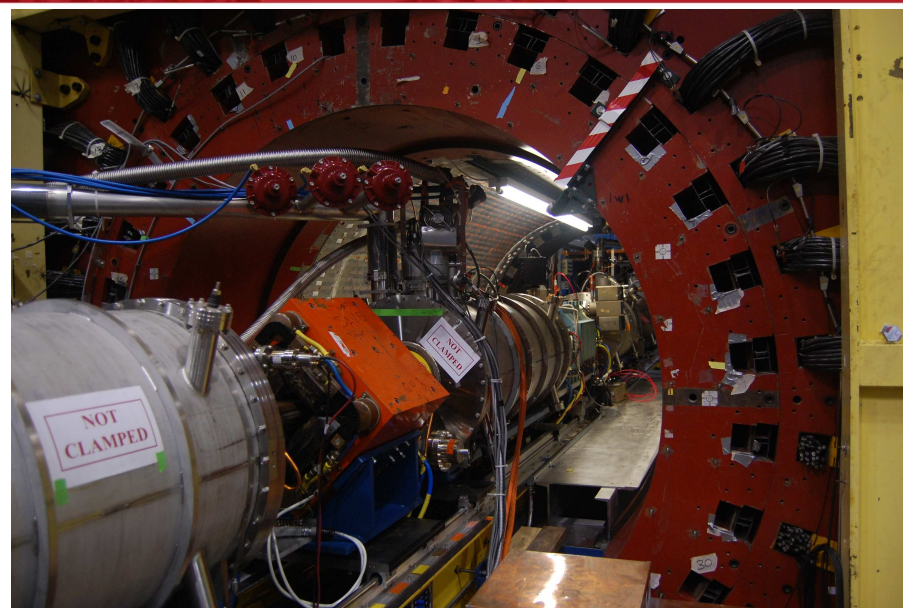
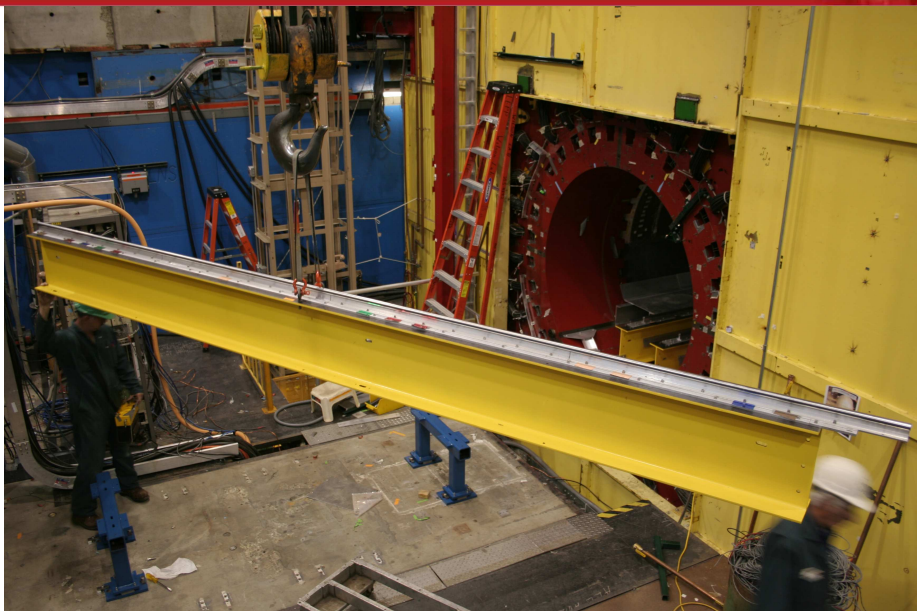


L0 Modifications





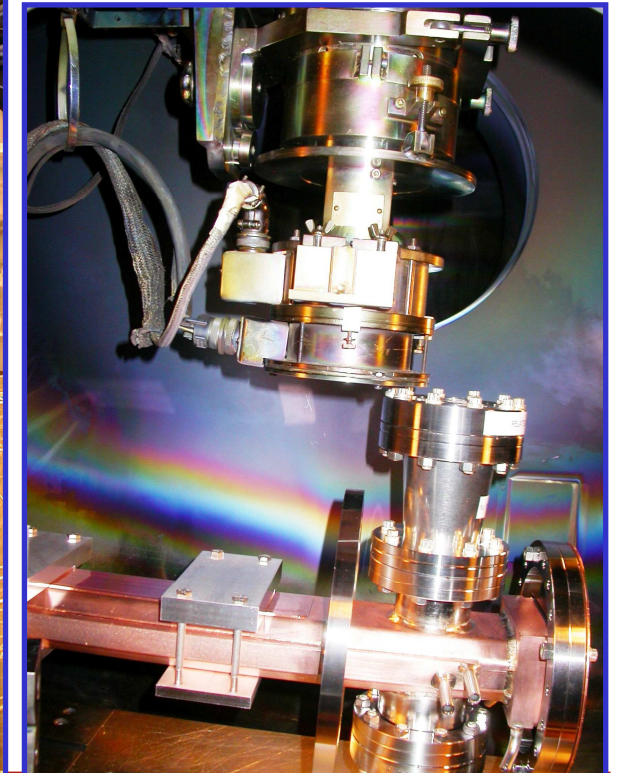
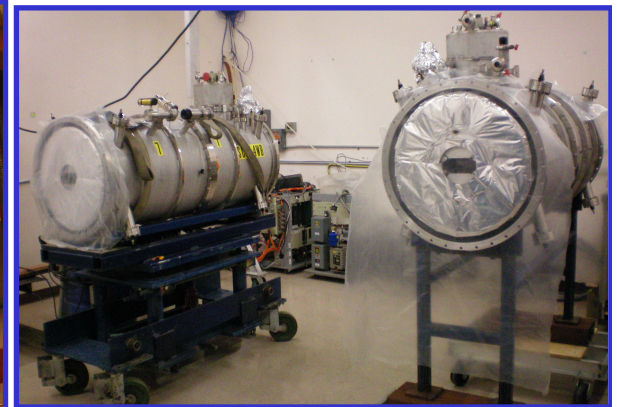
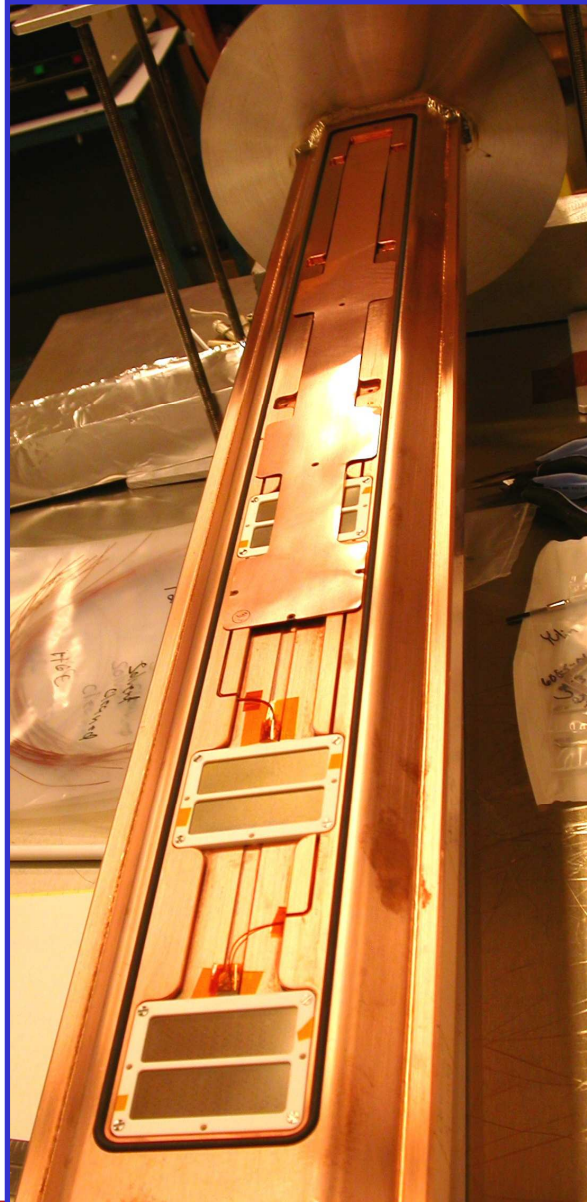
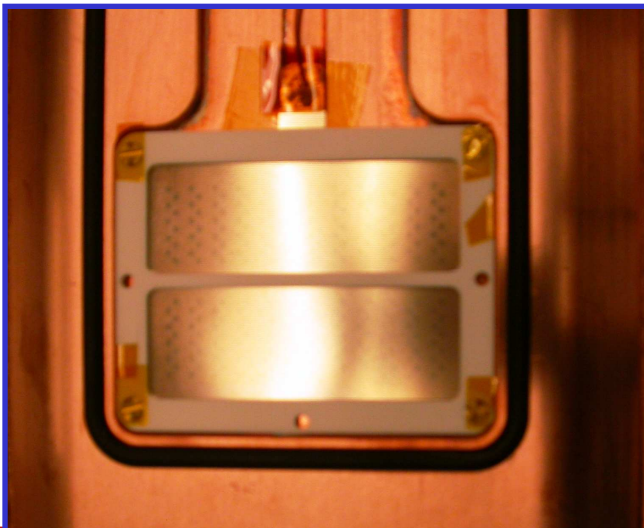
L0 Modifications





Wiggler EC Diagnostics

- RFAs assembled and checked for both VCs
- E-beam welding
 - 1st VC complete
 - 2nd VC on Oct 10
- Installed into cryostats week of Oct 14
- Installed into CESR Oct 23-24



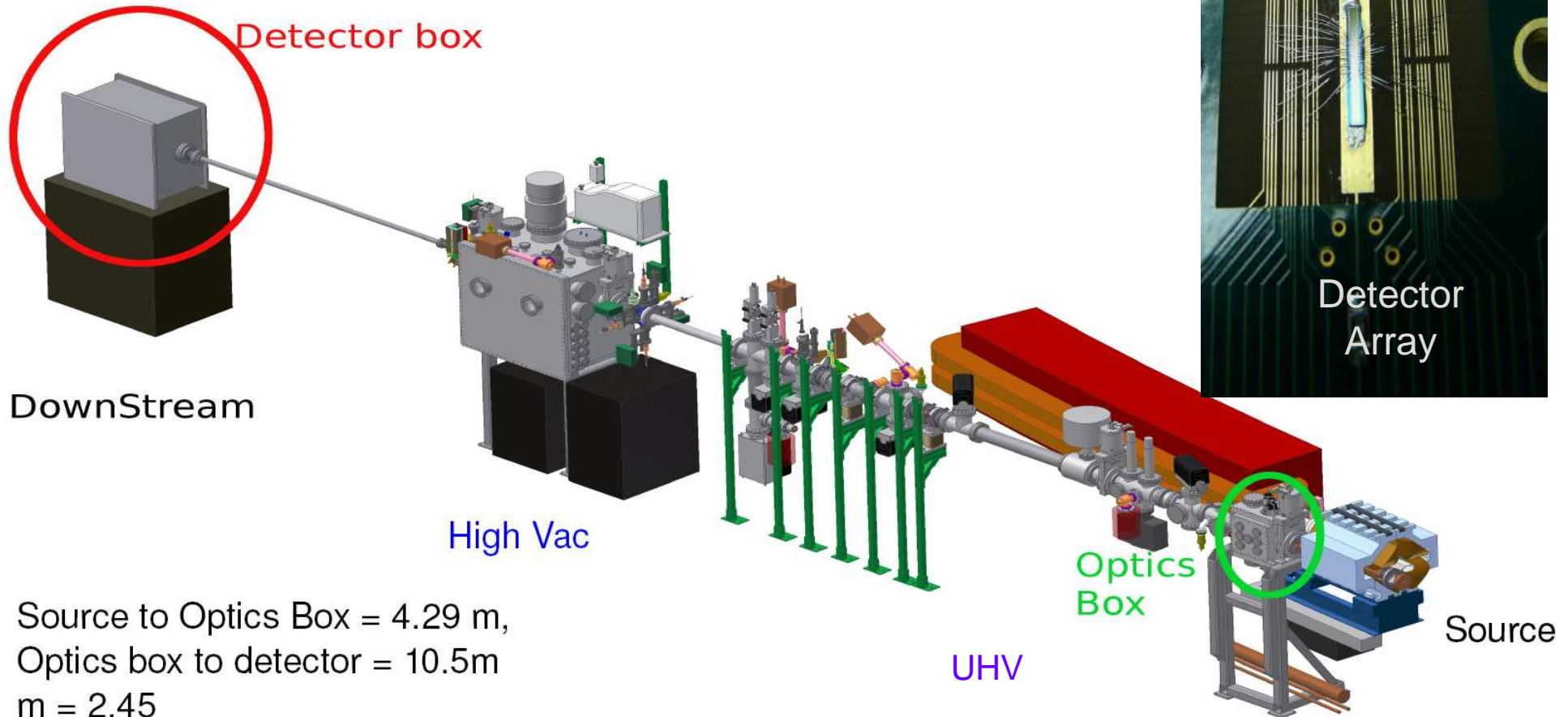


- **X-ray beam size monitor**
 - Upgrade of the positron x-ray beam line is complete (windowless line)
 - 4 μ m diamond window to isolate storage ring and beam line vacuum will be tested with a beryllium backup in January run
 - X-ray optics
 - “Simple” optics (adjustable slit and 3 slit coded aperture) for Oct-Nov 2008
 - Optical Array Element from Applied Nanotechnologies arrives in December
 - 230 ring fresnel zone plate and 41 element coded aperture on a single substrate
 - X-ray detectors
 - Detectors (from 3 different vendors) tested in Oct-Nov
- **X-ray beam size measurement**
 - Measurement of the size of the positron beam with few micron resolution beginning in January 2009 run
 - Single bunch/single pass measurements in May-June 2009 run



