

# Initial energy density of p+p and A+A collisions

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## Hydrodynamic Solutions

Exact, explicit solutions of accelerating relativistic hydrodynamics allow for a simple and natural description of highly relativistic p+p collisions and heavy ion collisions. An exact analytic solution for velocity field and pressure in Rindler coordinate  $(\tau, \eta_R) = (\sqrt{t^2 - r^2}, \frac{1}{2} \ln \frac{t+r}{t-r})$  reads

$$v = \tanh \lambda \eta_R, \quad (1)$$

$$p = p_0 \left( \frac{\tau_0}{\tau} \right)^{\lambda d \frac{\kappa+1}{\kappa}} \left( \cosh \left( \frac{\eta_R}{2} \right) \right)^{-(d-1)\phi_\lambda}, \quad (2)$$

where  $\lambda$  is the acceleration parameter.

Table I. Solution parameters

case	$\lambda$	$d$	$\kappa$	$\phi_\lambda$
(a)	2	$\in \mathbb{R}$	$d$	0
(b)	$\frac{1}{2}$	$\in \mathbb{R}$	1	$\frac{k+1}{k}$
(c)	$\frac{2}{3}$	$\in \mathbb{R}$	$\frac{4d-1}{3}$	$\frac{k+1}{k}$
(d)	1	$\in \mathbb{R}$	$\in \mathbb{R}$	0
(e)	$\in \mathbb{R}$	1	1	0

## Pseudo-rapidity Distribution

A finite pseudo-rapidity distribution  $dN/d\eta$  is yielded by these hydrodynamic solutions, which can be used to extract the QGP acceleration parameter  $\lambda$ .

$$\frac{dN}{d\eta} \simeq N_0 \frac{\bar{p}_T \cosh \eta}{\sqrt{m^2 + \bar{p}_T^2 \cosh^2 \eta}} \times \cosh^{-\frac{\alpha}{2}-1} \left( \frac{y}{\alpha} \right) \exp \left\{ -\frac{m}{T_f} \cosh^\alpha \left( \frac{y}{\alpha} \right) \right\}, \quad (3)$$

with  $\alpha = \frac{2\lambda-1}{\lambda-1}$ ,  $T_{eff} = T_f + \frac{m(\mu_t)^2}{1+m/T_f}$ ,  $\bar{p}_T = \frac{T_{eff}}{1+\sigma^2 y^2}$ , and  $N_0$  is a normalization parameter.

## Improved $\epsilon_{ini}$ Estimation

Bjorken's energy density estimation

$$\epsilon_{Bj} = \frac{\langle E \rangle dN}{(R^2 \pi) \tau_0 d\eta_0} = \frac{\langle E \rangle}{(R^2 \pi) \tau_0} \frac{dN}{d\eta} \Big|_{\eta=\eta_0}. \quad (4)$$

When the acceleration effect is taken into account, the corrected initial energy density estimation is:

$$\epsilon_{corr} = \epsilon_{Bj} (2\lambda - 1) \left( \frac{\tau_f}{\tau_0} \right)^{\lambda-1}. \quad (5)$$

## References

- [1] T. Csörgő, M. I. Nagy, and M. Csanád, *A New family of simple solutions of perfect fluid hydrodynamics*, Phys. Lett. B **663**, 306(2008), arXiv: 0605070.
- [2] M. Csanád, T. Csörgő, ZeFang. Jiang and ChunBin. Yang, *Initial energy density of  $\sqrt{s} = 7$  and 8 TeV p+p collisions at the LHC*, in: Universe, Special Issue On Quark-Gluon Plasma in the Early Universe and in Relativistic Heavy-Ion Collisions, arXiv:1609.07176.

## Acknowledgements

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## Pseudo-rapidity distribution at RHIC and at LHC

Acceleration effect is extremely sensitive to quantify QGP properties and to investigate the QGP initial state. One can extract **acceleration parameter characteristics** from precise  $dN/d\eta$  for different centralities (Eq.1-3 and Table I case.e) and obtained a series of **improved initial energy density estimation** (for RHIC and LHC).

