#### FCC WEEK 2016, Rome, April 11-15, 2016

# Italian view on HEP beyond LHC

Antonio Masiero
INFN and Univ. of Padova

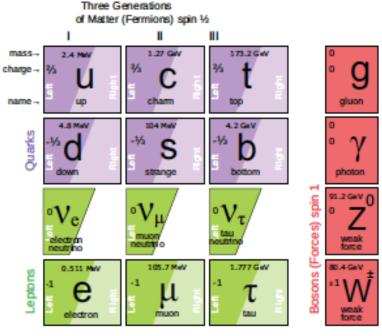
See also the talk "Italian & INFN activities towards future colliders" by A. Zoccoli on Monday

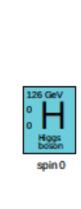
• By the end of the 20<sup>th</sup> century ... we have a comprehensive, fundamental theory of all observed forces of nature which has been tested and might be valid from the Planck length scale [10<sup>-33</sup> cm.] to the edge of the universe  $[10^{+28} \text{ cm.}]$ 

**D. Gross 2007** 

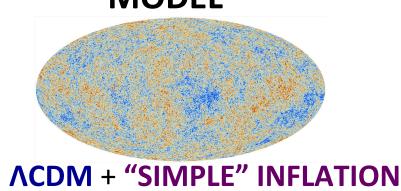
## 2013 – 2016: the triumph of the STANDARD

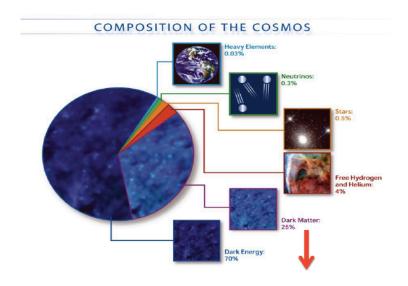
PARTICLE STANDARD
 MODEL





COSMOLOGY STANDARD MODEL



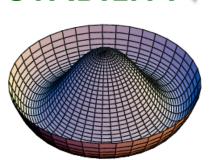


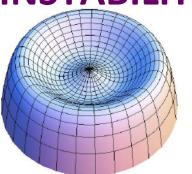
## PARTICLE STANDARD MODEL



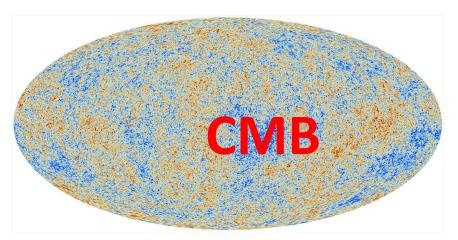
The Higgs boson and the destiny of the Universe

**STABILITY** INSTABILITY

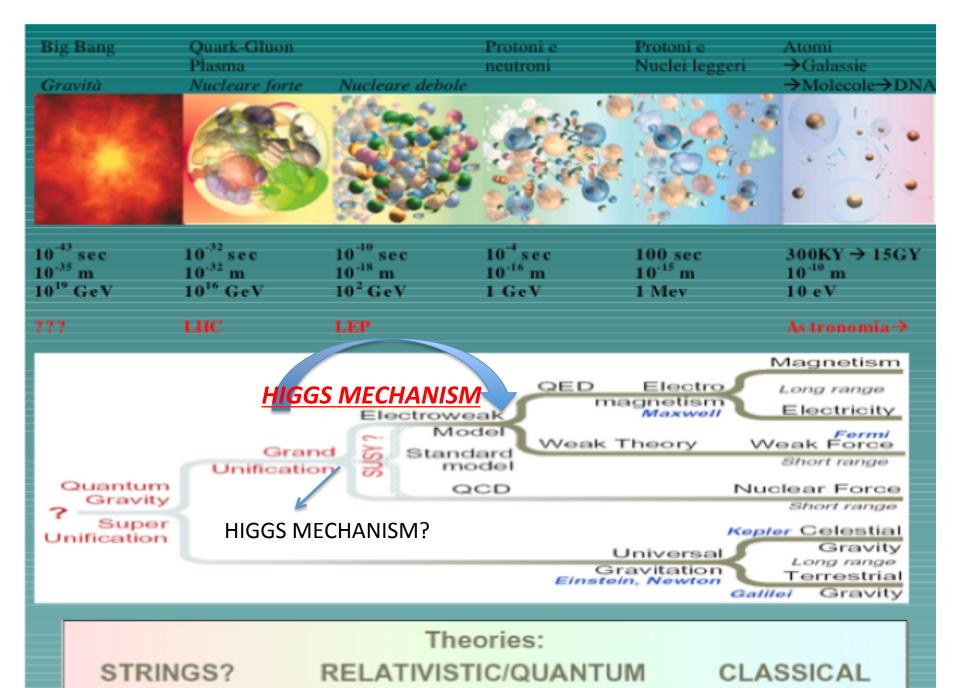




## COSMOLOGY STANDARD MODEL







## Are the SMs really STANDARD? G-W-S SM ACDM SM

- All the experimental results of both high-energy particle physics and high-intensity flavor physics are surprisingly (and embarrassingly ) in very good agreement with the predictions of the GSW SM
- All the cosmic observations are in agreement with the ~25% CDM, ~70% cosmological constant Λ, ~5% ordinary matter of the ΛCDM SM

- Only (possible) exceptions:
  - -- the anomalous magnetic moment of the muon (3.6 σ discrepancy w.r.t. the SM prediction);

 (Possible) exception: troubles with pure Cold DM from absence proto-galaxies, nonexistence of spikes in DM density at the centre of the galaxies

. .

diphoton peak at 750 GeV

#### THE FLAVOUR PROBLEMS

#### **FERMION MASSES**

What is the rationale hiding behind the spectrum of fermion masses and mixing angles (our "Balmer lines" problem)

### LACK OF A FLAVOUR "THEORY"

( new flavour – horizontal symmetry, radiatively induced lighter fermion masses, dynamical or geometrical determination of the Yukawa couplings, ...?)

