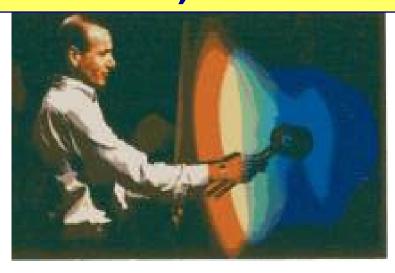


## Introduction to Particle Physics (for non physics students) 4. UNIFIED UNIVERSE (no strings attached)



PROFESSOR FRANK CLOSE EXETER COLLEGE UNIVERSITY OF OXFORD



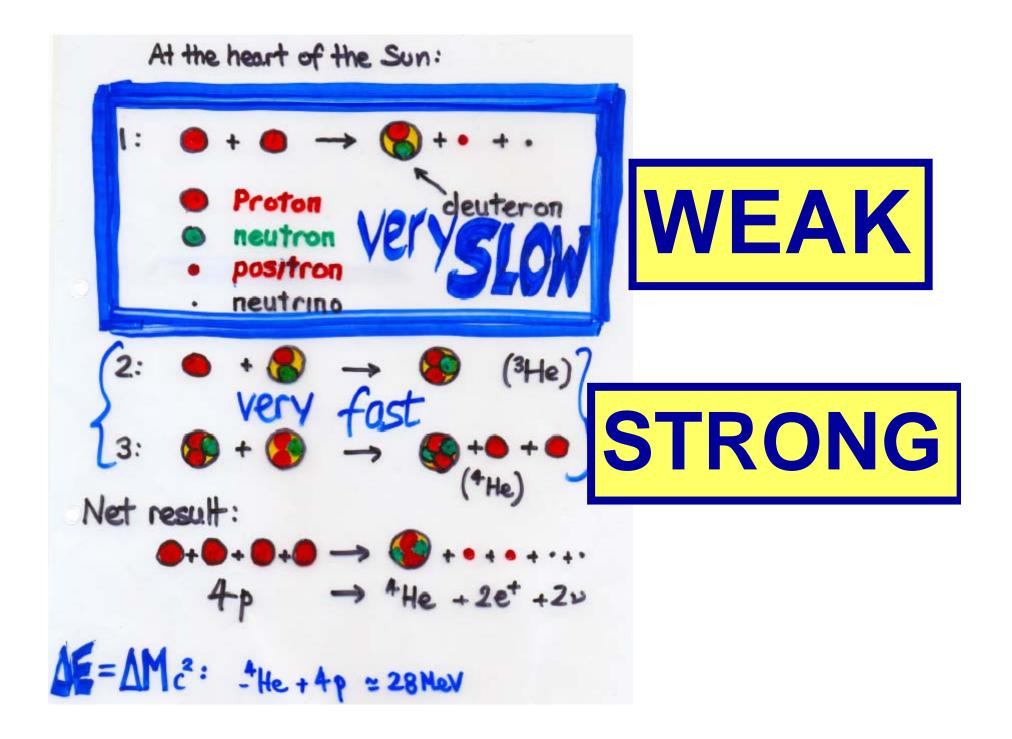
(remember	that waves -> particle	(2
NAME	action	CARRIE
Gravity	no su zassi	graviton ?
Electromo	solids stops us falling to cendre of Garth	by by by the by
Weak	β-radioactivity P→He in Sun	W+ W- Z
Strong	quatks glued inside p.n p.n in nuclei	gluons (g 8 different
Only the	weak force carriers	have MAS
M. ~	80 GeV/c2	
M~	91 Get/c2	

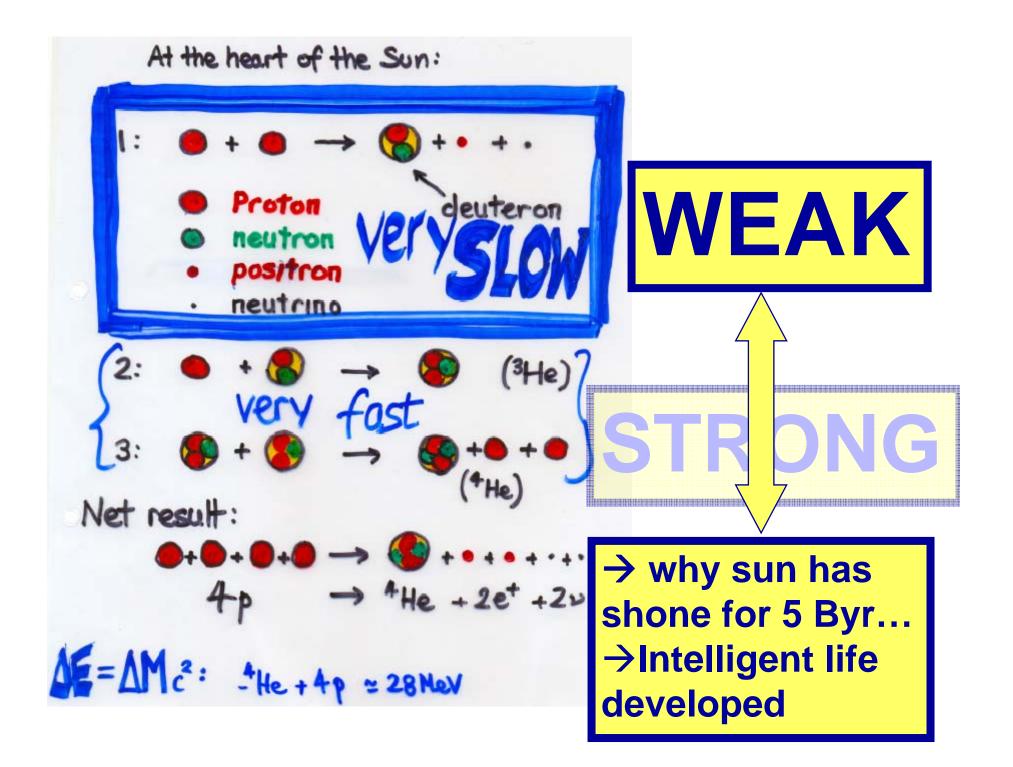
Feynman rules:

If energy E flows through the transmitted "virtual" particle (photon; Z) it costs 1/(E^2+M^2)

