

# Phase Transitions in Dense Matter

Veronica Dexheimer

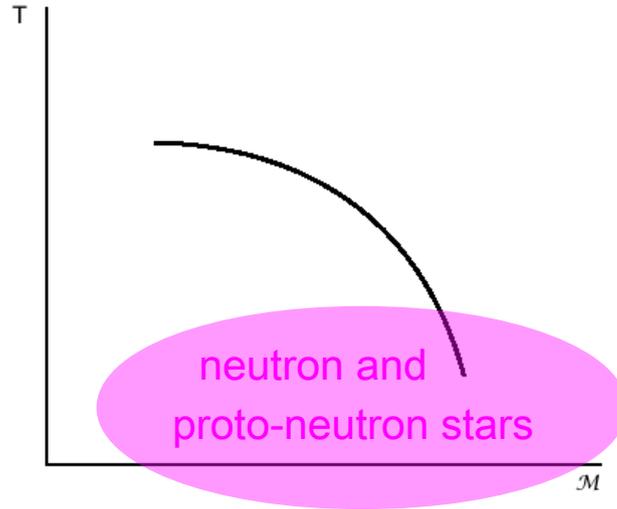
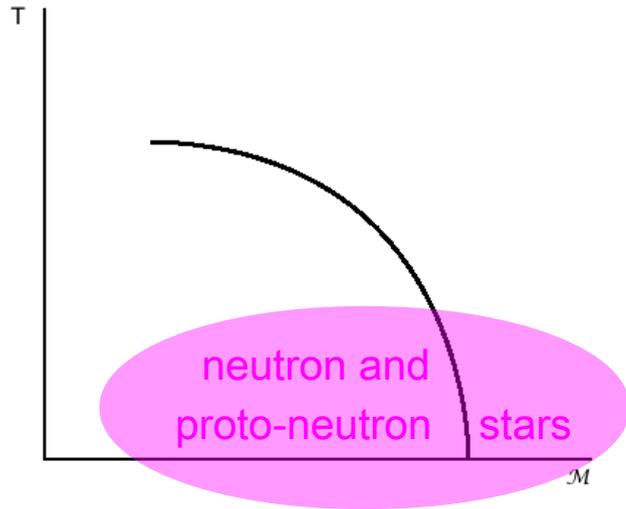
Collaborators: J. Roark, M. Hempel, I. Iosilevskiy and S.  
Schramm



# ★ Motivation:

- very little information about the QCD phase diagram at zero or low temperature

- strength of transition
- population (required for cooling simulations, etc)



## ★ CMF (Chiral Mean Field) Model:

- extended non-linear realization of SU(3) sigma model
- uses pseudo-scalar mesons as parameters of chiral transformation
- includes baryon octet, leptons and quarks
- fitted to reproduce nuclear, lattice QCD and astrophysical constraints
- effective masses

$$m_b^* = g_{b\sigma}\sigma + g_{b\delta}\tau_3\delta + g_{b\zeta}\zeta + \delta m_b + g_{b\Phi}\Phi^2$$

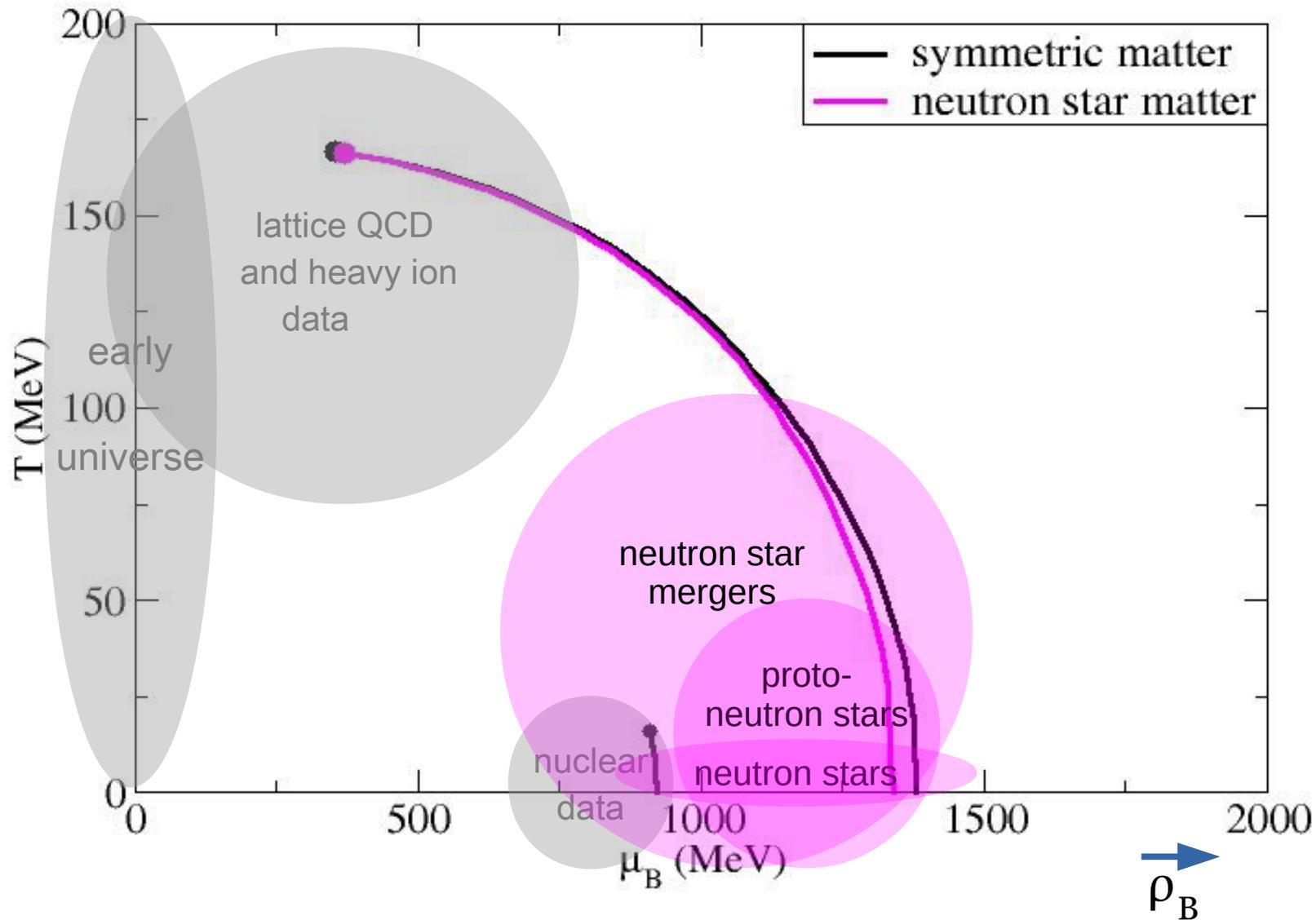
$$m_q^* = g_{q\sigma}\sigma + g_{q\delta}\tau_3\delta + g_{q\zeta}\zeta + \delta m_q + g_{q\Phi}(1 - \Phi)$$

