

Lattice Gauge Theory bounds on Composite Dark Matter

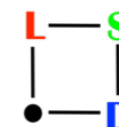
Enrico Rinaldi



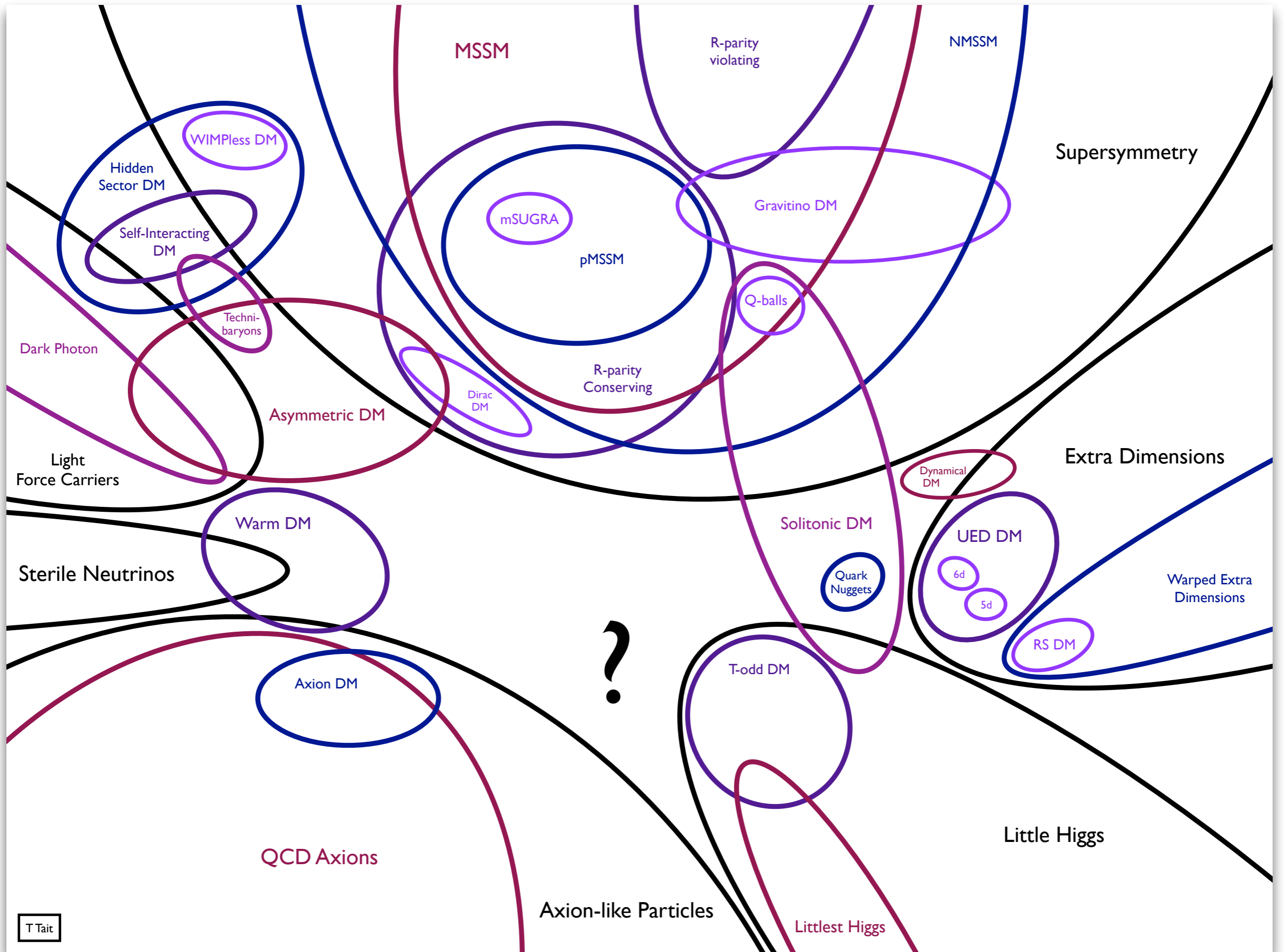
This research was performed under the auspices of the U.S. Department of Energy by Lawrence Livermore National Laboratory under Contract DE-AC52-07NA27344 and supported by the LLNL LDRD "Illuminating the Dark Universe with PetaFlops Supercomputing" 13-ERD-023.

Computing support comes from the LLNL Institutional Computing Grand Challenge program.

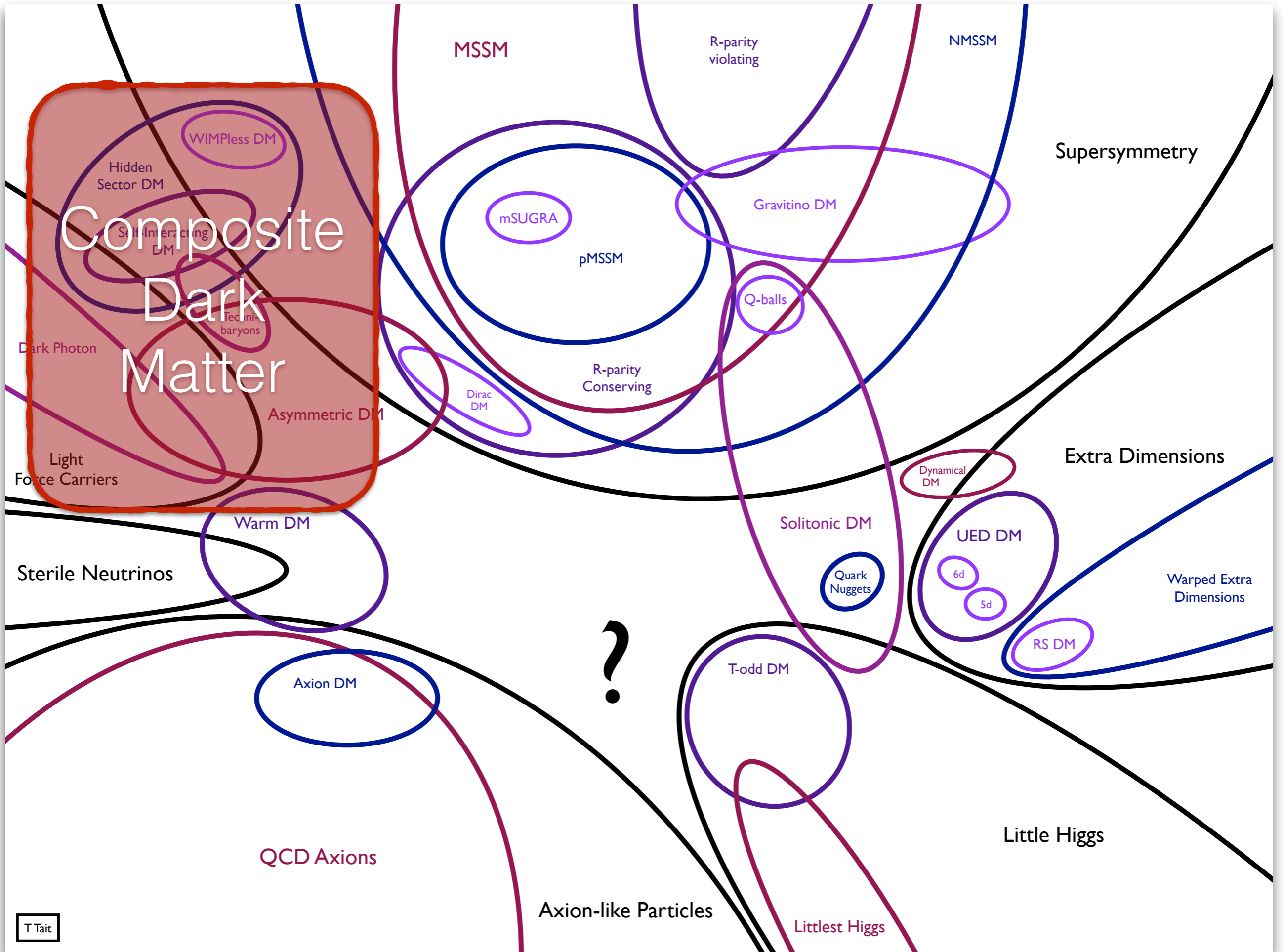
work with the **L**attice **S**trong **D**ynamics Collaboration



LLNL-PRES-669543



T Tait



Composite Dark Matter

Sterile Neutrinos

Dark Photon

Light Force Carriers

QCD Axions

T Tait

MSSM

R-parity violating

NMSSM

Supersymmetry

mSUGRA

pMSSM

Gravitino DM

Q-balls

R-parity Conserving

Dirac DM

Extra Dimensions

Dynamical DM

Warm DM

Solitonic DM

UED DM

6d

5d

Warped Extra Dimensions

Quark Nuggets

RS DM

?

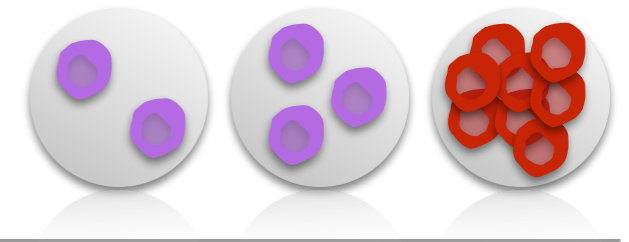
T-odd DM

Little Higgs

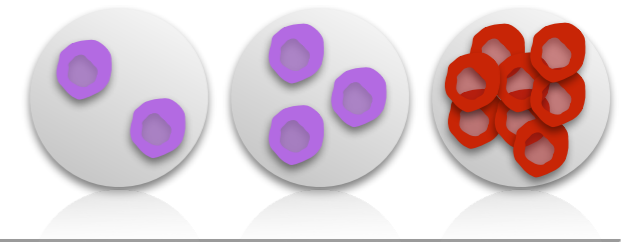
Axion-like Particles

Littlest Higgs

Composite Dark Matter

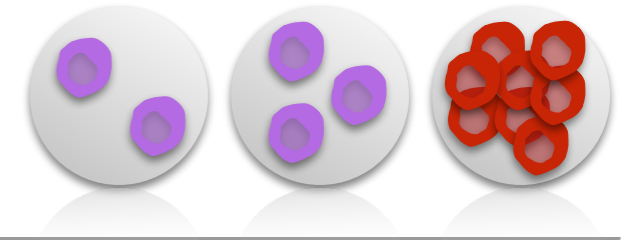


Composite Dark Matter



- ◆ Dark Matter is a **composite** object

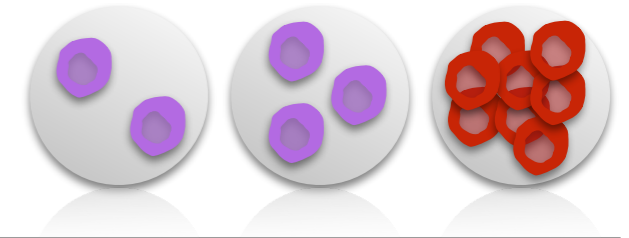
Composite Dark Matter



- ◆ Dark Matter is a **composite** object

e.g. **technibaryon** or
hidden **glueball**

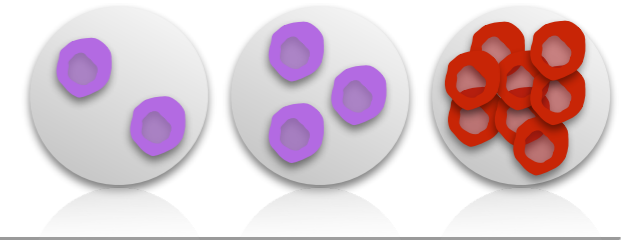
Composite Dark Matter



- ◆ Dark Matter is a **composite** object
- ◆ Interesting and complicated internal **structure**
- ◆ Properties dictated by **strong dynamics**
- ◆ **Self-interactions** are natural

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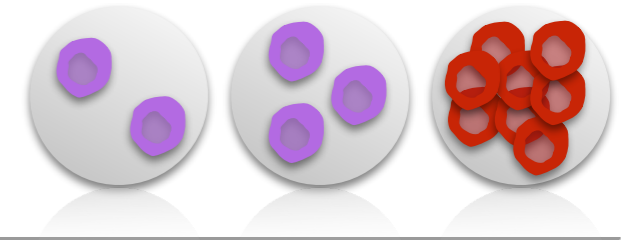


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Similar to **QCD**

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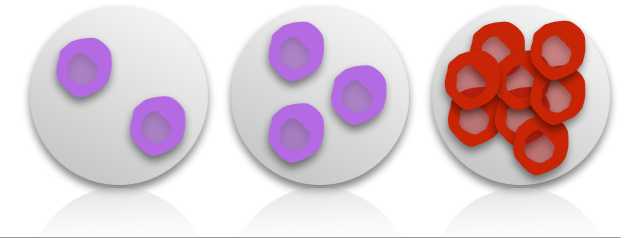


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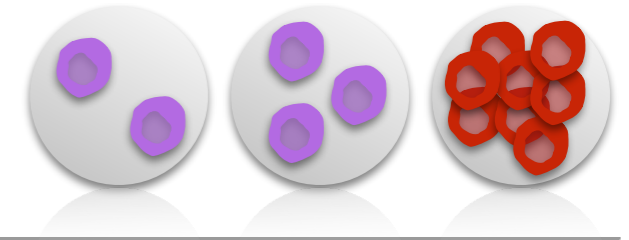
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Chance to **observe them**
in experiments and give the
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Composite Dark Matter



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e.g. **technibaryon** or
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◆ Interesting and complicated internal **structure**

Lattice Field Theory methods

◆ Properties dictated by **strong dynamics**

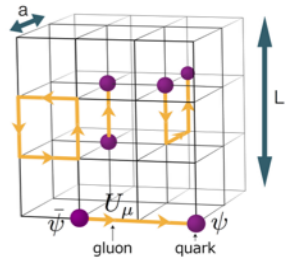
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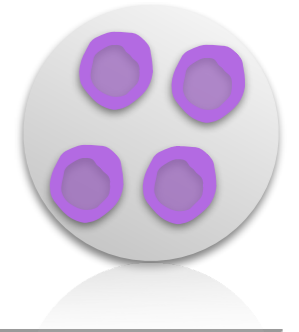
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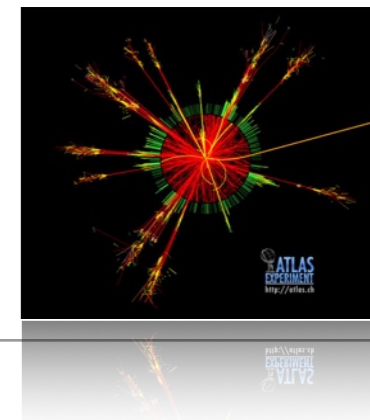
Importance of lattice field theory simulations

- ◆ *lattice simulations are needed to solve the strong dynamics*
- ◆ naturally suited for models where dark fermion masses are comparable to the **confinement scale**
- ◆ **controllable** systematic errors and room for **improvement**
- ◆ Naive dimensional analysis and EFT approaches can miss important **non-perturbative** contributions
- ◆ NDA is **not precise enough** when confronting experimental results and might not work for certain situations: there are uncontrolled theoretical errors



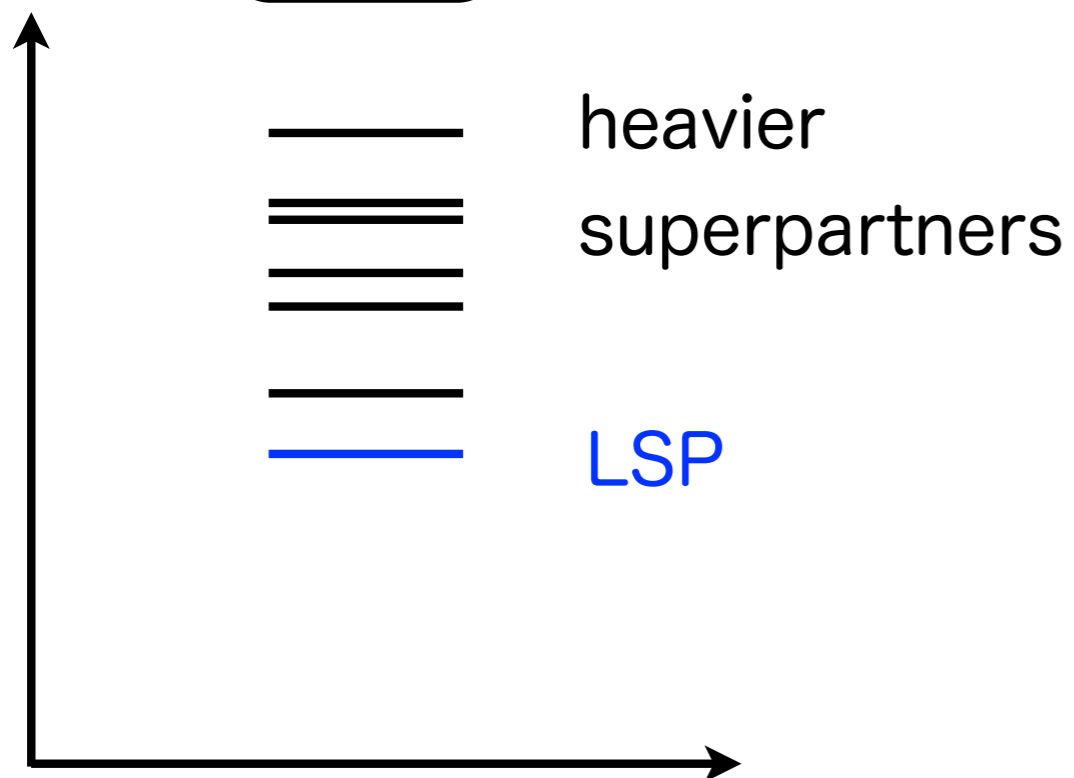
“Stealth Dark Matter” Model

- ◆ **New strongly-coupled SU(4) gauge sector** “like” QCD with a **plethora of composite states** in the spectrum: all mass scales are technically natural for hadrons
- ◆ New **Dark fermions**: have **dark color** and also have **electroweak charges** ($W/Z, \gamma$)
- ◆ Dark fermions have **electroweak breaking masses** (Higgs) and **electroweak preserving masses** (not-Higgs)
- ◆ A global symmetry naturally stabilizes the **dark lightest baryonic** composite states (e.g. dark neutron)



