

# TGCs sensitivities at 100 TeV from WW

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QCD, EW and tools at 100 TeV  
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# TGCs vs EFT approach

## EFT approach

$$\mathcal{L} = \mathcal{L}_{SM} + \sum_i \frac{c_i}{\Lambda^2} \mathcal{O}_i + \dots$$

$$V = \gamma, Z; W_{\mu\nu}^{\pm} = \partial_{\mu} W_{\nu}^{\pm} - \partial_{\nu} W_{\mu}^{\pm}, V_{\mu\nu} = \partial_{\mu} V_{\nu} - \partial_{\nu} V_{\mu}$$

$$\mathcal{O}_{WWW} = \text{Tr}[W_{\mu\nu} W^{\nu\rho} W_{\rho}^{\mu}]$$

$$\mathcal{O}_W = (D_{\mu}\Phi)^{\dagger} W^{\mu\nu} (D_{\nu}\Phi)$$

$$\mathcal{O}_B = (D_{\mu}\Phi)^{\dagger} B^{\mu\nu} (D_{\nu}\Phi)$$

## Anomalous TGCs

$$\mathcal{L} = ig_{WWW} \left( g_1^V (W_{\mu\nu}^+ W^{-\mu} - W^{+\mu} W_{\mu\nu}^-) V^{\nu} + \kappa_V W_{\mu}^+ W_{\nu}^- V^{\mu\nu} + \frac{\lambda_V}{M_W^2} W_{\mu}^{\nu+} W_{\nu}^{-\rho} V_{\rho}^{\mu} \right)$$

Can be related via

$$g_1^Z = 1 + c_W \frac{m_Z^2}{2\Lambda^2}$$

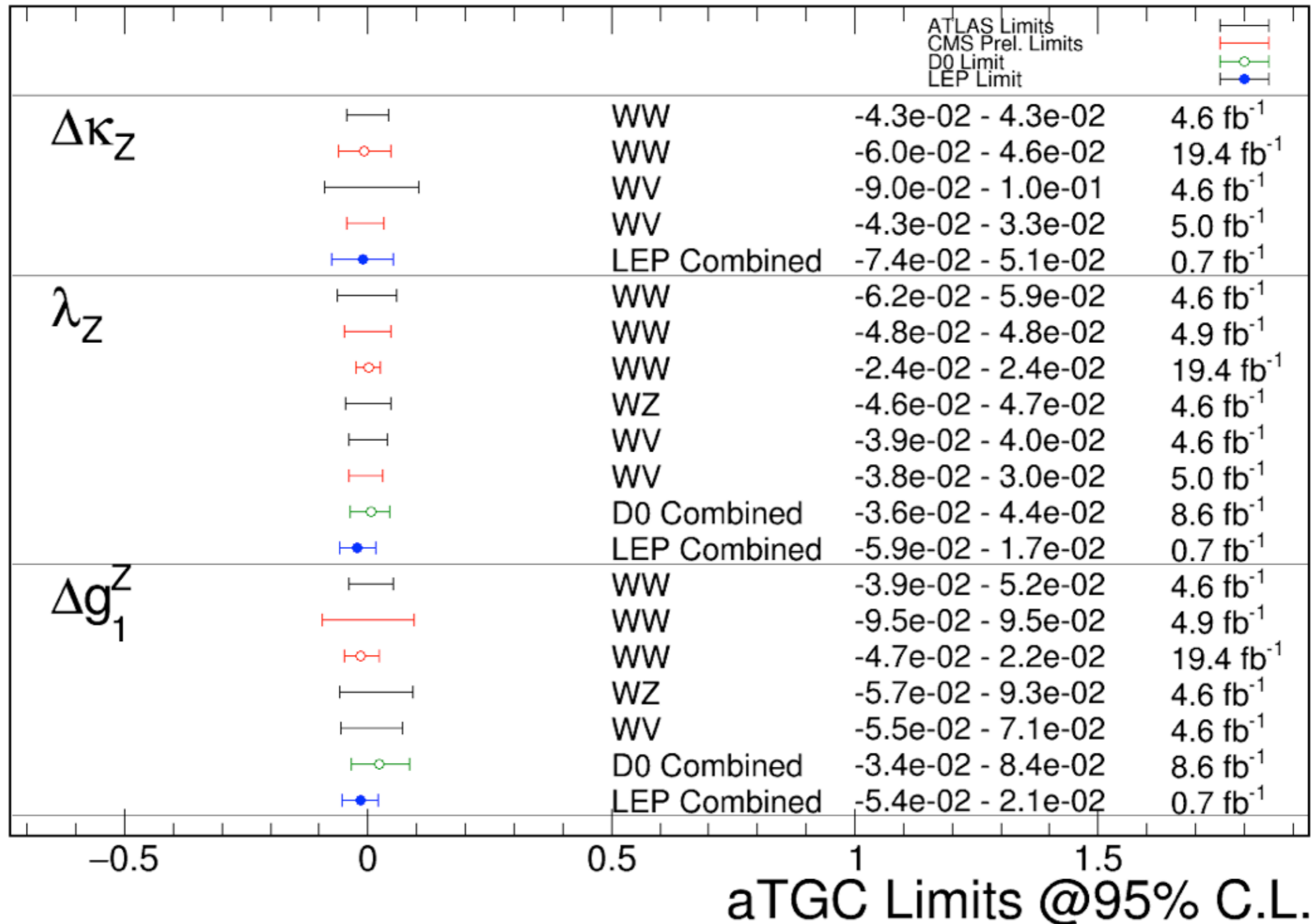
$$\kappa_{\gamma} = 1 + (c_W + c_B) \frac{m_W^2}{2\Lambda^2}$$

