

FAIR high intensity beam dynamics

Space charge and impedance effects in SIS-18/100



Topics:

- Space charge limit and related beam loss: Optimized working point
- Optimization of RF manipulations with space charge and beam-loading
- Transverse resistive wall and kicker impedances
- Transverse coherent beam instabilities with space charge
- Transverse Schottky noise with space charge

People involved:

O. Chorniy, G. Franchetti, I. Hofmann, R. Hasse, V. Kornilov, S. Paret, A. Parfenova, O. Boine-Frankenheim
GSI, Darmstadt, FAIR accelerator theory

B. Doliwa, L. Hänichen, Th. Weiland, TU-Darmstadt, TEMF

A. Al-khateeb, W. Daqa, University Irbid, Jordan

G. Rumolo, H. Damerau, E. Metral, F. Zimmermann, CERN



Required FAIR beam intensities

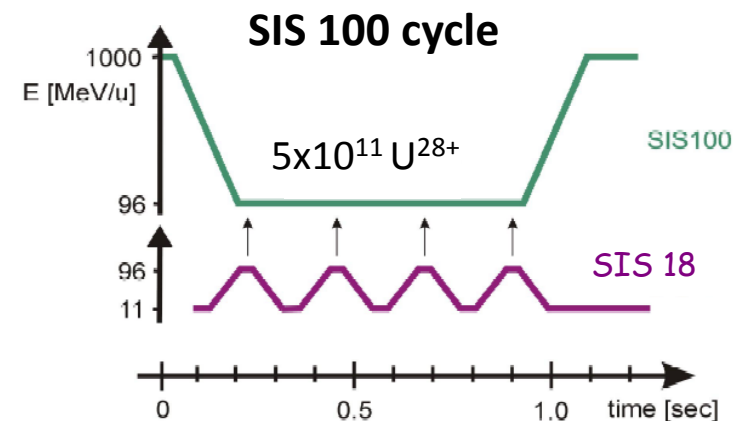
Heavy ions



	SIS-18 today*	After upgrade (2009)	After upgrade (to SIS 100)	SIS 100
Reference ion	U^{28+} / U^{73+}	U^{73+}	U^{28+}	U^{28+}
Maximum energy	0.2 / 1 GeV/u	1 GeV/u	0.2 GeV/u	2.7 GeV/u
Maximum intensity	$5 \times 10^9 / 3 \times 10^9$	2×10^{10}	1.5×10^{11}	5×10^{11}
Repetition rate	0.3 Hz	1 Hz	2.7 Hz	0.7 Hz
Bunch length	> 100 ns	50 ns	-	60 ns

FAIR specific beam dynamics challenges:

- Intensities at the 'space charge limit'
- High beam quality (weak or lost Landau damping)
- Long accumulation time (1 s) in SIS-100
- Bunch compression



SIS 100 dipole magnets

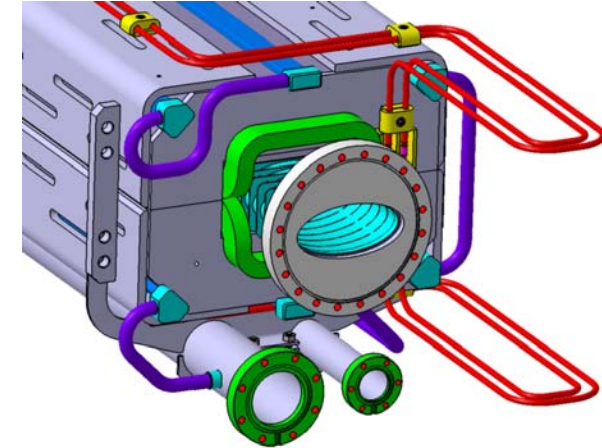
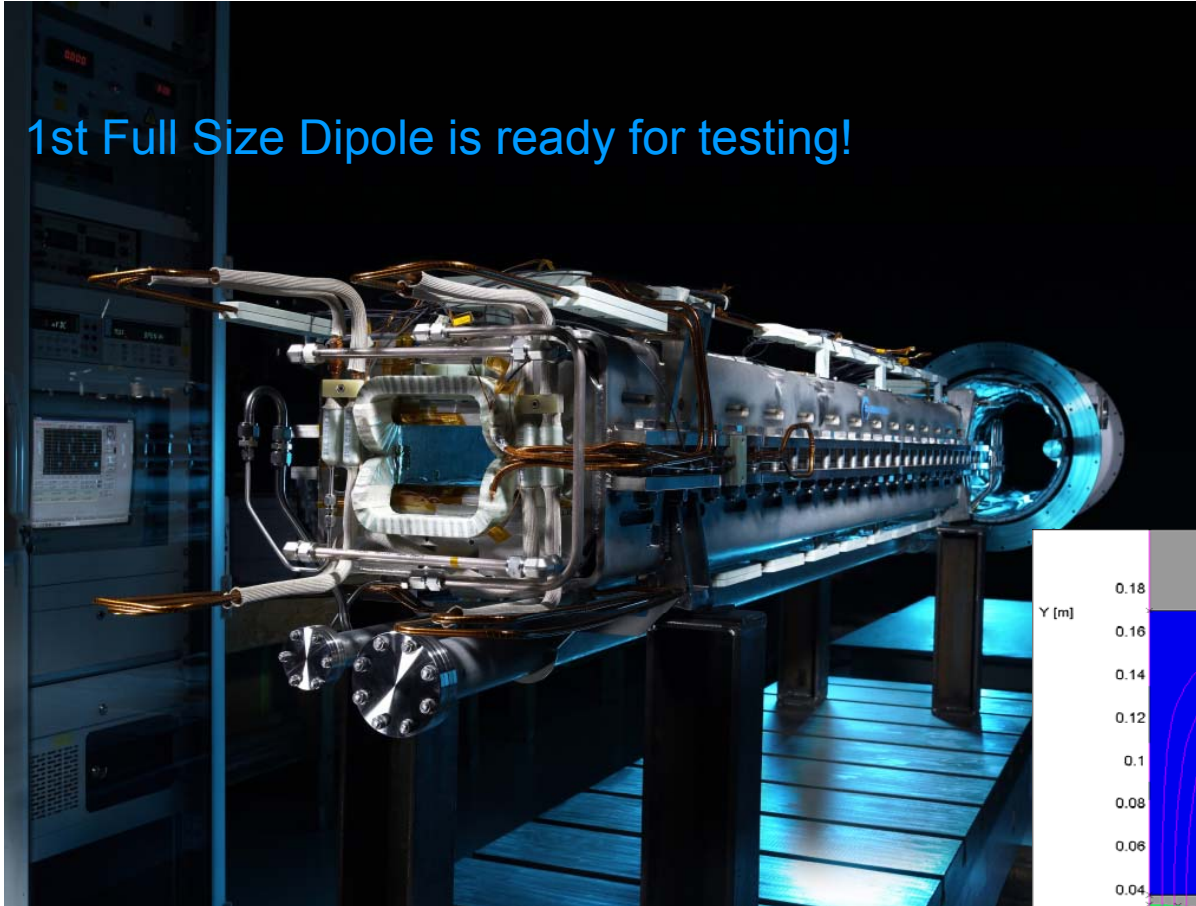
Field quality