New optics line in collaboration with CHESS: All vacuum line

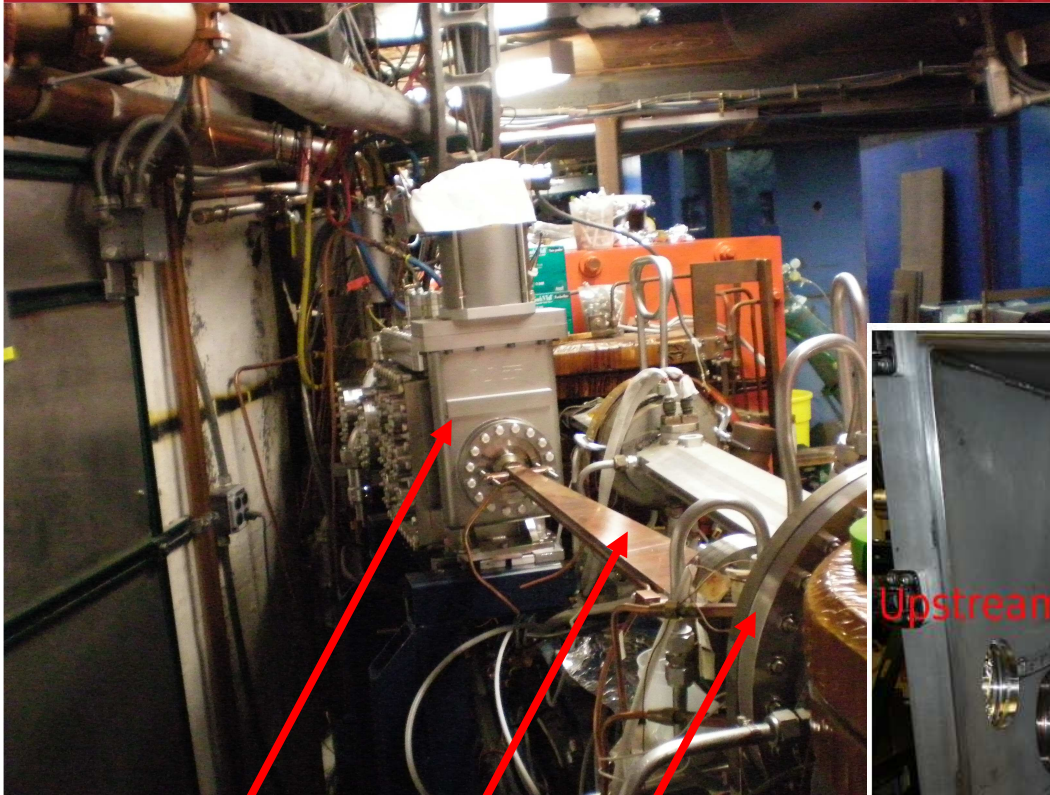
Helium or Vacuum



Source to Optics Box = 4.29 m,
Optics box to detector = 10.5m
m = 2.45



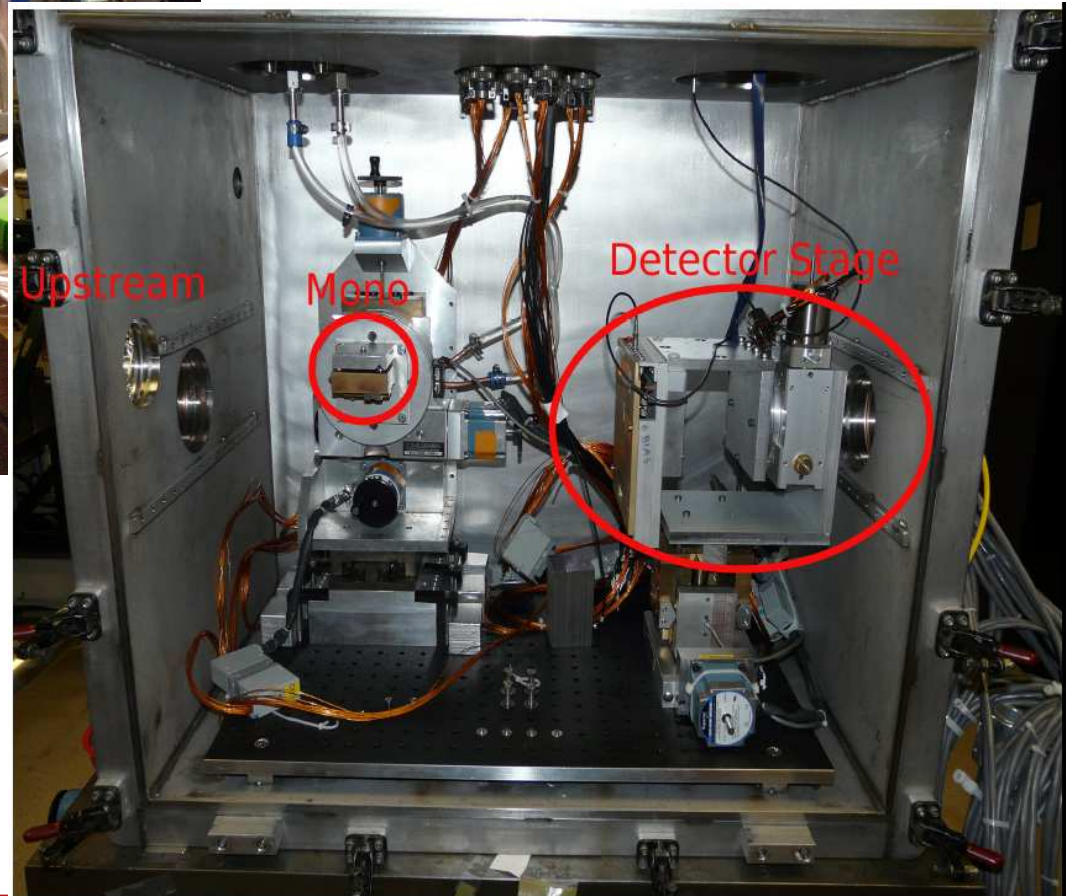
xBSM Optics Box



Optics Box

Cu Flare Chamber

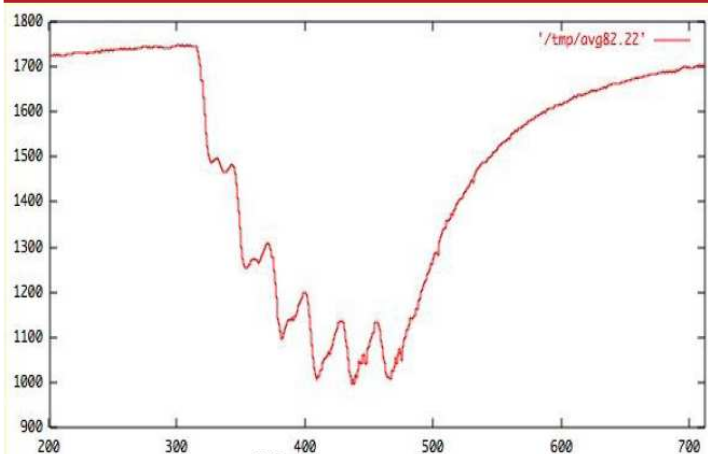
CHES D-line Crotch



Upstream

Mono

Detector Stage



Hamamatsu
100 turn average

Single-pass measurement capability:

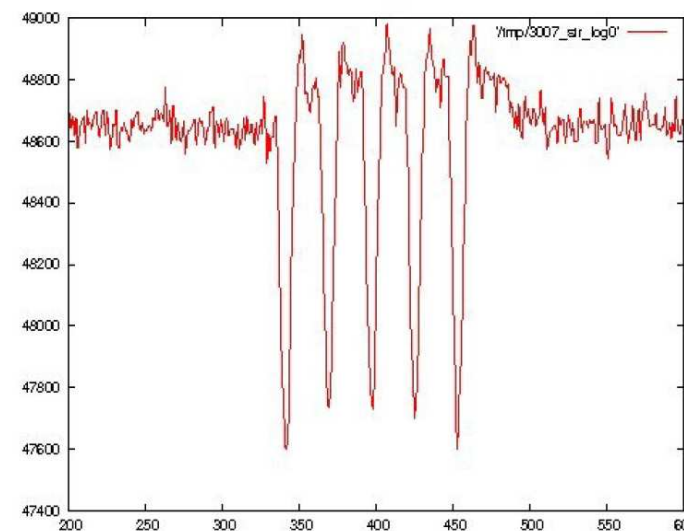
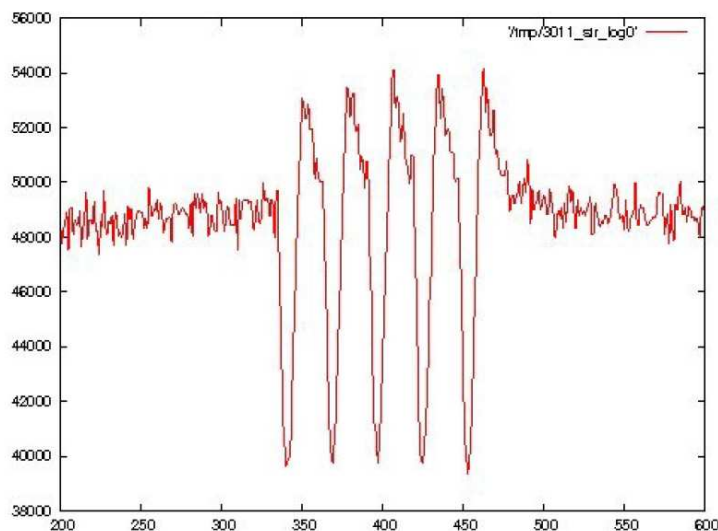
After previously discovering very slow recovery times with Hamamatsu diode arrays (fabrication issue – not fundamental), it was critical to test the response of alternate detector arrays. Excellent time response found from 2 suppliers!

Kyosemi

Amplitudes are average over
100 turns

Fermionics

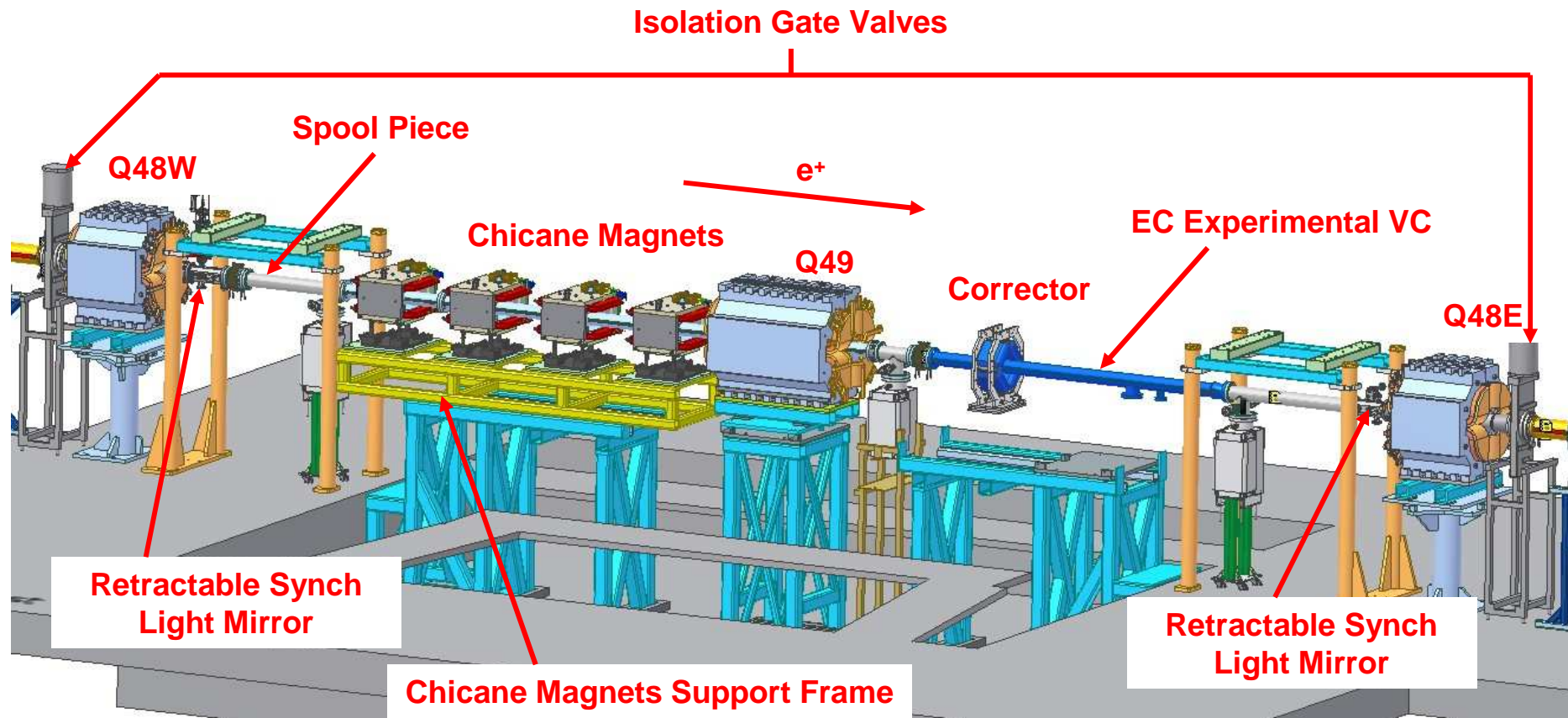
14 ns spaced
bunches





L3 Preparations

- In L3, removed vertical separators and installed gate valves to isolate experimental region
- PEP-II EC Hardware to arrive at Cornell by end of this week



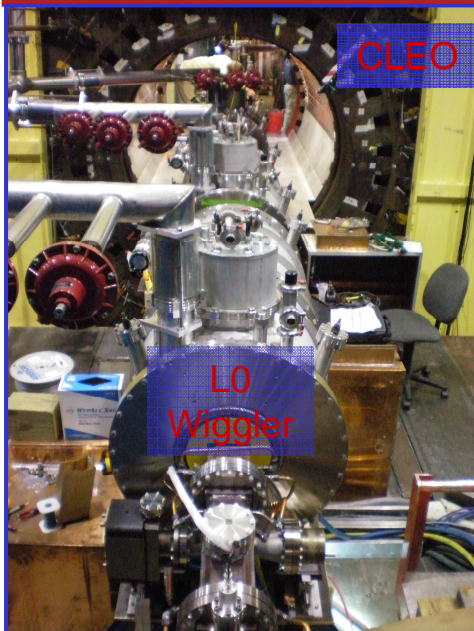
L3 Region Isometric View for February 2009



- **Short 2 week commissioning/machine processing run**
 - 1st week at 5.3 GeV (hardware commissioning)
 - 2nd week at 2.0 GeV (wigglers on/low emittance lattices)
- **Key Goals**
 - Electron Cloud Studies
 - Deploy and commission wiggler RFAs, carry out a minimal set of RFA experimental measurements
 - Study of systematic effects in multi-bunch tune measurements
 - Set up TE wave measurements in new wiggler straight
 - Optics and LET
 - Commission new ring layout and optics
 - Begin work on optics correction
 - Instrumentation and Feedback
 - X-Ray Beam Size Monitor (xBSM)
 - Commissioning of upgraded CHESS D-line, x-ray optics, and detectors
 - Beam Position Monitor System Tests
 - Feedback System (4ns)
 - Test upgraded transverse feedback system
 - Longitudinal instability threshold measurements and upgrade specifications
- **All tasks time-limited**
 - Initial checks only
 - Provide guidance for longer dedicated run in January 2009
- **A few highlights are shown on the following pages...**