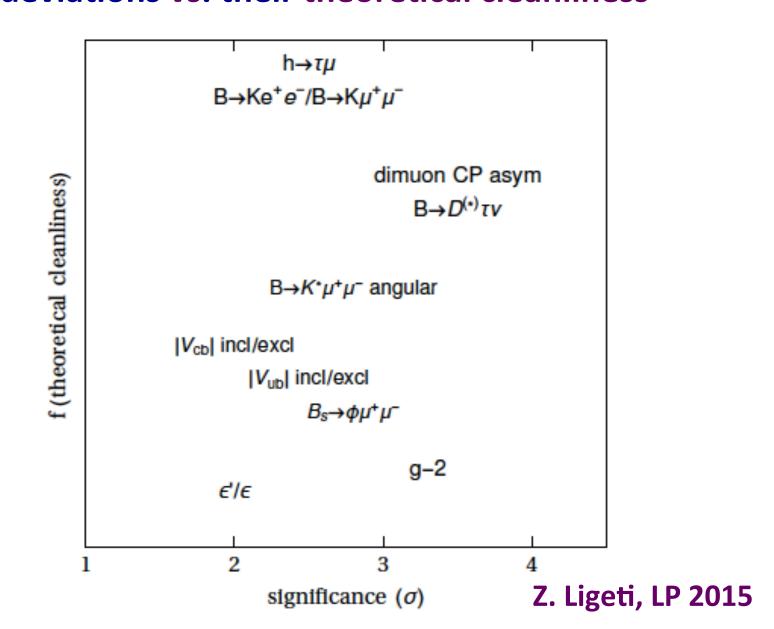
#### **FCNC**

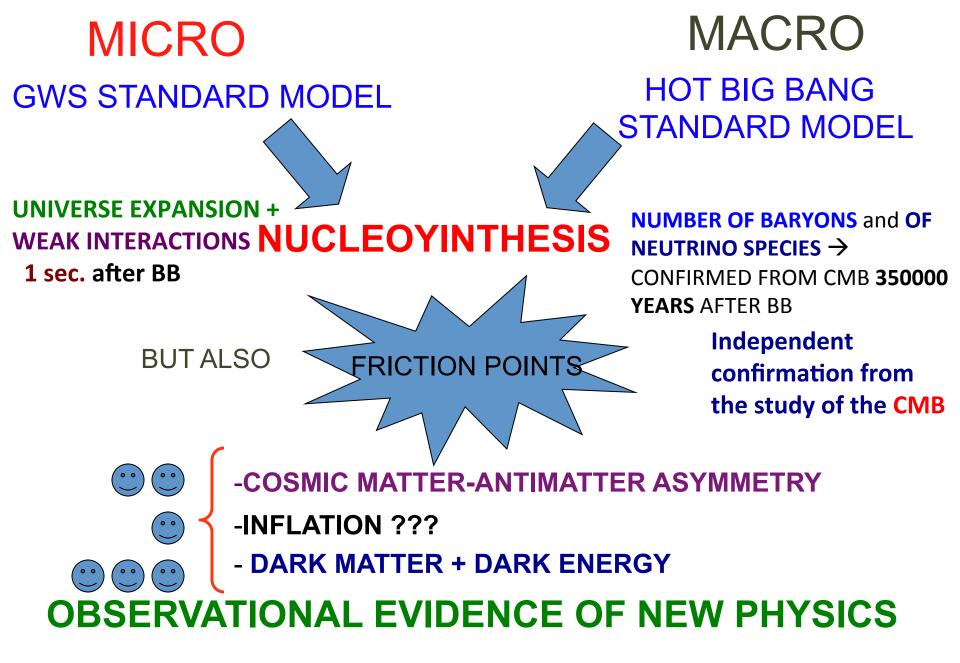
Flavour changing neutral current (FCNC) processes are suppressed.

In the SM two nice mechanisms are at work: the GIM mechanism and the structure of the CKM mixing matrix.

How to cope with such delicate suppression if the there is new physics at the electroweak scale?

## Deviations from the SM expectations: significance of such deviations vs. their theoretical cleanliness



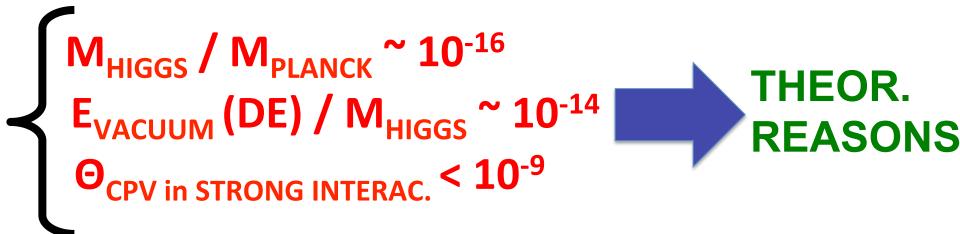


**BEYOND THE STANDARD** 

#### What the SM does not account for...

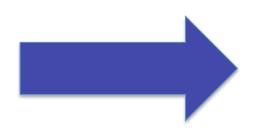
neutrino masses
dark matter
baryogenesis
inflation

OBSERVATION
AL REASONS



## The Energy Scale from the "Observational" New Physics

neutrino masses dark matter baryogenesis inflation



NO NEED FOR THE NP SCALE TO BE CLOSE TO THE ELW. SCALE

The Energy Scale from the "Theoretical" New Physics

 $\star$   $\star$  Stabilization of the electroweak symmetry breaking at M<sub>W</sub> calls for an ULTRAVIOLET COMPLETION of the SM

already at the TeV scale



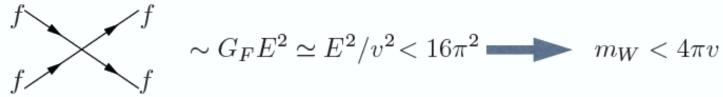


CORRECT GRAND UNIFICATION "CALLS" FOR NEW PARTICLES

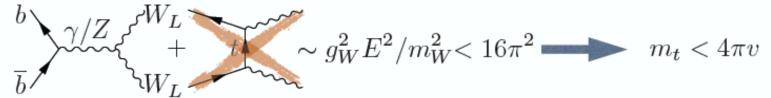
#### No-Lose Theorems

#### A number of guaranteed discoveries in the history of HEP

Beyond the Fermi Theory:



