

# Luminosity, Emittance Evolution & OP Scans

**M. Hostettler, G. Papotti**

Acknowledgements: F. Antoniou, A. Dabrowski, W. Herr, W. Kozanecki, M. Kuhn, P. Lujan, Y. Papaphilippou, T. Pieloni, M. Solfaroli, D. P. Stickland, G. Trad, J. Wenninger, LHC Shift Crews

**Evian, 2015-12-15**

## 1 OP scans

- technique, parameters
- errors, comparison to other instruments

## 2 emittance in stable beams

- emittance at the start of stable beams
- emittance evolution in collisions
- BCMS beam observations

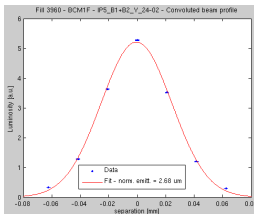
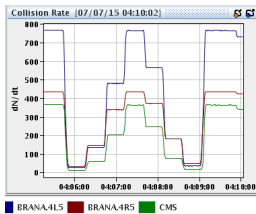
## 3 luminosity & luminosity lifetime

- luminosity & luminosity lifetime in 2015
- optimum fill length
- IP 1/5 luminosity imbalance

## 4 conclusions & outlook

# OP scans: technique, parameters

- estimate beam spot size by displacing beams at the IP and recording the luminosity change
  - fit Gaussian to relative luminosity vs beam separation
  - only dependences: knob accuracy and lumi linearity
- latest parameters & strategy
  - scan at the start, after 10h and before dump
  - IP 5 only, earlier IP 1 & 5
    - IP 1: instabilities, luminosity monitor non-linearities
    - IP 5: appreciated by CMS, bunch-by-bunch data available
  - 7 steps per plane, 10 s per step,  $\pm 3\sigma$  displacement
    - cost:  $\sim 40$  seconds at low luminosity per scan



- **errors: ~6% in the separation plane, ~20% in the crossing plane**
  - luminosity non-linearity (~5%),  $\beta^*$  (~3%), dynamic  $\beta^*$  (~2%), beam-beam kick (~2%)
  - crossing plane only: crossing angle (~15%), bunch shape (~10%)
  - bunch-to-bunch relative differences: only non-linearity
  - evolution in a fill: non-linearity and bunch shape

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- agreement with wire scanners within error bars
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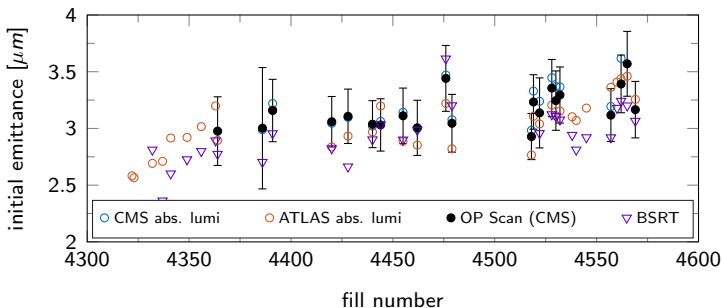
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- OP scan emittance in ATLAS and CMS compared during  $\mu$ -scan fill
  - near perfect relative bunch-by-bunch agreement
  - absolute values very sensitive to crossing angle, agree within error bars
    - extra information about geometric factors in ATLAS and CMS

# emittance at start of stable beams

- convoluted emittances at the “start” of stable beams
  - OP scans done up to ~2 h into stable beams
  - using measured  $\beta^*$  (0.84 m) and nominal crossing angle
  - emittances from luminosity, OP scans and BSRT within ~20%

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- emittances around  $3\ \mu\text{m}$