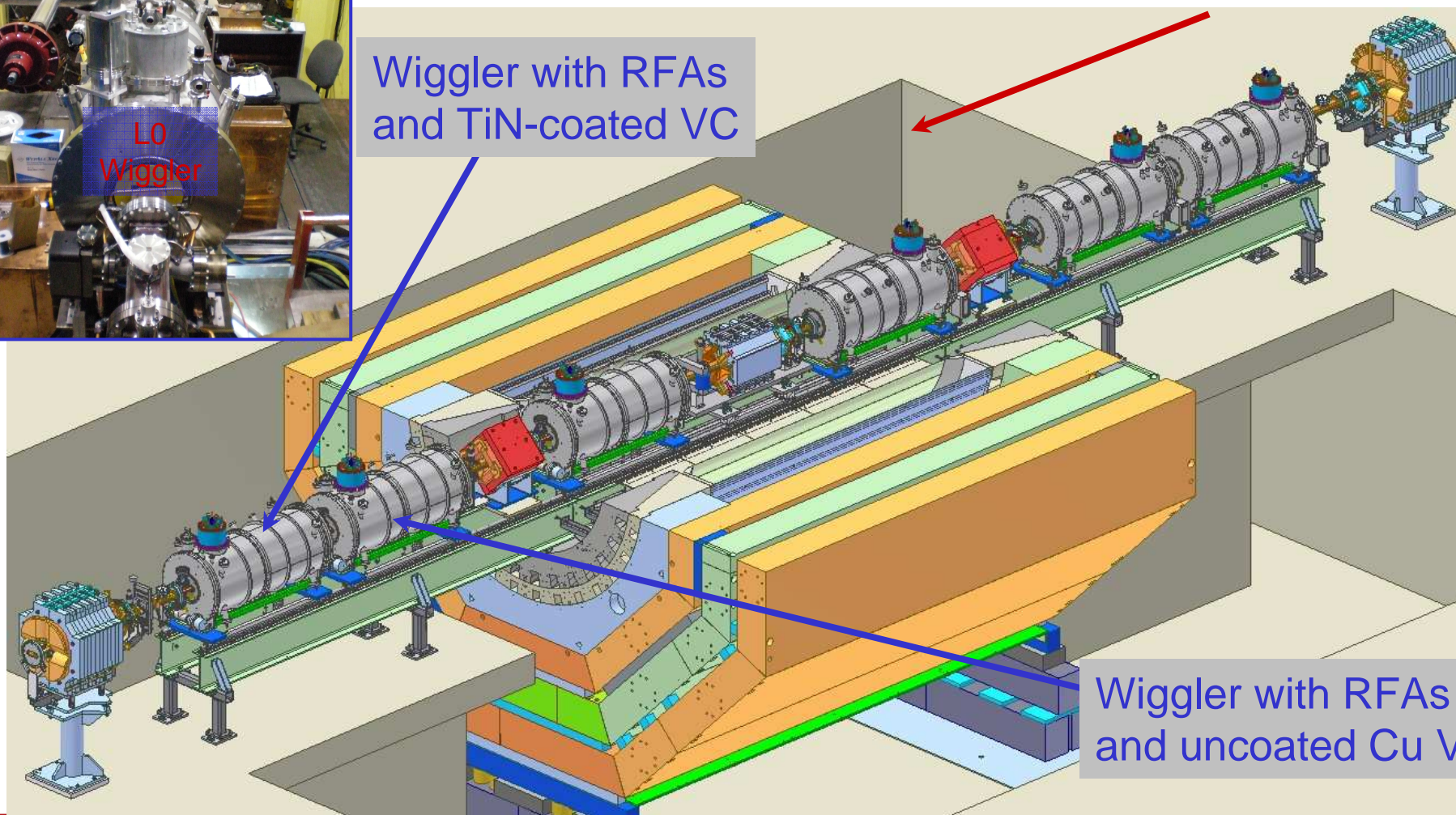


Wiggler Straight



- Zero dispersion straight for low emittance
- Wiggler experimental region – local EC measurements
 - Retarding Field Analyzers
 - TE Wave Transmission Experiments

Wiggler with RFAs
and TiN-coated VC



Wiggler with RFAs
and uncoated Cu VC

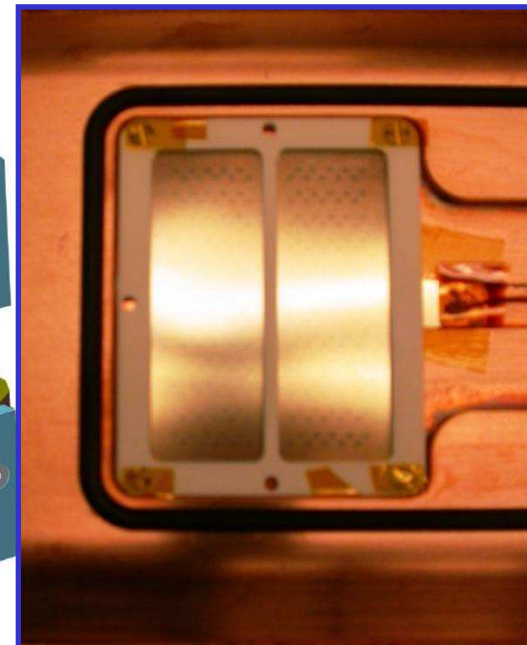
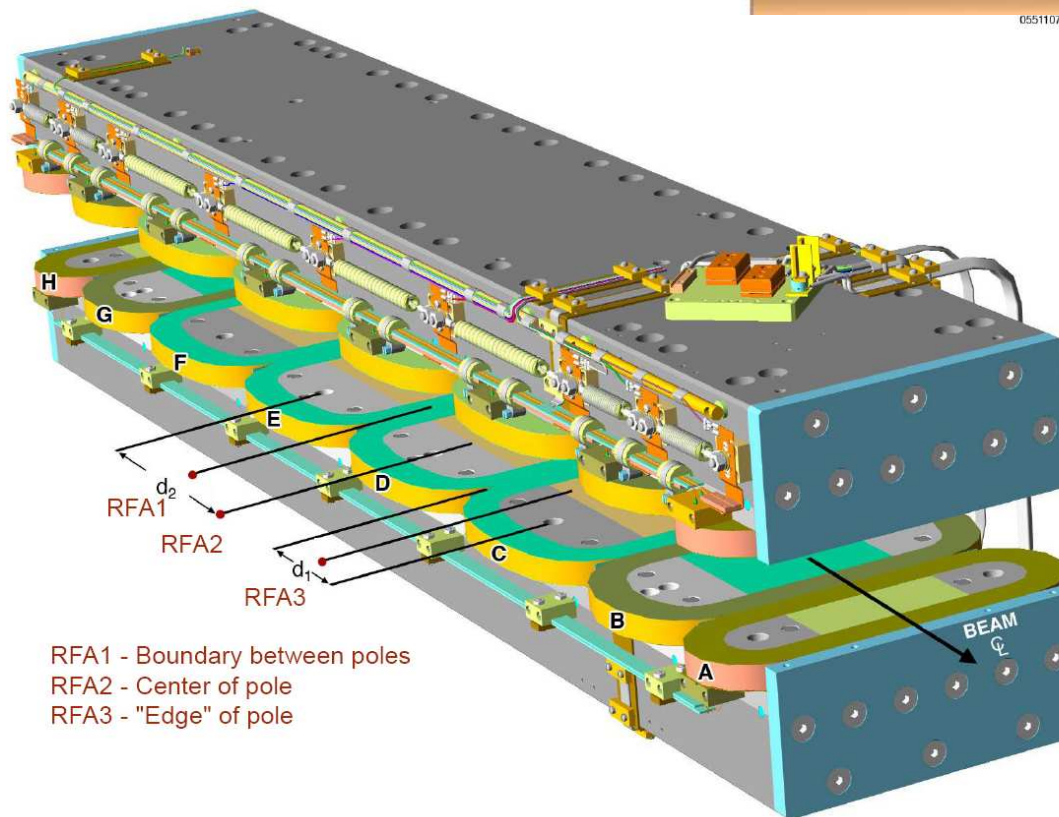
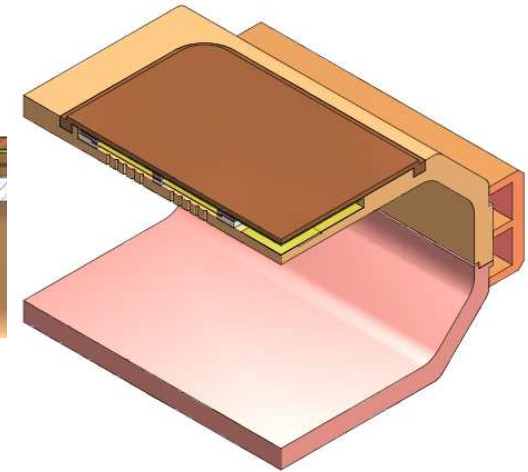
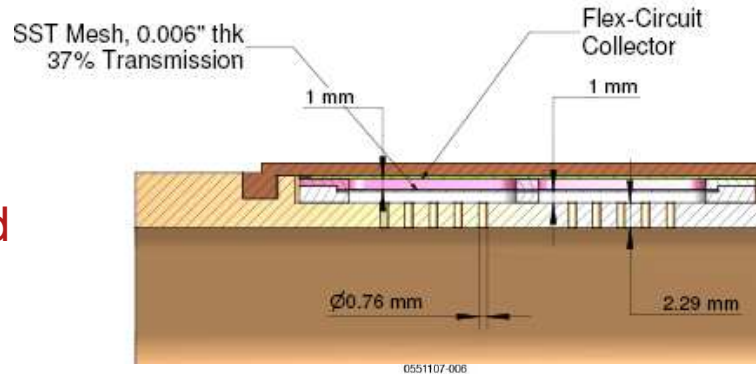


- A sequence of low energy optics was employed
 - Low energy “startup-optics”
 - $Q_x=10.57$, 2.085GeV
 - Arc quads same as June 08 low energy operation (differences only in L0 and L3)
 - Wiggler field 1.9T
 - High tune - intermediate emittance optics
 - $Q_h=14.57$, same L0 quads as “start-up”, 2.085GeV
 - High tune - low emittance (2.6nm)
 - Tune up electron and positron injection.
 - Begin process of correcting betatron phase, coupling, dispersion
 - Create and tune orbit bumps to direct light onto xBSM
 - Good injection efficiency indicates that dynamic aperture is adequate
- Commissioning emphasis during run provided time for loading optics and establishing initial conditions but essentially no time for low emittance tune-up.



Wiggler RFAs

- RFA chamber during assembly and locations of detectors in superferric wiggler. 3 RFAs in each vacuum chamber at different field locations. (CU/KEK/LBNL/SLAC)



12 collectors
across top of
vacuum
chamber

1 retarding grid
spans the 12
collectors

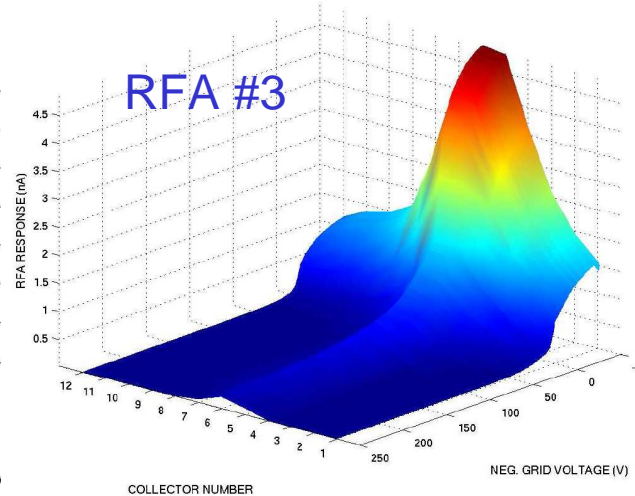
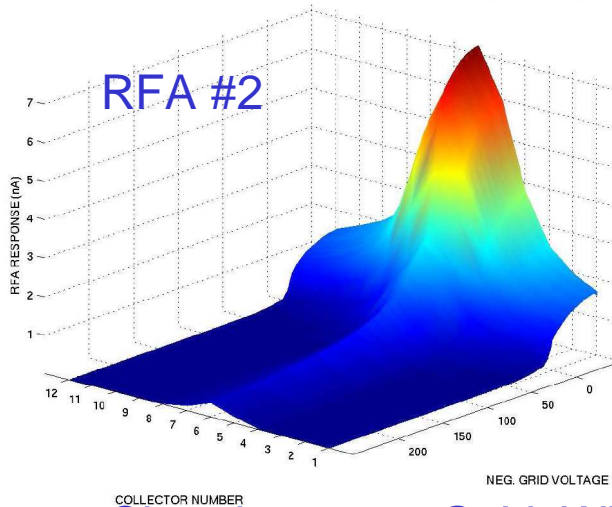
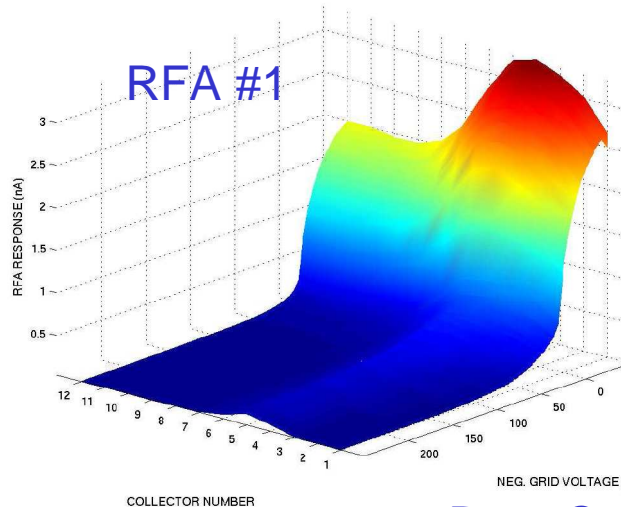


