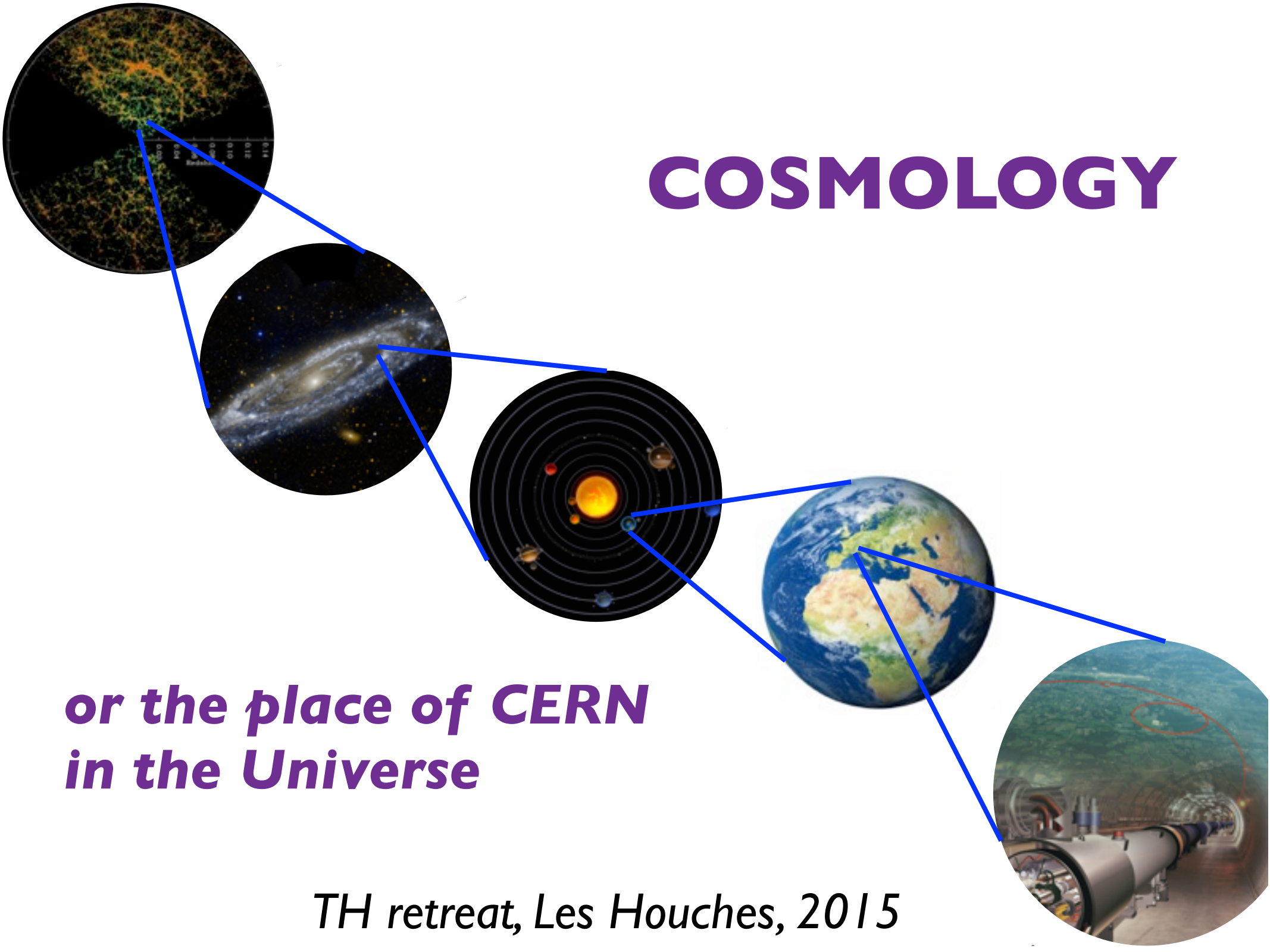


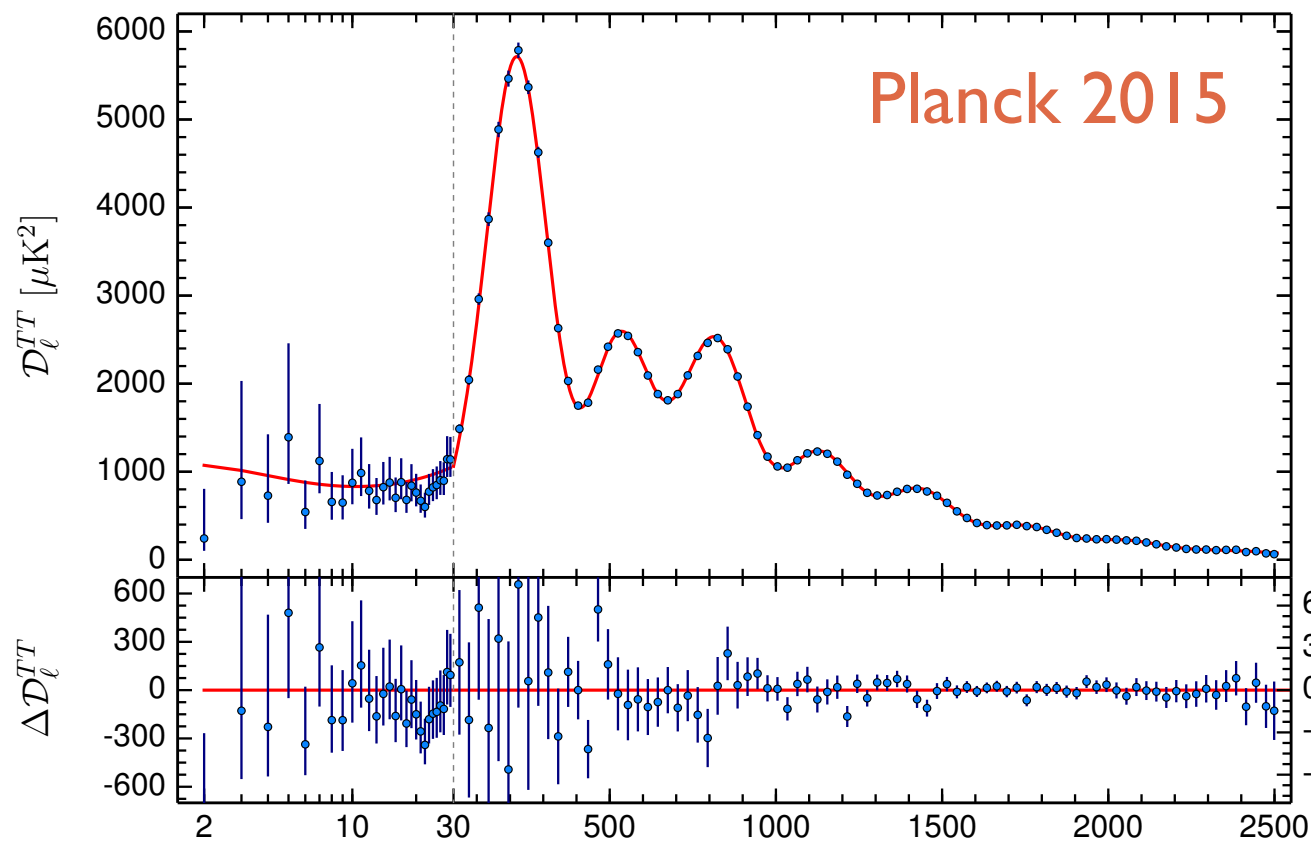
