

# Doping liquid xenon with light noble gases

Hugh Lippincott, Simone Rizzardini, Fermilab  
Tom Alexander, Andrew Hime, PNNL

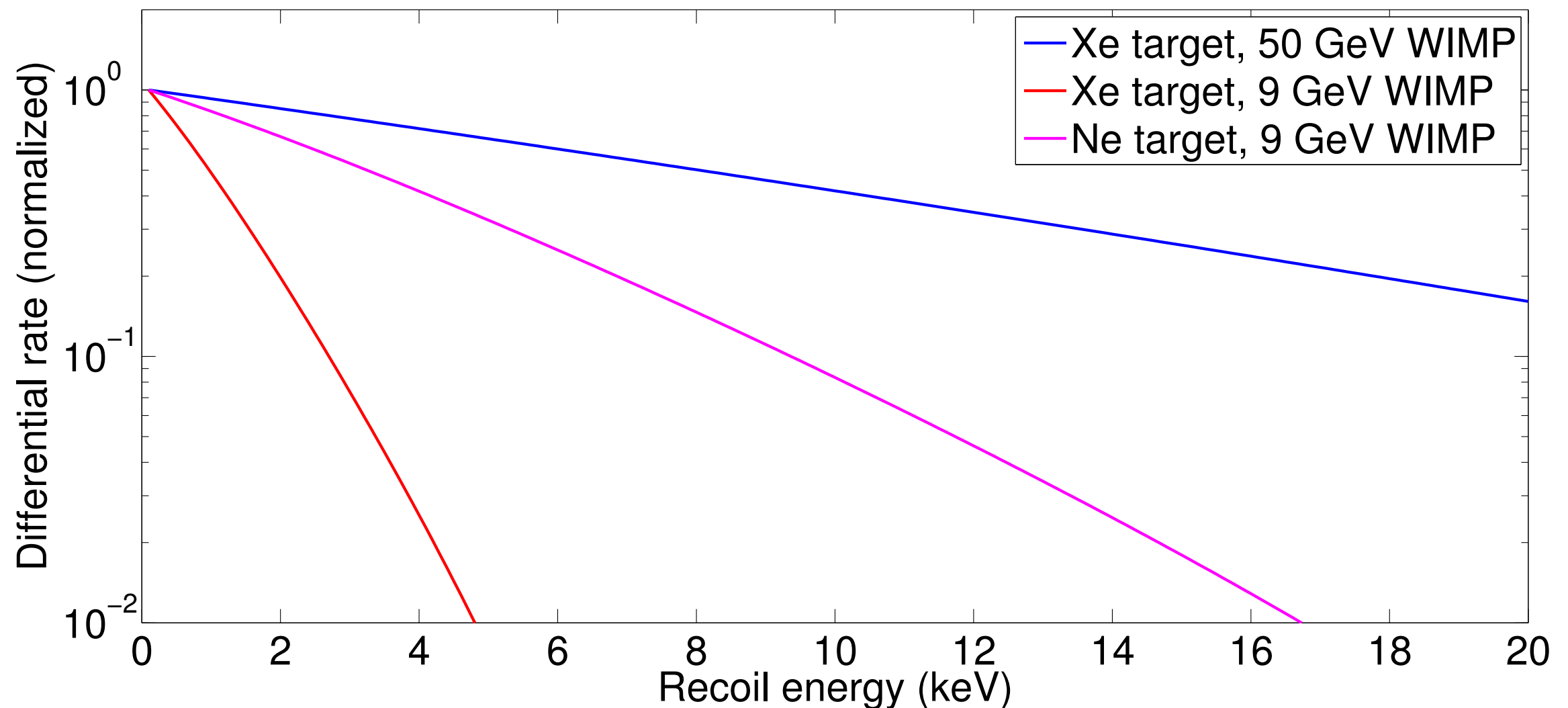
ICHEP 2016

# Low Mass Dark Matter ( $<10$ GeV)

- An original DM candidate (Lee and Weinberg 1977)
- “Ruled out” (or at least not thought about much) for many years
- DAMA excess brings them back - turns out they weren’t ruled out
  - In last decade, several anomalies add excitement
    - e.g. Pamela/FERMI CoGeNT/CDMS-Si/CRESST
- Many anomalies are now resolved, but excitement remains
  - “I think light WIMPs are *more* theoretically motivated than 10 years ago” - Neal Weiner, CIPANP 2015
- Supersymmetry, asymmetric dark matter, minimalist, dark sector, etc.
  - Many existing candidates that evade all constraints, including collider constraints

# What do you need for low mass?

$$\frac{dR}{dQ} = \frac{\rho_0}{m_\chi} \times \frac{\sigma_0 A^2}{2m_p^2} \times F^2(Q) \times \int_{v_m}^{v_{esc}} \frac{f(v)}{v} dv$$



# What do you need for low mass?

$$\frac{dR}{dQ} = \frac{\rho_0}{m_\chi} \times \frac{\sigma_0 A^2}{2m_p^2} \times F^2(Q) \times \int_{v_m}^{v_{esc}} \frac{f(v)}{v} dv$$

$$v_m = \sqrt{Qm_N/2m_r^2}$$

$$v_{esc} = 544 \text{ km/s (current value)}$$

$m_N$  is mass of nucleus

$$m_r = \frac{m_N m_\chi}{m_N + m_\chi}$$

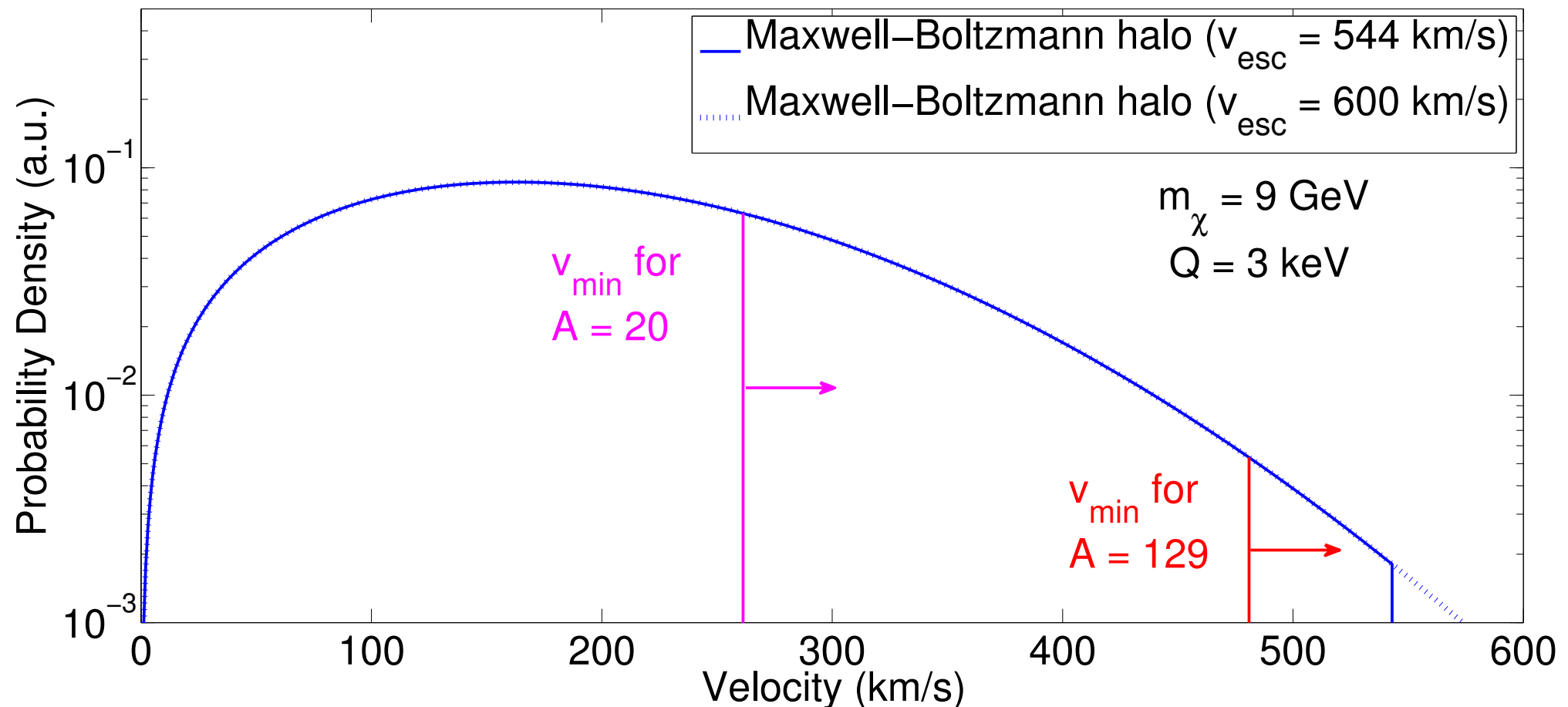
- Low threshold
- Low mass target (for better kinematic match to the dark matter mass)
- For given  $Q$ ,  $v_m$  is minimized when  $m_n = m_\chi$