If E >> M the cost is 1/E^2....like the case of the photon

Only appears weak at low energy. Unified at high energy





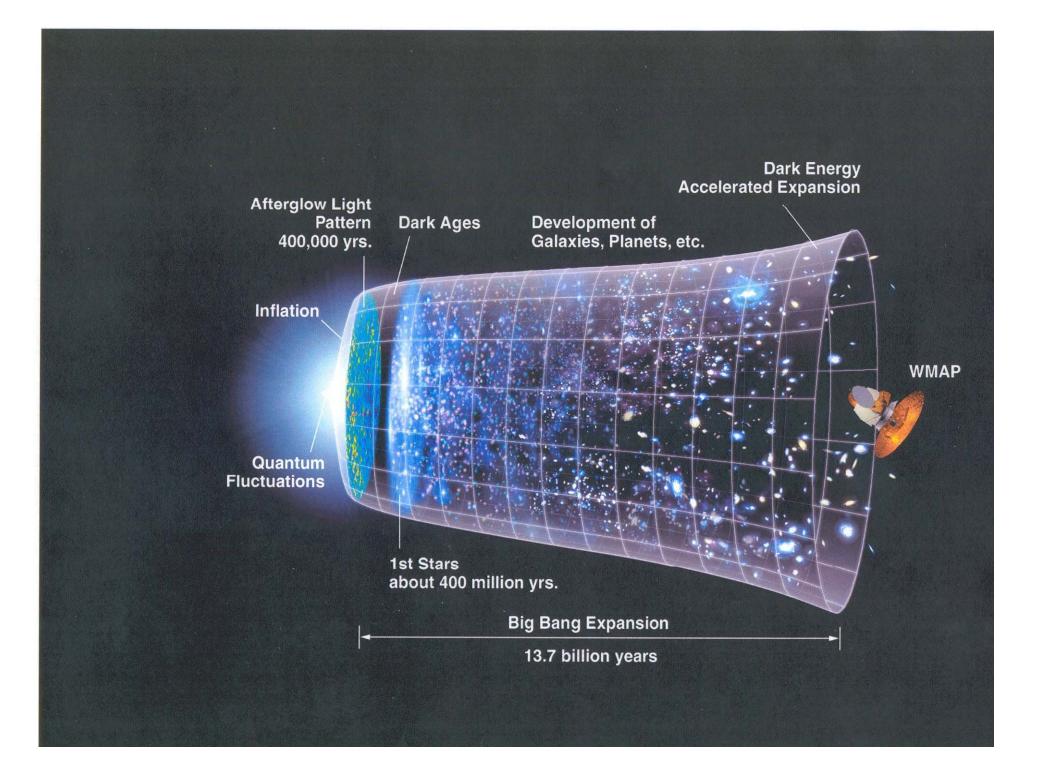
## The weak force is feeble in the Sun ...

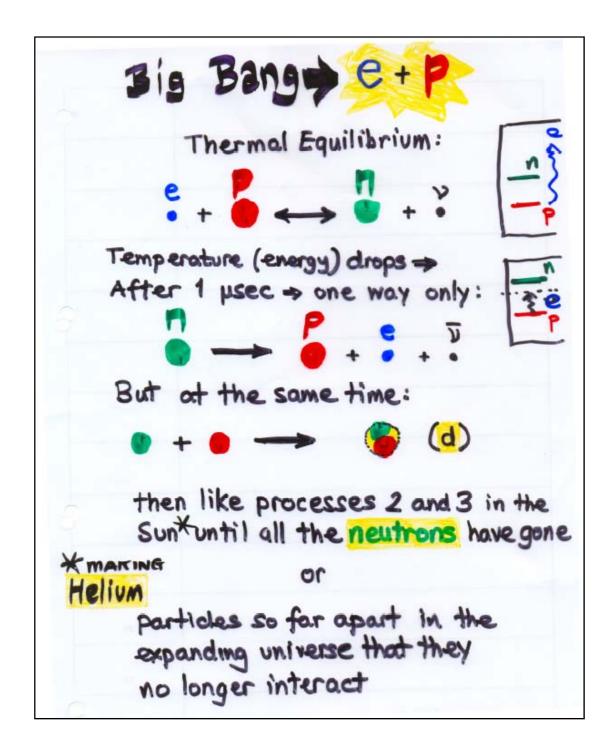
..because 10,000,000K ~ 1 keV << 80 GeV

...this is why the sun has stayed active long enough for us to have evolved and be having this conversation.

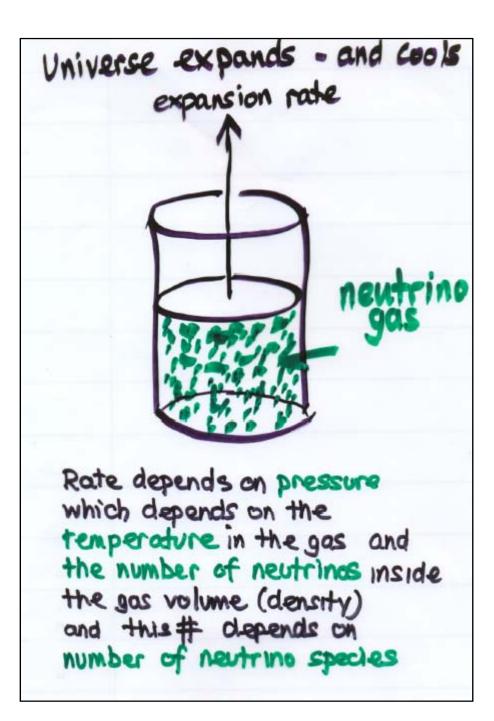
 $\rightarrow$ We exist because m(W) is not zero

→ Mass matters





T= I usec after BIG BANG ("the first fossils in the Universe) move at high speed and if they have mass they start clustering together -> contribute to formation of galaxies Billion > per atom > if m(v) > m(proton)/109 ~1eV they will dominate mass density of the Universe my big question for future of universe and its formation



T= 3 minutes after BIG BANG

75% protons 24% Helium Nuclei + small amount of deuterons + free electrons.

