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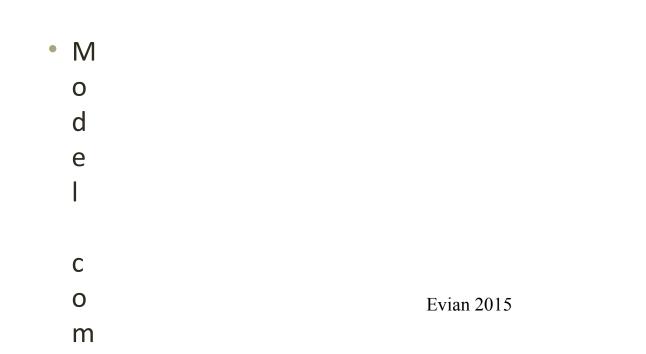
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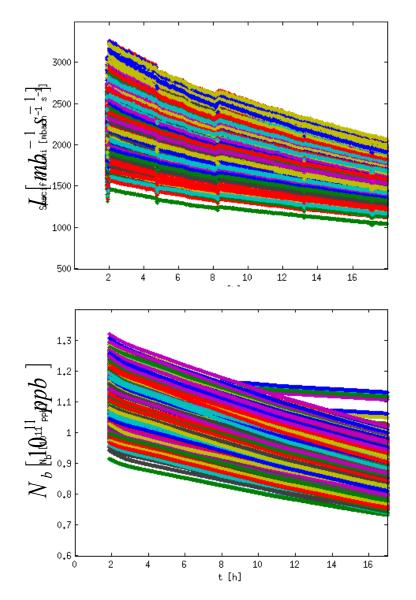
Outline

- Introduction: Luminosity
- Luminosity model components
- Luminosity performance: Runl Vs Runll
- Luminosity model Vs RunII data
- Optimal Fill times for 2016
- Summary and Outlook

Introduction: Luminosity



Luminosity decay



- During stable beams the interplay between different effects affects the bunch characteristics evolution:
 - Intra-beam scattering, Burnoff, Synchrotron Radiation, Beam-beam, noise, other unknown mechanisms, ...
- Many of the effects depend on the initial bunch brightness
- Big spread in the bunch by bunch behavior is observed

→ We need a model that takes the bunch-by-bunch variations into account

Model components (1)

Emittance and bunch length evolution at Flat Top energy:

- Intrabeam scattering (IBS):
 - Multiple Coulomb scattering effect leading to the redistribution of phase space and finally to emittance blow up in all three planes $\frac{d\varepsilon_{i}}{dt} = f(En, N_{b0}, \varepsilon_{x0}, e_{y0}, \sigma_{l0})$

 \rightarrow Analytical integrals

Iteration in time as the beam characteristics are evolving

Synchrotron Radiation (SR):

 ε_{r} =At high ehergies becomes important for proton beams as well, leading to emittance damping in all three planes

emittance damping time

The emittance evolution due to IBS and SR has been fully parameterized

 $\begin{bmatrix} \varepsilon_x(t_i), \varepsilon_y(l_i), \sigma_l(t_i) \end{bmatrix} = Compute IBSE mile vol (En, N_{b0}, \varepsilon_x(t_0), \varepsilon_y(t_0), \sigma_l(t_0), ilmestep) \\ Their effect in any plane can be calculated through a function: \\ \end{bmatrix} = Compute IBSE mile vol (En, N_{b0}, \varepsilon_x(t_0), \varepsilon_y(t_0), \sigma_l(t_0), ilmestep) \\ \end{bmatrix} = Compute IBSE mile vol (En, N_{b0}, \varepsilon_x(t_0), \varepsilon_y(t_0), \sigma_l(t_0), ilmestep) \\ \end{bmatrix} = Compute IBSE mile vol (En, N_{b0}, \varepsilon_x(t_0), \varepsilon_y(t_0), \sigma_l(t_0), \sigma_l($

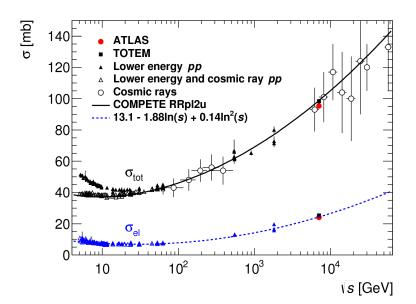
Model components (2)