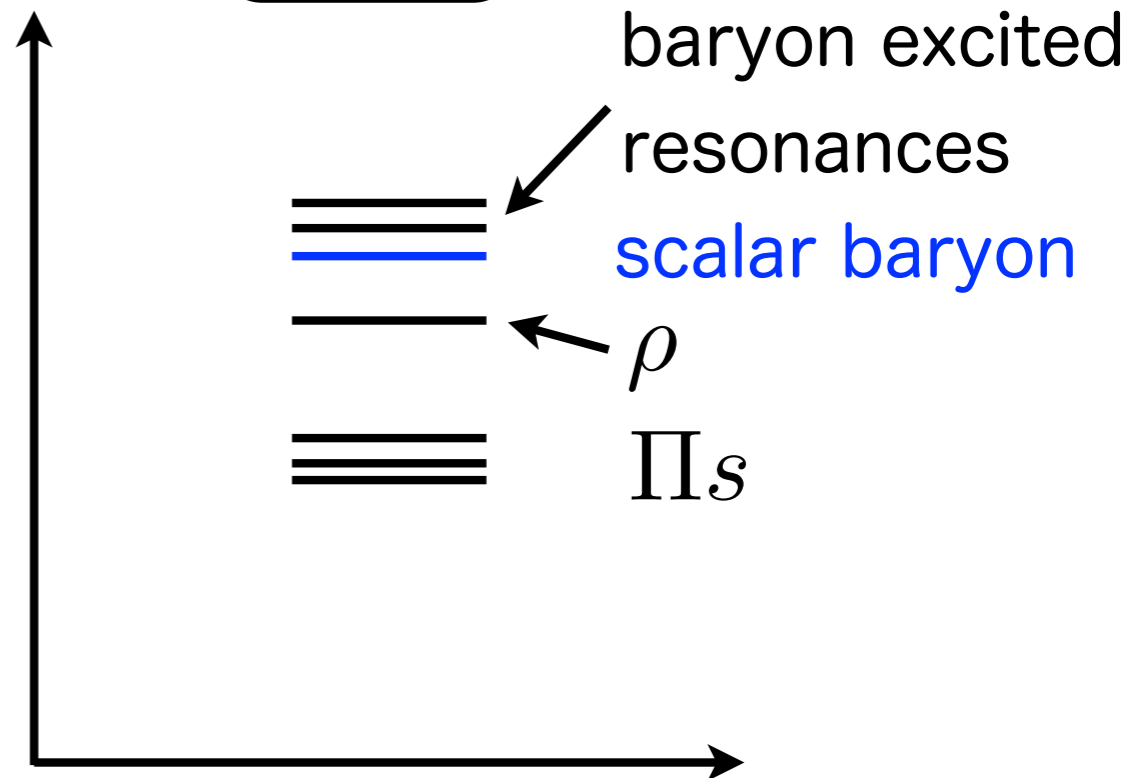
Stealth DM at colliders

SUSY

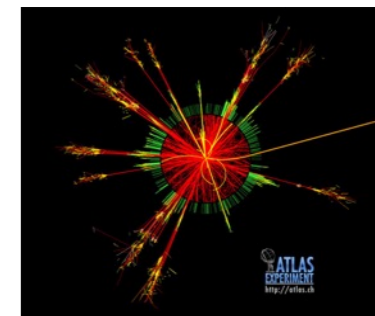


VS.

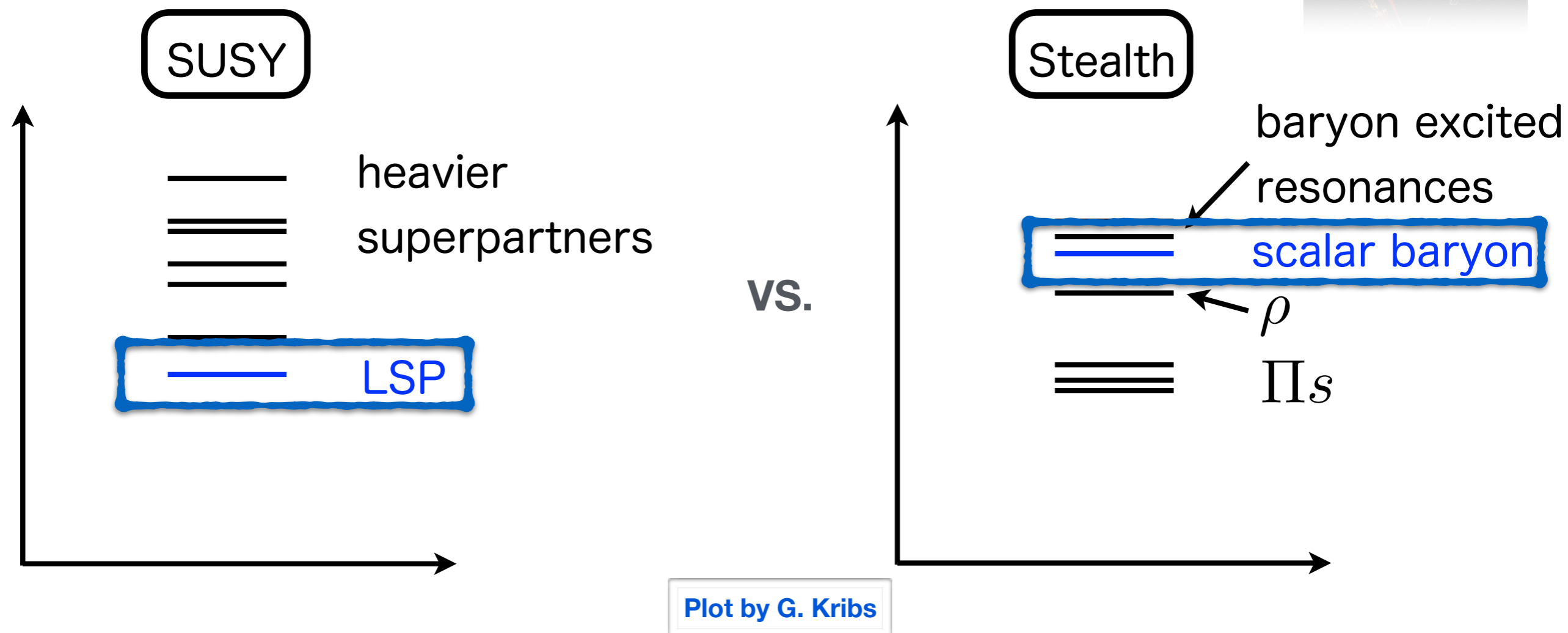
Stealth



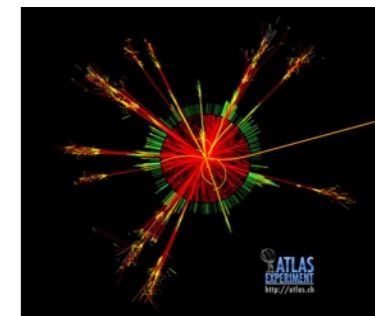
Plot by G. Kribs



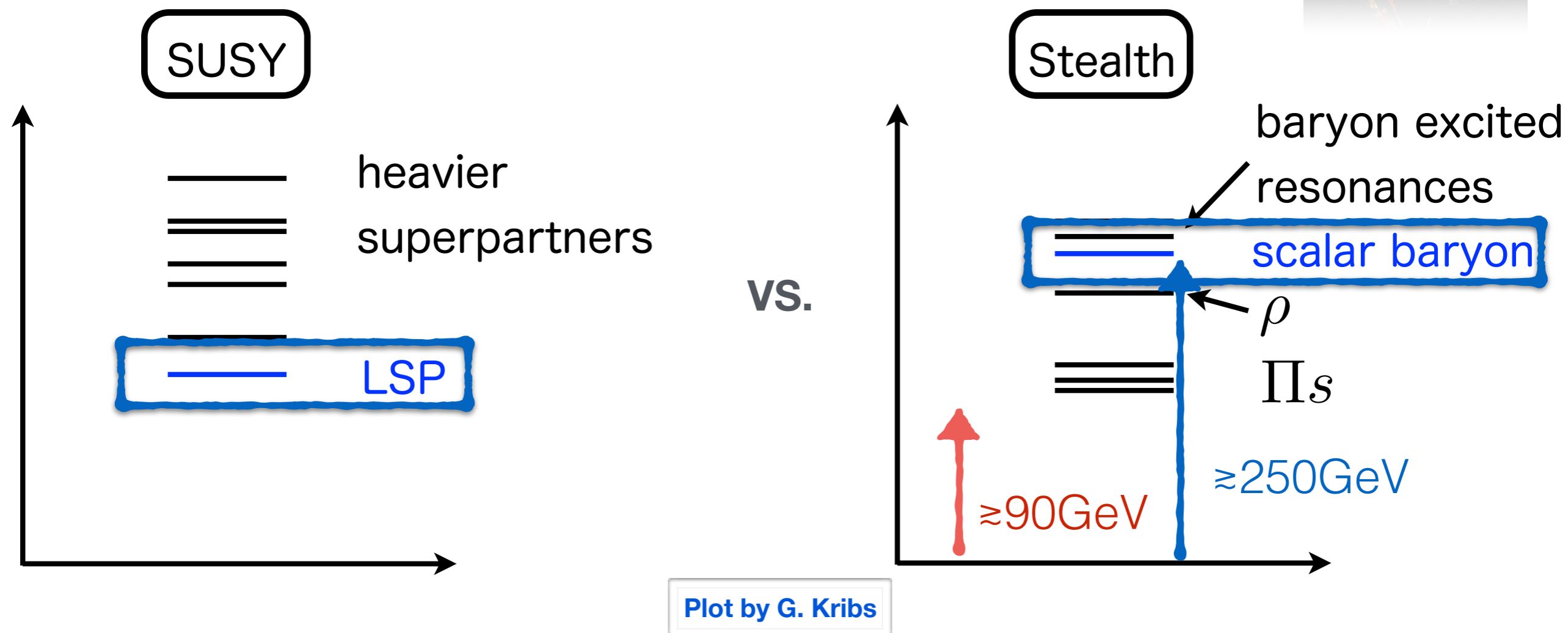
Stealth DM at colliders



- ◆ Signatures are not dominated by missing energy: **DM is not the lightest particle!** The interactions are suppressed (form factors)

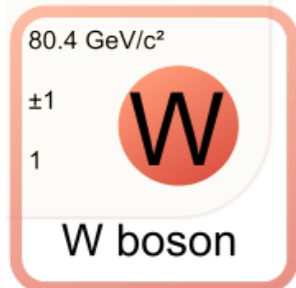
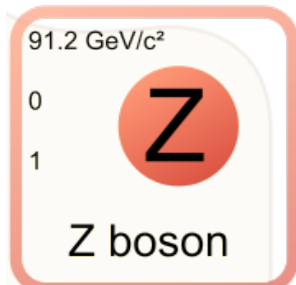
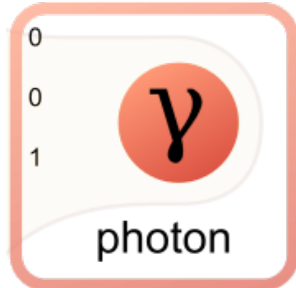


Stealth DM at colliders

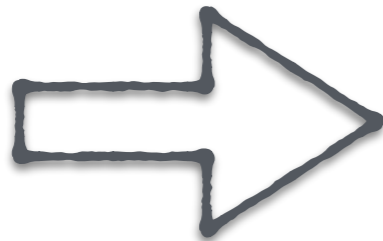
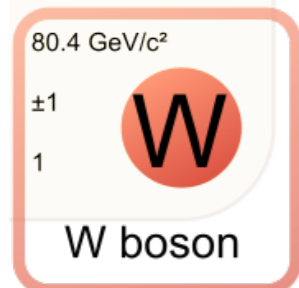
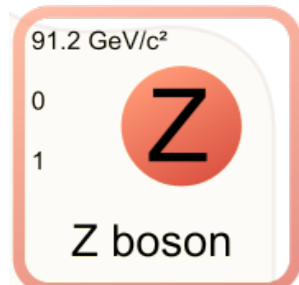
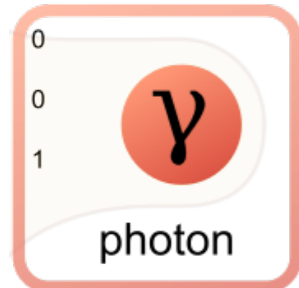


- ◆ Signatures are not dominated by missing energy: **DM is not the lightest particle!** The interactions are suppressed (form factors)
- ◆ Light meson production and decay give interesting signatures: **the model can be constrained by collider limits!** $m_{\Pi} \approx 90\text{GeV}$

The darkness of Composite Dark Matter



The darkness of Composite Dark Matter



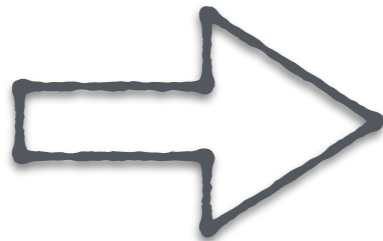
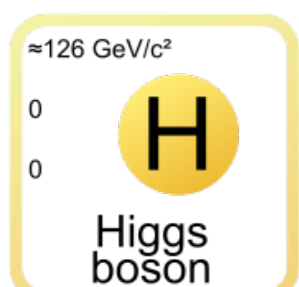
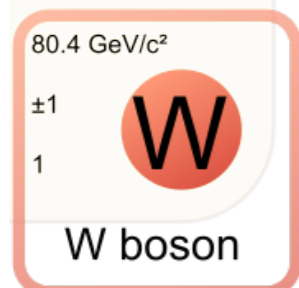
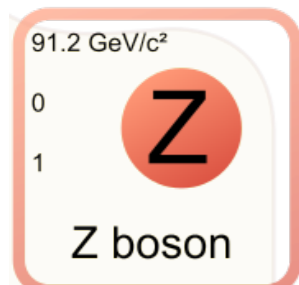
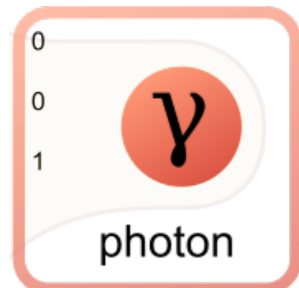
Lowest dimensional operators:

★ magnetic dipole (5)

★ charge radius (6)

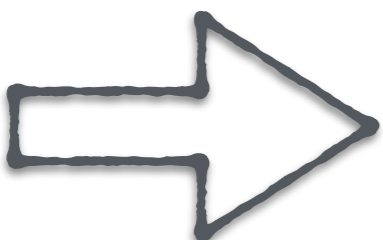
★ polarizability (7)

The darkness of Composite Dark Matter

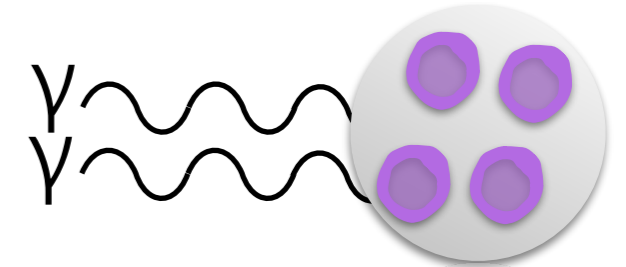


Lowest dimensional operators:

- ★ magnetic dipole (5)
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- ★ polarizability (7)

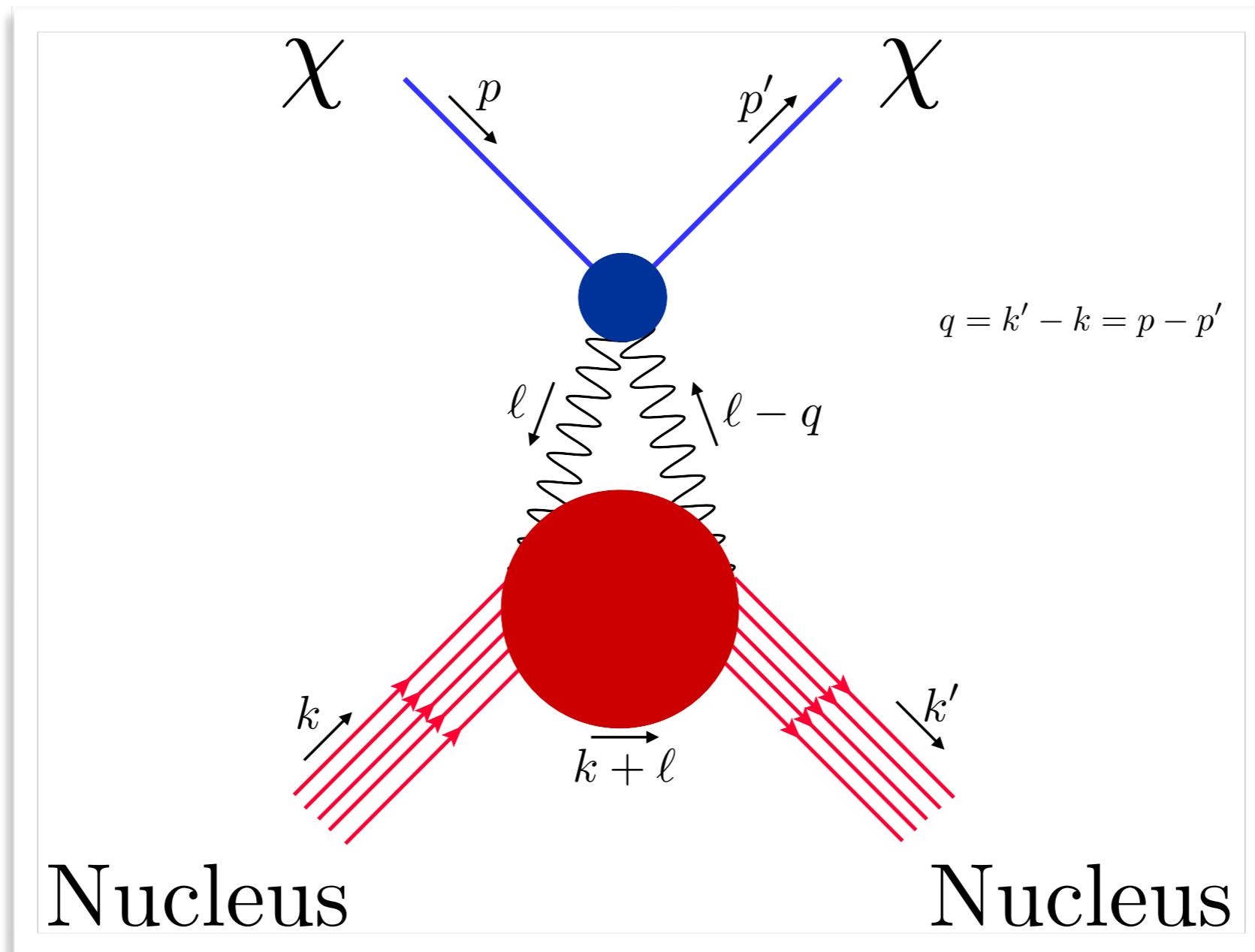


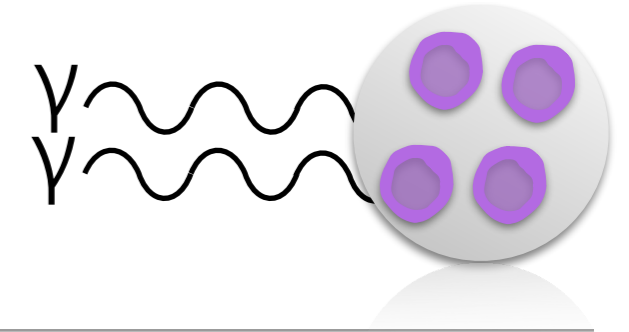
Most relevant interaction if constituents have Yukawa couplings!



