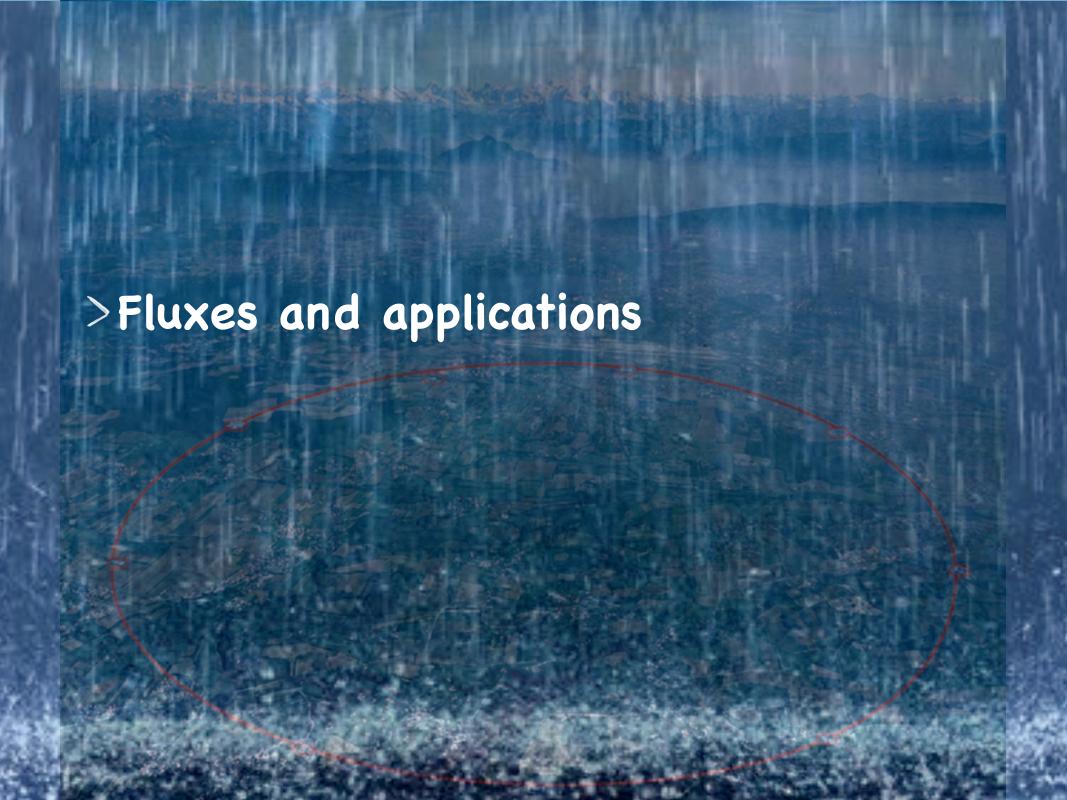
String Model Building for Particle Physics and Cosmology

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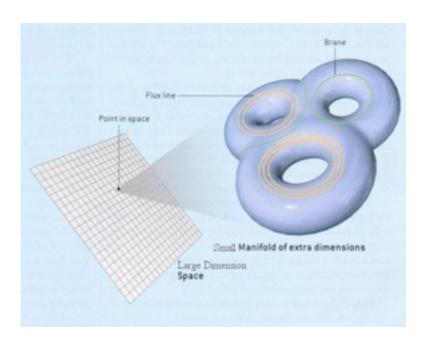
Plan

- >Flux compactifications
- > Moduli stabilization
- > Supersymmetry breaking
- >Inflation

Moduli stabilization

- Massless moduli mediate 5th forces. Unobserved
- Cosmological problems unless massive enough
- Need mechanisms for moduli stabilization

Focus on flux compactifications



Flux compactification

In addition to metric background, introduce backgrounds for NSNS and RR p-form fields

Due to gauge invariance, backgrounds for field strength

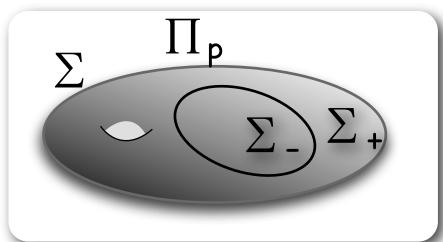
Fopological sector defined by cohomology class

In the absence of sources, $dF_{p+1}=0$

Must specify fluxes over basis of cycles

$$\int_{\Lambda_k} F_{p+1} = N_k$$

Flux quantization



Flux compactification

Fluxes introduce moduli dependence in potential energy

Closed string moduli are stabilized

Dating back to Freund-Rubin

Ex: AdS₅ \times S⁵

S⁵ volume not a modulus, but sits at a minimum of potential

Vacua need not be Minkowski, but also AdS (or dS??)