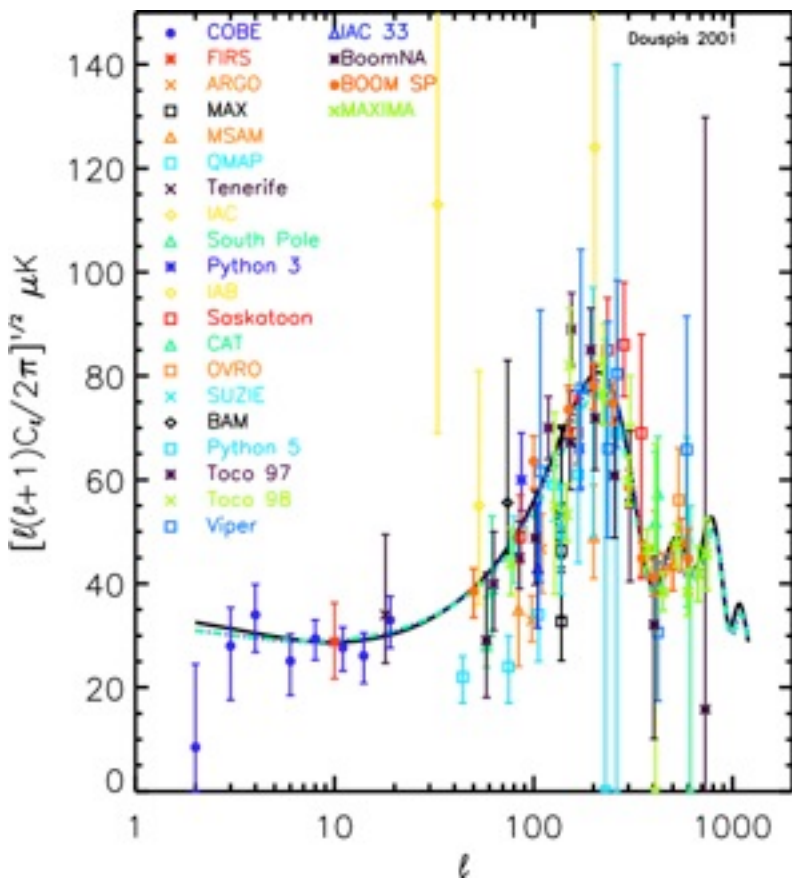
# COSMOLOGY