$$\kappa_Z = 1 + (c_W - c_B \tan^2 \theta_W) \frac{m_W^2}{2\Lambda^2}$$

$$\lambda_{\gamma} = \lambda_Z = c_{WWW} \frac{3g^2 m_W^2}{2\Lambda^2}$$

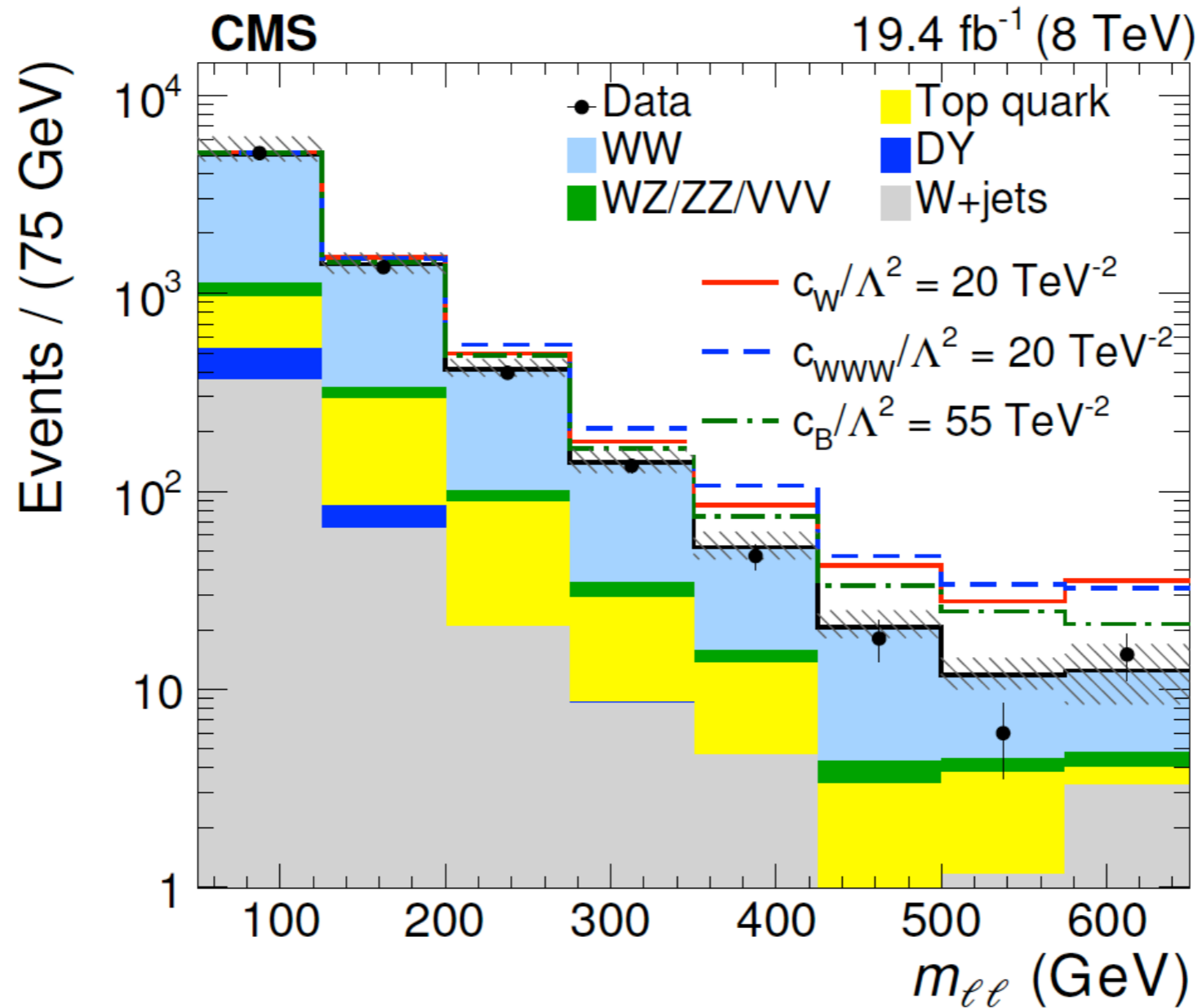
# Current bounds

July 2015



# Current CMS bounds from WW

CMS 1507.03268



The corresponding 95% confidence level intervals are  $-5.7 < c_{WWW}/\Lambda^2 < 5.9 \text{ TeV}^{-2}$ ,  $-11.4 < c_W/\Lambda^2 < 5.4 \text{ TeV}^{-2}$ ,  $-29.2 < c_B/\Lambda^2 < 23.9 \text{ TeV}^{-2}$ , in the HISZ basis.

# Our very preliminary 100 TeV study

Look at  $pp \rightarrow WW$  at NLO (LHE event level) with in the following setup