Focus on M_4 , or in (A)dS₄ with hierarchycal length scales

Prototypical example: type IIB with NSNS and RR 3-form flux.

There are no I- or 5-cycles on CY threefolds

Ind action

$$S_{ ext{IIB}} = rac{1}{2\kappa_{10}^2} \int d^{10}x \, \sqrt{-g} \left(R - rac{\partial_M \overline{ au} \, \partial^M au}{2(ext{Im} \, au)^2} - rac{1}{2} ig|F_1ig|^2 - rac{ig|G_3ig|^2}{2 ext{Im} \, au} - rac{1}{4} ig| ilde{F}_5ig|^2
ight) \ + rac{1}{2\kappa_{10}^2} \int C_4 \wedge rac{G_3 \wedge ar{G}_3}{4i ext{Im} \, au} + S_{ ext{local}} \, .$$

Useful combination $G_3 = F_3 - \tau H_3$; $\tau = C_0 + ie^{-\phi}$

Specify integrals, and local sources, satisfying RR tapole condition

$$Q_{flux} + Q_{D3} + Q_{O3} = 0$$

Flux "tension" and charge

$$\mathcal{L}_{G} = -\frac{1}{24\kappa_{10}^{2}} \int_{\mathbf{X}_{6}} d^{6}y \ g^{\frac{1}{2}} \frac{(G_{3})_{mnp} (\bar{G}_{3})^{mnp}}{\operatorname{Im} \tau} = -\frac{1}{8\kappa_{10}^{2}} \int_{\mathbf{X}_{6}} d^{6}y \ \frac{G_{3} \wedge *_{6}\bar{G}_{3}}{\operatorname{Im} \tau}$$

$$\mathcal{L}_{C_4} \,=\, rac{1}{2\kappa_{10}^2} \int C_4 \wedge rac{G_3 \wedge ar{G}_3}{4i \mathrm{Im}\, au}\,.$$

"BPS-like" for ISD flux

$$*_{6d}G_3 = iG_3$$

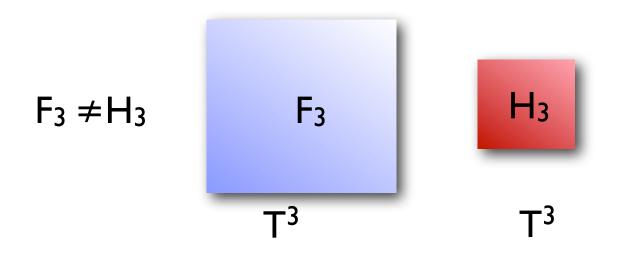
- Allows solving for compactifications with 4d Minkowski
- Stabilizes complex structure moduli (and dilaton)

$$F_3 = H_3$$
 F_3 H_3 T^3

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Warped geometries

- Fluxes gravitate. Backreaction.
- For ISD flux the I0d solution is warped M₄ x CY with warp and 5-form sourced by flux and local sources

$$\begin{split} ds^2 &= Z(x^m)^{1/2} \ ds_{4d}^2 + Z(x^m)^{1/2} \ g_{mn}^{CY} \ dx^m \ dx^n \\ F_5 &= (I + *_{I0d}) \ dZ^{-I} \ dx^0 \ dx^I \ dx^2 \ dx^3 \\ \nabla^2 \ Z(x^m) &= g_s \ |G_3|^2 + g_s \ \sum \delta(D3/O3) \end{split}$$

Analogous to 3-brane solution: Intuition: ISD flux works like "effective" D3-brane