Computing polarizability

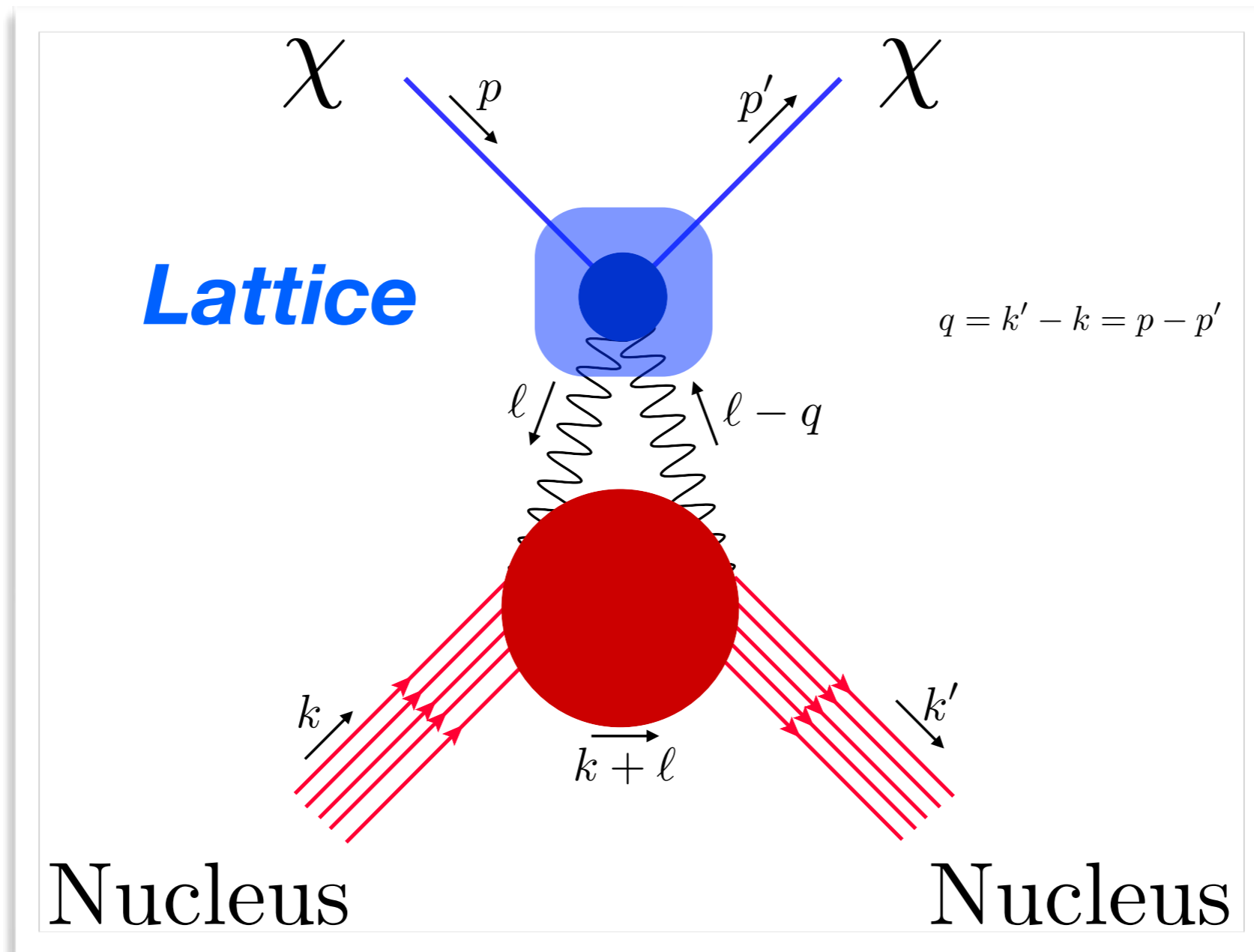
$$\frac{c_F e^2}{m_\chi^3} \chi^* \chi F^{\mu\alpha} F_\alpha^\nu v_\mu v_\nu$$

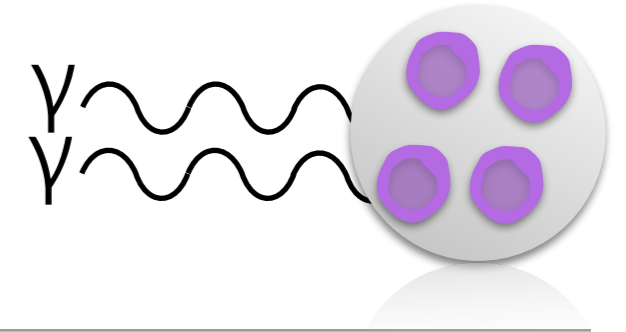




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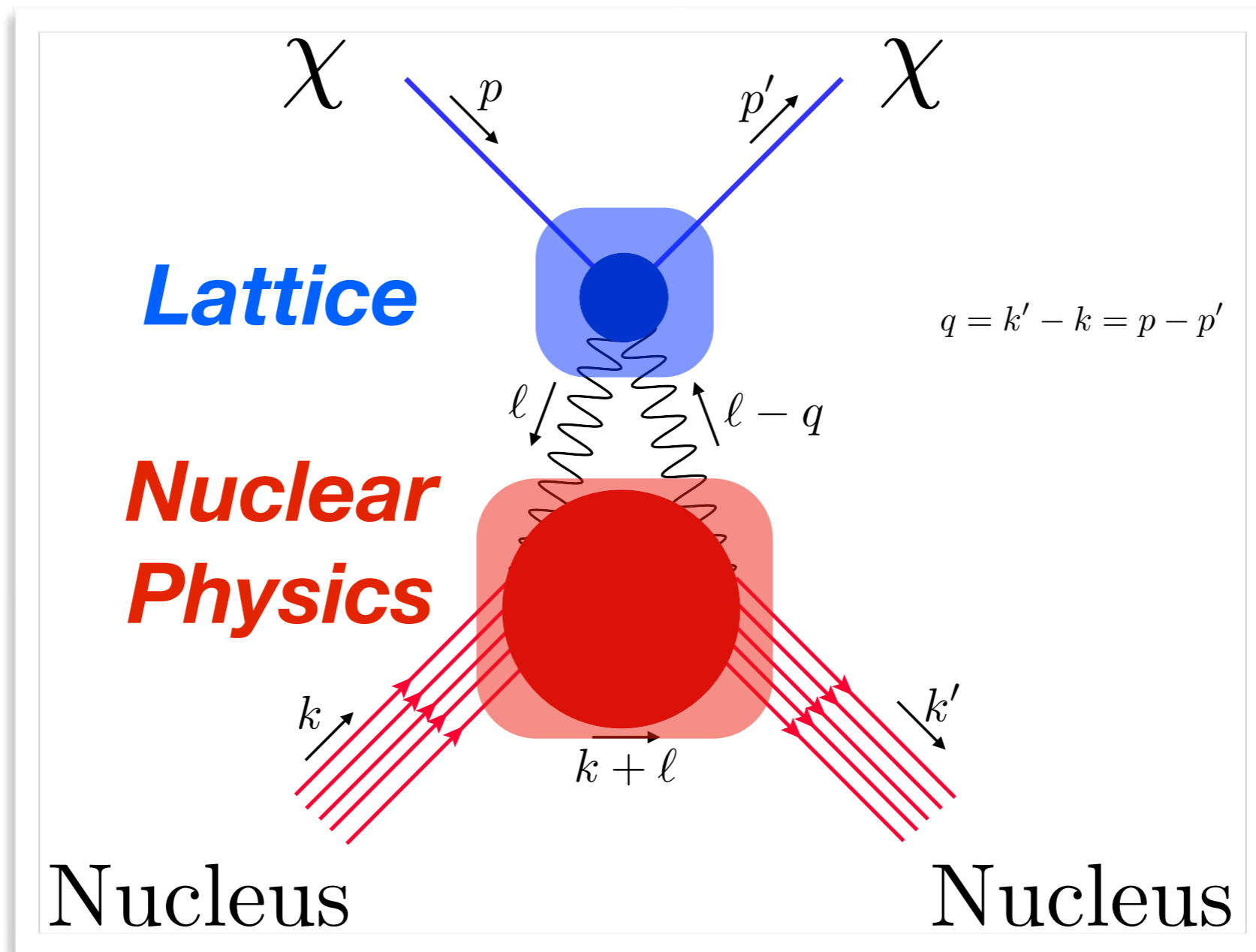
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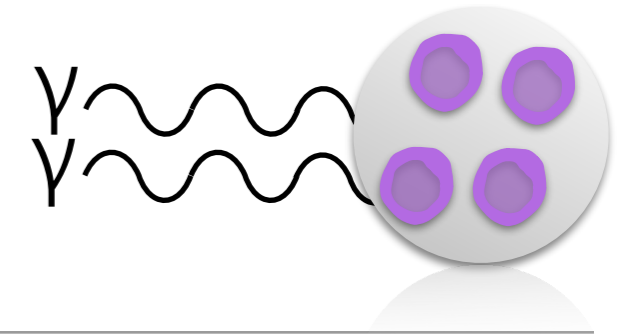




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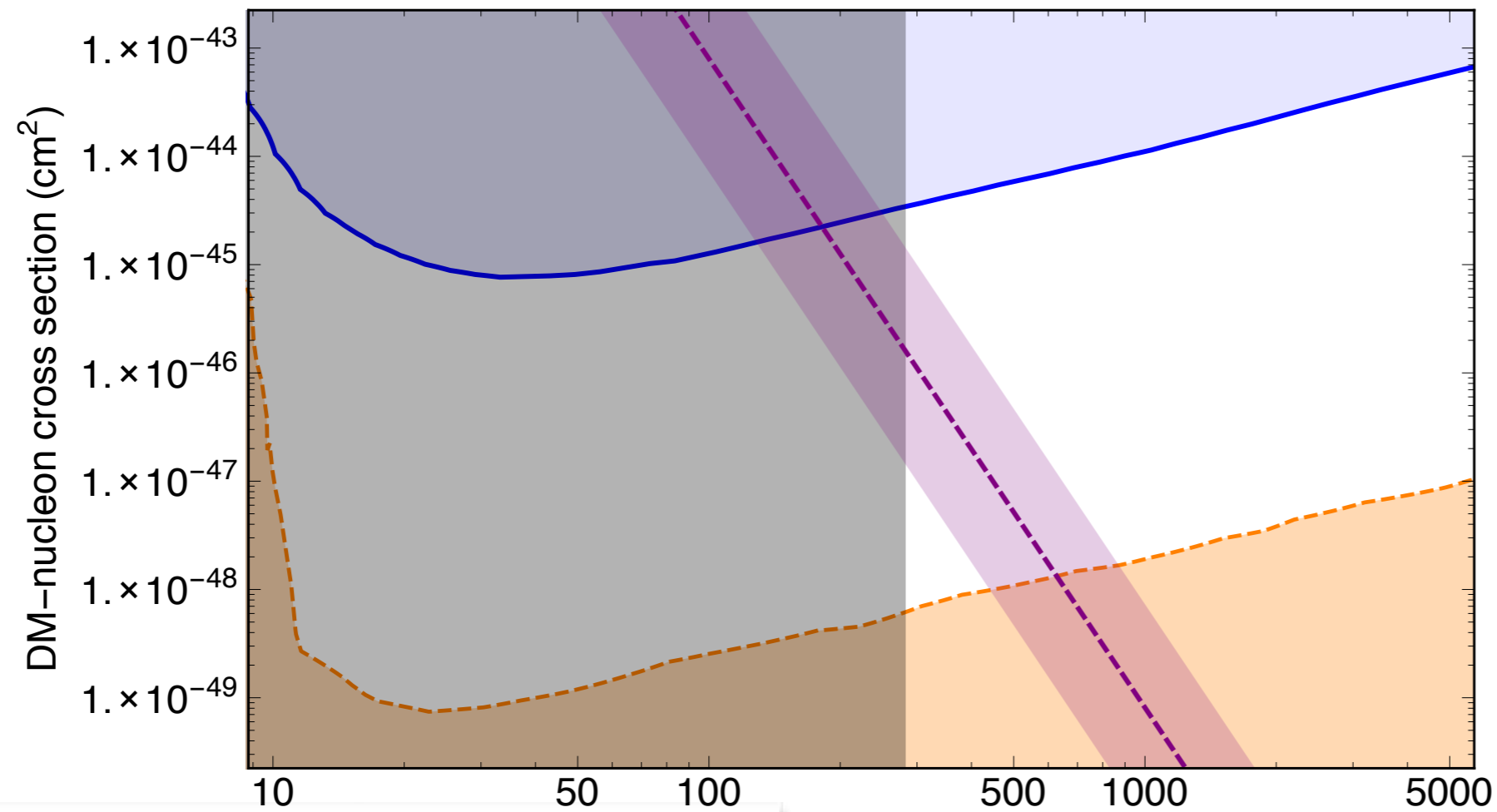
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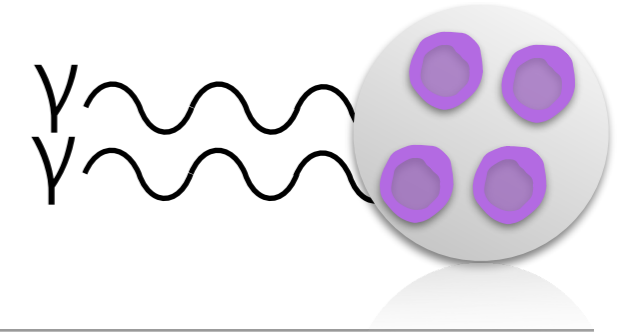


Lowest bound from EM polarizability

Electric polarizability from lattice simulations with background fields



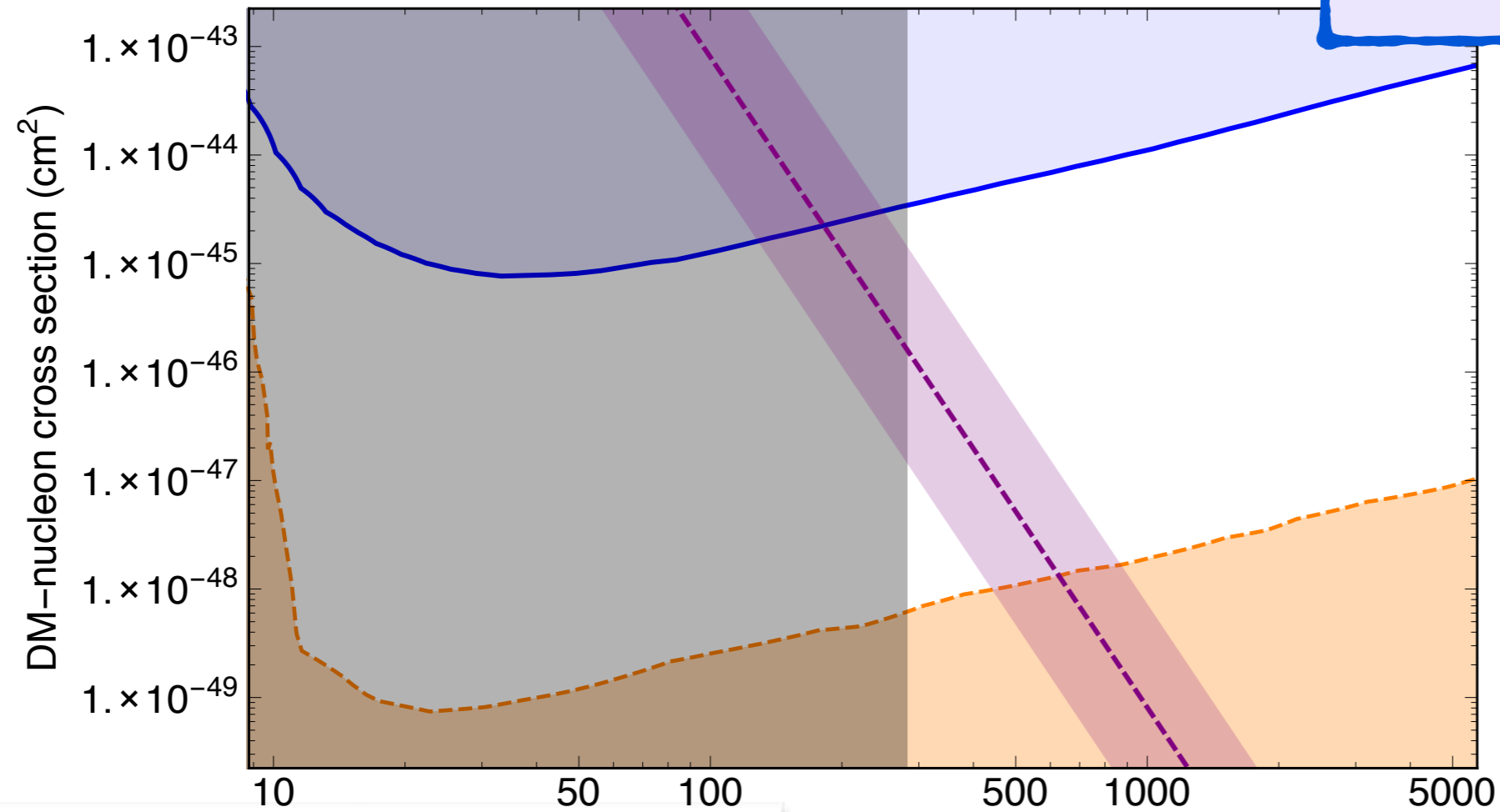
$$\sigma_{\text{nucleon}}(Z, A) = \frac{Z^4}{A^2} \frac{144\pi\alpha^4 \mu_{n\chi}^2 (M_F^A)^2}{m_\chi^6 R^2} [c_F]^2 M_\chi (\text{GeV})$$