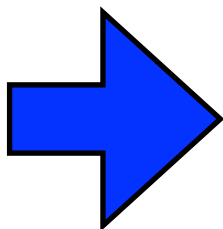
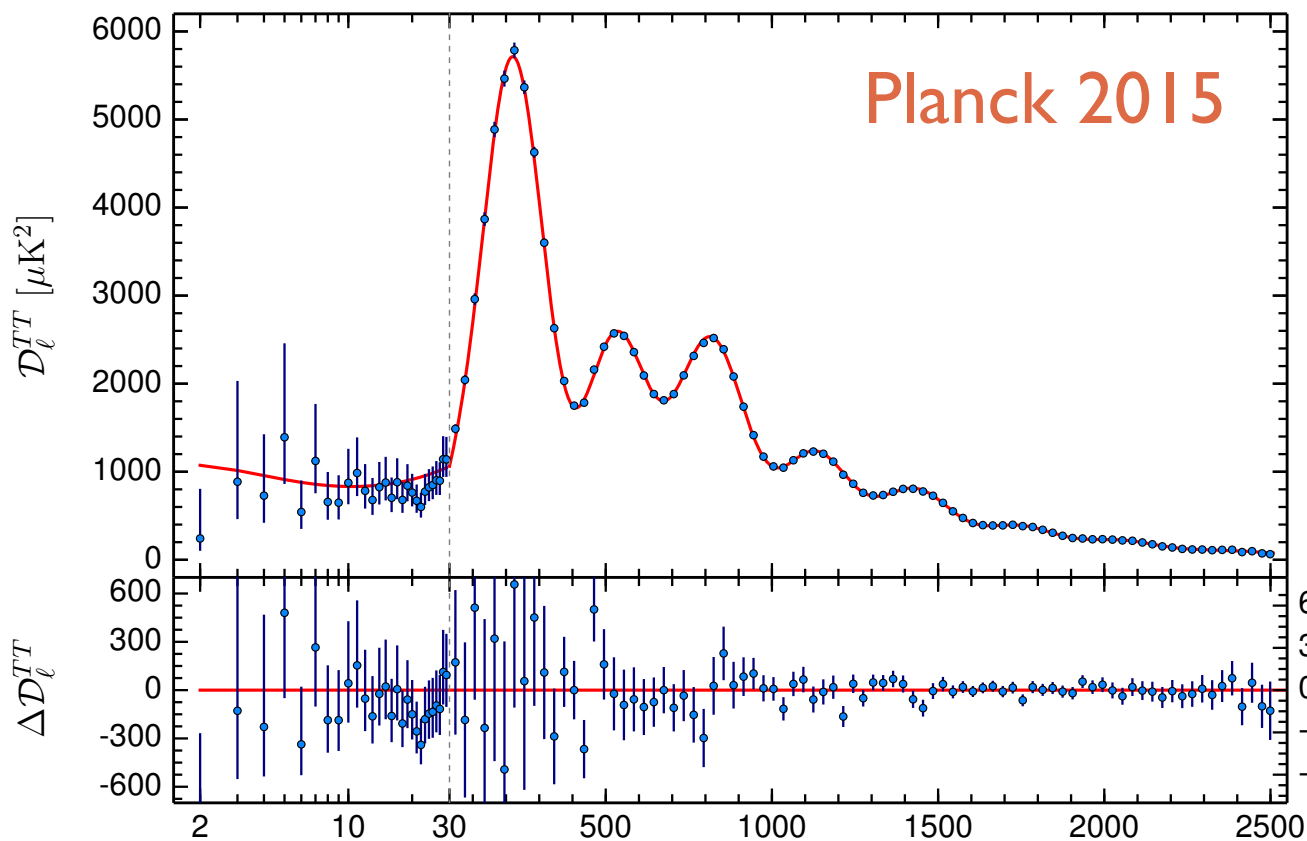
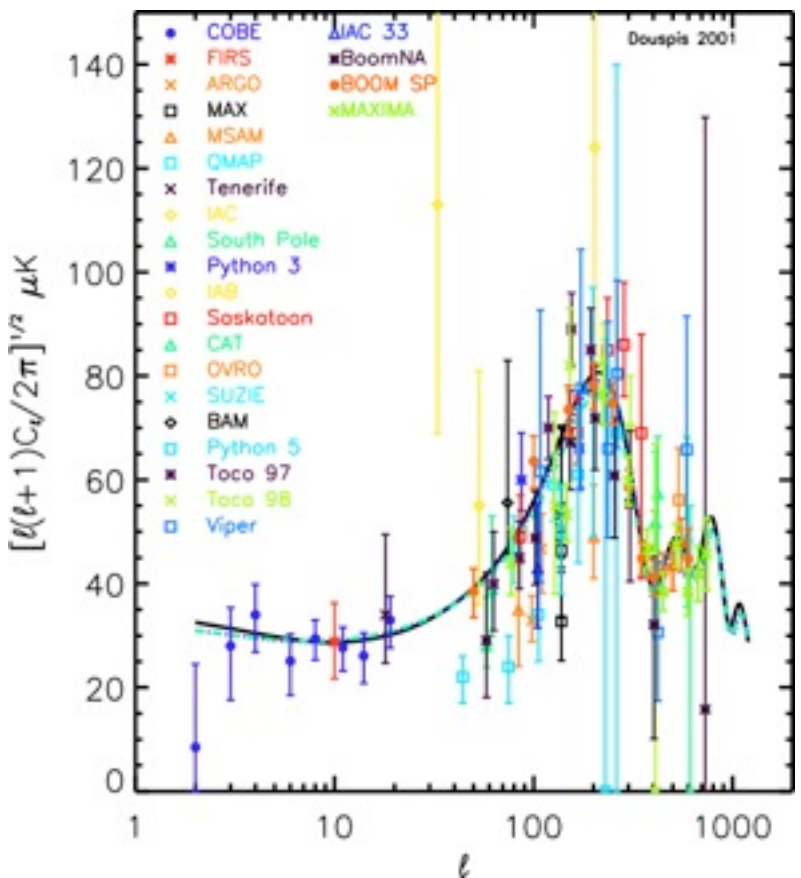
***or the place of CERN  
in the Universe***

*TH retreat, Les Houches, 2015*

# From stone age to precision cosmology

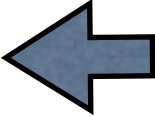


# From stone age to precision cosmology



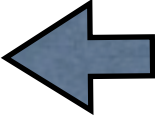
# Eldorado of data

## today:

- CMB
  - temperature anisotropies
  - polarization anisotropies  window to primordial GW
- wealth of astronomical / astroparticle data (halo profiles, X-ray,  $\gamma$  , cosmic rays, ...)