Helium abundance ; throces of the

Universe which depends on number of neutrino species

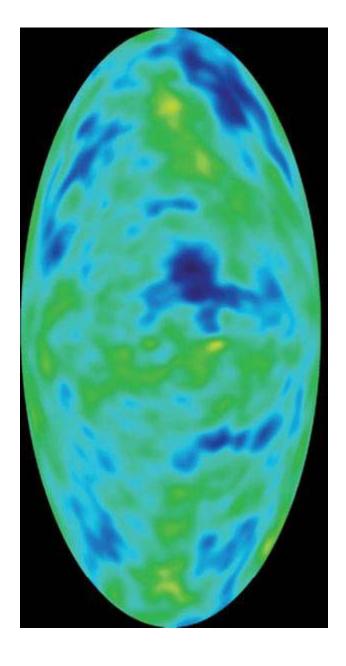
Deuterium abundance depends on density of "ordinary mather" in the Universe. T= 3 minutes after BIG BANG

75% protons 24% Helium Nuclei + small amount of deuterons + free electrons.

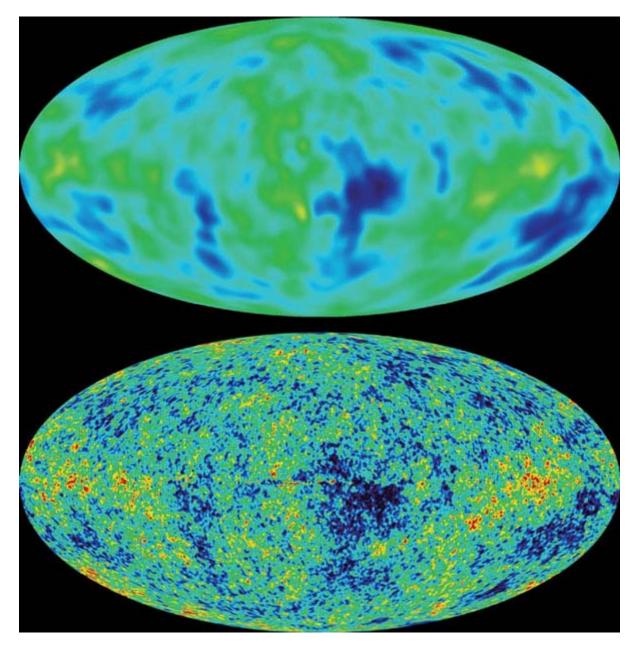
Helium abundance the light deal depends on expansion rate of the Universe which depends on number of neutrino species

Deuterium abundance depends on density of "ordinary mather" in the Universe. T= 3 minutes after BIG BANG 75% protons 24%. Helium Nuclei + small amount of deuterons + free electrons. Active abundance \*: " depends on expansion rate of the Universe which depends on number of neutrino species Deuterium abundance depends on density of "ordinary mather" in the Universe. density of ordinary matter << total in universe part of DARK MATTER puzzle

Time Passes. Temp drops 300,000 years later Exloced T< 104K electrons combine with nuclei and make neutral atoms electromagnetic radiation was set free Universe becomes transparent 10 years later Emag 2 stretched : Micromane Band. Black body background 3K (small fluctuations in Manuave rad = hints of proto structures, galaxier in early universe)



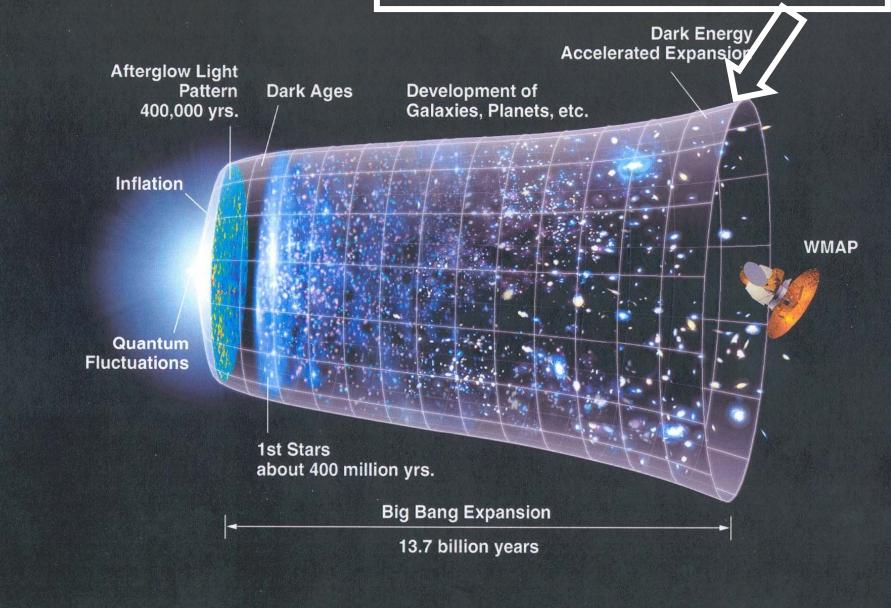
### 3K microwave bgnd now seen to have structure