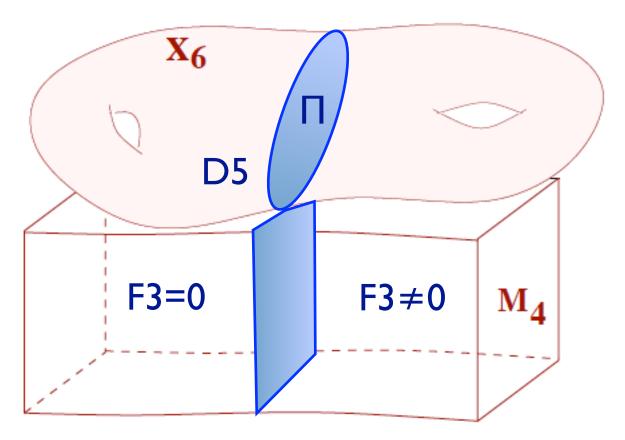
Underlying CY makes life easy

Many explicit models

Flux superpotential

Feet of flux on closed string moduli can be described in low energy effective field theory by W_{flux}

Compute W_{flux} from tension of domain wall introducing flux



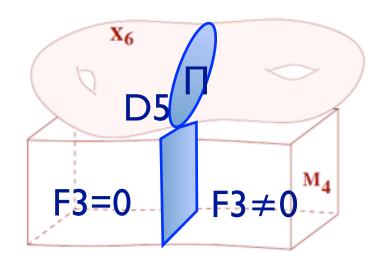
Flux superpotential



Compute W_{flux} from tension of domain wall introducing flux

D5 is magnetic source for C2

$$dF_3 = \delta_3(\Pi) \wedge \delta_0(x^3) dx^3$$



Integrate over dual 3-cycle Π' x interval in x3

$$\Delta \int_{\Pi'} F_3 = 1$$

From DW tension

$$W_{F_3} = \int_{\Pi} \Omega_3 = \int_{\mathbf{X}_6} F_3 \wedge \Omega_3$$

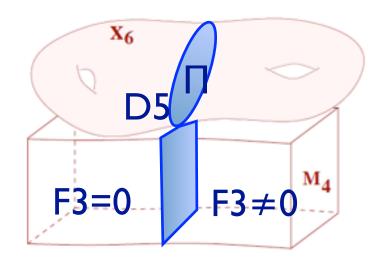
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From DW tension

$$W_{\rm GVW} = \int_{\mathbf{X}_e} G_3 \wedge \Omega_3$$

For general F3, H3 fluxes

Supersymmetry

- Can be studied from warped supergravity background
- Also from 4d effective theory

$$K = -\log(\int_X \Omega \wedge \overline{\Omega}) - \log(\tau - \overline{\tau}) - 3\log(\rho + \overline{\rho})$$

$$W_{\rm flux} = \int_X G_3 \wedge \Omega$$

Scalar potential
$$V = e^K \left(g^{a \bar{b}} D_a W D_{\bar{b}} \overline{W} - 3|W|^2 \right)$$

For complx. structure + dilaton $V = e^K \left(g^{i\bar{j}} D_i W D_{\bar{j}} \overline{W} \right)$

Vacuum:
$$D_i W_{\text{flux}} = 0$$
 $D_\tau W \sim \int_X \overline{G}_3 \wedge \Omega$ $G_3|_{(3,0)} = 0$

G3 is ISD
$$D_{z_i}W \sim \int_X G_3 \wedge \chi_{(2,1),i} \quad G_3|_{(1,2)} = 0$$

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Scalar potential $V = e^K \left(g^{a\bar{b}} D_a W D_{\bar{b}} \overline{W} - 3|W|^2 \right)$

Fo "No-scale structure" $V = e^K \left(g^{i \bar{j}} D_i W D_{\bar{j}} \overline{W} \right)$

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Supersymmetry

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 $G_3|_{(3,0)}=0$

G3 is ISD

 $D_{z_i}W \sim \int_X G_3 \wedge \chi_{(2,1),i} \quad G_3|_{(1,2)} = 0$

SUSY: G3 is (2,1) $D_{\rho}W \sim \int_{X} G_{3} \wedge \Omega$

 $G_3|_{(0,3)}=0$

F-theory description

- M-theory on elliptic CY4 in limit of vanishing fiber size
- Both NSNS and RR 3-form flux become M-theory 4-form flux

$$G_{(4)} = -\frac{G_{(3)}d\bar{w}}{\tau - \bar{\tau}} + \text{h.c.}$$

Stabilizes 4-fold complex str. moduli: some are 7-brane moduli!