- 1<sup>st</sup> order phase transitions or crossovers (order parameters  $\sigma, \Phi$ )
- potential for  $\Phi$  (deconfinement)

$$U = (a_0 T^4 + a_1 \mu^4 + a_2 T^2 \mu^2)\phi^2 + a_3 T_0^4 \ln(1 - 6\phi^2 + 8\phi^3 - 3\phi^4)$$

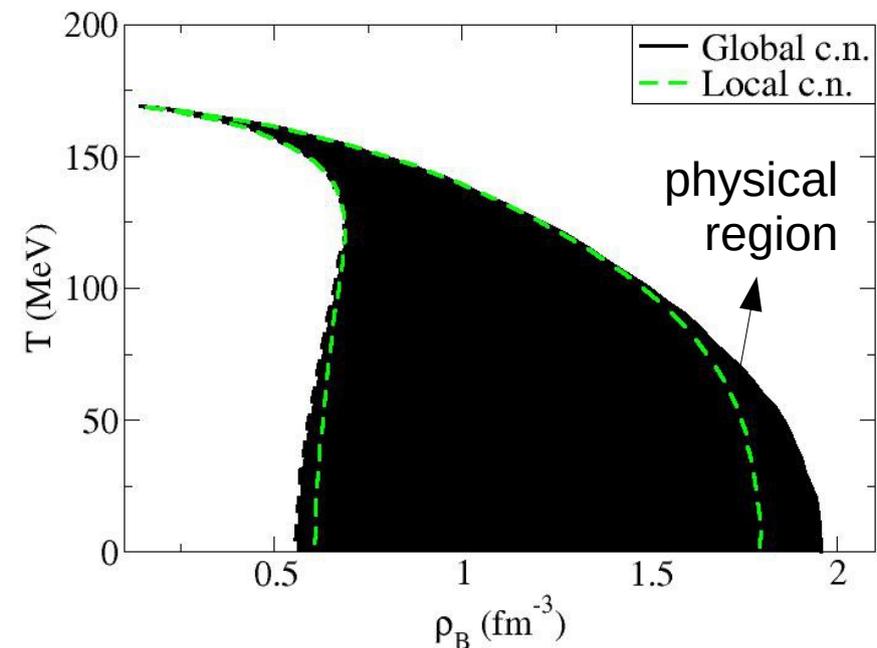
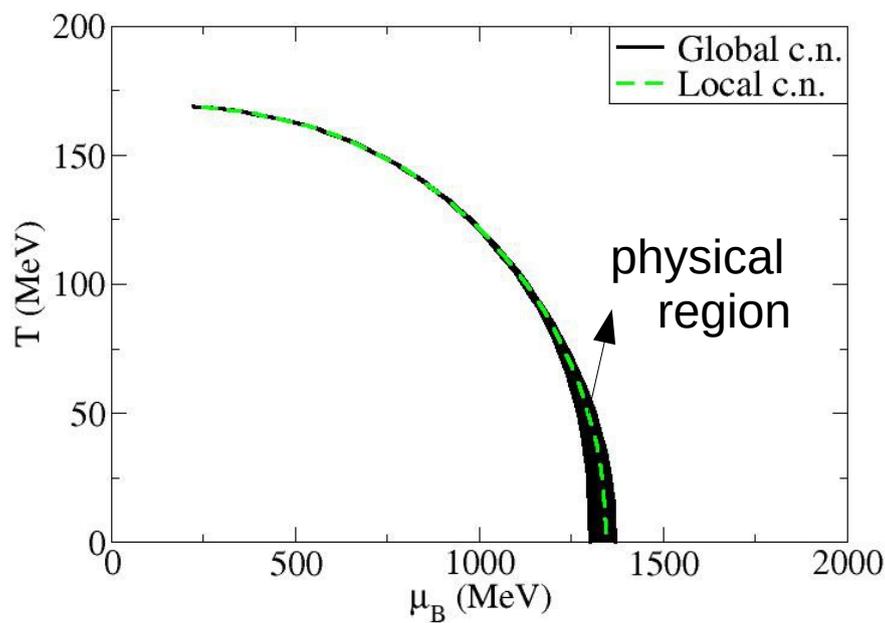
# ☆ General Picture:

