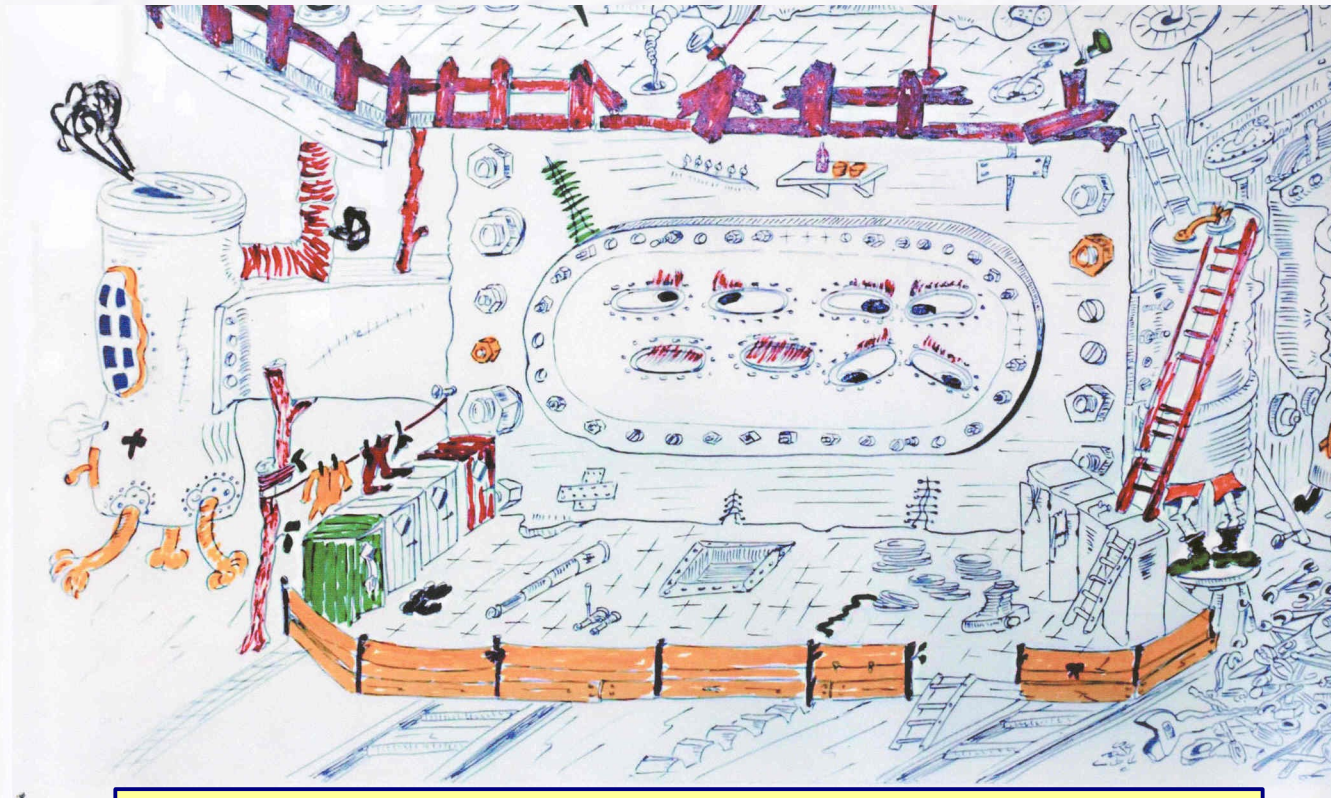
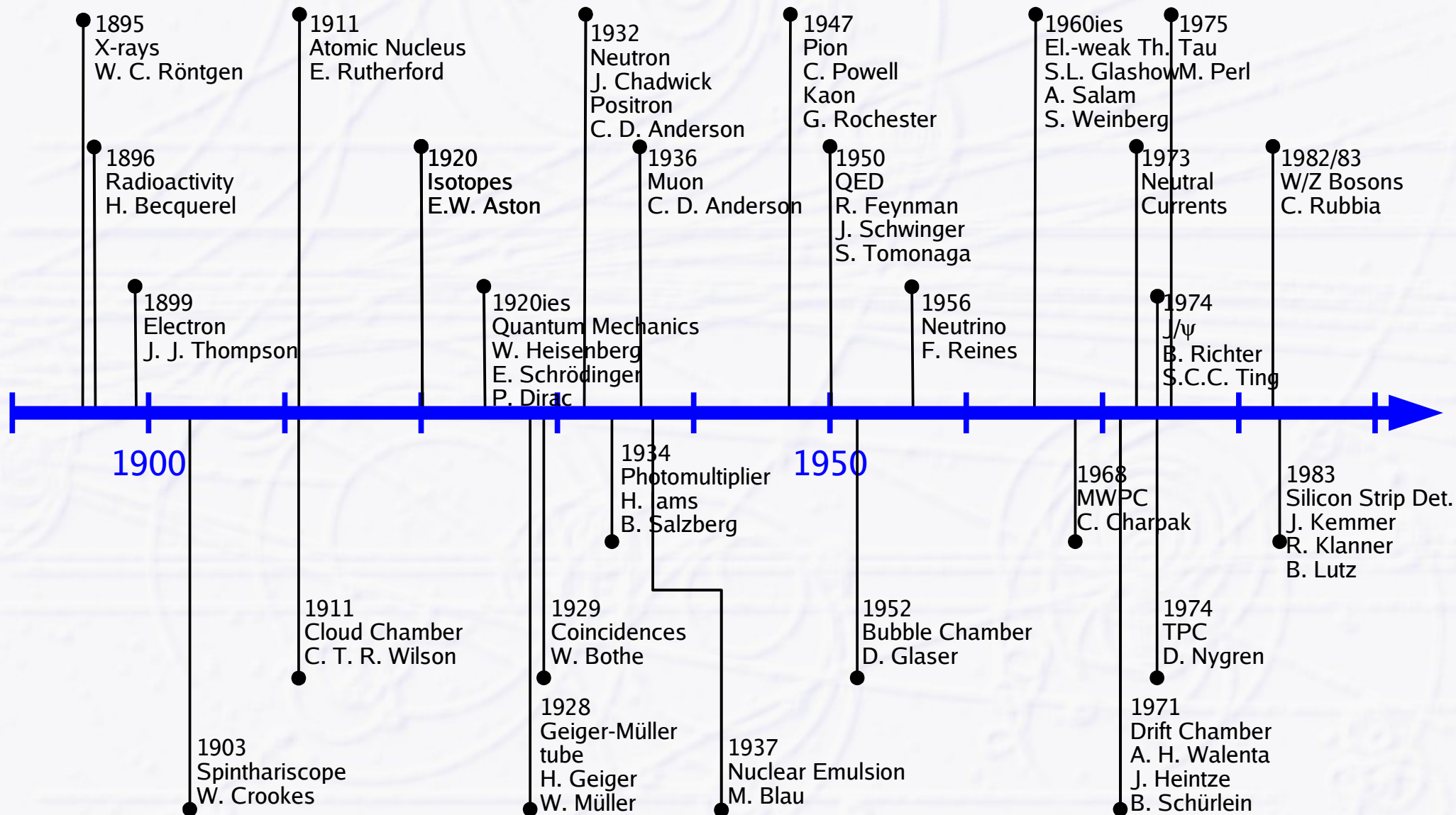


# ***History of Instrumentation***



**Artist's View of a Bubble Chamber by a CERN physicist**

# Timeline of Particle Physics and Instrumentation



# Early Image Detectors

## ● Second half of 19<sup>th</sup> century

→ growing interest in meteorological questions

- climate, weather phenomenon, **cloud formation**

→ people started to study condensation of water vapour in the lab

- also motivated by raising use of steam engines

## ● John Aitken built a “Dust Chamber” 1888

- water vapour mixed with dust in a controlled way

→ result: **droplets are formed around dust particles**

→ further speculations

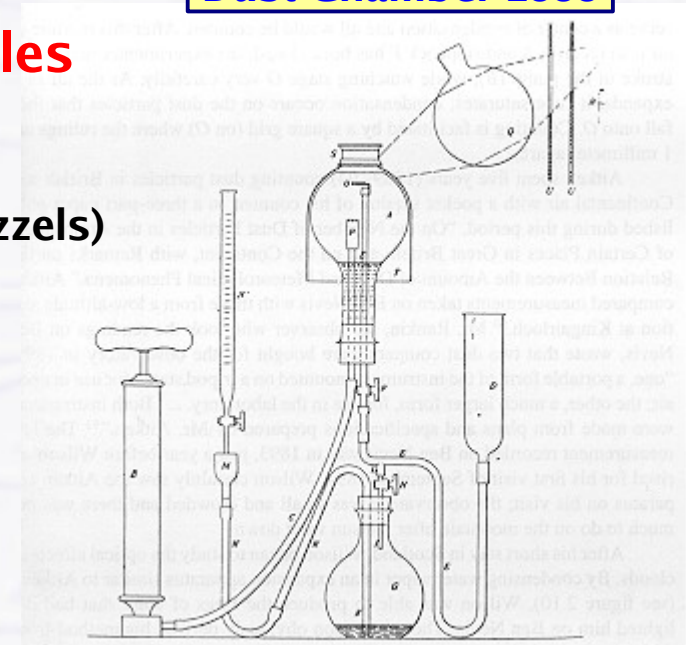
- electricity plays a role (from observations of steam nozzels)

## ● Charles T. R. Wilson became interested

→ first ideas to build a cloud chamber 1895  
to study influence of electricity/ions

- also to solve question why air shows natural slight conductivity

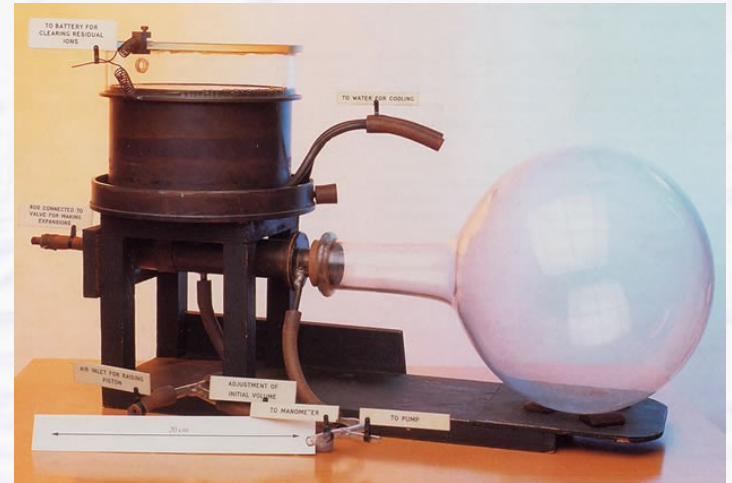
Dust Chamber 1888



# Cloud Chamber I

## ● Cloud chamber (1911 by Charles T. R. Wilson, Noble Prize 1927)

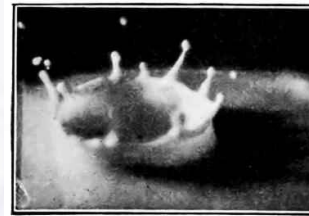
- chamber with saturated water vapour
- charged particles leave trails of ions
  - water is condensing around ions
- visible track as line of small water droplets