Flux superpotential:
$$W = \int_{\mathbf{X}_8} G_4 \wedge \Omega_4$$

Vacuum: G4 is SD

Susy: G4 is (2,2)

Towards the SM

MSSM from magnetized D7s

$$N_{\alpha}$$
 $(n_{\alpha}^{1}, m_{\alpha}^{1})$ $(n_{\alpha}^{2}, m_{\alpha}^{2})$ $(n_{\alpha}^{3}, m_{\alpha}^{3})$ $N_{a} = 3$ $(0, I)$ $(3, I)$ $(-3, I)$ $N_{b} = 1$ $(1, 0)$ $(0, I)$ $(-1, 0)$ $(-1, 0)$ $N_{c} = 1$ $(1, 0)$ $(-1, 0)$ $(0, I)$ $(-1, 0)$ $N_{d} = 1$ $(0, I)$ $(3, I)$ $(-3, I)$ $(-3, I)$



(need few extra branes, adding few extra matter)

Can add Susy or Susy breaking 3-form flux to stabilize moduli

Susy
$$G_3 = \frac{8}{\sqrt{3}} e^{-\frac{\pi i}{6}} \left(d\overline{z}_1 dz_2 dz_3 + dz_1 d\overline{z}_2 dz_3 + dz_1 dz_2 d\overline{z}_3 \right)$$

Stabilizes at $au_1 = au_2 = au_3 = au = e^{2\pi i/3}$

Non-susy
$$G_3 = 2(d\overline{z}_1dz_2dz_3 + dz_1d\overline{z}_2dz_3 + dz_1dz_2d\overline{z}_3 + d\overline{z}_1d\overline{z}_2d\overline{z}_3)$$

Stabilizes e.g. at
$$au_1 = au_2 = au_3 = au = i$$



Corrections

- Earlier no-scale structure disappears upon including corrections
- Perturbative and non-perturbative, in α and g_s
- Corrections may be small in large volume, small g_s regime, but not compared to zero!

Generalized fluxes

- Above set of fluxes is not really the most general
- Geometric, non-geometric, U-dual fluxes
- Superpotentials depending on all moduli

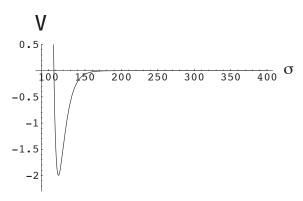
Use non-perturbative effects from D3-brane instantons

One-modulus toy model

$$W = W_{0,\text{flux}} + A(z_0)e^{-T}$$

Susy AdS vacua with stabilized moduli

[Kachru, Kallosh, Linde, Trivedi]

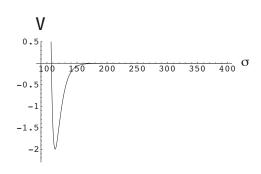


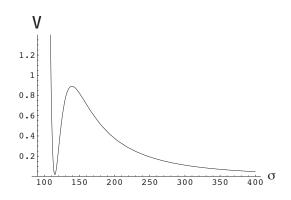
- Must tune W₀ t<< I to achieve controlable regime
 Can be relaxed in other scenarios (Large Volume Stabilization)
- Generalization: D3-instantons for each independent Kahler modulus?
- Clash with SM arising from D7's

Intersections of instanton with D7's lead to supos involving SM fields

$$W = e^{-T} \rightarrow W = e^{-T} \Phi_1 \dots \Phi_n$$

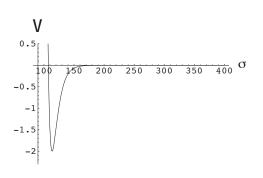
Proposal to add sources of extra tension for uplifting to deSitter

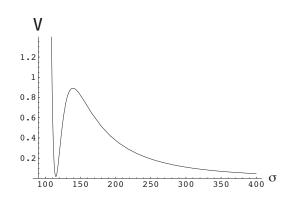




- Anti D3-branes

Proposal to add sources of extra tension for uplifting to deSitter





- Anti D3-branes
- D-terms: anti-instantons on D7s
- F-terms fon flux supo itself
- DSB sectors
- D-terms: anti-instantons on D7s

Generalized fluxes

- Flux compactifications in type IIA involve even and odd forms
- → Talk both to Kahler and complex structure moduli!