Dexheimer et al. Phys. Rev. C 2010



# ★ Neutron Star Matter: Local and Global Charge Neutrality:

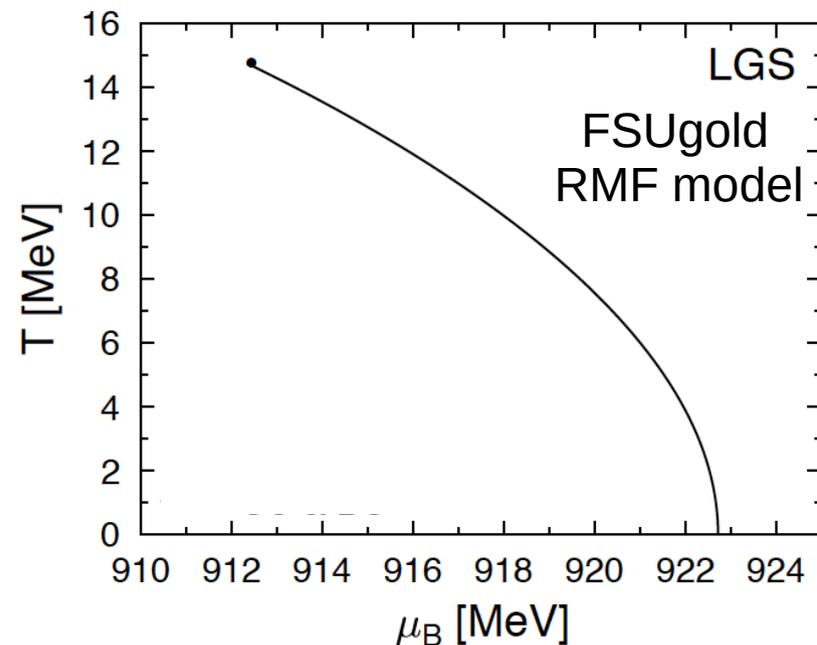
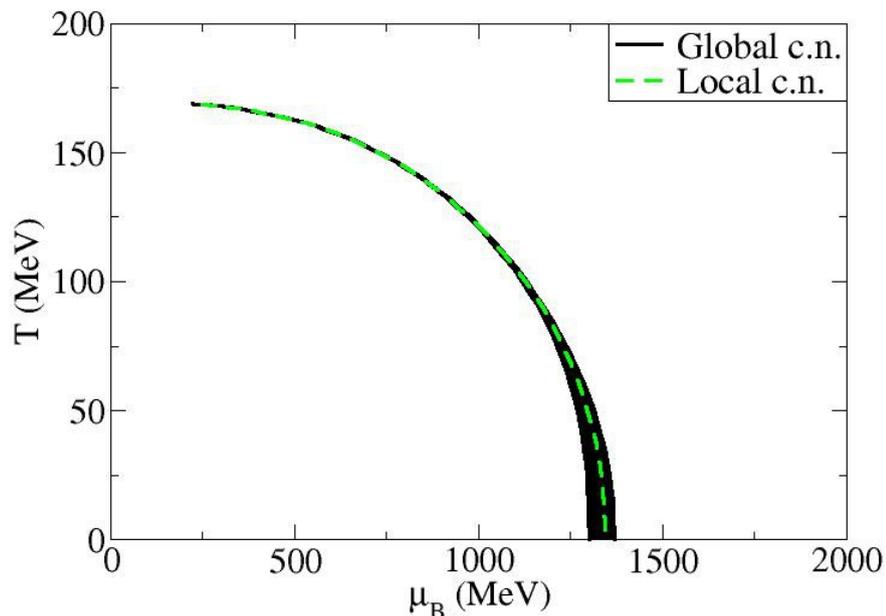
- absence / presence of mixture of phases: surface tension ???
- “mixed” quantities like  $\rho_B = \lambda \rho_B^Q + (1 - \lambda) \rho_B^H$



Hempel et al. Phys. Rev. C 2013

# ★ Non-congruent Phase Transitions:

- more than one globally conserved charge within 2 macroscopic phases (in a Coulomb-less model): baryon #, electric charge
- local concentration of charge varies during phase transition
- same chemical potential (assoc. to charge) in both phases ( $\mu_q$ )
- non-congruent features vanish around critical point
- different from symmetric matter liquid-gas