# What do you need for low mass?

$$\frac{dR}{dQ} = \frac{\rho_0}{m_\chi} \times \frac{\sigma_0 A^2}{2m_p^2} \times F^2(Q) \times \int_{v_m}^{v_{esc}} \frac{f(v)}{v} dv$$

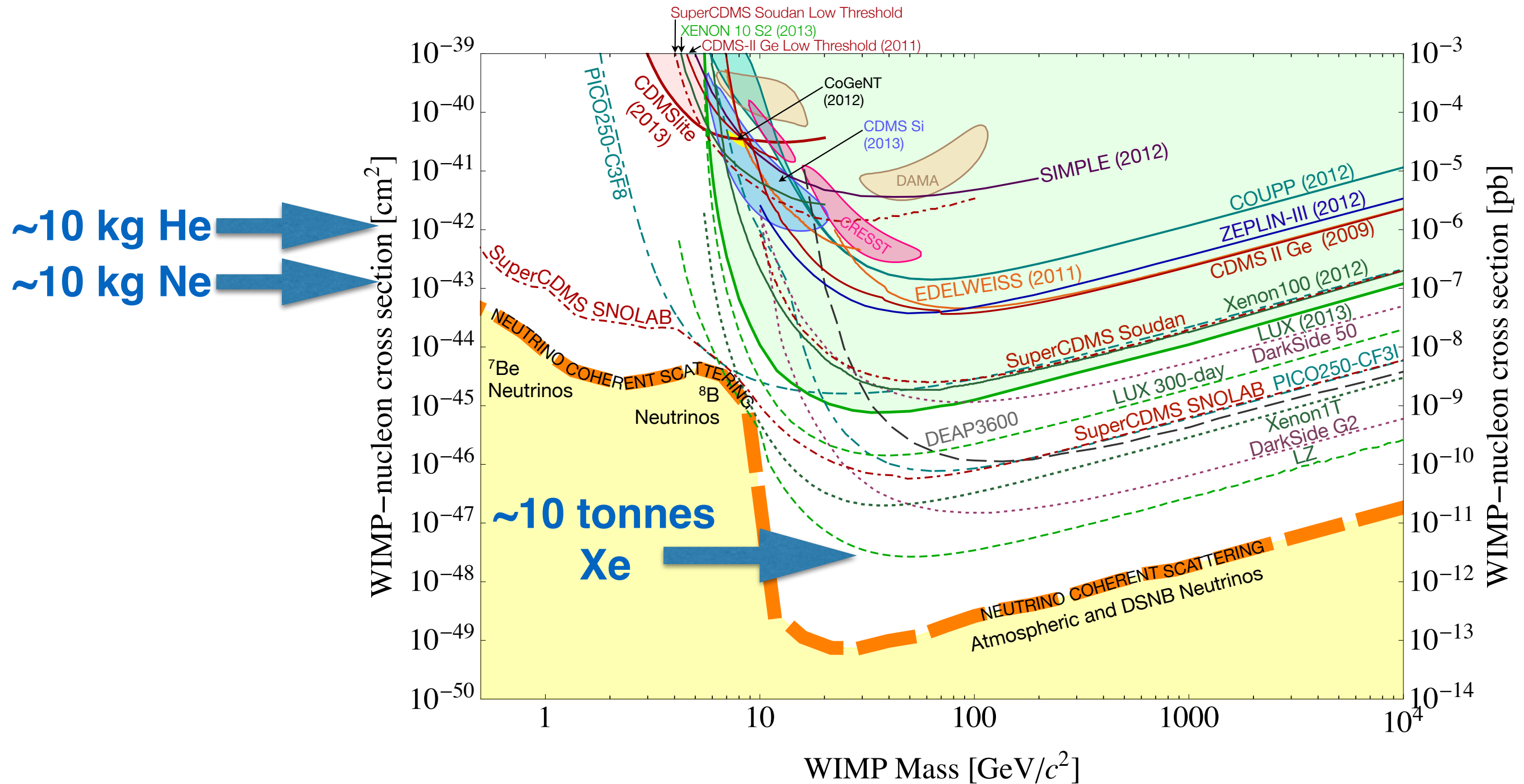
$$v_m = \sqrt{Qm_N/2m_r^2}$$

$$v_{esc} = 544 \text{ km/s (current value)}$$



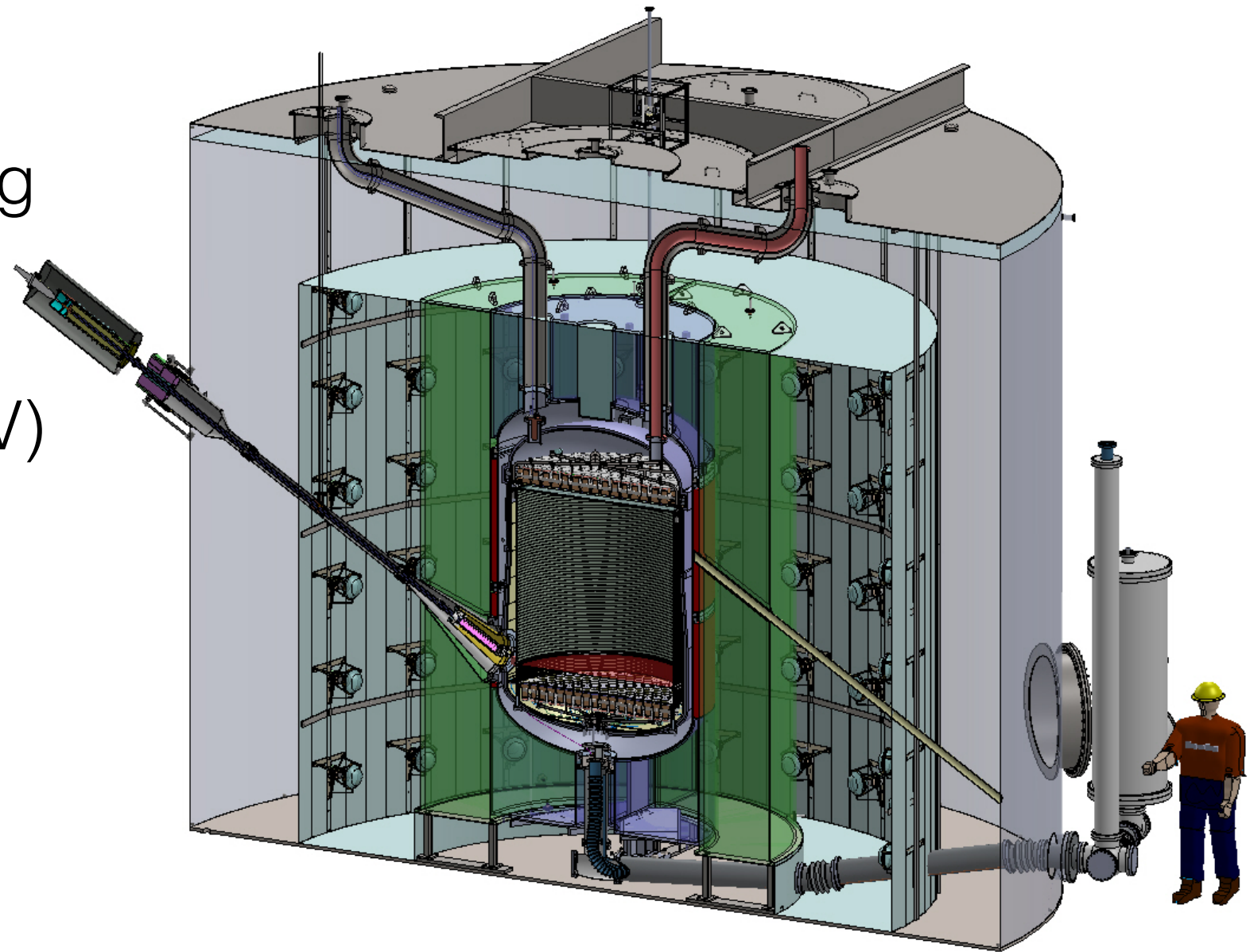
# What don't you need for low mass?