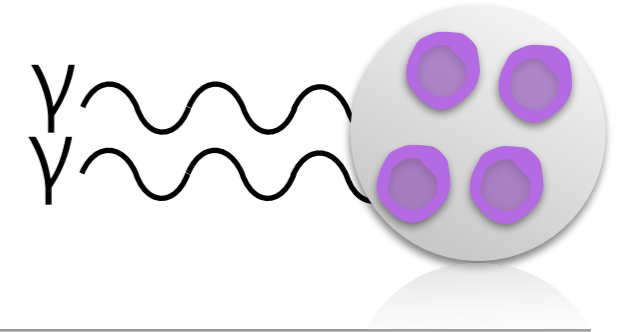
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Electric polarizability from lattice simulations with background fields

LUX exclusion bound for spin-independent cross section



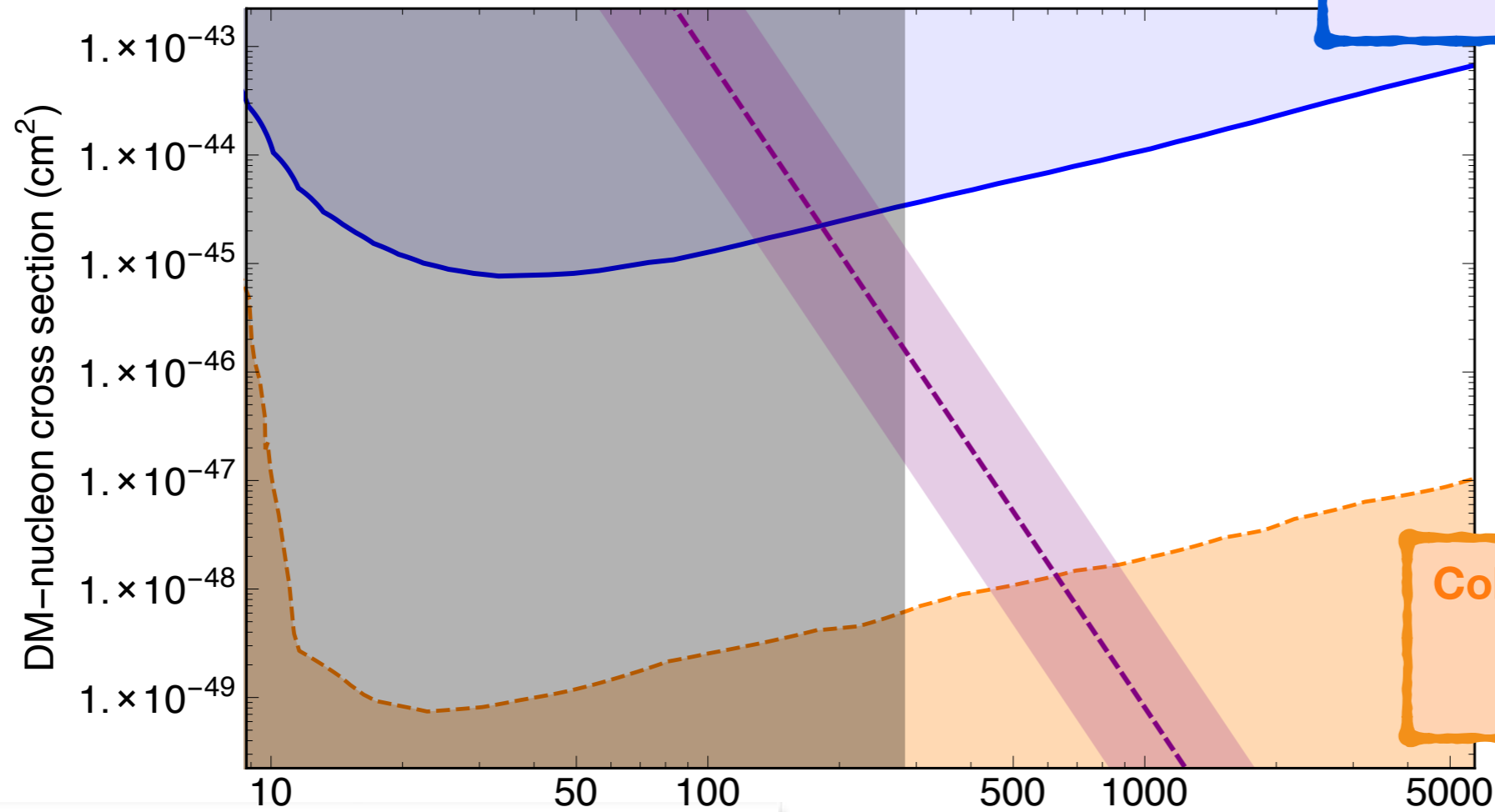
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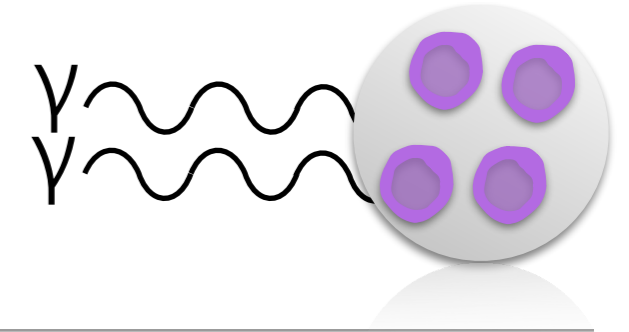
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LUX exclusion bound for spin-independent cross section



Coherent neutrino scattering background

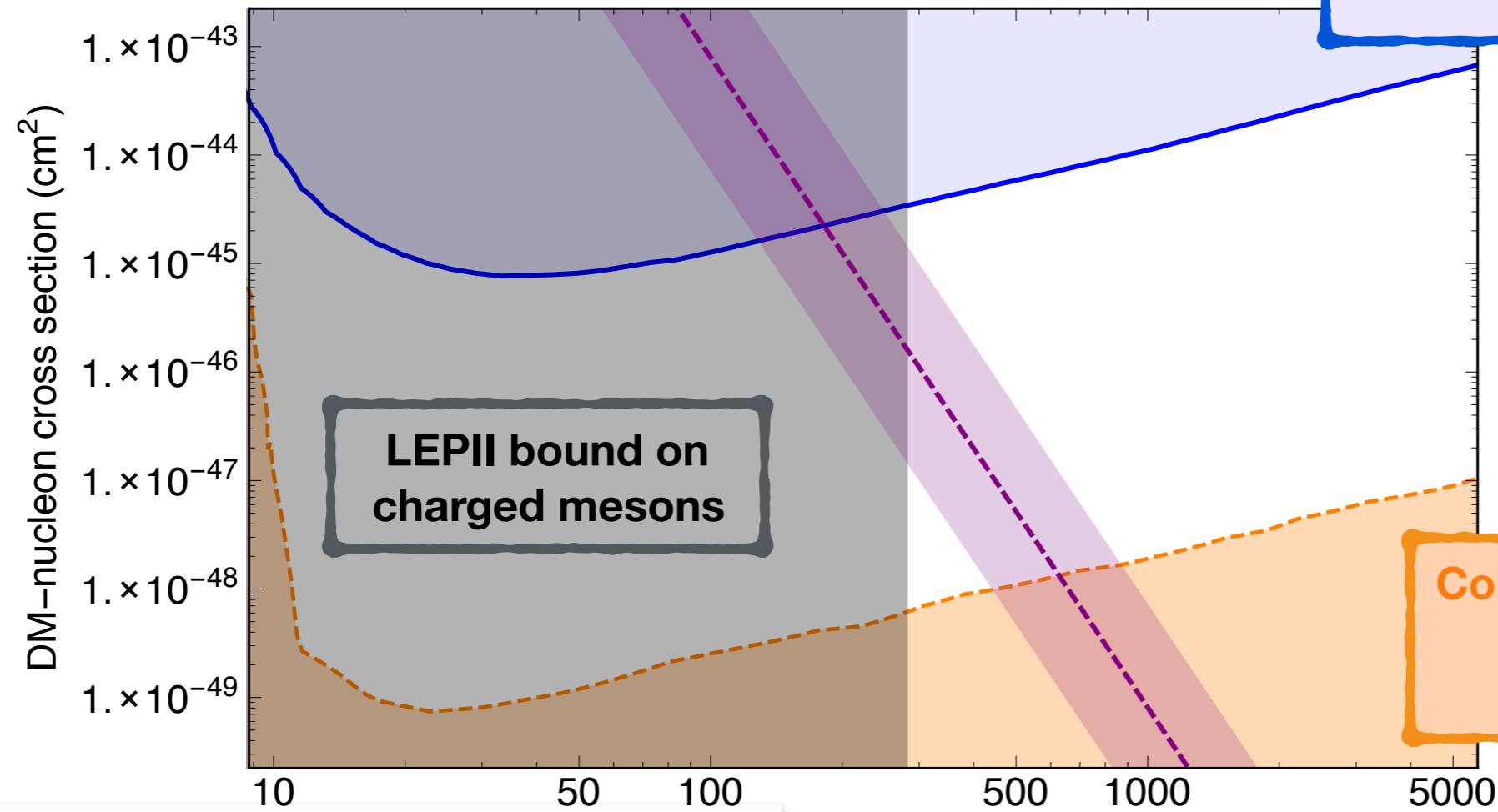
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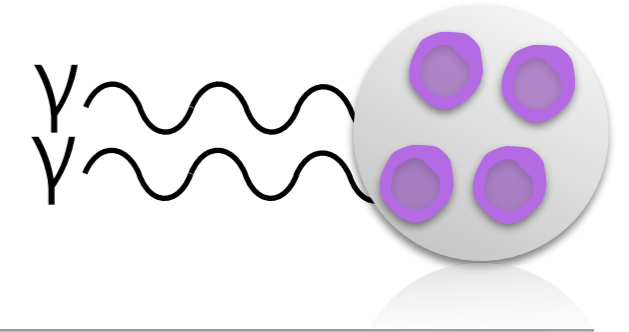
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LEP II bound on charged mesons

Coherent neutrino scattering background

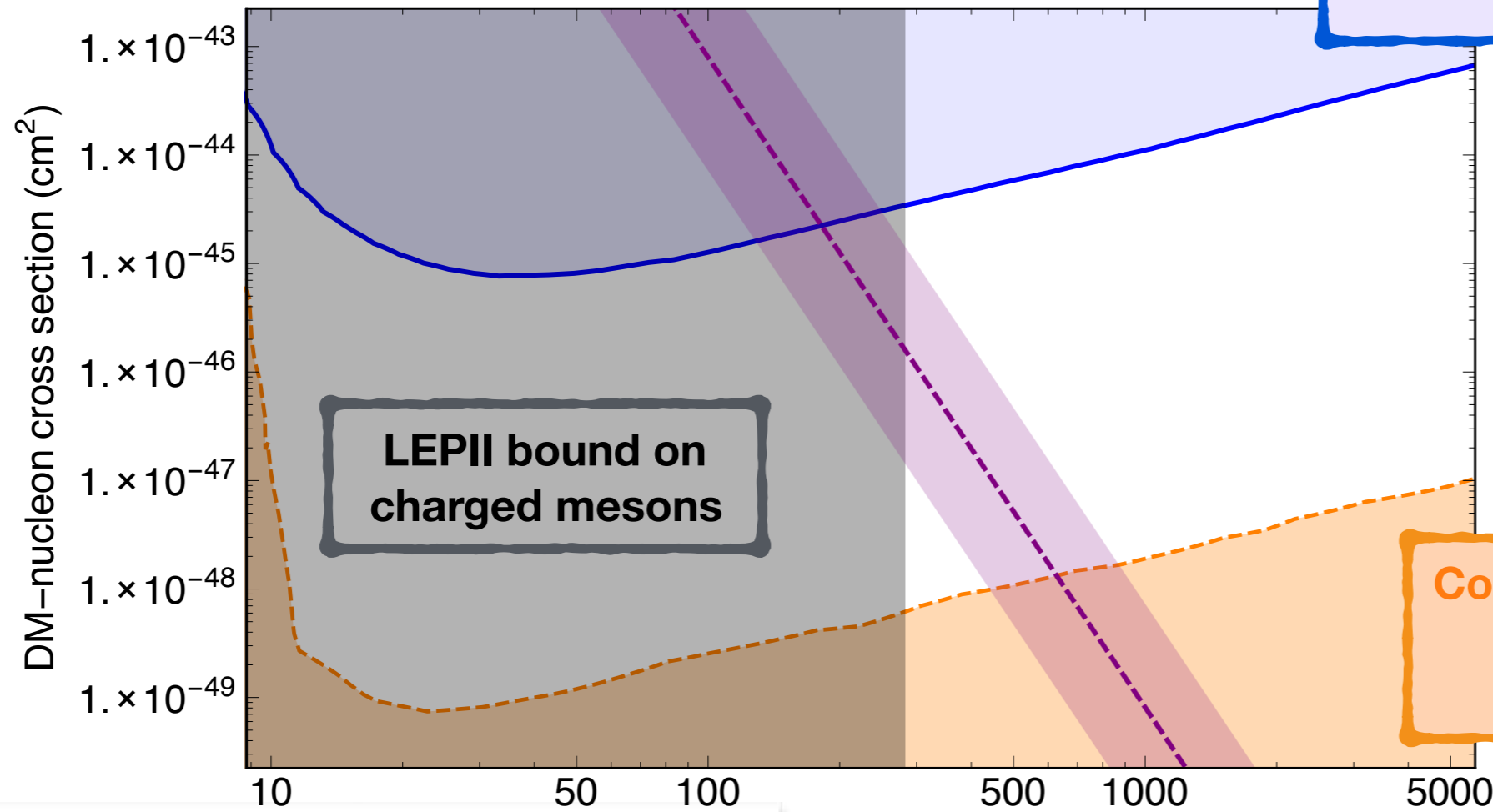
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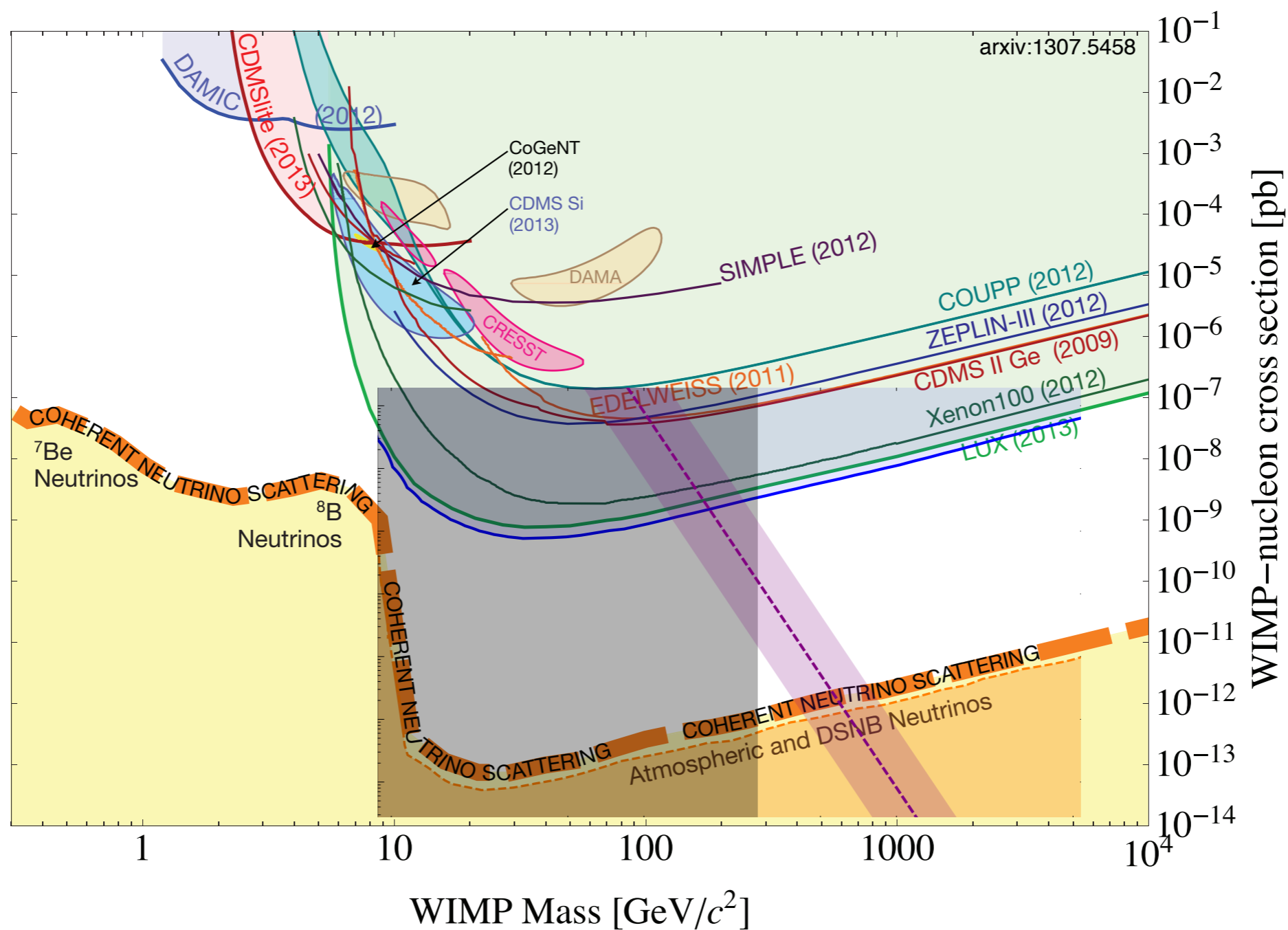
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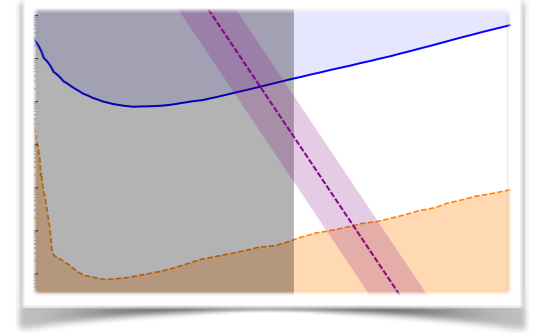
lowest allowed direct detection cross-section for composite dark matter theories with EW charged constituents

Stealth DM Polarizability



Stealth Dark Matter is excluded for $M_\chi \lesssim 300 \text{ GeV}$

Direct detection signal is below the neutrino coherent scattering background for $M_\chi \gtrsim 700 \text{ GeV}$