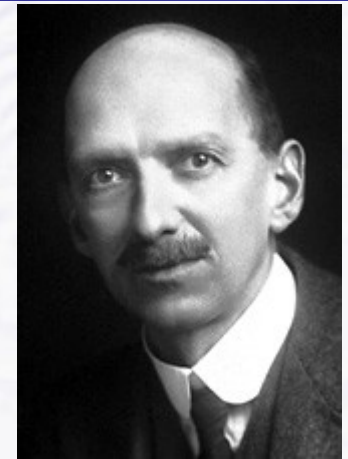
UK Science Museum

## ● Also required

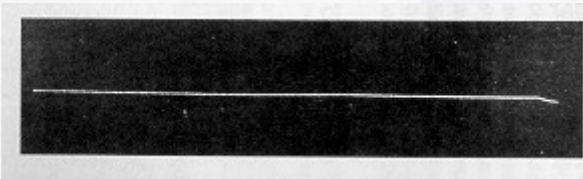
- high speed photographic methods
  - invented by Arthur M. Worthington 1908 to investigate the splash of a drop
  - ultra short flash light produced by sparks



Charles T. R. Wilson

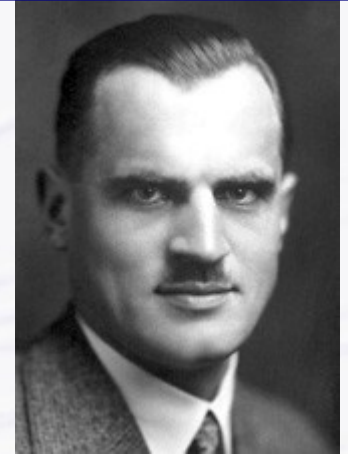


## ● First photographs of $\alpha$ -ray particles 1912



# Cloud Chamber II

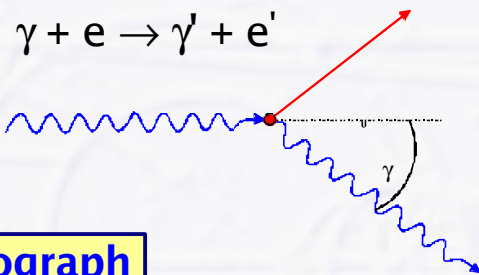
Arthur H. Compton



- Arthur H. Compton used the cloud chamber in 1922 to discover scattering of photons on electrons (Compton effect) (Nobel Prize 1927 together with Charles T. R. Wilson)

→ X-rays emitted into cloud chamber

- photon scattered on electrons (recoiling electron seen in cloud chamber)
- photon with reduced energy under certain angle visible by photo effect or Compton effect again



original photograph

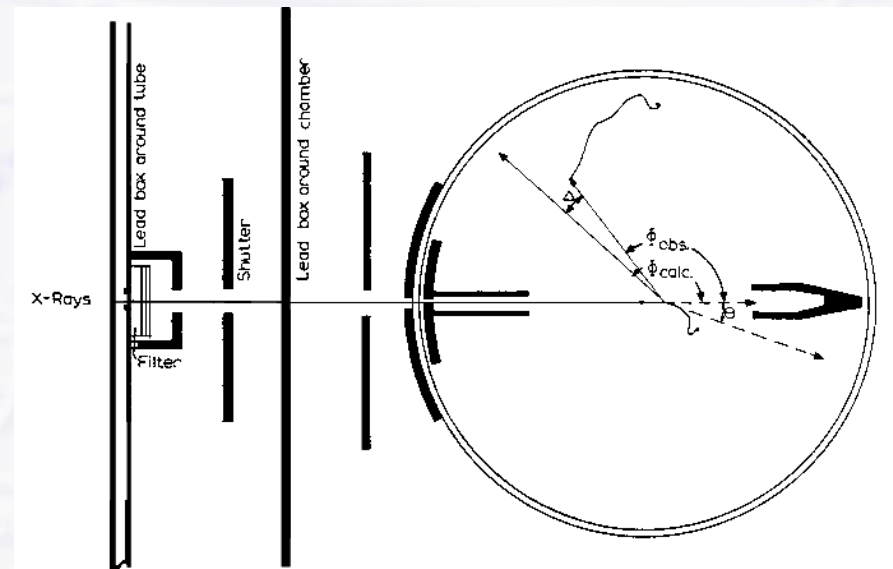
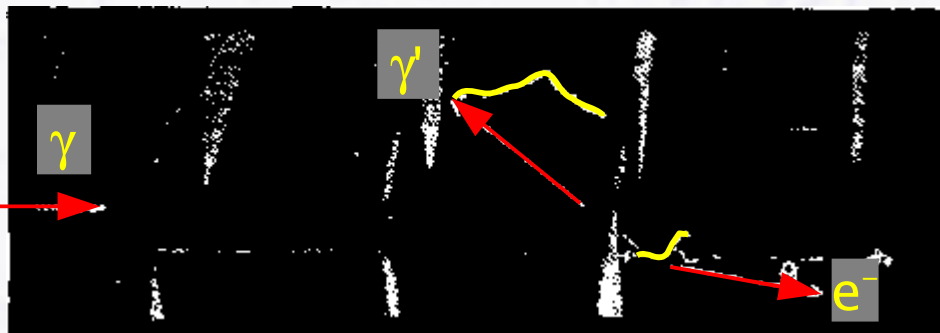
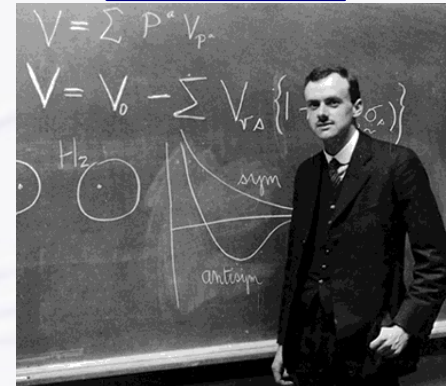


Fig. 10. An electron recoiling at an angle  $\theta$  should be associated with a photon deflected through an angle  $\phi$ .

# Cloud Chamber III

- Was also used for the discovery of the **positron**
  - predicted by Paul Dirac 1928 (Nobel Prize 1933)
  - found in cosmic rays by Carl D. Anderson 1932 (Nobel Prize 1936)

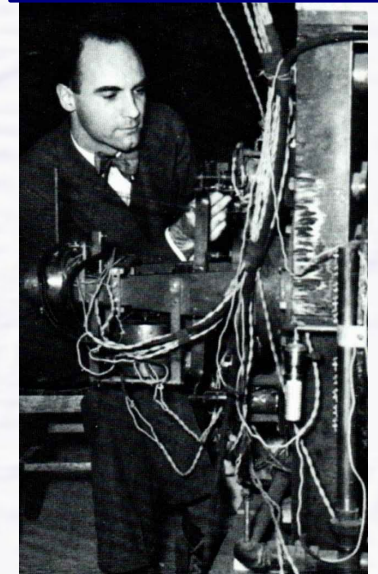
Paul Dirac



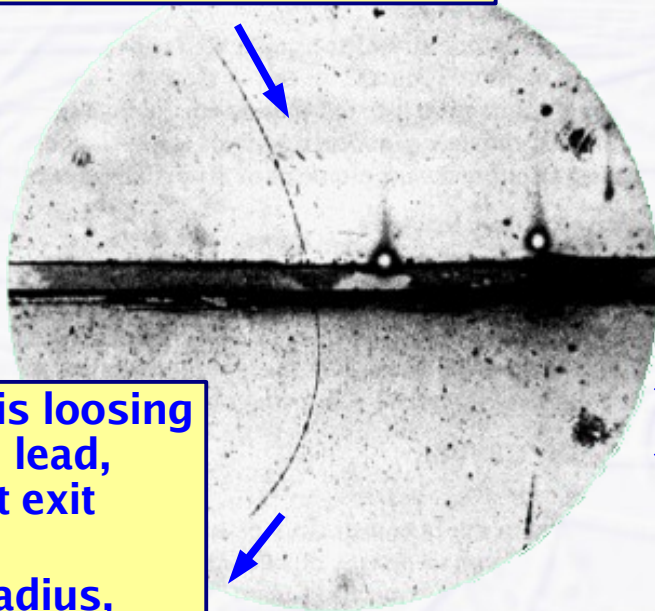
Anderson also found the **muon** in 1936, the first 2<sup>nd</sup> generation particle in the Standard Model

Isidor Isaac Rabi said: "Who ordered that?"

Carl D. Anderson



downward going positron, 63 MeV



positron is losing energy in lead, 23 MeV at exit

→ smaller radius, this defines the track direction!

6 mm lead plate



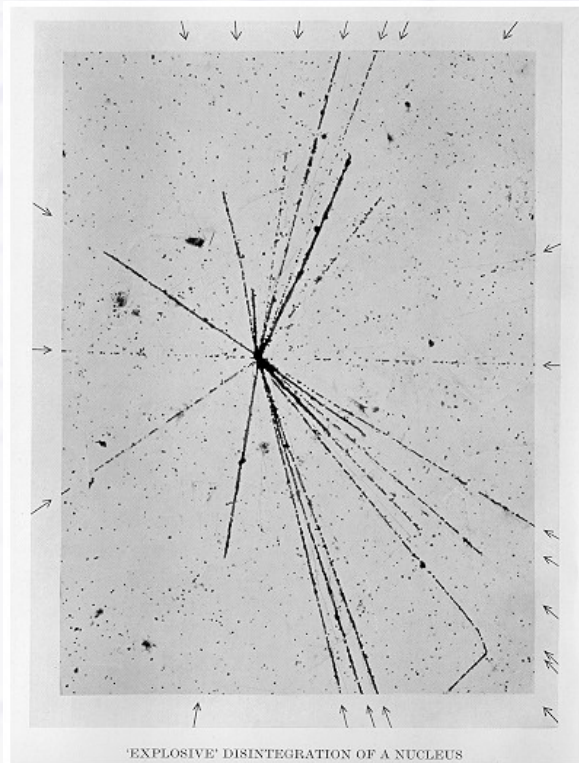
1.5 T magnetic field

# Nuclear Emulsion I

- **Pioneered by Marietta Blau between 1923 – 1938 (no Nobel Prize)**

- ➔ **photographic emulsion layer, 10 – 200  $\mu\text{m}$  thick, uniform grains of 0.1 – 0.3  $\mu\text{m}$  size**
- ➔ **very high resolution for particle tracks**
  - **analysis of developed emulsion by microscope**

Marietta Blau



nuclear disintegration from cosmic rays, observed 1937 for the first time

- **Since early 20<sup>th</sup> century**

- ➔ **important role of photography to study radioactivity**
- ➔ **but capability to make individual tracks visible not seen until nuclear emulsion technique was developed**