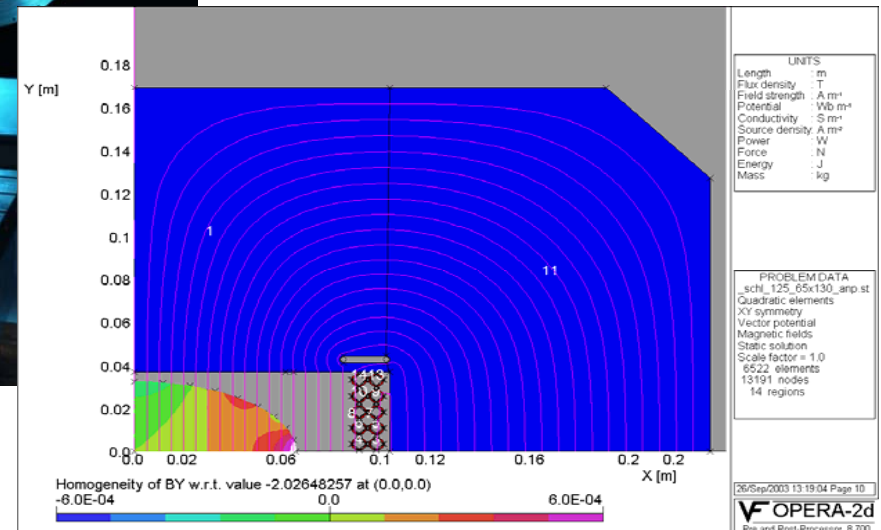
Superconducting magnet with 4 T/s and $B_{\max} = 1.9$ T

Dipole magnet with elliptical beam pipe

1st Full Size Dipole is ready for testing!



2D/3D field calculations



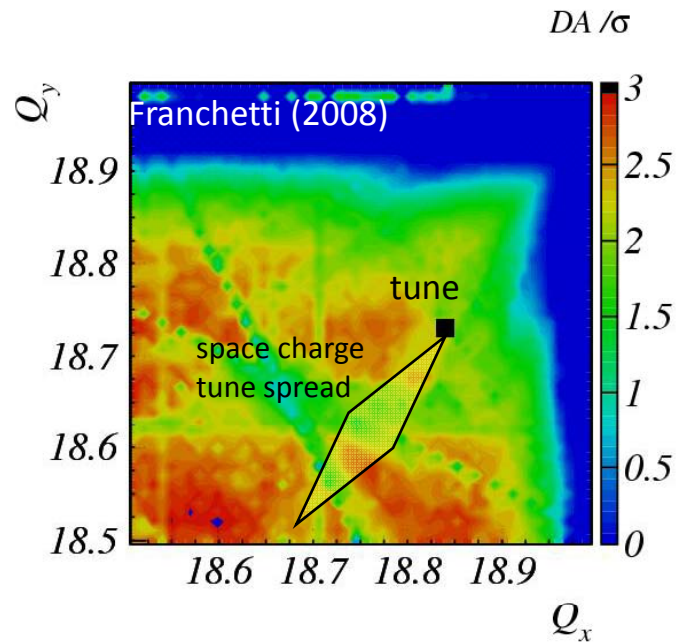
P. Spiller talk !



SIS-100 field quality and beam loss at injection energy 0.2 GeV/u



Simulation scan of the SIS-100 dynamic aperture

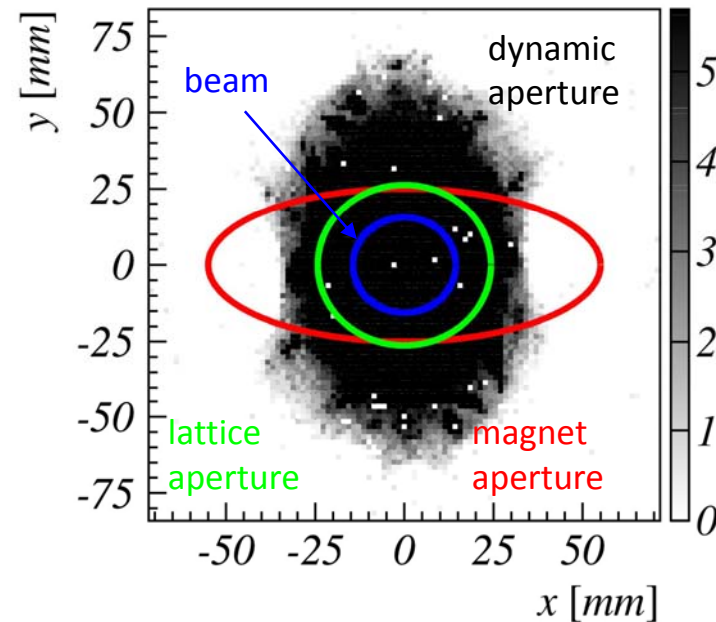


Long-term (up to 1 s) 3D particle tracking studies with 'frozen' space charge indicate a space charge limit at 3×10^{11} U^{28+} (design 5×10^{11}).