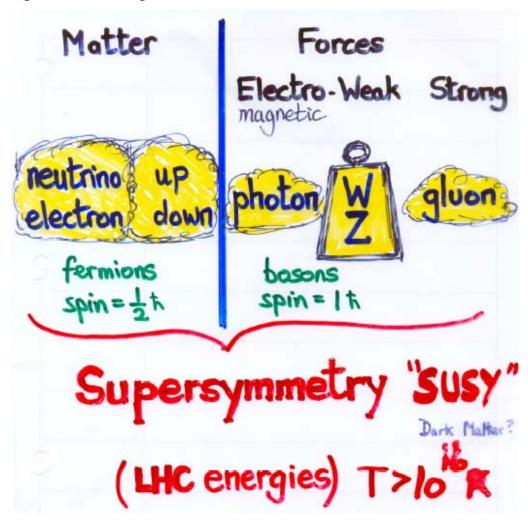
#### **COBE 2000**

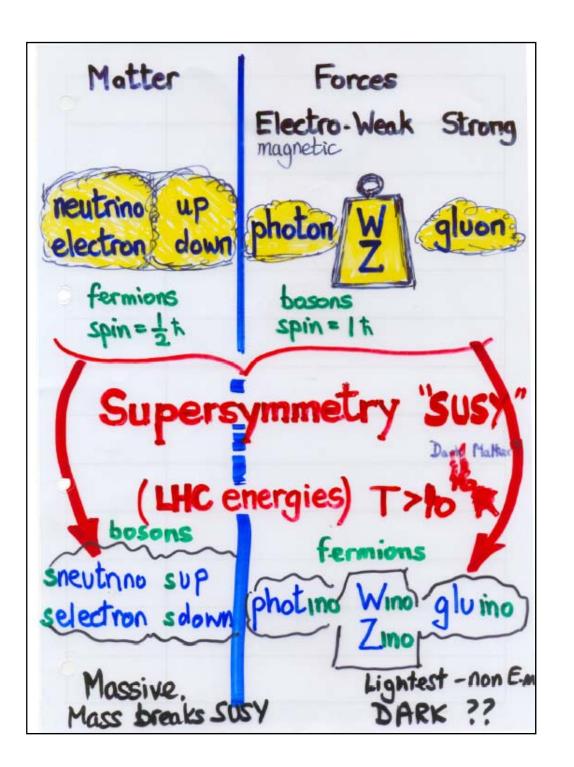
### **WMAP 2003**

## 5byr ago accelerated expansion = Dark Energy. What? Why?.....



#### **One further symmetry??**





## Whole new families to be found

Particle Physics @ CERN.  
Standard Model of PHes + Forces.  
Quarks + Leptons. Spin 1/2 fermions  
V W<sup>±</sup>Z gluons Spin 1 gauge basons  
Higgs Spin 0 bason  
High Energy 
$$\iff$$
 Early Universe  
Origins of mather.  
Structures + patterns at ESITEV  
Symmetry revealed at EZ O(TeV)  
Forces (and particles) unified - SUSY.  
Some current big Ph22les.  
Dark Mather, Solar V, Massive V?  
(all the same?)  
Why 3 generations 1 the  
What is difference between Mand M? [ Jame?  
. The Fifth Dimension

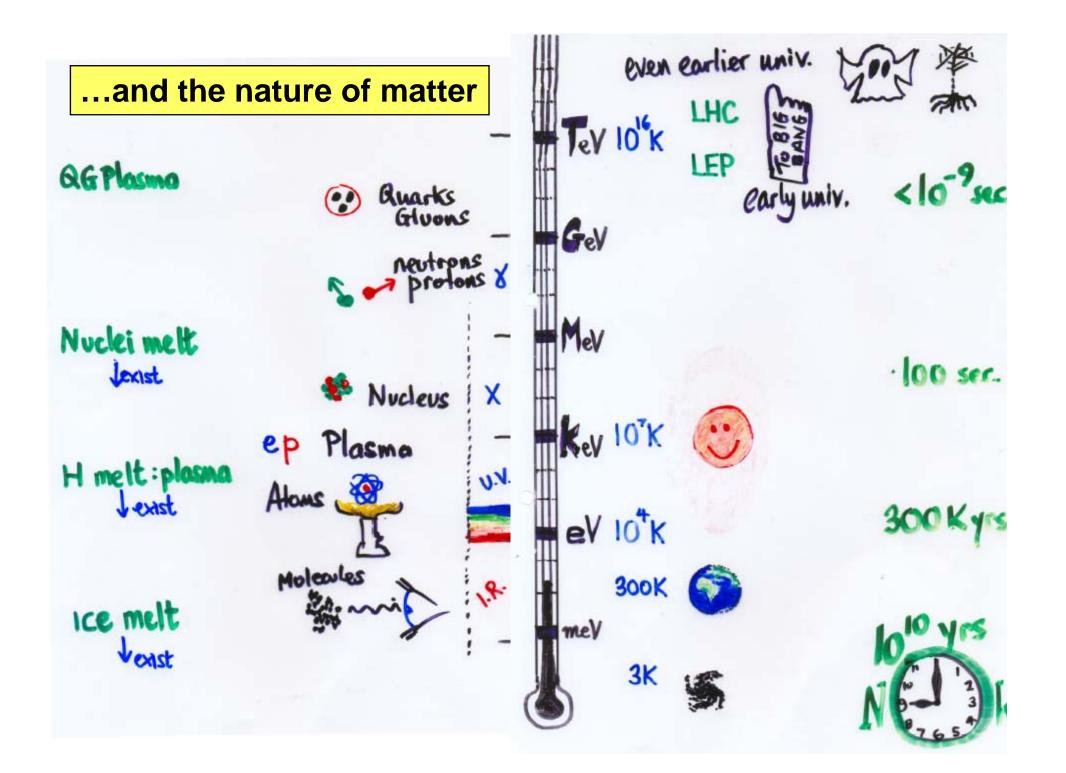
# Finale: A glimpse of the future recall from lecture 1.....