Lepton cuts:  $p_{t,l} > 20\text{GeV}$   $|\eta_l| < 2.5$

Jet veto (to reduce top background):  $p_{t,j} < 30\text{GeV}$

PDFs: NNPDF30\_nlo\_as\_0118

Start with three cases

$$\frac{c_W}{\Lambda^2} = 0.1 \text{TeV}^{-2}$$

$$g_1^Z \sim 4 \cdot 10^{-4}$$

$$\kappa_\gamma = \kappa_Z \sim 3 \cdot 10^{-4}$$

$$\frac{c_B}{\Lambda^2} = 0.3 \text{TeV}^{-2}$$

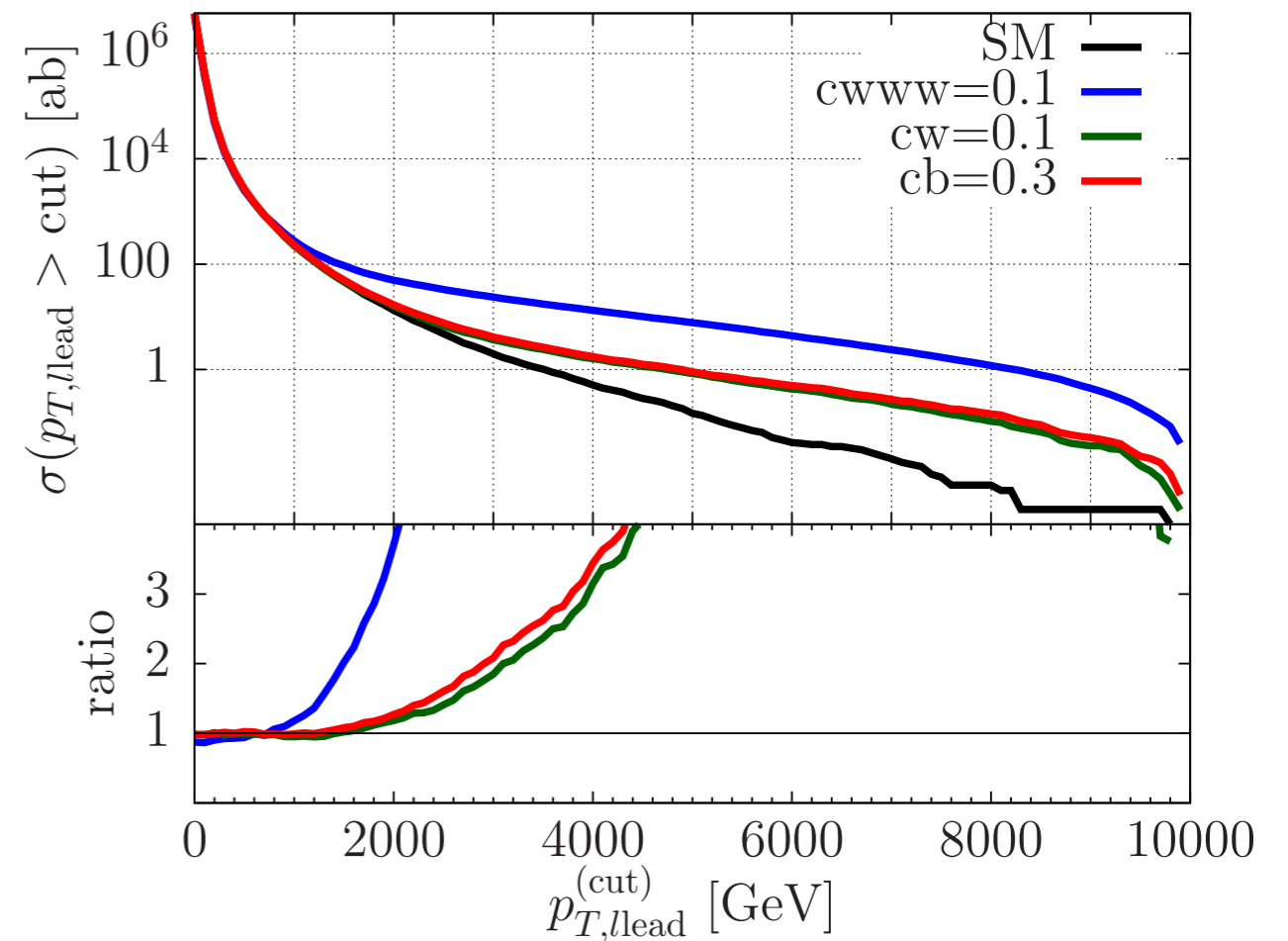
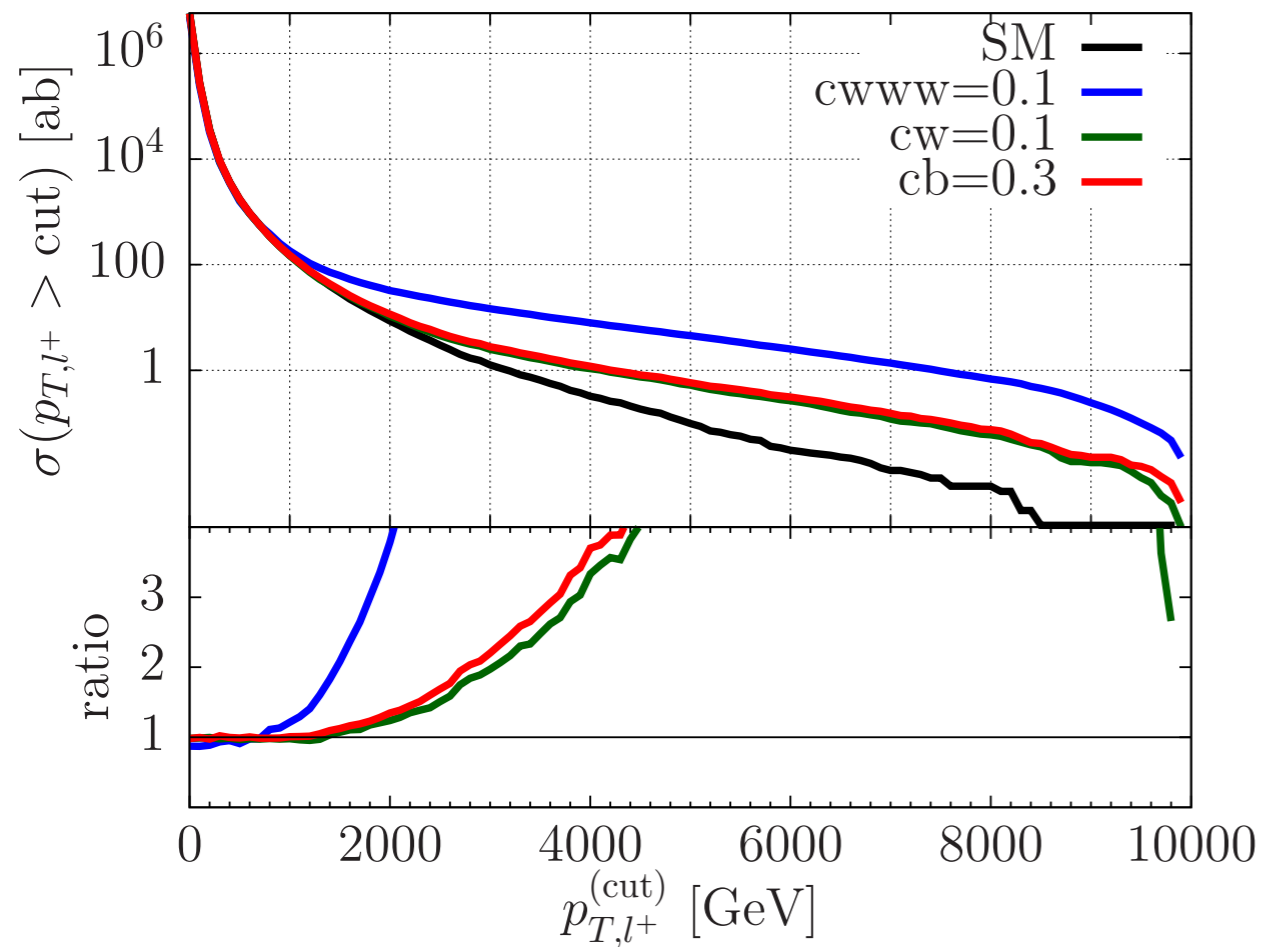
$$\kappa_\gamma \sim 9 \cdot 10^{-4}$$