- A lot of mass



# LUX-Zeplin (LZ)

- 7 tonne active LXe TPC
  - Heavy target
  - Excellent self shielding
  - Good discrimination
  - Low threshold ( $<3$  keV)
- 31 institutions, ~200 people
- To be located at Sanford Lab in SD

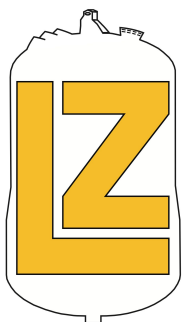
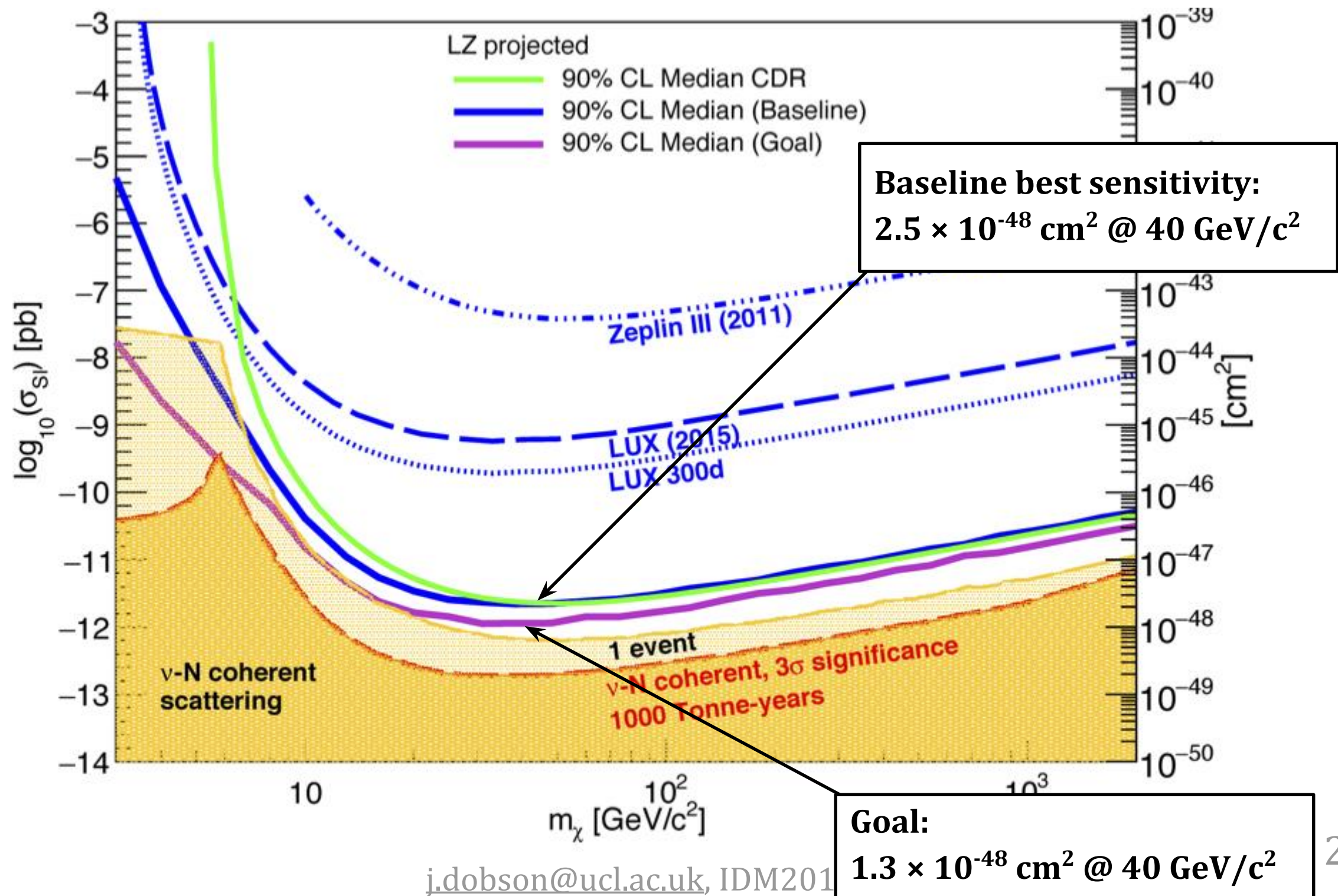