G. Franchetti et al. EPAC 2006

G. Franchetti, SIS-100 technical design report (2008)

SIS-100 transverse apertures at injection (0.2 GeV/u)



High intensity challenges for SIS-100:

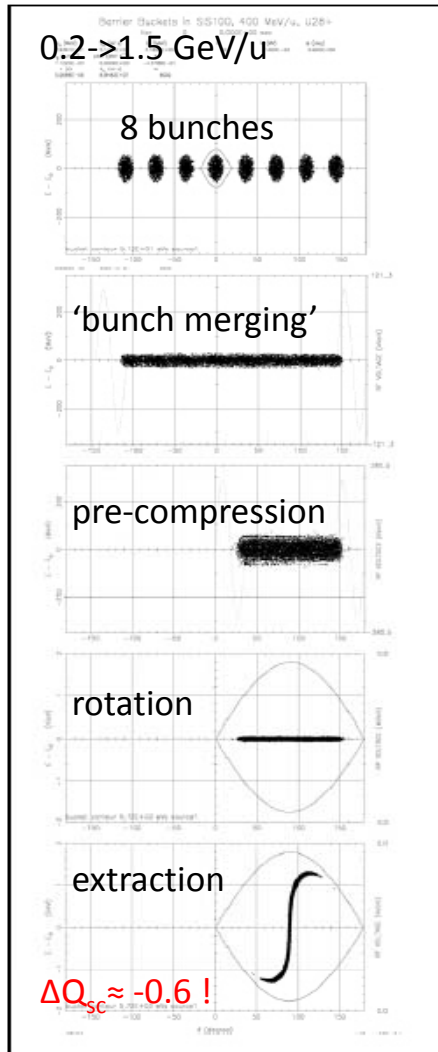
- Long time scales up to 10^6 turns (1 s)
- Space charge tune shift of $\Delta Q_{sc} = -0.25$
- 'Thick' medium energy beams (2/3 filling factor)
- Optimized working point for < 5 % loss.



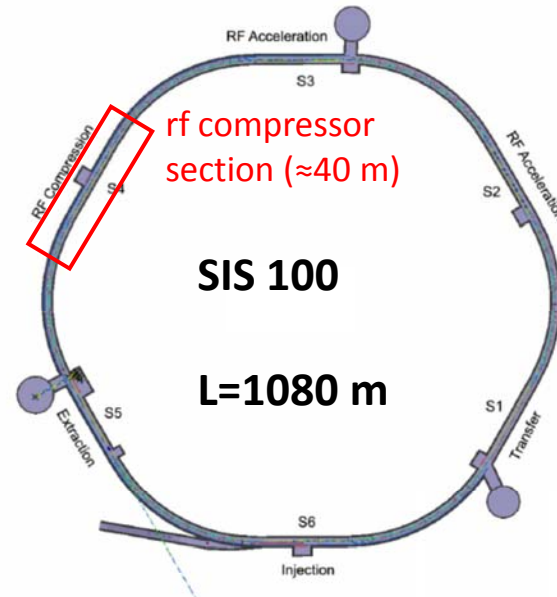
SIS-100 bunch compression



Single bunch formation



rf bunch compressor loaded with 20 MA cores



RF compression systems in SIS 100:

	#cavities	Voltage [kV]	Frequency [MHz]	Concept
Barrier rf	2	30	1.5 (h≈10)	MA
Pre-compresson				
Compression	16	600	0.4-0.5 (h=2)	MA (low duty)

Final bunch parameters:

	Particles/bunch	bunch length
1.5 GeV/u U ²⁸⁺	5x10 ¹¹	60 ns
29 GeV/u p	2-4x10 ¹³	50 ns



SIS-100 bunch pre-compression

barrier rf bucket with beam-loading and space charge



Longitudinal space charge impedance:

$$\frac{Z_l^{sc}}{n} = \frac{igZ_0}{2\beta_0\gamma_0^2} \approx i200 \Omega$$

	# cavities	R_s/cavity [k Ω]	Q_s	f_{res}
acceleration	20	3	10	$10f_0$
barrier	2	1	0.4	1.5 MHz
compression	16	1	2	1.5 MHz

Haissinski equation for the matched bunch form:

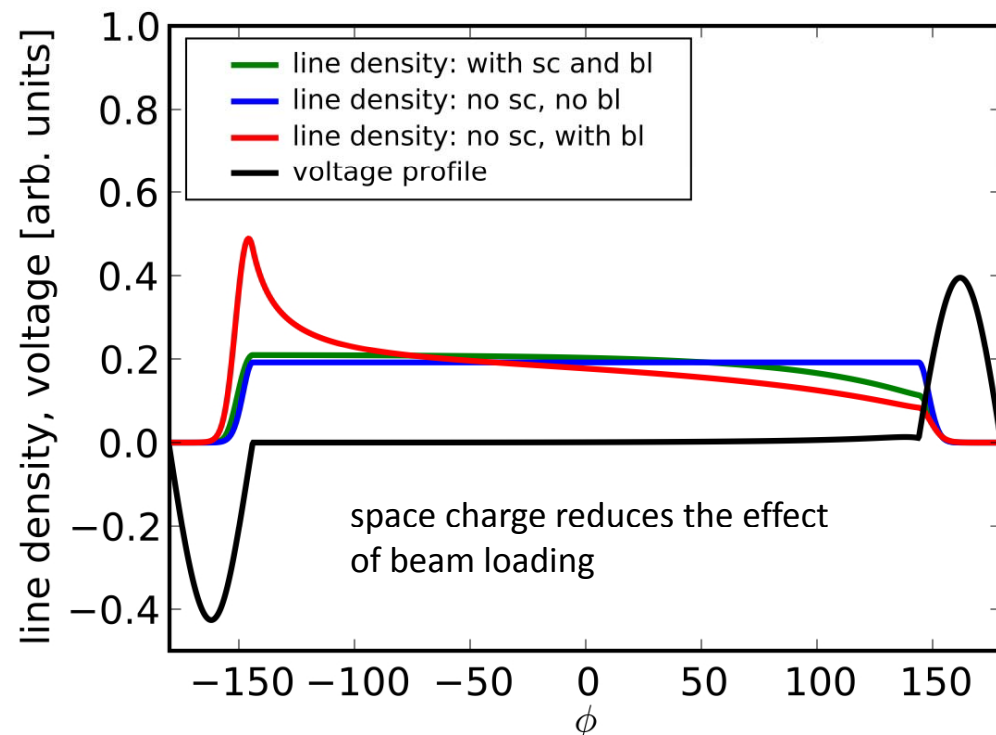
$$\lambda(z) = \lambda_m \exp\left(\frac{qV_0 Y(z)}{m^* L v_{rms}^2}\right)$$