# Nuclear Emulsion II

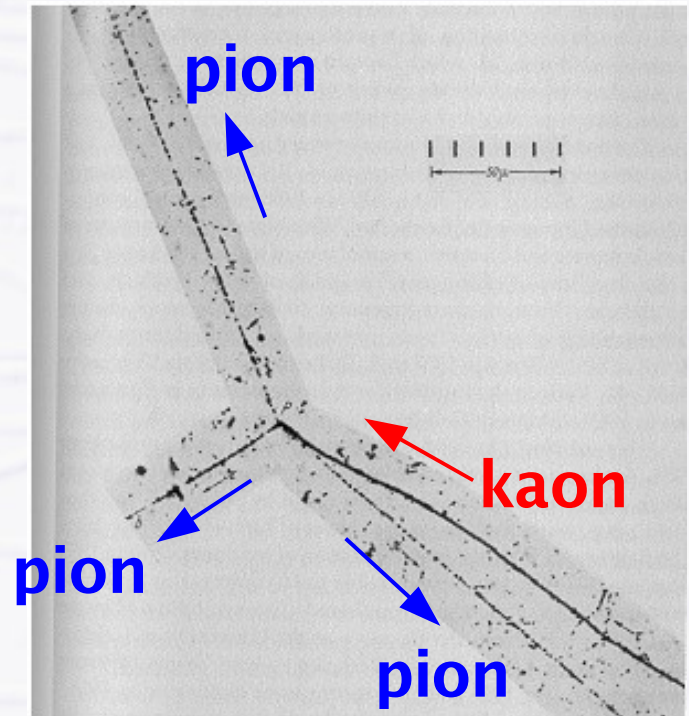
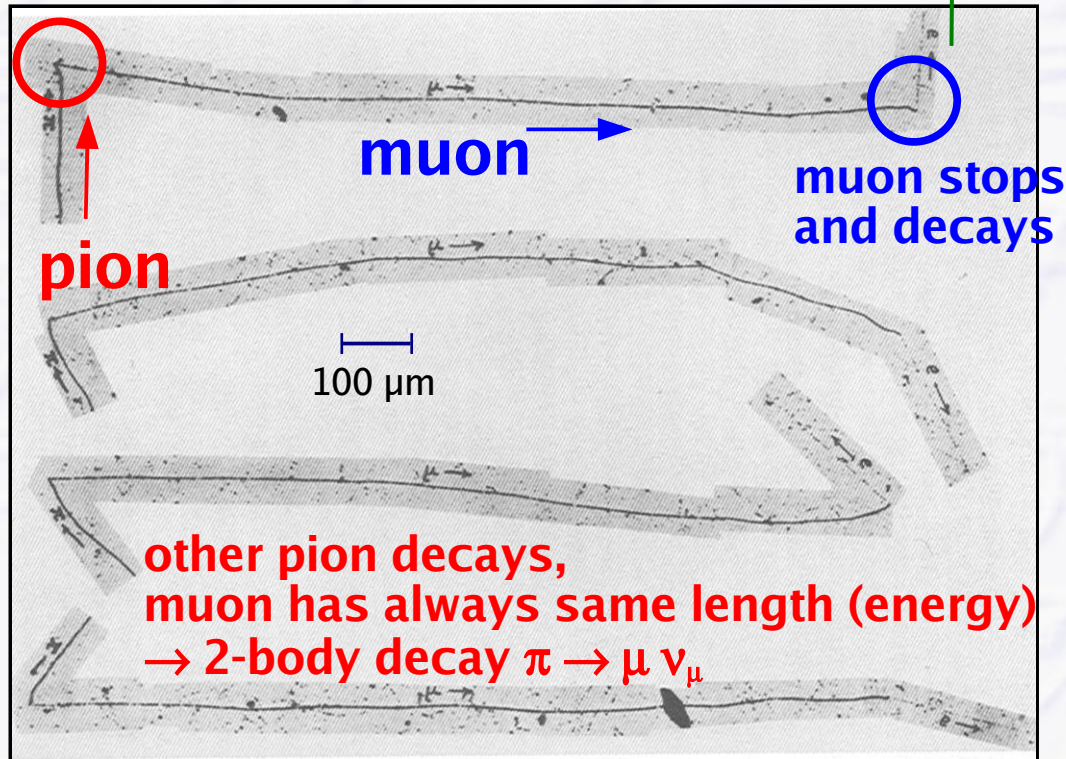
- Discovery of the **pion** in cosmic rays by Cecil Powell 1947 (Nobel Prize 1950)
- Discovery of the **kaon** 1949 (G. Rochester)

Cecil Powell



pion stops and decays

electron





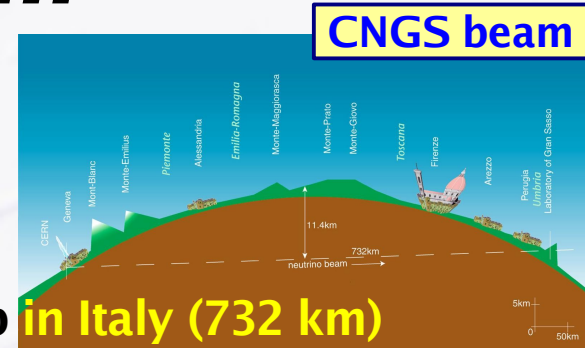
# Nuclear Emulsion III

- Still used in actual experiments with highest precision requirements over a large volume

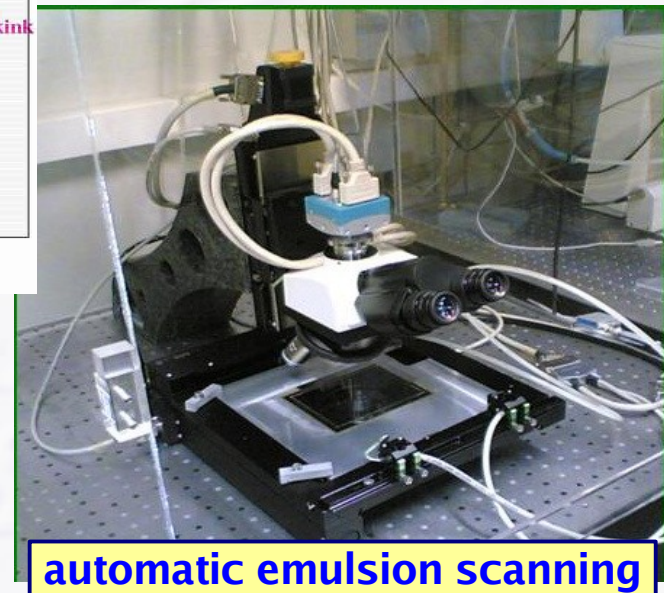
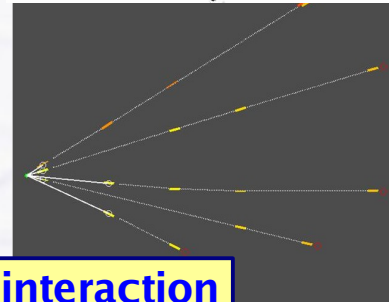
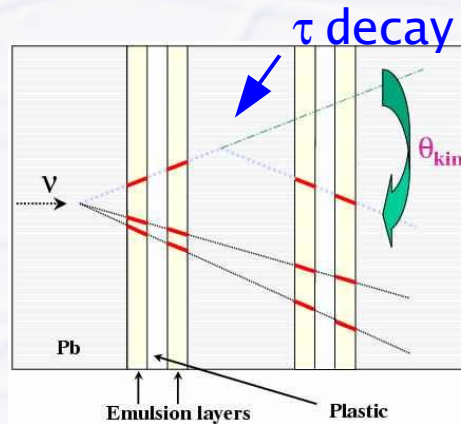
→  $\nu_\mu$  beam sent from CERN to Gran Sasso Underground lab

→ OPERA experiment is searching for  $\nu_\tau$  appearance after neutrino oscill.  $\nu_\mu \rightarrow \nu_\tau$

- need to reconstruct  $\tau$  decays ( $\nu_\tau + N \rightarrow \tau^- + X$ ) (few  $\sim 100 \mu\text{m}$  track length)
- 235'000 “bricks” (1.7 ktons) of lead + emulsion sheets



OPERA at Gran Sasso



# Bubble Chamber I

● Intended 1952 by Donald Glaser (Noble Prize 1960)

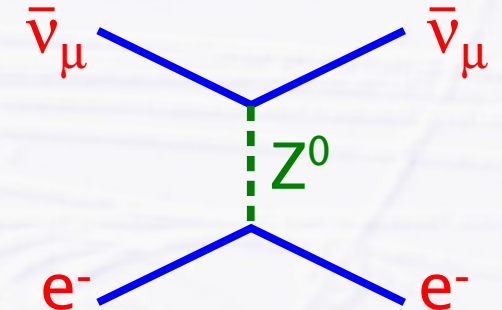
Donald Glaser



LBNL Image Library

- similar to cloud chamber
- chamber with liquid (e.g.  $H_2$ ) at boiling point (“superheated”)
- charged particles leave trails of ions
- formation of small gas bubbles around ions

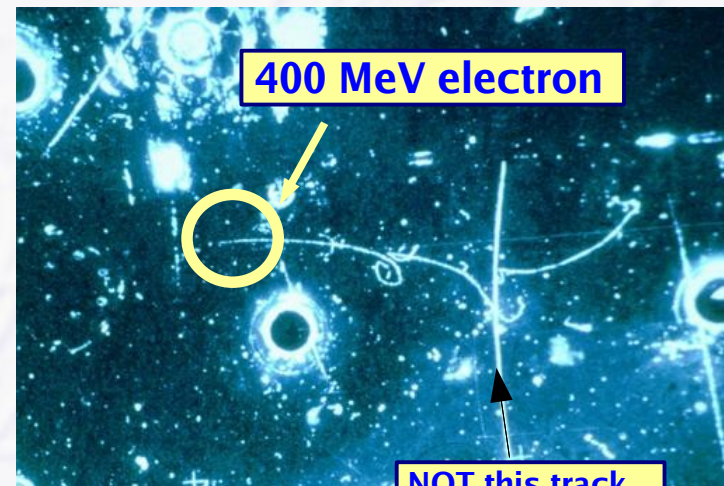
was used at discovery of the “neutral current”  
(1973 by Gargamelle Collaboration, no Noble Prize yet)



Gargamelle bubble chamber

CERN

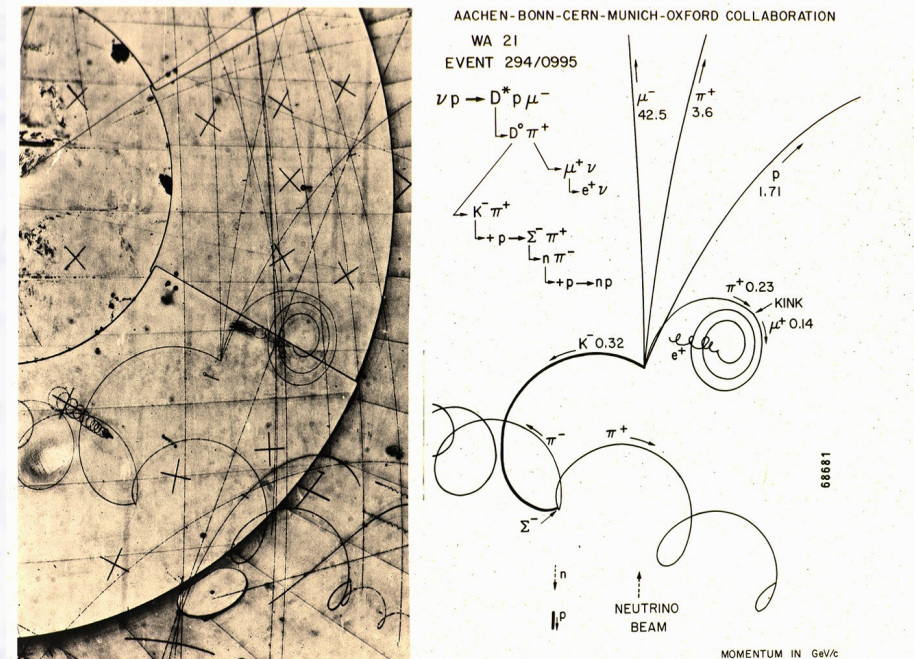
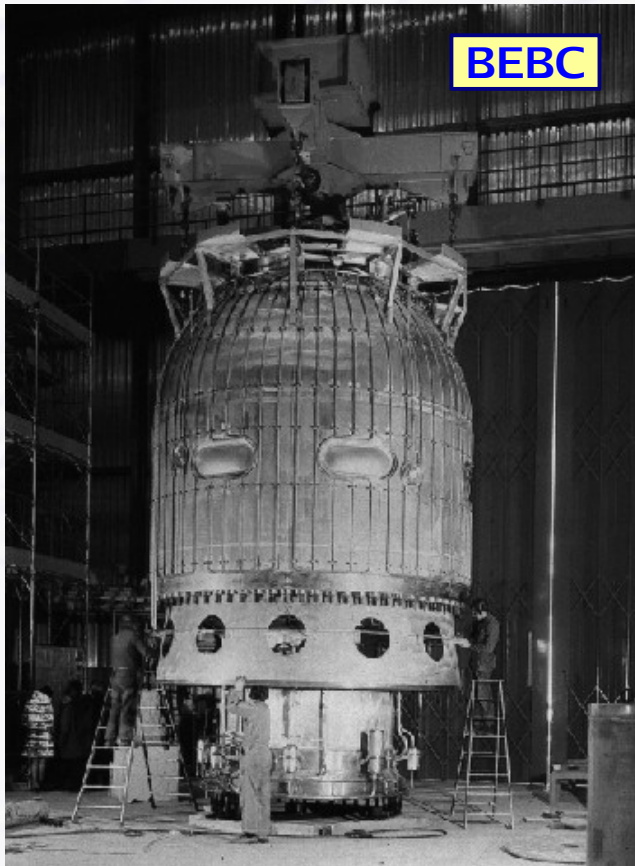
$\bar{\nu}_\mu \rightarrow$