# LUX-Zeplin (LZ)

5600 kg fiducial, 1000 live-day exposure



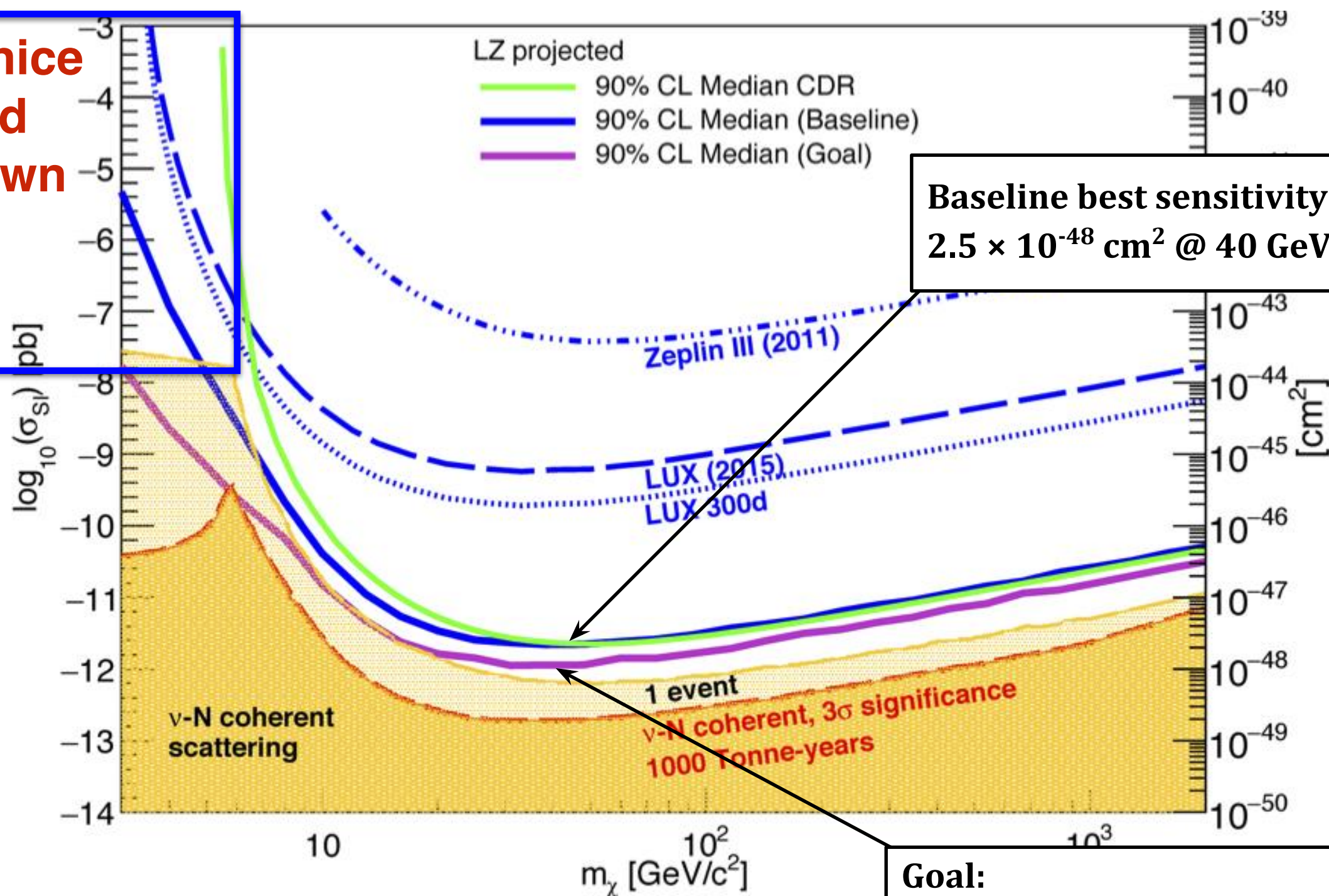




# LUX-Zeplin (LZ)

5600 kg fiducial, 1000 live-day exposure

Would be nice  
to extend  
further down  
here!



Baseline best sensitivity:  
 $2.5 \times 10^{-48} \text{ cm}^2 @ 40 \text{ GeV}/c^2$

Goal:  
 $1.3 \times 10^{-48} \text{ cm}^2 @ 40 \text{ GeV}/c^2$

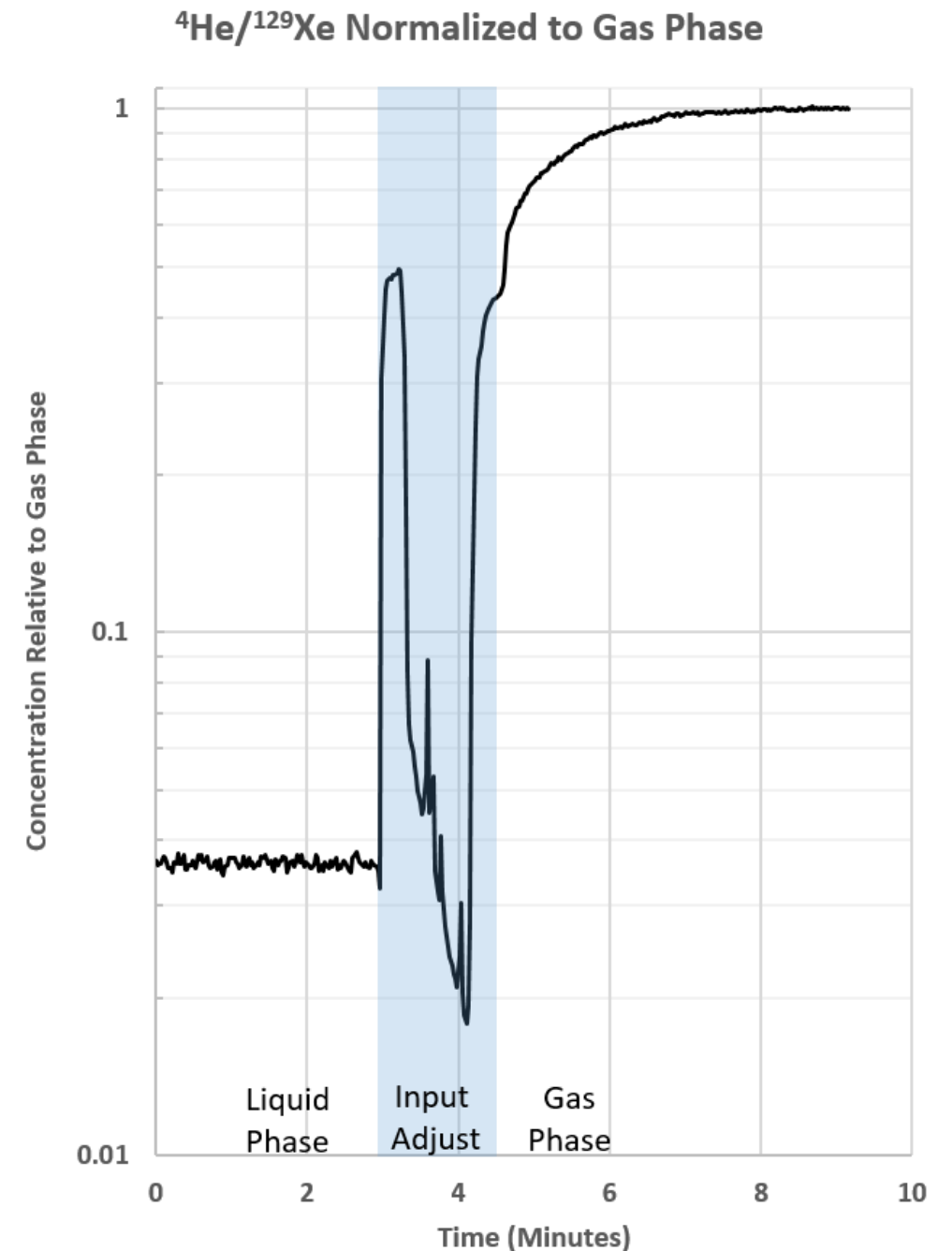


# Can we add He/Ne to LXe?

- Dissolve small quantities of He or Ne in liquid xenon
- Extend the reach of a detector like LZ (or Xenon1T or PandaX, etc)
- Add new targets to field of direct detection
  - No current experiments using helium or neon
- Capitalize on investment in large detectors by adding flexibility