Wiggler RFA Characterization

SCW02WA_RFA1_20081102_0385 COLLECTORS 1x45x1mA e+ 5.3 GeV Wignlers Off Copper

SCW02WA_RFA2_20081102_0385 COLLECTORS 1x45x1mA e+ 5.3 GeV Wignlers Off Copper

SCW02WA_RFA3_20081102_0387 COLLECTORS 1x45x1mA e+ 5.3 GeV Wignlers Off Copper



Bare Copper Chamber – 5.3 GeV, Wiggler OFF

1x45x1mA positron beam, 14 ns bunch spacing

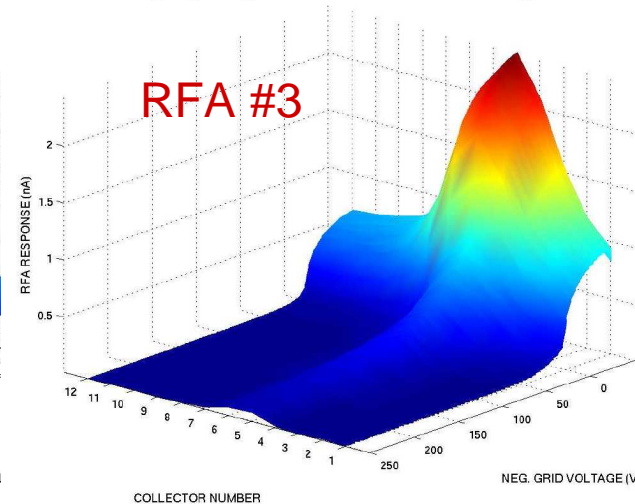
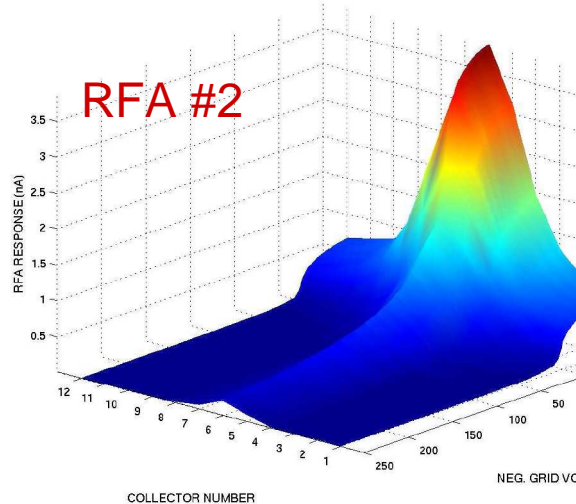
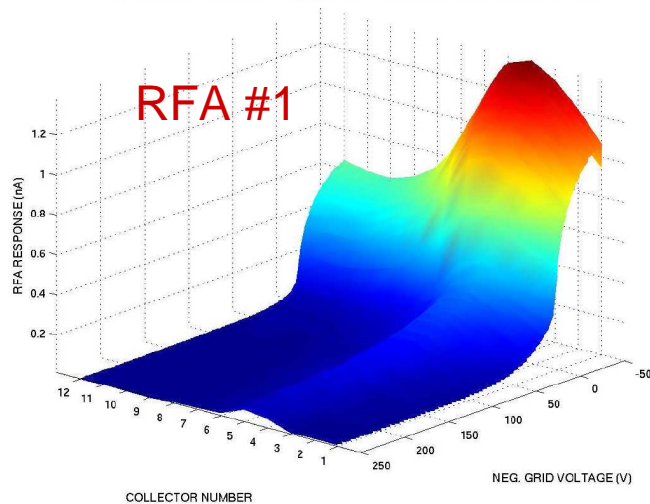
Some Residual Field Present

TiN-coated Chamber – 5.3 GeV, Wiggler OFF

SCW02WB_RFA1_20081102_0391 COLLECTORS 1x45x1mA e+ 5.3 GeV Wignlers Off TiN

SCW02WB_RFA2_20081102_0391 COLLECTORS 1x45x1mA e+ 5.3 GeV Wignlers Off TiN

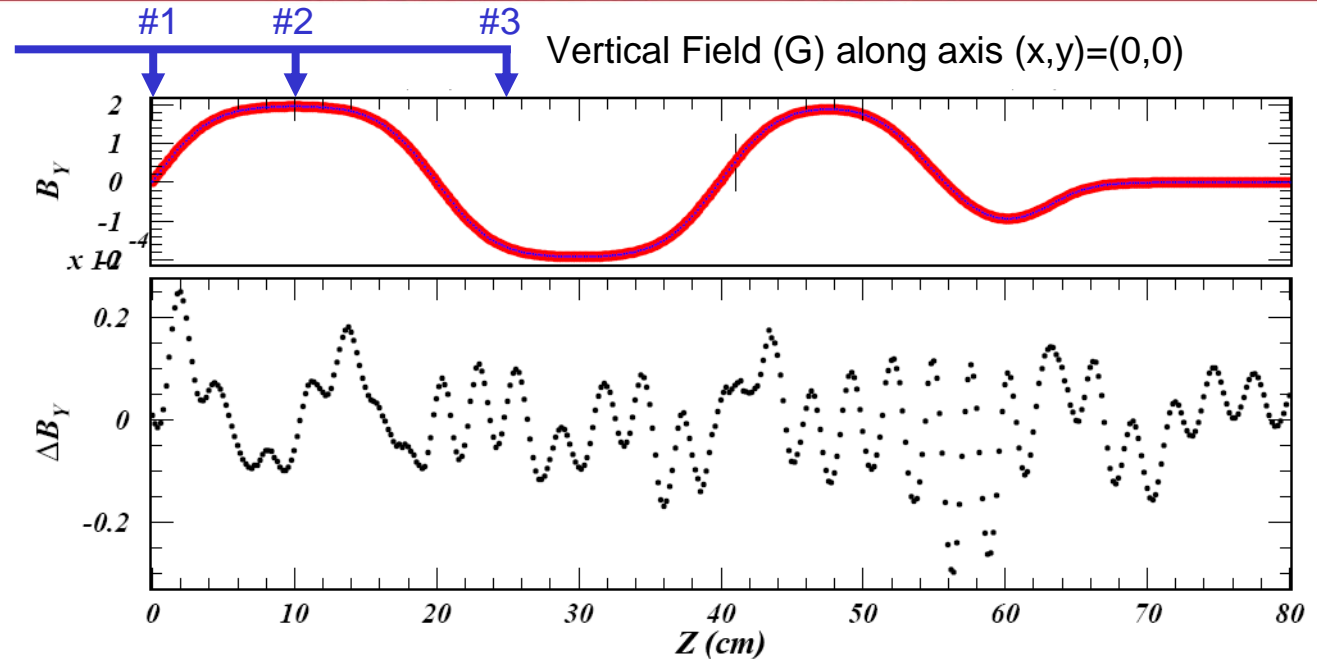
SCW02WB_RFA3_20081102_0396 COLLECTORS 1x45x1mA e+ 5.3 GeV Wignlers Off TiN





Fields in the Wiggler

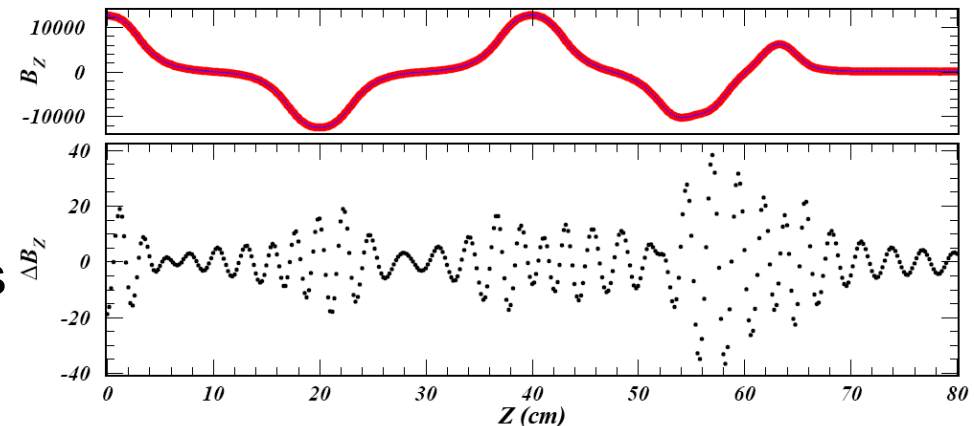
Wiggler RFA Locations:



• Detectors

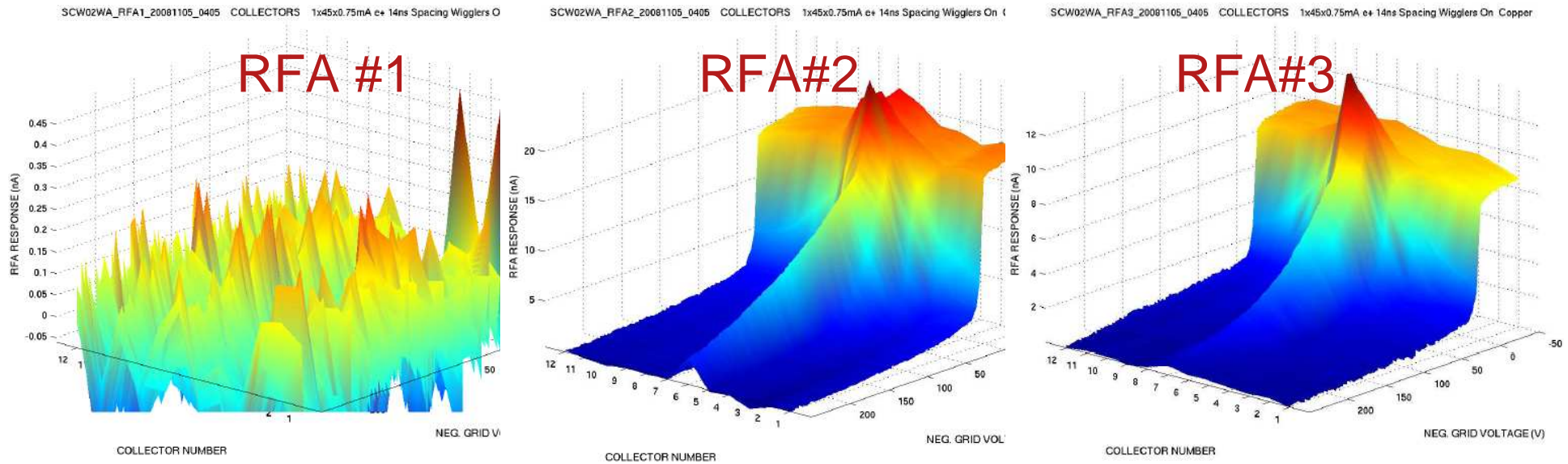
- RFA #2 - choice of detector at peak field is obvious.
- RFA #1 - should be a “canary”, ie, no EC electrons expected to penetrate longitudinal field lines with field applied.

Longitudinal Field (G) along axis (x,y)=(0,2.4) [cm]





Cu Wiggler Chamber RFAs at 2 GeV:

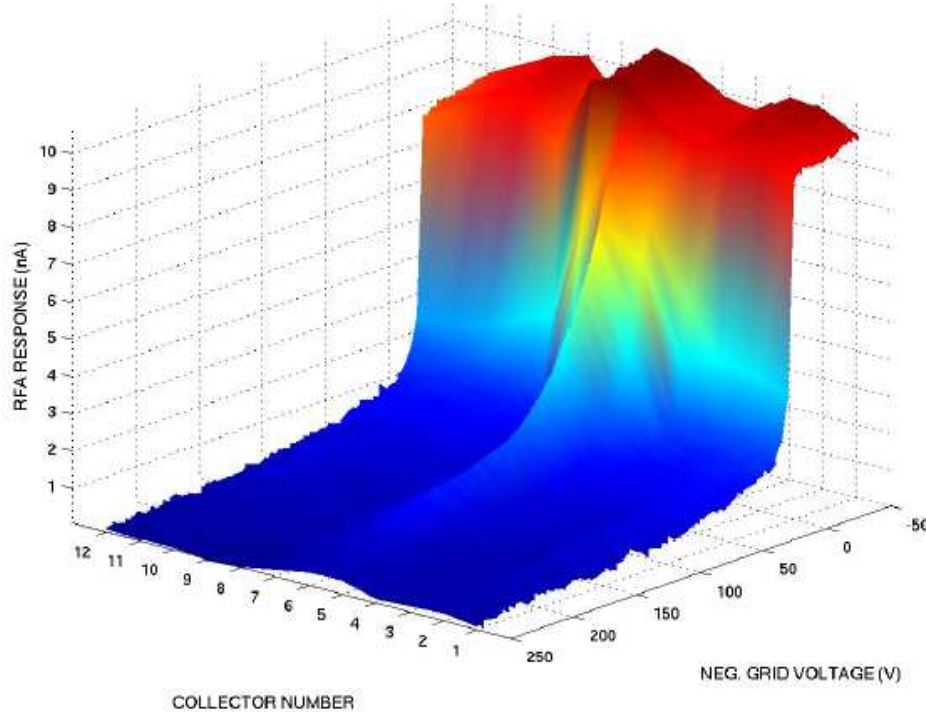


1x45x0.75mA (1.2×10^{10} /bunch)