\* excluded CMS points for fills with magnet off or ramping

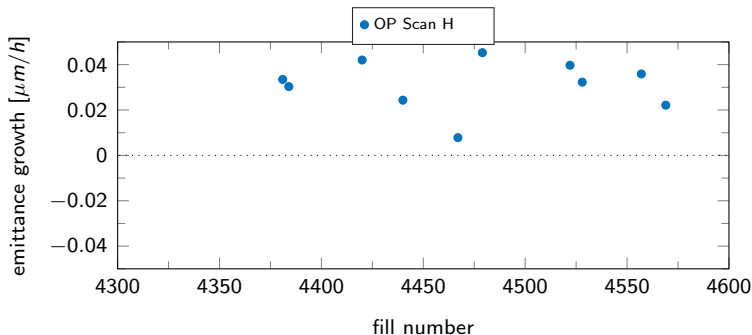


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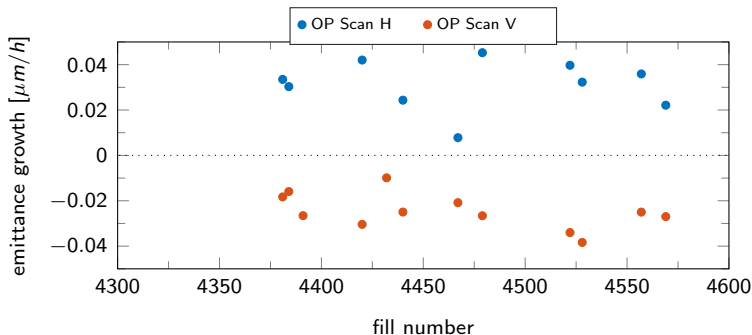
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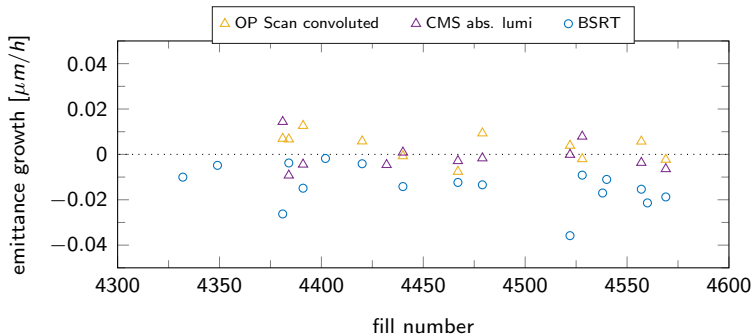
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- vertical emittance shrinkage,  $\sim 0.02 \mu\text{m}/\text{h}$
- convoluted emittance: constant within error bars
  - BSRT sees small shrinkage, difference in horizontal plane



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# BCMS beam observations (fill 4555, 601 bunches)

- mean emittances in stable beams
  - start:  $\sim 0.5 \mu\text{m}$  lower than nominal
  - horizontal growth, vertical constant

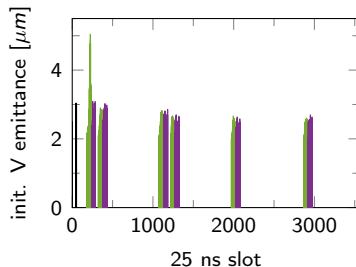
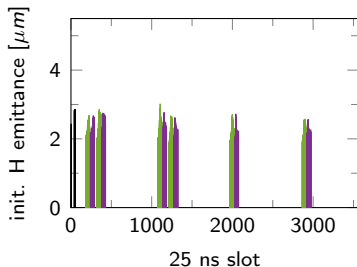
| time in SB | <b>H</b>          | <b>V</b>          |
|------------|-------------------|-------------------|
| 14 min     | $2.3 \mu\text{m}$ | $2.5 \mu\text{m}$ |
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- bunch-by-bunch observations
  - selective blow-up on first train in vertical plane
  - large spread in emittances over trains,  $\sim 1 \mu\text{m}$ 
    - conditioning possible?
  - **first** and **second** trains in a batch behave differently