Between barriers (no rf):

$$\lambda(z) = \lambda_m \exp\left(\alpha_s \lambda + \alpha_r \int_0^z \lambda dz\right)$$

with $\alpha_s \propto \frac{Z_l^{sc}}{n}$ $\alpha_r \propto R_s$

$$\lambda(z) = \frac{\lambda_m}{1 - \alpha_s(\lambda - \lambda_m) - \alpha_r \lambda_m z}$$



Barrier Bucket Pre-Compression in SIS-100



Non-adiabatic barrier rf phase ramp:

$$T=100 \text{ ms} \quad T_{s0} > 100 \text{ ms}$$

(ramp time) ('synchrotron period')

Space charge factor:

$$\Sigma = \frac{c_s}{v_{rms}} > 1$$

Cold fluid equation (Rumolo et al. 1999):

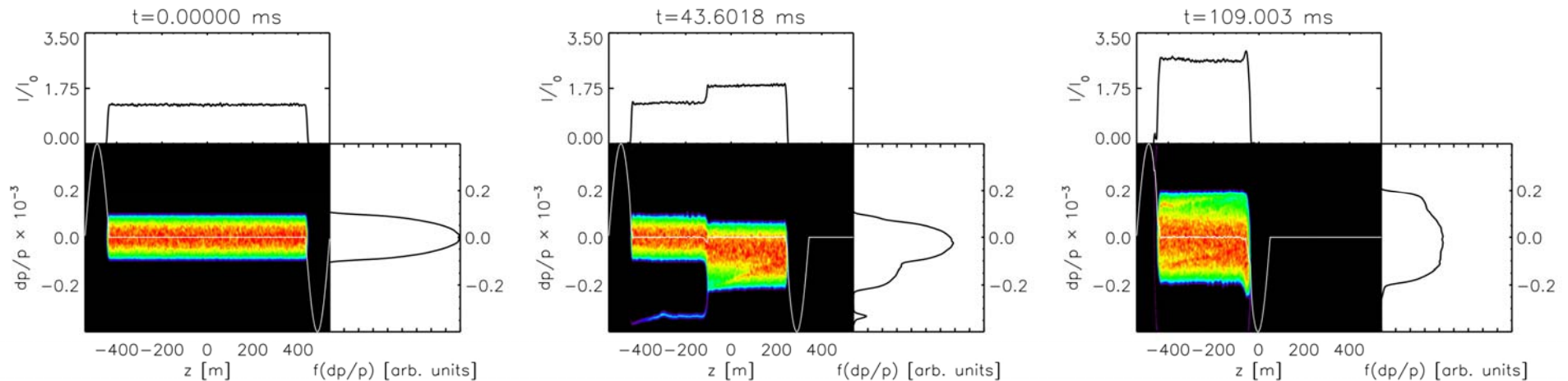
$$\frac{\partial \lambda}{\partial t} + \frac{\partial}{\partial z}(\lambda u) = 0 \quad \frac{\partial u}{\partial t} + u \frac{\partial u}{\partial z} + \chi \frac{\partial \lambda}{\partial z} = 0$$

with $\chi = \frac{qX_s \beta_0 c}{2\pi m^*} \quad \frac{Z_{sc}}{n} = -iX_{sc}$

'speed of sound': $c_s \approx \sqrt{\lambda_m \chi}$

Shock compression:

The barrier acts like a piston moving with velocity v in a gas, creating a shock wave of speed $c_s > v$



The shock compression is dilution-free !