Beyond the Bottom Quark:



Beyond the (Higgsless) EW Theory:

$$W_L \sim W_L + \dots \sim g_W^2 E^2/m_W^2 < 16\pi^2 \longrightarrow m_H < 4\pi v$$

$$W_L \sim W_L$$

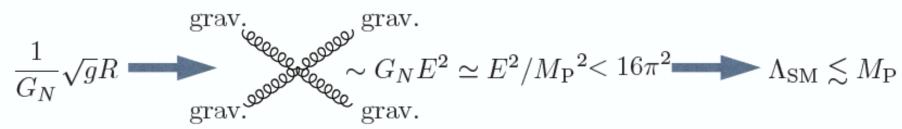
Each (secretly) due to d=6 non-renormalizable operators, signalling nearby new physics.

A. Wulzer 2015

#### No-Lose Theorems

#### A. Wulzer

Only one d>4 is left after Higgs discovery ...



... the last, impractical, No-Lose Theorem is Q.G. at  $M_P$ !

We do have exp. evidences of BSM, but none necessarily pointing to light/strongly-coupled enough new physics:

"No guaranteed discoveries" = "post-Higgs depression"

However, one d<4 comes with the Higgs discovery:

$$\frac{m_H^2}{2} H^\dagger H$$
  $\longrightarrow$  The Naturalness Problem: Why  $m_H \ll \Lambda_{\rm SM}$ ?

#### THE "COMPREHENSION" OF THE ELECTROWEAK SCALE

$$m_H^2 \sim -2\mu^2 + \frac{g^2}{(4\pi)^2} M^2 \quad {\rm Only \ for \ scalars; \ sm \ fermions \ and \ \ gauge \ boson \ masses \ are \ protected \ by \ the \ su(2) \times u(1) \ symmetry \ !}$$

To comprehend (i.e. stabilize) the elw. scale need **NEW PHYSICS (NP) to be operative at a scale** 



#### Naturalness or

 New SYMMETRY giving rise to a cut-off at



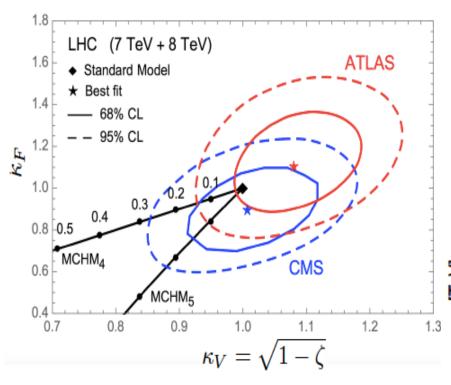
Low-energy **SuperSymmetry** 

- Space-time modification (extra-dim., warped space)
- COMPOSITE HIGGS: the Higgs is a pseudo-Goldstone boson (pion-like) → new interaction getting strong at

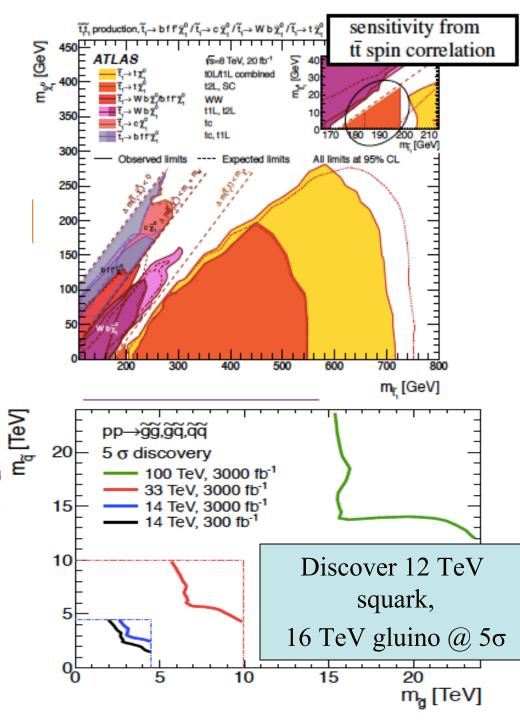
#### **Un-naturalness?**

- The scale at which the electroweak symmetry is spontaneously broken by <H> results from COSMOLOGICAL EVOLUTION
- H is a fundamental (elementary) particle > we live in a universe where the fine-tuning at M arises (anthropic solution, multiverse, Landscape of string theory)

## Higgs boson: elementary or composite?



Current bound  $\zeta < 0.12 \rightarrow$  already some tuning on the **composite** models to look like SM

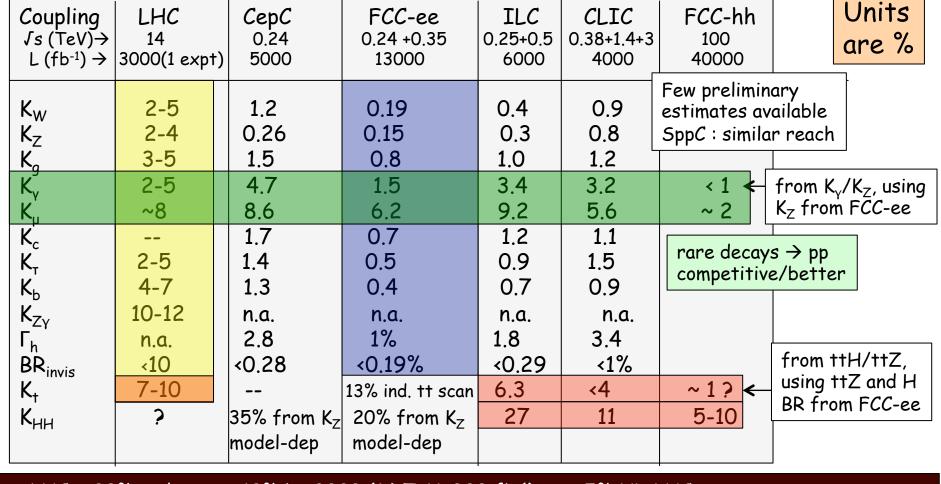


## What remains to be learnt on the SM from LHC and future accelerators

Higgs boson couplings to bosons and fermions: precisions
 ≤ 10% attainable with 300 fb<sup>-1</sup>;
 precisions 2% - 5% in the High Luminosity phase
 uncertainties O(1%) at ILC and <1% at FCC-ee</li>