# How much could we get in?

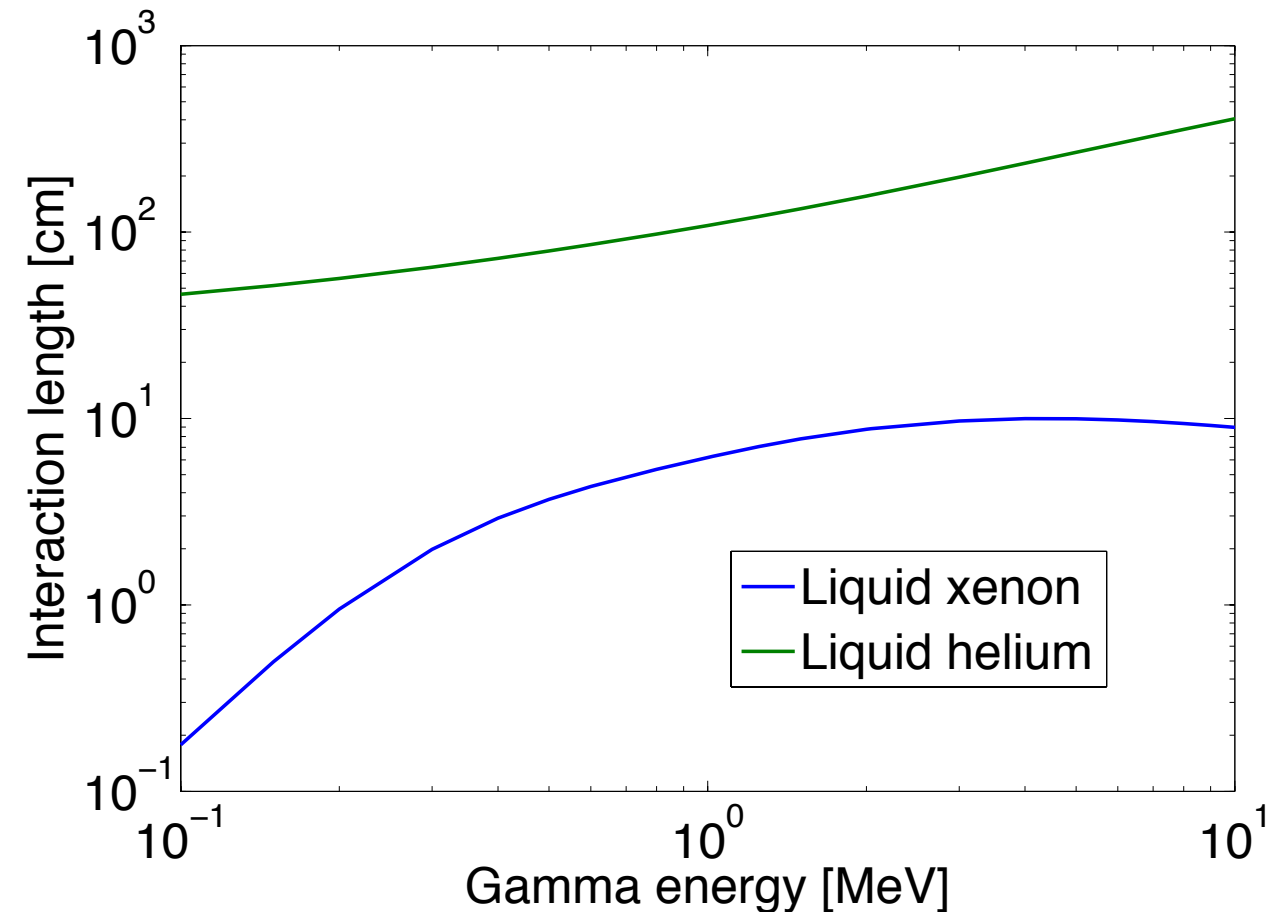
- No published measurements
- Preliminary test at Fermilab shows 0.1% He in LXe by mass is easily achievable
  - 1 bar of partial pressure
  - Consistent with measurements from LUX
- Expected to scale with mass ratio (e.g. 0.5% Ne)
- Can we get more in?
  - Temperature dependence?



$$0.037 \text{ mol He/mol Xe} \times \frac{M_{\text{He}}}{M_{\text{Xe}}} \sim 0.1\%$$

# Backgrounds

- Helium and neon have no long lived isotopes
  - No new backgrounds introduced
- Detector is already built of low background materials
- Keep excellent self shielding of LXe (not possible with LNe or LHe-only detector)



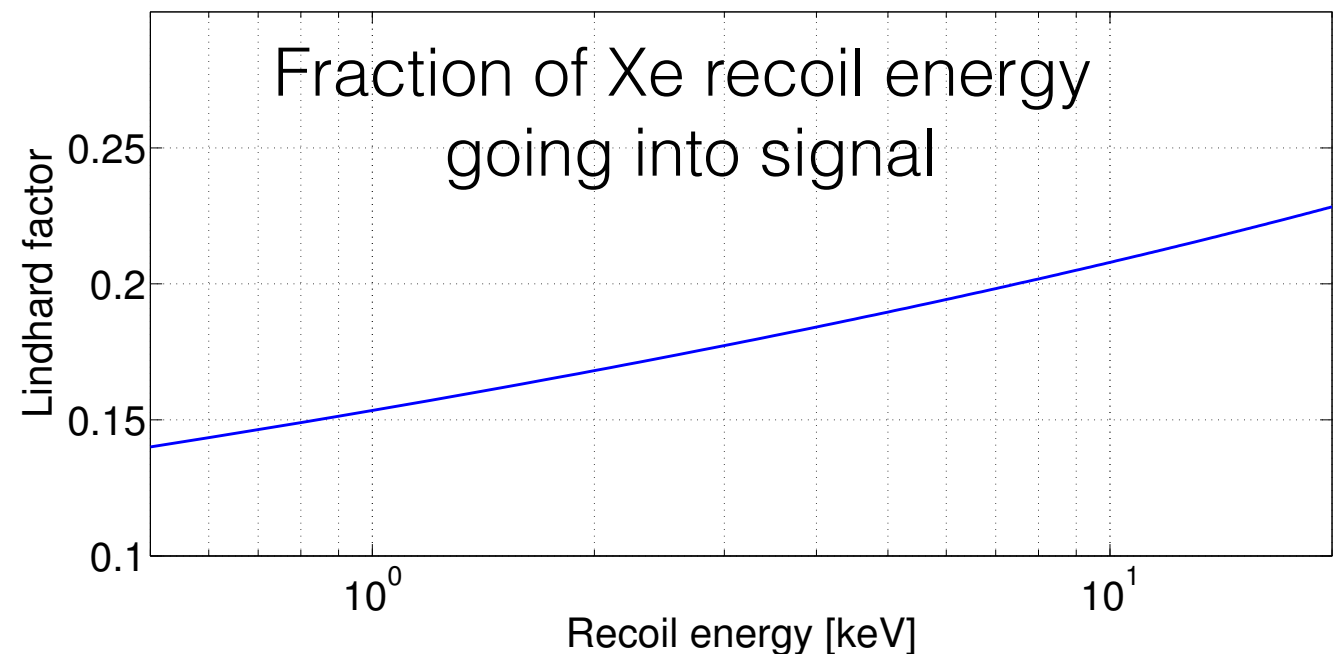
# Signal detection

- Helium and neon scintillate in harder UV
  - 80 nm vs 175 nm in LXe
- Those photons will wavelength shift in the xenon to 175 nm
  - See, for example, xenon doped in argon (JINST 9, P06013, among others)
- Keep same photon detection scheme!



# Signal yield

- Strong quenching factor for nuclear recoils in liquid xenon (Lindhard factor)
  - Less than 20% of a 7 keV recoil event goes into detectable signal
  - The rest goes into nuclear collisions that lead to heat
- Helium/Neon are light nuclei, meaning more energy goes into electronic channels -> more signal