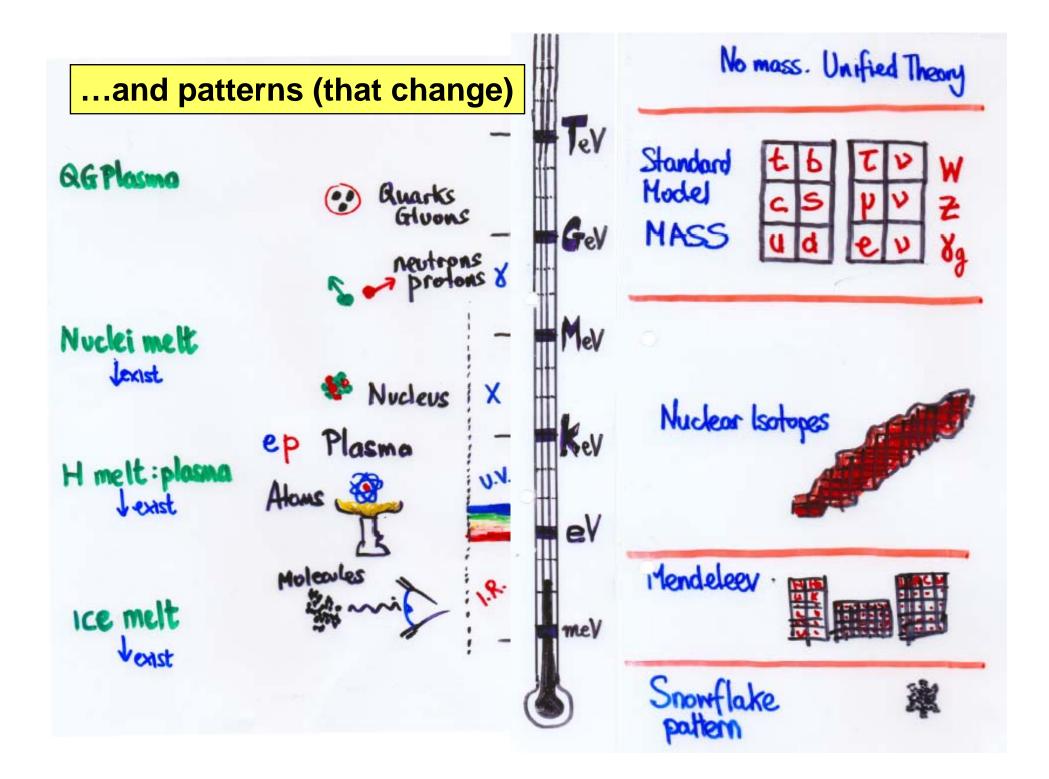
## **The Universe**

in

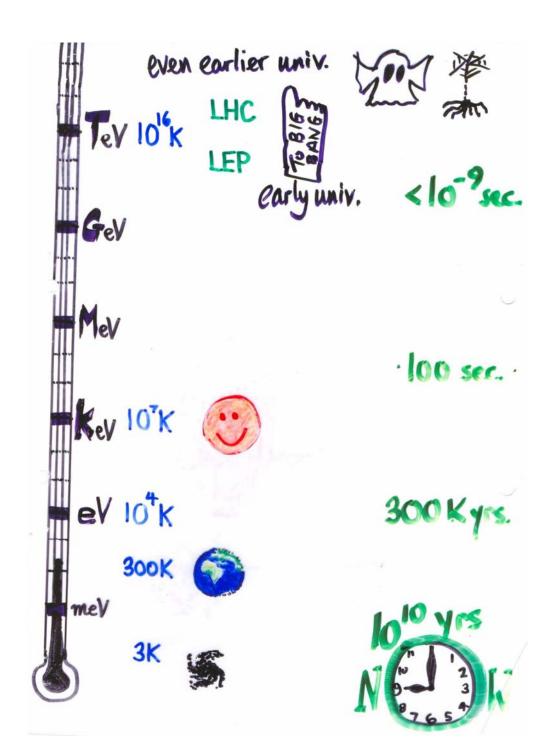
Temperature Energy and Time

even earlier univ. LHC ev 10"K LEP Carly univ. GeV 100 sec. Key 10 K eV 10tK 300K mey 3K



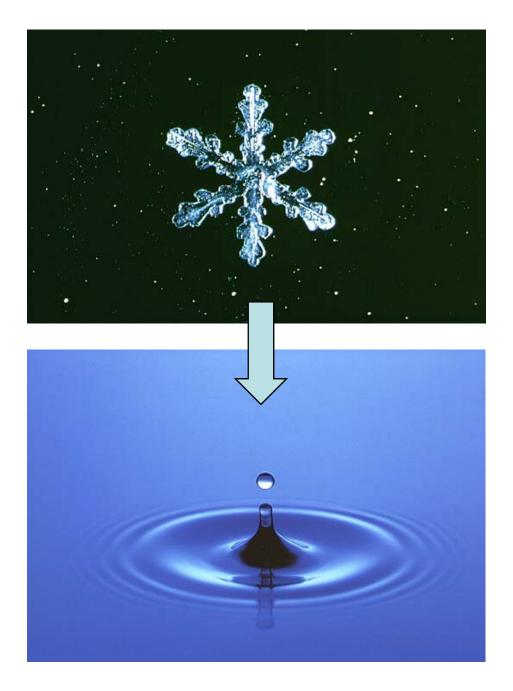


Temperature and symmetry in the universe



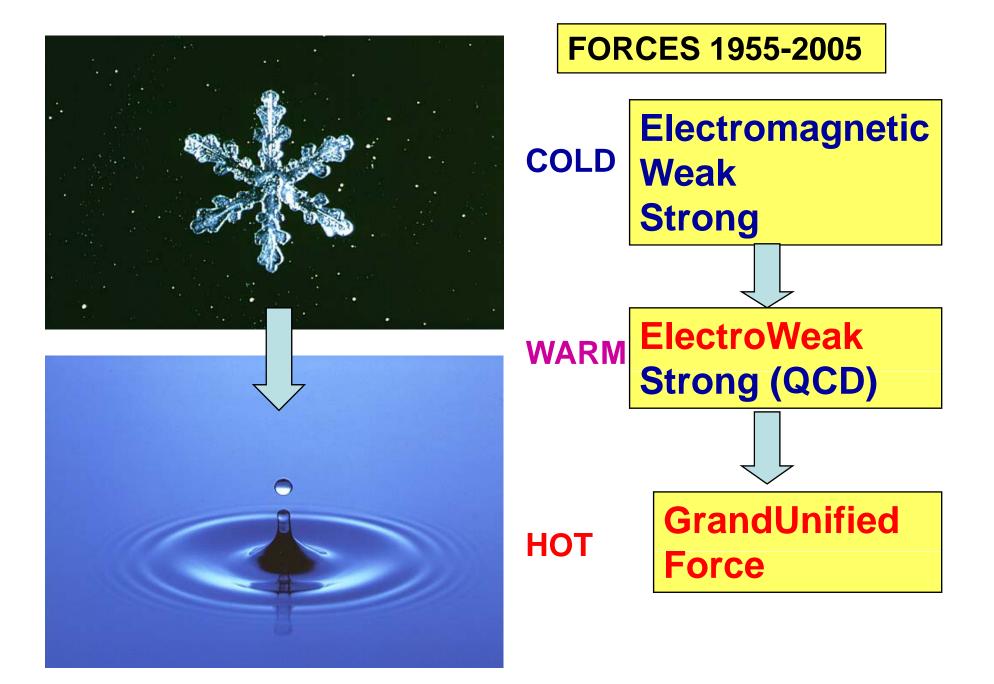


(I will tell you when to be cautious about inhaling)