What about mirror symmetry?

- \checkmark Using T-duality in local T³ fibration, H₃ turns e.g. into geometric twist
- ⇒ Geometric fluxes
- Compactification on non-CY geometries, possibly 4d N=1
 - SU(3) holonomy $\rightarrow SU(3)$ structure
 - ⇒ Generalized complex geometry, mirror symmetry, ...
 - Painful lack of explicit compact examples

Generalized fluxes

Generalized geometric and non-geometric fluxes from T-duality

Regard T³ as T² (trivially) fibered over S¹

H₃ is monodromy b \rightarrow b+1 for b= \int_{T_2} B

Particular SL(2,Z) monodromy on T=A_{T2}+ib

One T-duality along T2 gives $\tau \rightarrow \tau + 1$ in SL(2,Z) of τ

⇒ Geometric twisting, geometric flux

One SI non-trivially fibered over two directions ω^a_{bc}

Two T-dualities give non-geometric SL(2,Z) monodromy on T

 \Rightarrow Non-geometric twisting, non-geometric flux Q^{ab}_c

Full T-duality covariance suggests

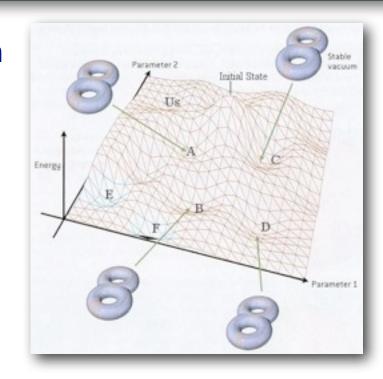
⇒ locally non-geometric R^{abc}



Flux landscape

The general picture is compelling enough Internal data determine 4d physics symmetries, spectrum, couplings

Many choices: "Landscape"



Flux landscape, part of full string landscape

Various estimates by flux counting

Revision of "naturalness": cosmological constant, hierarchy,...

Yeah, but...



What about some real Physics?



Keep rocking

After all, we are dealing with a theory which is much more clever than any of us

Fluxes and D-branes

- Must consider the interplay of different ingredients
- Focus on inter-relations between (field strength) fluxes and D-branes
- Efects at several levels
 - Topological: Freed-Witten consistency conditions
 - (Susy) Open string moduli stabilization
 - Susy breaking

Freed-Witten consistency conditions

- Certain D-branes are inconsistent in presence of NSNS 3-form flux
- Certain D-branes can decay in the presence of NSNS 3-form flux

Take D6-brane on 3-cycle with non-trivial H₃

$$\int_{\mathrm{D6}} H_3 \wedge \tilde{A}_4$$

Flux is magnetic source for D6 U(I)

Must be cancelled by boundaries of outgoint D4's



Ensures D-brane instantons respect gaugings in fluxes

Analogous statements for RR fluxes

Freed-Witten and discrete symmetries

- D-branes Z-valued in homology but Zn-valued in "K-theory"
- Torsion classes signal discrete gauge symmetries

Discrete gauge symmetries from flux cathalysis

Ex: Freund-Rubin (analogous to Zk center in AdS5/CFT4)

$$\int_{4d\times X_6} \overline{F}_6 \wedge B_2 \wedge F_2 \to k \int_{4d} B_2 \wedge F_2$$

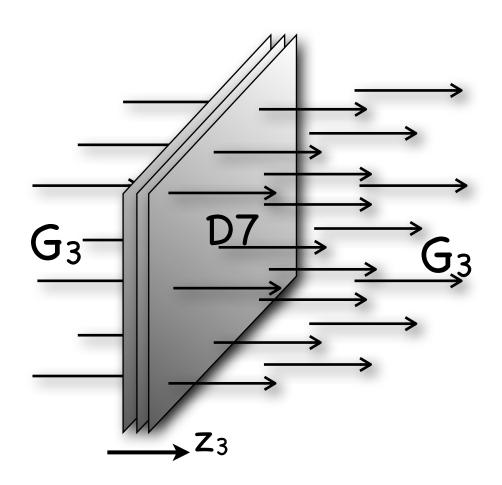
Instanton: NS5 on X6. Emits k D0-branes (Zk particles)