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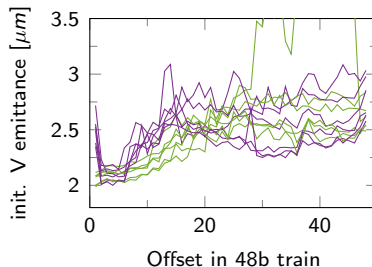
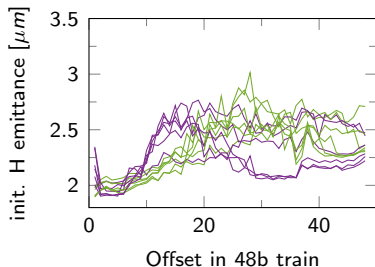
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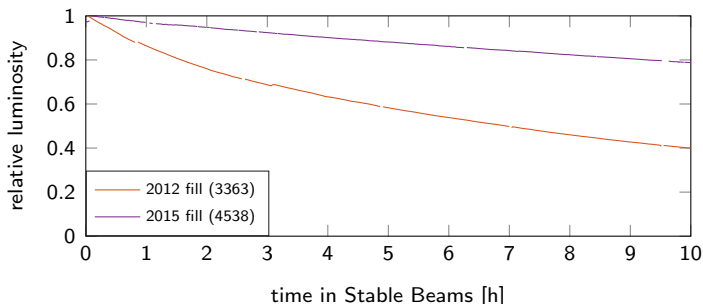
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# luminosity & luminosity lifetime in 2015

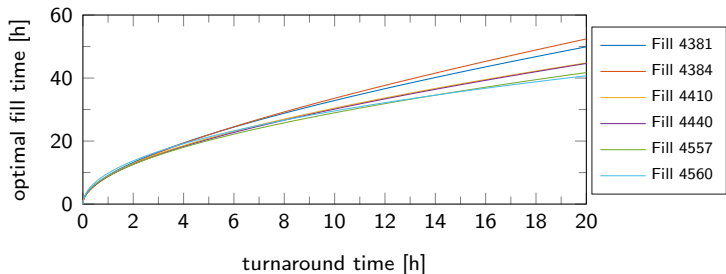
- luminosity lifetime very good w.r.t. 2012
  - 2015: 30 – 60 h, peak  $\sim 5000 \text{ Hz}/\mu\text{b}$
  - 2012: 5 – 10 h, peak  $\sim 7500 \text{ Hz}/\mu\text{b}$
- 6.5 TeV: synch light damping, emittance shrinking
- 25 ns: lower initial brightness, intensity and luminosity per bunch
- luminosity decay strongly dominated by intensity decay





# optimum fill length (from 2015 lumi data)

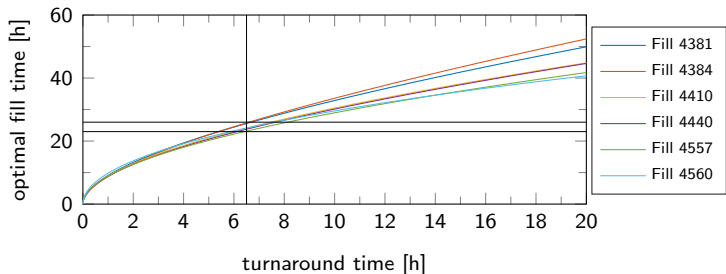
- optimum fill length calculated from direct fit to luminosity curves
  - “turnaround time” = time from dump to next physics
  - see LBOC presentation of 2012-10-30



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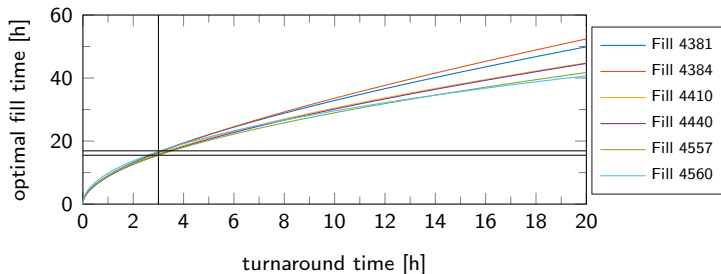
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- for turnaround times in 2015 (*M. Solfaroli*)
  - most probable case: ~6.5 h
    - optimum fill length: ~25 h