 Higgs total width: too narrow (~4 MeV) to be measured at LHC – at HL-LHC try using the interference of a specific mode with the continuum; at ILC/FCC-ee through HZ

Higgs boson rare production and rare decay modes: HH
production important → related to Higgs self-couplings →
need full HL-LHC phase



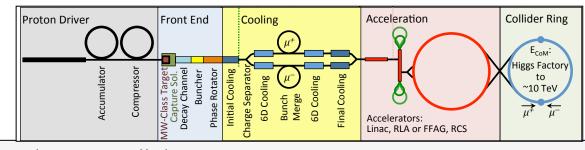
- □ LHC: ~20% today  $\rightarrow$  ~ 10% by 2023 (14 TeV, 300 fb<sup>-1</sup>)  $\rightarrow$  ~ 5% HL-LHC
- $\Box$  HL-LHC: -- first direct observation of couplings to 2<sup>nd</sup> generation (H $\rightarrow$  µµ)
- -- model-independent ratios of couplings to 2-5%
- where high energy needed  $\rightarrow$  linear colliders, high-E pp colliders

 $\square$  Best precision (few 0.1%) at FCC-ee (luminosity!), except for heavy states (ttH and HH)

□ Complementarity/synergies between ee and pp F. Gianotti, EPS '15

Theory uncertainties (presently few percent e.g. on BR) need to be improved to match expected superb experimental precision

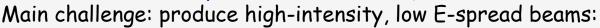
#### Muon colliders



Synergies with neutrino factories

Main advantage compared to e<sup>+</sup>e<sup>-</sup> colliders: m<sub>u</sub> ~ 200 m<sub>e</sub>

- $\rightarrow$  negligible SR  $\rightarrow$  can reach multi-TeV with (compact!) circular colliders:
- 300 m ring for  $\sqrt{s}$  = 125 GeV, 4.5 km for  $\sqrt{s}$  = 3 TeV
- → negligible beamstrahlung → much smaller E spread
- $\rightarrow$   $\sigma$  ( $\mu\mu \rightarrow H$ ) ~ 20 pb (s-channel resonant production)  $\rightarrow$  H factory

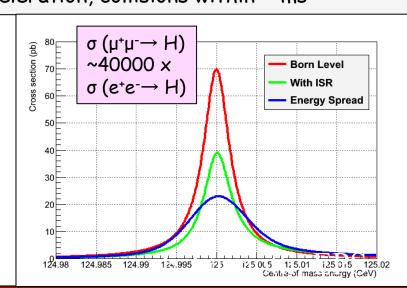


- $\square$  m<sub>u</sub> ~ 200 m<sub>e</sub>  $\rightarrow$  SR damping does not work  $\rightarrow$  novel cooling methods (dE/dx based) needed to reach beam energy spread of  $\sim 3\times10^{-5}$  (for precise line shape studies) and high L
- $\Box$   $\tau_u \sim 2.2 \ \mu s \rightarrow$  production, collection, cooling, acceleration, collisions within  $\sim$  ms

Beam spread of  $\sim 3 \times 10^{-5}$  would allow  $\Gamma_{H}$ measurement from line shape to 5% (0.2 MeV) → resolve (possible) resonances

However, with currently projected L (~10<sup>32</sup>): ~ 20000 H/year → not competitive with ete colliders for coupling measurements (except Hμμ ~ 1%)

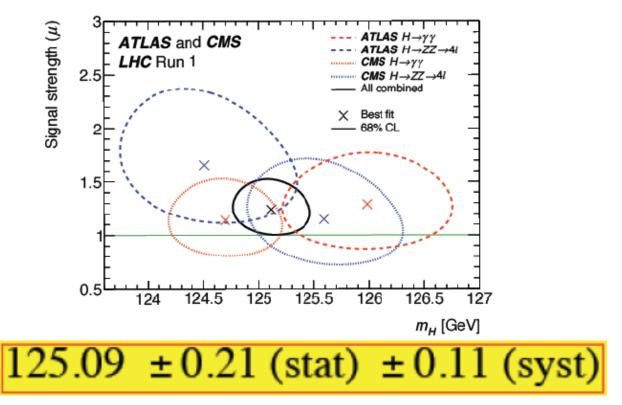




More R&D needed to demonstrated feasibility, in particular cooling:

linear systems (MICF at RAL) rings (recently re-ignited by C. Rubbi

## Higgs Mass measurements ATLAS + CMS ZZ\* and γγ final states

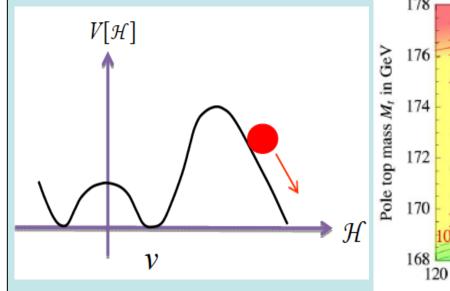


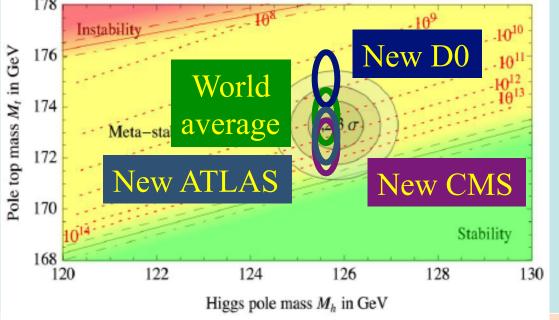
The values of the TOP and HIGGS masses are crucial to establish the stability of the