# ★ More Comparison with Liquid-Gas:

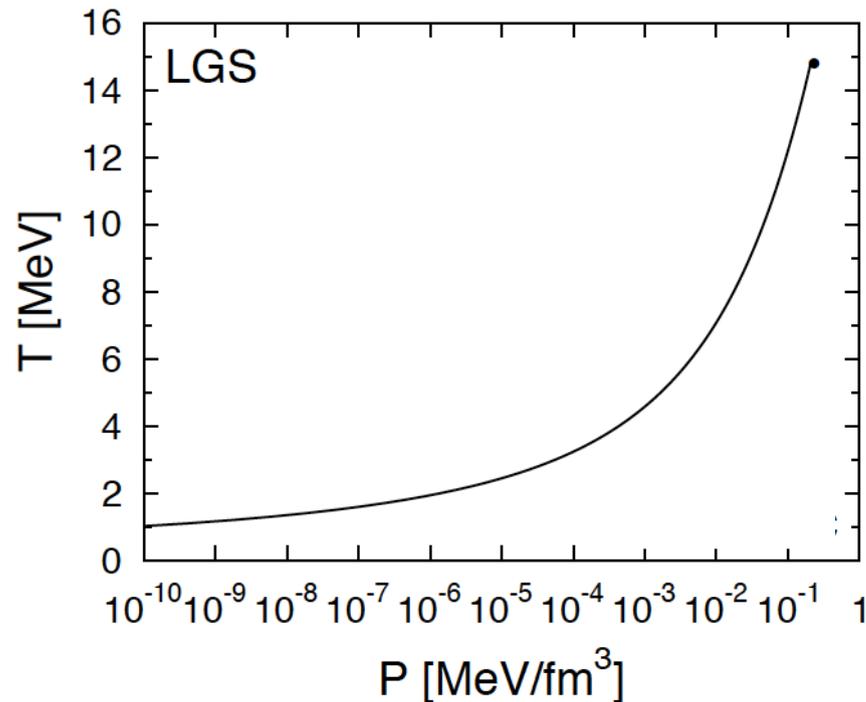
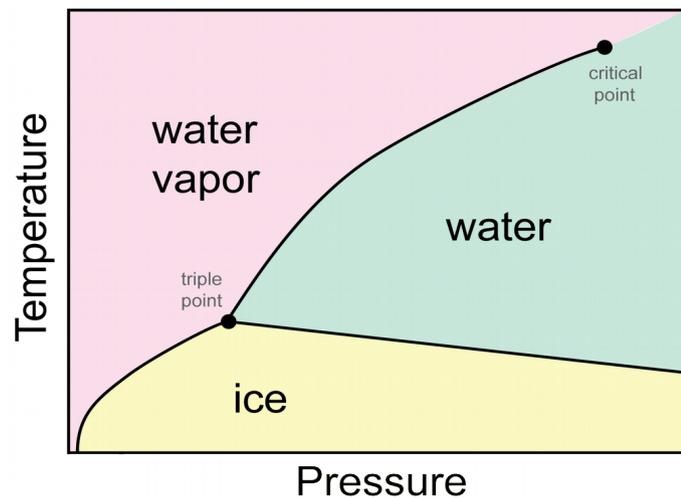
- Clausius-Clapeyron equation 
$$\frac{dP}{dT} = \frac{s^I - s^{II}}{1/\rho_B^I - 1/\rho_B^{II}}$$