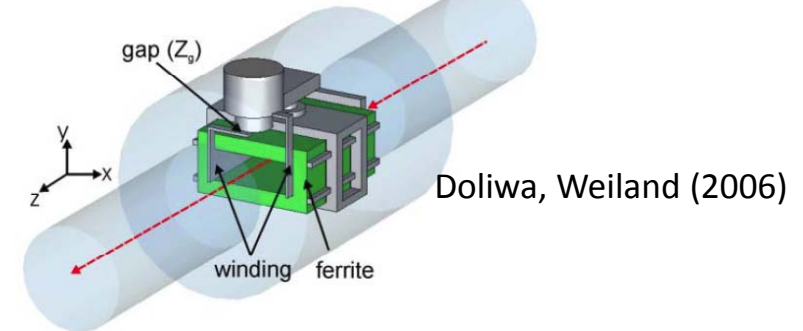
Transverse impedances in SIS-18/100



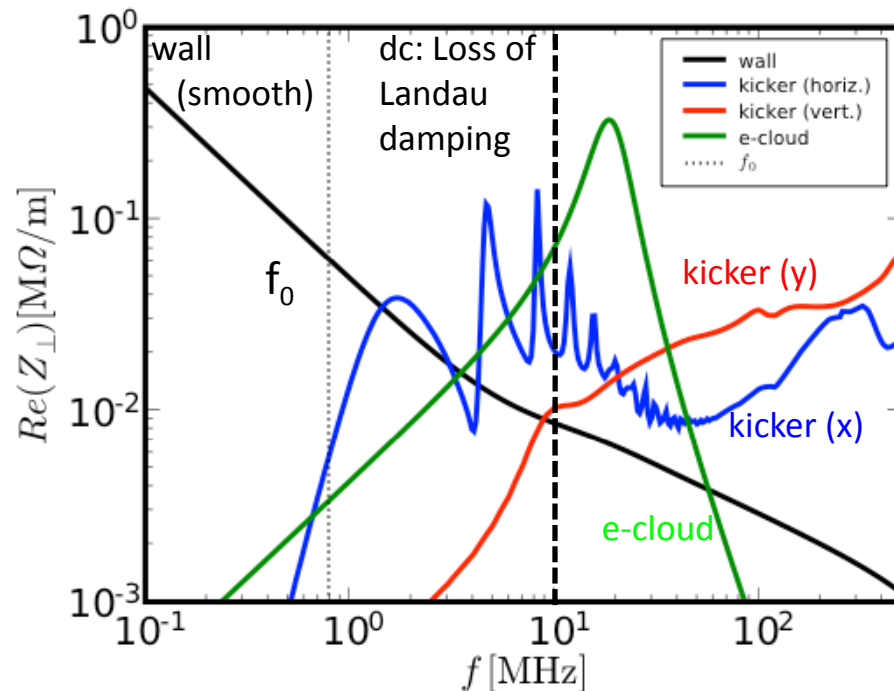
Specific SIS-18/100 impedance issues:

- Low frequencies and beam energies
- Thin (0.3 mm) stainless steel pipe (**optional** corrugated)
- Ferrite or magnetic alloy loaded ring components
- Distributed collimation system

Ferrite loaded kicker modules



SIS-18 transverse impedance spectrum (200 MeV/u)



‘Space charge impedance’: $Z_{\perp}^{sc} = -i \frac{Z_0 R}{\beta_0^2 \gamma_0^2} \frac{1}{b^2} \approx -i 10 \text{ M}\Omega/\text{m}$

e-clouds ?:

- electrons from residual gas ionization and beam loss.
- e-accumulation: survival rate in bunch gaps ?
- transv. ‘impedance’: Ohmi, Zimmermann, PRL (2000)
- long. impedance: Al-khateeb et al. NJP (2008)

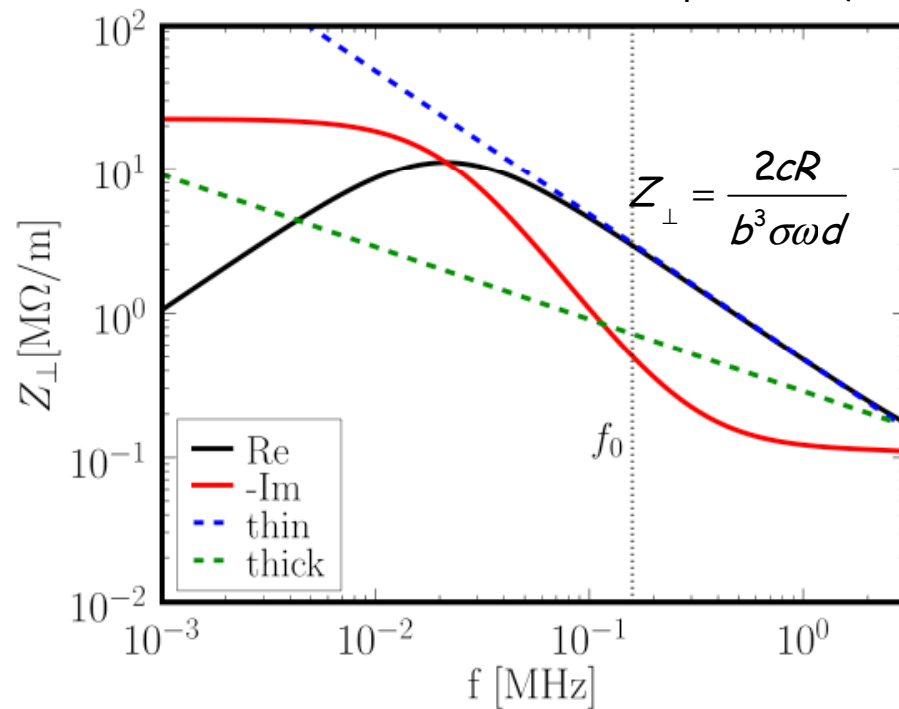
Resistive (thin) beam pipe in SIS-100

Transverse impedance and shielding

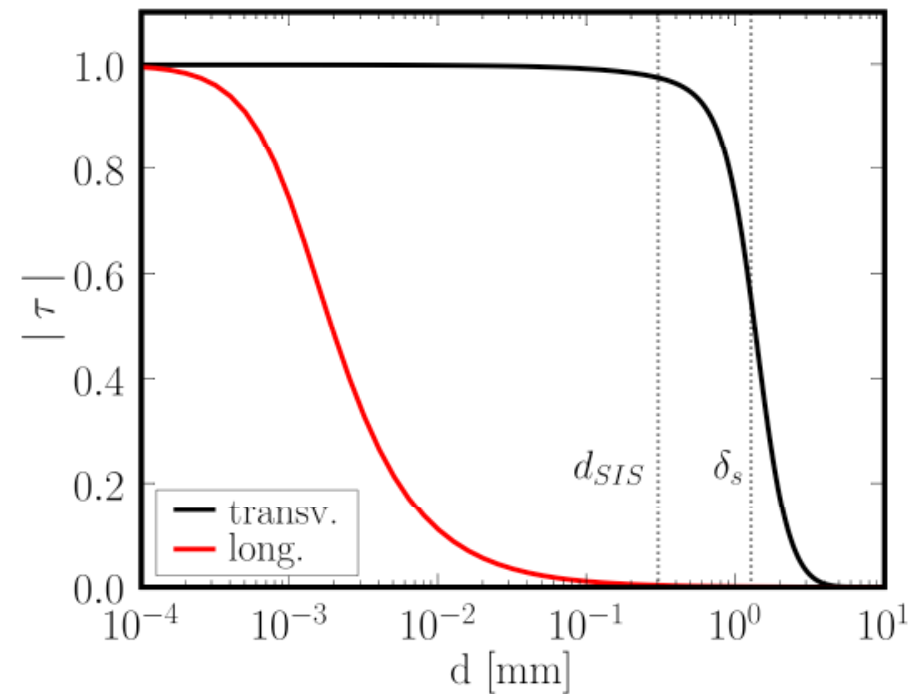


In SIS-100 the thin (0.3 mm) beam pipe is the dominant transverse impedance contribution !