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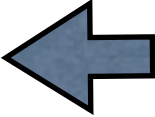
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## tomorrow: galaxy surveys

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## after tomorrow:

- 21 cm deep surveys
- spectral distortions
- multi-messenger astronomy ( $\nu$ , GW)

# Theorist's tasks

1. Learn how to extract useful information from the data
2. Understand what Universe is made of at scales from  $10^{21}$  to  $10^{26}$  m
3. Provide self-contained description of the early universe before nucleosynthesis ( $\sim 1$  min after the Big Bang)

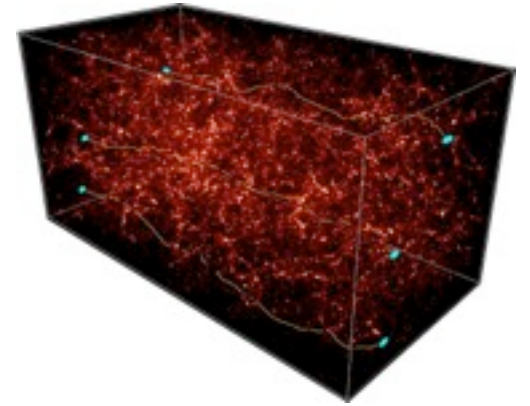
# **Theorist's tasks**

I.Learn how to extract useful information from the data



# Theorist's tasks

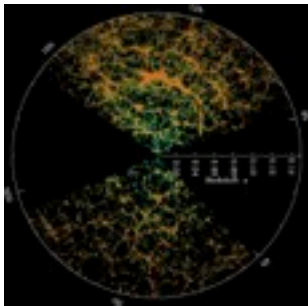
- I. Learn how to extract useful information from the data
- Control and exploit the relation between the observed signal and “primary” quantities (redshift space distortion, lensing, halo bias, ...)



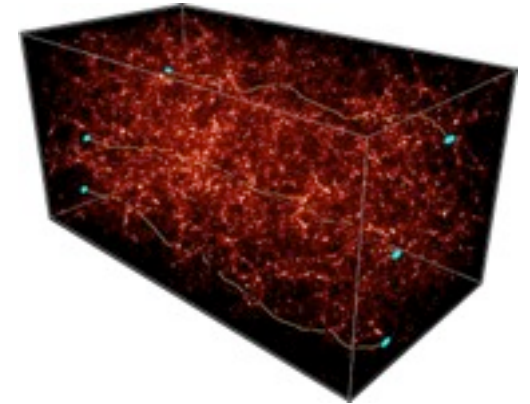
# Theorist's tasks

I. Learn how to extract useful information from the data

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- Describe structure formation in
  - perturbative (LSS)



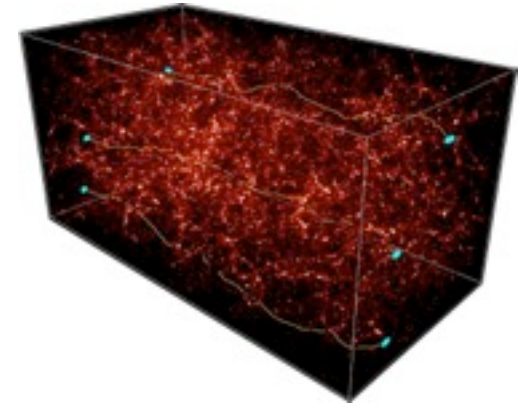
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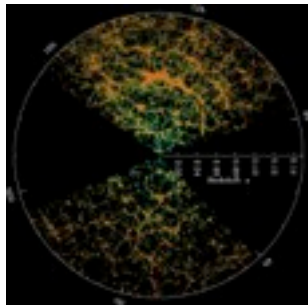
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methods of perturbative  
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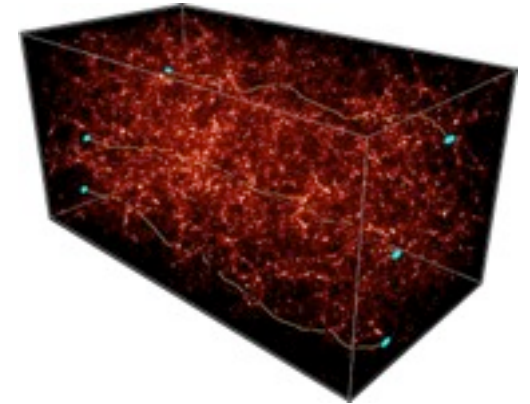
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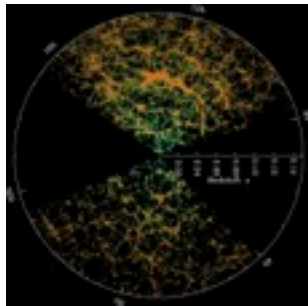
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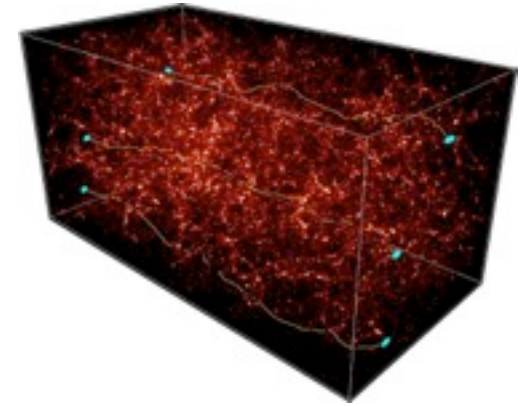