-  $s_q^{II} > s_h^I, \rho_{Bq}^{II} > \rho_{Bh}^I$

so  $dP/dT < 0$  for deconfinement!

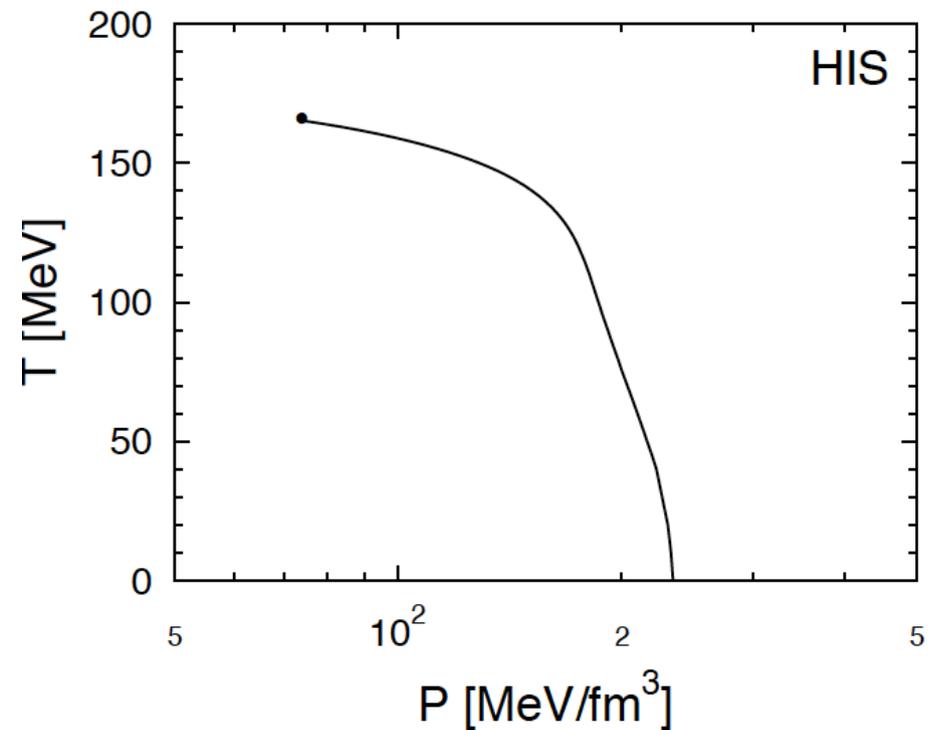
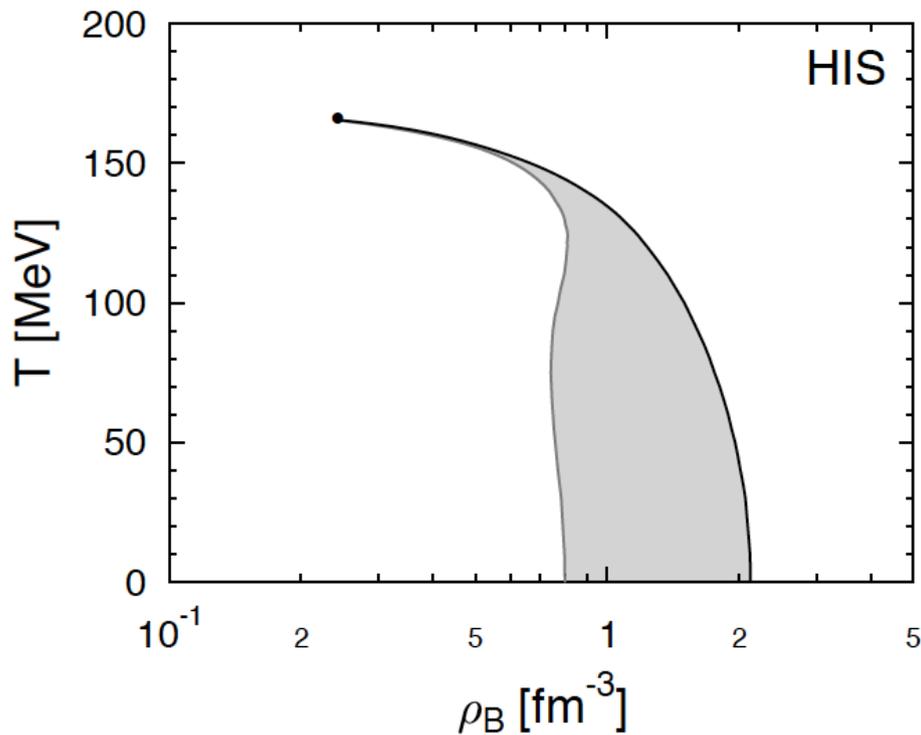
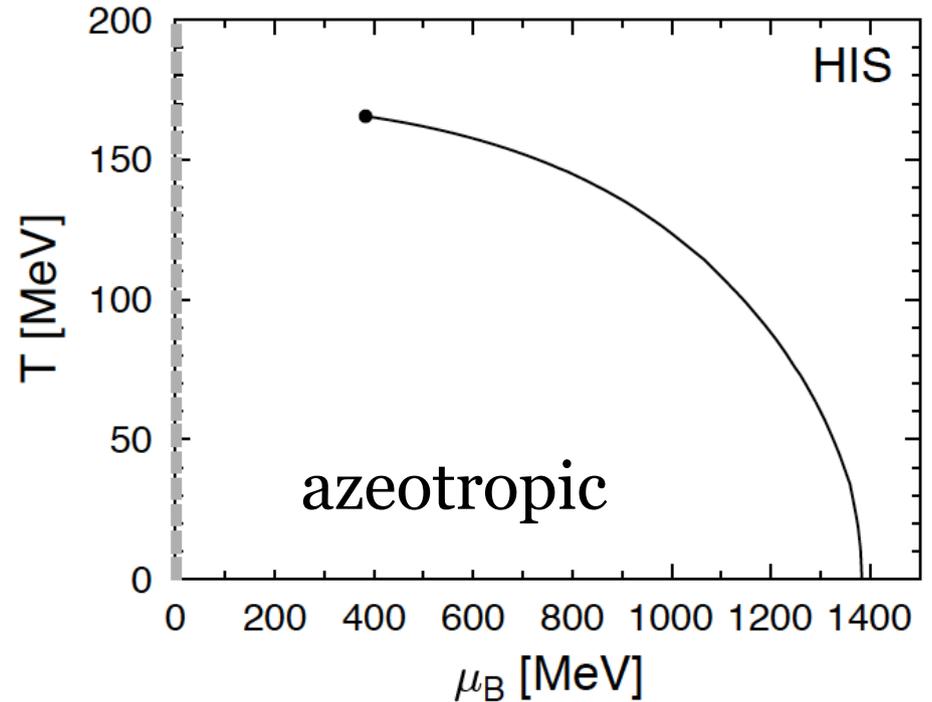
-  $s_L^{II} < s_G^I, \rho_{BL}^{II} > \rho_{BG}^I$