SIS 100 transverse resistive wall impedance (200 MeV/u)



Transmission coefficient !



Transmission: Structures behind the pipe might contribute !

SIS-18: Measured growth rate of the resistive wall instability agrees with impedance formula.

Al-khateeb, Hasse, Boine-F., Daqa, Hofmann, Phys. Rev. ST-AB (2007)

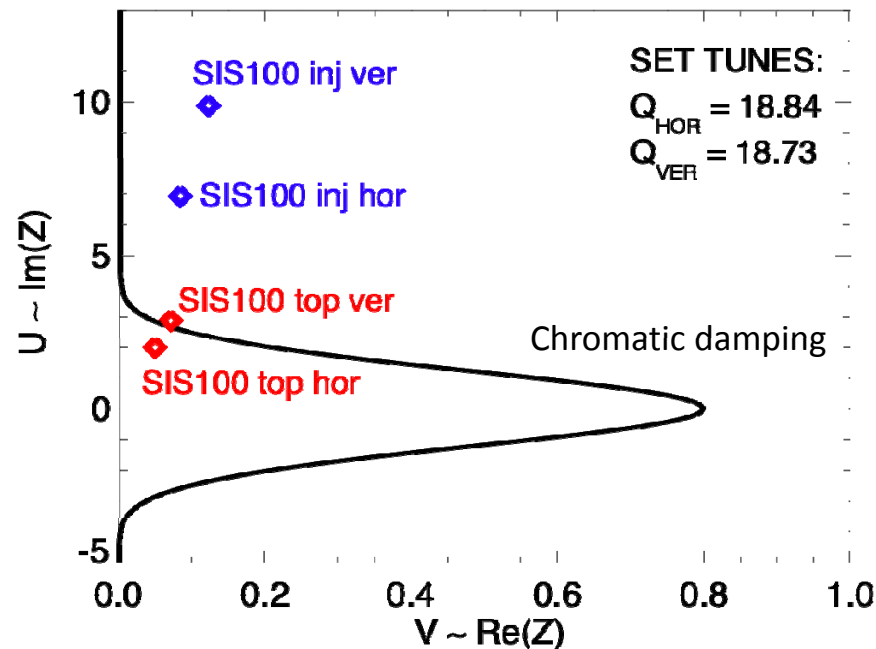


Transverse Impedance Budget with Space Charge



coasting beam stability boundaries (from dispersion relation)

$$Z_{\perp} = Z_{\perp}^{rw} + Z_{\perp}^{sc} \quad Z_{\perp}^{sc} = -i \frac{Z_0 R}{\beta_0^2 \gamma_0^2} \left(\frac{1}{a^2} - \frac{1}{b^2} \right)$$



Specific issues for SIS-100:

- Long time scales (up to 1 s)
- Image current and space charge effects
- Long bunches and 'dc-like' barrier rf beams

Potential cures:

- octupoles and nonlinear space charge
(Kornilov et al. PRST-AB 2008)
- feedback systems

Challenges:

- Octupoles and dynamic aperture
- 3D space charge and image current effects

Head-tail instability in SIS-100 (talk by V. Kornilov this Monday afternoon):

- Sacherer's formula predicts head-tail modes with 70 ms growth rate in SIS-100 !
- 3D simulation studies including: resistive wall impedance, image currents, space charge, nonlinear rf.
- 'Mode coupling regime': $\Delta Q_{incoh}^{sc} > \Delta Q_{coh}^{sc} > \nu_s$ (ν_s : synchrotron tune)



TWO SIMULATION TOOLS



V. Kornilov

**our strategy on the way to a correct description
of the effects relevant in SIS-100 for the Head-Tail instability:
two different tools**

PATRIC

(development at GSI)

short-term simulation:
resolve betatron oscillations
impedances $Z(\omega)$ [add. $W(s)$]
self-consistent 3D space charge
coasting / long bunches

Coasting beam modes: comparison with analytic
dispersion relation
Schottky noise: comparison with analytic theory.

HEADTAIL

(joint work with G.Rumolo)

long-term simulations:
once per turn
wake fields $W(s)$
beam-beam kick
short bunches

Head-Tail modes: consistency with the Sacherer theory
was confirmed in the recent work
(E.Metral,G.Rumolo,R.Steerenberg,B.Salvant,PAC'07)