Junction: D6 on X6. Emits k F1's (Zk strings)

Can also do Zk domain walls. See later

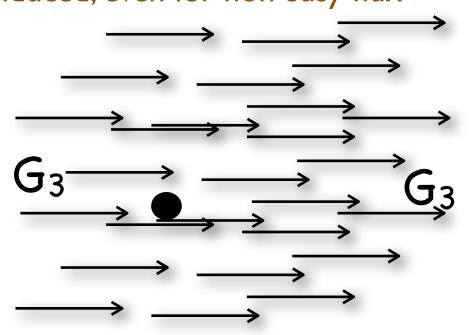
- An appealing scenario: Susy MSSM D-brane sector and non-susy flux
- Soft terms arise from effect of non-susy flux on susy D-branes

Explicitly computable using D-brane world-volume action in general supergravity background, or using 4d effective theory approach



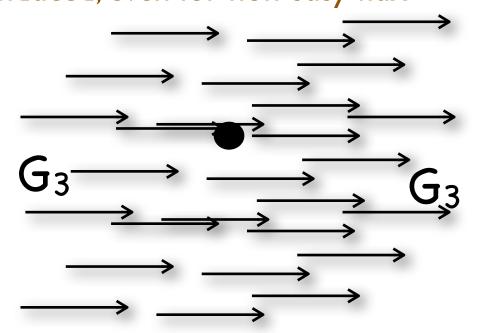
D3s in ISD 3-form flux background

ISD flux background: same tension and 4-form charge as a D3-brane Gravitational attraction cancels against Coulomb repulsion: No terms induced, even for non-susy flux



Solution Description Descripti

ISD flux background: same tension and 4-form charge as a D3-brane Gravitational attraction cancels against Coulomb repulsion: No terms induced, even for non-susy flux



Anti-D3s in ISD 3-form flux background

Effects add up instead of cancelling Scalars stabilize at max of flux density

D3s in ISD 3-form flux background

ISD flux background: same tension and 4-form charge as a D3-brane Gravitational attraction cancels against Coulomb repulsion: No terms induced, even for non-susy flux

$$m^2 \sim |G_{(0,3)}|^2$$
 $M \sim G_{(0,3)}$
 $A \sim G_{(0,3)}$

Anti-D3s in ISD 3-form flux background

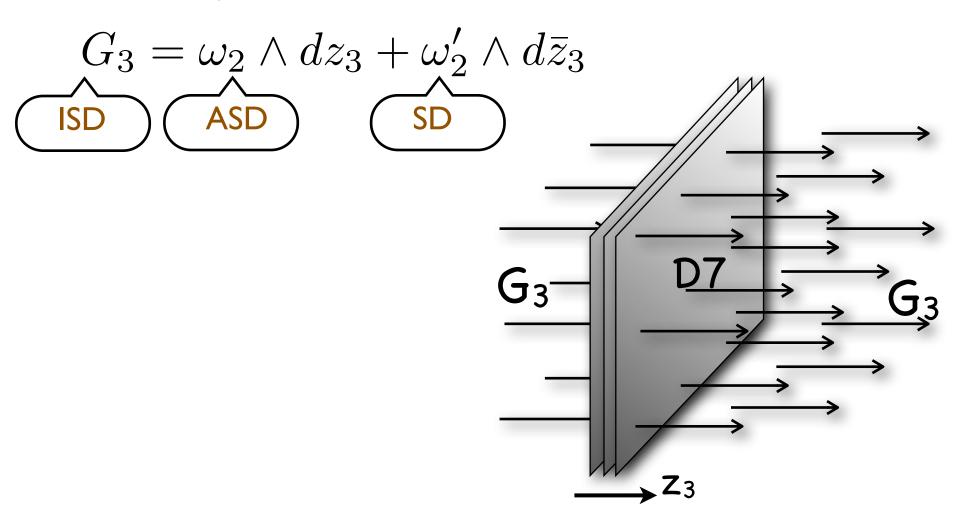
Effects add up instead of cancelling

Scalars stabilize at max of flux density

Fluxes, susy breaking and soft terms

PD7s in ISD 3-form flux background

Same tension and 4-form charge as a D3-brane: naively cancellation But flux itself, induces D3 or antiD3 on D7 volume