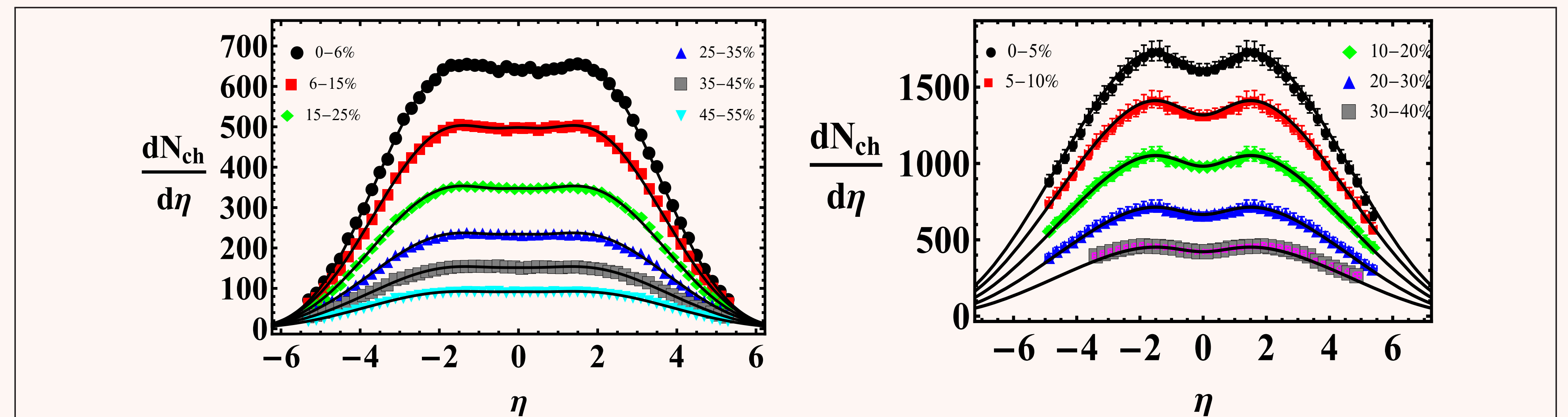


Fig. 1: Left: 200 GeV Au+Au Collision. Right: 2.76 TeV Pb+Pb collision. Charged particle pseudorapidity distribution: Relativistic hydrodynamic solution (solid line)+The data from RHIC and LHC (color dots).

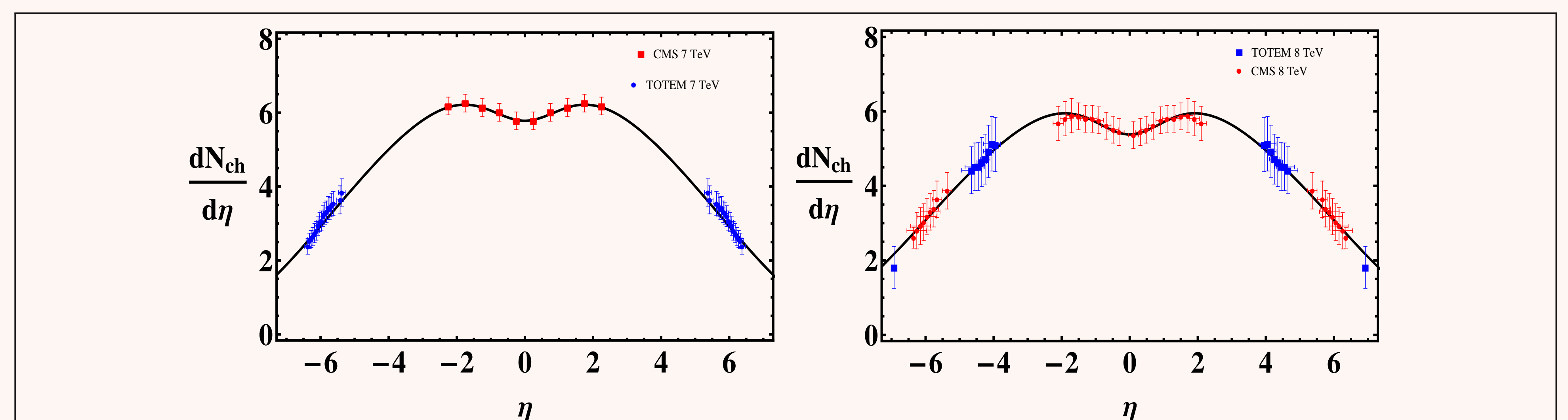


Fig. 2: 7 and 8 TeV p+p collision. Extracted  $\lambda$  with  $dN/d\eta$ : Relativistic hydrodynamic solution (solid line)+The data from LHC (color dots).

## Improved initial estimation combined RHIC and LHC

In our previous work [1], we given an advanced estimation of the initial energy density in high energy collisions. Here we focus on the more **systematic domain of initial estimation** at RHIC and LHC, which takes acceleration effects (i.e. the work done by the pressure, and the modified change of the volume elements) into account (Eq. 4-5).