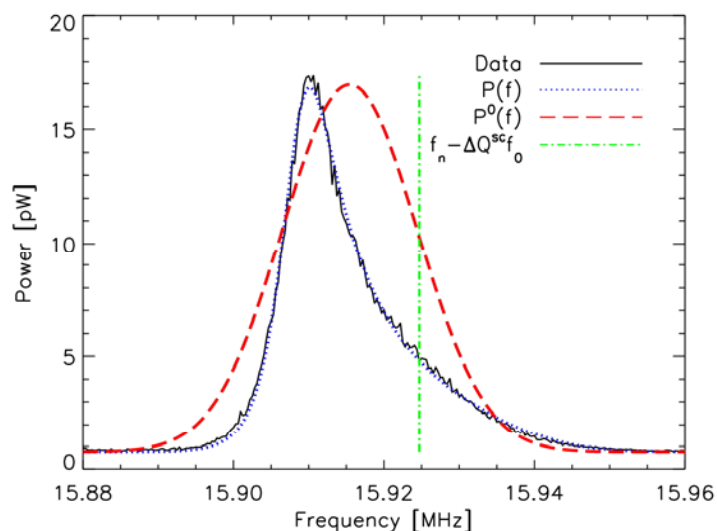


Schottky noise measurement in SIS-18

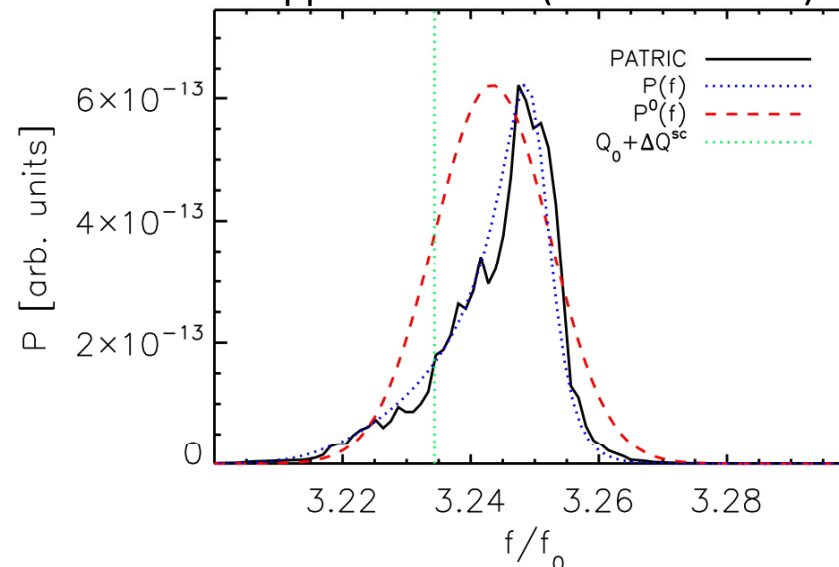
coasting beam with space charge



Measured lower sideband (n=75)



Upper sideband (PIC Simulation)



Transverse Schottky signal:

$$d(t) = \sum_{j=0}^N x_j(t) I_j(t)$$

coasting beam coherent betatron lines:

$$\Omega_n = (n \pm Q_0) f_0$$

Low intensity Schottky band: $P^0(\Omega)$

Modified Schottky band (analytic approach):

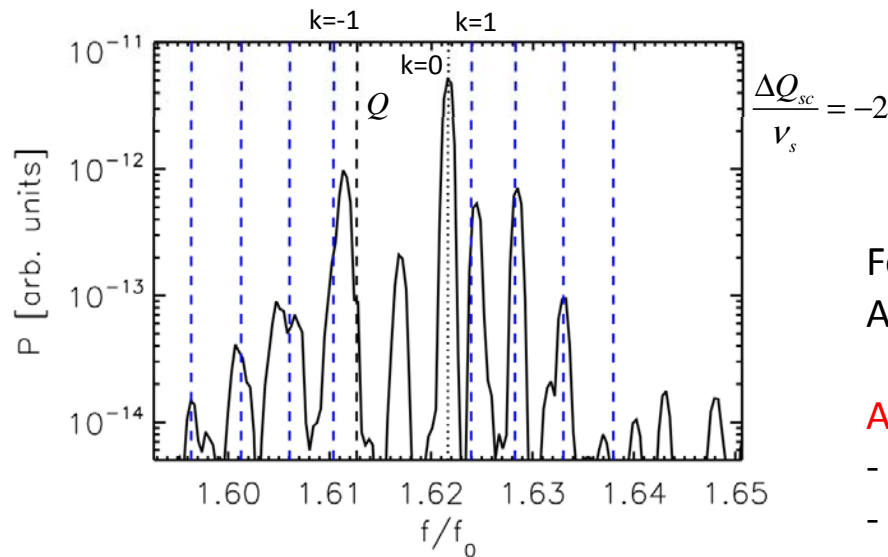
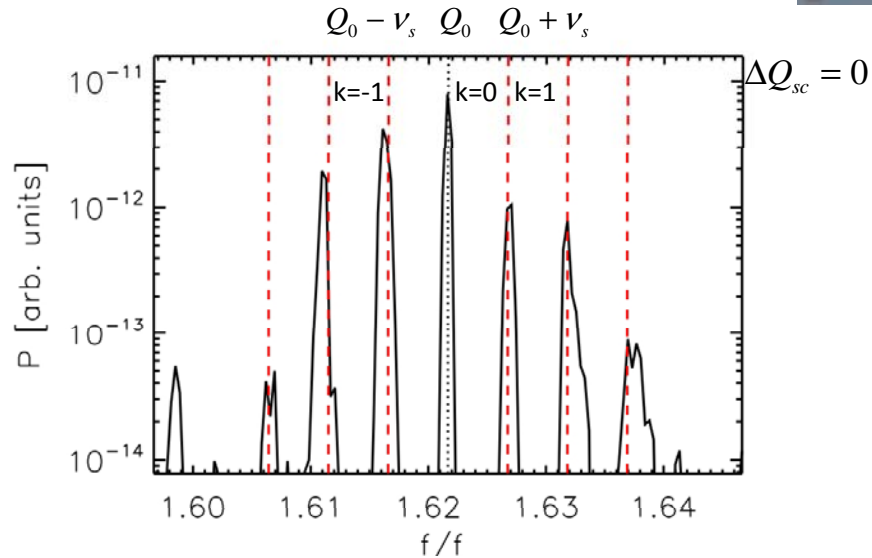
- Assuming linear space charge (KV-dist.)
- Dispersion relation $D(z)$ with space charge:

$$\varepsilon(z) = 1 - D(z) \quad P(z) = \frac{P_0(z)}{|\varepsilon(z)|^2}$$

Comparing the measured Schottky bands with $P(z)$ we can obtain the space charge parameter.