so  $dP/dT > 0$  for L-G!



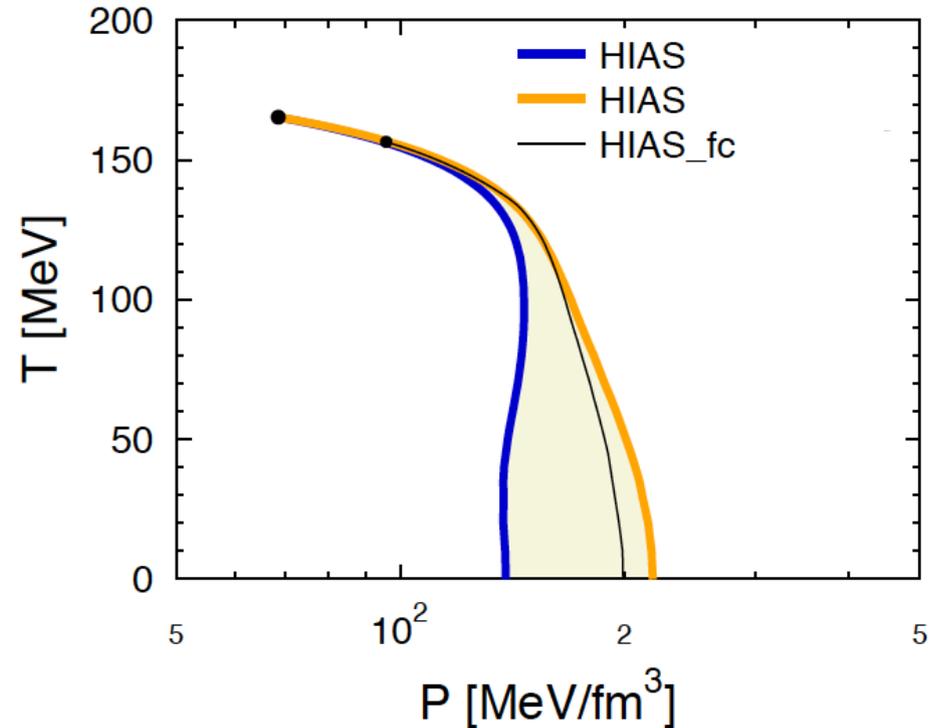
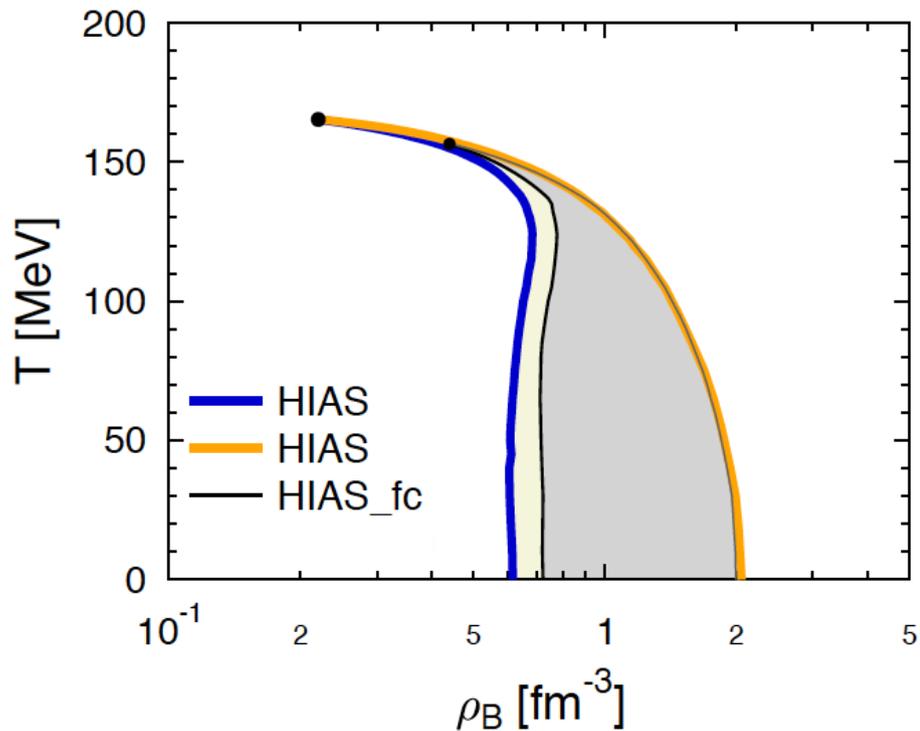
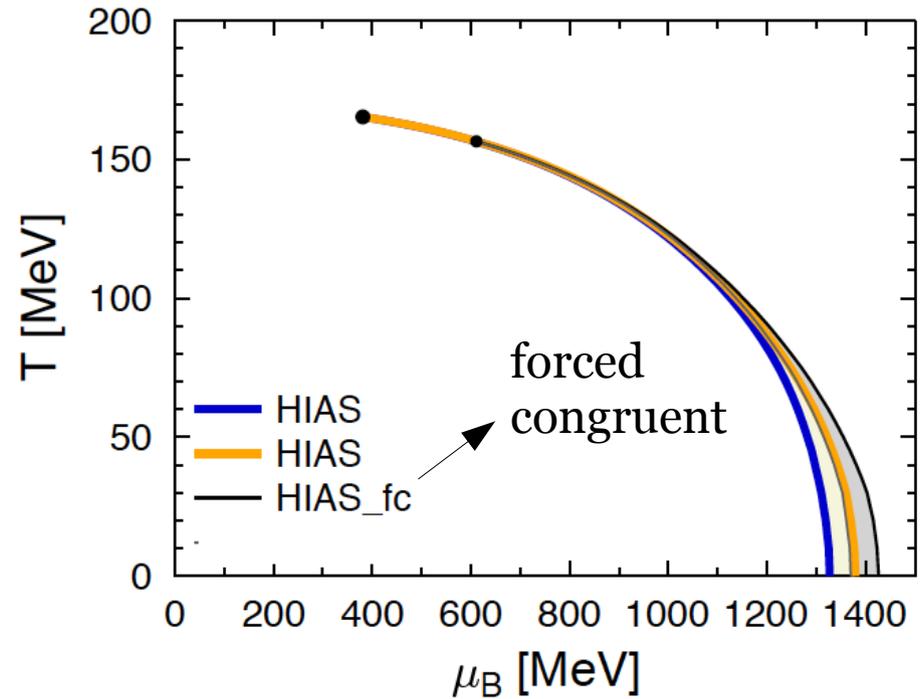
# ★ Symmetric Matter:

- heavy ion collisions  
(no net strangeness)
- more than one conserved charge (baryon #, isospin)  
but congruent phase transition! ( $\mu_q=0$ )



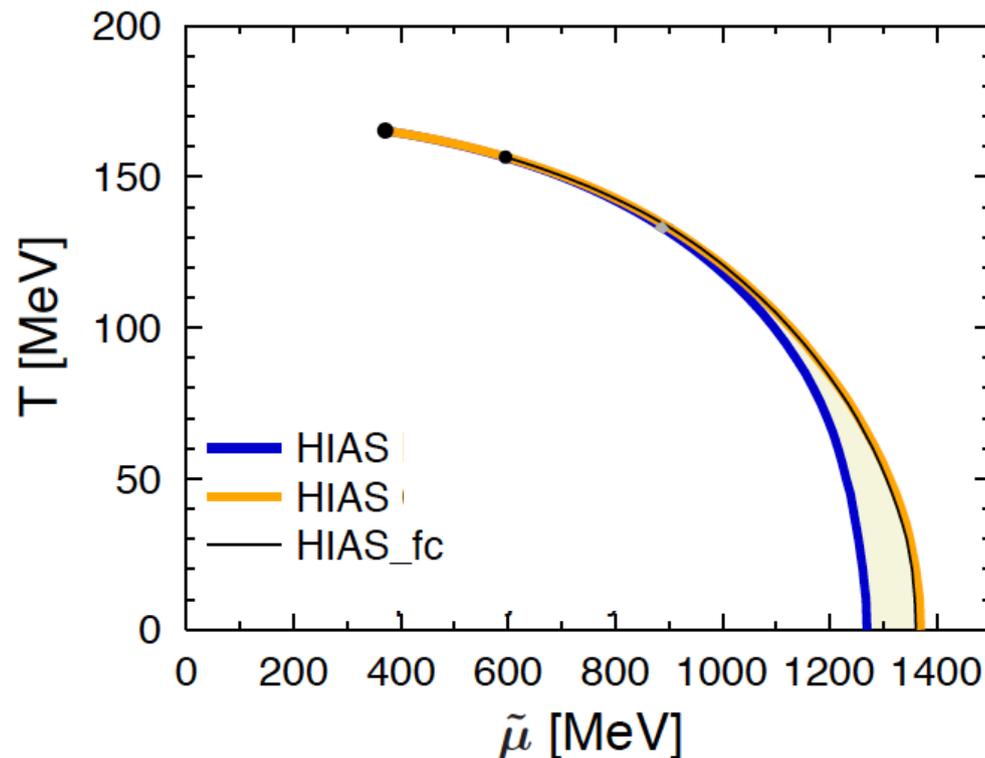
# ★ Asymmetric Matter:

- HI with  $Y_Q=0.3$
- more than one conserved charge (baryon #, charge fraction): non-congruent phase transition!
- $dP/dT < 0$



# ★ Modified Chemical Potential:

- in mixture of phases  $\tilde{\mu} = Y_Q \mu_Q + \mu_B$ , since  $\tilde{\mu}^I = \left. \frac{\partial F^I}{\partial B^I} \right|_{T, V^I, S^I, Y_Q^I} = \mu_B^I + Y_Q^I \mu_Q^I$ , is the only chemical potential which is the same in both phases
- important around phase transition
- HI forced congruent inside mixed region
- not relevant for charge neutral case



# ★ Conclusions and Outlook

- More investigation of high density part of phase diagram is required!

  - Astrophysical signature for 1<sup>st</sup> order phase transition?

- Description of compact stars requires finite temperature description

- We need a realistic EOS that covers large portion of phase diagram and provides population for simulations: only a unified EOS (usually used for L-G transitions) description of phases can provide critical points and crossovers

- congruent/not-congruent deconfinement phase transitions still being understood and might have associated signatures

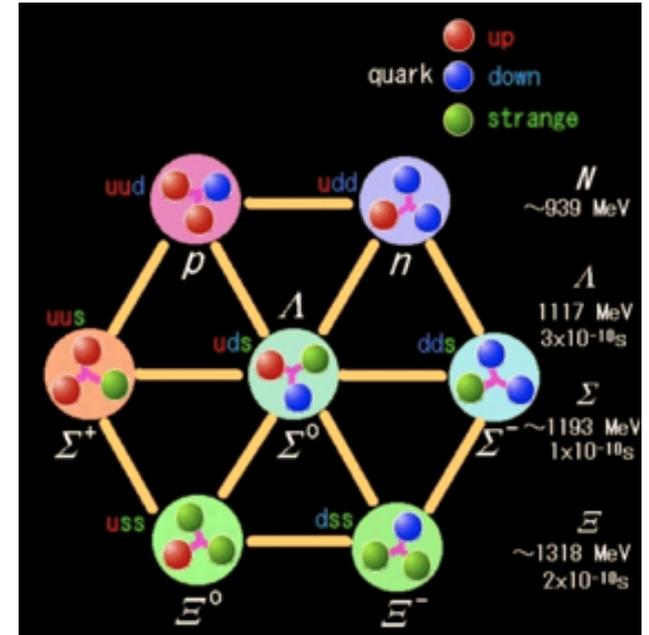
- we still need to include neutrino trapping at finite temperature

- we still need to include magnetic field effects at finite temperature

- we still need to include quark pairing effects

# ★ Ingredients for Core Description:

- baryon octet:  $p$ ,  $n$ ,  $\Lambda$ ,  $\Sigma^+$ ,  $\Sigma^0$ ,  $\Sigma^-$ ,  $\Xi^0$ ,  $\Xi^-$
- leptons ensure charge neutrality:  $e$ ,  $\mu$
- amount of each particle not constant (chemical equilibrium)
- nuclear physics constraints

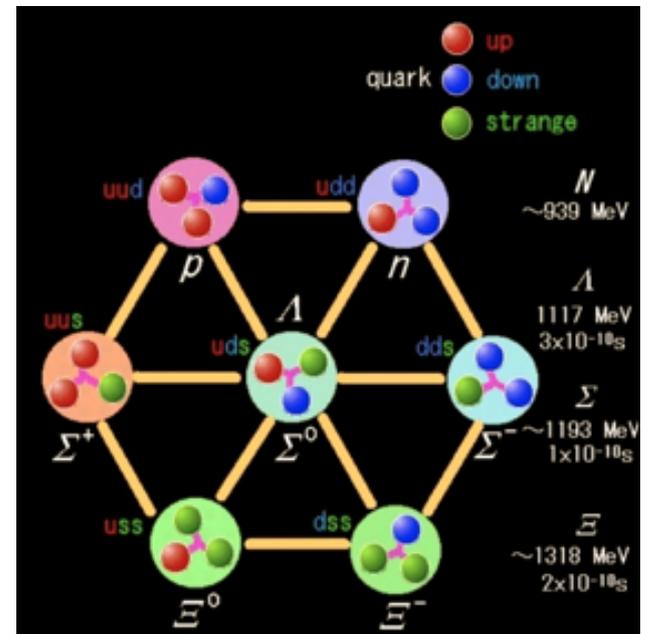


- vacuum masses of baryons and mesons
- pion and kaon decay constants
- saturation density
- binding energy at saturation
- compressibility at saturation
- symmetry energy and derivative at saturation
- hyperon optical potentials at saturation