- Bunch intensity degradation
 - Luminosity burn-off: Luminosity decay due to the collisions themselves:

Evian 2015

$$\tau_{nuclear} = \frac{N_{tot,0}}{L_0 \sigma_{tot} k} \qquad N_{tot}(t) = \frac{N_{tot,0}}{1 + t/\tau_{nuclear}}$$

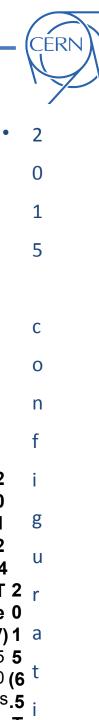
- $N_{\scriptscriptstyle tot\,,\,0}$: the initial beam intensity
- L_0 : the initial Luminosity
- $\sigma_{\it toi}$ the total cross section
- *k* the number of interaction points
- It can be easily folded into the emittance evolution function in order to have a self consistent evolution



Luminosity model summary

- The **basic model** includes the three main mechanisms of the transverse emittance, bunch length, bunch intensity and luminosity evolution due to **IBS**, **SR** and **Burn-off**
- It can be easily applied and compared with the data for bunch by bunch and averaged quantities studies
 - On going effort to find correlations from the data from average and bunch by bunch behavior (brightness, long ranges, losses, blow-up,...)
- Other sources need to be considered
 - Non-linearities of the machine
 - Noise effects
 - Scattering on residual gas

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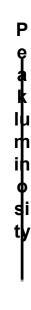
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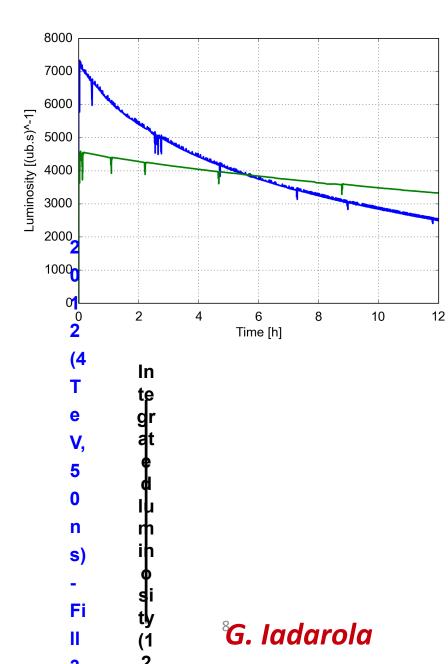
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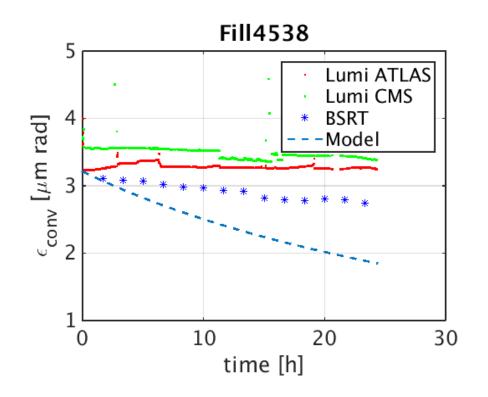




Analyzing RunII data

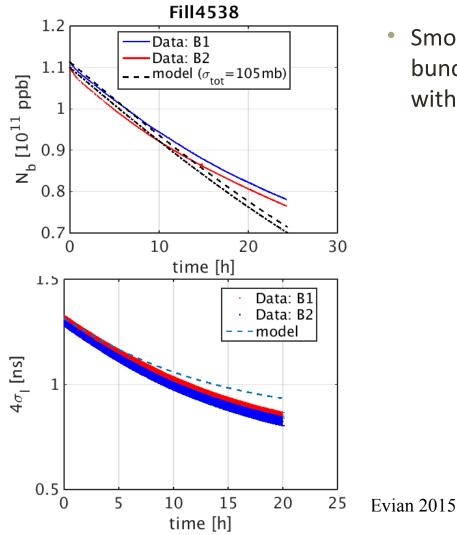
- In Run II we have emittance measurements both at Flat Bottom and Flat Top
 - **BSRT** data for both beams and both planes
 - Convoluted emittance from luminosity from the **experiments**
 - Convoluted horizontal and vertical emittance from **OP scans**
- Comparison between the different methods not always in good agreement
 - Work in progress to understand the data
- The data were compared with the model predictions for all Fills that arrived at Stable Beams