CERN

# Bubble Chamber II

- BEBC (Big European Bubble Chamber) at CERN, 1973 – 1984
  - largest bubble chamber ever built (and the last big one...),  $\varnothing$  3.7 m
  - 6.3 million photographs taken, 3000 km of developed film
  - now displayed in permanent exhibition at CERN



production of  $D^*$  meson  
with long decay chain

# Bubble Chamber III

## Advantages of bubble chambers

- liquid is BOTH detector medium AND target
- high precision

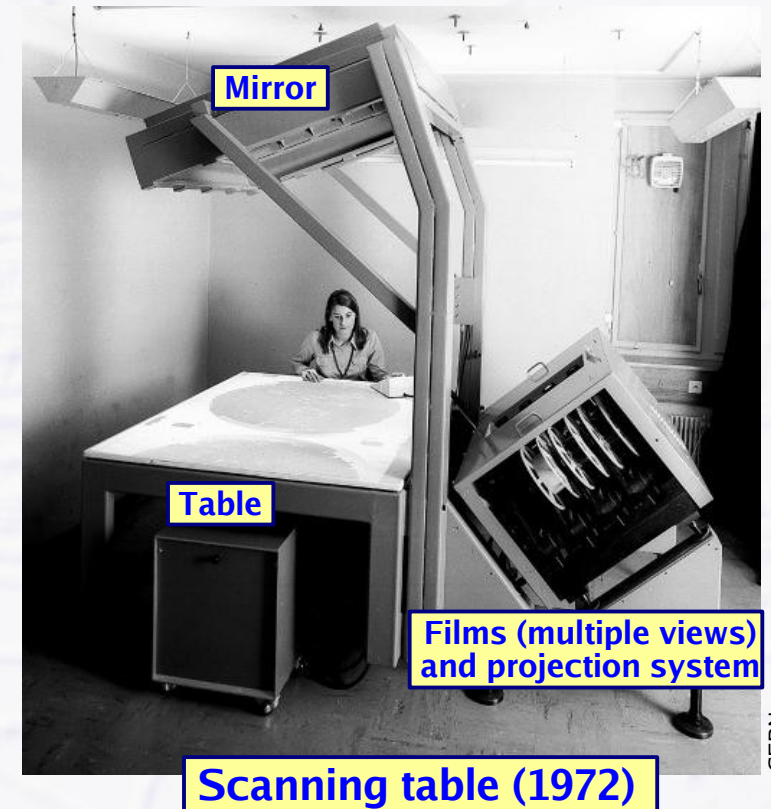
## Disadvantages

- **SLOW!!!**
  - event pictures taken with cameras on film
  - film needs to be developed, shipped to institutes and optically scanned for interesting events
- Need FASTER detectors (electronic!)

## However:

### Some important social side effects of bubble chamber era...

- scanning was often done by young “scanning girls” (students)...
- ...who later got married with the physicists...



# Early “Electronic” Detectors - Spintharoscope

- 1911: Ernest Rutherford + studied (elastic) scattering of  $\alpha$  particles on gold atoms (famous Rutherford experiment)

→ discovery of atomic nucleus:  
small (heavy) positively charged nucleus orbited by electrons

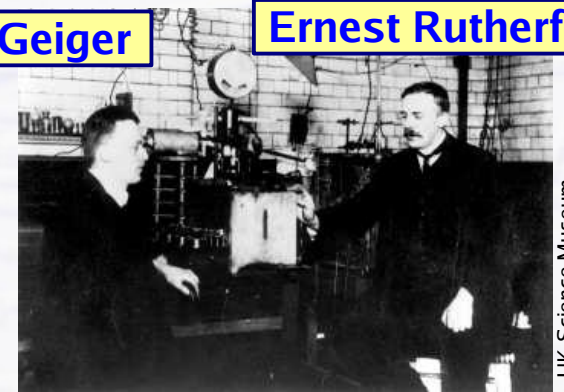
- Zinc sulfide screen with microscope (spintharoscope by William Crookes 1903) was used to detect scattered  $\alpha$  particles

→ light flash was observed by eye

- to increase light sensitivity, “bella donna” (from the deadly night shade plant = Tollkirsche) was often used to open eye's pupil

Hans Geiger

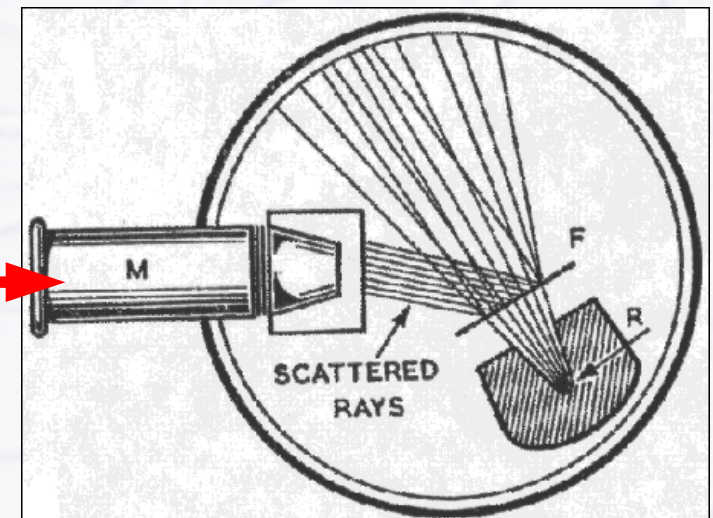
Ernest Rutherford



UK Science Museum

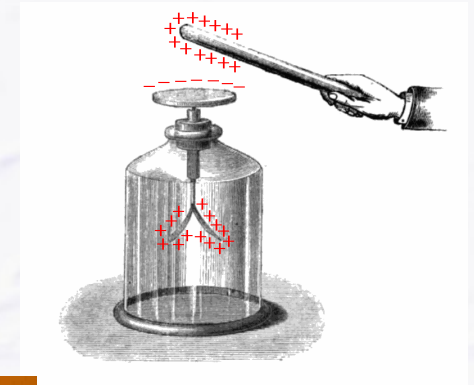


deadly night shade



# Early Electronic Detectors - Electroscope

- Gold-leaf electroscope already invented 1787 by Abraham Bennet
- End of 19<sup>th</sup> century raising interest on electricity in gases



→ cathode ray tubes, glow discharges



early cathode ray tube

→ observation:

charged electroscope is losing its charge in dry air after some time

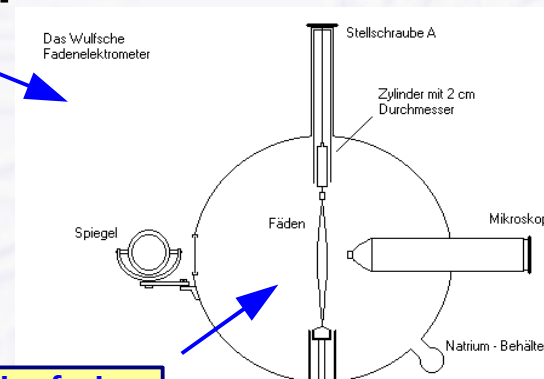
- source of conductivity? ionisation by recently discovered radioactivity?

- Victor Hess discovered cosmic rays 1912 (Nobel Prize 1936)

→ used calibrated string electrometer by Theodor Wulf

→ found increasing ionisation at higher altitudes at a series of balloon ascents

- not related to sun radiation!



pair of wires

Victor Hess in balloon

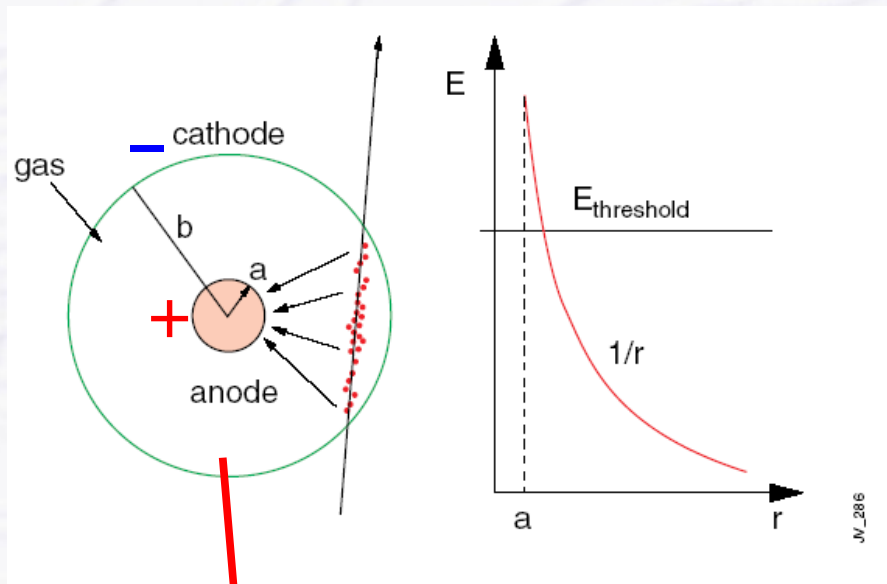


# Geiger-Müller Tube

## ● The Geiger-Müller tube (1928 by Hans Geiger and Walther Müller)

→ Tube filled with inert gas (He, Ne, Ar) + organic vapour

→ Central thin wire (20 – 50  $\mu\text{m}$   $\varnothing$ ), high voltage (several 100 Volts) between wire and tube



→ Strong increase of E-field close to the wire

- electron gains more and more energy

→ above some threshold ( $>10$  kV/cm)

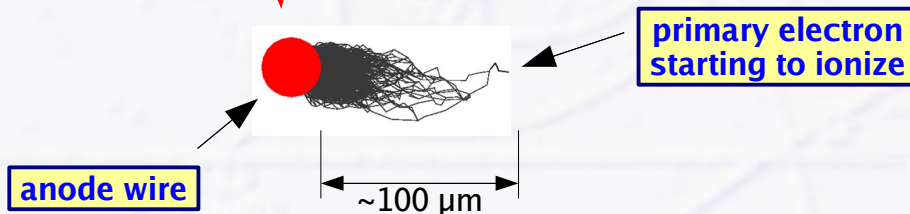
- electron energy high enough to ionize other gas molecules

- newly created electrons also start ionizing

→ **avalanche effect**: exponential increase of electrons (and ions)