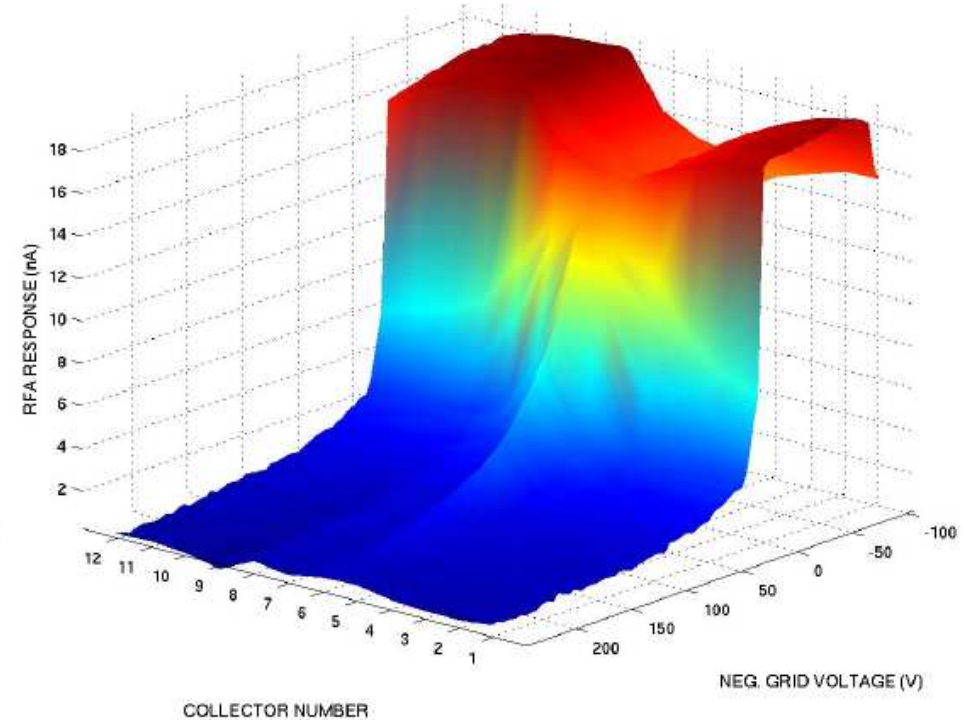


More Cu vs TiN Comparisons

1 × 45 × 0.4 mA (0.64×10^{10} /bunch), 14ns, Cu



1 × 45 × 0.5 mA (0.8×10^{10} /bunch), 14ns, TiN



- Low bunch current checks
- RFA in center of pole for each wiggler

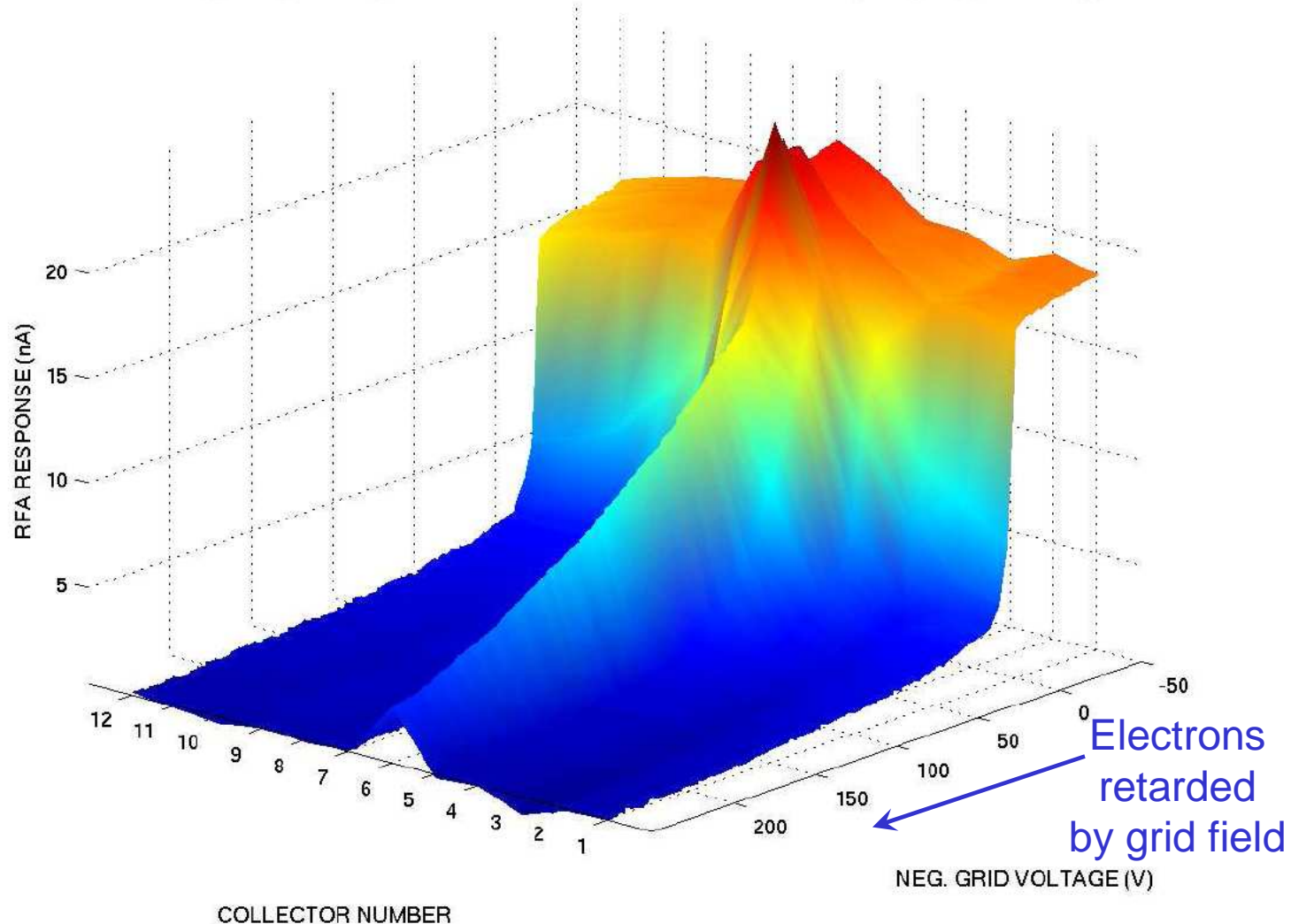
- Activity in central region appears suppressed



Wiggler RFA Data

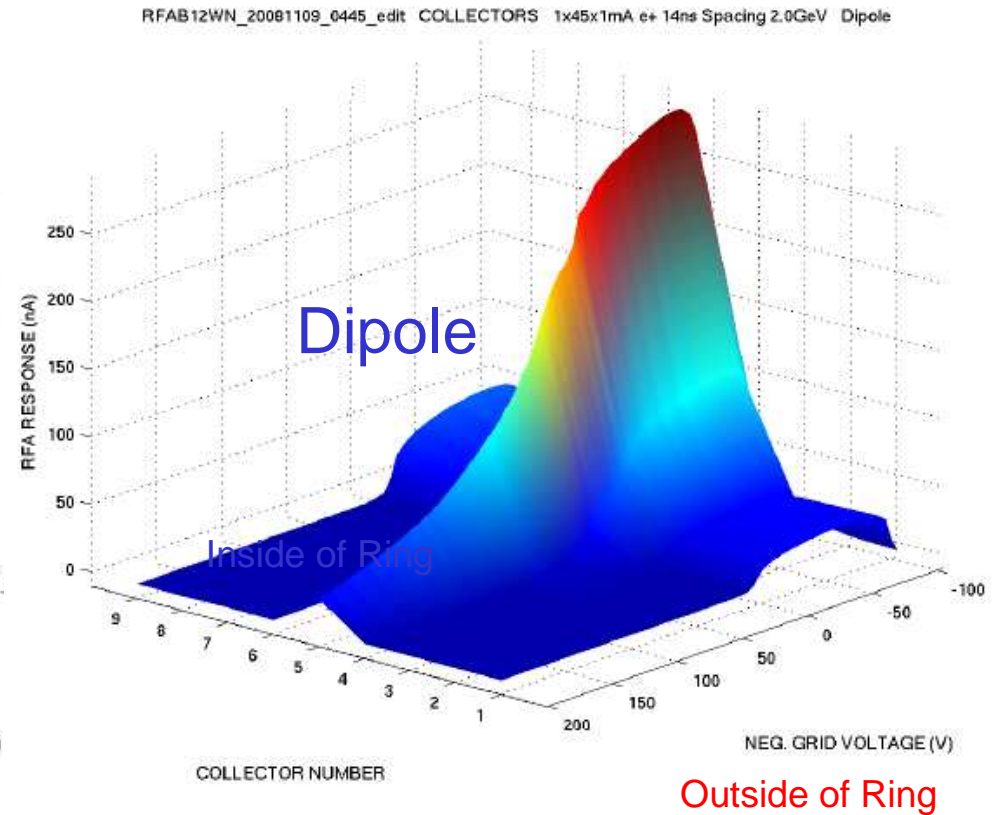
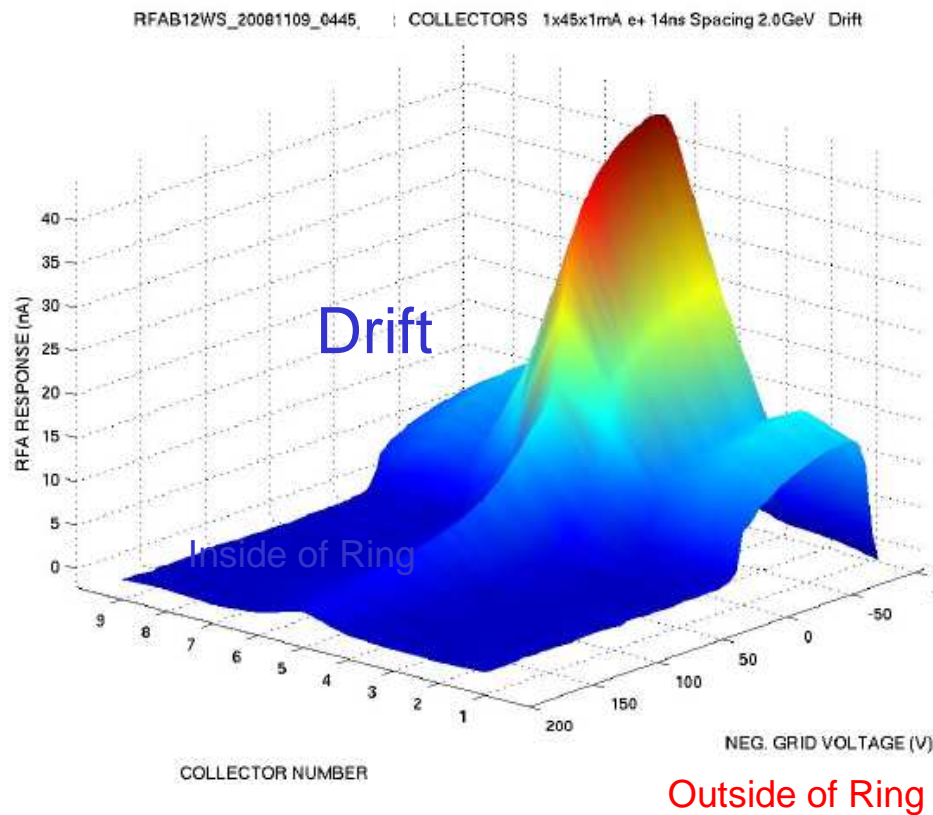
- 1 Train, 45 bunches, 1.2×10^{10} positrons/bunch

SCW02WA_RFA2_20081105_0405 COLLECTORS 1x45x0.75mA e+ 14ns Spacing Wigglers On Copper



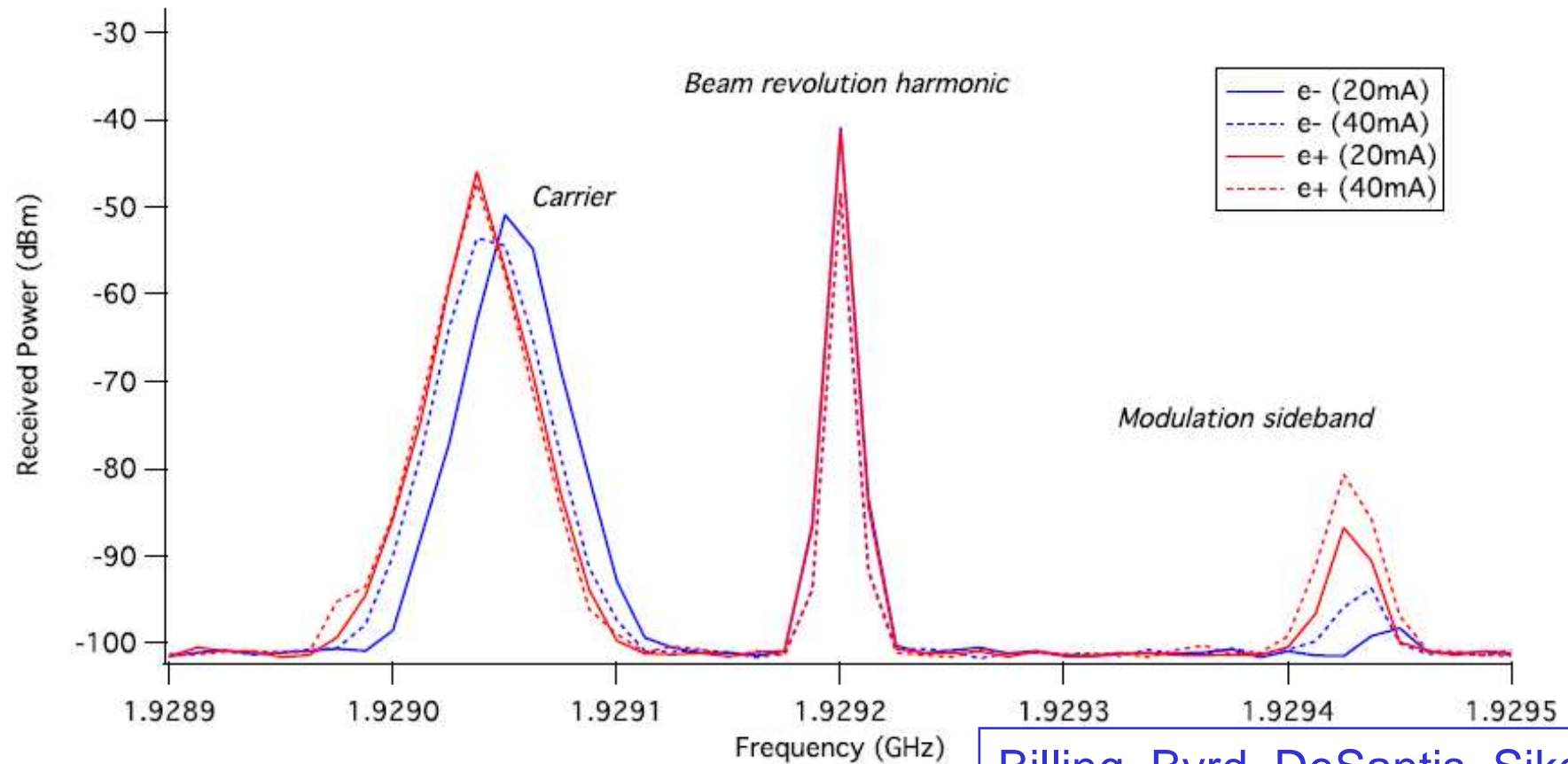


- Look at CESR vacuum chamber RFAs in drift and dipole
- $1 \times 45 \times 1 \text{ mA}$ (1.6×10^{10} /bunch) at 2 GeV





Observed modulation signal for both electron and positron beams at 14W test location in CESR.