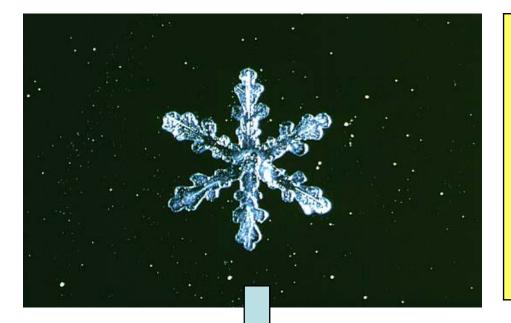
patterns and structures when cold (low energy)

Symmetry when warm (high energy)



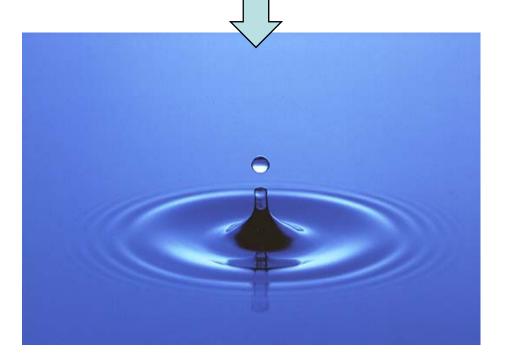


Standard Model of Quarks Leptons and forces = pattern based on Mass "cold" ="low" energy = below 1 TeV



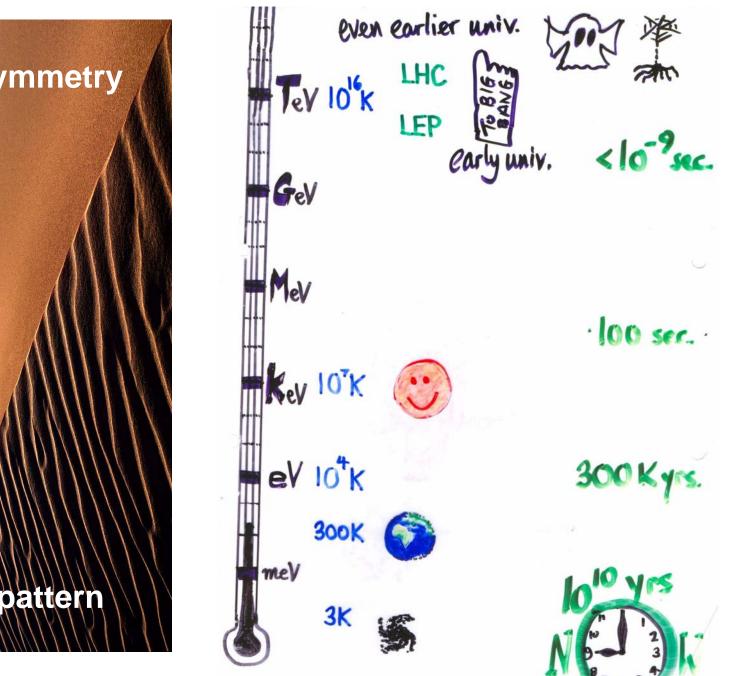
Standard Model of Quarks Leptons and forces

= pattern based on Mass
"cold" ="low" energy
= below 1 TeV



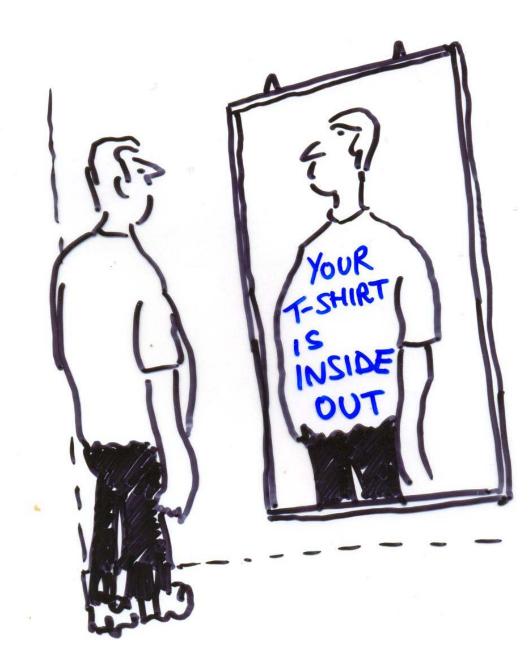
superSymmetry
when "warm"
(= high energy > 1TeV)