# ★ Non-Linear Realization SU(3) Sigma Model:

- includes baryon octet and leptons
- constructed from symmetry relations → allow it to be chirally invariant → masses from interaction with medium
- pseudo-scalar mesons as parameters of chiral transformation
- $\sigma$  signals chiral symmetry restoration
- describes hadrons interacting via meson exchange ( $\sigma, \delta, \zeta, \omega, \rho, \phi$ )
- fitted to reproduce nuclear and astrophysical constraints

Dexheimer et al. *Astrophys.J.* 2008  
Negreiros et al. *Phys. Rev. C* 2010



# ★ Lagrangian Density in MFT:

$$L = L_{Kin} + L_{Int} + L_{Self} + L_{SB}$$

$$L_{Int} = - \sum_i \bar{\psi}_i [\gamma_0 (g_{i\omega} \omega + g_{i\phi} \phi + g_{i\rho} \tau_3 \rho) + M_i^*] \psi_i,$$

$$L_{Self} = \frac{1}{2} (m_\omega^2 \omega^2 + m_\rho^2 \rho^2 + m_\phi^2 \phi^2) \quad \text{reproduces vector meson vacuum masses}$$

$$+ g_4 \left( \omega^4 + \rho^4 + \alpha^2 \frac{\phi^4}{2} + 3\alpha (\omega^2 + \rho^2) \phi^2 \right)$$

$$- k_0 (\sigma^2 + \zeta^2 + \delta^2) - k_1 (\sigma^2 + \zeta^2 + \delta^2)^2$$

$$- k_2 \left( \frac{\sigma^4}{2} + \frac{\delta^4}{2} + 3\sigma^2 \delta^2 + \zeta^4 \right) - k_3 (\sigma^2 - \delta^2) \zeta$$

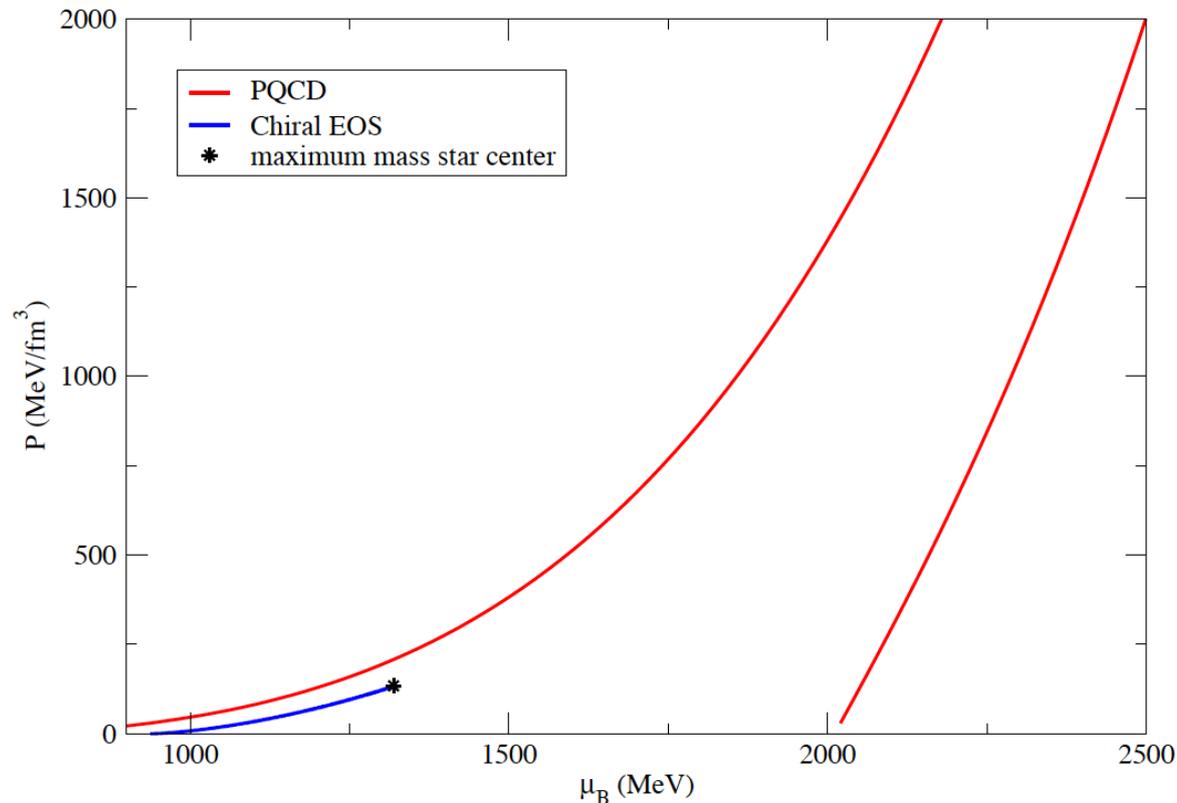
$$- k_4 \ln \frac{(\sigma^2 - \delta^2) \zeta}{\sigma_0^2 \zeta_0},$$

explicit  
symmetry  
breaking

$$L_{SB} = -m_\pi^2 f_\pi \sigma - \left( \sqrt{2} m_k^2 f_k - \frac{1}{\sqrt{2}} m_\pi^2 f_\pi \right) \zeta$$

# ★ Perturbative limit comparison

- PQCD data from: Fraga, Kurkela and Vuorinen, *Astrophys. J.* 2014 for 3-flavor QGP at zero temperature including  $\beta$ -equilibrium and charge neutrality
- For  $T=0$  things look good! Larger temperature results coming soon ...



# ★ More Comparison with Liquid-Gas:

- different behavior at  $T=0$  for hadronic matter and nuclei:  
Fermi-Dirac statistics
- all features vanish around critical point for deconfinement phase transitions