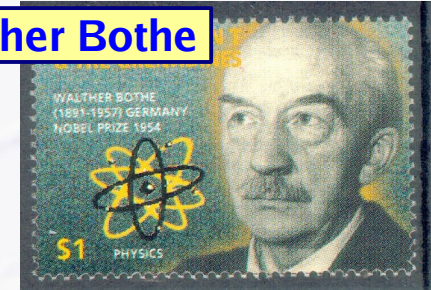
→ measurable signal on wire

- organic substances responsible for “quenching” (stopping) the discharge



# Coincidence Units

Walther Bothe

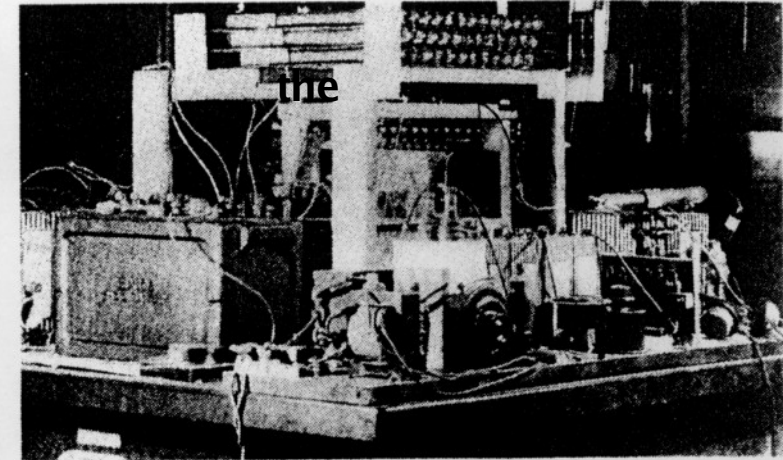


- “Zur Vereinfachung von Koinzidenzzählungen”, Walther Bothe 1929 (Nobel Prize 1954)

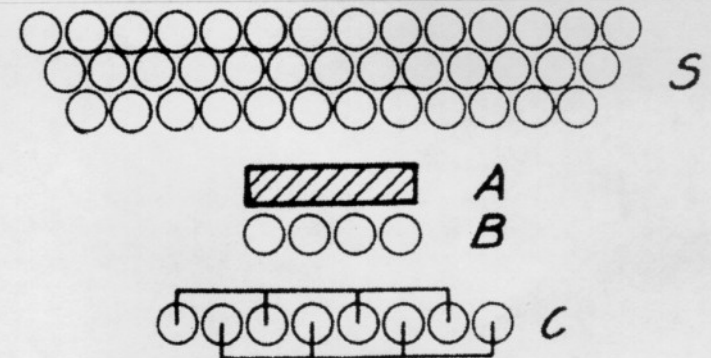
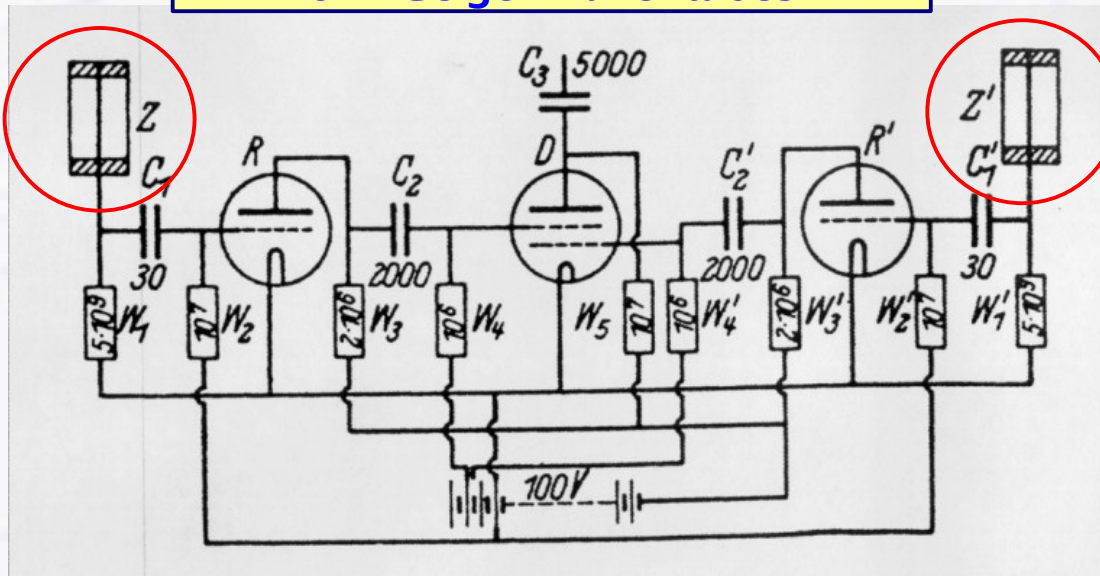
→ single tube has no information on direction of incoming particle

- two or more tubes giving signals within the same time window give direction
- also information if two particles come from the same decay

cosmic ray telescope 1934



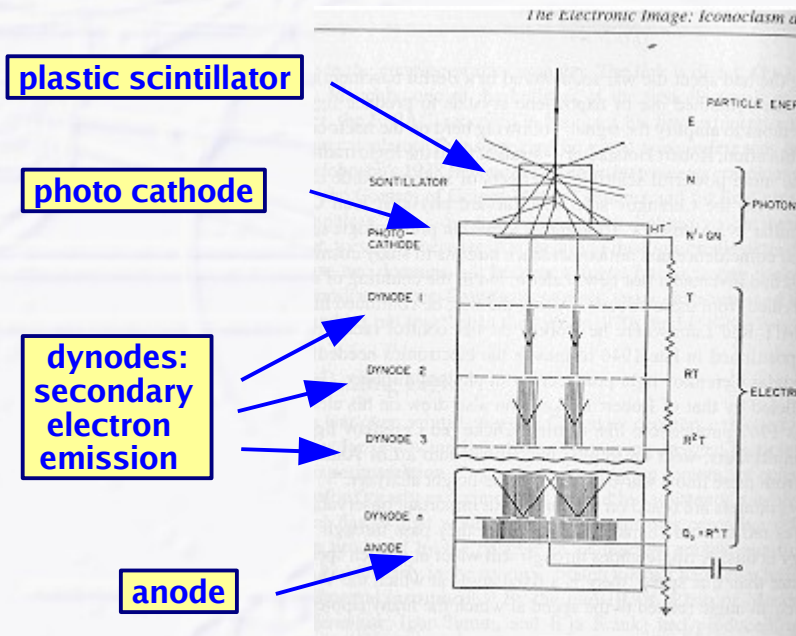
coincidence unit with vacuum tubes for 2 Geiger-Müller tubes





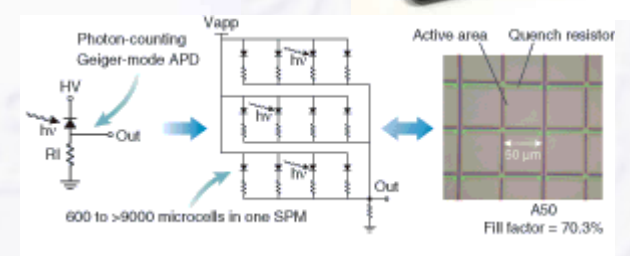
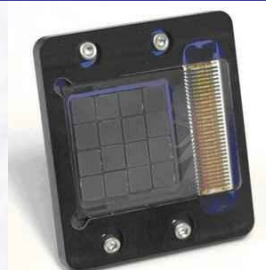
# Photo Multiplier Tubes (PMT)

- **Invented 1934 by Harley Janssen and Bernard Salzberg (RCA Cooperation)**
  - based on photo effect and secondary electron emission
  - sensitive to single photons, replaced human eye + belladonna at scintillator screen
- ➔ **first device had gain  $\sim 8$  only but already operated at  $>10$  kHz (human eye: up to 150 counts/minute for a limited time)**
- nowadays still in use everywhere, gain up to  $10^8$
- recent developments: multi-anode (segmented) PMTs, hybrid and pure silicon PMs



**classic PMT**

**Silicon PM = array of avalanche photo diodes**



# Multi Wire Proportional Chambers I

- Geiger-Müller tube just good for single tracks with limited precision (no position information inside tube)

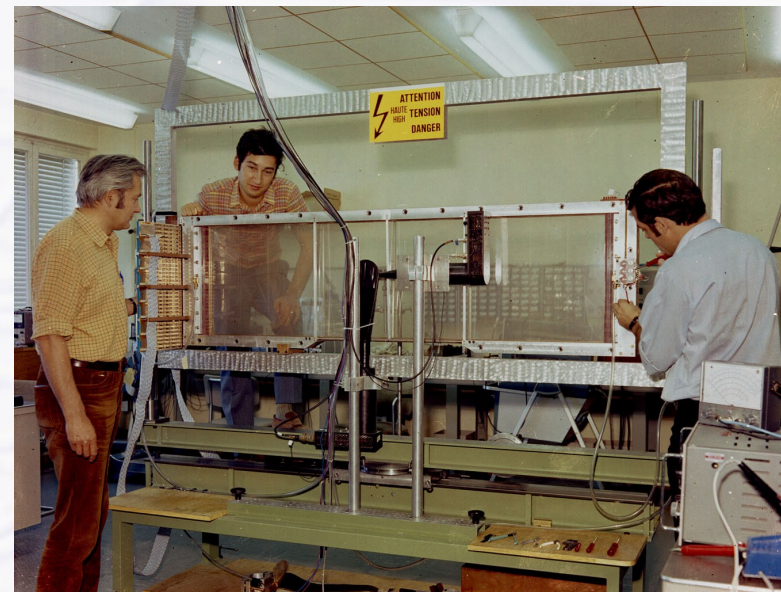
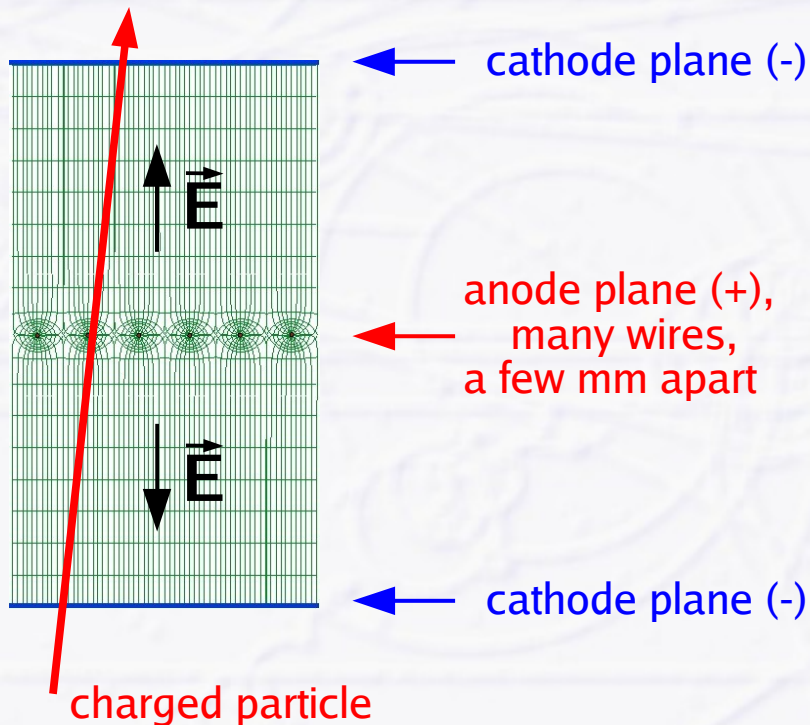
→ in case of more tracks more tubes are needed or...