**ELECTROWEAK VACUUM** 

#### Vacuum Instability in the Standard Model

• Very sensitive to m<sub>t</sub> as well as M<sub>H</sub>





J. Ellis, LP 2015

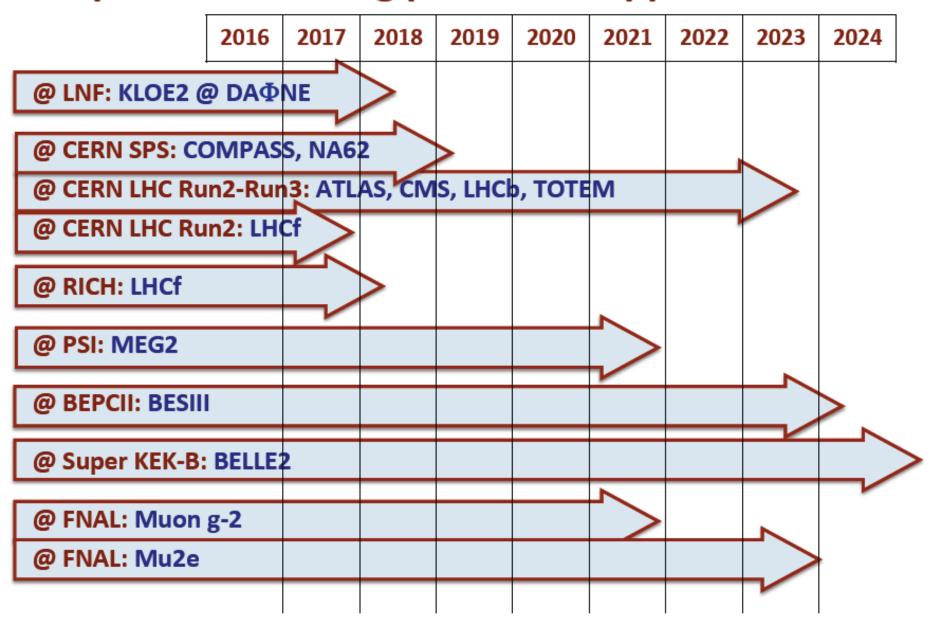
• Instability scale.

Buttazzo, Degrassi, Giardino, Giudice, Sala, Salvio & Strumia, arXiv:1307.3536

$$\log_{10} \frac{\Lambda_I}{\text{GeV}} = 11.3 + 1.0 \left( \frac{M_h}{\text{GeV}} - 125.66 \right) - 1.2 \left( \frac{M_t}{\text{GeV}} - 173.10 \right) + 0.4 \frac{\alpha_3(M_Z) - 0.1184}{0.0007}$$

 $m_t = 173.3 \pm 1.0 \text{ GeV} \rightarrow \log_{10}(\Lambda/\text{GeV}) = 11.1 \pm 1.3$ 

#### Complete data taking plans with approved detectors



#### Full exploitation of our ability of 'understanding' the SM

- Still relevant investigations on the SM importance of the SM precision physics for
  - i) a better and deeper understanding of the SM,
  - ii) indirect hints of new physics

 'Sociological' remark: it may be difficult to attract and keep young people with the horizon of performing precious SM precision physics

#### much depends on the next 5 years ...

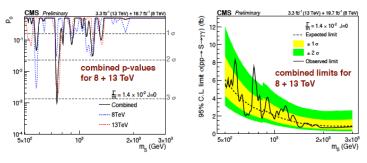
- LHC14 (high energy: ATLAS, CMS; flavor: LHCb; quark-hadron phase transition: ALICE)
- Flavor: NA62; upgraded MEG, Mu-e; BELLEII; EDMs; g-2
- **DM** 1-ton exps.  $\rightarrow$  10<sup>-10</sup> 10<sup>-11</sup> pb
- Neutrinoless double  $\beta \rightarrow v$  mass degenerate region; enter IH region
- SBN  $\rightarrow$  sterile v?
- Gravitational waves → discovery to pave the way to gravitational wave astronomy
- DE: BOSS → DESI; DES → LSST
- CMB: final PLANCK; B-modes of the polariz.+ black-body spectrum: EU exps. QUBIC, LSPE, QIJOTE + many others on ground and balloons in US, Japan



#### DIPHOTON RESONANCES



- Combined 8 TeV + 13 TeV results
  - Largest excess is observed for 750 GeV, spin-0, narrow width
    - local significance of 3.4σ, 1.6σ after look-elsewhere effect

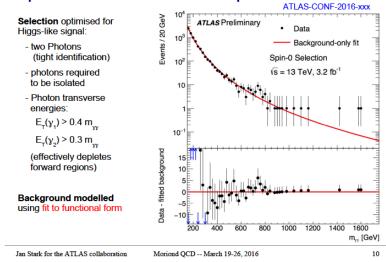


- Dec '15 result: largest excess at 760 GeV for Γ/M=1.4x10<sup>-2</sup>
  - local significance of  $\sim 3\sigma$ ,  $< 1.7\sigma$  after look-elsewhere effect

New Physics with light SM particles at CMS – JPC – Rutgers University – Sunday, March 20th, 2016

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#### Di-photons: search for spin-0 resonance



or some other NP signal shows up at LHC14, then, obviously, "jump" into such NP area  $\rightarrow$  for instance if the Diphoton peak proves to be real, then doubling LHC energy may become an attractive option

ATLAS

#### if...

 No new NP signal shows up at LHC14 and at the new generation astroparticle projects → vigorous study of the various post-LHC options, while

#### WHAT NEXT

In view of the complex landscape
we have to confront, INFN has recently
started a process to identify the most
important research themes that we
should focus on amongst those that
in this moment do not receive enough
attention (people, funding).
Astroparticle physics is a sector that
is at the center of attention.