Billing, Byrd, DeSantis, Sikora



- Major effort underway to understand recent data as well as to benchmark codes in the CsrTA parameter regime:

- RFA response
- Data-simulation comparisons
- Detailed comparisons between codes

G. Dugan, J. Calvey,
J. Crittenden, J. Livezey,
S. Greenwald, P. Jain

Comparison with data of 4/2/07 Coherent vertical tune shifts at 1.9 GeV

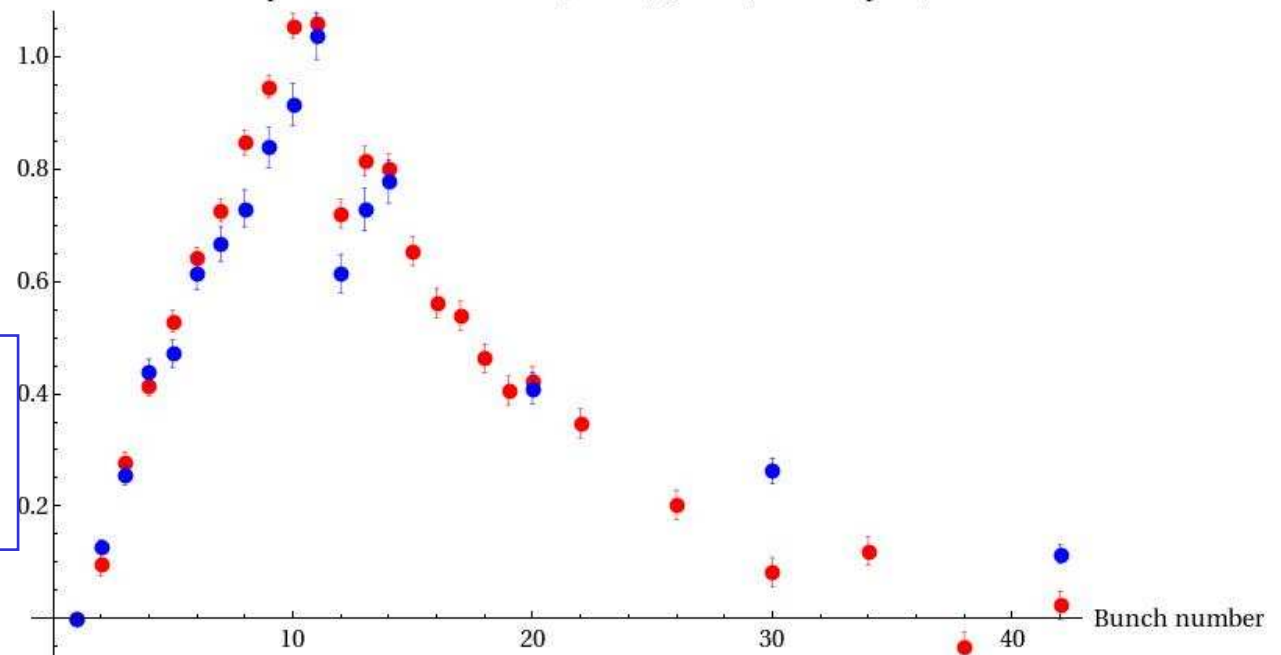
Vertical tune shift (kHz) vs bunch number

Red: Tune shift data 1.885 GeV 10 bunch train 0.75 mA/bunch positrons 4/2/07

Blue: simulation

CESR-TA drift at 1.885 GeV: SEY=2.0, $r=15\%$, QE=12%, 51 nicks, pa=1, 120000

CESR-TA dipole at 1.885 GeV: SEY=2.0, $r=15\%$, QE=12%, 51 nicks, pa=1, 20000

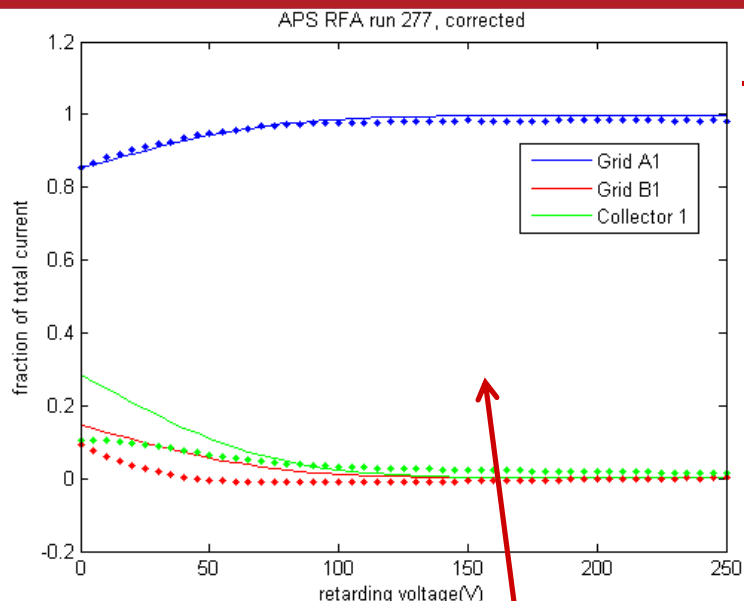




- **Validate EC buildup and dynamics simulations using CestrTA data**
 - Develop confidence in the application of these simulations to predict cloud behavior in the ILC damping ring
- **Tools:**
 - Simulation codes: POSINST, ECLLOUD, CLOUDLAND
 - Analytic and numerical estimates of response of beam to cloud
 - RFA response models
- **Initial steps:**
 - Benchmark simulation codes using simple cases relevant to CestrTA and ILC DR conditions.
 - Simulate cloud buildup in RFA-instrumented chambers, and RFA instrumental response, to guide RFA experiments as probes of average cloud density.
 - Simulate coherent tune shifts, to guide tune shift measurements as probes of cloud density and dynamics
 - Compute EC-related parameters for all beamline elements in CestrTA
 - Simulate ring-averaged cloud buildup and compute coherent tune shifts

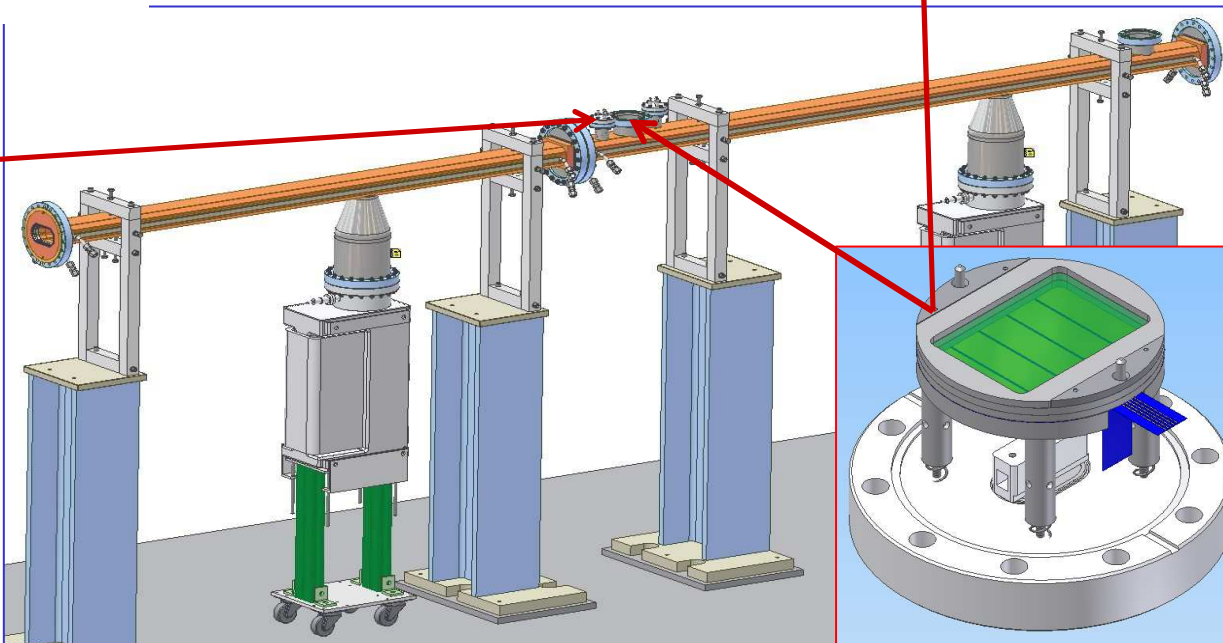
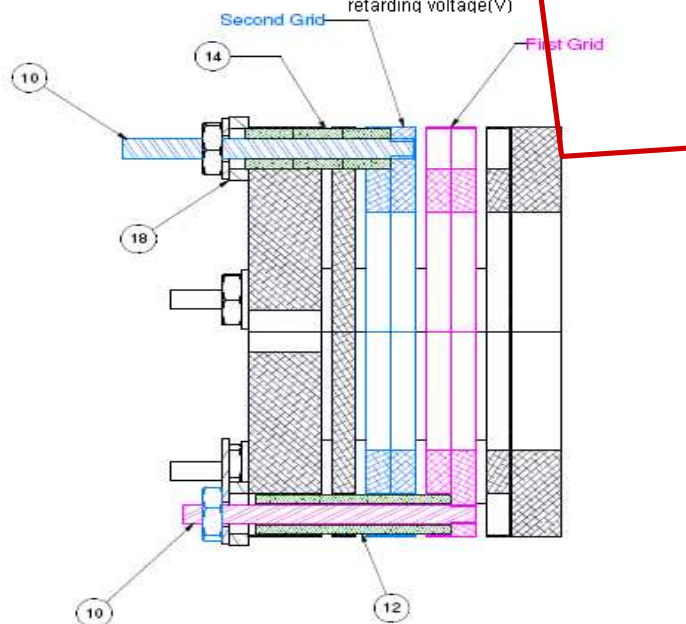
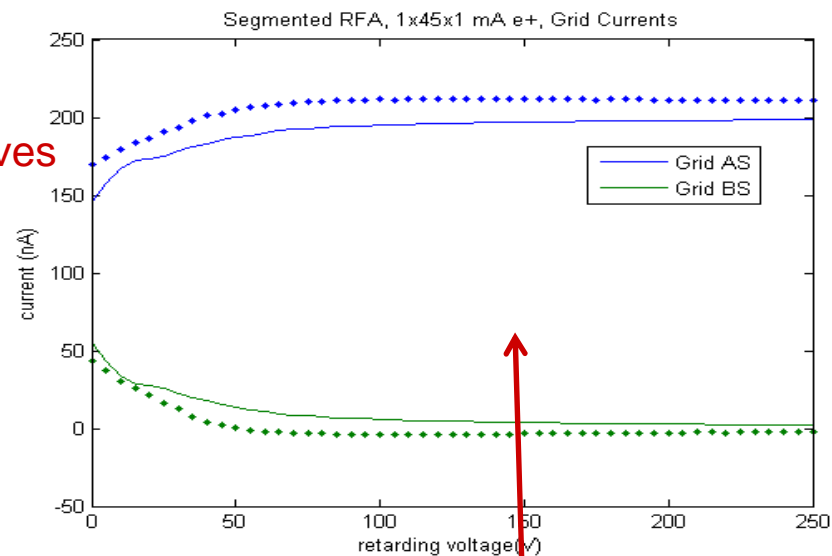


RFA Data and Simulation I



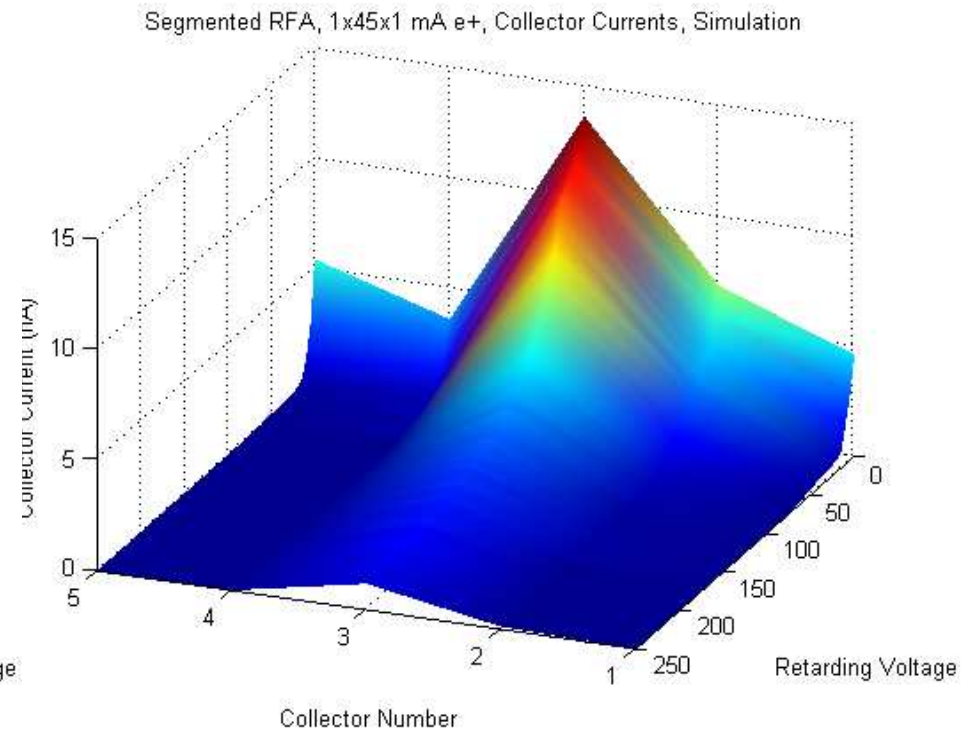
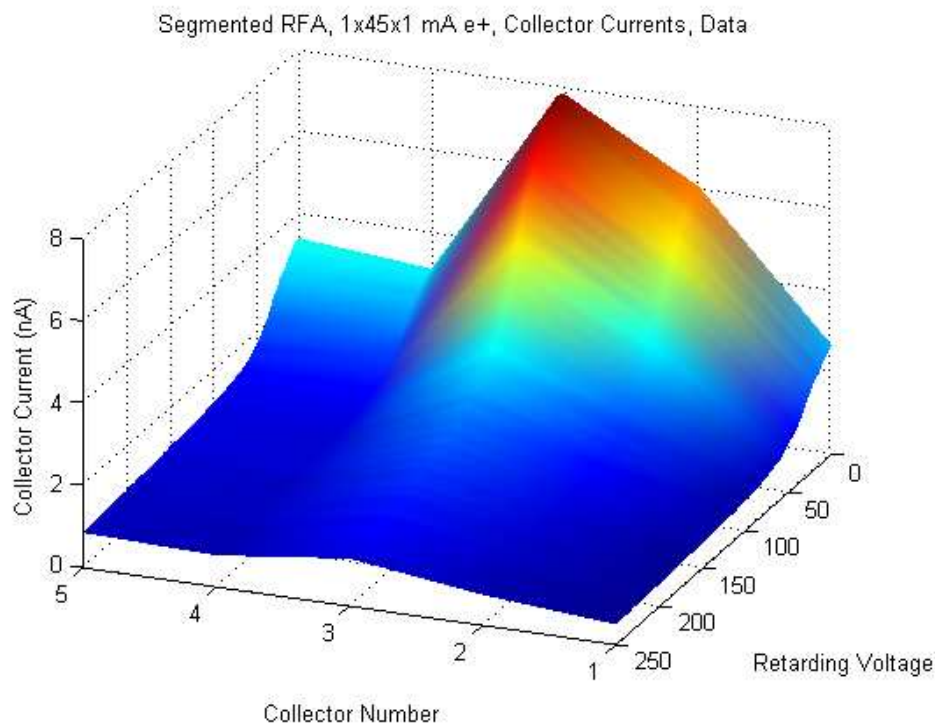
1x45x1mA e⁺
Data – points
Simulation – curves
(POSINST plus
RFA model)