- Multi Wire Proportional Chamber (MWPC) (1968 by Georges Charpak, Nobel Prize 1992)

→ put many wires with short distance between two parallel plates



Georges Charpak



Georges Charpak, Fabio Sauli and Jean-Claude Santiard

# Multi Wire Proportional Chambers II

## ● Multi Wire Proportional Chamber (MWPC)

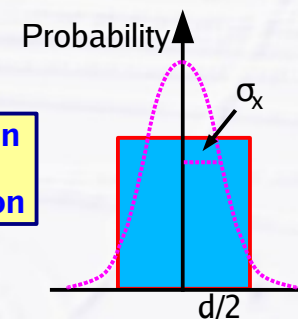
- was first electronic device allowing high statistics experiments
- with multiple channels and reasonable resolution

## ● Typically several 100 – 1000 wires, ~ 1 mm spacing

- if charged particle is passing the MWPC → one wire gives signal

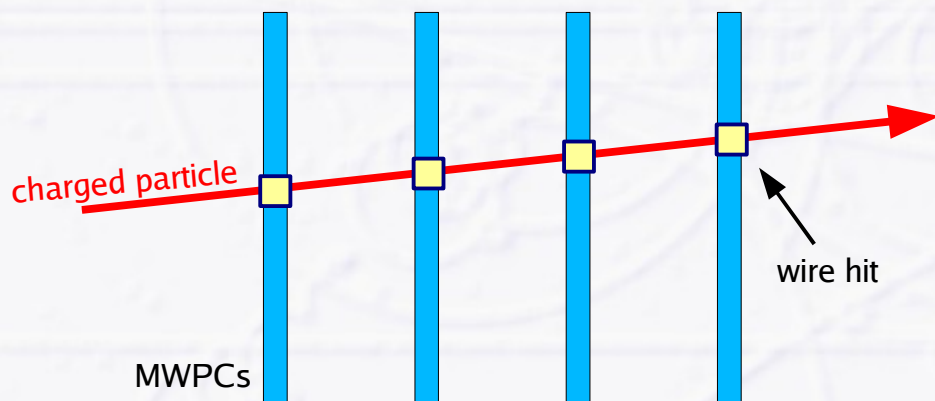
- resolution:  $\sigma_x \approx \frac{d}{\sqrt{12}}$  e.g. for  $d = 1 \text{ mm} \rightarrow \sim 300 \mu\text{m}$

we don't know where the particle went through within the 1 mm spacing = "flat" probability distribution, this is the width of an equivalent Gaussian distribution



## ● If many MWPCs are put one after each other

- each particle creates one point per MWPC ( $\sim 300 \mu\text{m}$  resolution per point)



can reconstruct track with e.g. 4 points

one coordinate only, use additional MWPCs tilted by  $90^\circ$  to get other coordinate

# Drift Chamber

## Resolution of MWPCs limited by wire spacing

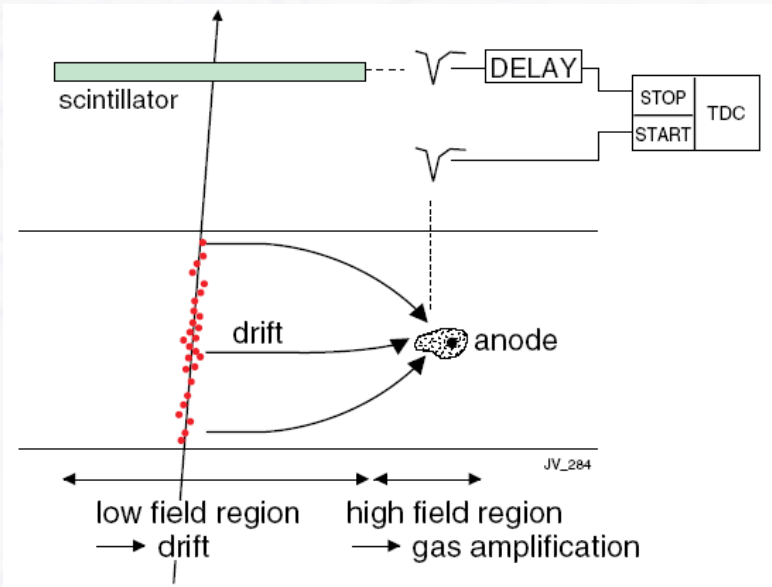
→ better resolution → shorter wire spacing → more (and more) wires...

- larger wire forces (heavy mechanical structures needed)
- (too) strong electrostatic forces when wires too close to each other

## Solution by A. H. Walenta, J. Heintze, B. Schürlein 1971

→ obtain position information from drift time of electrons (fewer wires needed)

- drift time = time between primary ionization and arrival on wire (signal formation)



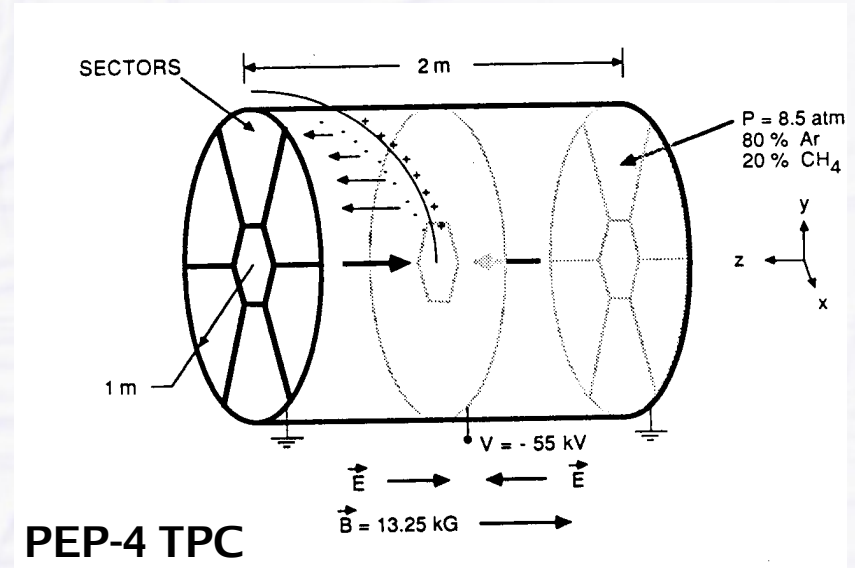
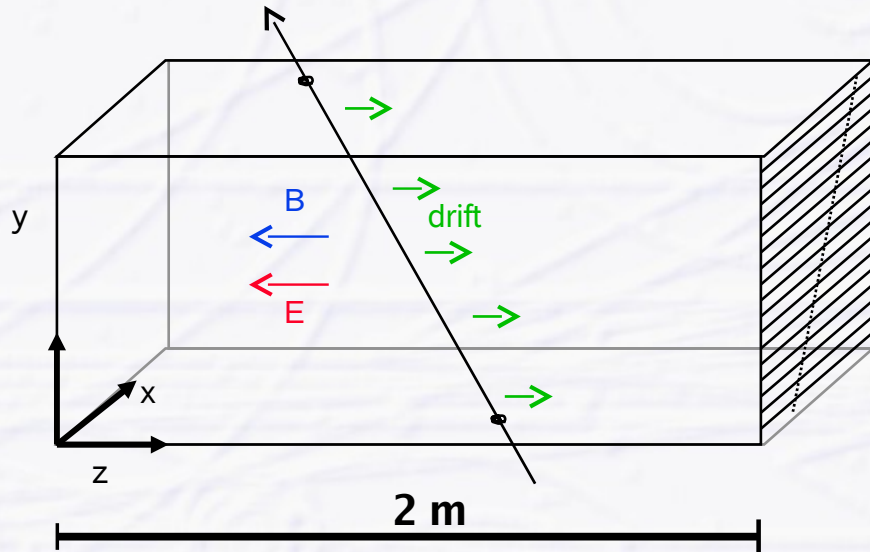
**start signal** (track is passing drift volume)  
has to come from **external source**:  
scintillator or beam crossing signal

- Need to know drift velocity  $v_D$  to calculate distance  $s$  to wire (= track position within the detector)

$$s = \int_{t_{start}}^{t_{stop}} v_D dt$$

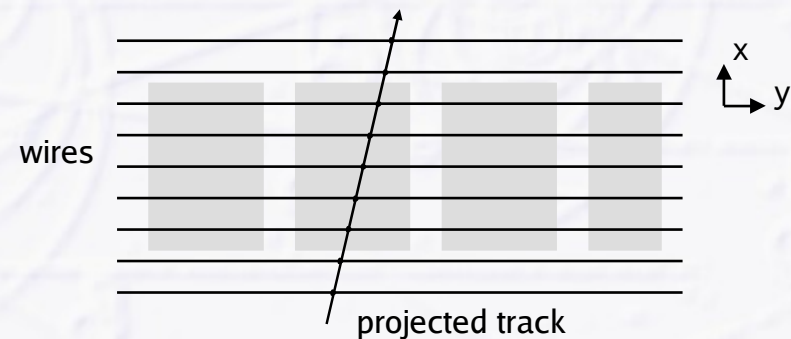
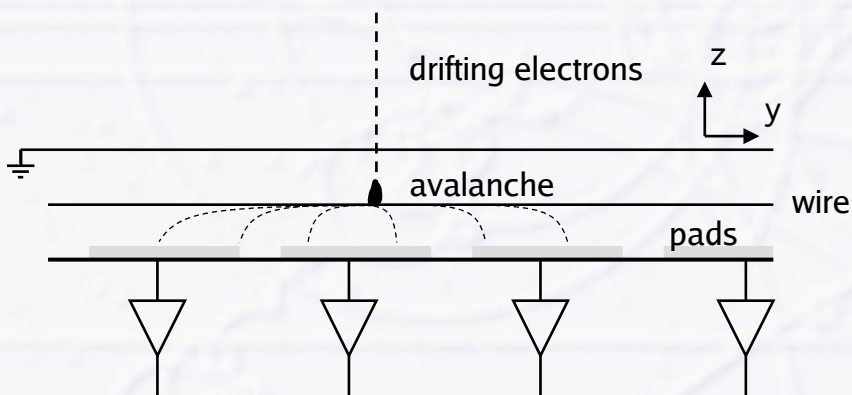
# Time Projection Chamber (TPC)