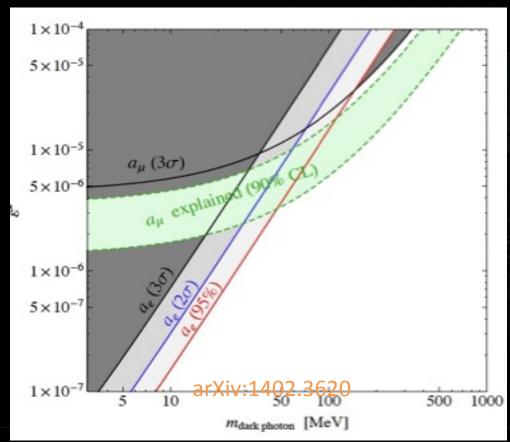
HIGH ENERGY, HIGHINTENSITY,
ASTROPARTICLE
PHYSICS →
COMPLEMENTARY
ATTACK TO THE NEW
PHYSICS FORTRESS



#### **PADME at Frascati**

PADME aims to detect U boson produced in e<sup>+</sup>e<sup>-</sup> annihilation and decaying into invisibles

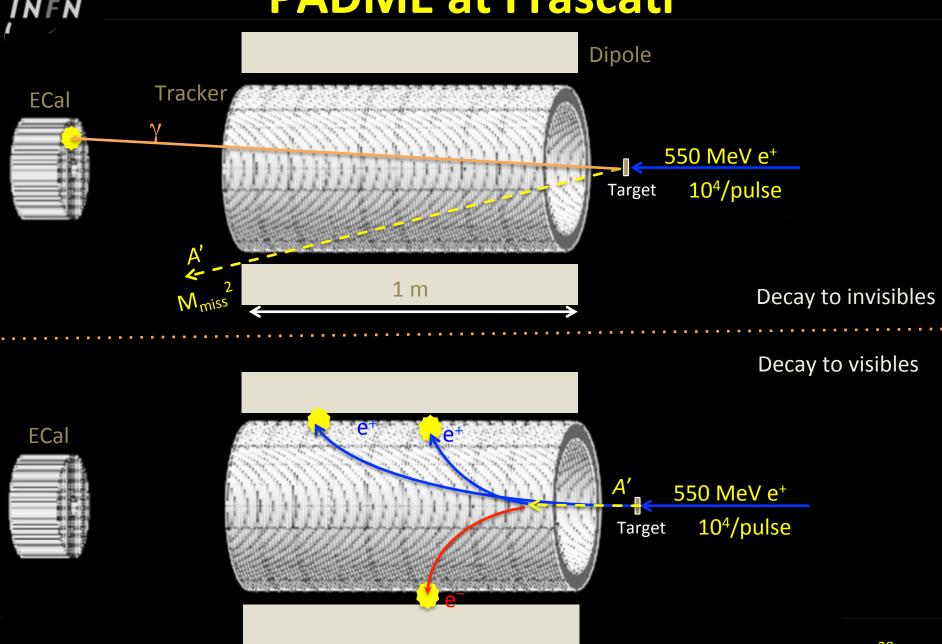
- No assumption on the U decays products and coupling to quarks
- Only minimal assumption: U bosons couples to leptons
  No experimental results yet with this approach (VEPP3 planned experiment)



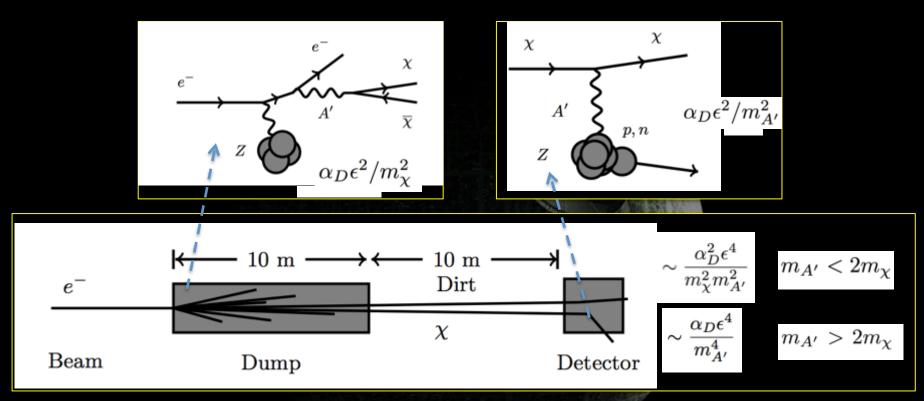


# **ECal**

#### **PADME at Frascati**



## New ideas for dump experiment: BDX at JLAB



Scintillator 1 m<sup>3</sup>
1 MeV/10 MeV e<sup>+</sup>e<sup>-</sup> detection threshold

#### Backgrounds:

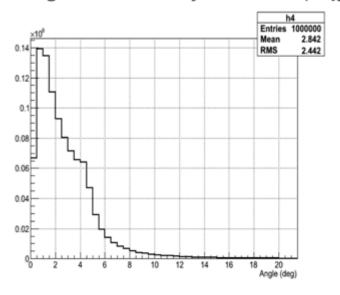
- Neutrino production
- Cosmogenic muons and neutrons



#### **BDX at Frascati?**

#### χ production and detection

- 1.5 GeV electron beam
- 7 10<sup>19</sup> EOT/year
- 1 year run (50% efficiency)
- Repetition rate: 50 Hz, (0.7A in 10 ns bunch)
- · Negligible cosmogenic BG with timing cut
- Expected ~20 counts in 1m³ plastic scintillator detector (1 MeVee threshold)
- Significant sensitivity to low mass (A'/ $\chi$ ) region



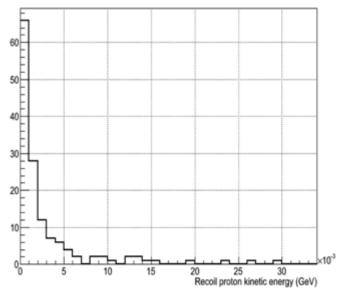
#### **Parameters:**

M A' = 50 MeV

M\_Chi = 10 MeV

Alpha\_dark = 0.1

Epsilon =  $10^-3$ 



Very preliminary study. Results look very promising and should be investigated further.

quale minima estensione del SM posso fare per "curare" le sue tre malattie "osservative":dar massa ai neutrini, avere materia oscura e produrre una asimmetria materia-antimateria Modello di Shaposhnikov et al. con 3 neutrini right-handed "leggeri"

Role of  $N_1$  with mass in keV region: dark matter

inflate the Universe.