self-similar dynamics, phenomenological models

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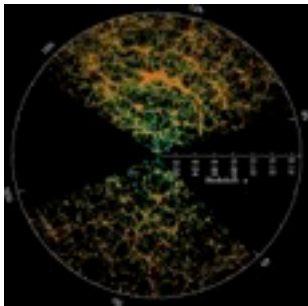
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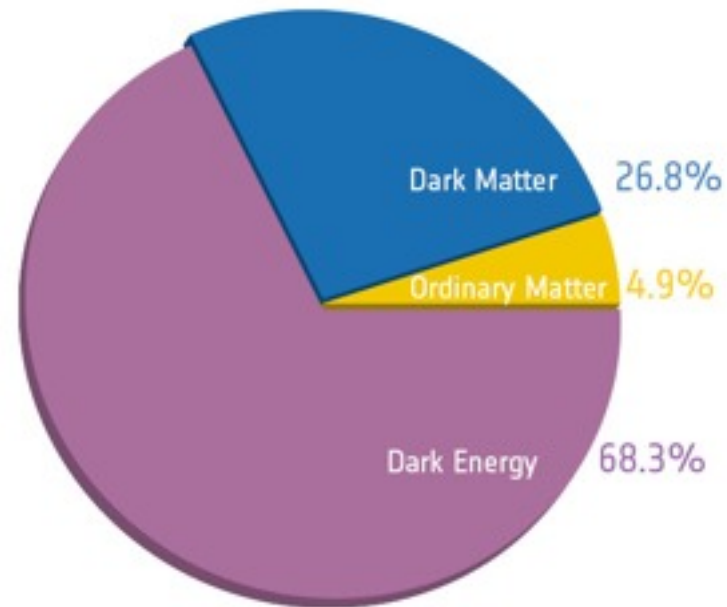


self-similar dynamics, phenomenological models

numerical N-body simulations

# Theorist's tasks

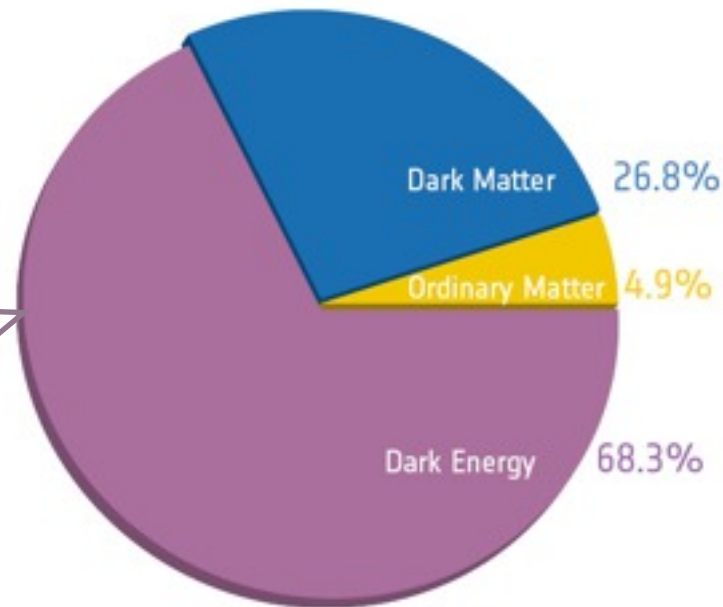
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- is it constant or evolving ?
- is it just another form of matter or gravity is different at the ultra-long scales ?