Boine-Frankenheim, Kornilov, Paret, PRST-AB (2008)





Shift of head-tail modes with space charge

M. Blaskiewicz, Phys. Rev. ST Accel. Beams 1, 044201 (1998)

for a barrier air-bag distribution.

$$\text{For } k \neq 0: \quad \Delta Q_k = -\frac{\Delta Q_{sc}}{2} \pm \sqrt{(\Delta Q_{sc} / 2)^2 + (k v_s)^2}$$

$$\Delta Q_{sc} = v_s \quad \Delta Q_k = -\frac{\Delta Q_{sc}}{2} \pm k v_s$$

$$\Delta Q_{sc} \neq v_s \quad \Delta Q_k = \frac{(k v_s)^2}{\Delta Q_{sc}} \approx 0, \text{ for } k > 1$$

$$\Delta Q_k = -\Delta Q_{sc}, \text{ for } k < 1$$

For barrier air-bag: [code benchmarking](#)

Analytic results apply for **realistic bunches** (e.g. elliptic).

Application:

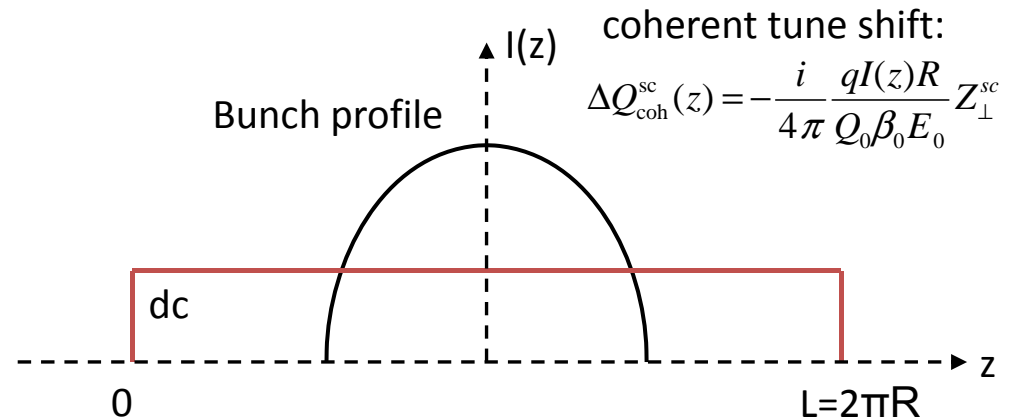
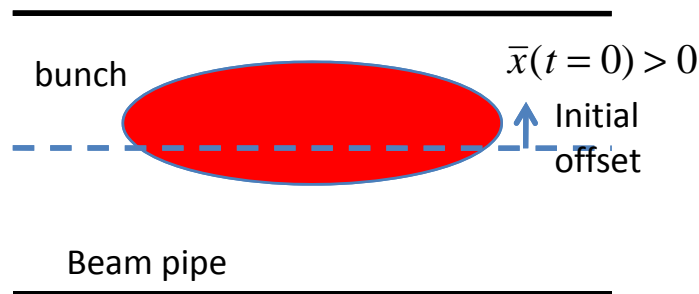
- Measure space charge tune shifts in bunches
- Understand thresholds for head-tail instabilities

Decoherence with image currents and space charge

3D simulation studies with space charge and impedances



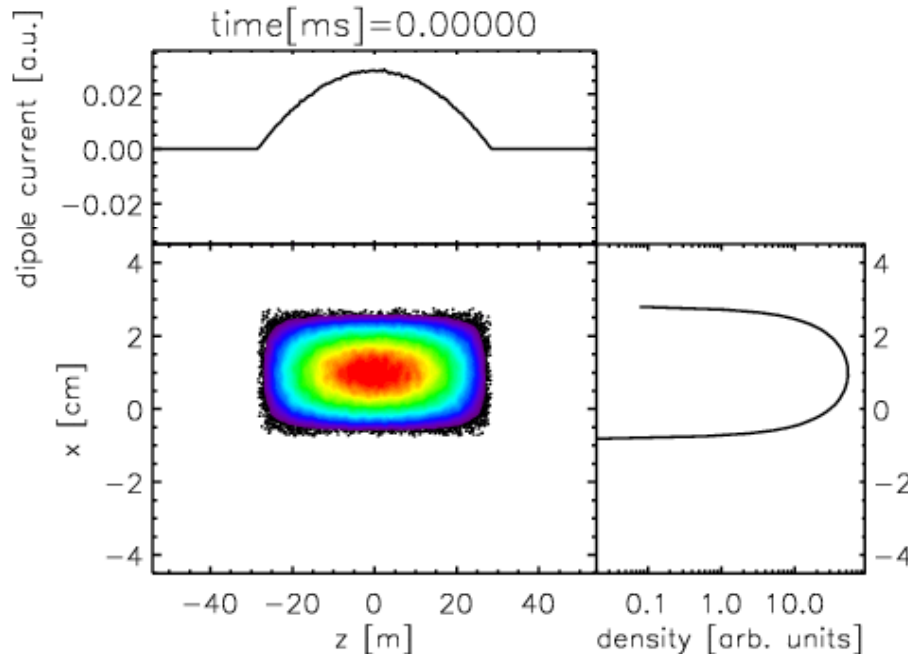
Injected bunch in SIS-100 with a transverse offset.



$$\Delta Q_{\text{coh}}^{\text{sc}}(z) = -\frac{i}{4\pi} \frac{qI(z)R}{Q_0\beta_0 E_0} Z_{\perp}^{\text{sc}}$$

Coherent and incoherent tune spread along the bunch:

$$\Delta Q_{\text{coh, incoh}}(z) = \Delta Q_{\text{coh, incoh}}^{\text{max}} \left(1 - \frac{z^2}{z_m^2} \right)$$



Interplay of space charge and image currents:

- mode coupling (due to image currents)
- persistent bunch modes
- halo formation !



Work performed in the framework of the CARE-HHH network activity.

- Long-term particle tracking studies for SIS-100 including ‘frozen’ space charge for different working points: Beam loss estimates.
- Analysis of space charge and beam-loading effects in barrier rf buckets: Optimized pre-compression scheme with space charge.
- Transverse impedance spectrum in SIS-18/100 from analytic and from simulation models.
- Transverse impedance budget for dc-like beam and for bunches including space charge.
- Experimental and simulation studies of transverse Schottky noise and decoherence with space charge.