- A 3D-imaging chamber with rather long drift length



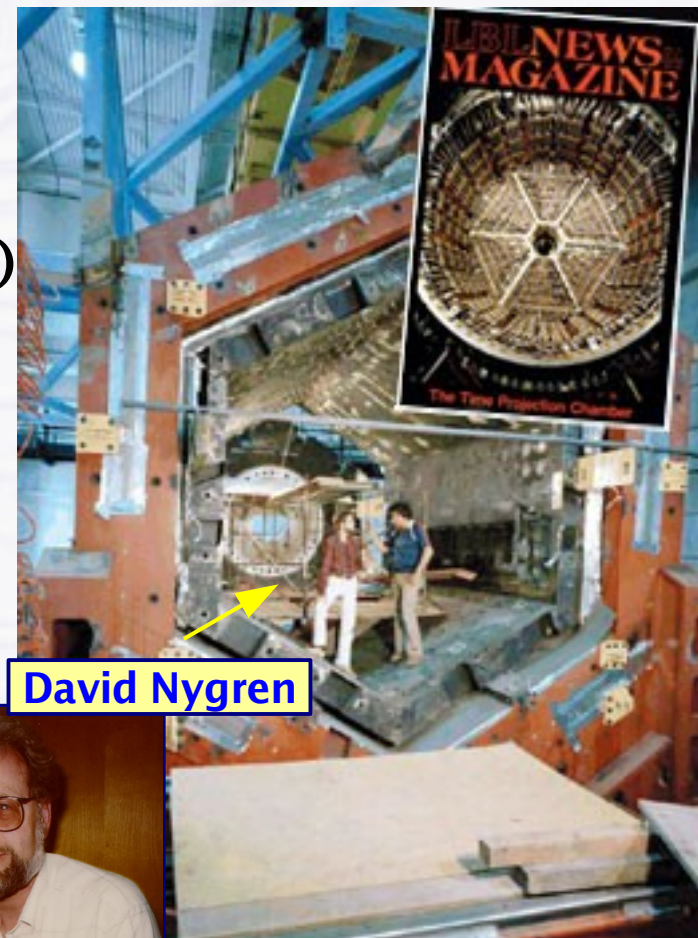
→ homogeneous B- and E-fields

→ anode plane equipped with MWPC wire chambers

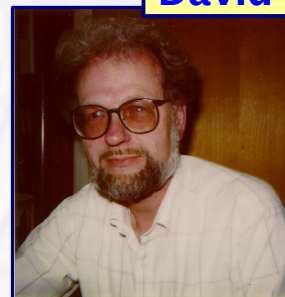


# Time Projection Chamber (TPC)

- Invented by David Nygren (Berkeley) in 1974
- Proposed as central tracking device for the PEP-4 detector at the PEP  $e^+e^-$  collider at SLAC 1976
- More (and even larger) TPCs were built or are planned at other colliders
  - TRISTAN (KEK, 2 x 32 GeV  $e^+e^-$ , 1986 – 1995)
    - TOPAZ
  - LEP (CERN, 2 x 104 GeV  $e^+e^-$ , 1989 – 2000)
    - ALEPH, DELPHI
  - RHIC (BNL, 2 x 100 GeV/nucleus, 2001 – )
    - STAR
  - LHC pp and Pb-Pb collider (CERN)
    - ALICE
  - ILC  $e^+e^-$  collider
    - ILD



David Nygren



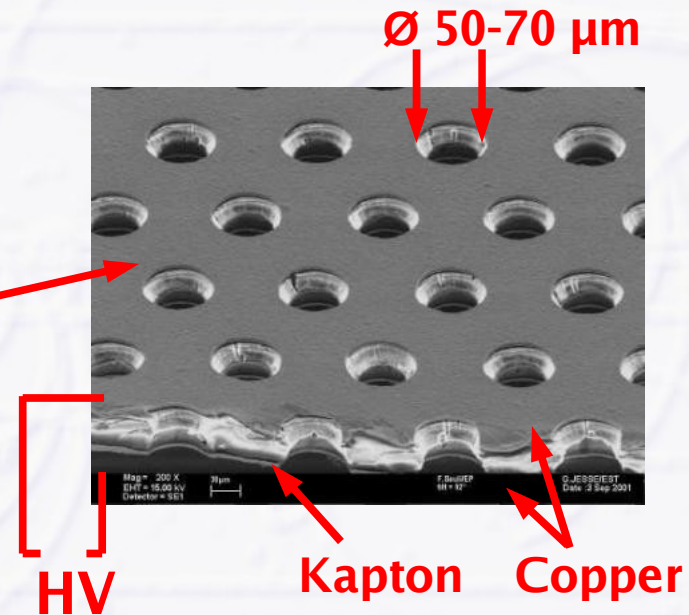
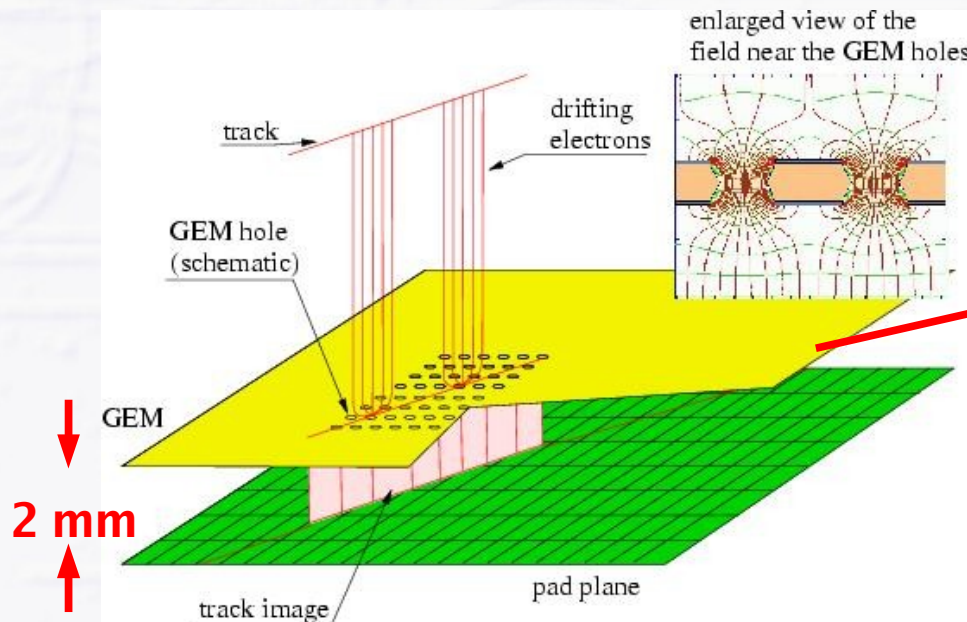
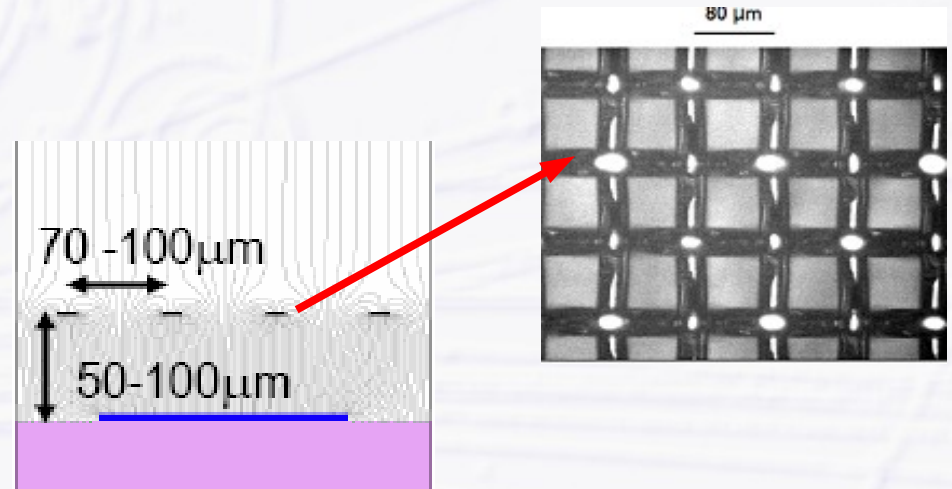
# Recent Developments: Micro Pattern Gas Detectors (MPGD)

## ● Replace wires at TPC with Micro Pattern Gas Detectors

- **MicroMegas** (metallic micromesh)
- **GEM** (Gas Electron Multiplier)

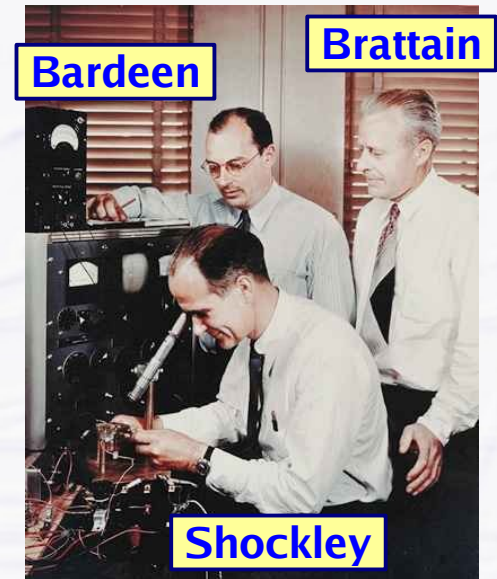
## ● Concept

- **2D structures** with holes + underlying pads
- **Gas amplification inside holes**, collect electrons on small pads, few mm<sup>2</sup>

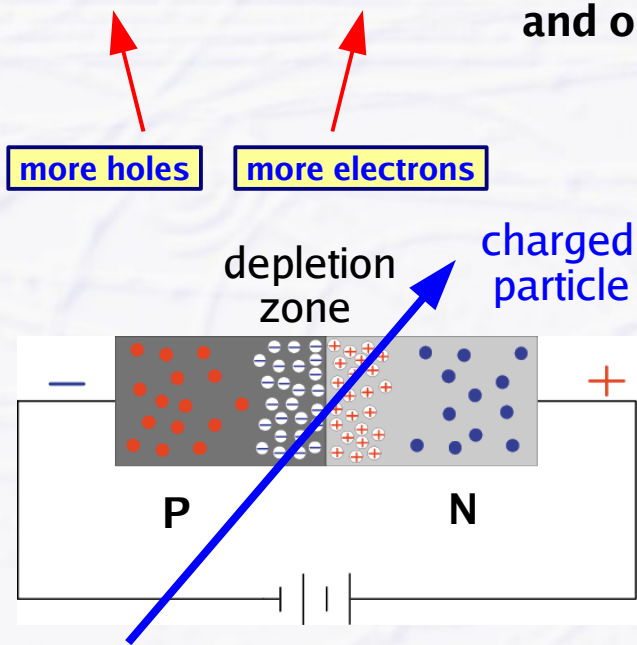


# Solid State Detectors

- First transistor was invented 1947 by William B. Shockley, John Bardeen and Walter Brattain (Nobel Prize 1956)



- transistors and diodes became common soon after
- ➔ Germanium diodes were used for particle detection
- p-type and n-type doped silicon material is put together and operated with **reversed voltage**



- ➔ around junction of p- and n-type material depletion zone is created
- ➔ zone free of charge carriers
  - no holes, no electrons
  - thickness of depletion zone depends on voltage, doping concentration

charged particle typically creates 20'000 – 30'000 electron/hole pairs in 300  $\mu\text{m}$  thick material -> sufficient signal size