Fluxes, susy breaking and soft terms

PD7s in ISD 3-form flux background

Same tension and 4-form charge as a D3-brane: naively cancellation But flux itself, induces D3 or antiD3 on D7 volume

$$G_3 = \omega_2 \wedge dz_3 + \omega_2' \wedge d\bar{z}_3$$
ISD
ASD
SD

Non-trivial terms, even for ISD flux

Supersymmetric masses (mu-term): 3-form flux fix D7 moduli

Soft terms: Similar to those of antiD3s

Can recover both D3 and D7 results from effective theory

Flux components are vevs for moduli auxiliary fields: spurions

$$D_{\tau}W \sim \int_{X} \overline{G}_{3} \wedge \Omega \quad D_{z_{i}}W \sim \int_{X} G_{3} \wedge \chi_{(2,1),i} \quad D_{\rho}W \sim \int_{X} G_{3} \wedge \Omega$$

Fluxes, susy breaking and soft terms

Gravity mediation (in general, not universal, no mSUGRA)

Scales

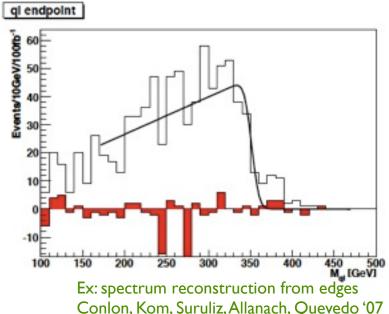
$$M_{SUSY} \sim f \frac{\alpha'}{R^3} \sim f \frac{M_c^2}{M_p}$$
 f: possible local suppression

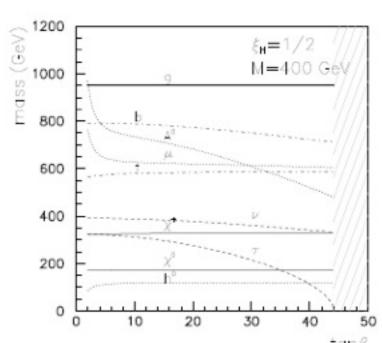
LowE Susy:

TeV soft terms from $M_c \sim 10^{11} {\rm GeV}$

Choose $M_c \sim 10^{14} {\rm GeV}$ then $M_{SUSY} \sim 10^{10} {\rm GeV}$

Can get to make plots



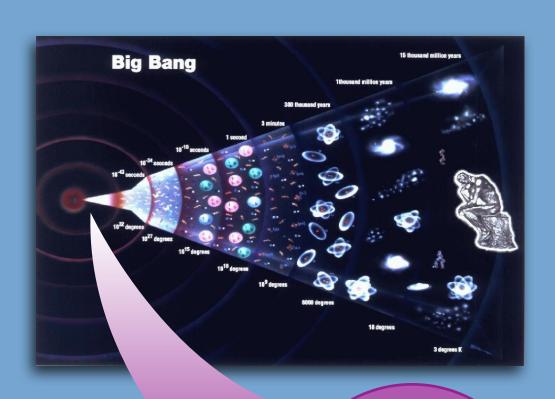


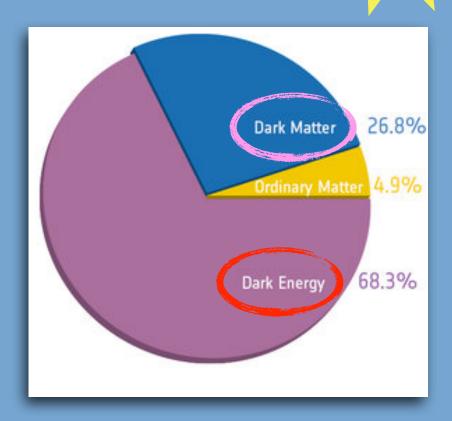
Ex: MSSM parameters Aparicio, Cerdeno, Ibanez '08