# Signal yield

Recoil	Lindhard	SRIM
Xenon	0.02	0.02
Neon	0.20	0.09
Helium	0.68	0.69

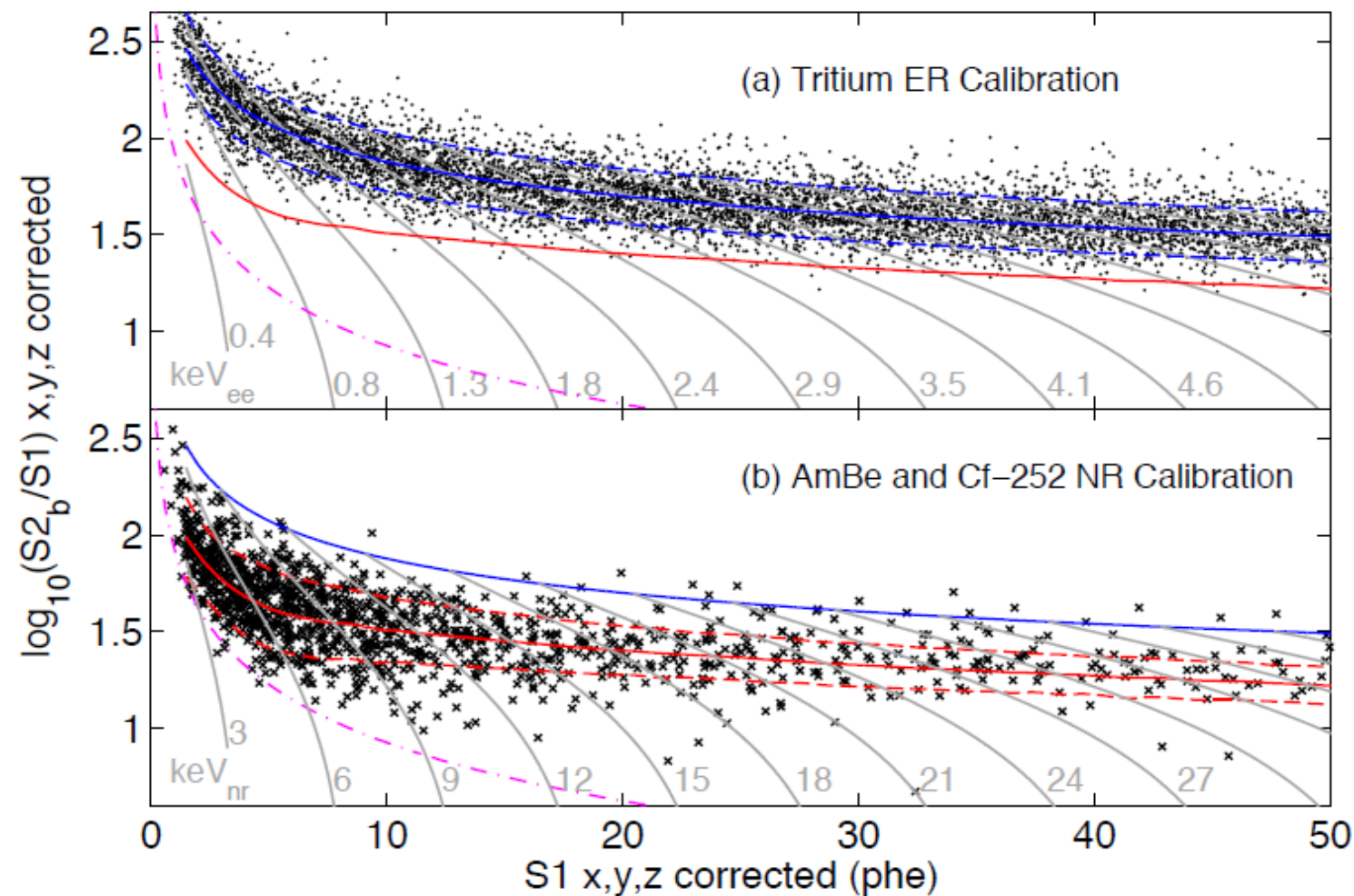
Table 1: Estimated fraction of energy given to electronic stopping for nuclear recoils (not accounting for secondary cascades) from Xe, He, and Ne recoils in LXe, calculated using Lindhard theory [41] or the SRIM simulation package [42].

- At worst, we can expect a factor of 3.5 more signal for helium recoils in LXe

**Even lower thresholds with the light target!**

# Key questions

- What is the true signal yield?
- What happens to discrimination?