J. Calvey



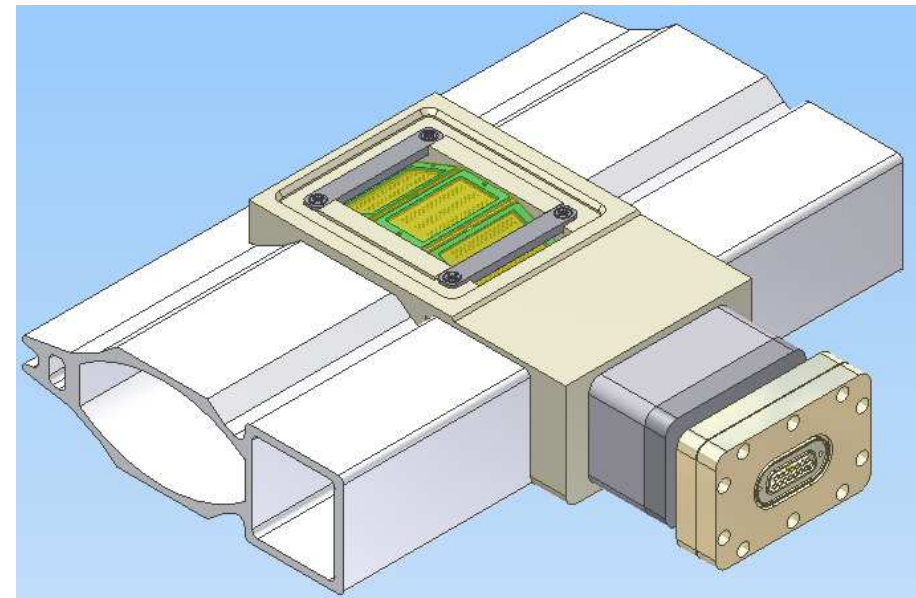
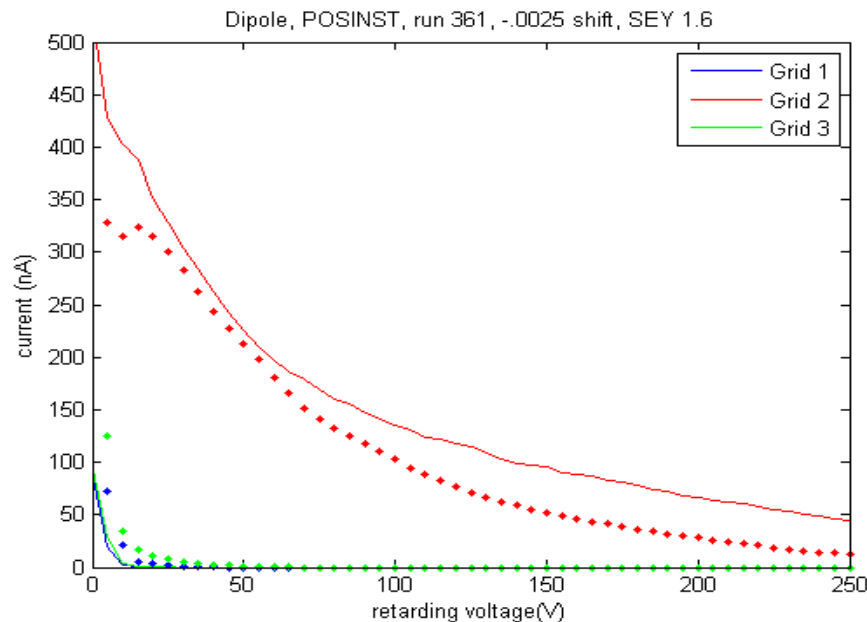


- **Collector currents match qualitatively**
 - Plots show collector current vs collector number (collector 1 is opposite source point) and retarding voltage
 - Data on left, simulation on right
- **We will try to match the azimuthal distribution better by adjusting simulation parameters**





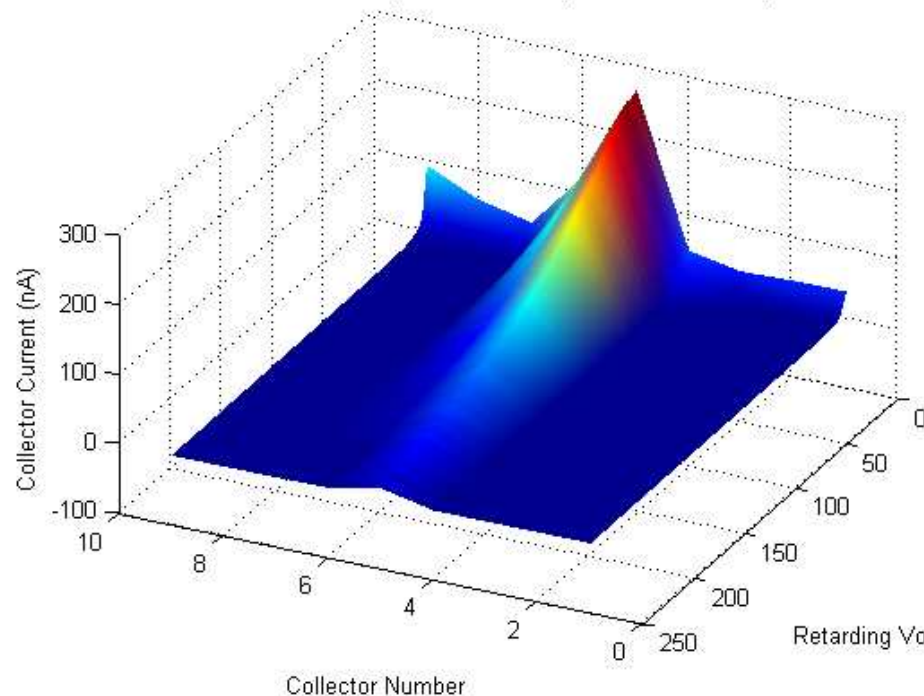
- 3 retarding grids, 9 collectors
 - One presently deployed in a dipole and one in a drift
- Attempting basic adjustments to model to obtain best data-simulation comparison
 - Adjust SEY parameters in POSINST (these plots are for an SEY peak of 1.6)
 - Shift RFA holes to represent a non-central beam



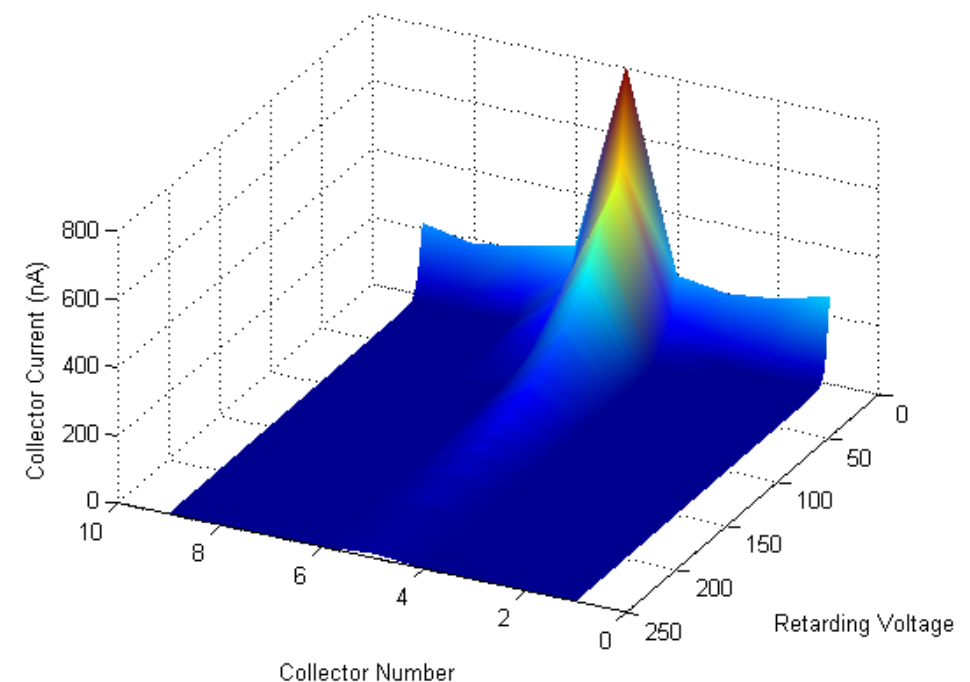


- Collector data matches qualitatively between data and simulation
- Optimization of EC model parameters and further studies of RFA efficiencies underway to understand remaining differences between data and simulation

Thin Insertable RFA, 1x45x1 mA e+, Collector Currents, Data



Thin Insertable RFA, 1x45x1 mA e+, Collector Currents, Simulation





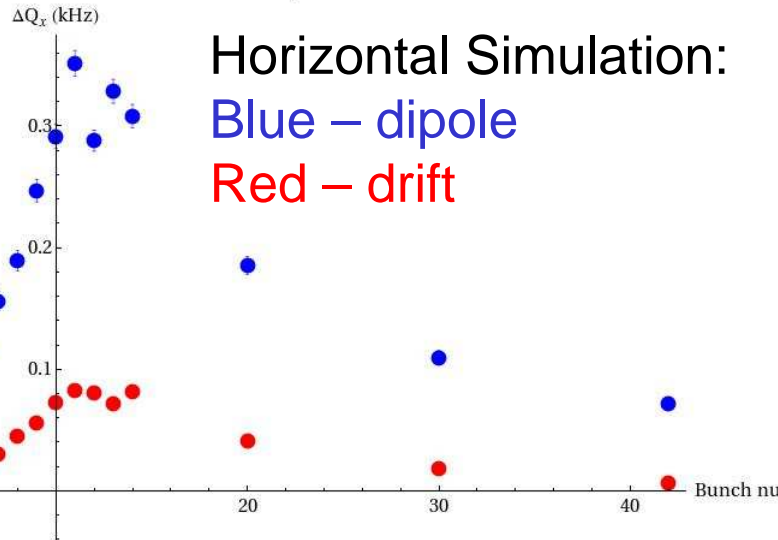
- “Witness bunch” technique:
 - a train of “loading bunches” generates a cloud density around the ring
 - “witness bunches” are placed at variable times after the loading train, and the coherent tune of the witness bunch is measured. The coherent tune shift is a measure of the beam-averaged field gradient due to the cloud charge density at the time of the witness bunch
- Coherent tune shift measurements (both vertical and horizontal tune) using the witness bunch technique have been done in a variety of conditions
 - Electrons and positrons
 - 1.9-2.1 GeV and 5.3 GeV
 - Various loading trains
- We have also made measurements of the systematic variation of tune shift along a train vs. bunch current
- Will show a few examples...



CESR-TA dipole at 1.885 GeV: SEY=2.0, r=15%, QE=12%, 51 nicks, pa=1, 20000
CESR-TA drift at 1.885 GeV: SEY=2.0, r=15%, QE=12%, 51 nicks, pa=1, 120000

Horizontal tune shift vs bunch number

Red: drift
Blue: dipole

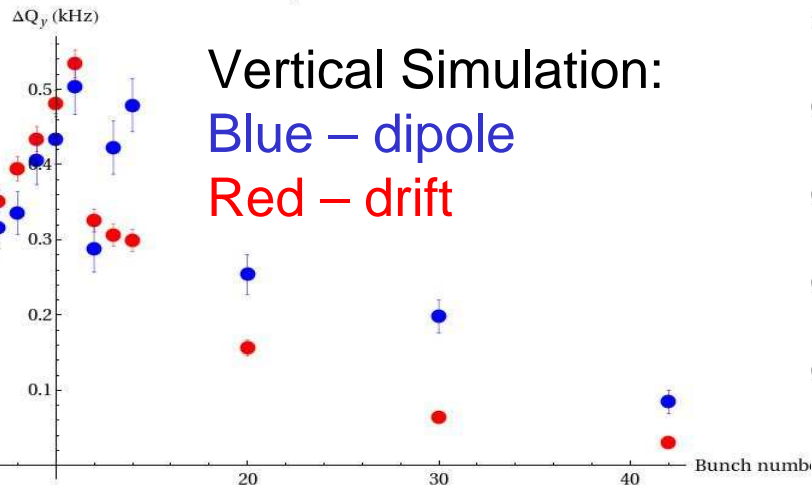


Horizontal Simulation:
Blue – dipole
Red – drift

CESR-TA dipole at 1.885 GeV: SEY=2.0, r=15%, QE=12%, 51 nicks, pa=1, 20000
CESR-TA drift at 1.885 GeV: SEY=2.0, r=15%, QE=12%, 51 nicks, pa=1, 120000

Vertical tune shift vs bunch number

Red: drift
Blue: dipole



Vertical Simulation:
Blue – dipole
Red – drift

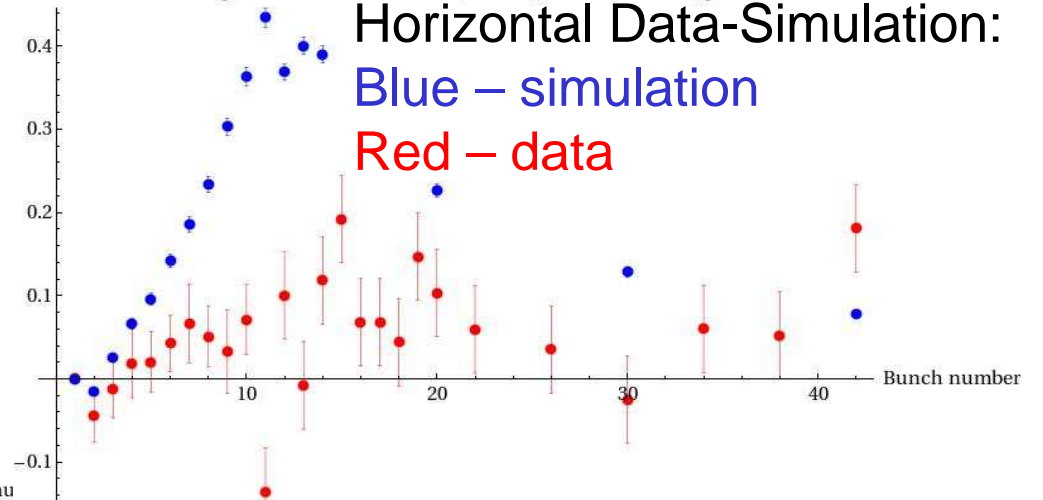
Horizontal tune shift (kHz) vs bunch number