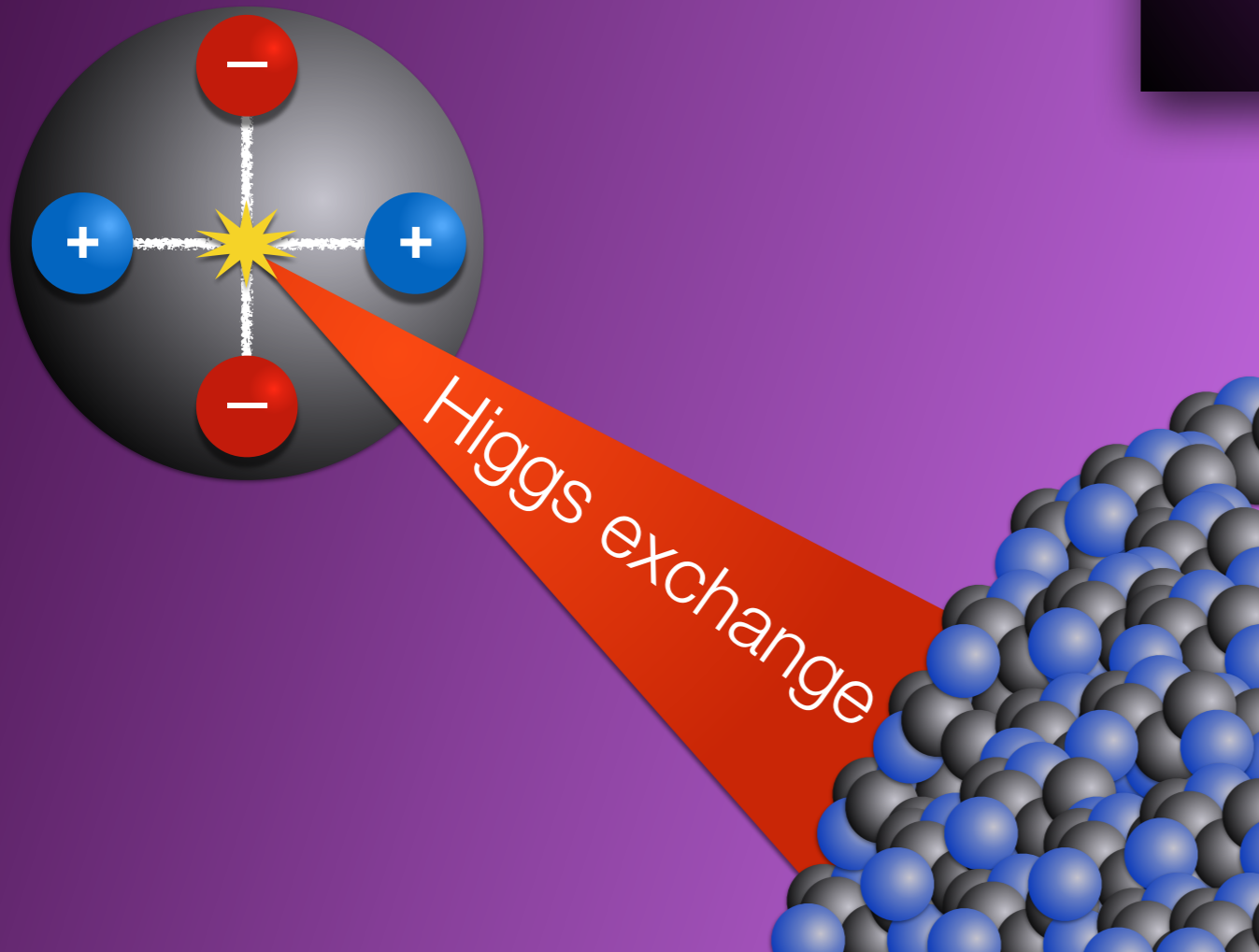
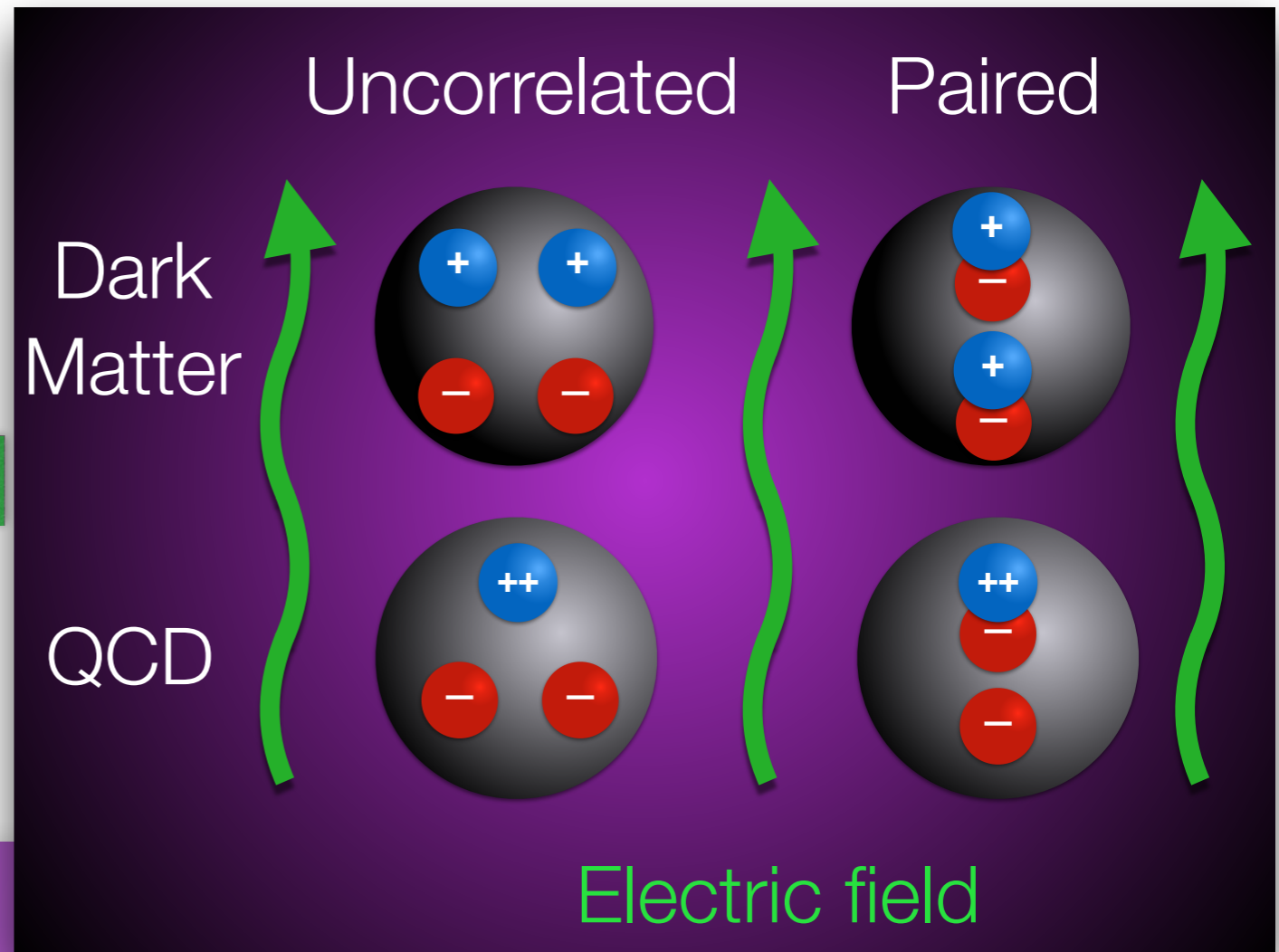
Concluding remarks

- ◆ **Composite** dark matter is a viable interesting possibility with rich phenomenology
- ◆ **Lattice** methods can help in calculating direct detection cross sections, production rates at colliders, and self-interaction cross sections. Direct phenomenological relevance.
- ◆ Dark matter constituents can carry electroweak charges and still the stable composites are currently undetectable. **Stealth cross section.**
- ◆ **Lowest bound for composite dark matter models: ~ 300 GeV** (colliders+direct detection+lattice)

extra

PRL Editors' Suggestion: Polarizability

[LSD collab., Phys. Rev. Lett. 115 (2015) 171803]



PRD Editors' Suggestion: Higgs exchange

[LSD collab., Phys. Rev. D92 (2015) 075030]

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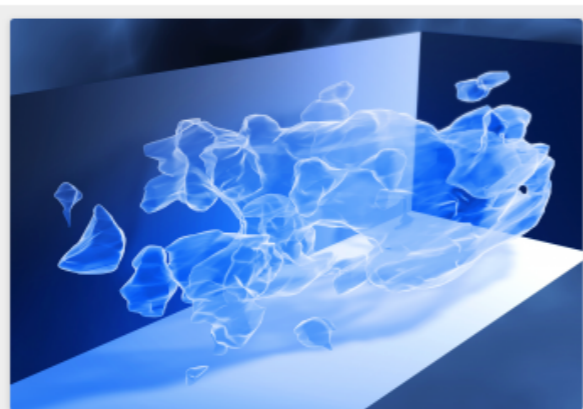
Materia oscura "stealth"

Quark oscuri tenuti insieme da un'interazione forte a sua volta oscura. Ecco come la dark matter riuscirebbe a eludere a ogni tentativo d'incastarla. Enrico Rinaldi (LLNL): «Esiste la possibilità che questo "mondo oscuro", con le sue nuove particelle, possa essere rivelato dagli esperimenti in corso al Large Hadron Collider al CERN di Ginevra»

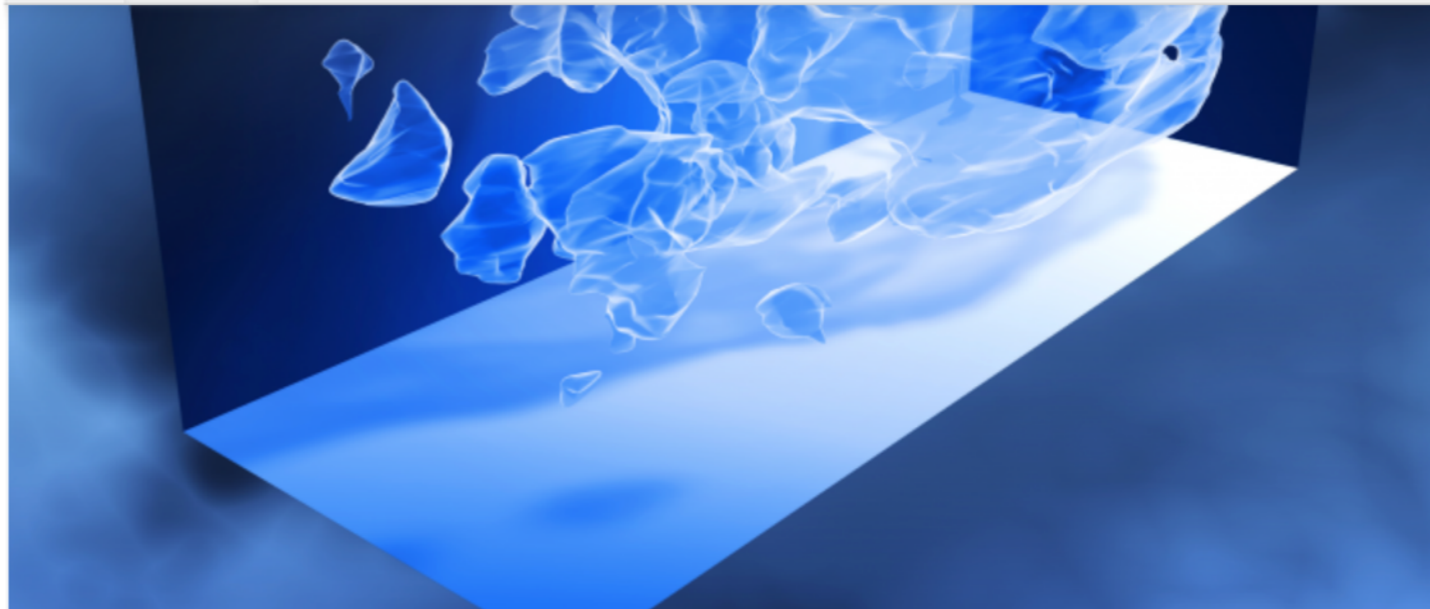
di Marco Malaspina | Segui @malamiao

venerdì 25 settembre 2015 @ 16:15

Stealth come furtiva. Stealth come imprevedibile. Stealth come quei minacciosi aerei da guerra dal profilo sagomato così da essere invisibili ai radar. Da quanto emerge dai calcoli dei fisici dell'LLNL, il Lawrence Livermore National Laboratory californiano, e dai modelli dati in pasto a Vulcan (un supercomputer per il calcolo parallelo in grado di masticare numeri al ritmo dei petaflop), sarebbe questa la natura della materia oscura: *stealthy*, appunto. Per forza non c'è ancora esperimento che sia riuscito a incastrarla.



Mappa 3D della distribuzione su larga scala della materia oscura ricostruita da misure di lente gravitazionale debole utilizzando il telescopio spaziale Hubble



This 3D map illustrates the large-scale distribution of dark matter, reconstructed from measurements of weak gravitational lensing by using the Hubble Space Telescope. (Download Image)

New 'stealth dark matter' theory may explain mystery of the universe's missing mass



Lawrence Livermore National Laboratory (LLNL) scientists have come up with a new theory that may identify why dark matter has evaded direct detection in Earth-based experiments.

Anne M Stark
stark8@llnl.gov
925-422-9799

Detecting Stealth Dark Matter Directly through Electromagnetic Polarizability.

Overview of attention for article published in Physical Review Letters, October 2015



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Title Detecting Stealth Dark Matter Directly through Electromagnetic Polarizability.
Published in Physical Review Letters, October 2015
DOI 10.1103/physrevlett.115.171803
Pubmed ID 26551103
Authors T. Appelquist, E. Berkowitz, R. C. Brower, M. I. Buchoff, G. T. Fleming, X.-Y. Jin, J. Kiskis, G. D...
Abstract We calculate the spin-independent scattering cross section for direct detection that results from...
Abstract We calculate the spin-independent scattering cross section for direct detection that results from...

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Materia oscura
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In edicola dal 1 settembre

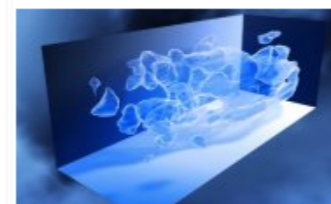
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ZOOM SU | comportamento | cosmologia | neuroscienze | alimentazione

iflammenco! | Festival Flamenco 5-11 ottobre

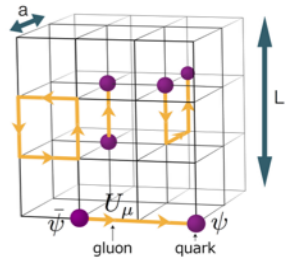
28 settembre 2015

Un nuovo modello per la materia oscura

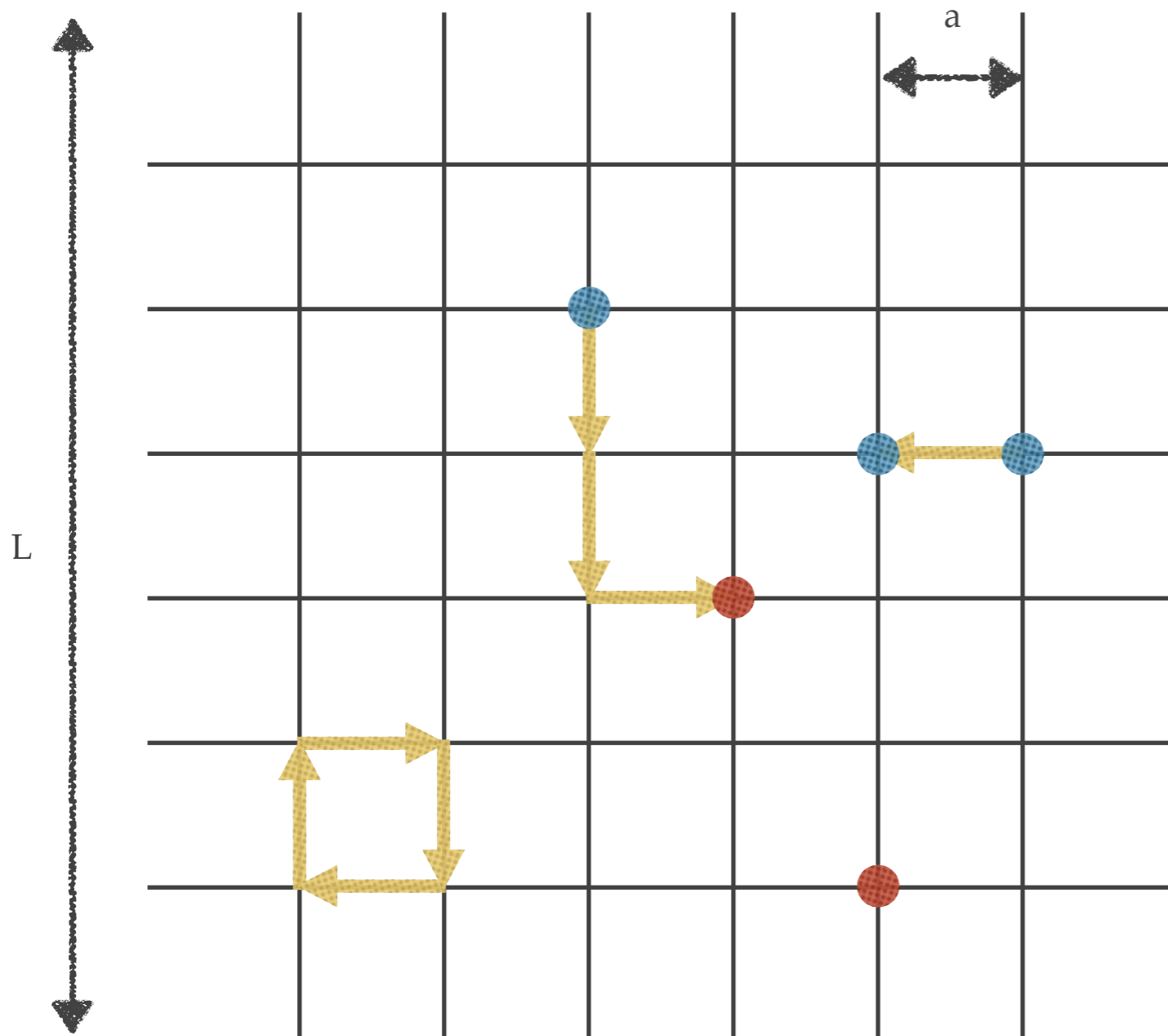


Cortesia Lawrence Livermore National Laboratory

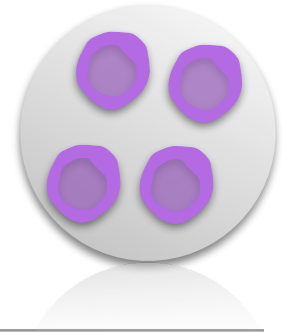
Questa forma misteriosa di materia potrebbe avere una struttura composita come la materia ordinaria, con "quark oscuri" aggregati e tenuti insieme da un analogo della forza che permette ai normali nuclei di rimanere stabili. I componenti di questo tipo di materia oscura, definita stealth matter, potrebbero essere studiati in modo indiretto dal collisore Large Hadron Collider del CERN di Ginevra