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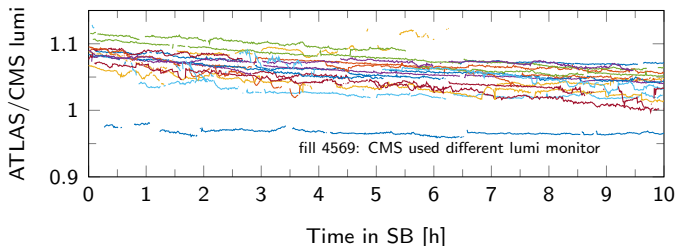
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  - optimistic case: ~3 h
    - optimum fill length: ~16 h



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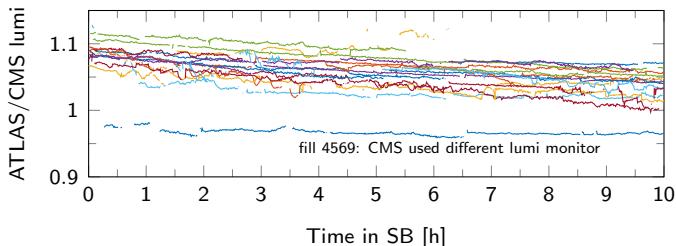
# luminosity imbalance between ATLAS and CMS

- CMS luminosity consistently lower than ATLAS
  - ~9% at the start of stable beams, ~4% in the end



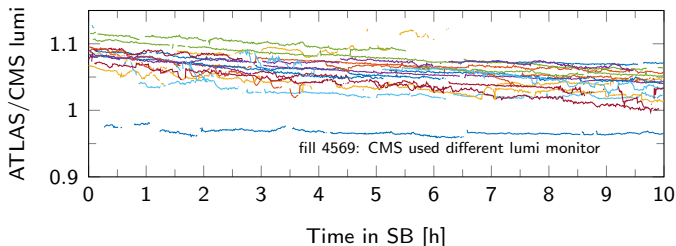
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  - $\beta^*$ , dynamic  $\beta$ , form factor (*R. Tomas et al., LBOC 2015-10-27*)
  - crossing angle (*J. Wenninger, LBOC 2015-10-27*)
  - indication from  $\mu$  scan fill: luminous area ~2% bigger in CMS



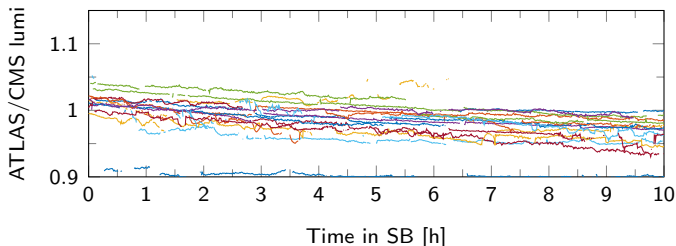
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- latest results from VdM scans in August
  - ATLAS high by 3.3%, residual error on calibration 5% (*W. Kozanecki*)
  - CMS low by ~4%, residual error on calibration 4.6% (*P. Lujan*)



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  - indication from  $\mu$  scan fill: luminous area ~2% bigger in CMS
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  - ATLAS high by 3.3%, residual error on calibration 5% (*W. Kozanecki*)
  - CMS low by ~4%, residual error on calibration 4.6% (*P. Lujan*)
  - luminosity imbalance down to ~1%



- luminosity & luminosity lifetime
  - peak  $\sim 5000 \text{ Hz}/\mu\text{b}$ , lifetime 30 – 60 h
  - 25 h optimal fill length for 6.5 h turnaround time
- emittance in stable beams
  - nominal beams:
    - start of stable beams:  $\sim 3 \mu\text{m}$
    - horizontal growth, vertical shrinking, convoluted constant
  - BCMS beams (only one fill!)
    - start of stable beams:  $\sim 2.5 \mu\text{m}$ , bunch-by-bunch spread
    - horizontal growth, vertical constant, convoluted growth
- OP scans
  - complementary emittance measurements in 2015
    - self-consistent and in agreement with wire scanners & BSRT
  - proposal for 2016:
    - IP5 only, end-of-2015 parameters (7 steps, 10 seconds per step)
    - scan before programmed beam dumps or after  $\sim 15 \text{ h}$
    - scan at the start of stable beams: frequency to be discussed with CMS
    - dedicated tab and online analysis in the new lumi scan application