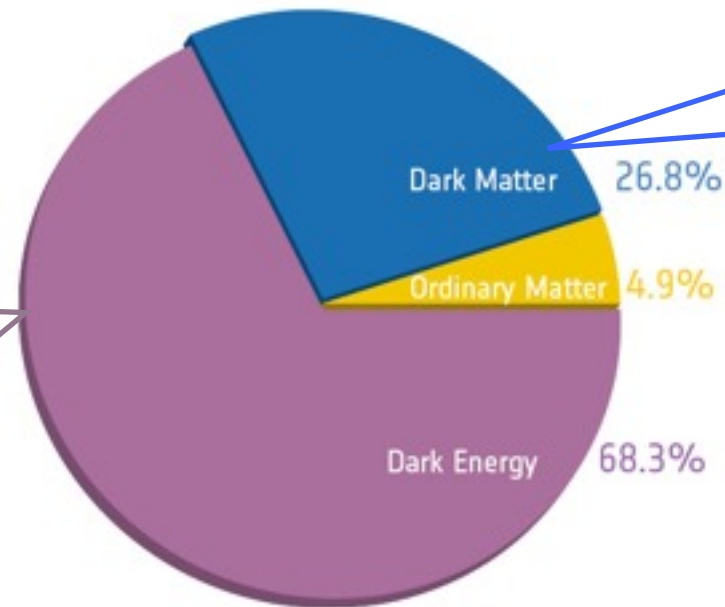


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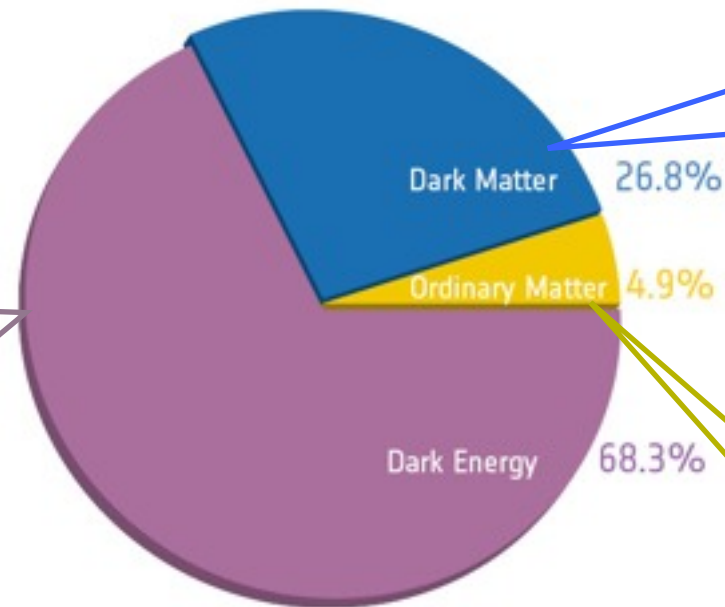


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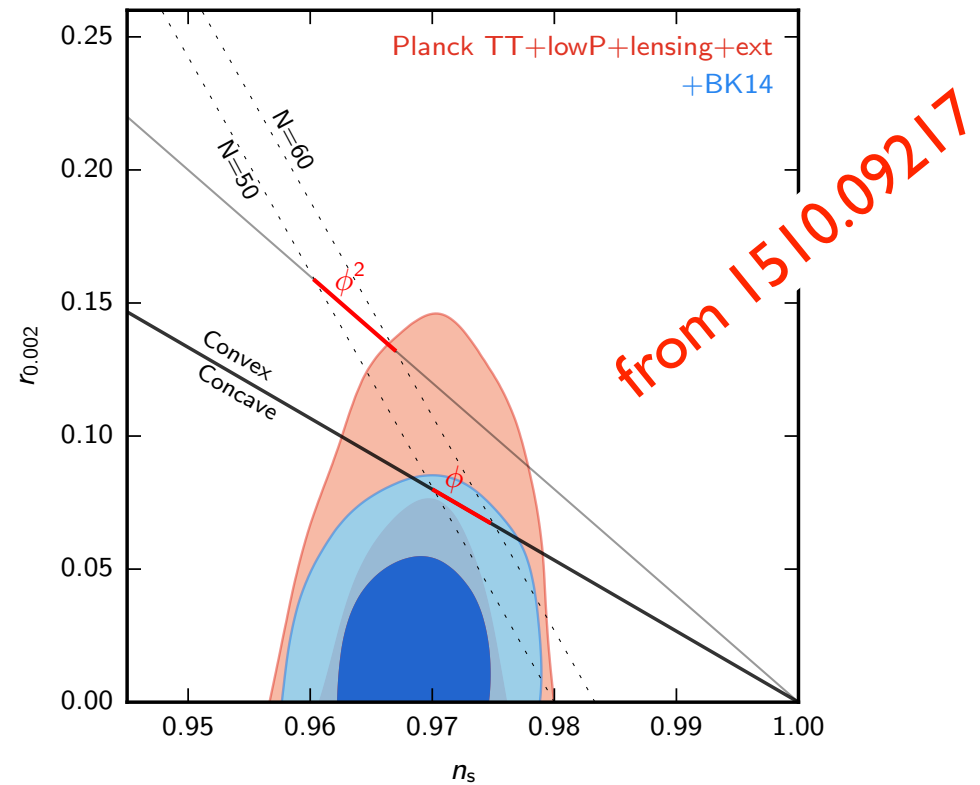
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- what are masses of neutrinos ?