$$\kappa_Z \sim -3 \cdot 10^{-4}$$

$$\frac{c_{WWW}}{\Lambda^2} = 0.1 \text{TeV}^{-2}$$

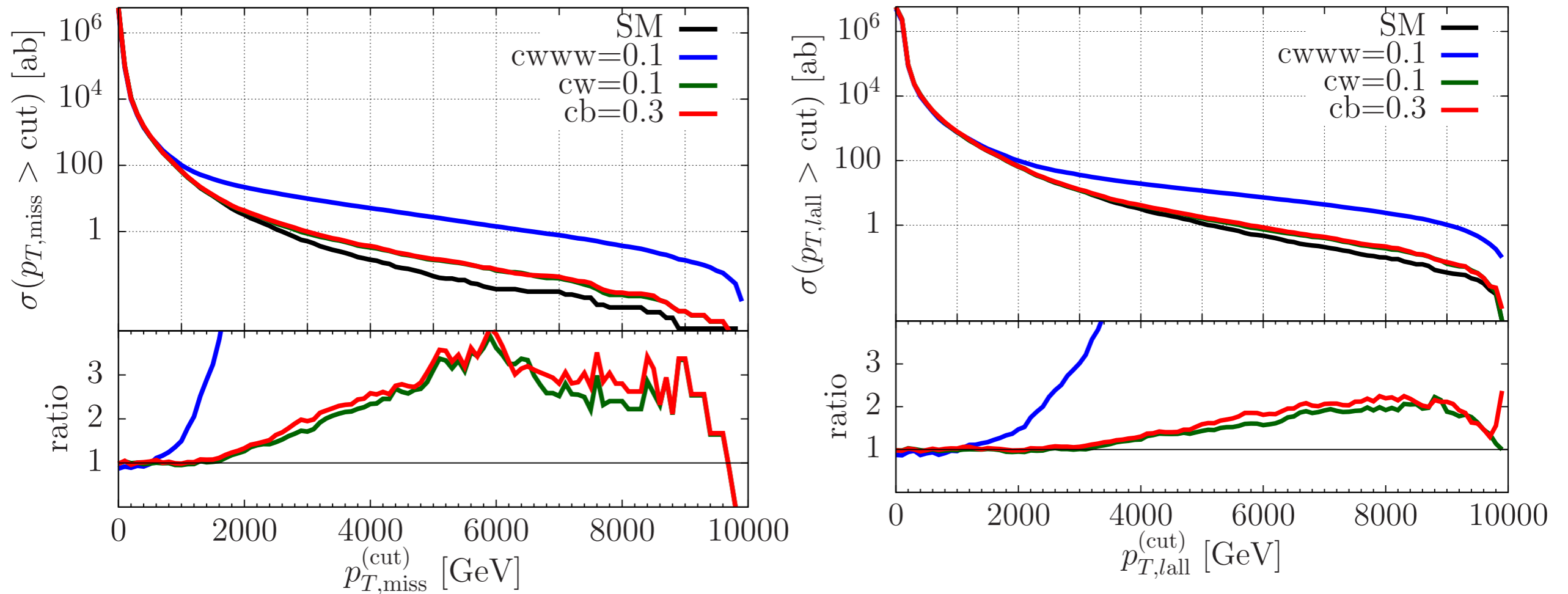
$$\lambda_\gamma = \lambda_Z \sim 4 \cdot 10^{-4}$$