Luminosity model Vs RunII data: Emittance @ SB



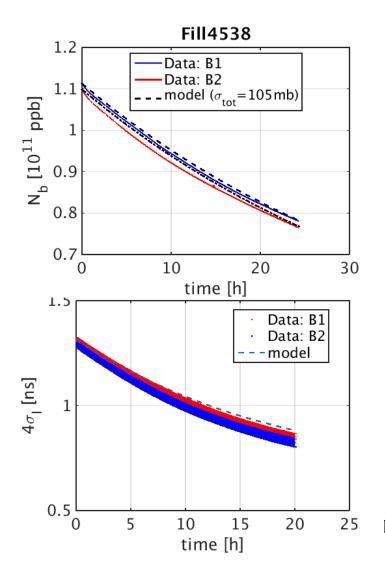
- Fill 4538 is used as an example here
- Emittance evolution (averaged) during SB from BSRT, Lumi ATLAS and Lumi CMS
 - Different evolution
- Blow up observed, with respect to the model
- We need to understand the data and include other sources of emittance blow-up

Luminosity model Vs RunII data: Bunch current & bunch length @ SB



 Smoother current decay and more bunch length damping is observed with respect to the model prediction

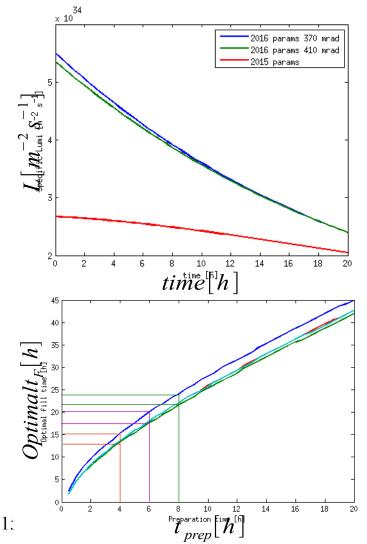
Luminosity model Vs RunII data: Bunch current & bunch length @ SB



- Smoother current decay and more bunch length damping is observed with respect to the model prediction
- Using the emittance evolution from the data (mean) → Very good prediction for the current and bunch length evolution
- Identify and add other sources of emittance evolution to the model is very important!
- The emittance blow up results in an **integrated luminosity** Evian 2016 Evian feduction of the order of 20%

Optimal Fill times for 2016

- 2016 proposed parameters:
 - $\beta^*=40$ cm in IP1/5
 - 410 μrad (11 σ) or 370
 μrad (10 σ) in IR1 and 5
 - Similar bunch brightness and bunch length as in 2015 (1.2e11, 3µm, 1.3ns)
- Most probable turnaround time (based on 2015) of 6-8h (see M. Solfaroli) Evian 201:
- Ilsing different emittance



Summary and Outlook

- A model including IBS, SR and Burn-off at Flat Top (4TeV, 6.5TeV and 7TeV) and Flat Bottom energy is ready
 - Can be easily applied bunch-by-bunch

- The model is based on analytical formulas which assume Gaussian distributions
 - This is not always the case for the LHC (especially in the longitudinal plane)
 - → Work in progress to understand the effect of the beam distribution on the IBS evolution of the bunch characteristics (S. Papadopoulou) Evian 2015

Summary and Outlook

- Observations from RunII data:
 - Differences have been observed on the emittance evolution from the different methods of measurement → Need to be understood
 - Using the emittance from the data, good prediction for the bunch length and bunch current evolution
 - Modeling the emittance evolution is a very important component of the model
- Based on the model and the observations, in 2016 long Fills should still be favorable
- Bunch-by-bunch analysis is in progress
 - Aims to identify correlations between the observed emittance blow-up and the bunch "lifestyle"

THANK YOU FOR YOUR ATTENTION!