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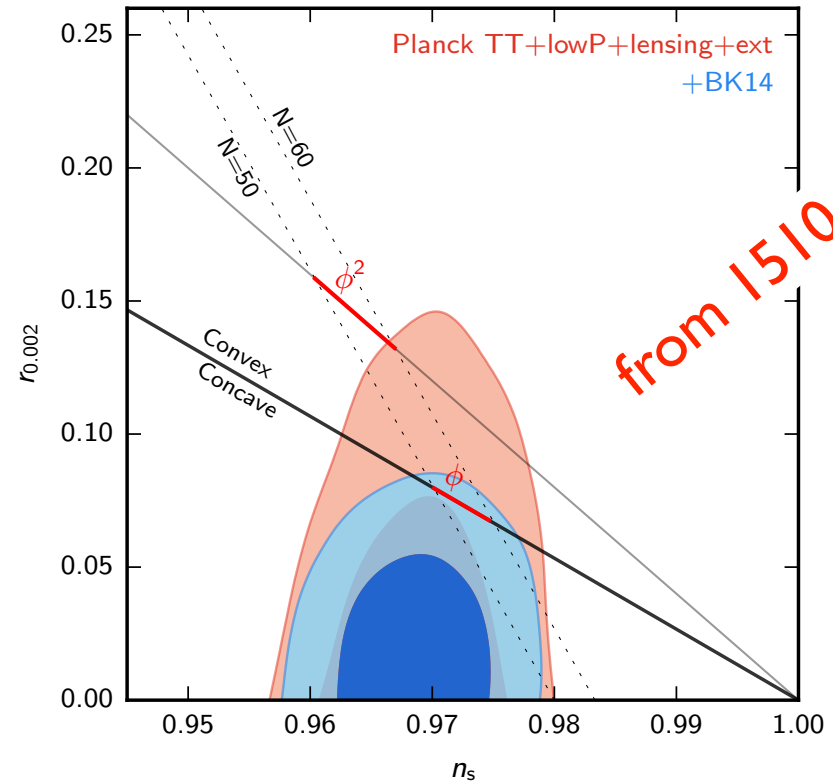
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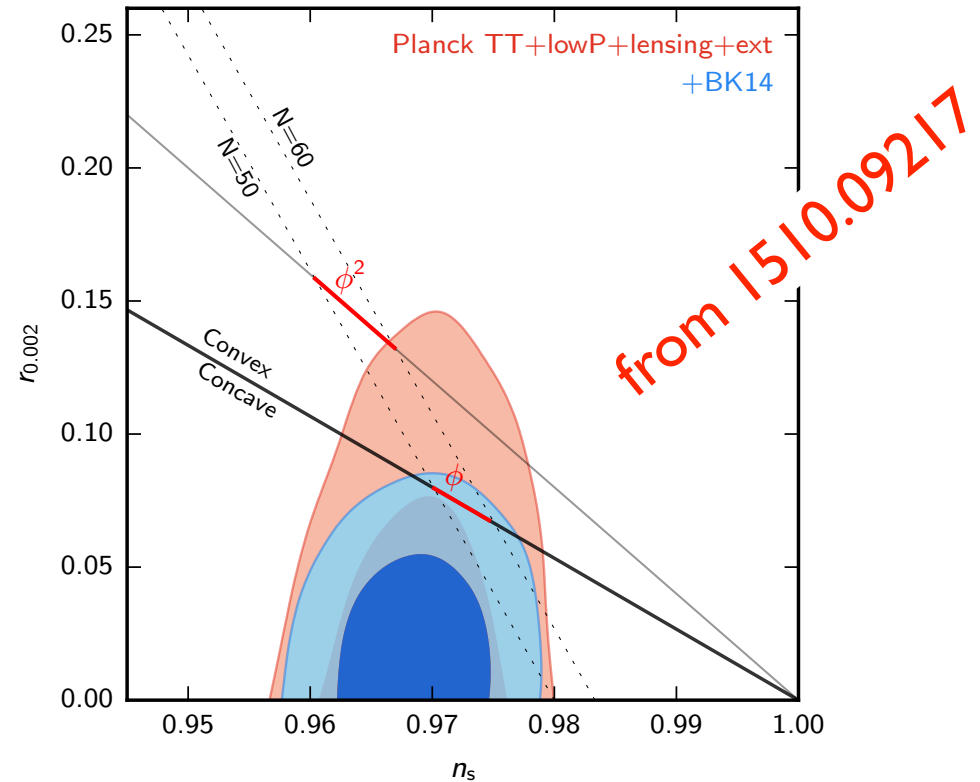
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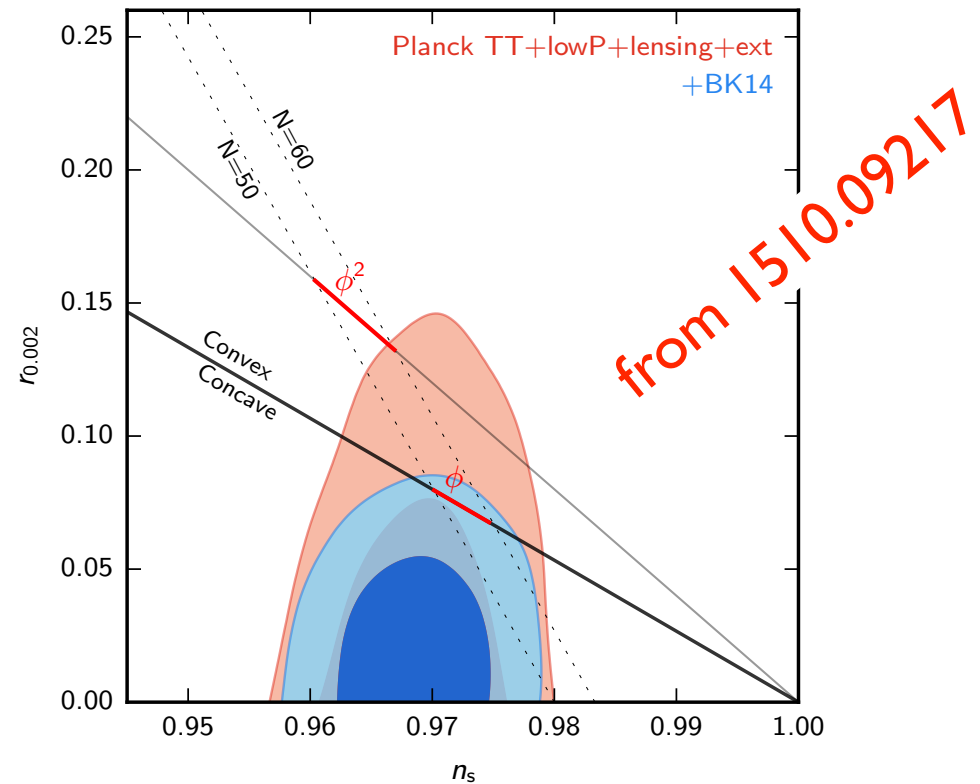
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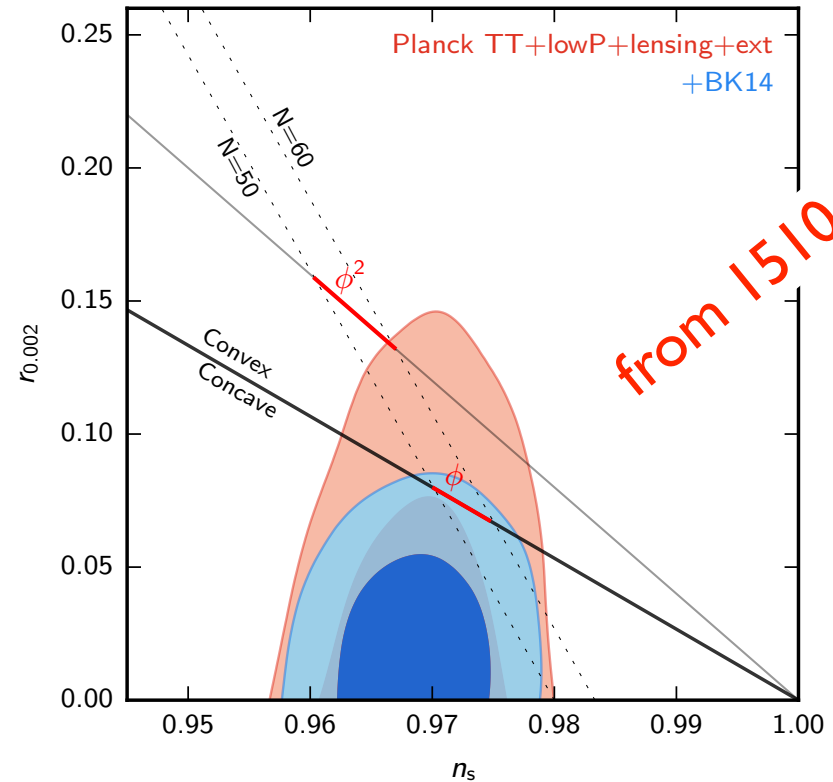