Role of  $N_2$ ,  $N_3$  with mass in 100 MeV – GeV region: "give" masses to neutrinos and produce baryon asymmetry of the Universe Role of the Higgs: give masses to quarks, leptons, Z and W and

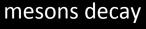
Role of scale invariance and unimodular gravity: dilaton gives mass to

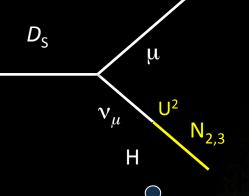
the Higgs and  $N_{1,0,0}$  and provides dynamical dark energy

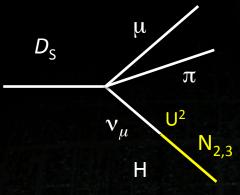


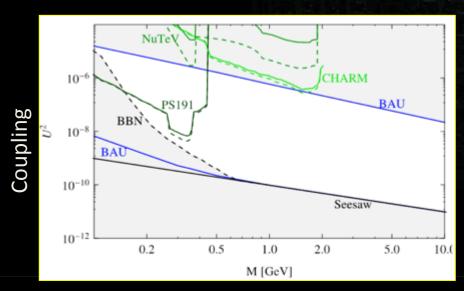
#### **SHiP at CERN SPS**

Look for HNL through the only possible interaction (Yukawa coupling to the Higgs) in the D







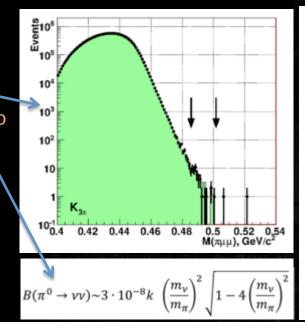


- Small couplings, long decay length (km!)
- Decay channels in  $\mu/e$  π,  $\mu/e$  ρ,  $\nu$   $\mu$  e



#### **NA62**

- Of course NA62 main goal is already searching for BSM effects
- What can be done with  $10^{13}$  K decay and  $2.5 \cdot 10^{12}$   $\pi^0$  decays?
- Search for LFV modes
- Dark photon
- Majorana neutrino ~
- Right-handed neutrino



D. Beyond  $K^+ \to \pi^+ \nu \overline{\nu}$ : Use all the rare K decays for a better identification

EW Penguin	SM and/or example of SUSY effect	Contributes to
$\vec{s}_L$ $d_L$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{split} K &\to \pi \nu \overline{\nu} \\ K_L &\to \pi^0 \ell^+ \ell^- \\ K_L &\to \ell^+ \ell^- \end{split}$
ξ <sub>γ</sub> d <sub>L</sub>	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$K_L \to \pi^0 \ell^+ \ell^-$
$\frac{\xi_H^0}{\bar{s}_L}$	$h^{0}, H^{0}, A^{0} $ $\widetilde{S}_{L,R} $ $d_{R,L} $ $M^{0}$ $\widetilde{U}_{R} $ $\widetilde{U}_{L} $ $\widetilde{U}_{R} $ $\widetilde{U}_{R} $ $\widetilde{U}_{R} $ $\widetilde{U}_{R} $	$K_L  ightarrow \pi^0 \mu^+ \mu^-$ $K_L  ightarrow \mu^+ \mu^-$ (helicity-suppressed)

NP to be identified by looking at patterns of deviations!



# A) Multimessenger astronomy,B) neutrino properties,C) dark side of the Universe and CMB

- A) Photon, cosmic ray, neutrino, gravitational astronomies (some in their maturity, some in their youth, some just baby or even at the embryonic level
- B) Neutrino mass and its relation to the global symmetry of the SM, Lepton number (Dirac vs. Majorana natur of the neutrinos); measuring the full neutrino mass parameters (neutrino mass hierarchy, CP violation)
- C) Dark Matter; Dark Energy and their role in the evolution of the Universe; CMB (primordial inflation, elw. Phase transition, quark-hadron phase transition, nucleosynthesis, matter-antimatter cosmic asymmetry) -

#### A memorable past decade for astroparticle physics...

- Multimessenger astronomy: 2 new entries, i.e. 2 new cosmic messengers are DISCOVERED, HE cosmic neutrinos and gravitational waves. Important progress in gamma- and charged cosmic ray – astronomy
- Impressive progress in our knowledge of neutrino properties through a combined action of astroparticle physics and cosmology
- CMB: extraordinary achievements by the Planck satellite on our knowledge of CMB temperature fluctuations as well as the CMB polarization modes
- The dark side of the Universe: amazing progress in our bounds especially on WIMP DM, but the DM mystery still remains. In spite of our better knowledge of some DE properties, still its nature remains completely obscure.

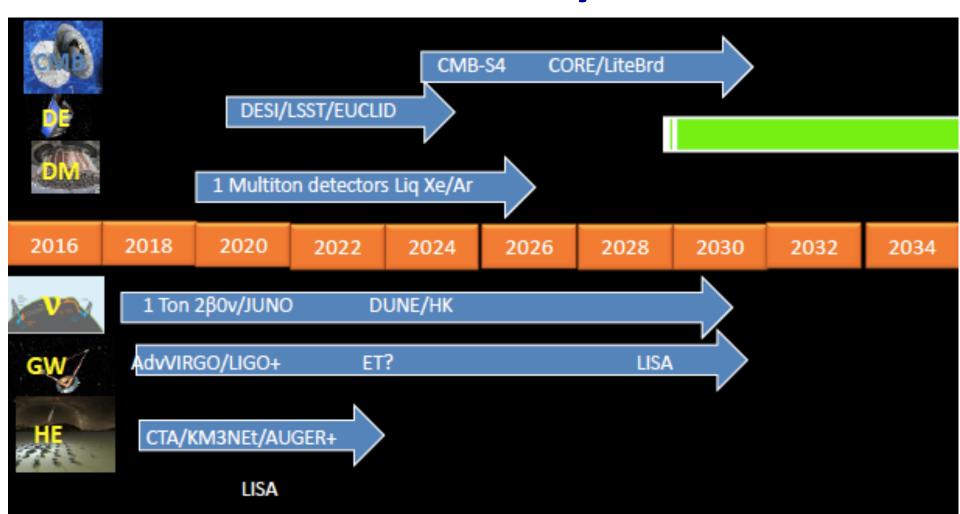
#### ... and a thrilling decade in front of us