LUX data  
PRL 112, 091303

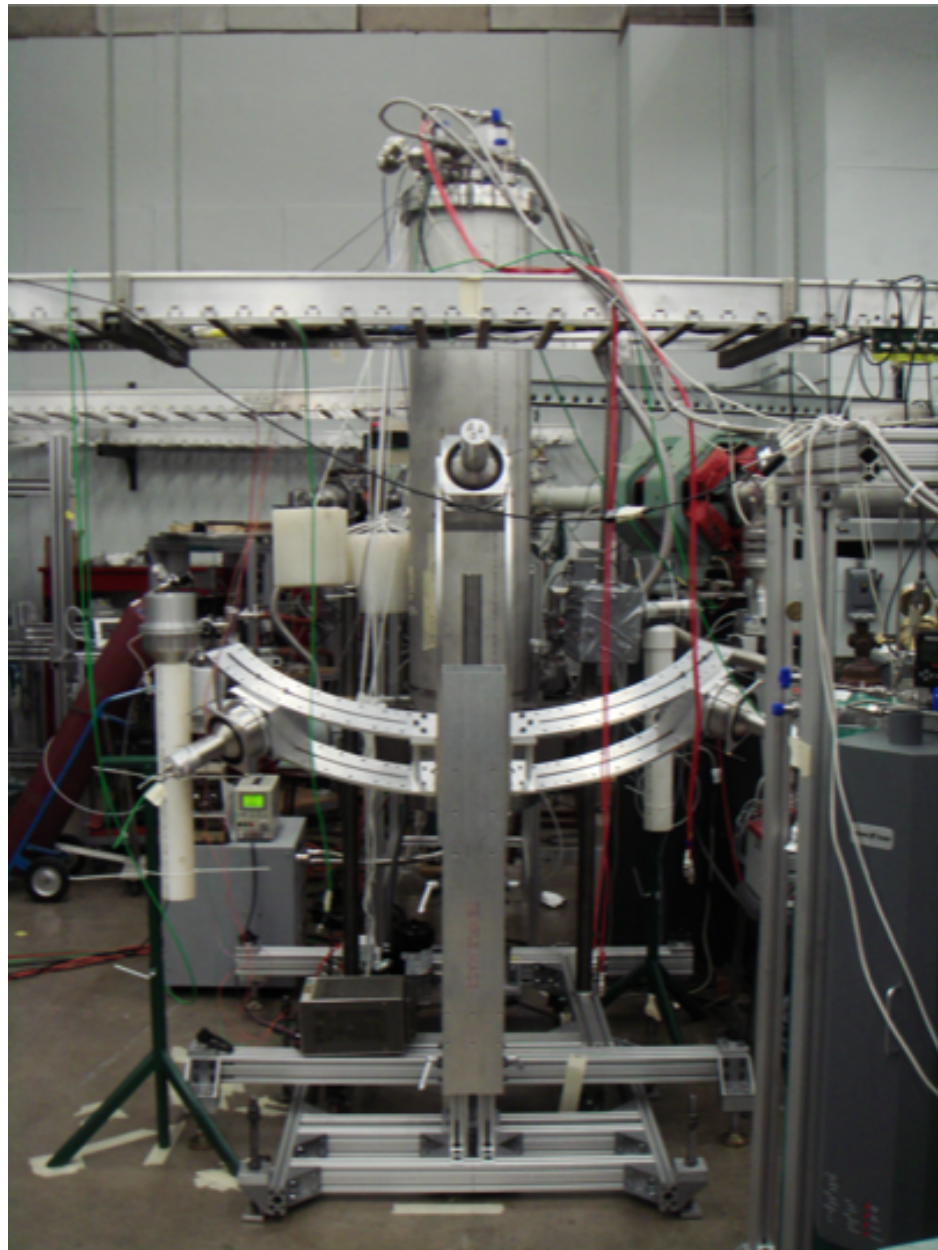
# SCENE

Pulsed, mono-energetic neutrons

LAr TPC

Scattering angle,  $\theta$

Neutron detector

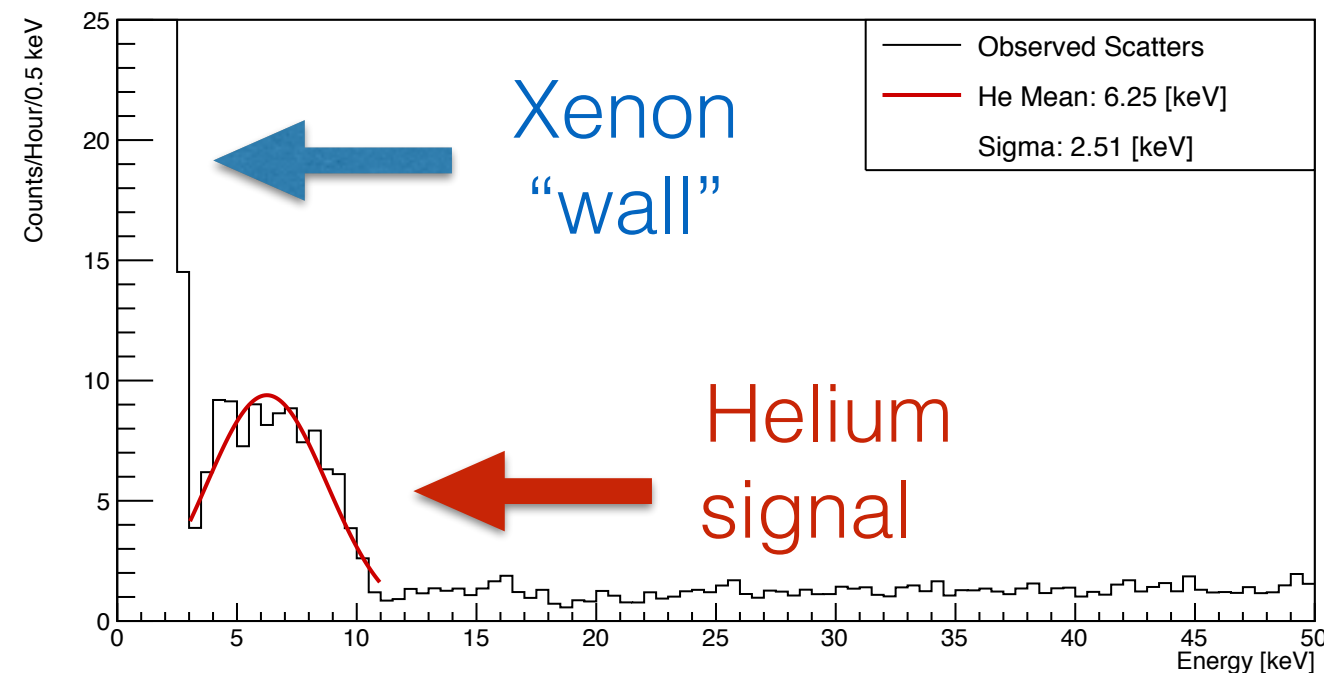


- Pulsed, monoenergetic beam (at Notre Dame) to measure response of to nuclear recoils of known energy
- Tunable nuclear recoil energy by changing the neutron energy and the scattering angle
  - Neutrons of 500 keV - 1.5 MeV
  - Recoils of a few keV up to 50 keV
  - Successful measurements in LAr (1406.4825, 1306.5675)

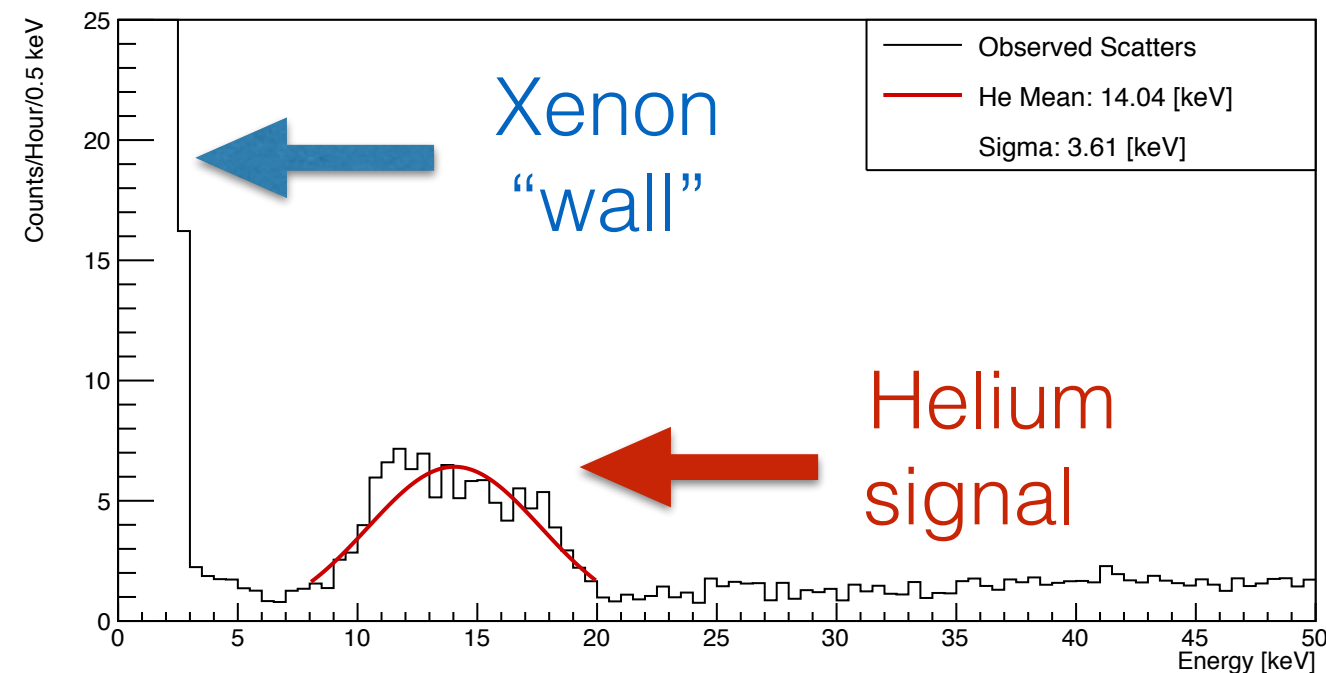
# SCENE

- For doping measurement, for a given scattering angle, He/Ne recoils have more energy
  - Before accounting for increased signal
- Pushes the peak out past the xenon background

Simulated Energy Deposition: Xenon/Helium(0.1%), 90% multiple scatter rejection  
Neutron Beam E=100keV, 22.5 degree scattering angle



Simulated Energy Deposition: Xenon/Helium(0.1%), 90% multiple scatter rejection  
Neutron Beam E=100keV, 45 degree scattering angle



SCENE-like measurement measures yield and S1/S2 response v. energy!