Lattice Gauge Theory - basics



- Discretize space and time
 - lattice spacing “a”
 - lattice size “L”
- Keep all d.o.f. of the theory
 - not a model!
 - no simplifications
- Amenable to numerical methods
 - Monte Carlo sampling
 - use supercomputers
- Precisely quantifiable and improvable errors
 - Systematic
 - Statistical



“Stealth Dark Matter” model

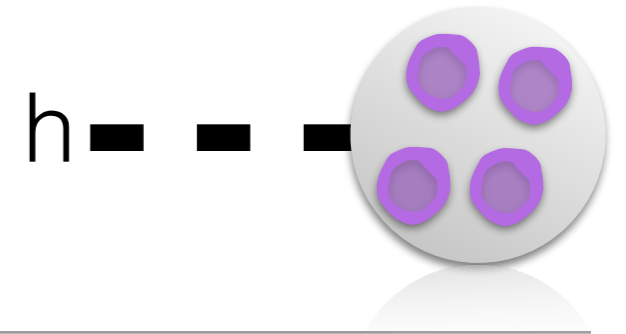
- The field content of the model consists in **8 Weyl fermions**
- Dark fermions interact with the SM Higgs and obtain **current/chiral masses**
- Introduce **vector-like masses** for dark fermions that do not break EW symmetry
- Diagonalizing in the mass eigenbasis gives **4 Dirac fermions**
- Assume **custodial SU(2) symmetry** arising when **$u \leftrightarrow d$**

Field	$SU(N)_D$	$(SU(2)_L, Y)$	Q
$F_1 = \begin{pmatrix} F_1^u \\ F_1^d \end{pmatrix}$	\mathbf{N}	$(\mathbf{2}, 0)$	$\begin{pmatrix} +1/2 \\ -1/2 \end{pmatrix}$
$F_2 = \begin{pmatrix} F_2^u \\ F_2^d \end{pmatrix}$	$\overline{\mathbf{N}}$	$(\mathbf{2}, 0)$	$\begin{pmatrix} +1/2 \\ -1/2 \end{pmatrix}$
F_3^u	\mathbf{N}	$(\mathbf{1}, +1/2)$	$+1/2$
F_3^d	\mathbf{N}	$(\mathbf{1}, -1/2)$	$-1/2$
F_4^u	$\overline{\mathbf{N}}$	$(\mathbf{1}, +1/2)$	$+1/2$
F_4^d	$\overline{\mathbf{N}}$	$(\mathbf{1}, -1/2)$	$-1/2$

$$\mathcal{L} \supset + y_{14}^u \epsilon_{ij} F_1^i H^j F_4^d + y_{14}^d F_1 \cdot H^\dagger F_4^u - y_{23}^d \epsilon_{ij} F_2^i H^j F_3^d - y_{23}^u F_2 \cdot H^\dagger F_3^u + h.c.$$

$$\mathcal{L} \supset M_{12} \epsilon_{ij} F_1^i F_2^j - M_{34}^u F_3^u F_4^d + M_{34}^d F_3^d F_4^u + h.c.$$

$$y_{14}^u = y_{14}^d \quad y_{23}^u = y_{23}^d \quad M_{34}^u = M_{34}^d$$



Computing Higgs exchange

- ◆ Need to **non-perturbatively** evaluate the dark **σ -term**
- ◆ **Effective Higgs coupling** non-trivial with mixed chiral and vector-like masses
- ◆ Model-dependent answer for the cross-section
- ◆ **Lattice input** is necessary: compute mass and form factor

$$\mathcal{M}_a = \frac{y_f y_q}{2m_h^2} \sum_f \langle B | \bar{f} f | B \rangle \sum_q \langle a | \bar{q} q | a \rangle$$

1. effective Higgs coupling with dark fermions and quark Yukawa coupling
2. dark baryon scalar form factor: need lattice input for generic DM models!
3. nucleon scalar form factor: ChPT and lattice input [Plenary talk by Collins, Tue@10:15]

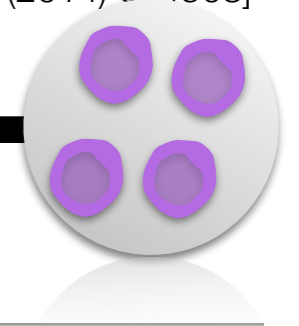
$$y_f \langle B | \bar{f} f | B \rangle = \frac{m_B}{v} \sum_f \left. \frac{v}{m_f} \frac{\partial m_f(h)}{\partial h} \right|_{h=v} f_f^{(B)}$$

$$m_f(h) = m + \frac{y_f h}{\sqrt{2}}$$

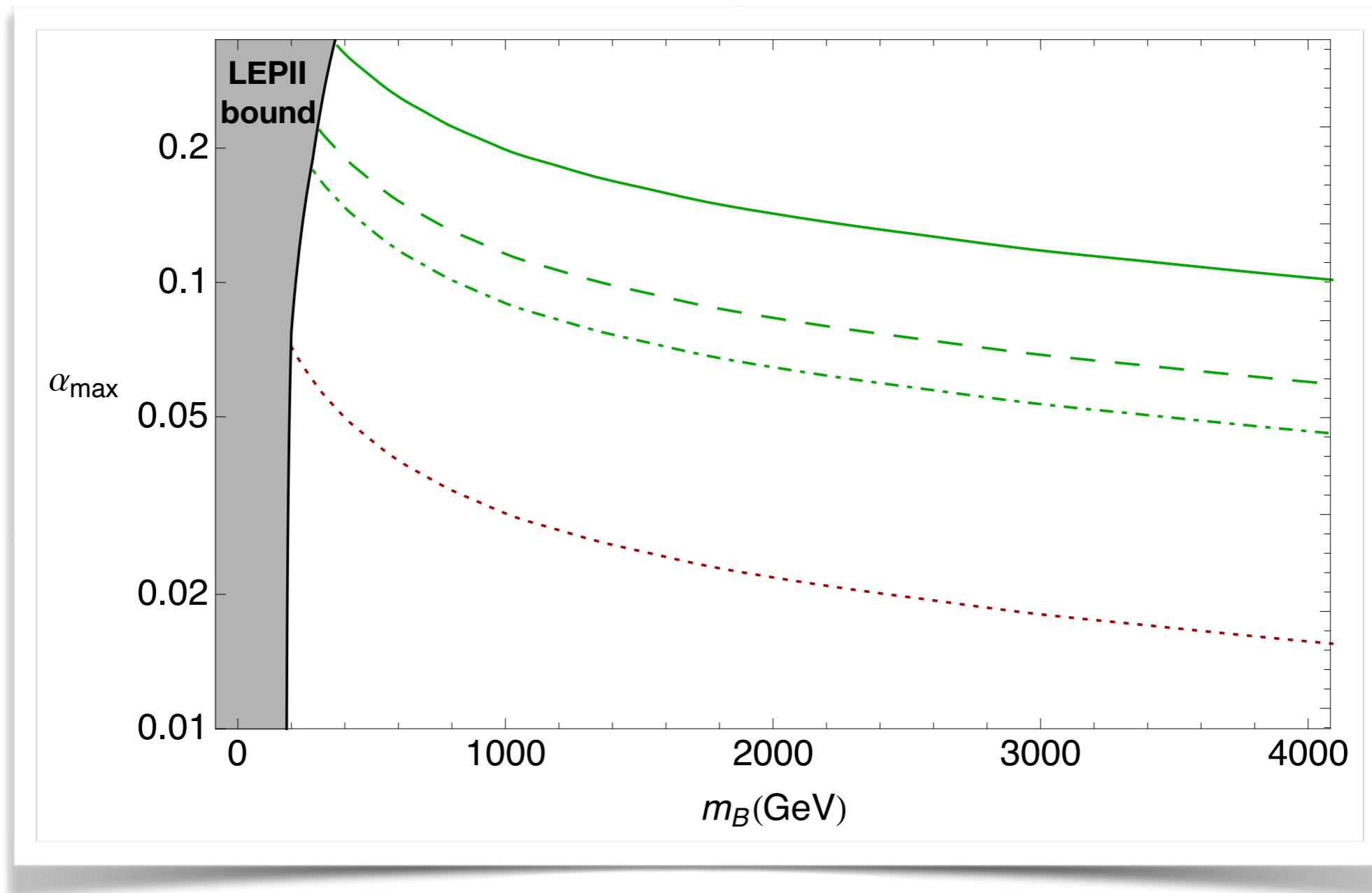
$$\alpha \equiv \left. \frac{v}{m_f} \frac{\partial m_f(h)}{\partial h} \right|_{h=v} = \frac{y v}{\sqrt{2} m + y v}$$

Lattice!

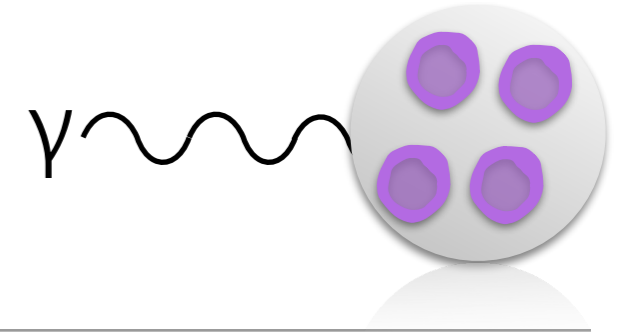
h ■ ■ ■



Bounds on the Yukawa coupling



$$\alpha \equiv \frac{v}{m_f} \left. \frac{\partial m_f(h)}{\partial h} \right|_{h=v} = \frac{yv}{\sqrt{2}m + yv}$$

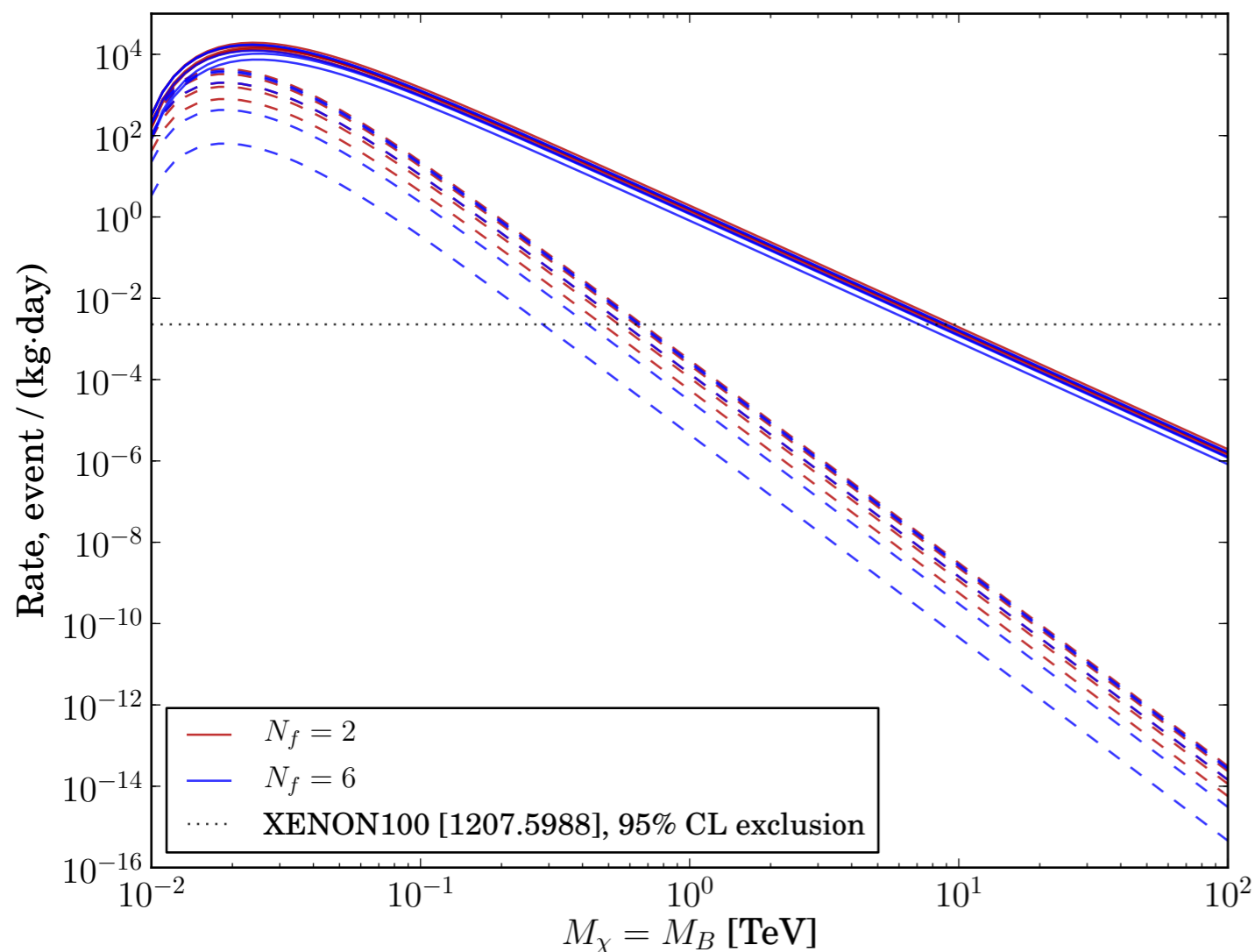


Bounds from EM moments

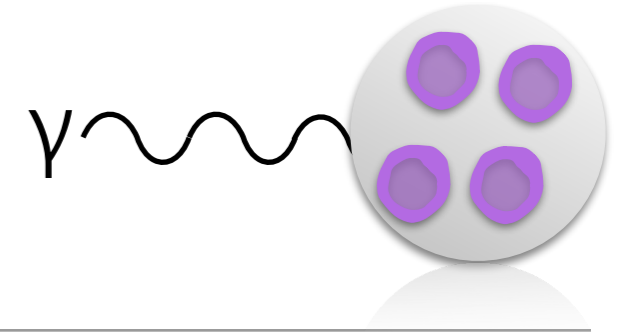
Mesonic and Baryonic EM form factors
directly from lattice simulations

SU(3) $N_f=2,6$ dark fermionic baryon

[LSD, 1301.1693]



- ★ baryon similar to QCD neutron
- ★ dark quarks with $Q=Y$
- ★ calculate connected 3pt
- ★ scale set by DM mass
- ★ magnetic moment dominates
- ★ results independent of N_f

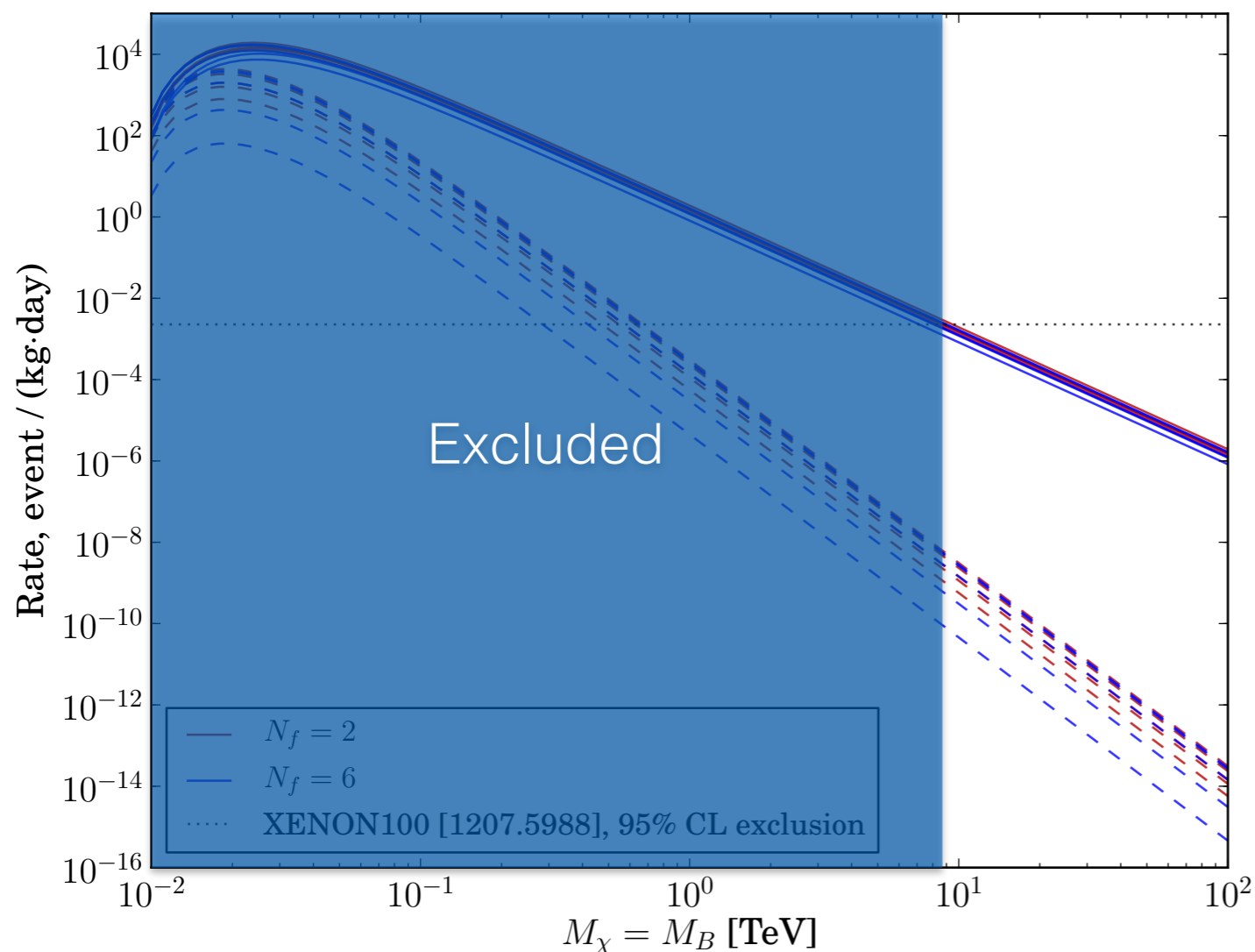


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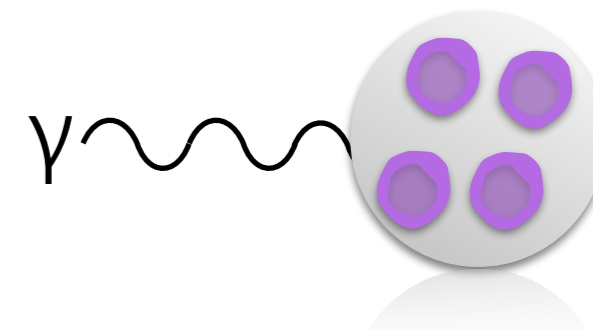
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$M_B > \sim 10$ TeV