Higgs Boson Supersymmetry Nature of Reality

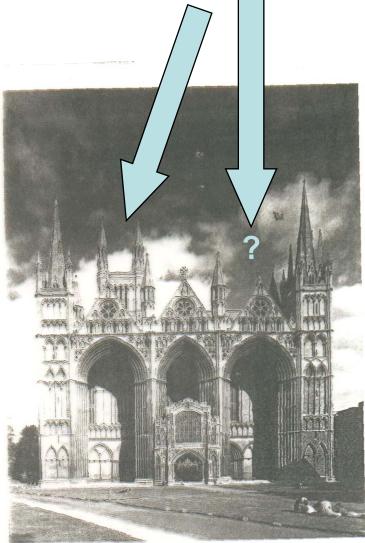


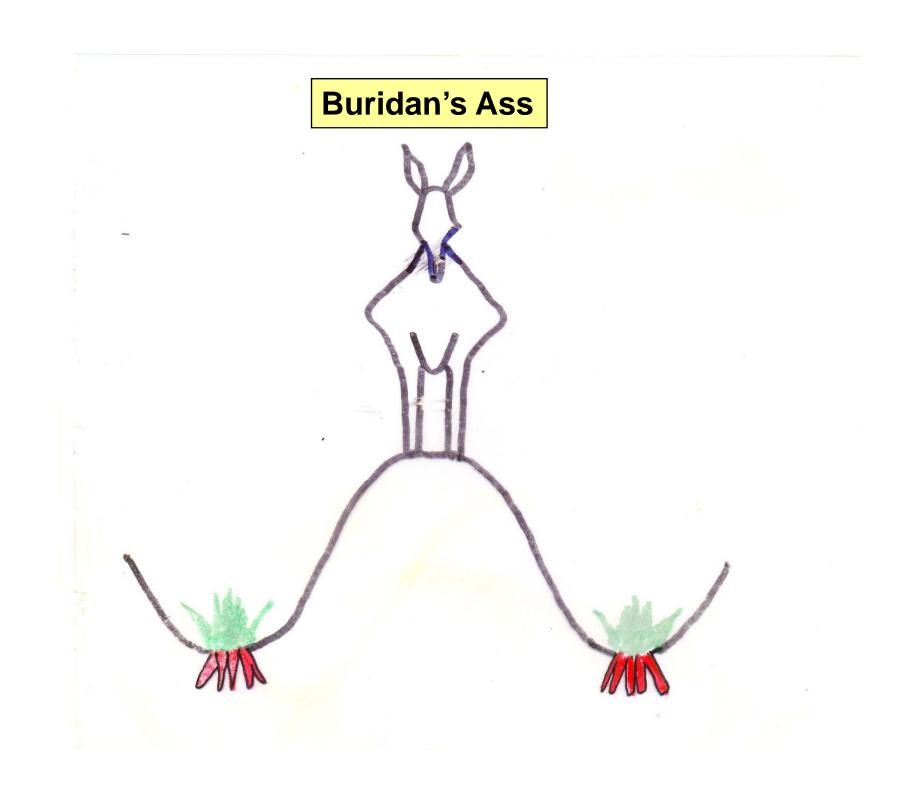


5. symmetries can disappear or change

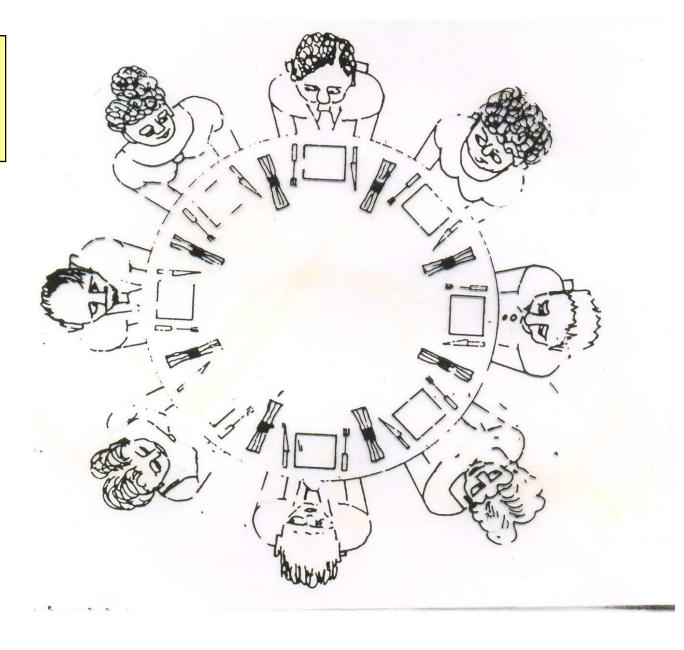


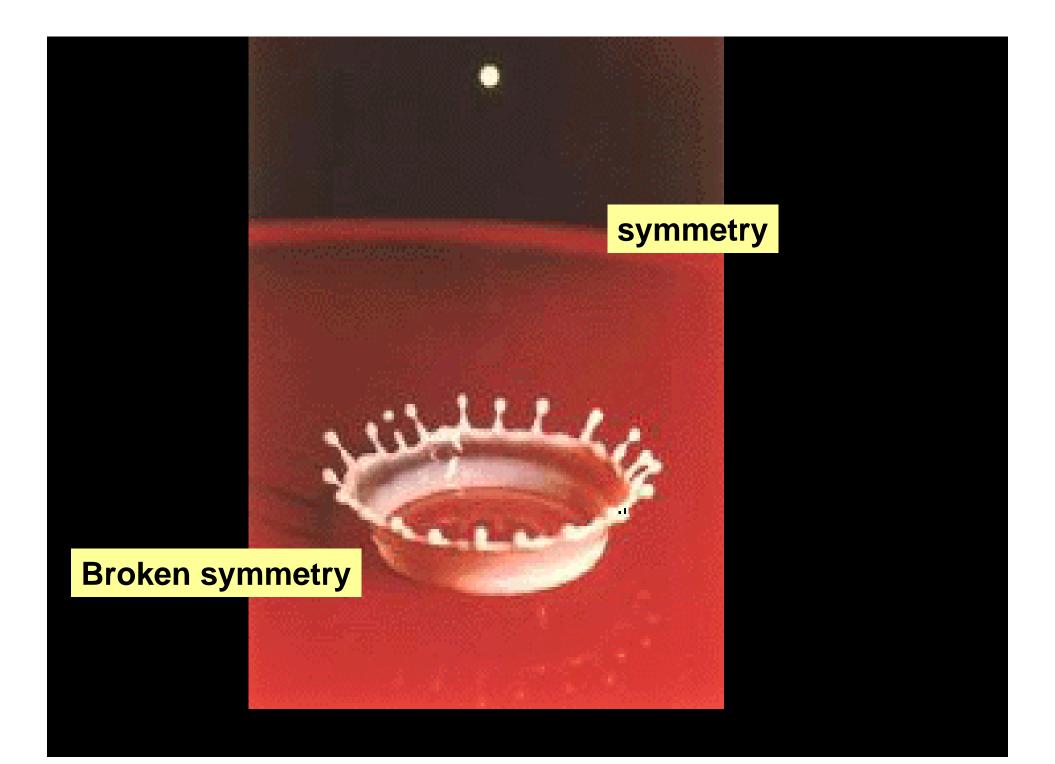
## We like symmetry and when its absent we want to know why