# He/Ne doping in LXe

- Physically possible
- Keep low background level achieved in LXe TPC
- Same signal readout with LXe sensitive light detectors
- Increased signal yield from He/Ne recoils
  - Lower energy thresholds for WIMP-He/Ne scattering
- Properties measurable using existing techniques



# What's the catch?

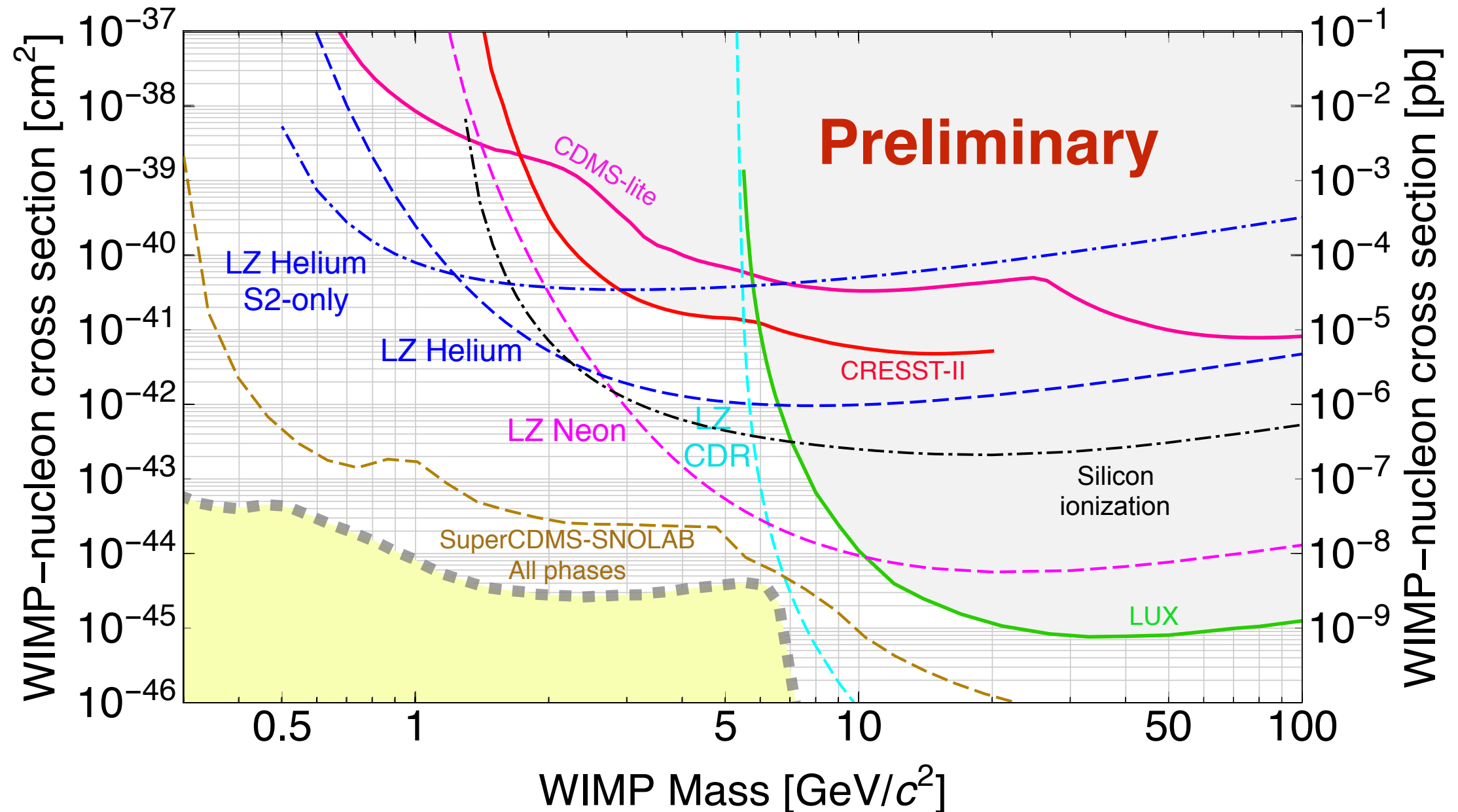
- At very low thresholds (where we want to go), we hit coherent scattering of neutrinos

$$\frac{R_{\nu,\text{coh}}}{R_{\chi}} \sim N^2 / A^2$$

- In doped LXe, N is still  $\sim 70$ , but A is now 4 or 20, instead of  $\sim 130$
- Hit the neutrino background at x1000 higher WIMP cross section for helium



# Preliminary projection



- Location of LZ Helium lines depends critically on assumed signal yield (conservative assumption of factor of 3 shown here)
  - Can get around neutrino backgrounds with more He signal
- Currently not using any spectral information (cut and count is not ideal)

# He/Ne doping in LXe

- Physically possible
- Keep low background level achieved in LXe TPC
- Same signal readout with LXe sensitive light detectors
- Increased signal yield from He/Ne recoils
  - Lower energy thresholds for WIMP-He/Ne scattering
- Properties measurable using existing techniques
- Coherent neutrino scattering background
- Possible mitigation with higher signal or spectral analysis



# He/Ne doping in LXe

- Physically possible
- Keep low background level achieved in LXe
- Same signal readout with different light detectors
- Increased sensitivity from He/Ne recoils
  - Lower energy thresholds for WIMP-He/Ne scattering
- Properties measurable using existing techniques
- Coherent neutrino scattering background
- Possible mitigation with higher signal or spectral analysis

**Worth further exploration!**

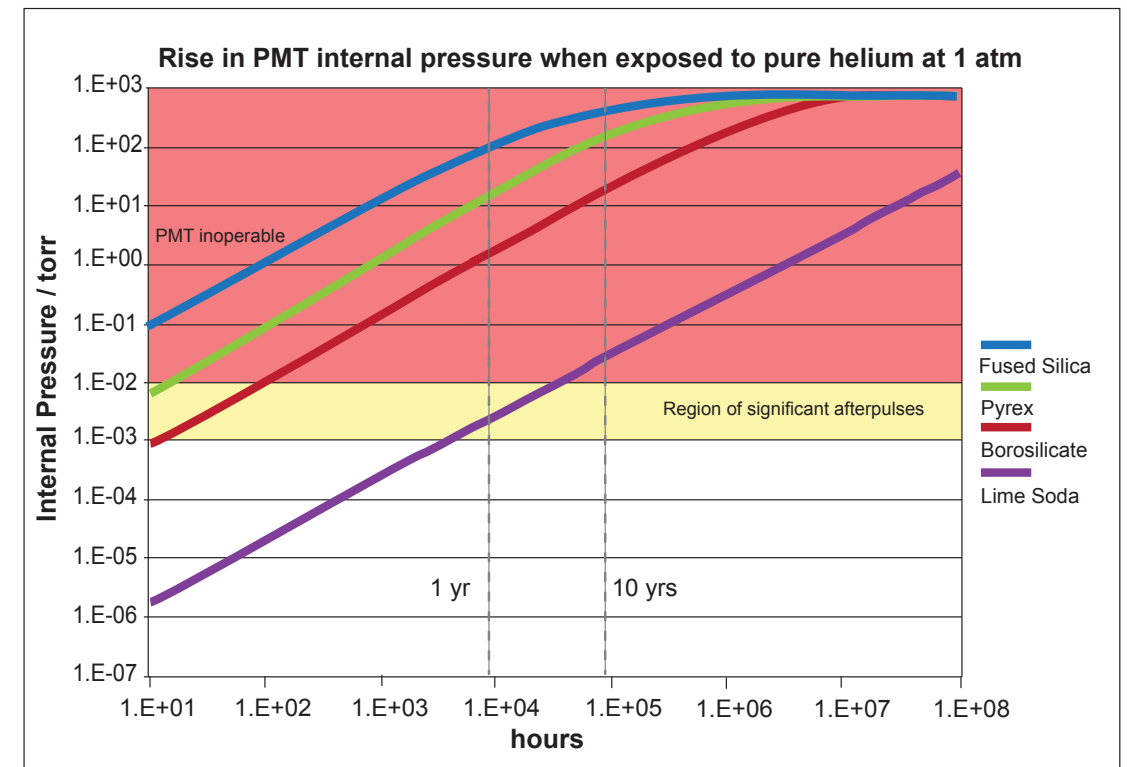


Backup

# He diffusion through PMT

Example for ET9226 PMT

- After pulsing from helium diffusion in tubes is a well known problem
- Diffusion exponentially suppressed by temperature (Arrhenius relationship)
- R11410 has a surprisingly thick window (3 mm)
- Calculation suggests 10 years at 1 bar/165 K before reaching significant afterpulses
- Needs confirmation...



**R11410**



**R8520**