- Multi-Messenger Astronomy (advent of the cosmic HE neutrino and gravitational waves astronomies, the CTA tremendous leap in gamma astronomy, the new horizon in charged cosmic ray astronomy with the upgrade of AUGER);
- Impressive progress in unveiling (some of) the neutrino mysteries: Dirac vs. Majorana (1-ton (ββ)<sub>0νν</sub> exps.); v mass hierarchy (the race: see fig.); v CP violation (new long baseline v exps.); v masses (direct exps., amazing input from cosmology)

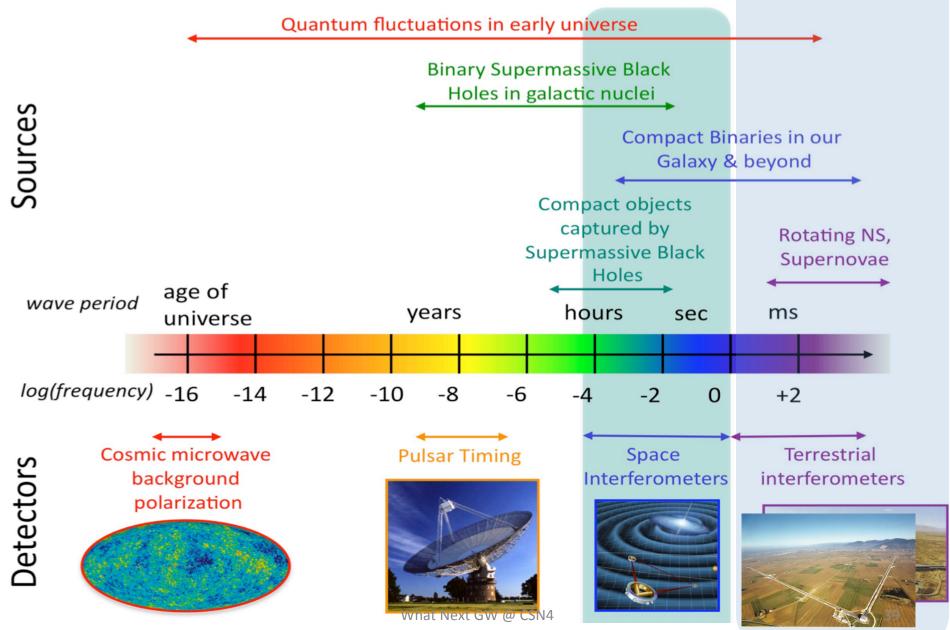
#### ... and a thrilling decade in front of us

- CMB in the post-Planck (satellite) era → tremendous progress in ground, balloon and space exps.
- Shedding (an impressive amount of) light on the dark side of the Universe: DM → multi-ton exps. towards the ultimate v background (attempting to even overcome it); DE: remarkable leap in our knowledge of the history of the expansion rate of the Universe and the rate of growth of the cosmic structures through new ground and space exps

# The magnificent 9 of astroparticle physics and cosmology in the next 20 years



#### The Gravitational Wave Spectrum



## ASTROPARTICLE PHYSICS the European strategy 2016

RNOWN UNKNOWN:

DM DE L B CP

INFLATION ...

unknown UNKNOWN: beyond QM – GR, ?



## The importance of being **SMALL**

My recommendation: beware the temptation of going ONLY for LARGE enterprises

The protective shield of large, Big Science: too big to fail!

Richness of small, "unorthodox" projects based more on clever ideas than on muscular, managerial strength!

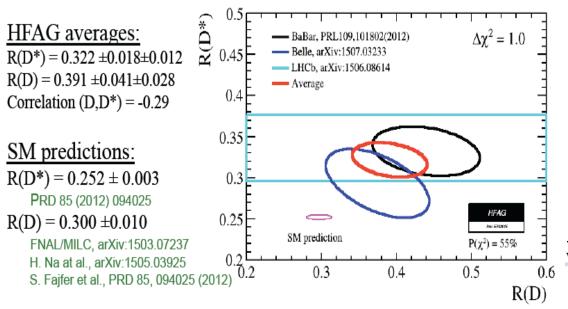
the two Standard Models are an extraordinary step forward in our knowledge of the Universe: but, beware, Nature is rich of "unknown unknown" after all Physics had already produced a "comprehensive, fundamental theory of all observed

Maybe the DM and the DE mysteries could represent the XXI century black-body and photoelectric problems

forces of nature" at the end of the XIX century...

#### Puzzling deviations: $R(D^{(*)}) = BR(\bar{B} \rightarrow D^{(*)}\tau\bar{\nu})/BR(\bar{B} \rightarrow D^{(*)}l\bar{\nu})$

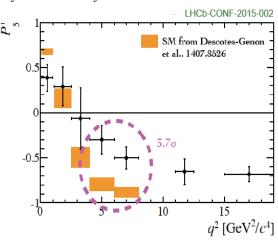
New HFAG average of  $R(D^*)$  and R(D):



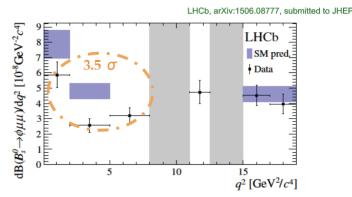
Difference with SM predictions at 3.9  $\sigma$  level.

G. Lanfranchi, LP 2015

Recently confirmed by LHCb with the full Run I dataset (3 fb<sup>-1</sup>)



..and recently also in the differential BR of  $B^0_{\ s} \to \!\! \varphi \ \mu^+\mu^-$  with full Run I dataset (3 fb-1)



SM predictions based on W. Altmannshofer and D. Straub, arXiv:1411.3161

A. Bharucha, D. Straub, R. Zwicky: arXiv:1503.05534