## The problem of the symmetric dinner party

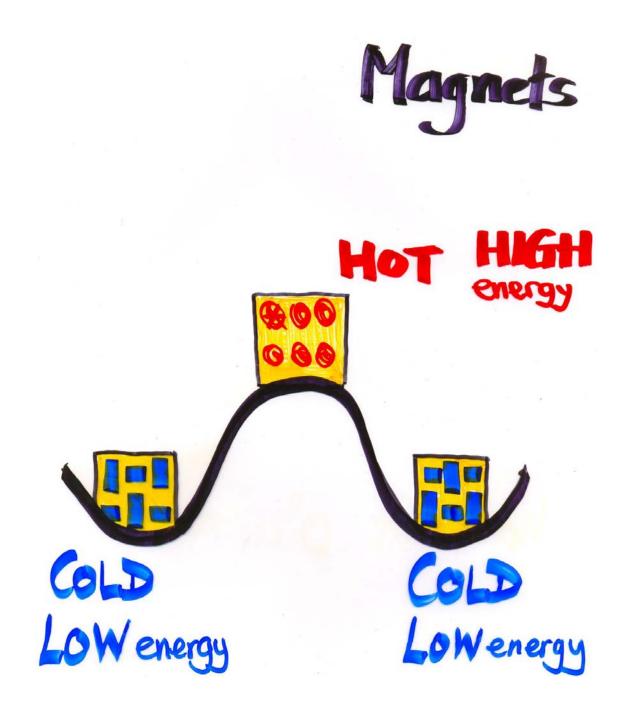


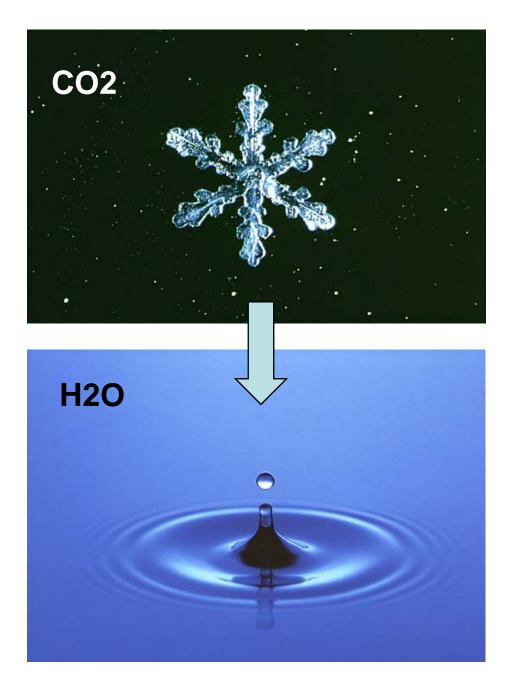


#### Why is this a peak and not a trough?

Answer: random chance But given it's a peak here it dictates where the other peaks are







patterns and structures when cold (low energy)

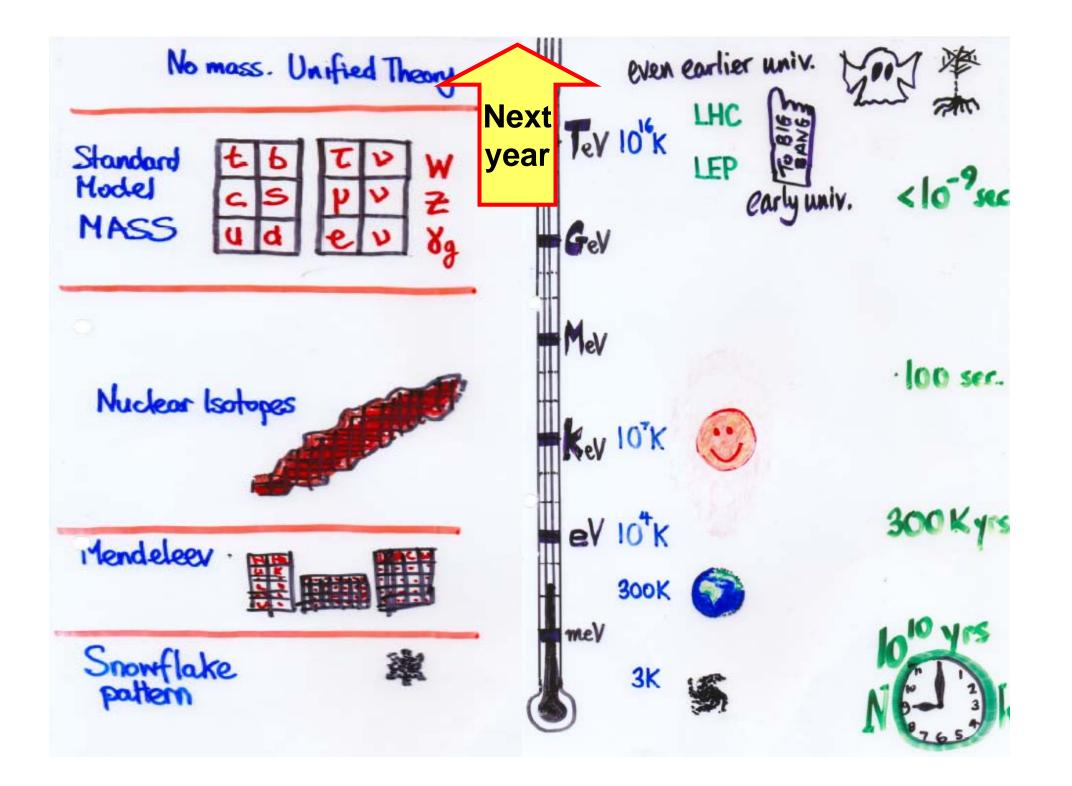
Symmetry when warm (high energy) As the universe cooled after the hot big bang..... We think that an elegant symmetry.....

... "froze" into structures .... And patterns

Such as Atoms ...... Mendeleev's periodic table, And particles .... Quarks, forces and the Standard Model

... which is a pattern based on MASS

2008: heat up to energies above 1000 GeV = "1 TeV" and discover the origin of MASS (= Higgs?)

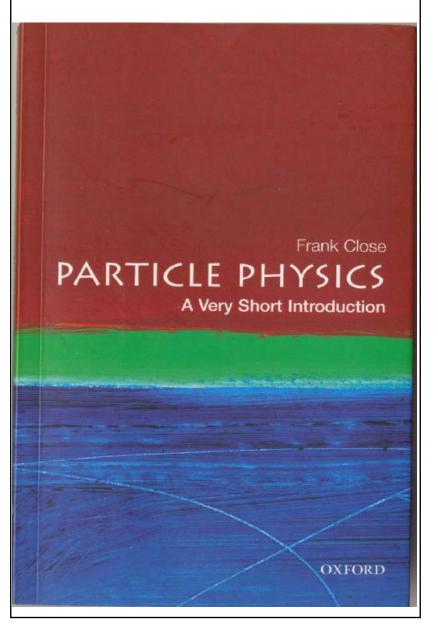


## The Particle Odyssey

## A Journey to the Heart of matter

Frank Close, Michael Marten, Christine Sutton

## **A Very Short Introduction**



#### **Coming out in December**