Red: Tune shift data 1.885 GeV 10 bunch train 0.75 mA/bunch positrons 4/2/07

Blue: simulation

CESR-TA drift at 1.885 GeV: SEY=2.0, r=15%, QE=12%, 51 nicks, pa=1, 120000

CESR-TA dipole at 1.885 GeV: SEY=2.0, r=15%, QE=12%, 51 nicks, pa=1, 20000



Horizontal Data-Simulation:
Blue – simulation
Red – data

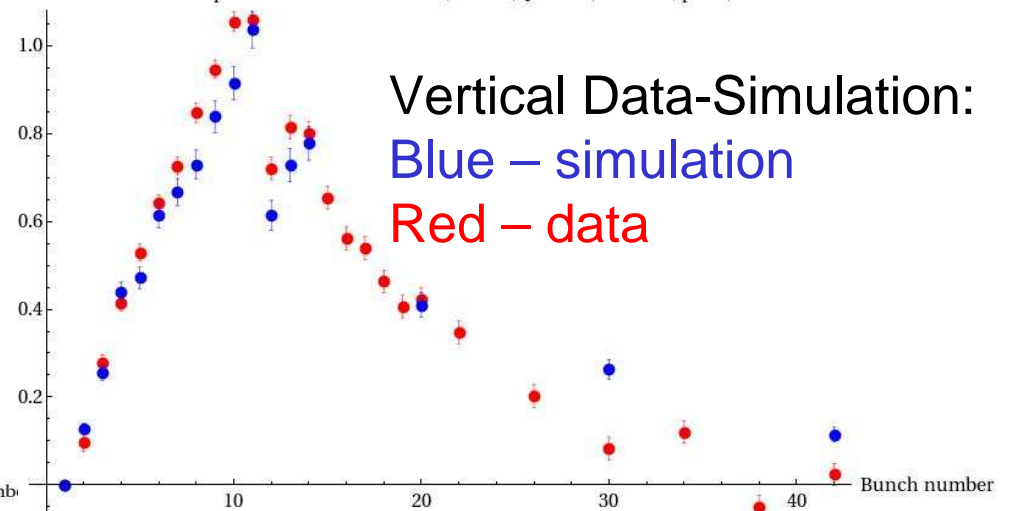
Vertical tune shift (kHz) vs bunch number

Red: Tune shift data 1.885 GeV 10 bunch train 0.75 mA/bunch positrons 4/2/07

Blue: simulation

CESR-TA drift at 1.885 GeV: SEY=2.0, r=15%, QE=12%, 51 nicks, pa=1, 120000

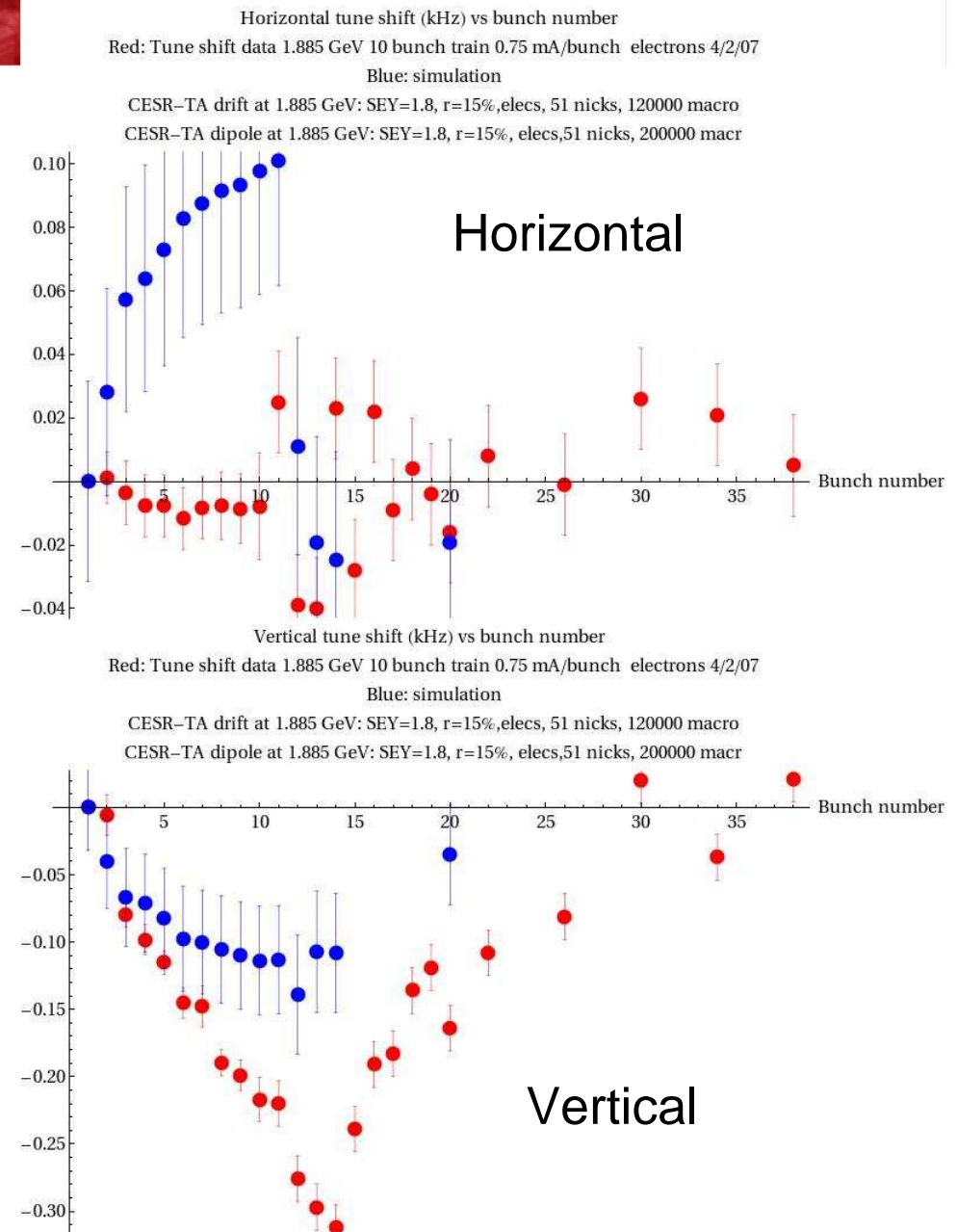
CESR-TA dipole at 1.885 GeV: SEY=2.0, r=15%, QE=12%, 51 nicks, pa=1, 20000



Vertical Data-Simulation:
Blue – simulation
Red – data



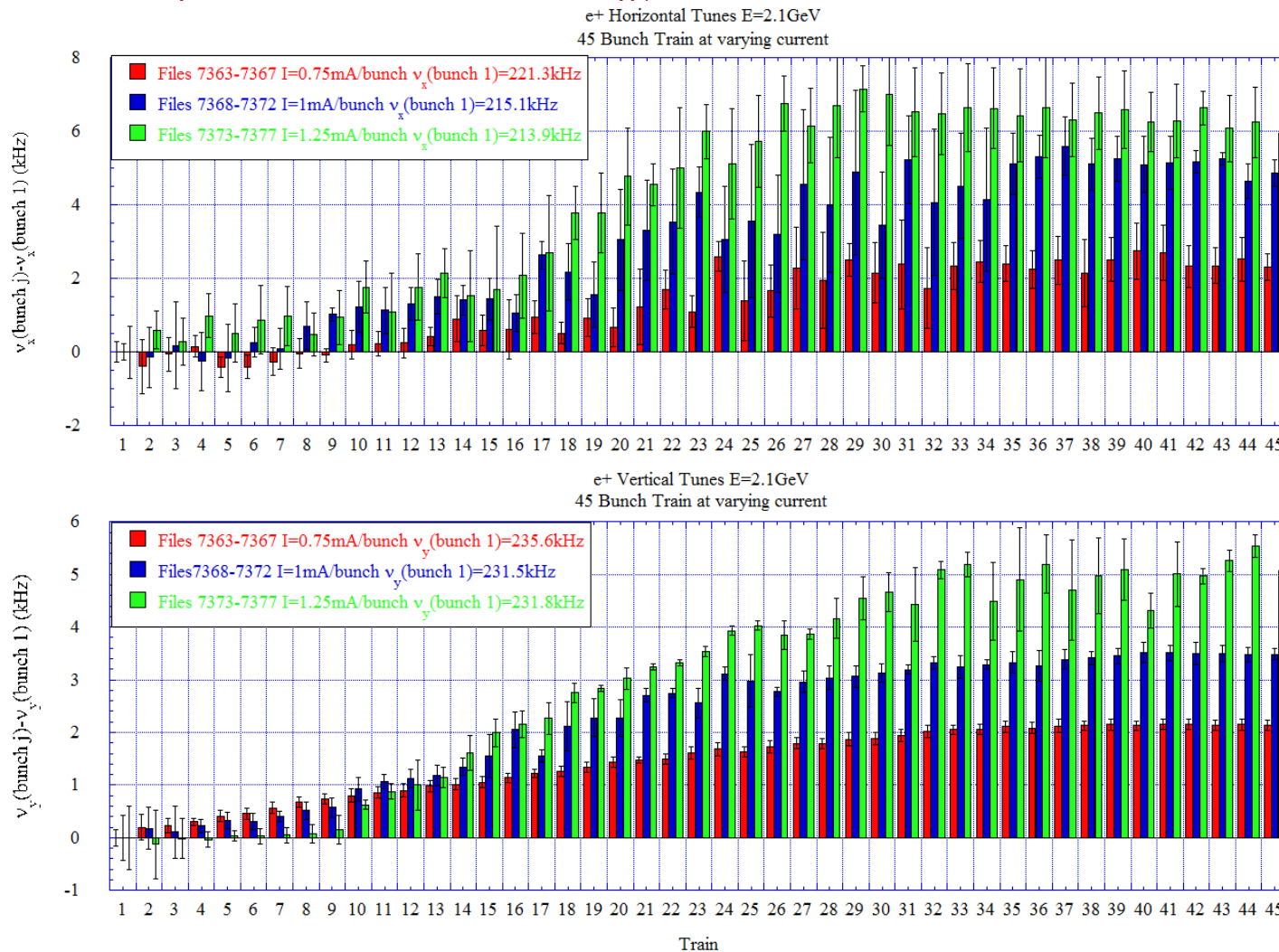
- 1.9 GeV Data
- To date have been unable to get good data-simulation comparisons for the electron beam
- Plots
 - Blue Points – simulation
 - Red Points – Data





A Surprise

- During a brief checkout at low energy, took tune data for 45 bunch trains versus bunch current:



- Observe significant horizontal tune shift for the first time!!
- What is driving the difference between data taken in the CESR-c configuration and the new CesrTA configuration???



- **Basic Goals for January Run:**
 - Low Emittance Tuning
 - Begin detailed tuning of baseline lattice
 - Beam-based alignment
 - Integrate improved BPM system
 - Electron Cloud
 - Systematic studies of EC growth in CESR drifts, dipoles and wigglers
 - Growth versus wiggler parameters and primary photon flux (all adjustable in wiggler straight)
 - Impact of detectors on measurements
 - Growth versus bunch spacing
 - Growth versus energy
 - Detailed comparison of measurements with RFAs, measurements with TE wave transmission, and simulations
 - Electrons versus positrons
 - Continue dynamics studies
 - Explore recent observation of conditions with significant ΔQ_x
 - Explore new measurement techniques
 - Further checks of systematic effects
 - Instability measurements at lowest emittances achieved during January run
 - » Bunch-by-bunch tunes, bunch-by-bunch vertical beam size, and mode spectra for various train configurations



- **Simulation**
 - Code Benchmarking
 - Particularly SEY models in E-CLOUD, CLOUDLAND and POSINST.
 - Improved model of the RFA response is in progress.
 - Include dynamic effects in the tune shift calculations (requires integration of beam motion into the simulation codes).
 - We need a full 3D simulation of the damping wiggler . The present plan is to use WARP-POSINST, relying on our LBNL collaborators.
 - Simulation of incoherent emittance growth – to be compared with measurements using visible BSM and xBSM.
 - Simulations exploring the dependence of cloud effects on the beam as a function of energy, species, bunch population, bunch spacing, and emittance and comparison with measurement.
- **February Down**
 - 3 items directly associated with EC portion of program
 - Installation of PEP-II EC hardware
 - Wiggler photon stop to provide more flexible energy range for wiggler tests
 - Possible installation of collaborator test chambers
- **Preparation of Test Chambers**
 - Wiggler chambers with grooves and low-profile clearing electrode (CU/KEK/LBNL/SLAC)
 - Carbon coated chamber (CERN)
 - Groove chamber for chicane test (SLAC)
 - Enamel clearing electrode chamber (FNAL/Project X)
 - SEY sample chamber (SLAC/FNAL)



- **Visiting Collaborators**
 - Who to contact:
 - Mark Palmer for EC studies (map36@cornell.edu)
 - Dave Rubin for LET work (dlr10@cornell.edu)
 - Forms and computer registration:
<https://wiki.lepp.cornell.edu/ilc/bin/view/Public/CesrTA/VisitorsPage>
 - Schedule Information:
<https://wiki.lepp.cornell.edu/ilc/bin/view/Public/CesrTA/CesrTASchedule>
- **Where to look:**
 - Main CesrTA Wiki Page:
<https://wiki.lepp.cornell.edu/ilc/bin/view/Public/CesrTA/>
 - CesrTA Collaboration Meetings Page:
<https://wiki.lepp.cornell.edu/ilc/bin/view/Public/CesrTA/CollabMeetings>
- **Many areas restricted to the “CesrTA Collaboration”**
 - Need to register for ILC Wiki
- **Mailing List**
 - CesrTA Collaboration Mailings via the new mailing list
 - Can subscribe from the main CesrTA Wiki page



- **CesrTA is an ILC-driven R&D Program**
 - Recent news:
 - Formal collaboration between CLIC-ILC DR efforts now in place (co-conveners of group: Yannis Papaphilippou and myself)
 - Major short term emphasis on exploiting EC experimental program
 - Integrated set of monthly WebEx meetings: ILC DR, CLIC DR, CesrTA
- **CesrTA commissioning run is now complete**
 - CESR is now operational in a damping ring configuration
- **Critical data collected during a short 2-week run**
 - New systems are successfully coming on line
 - Initial data will provide guidance for January experimental run as well as ongoing upgrade work
 - The initial commissioning work has yielded both expected results as well as some surprises
 - Ongoing collaborative effort critical to achieving full understanding and a productive project!