- Reheating
- Baryogenesis

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- Reheating
  - Baryogenesis
- } - how do they constrain inflationary models ?  
- observable signatures ?

# Our group

## Staff



Diego Blas



Sergey Sibiryakov

## Associate



Filippo Vernizzi

## Fellows



Camille Bonvin



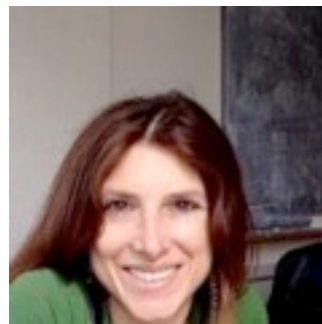
Guido D'Amico



Daniel Figueroa



Mathias Garny



Diana Lopez Nacir



Wessel Valkenburg

# Our group

Close interactions / long-term collaborations:

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- University of Geneva
- Geneva Observatory
- LAPTh, Annecy
- many others further away



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Wednesday, at 11.30 in TH Common Room:



informal seminar about  
anything related to cosmology

4) General parametrization of effects of dark energy and modified gravity on structure formation: constraints, forecasts, numerical Boltzmann code, etc

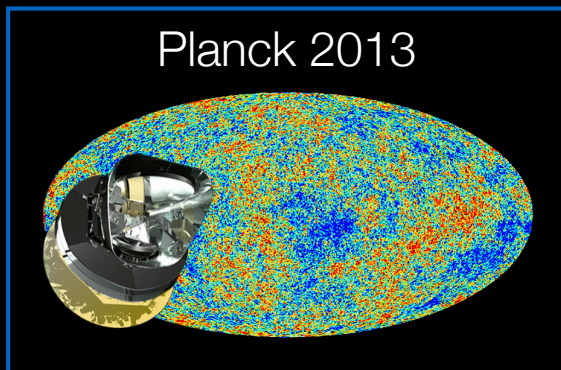


5) New classes of theoretically consistent scalar-tensor theories of gravity

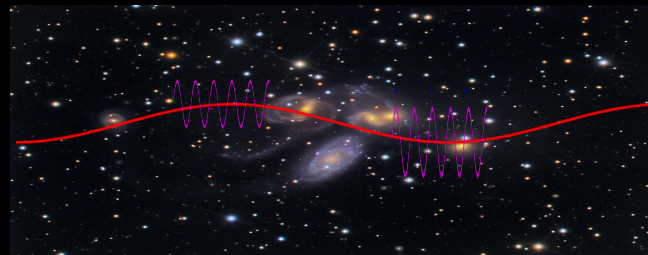
Inflation

1) Primordial non-Gaussianity from inflation

2) "Intrinsic" non-Gaussianity in the CMB



3) Consistency relations in the LSS and tests of EP violation



# About myself

- Staff at CERN & EPFL since 12.2013, on leave from INR RAS (Moscow)
- Organization of CosmoCoffee, Wednesday seminar, Academic Training
- Current research interests:
  - application of QFT / statistical physics methods to LSS
  - the role of Higgs in the early universe
  - non-stringy UV completions of gravity as renormalizable QFT (have to sacrifice Lorentz invariance or unitarity)
  - tests of Lorentz invariance in visible and dark sectors
  - RG flows with emergent space-time symmetries (Lorentz invariance, SUSY)
  - semiclassical methods for description of black hole production