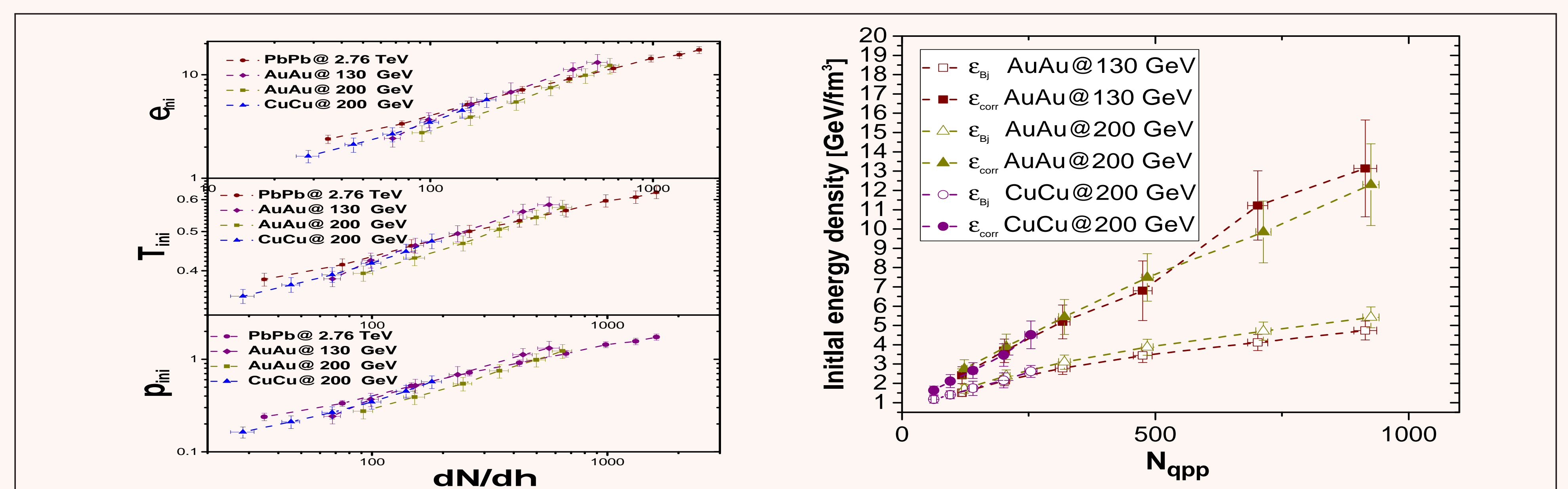


Fig. 3: Left: Initial energy density, temperature and pressure as a function of multiplicity density. Right: Initial energy as a function of quark participant pairs  $N_{qpp}$ .

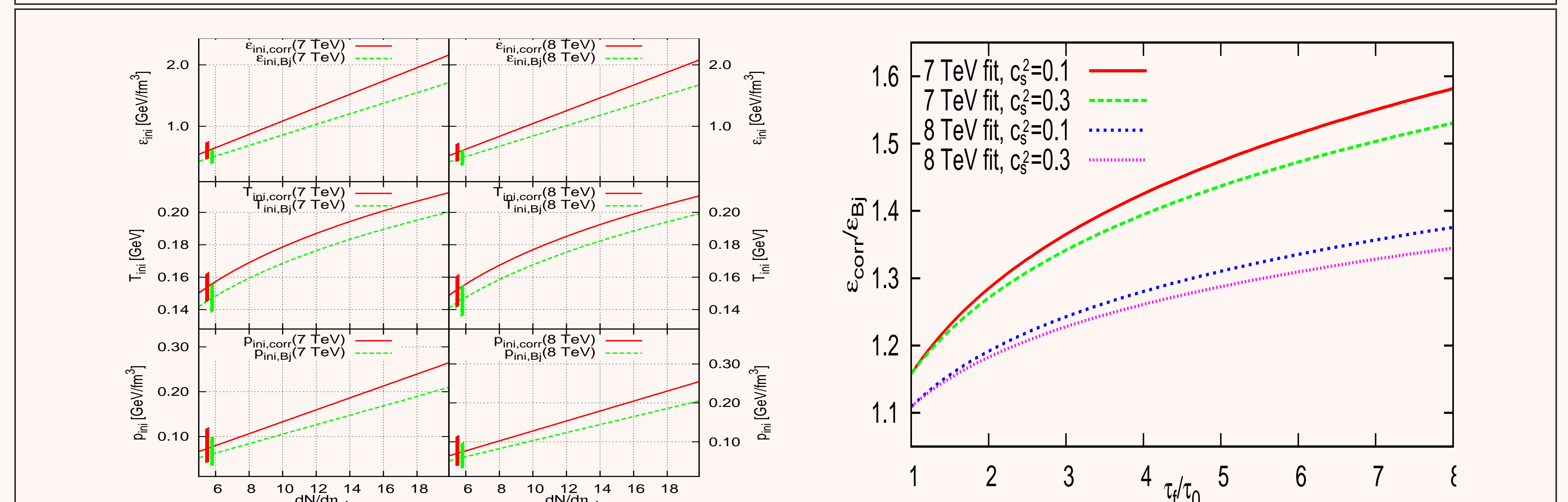


Fig. 4: Left: 7 and 8 TeV p+p collision large pseudo-rapidity distribution and initial energy density correction factor as a function of thermalization time [2]. Right: Initial energy density, temperature and pressure as a function of central multiplicity density at 7(left)and 8(right) TeV.