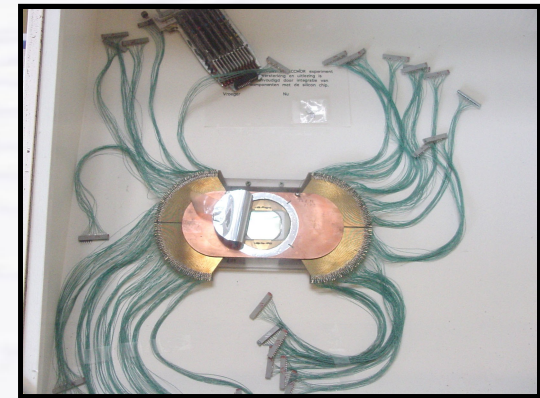
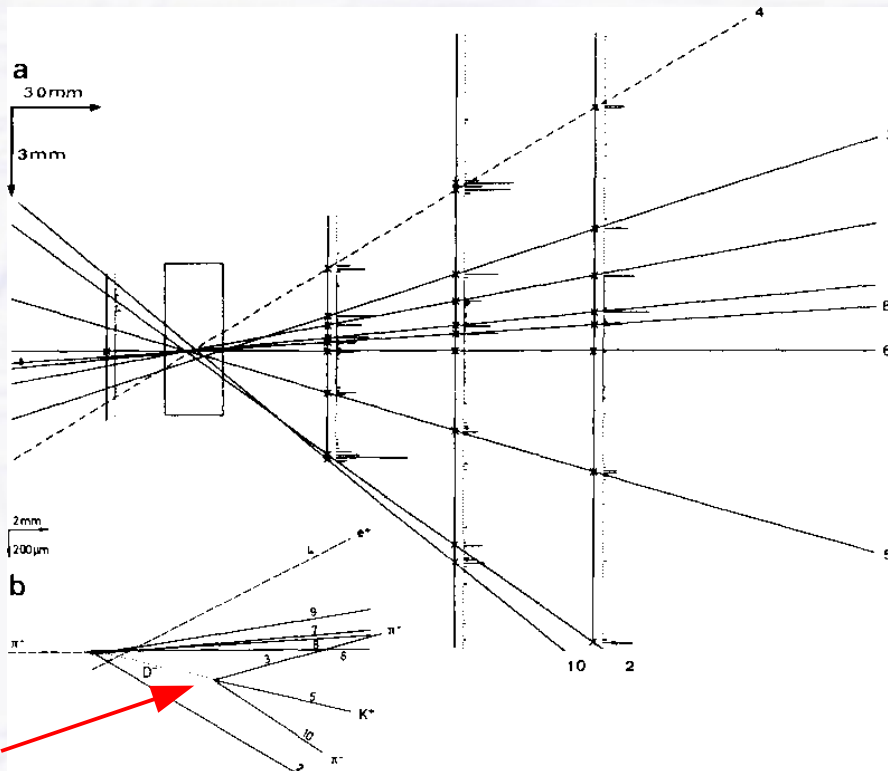


# The first Silicon Strip Detector

- First operational silicon strip detector used in an experiment (NA11 at CERN) by J. Kemmer, R. Klanner, G. Lutz et al. 1983

→ G. Lutz was founder of MPI Halbleiterlabor in Munich Max-Planck-Institut  
→ NA11 aimed to search for new short lived particles Halbleiterlabor

- first observation of  $D_s$   
many branching ratio and lifetime measurements



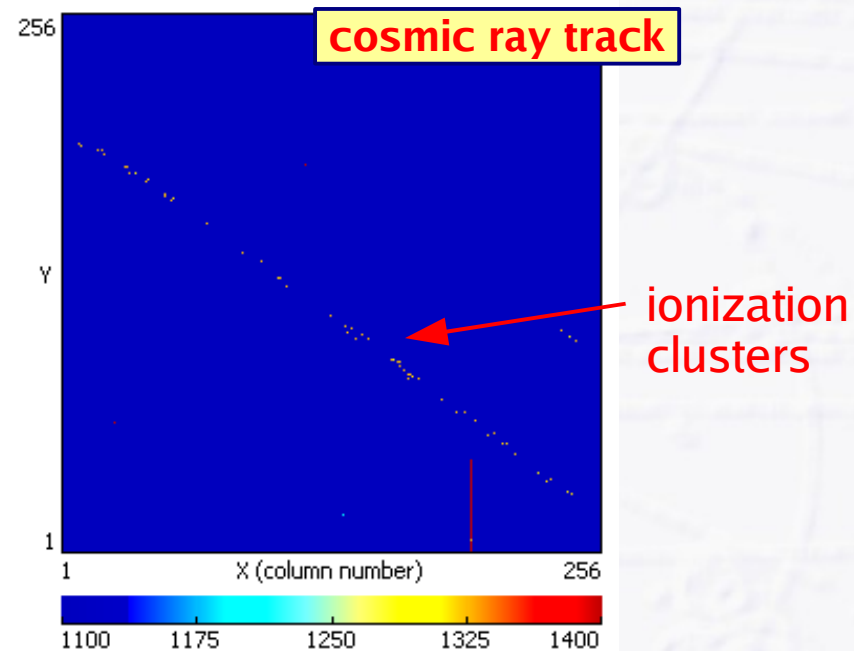
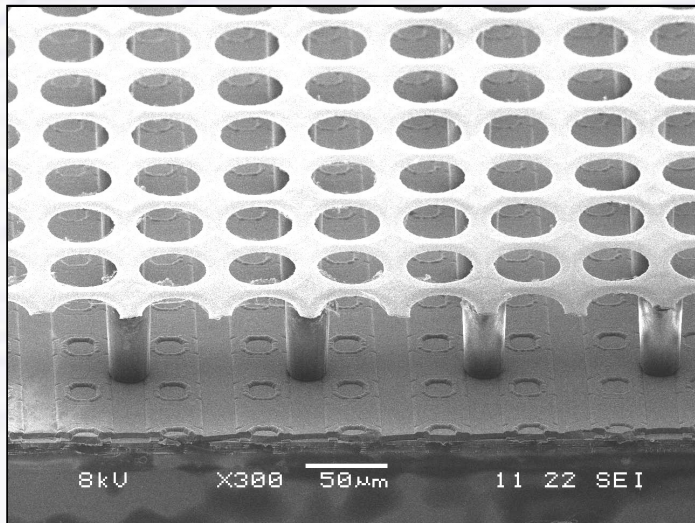
8 silicon strip planes  
(4 groups of 2 planes each  
with tilted strips to measure  
xy coordinate)

24 x 36 mm<sup>2</sup> size per chip  
1200 strips, 20 µm pitch  
240 read-out strips  
4.5 µm single hit resolution

$D^-$  decay

# Recent Developments: Hybrid Technologies

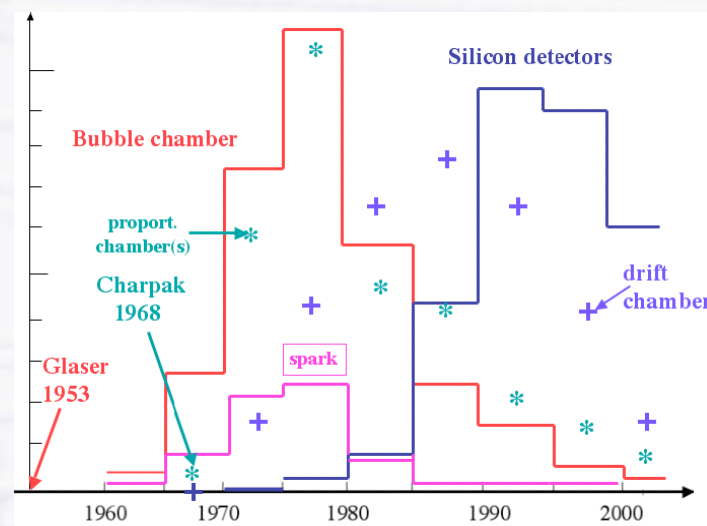
- Combine MPGD gaseous detector with silicon pixel detector
- Use MediPix2/TimePix chip as active TPC “padplane” for ILC detector
  - MediPix2 = 256x256 pixels with 55x55  $\mu\text{m}^2$  size for medical applications (X-ray film replacement)
  - MicroMegas mesh (provides gas amplification) integrated on top of pixel chip



- Individual ionization visible:  
the **digital Bubble Chamber** is in reach

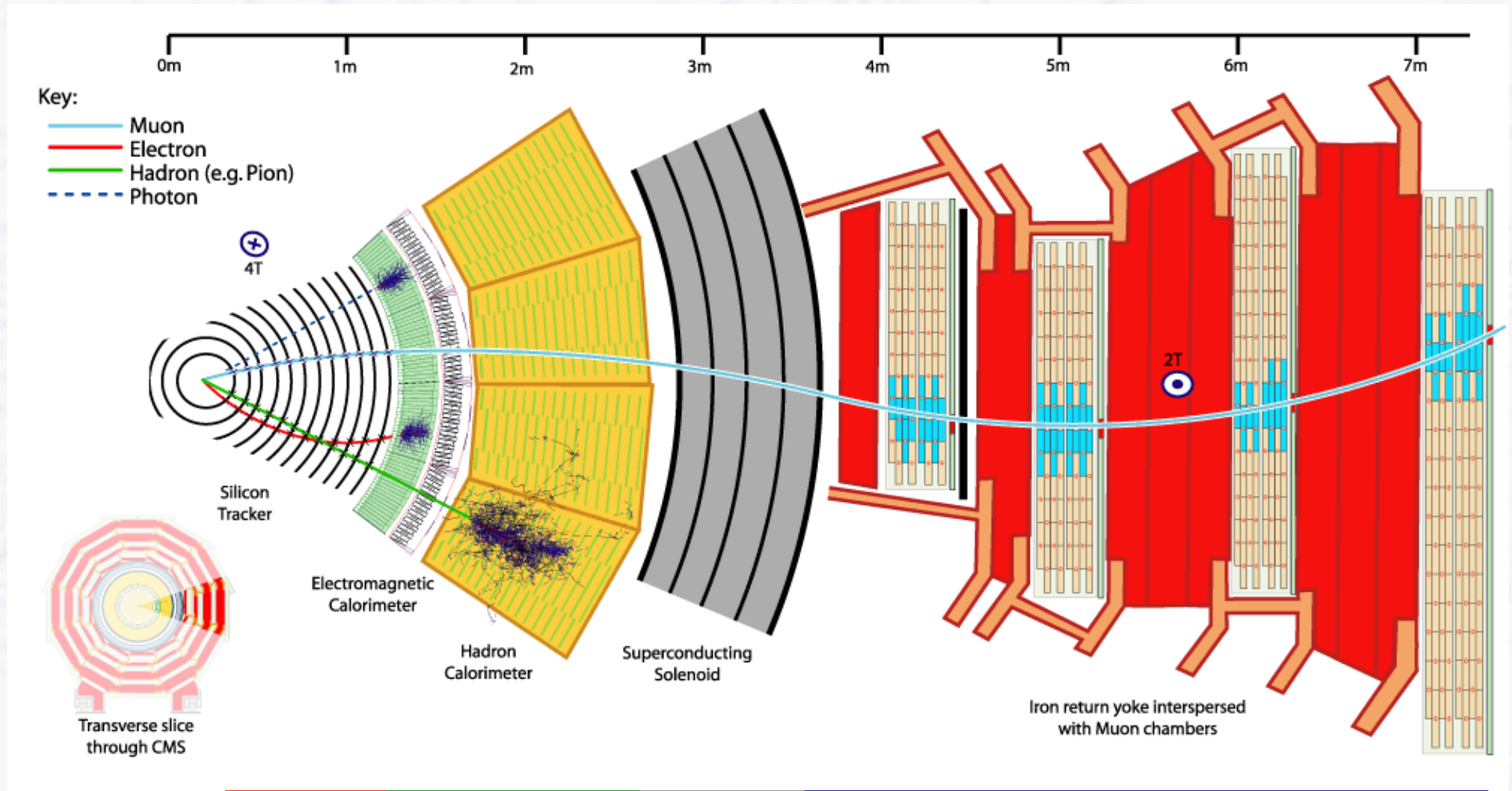
# Detector History

- **Cloud Chambers, Nuclear Emulsions + Geiger-Müller tubes dominated until the early 1950s**
  - Cloud Chambers now very popular in public exhibitions related to particle physics
- **Bubble Chambers had their peak time between 1960 and 1985**
  - last big bubble chamber was BEBC at CERN
- **Wire Chambers (MWPCs and drift chambers) started to dominate since 1970s**
- **Since late 1980s solid state detectors are in common use**
  - started as small sized vertex detectors (at LEP and SLC)
  - now ~200 m<sup>2</sup> silicon surface in CMS tracker
- **Most recent trend: hybrid detectors**
  - combining both gaseous and solid state technologies



# A typical Today's Particle Detector

## ● Cut-away view of CMS



Tracker

Calorimeter

Coil

Muon Detector and iron return yoke