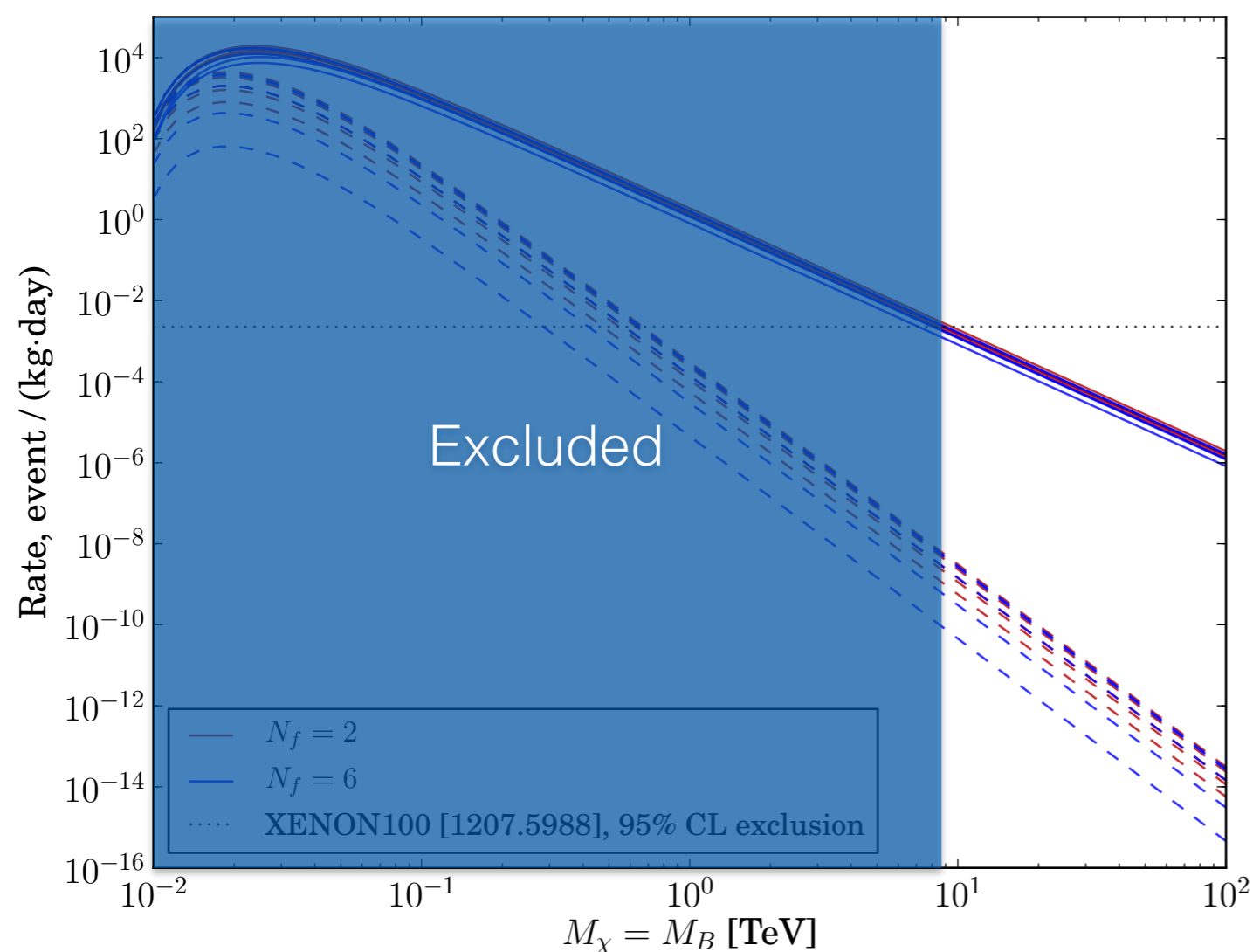


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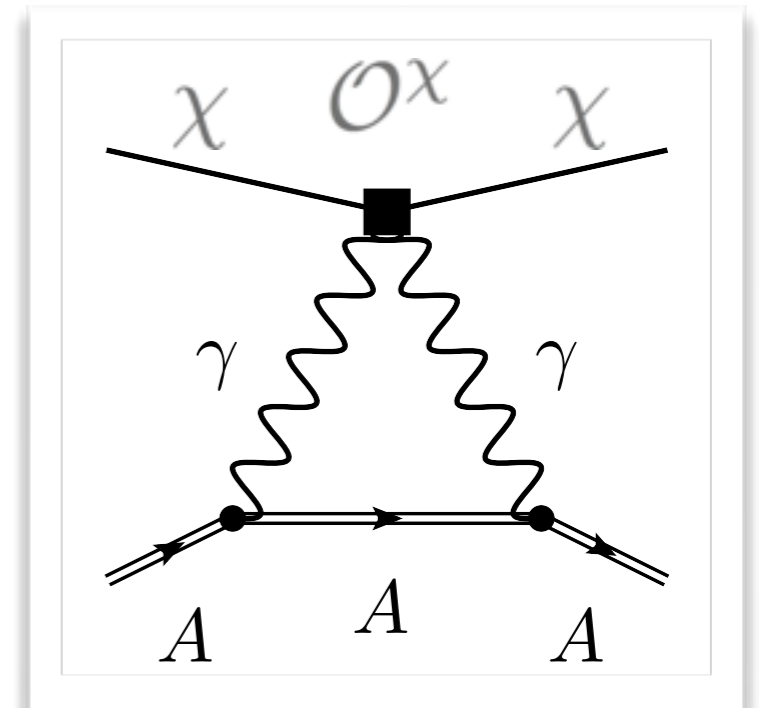
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$M_B > \sim 10$ TeV

pushed to ~ 100 TeV
with new LUX

Nuclear: Rayleigh scattering

- it is hard to extract the momentum dependence of this nuclear form factor
- similarities with the double-beta decay nuclear matrix element could suggest large uncertainties ~ orders of magnitude
- to assess the impact of uncertainties on the total cross section we start from naive dimensional analysis
- we allow a “magnitude” factor M_F^A to change from 0.3 to 3



$$f_F^A = \langle A | F^{\mu\nu} F_{\mu\nu} | A \rangle$$

$$f_F^A \sim 3 Z^2 \alpha \frac{M_F^A}{R}$$

$$\sigma \simeq \frac{\mu_{n\chi}^2}{\pi A^2} \left\langle \left| \frac{c_F e^2}{m_\chi^3} f_F^A \right|^2 \right\rangle$$

[Pospelov & Veldhuis, Phys. Lett. B480 (2000) 181]

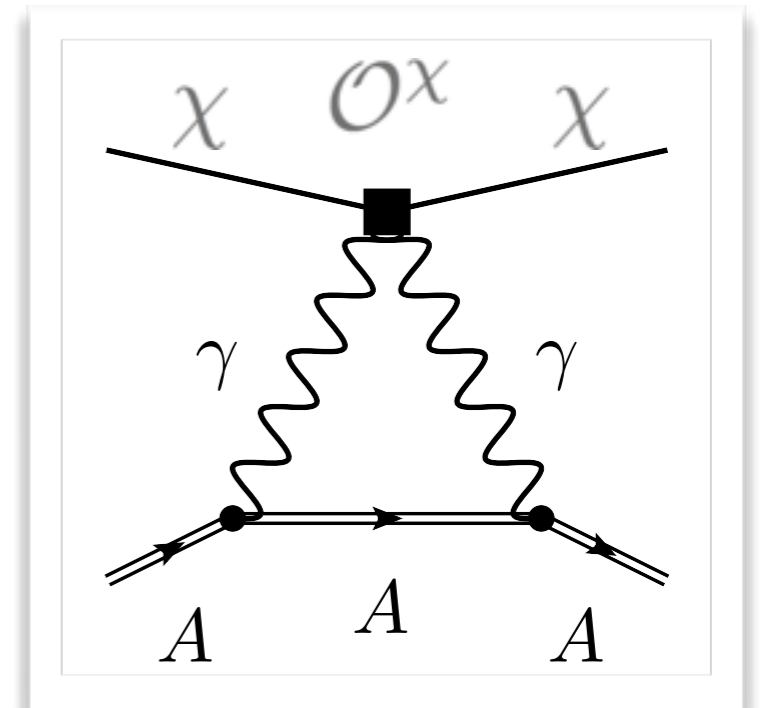
[Weiner & Yavin, Phys. Rev. D86 (2012) 075021]

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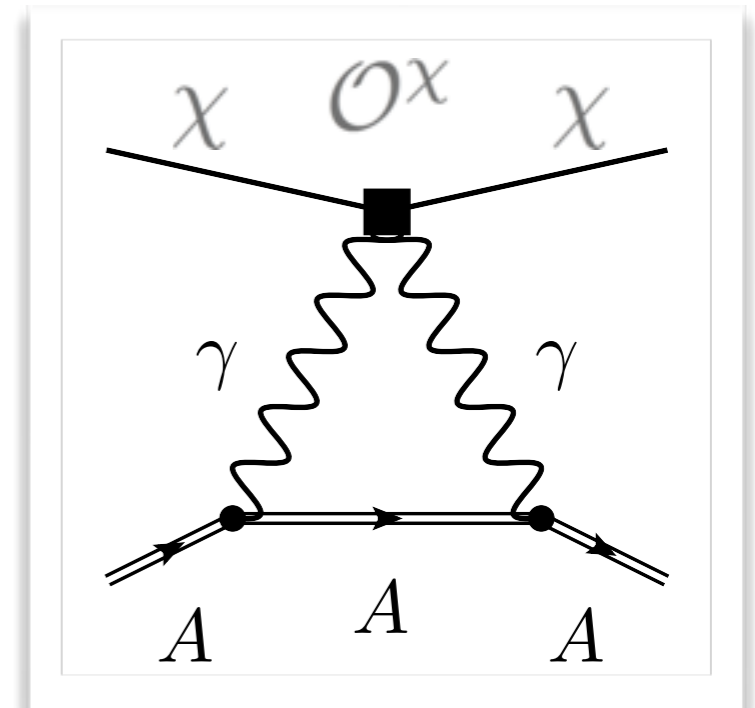
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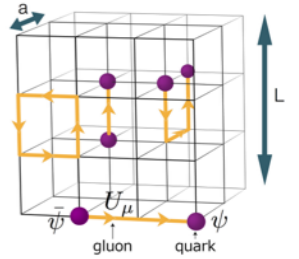
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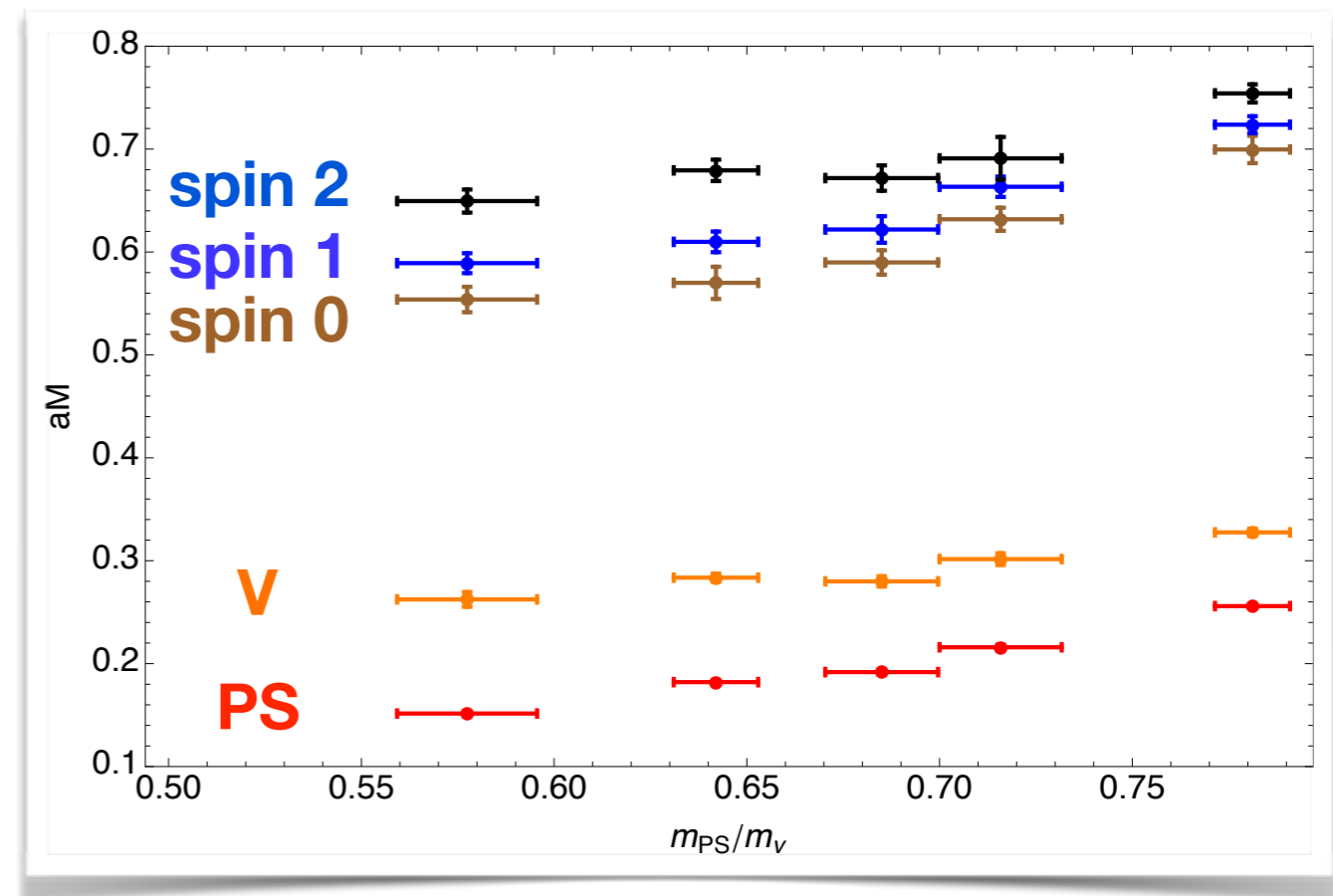
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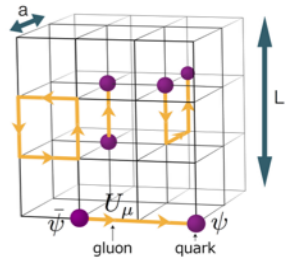
Lattice Stealth DM

- Non-perturbative lattice calculations of the spectrum confirm that **lightest baryon has spin zero**
- The ratio of **pseudoscalar (PS)** to **vector (V)** is used as probe for different dark fermion masses
- The meson to baryon mass ratio allows us to translate LEP II bounds on charged meson to **LEP bounds on composite bosonic dark matter**



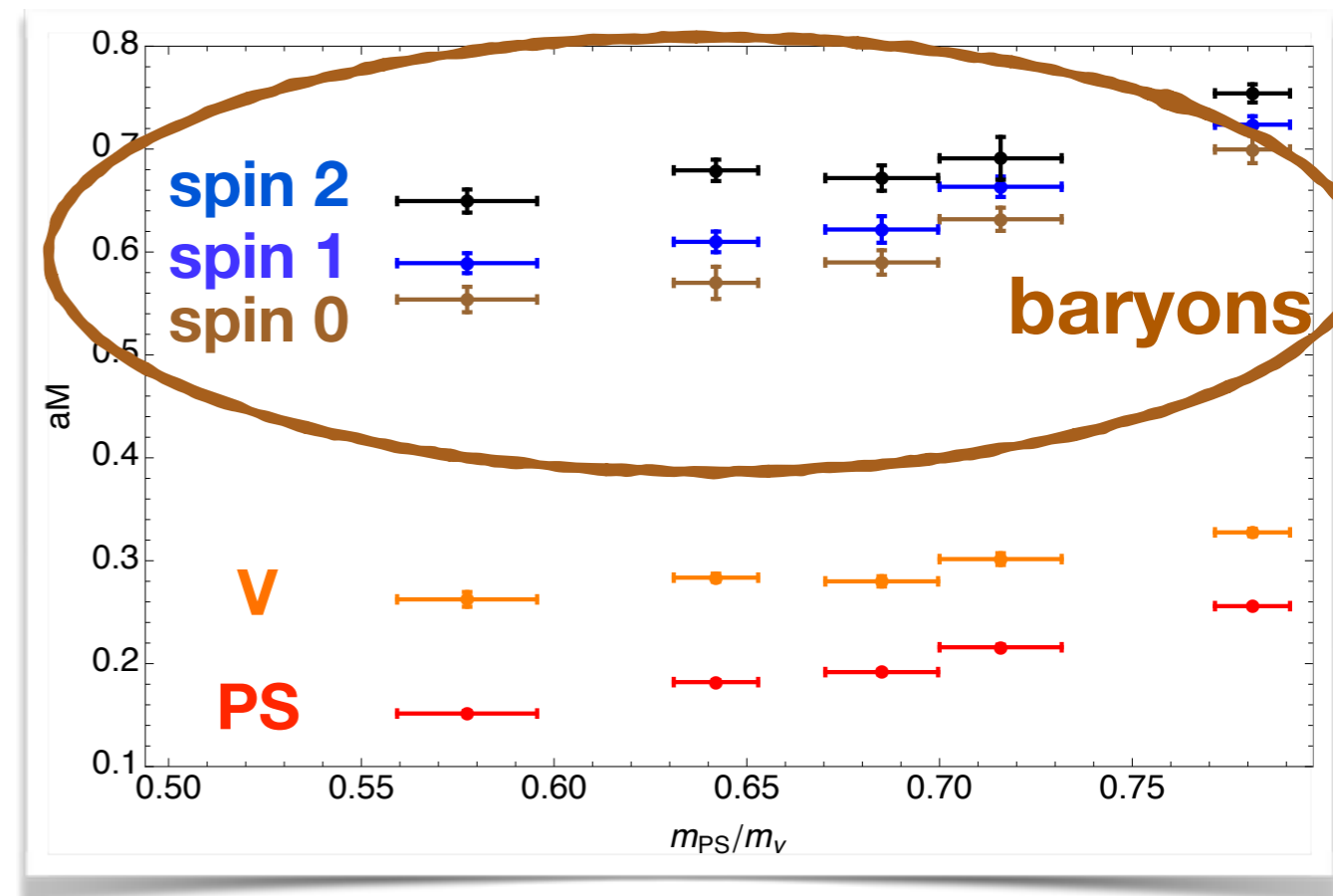
[LSD collab., Phys. Rev. D89 (2014) 094508]