# Integrated distributions



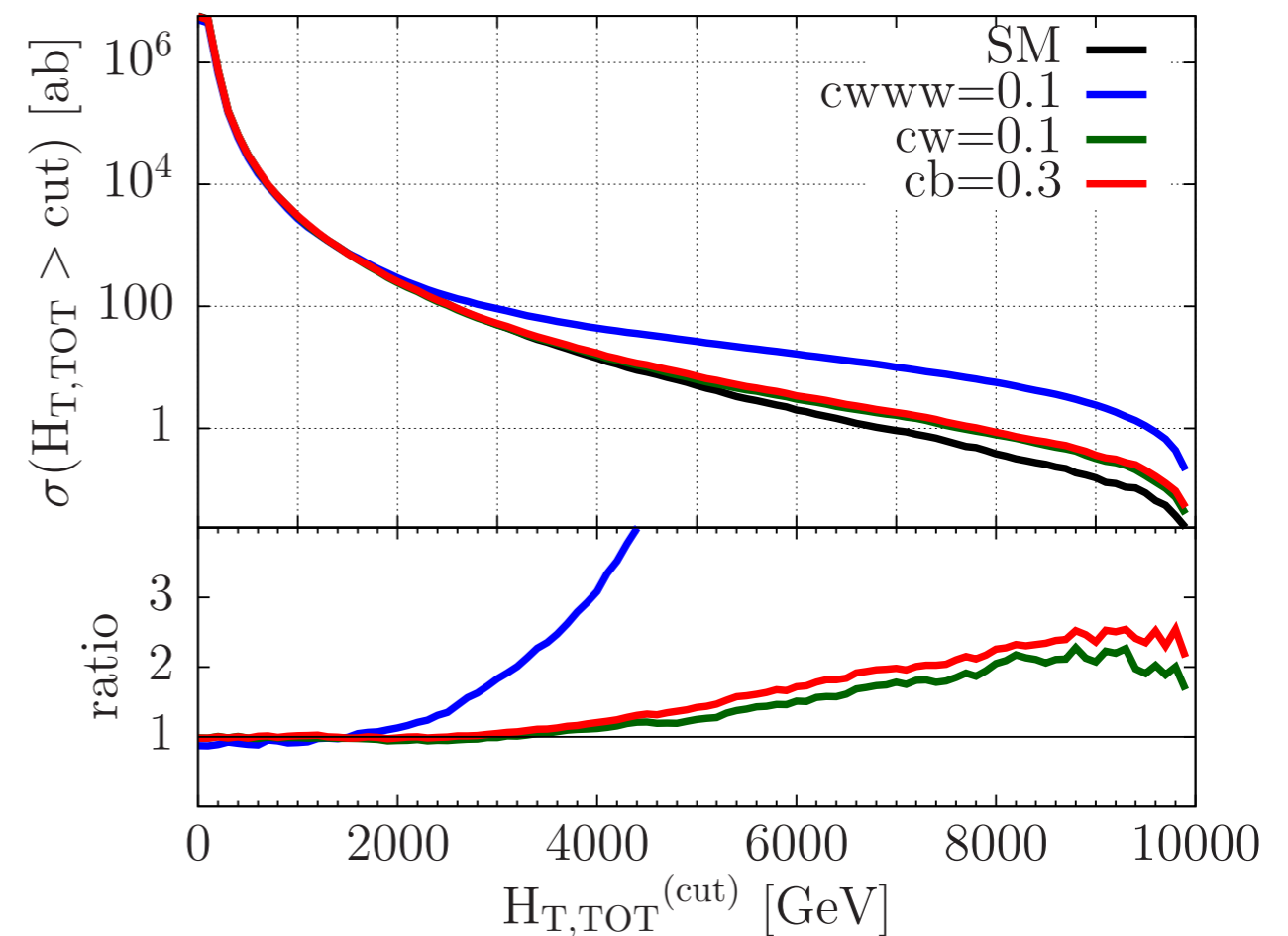
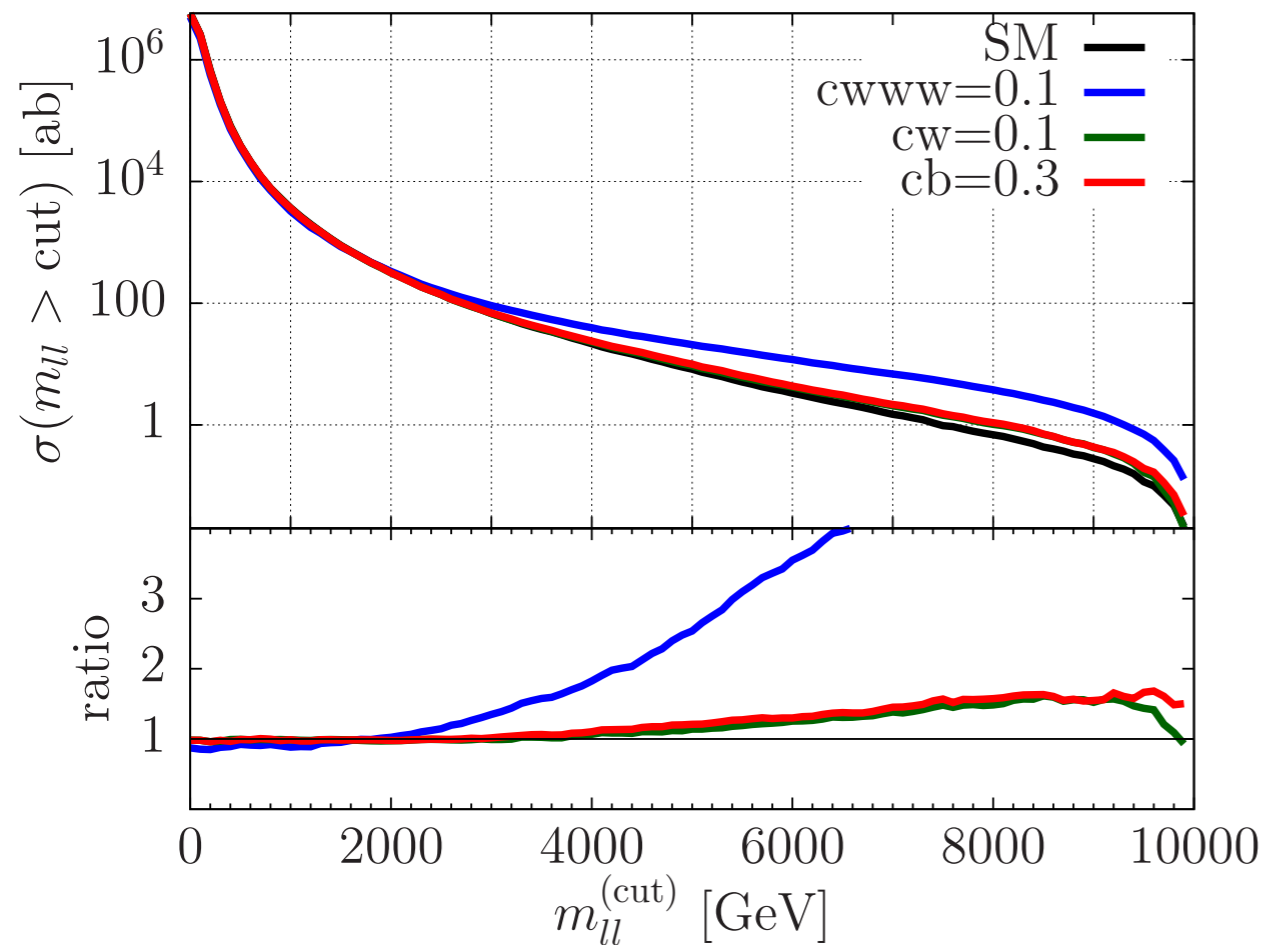
Look at integrated rates. Shown number of events assuming one  $\text{ab}^{-1}$  of integrated luminosity

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# Prospects

- Look at accessible limit for  $c_{WWW}$ . What is the lowest value for which you see one event per  $\text{ab}^{-1}$ ?
- Is our jet-veto sufficient to get rid of backgrounds?
- For  $O_W$  and  $O_B$ , does the dominant contribution come from terms with derivatives of the Higgs field?
- Look at other processes ( $WZ, W\gamma, \dots$ ). Any volunteer?

# Other benchmark results

hep-ph/0204087

Coupling	14 TeV 100 fb <sup>-1</sup>	14 TeV 1000 fb <sup>-1</sup>	28 TeV 100 fb <sup>-1</sup>	28 TeV 1000 fb <sup>-1</sup>	LC 500 fb <sup>-1</sup> , 500 GeV
$\lambda_\gamma$	0.0014	0.0006	0.0008	0.0002	0.0014
$\lambda_Z$	0.0028	0.0018	0.0023	0.009	0.0013
$\Delta\kappa_\gamma$	0.034	0.020	0.027	0.013	0.0010
$\Delta\kappa_Z$	0.040	0.034	0.036	0.013	0.0016
$g_1^Z$	0.0038	0.0024	0.0023	0.0007	0.0050