- Study **systematic effects** due to lattice discretization and finite volume due to the relative unfamiliar nature of the system



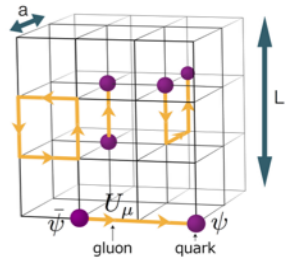
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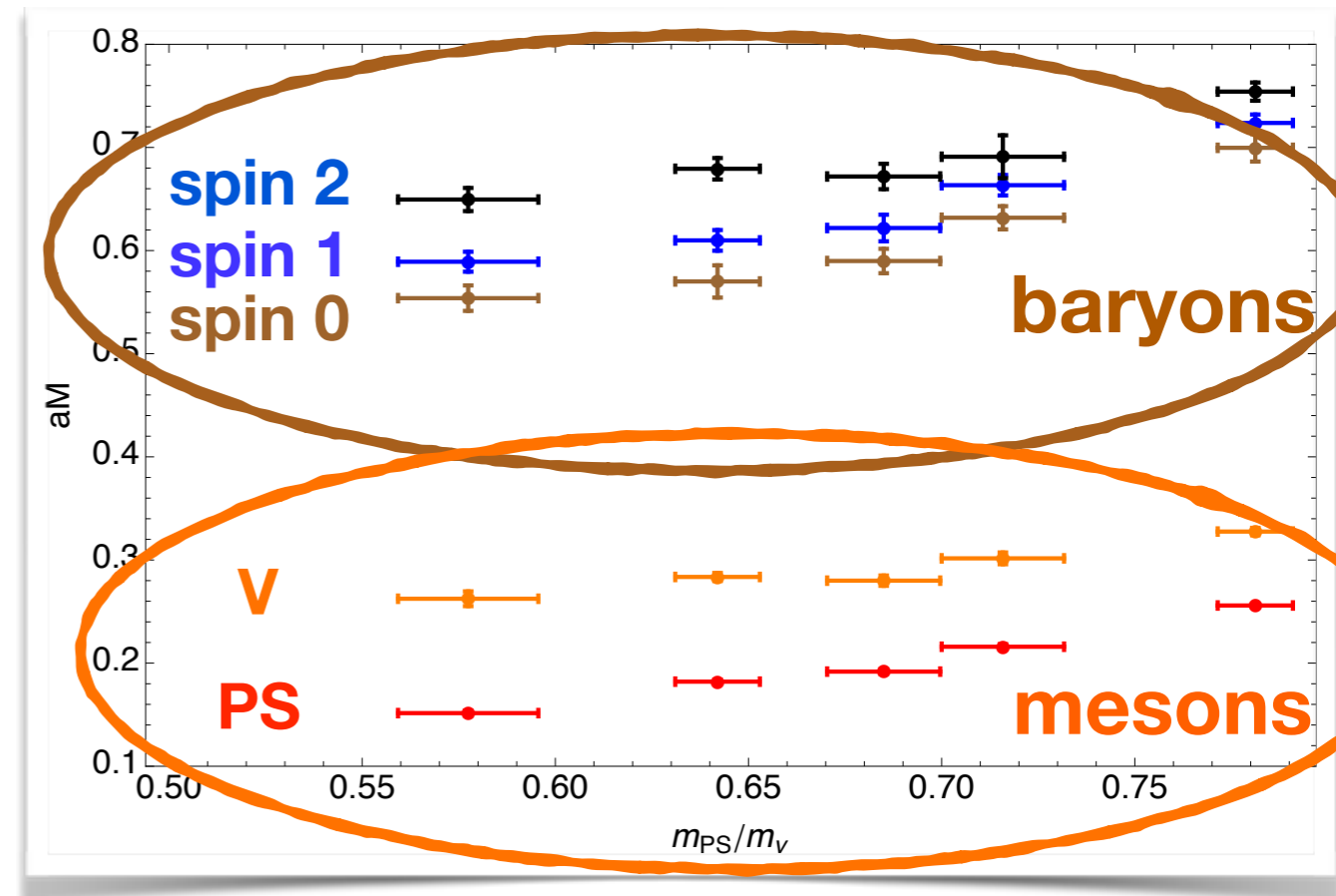
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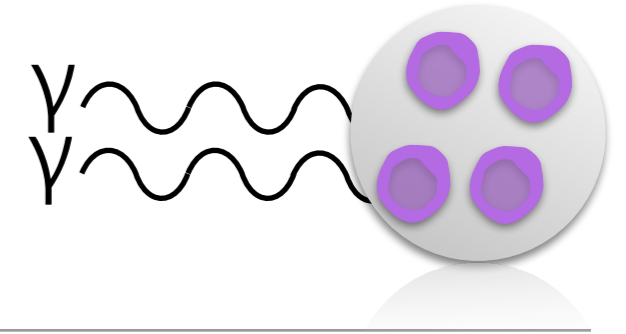
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Lattice: Polarizability of DM

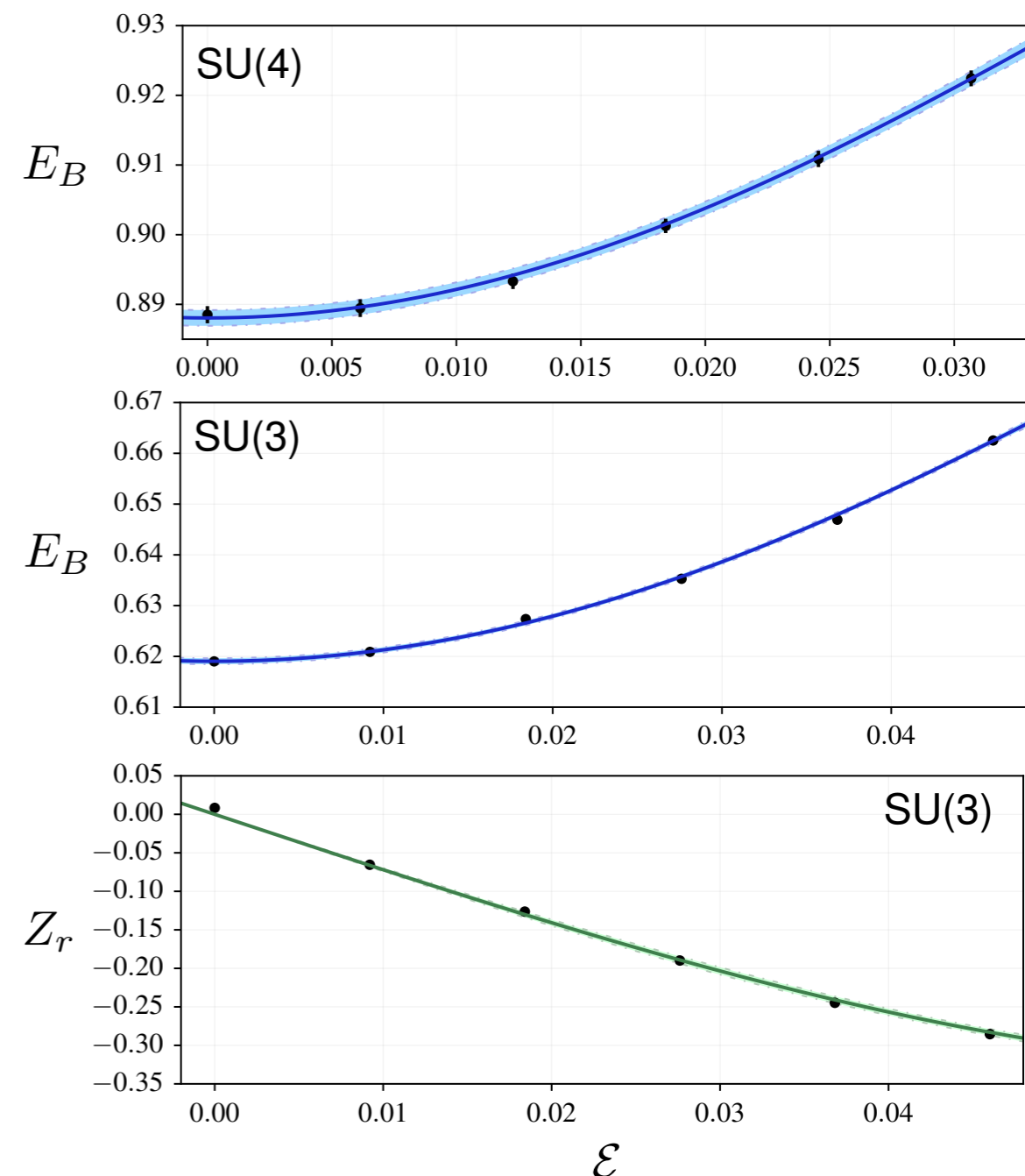
- **Background field method:**
response of neutral baryon to external electric field \mathcal{E}
- Measure the shift of the baryon mass as a function of \mathcal{E}

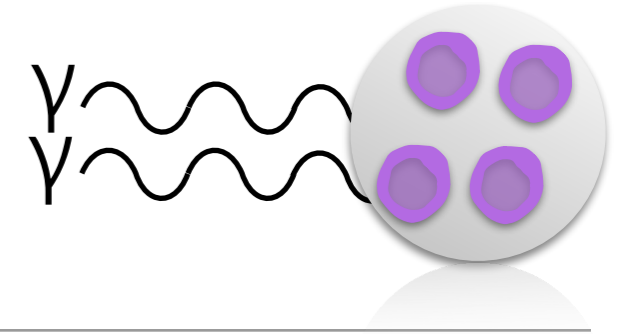
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$$Z_r = \frac{\mathcal{E} \mu_B(\mathcal{E})}{2m_B^2}$$

*32³x64 quenched lattices (large volume)
one lattice spacing and two masses (matched)
40 sources on 200 independent configurations
multi-exponential fits with 3 states for the baryon*





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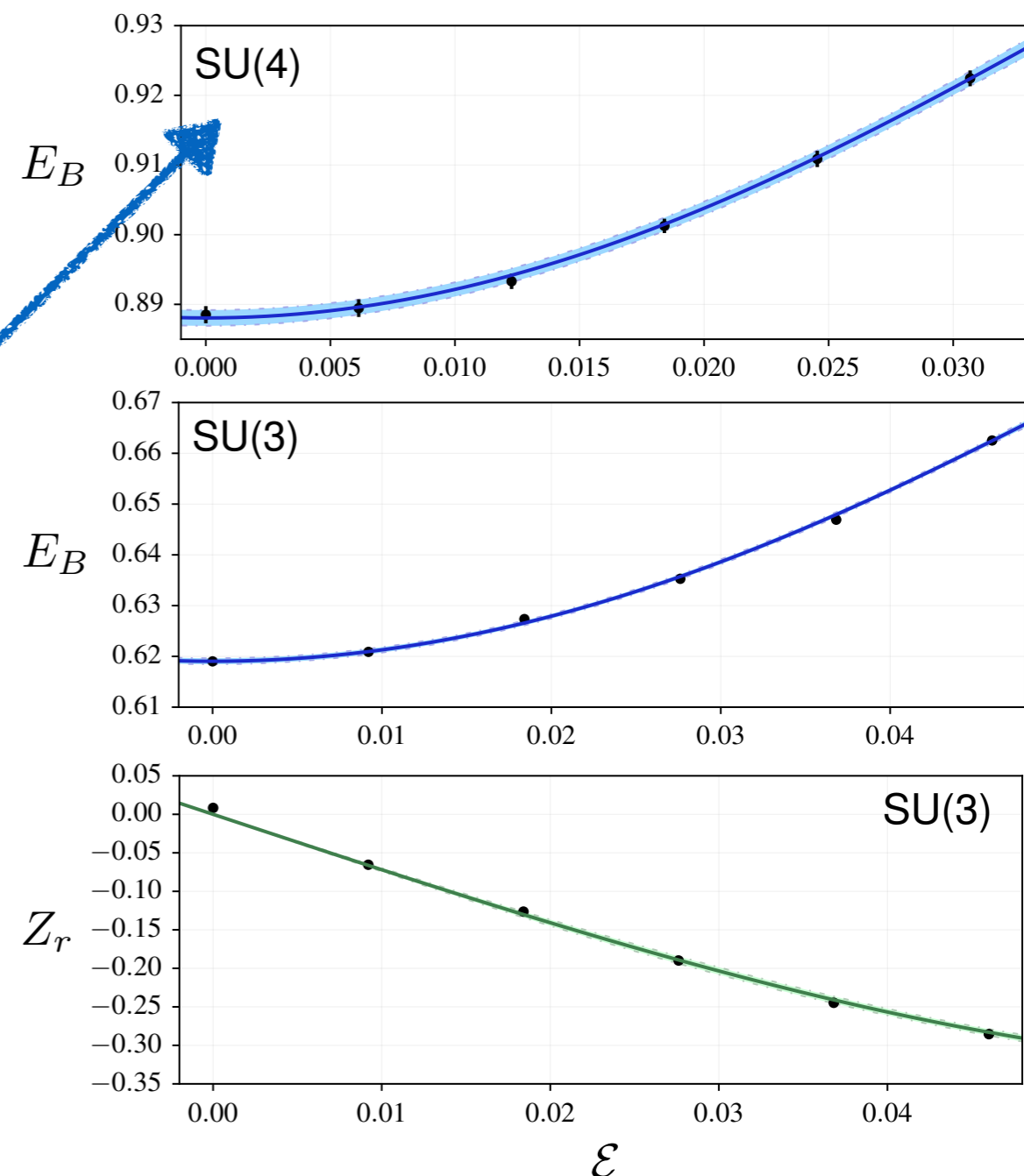
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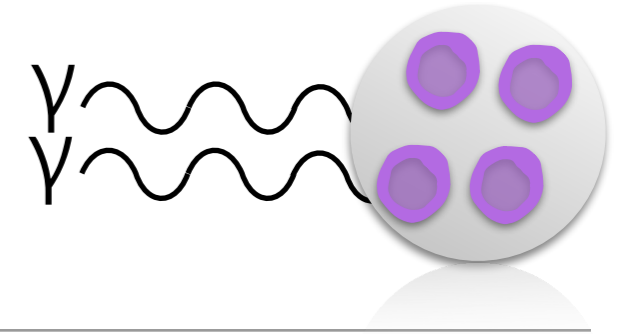
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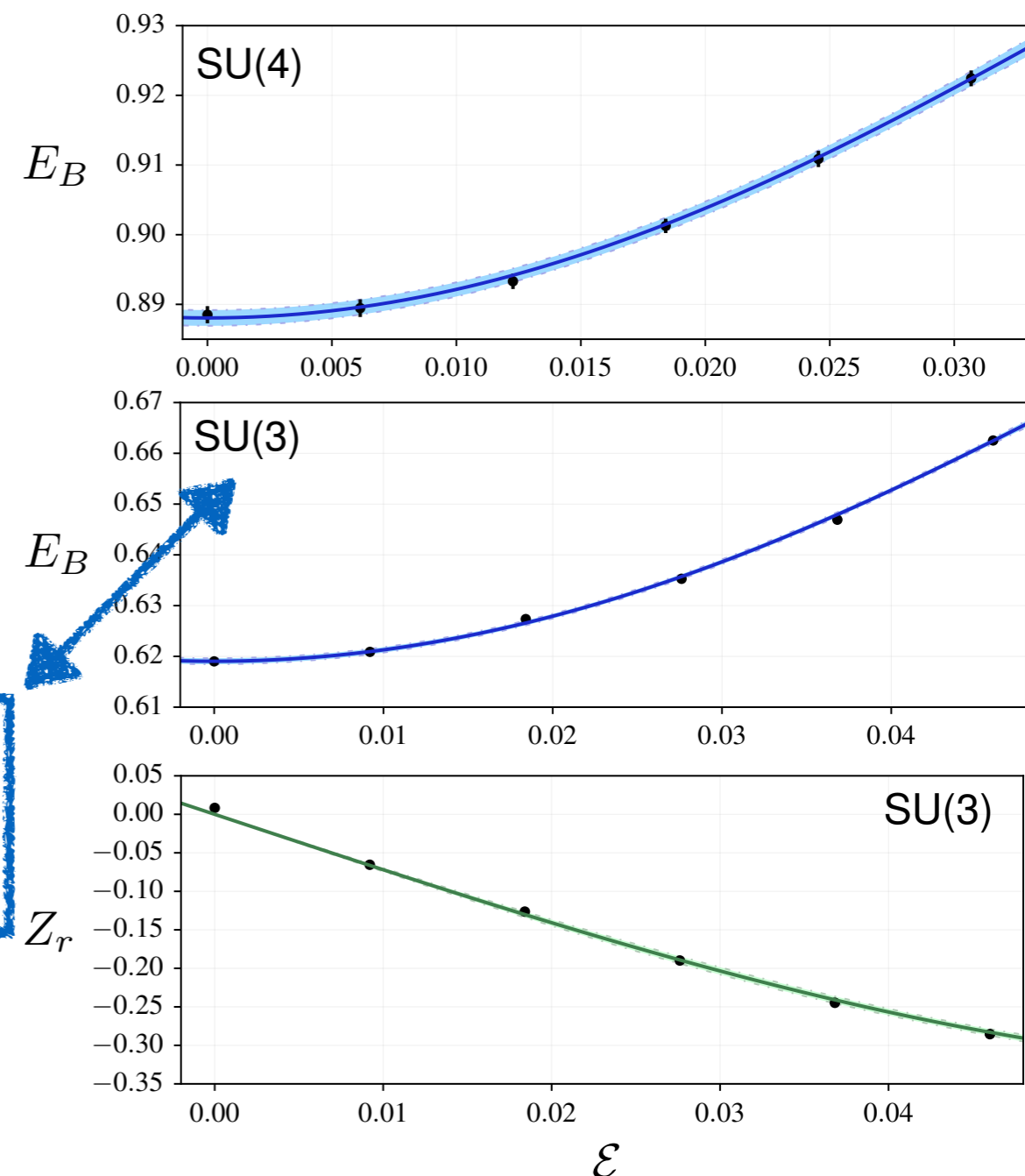
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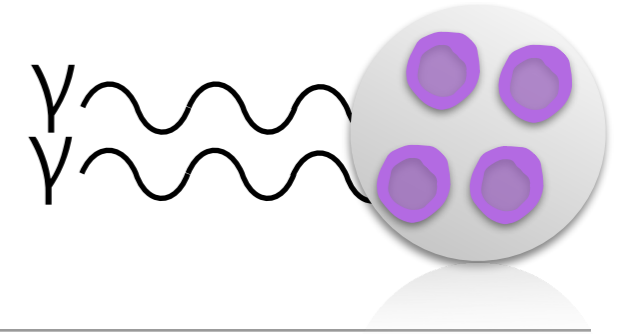
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