Cosmological Standard Model

(Λ CDM, "concordance model")





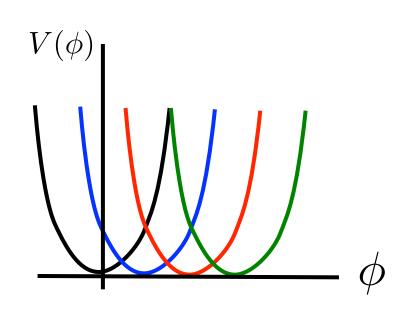


inflation

- Recent interest in large field inflation
- Scalars with shift symmetry (axions) are well protected continuous symmetry broken by non-pert effects to a discrete periodicity
- String theory axions have sub-Planckian decay constant

Axion monodromy: Potential is periodic but multivalued

Field theory analogue:
theta dependent vacuum energy
in large N pure gluodynamics Witten



Discrete Zn gauge symmetries



Discrete gauge symmetries

Zn particles, Zn strings, ...



Figure 1 There are Zn gauge symmetries associated to 4d domain walls

Zn symmetry of a 3-form that eats up a 2-form

$$\frac{1}{2}|F_4|^2 + |db_2 - n c_3|^2$$

Gauge invariance

$$c_3 \rightarrow c_3 + d\Lambda_2$$
 ; $b_2 \rightarrow b_2 + n\Lambda_2$

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Dualizing b2 to an axion, get Kaloper-Sorbo description of axion monodromy models.

Discrete Zn gauge symmetries

Can consider other Zn charged objects in 4d Lagrangian for 3-form eating up a 2-form

$$(|F_4|^2 + |dC_2 - nC_3|^2)$$

Gauge transformation
$$C_3 o C_3 + d\Lambda_2$$
 ; $C_2 o C_2 + n\Lambda_2$

Can be dualized to Φ F4 theory

$$|F_4|^2 + n \phi F_4 + |d\phi|^2$$

Dualizing also 3-form (to "(-1)form"), we get

$$|d\phi|^2 + \phi^2$$
 Massive axion

Can arise in D-brane & flux models



Structure is automatic in flux compactifications

After all, fluxes produce the stabilization of axions in moduli!!

10d Chern-Simons ⇒ modified field strengths

$$\int_{10d} B_2 \wedge F_p \wedge F_{p+2} \qquad \Rightarrow \qquad \tilde{F}_{p+2} = dC_{p+1} + B_2 \wedge F_p$$

Integrating over fluxed CY with $\ \phi = \int_{\Sigma_2} B_2$, $\ M = \int_{\Pi_p} F_p$

Change in axion induces extra flux

$$\Delta \phi \to \Delta \int_{\Sigma_2 \times \Pi_p} \tilde{F}_{p+2} = \phi M$$

A nice class: axions in flux compactifications

Change in axion induces extra flux

$$\Delta \phi \to \Delta \int_{\Sigma_2 \times \Pi_p} \tilde{F}_{p+2} = \phi M$$

Kinetic term for (p+2)-form leads to axion potential

In 4d terms, Kaloper-Sorbo lagrangian

$$\int_{10d} B_2 F_p F_{p+2} \to \int_{4d} M \, \phi \, F_4$$



Monodromy

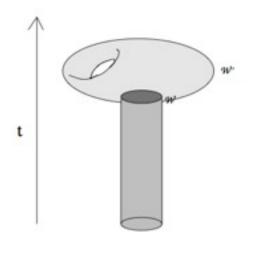
Multiple branches connected by domain walls changing (p+2)-form flux. They are D(6-p) on (4-p)-cycle



Many other realizations

Ex: IIB with NSNS flux on A-cycle

$$F_4 = \int_B F_7$$
; $\int_A H_3 = n$
$$\int_{4d \times X_6} \overline{H}_3 C_0 F_7 \to \int_{4d} n C_0 F_4$$



I0d IIB axion has a monodromy. Origin of energetic cost? GVW supo $W=\int_{X_6}(F_3-\tau\overline{H}_3)\wedge\Omega$

Period of C0 changes n units of F3 flux on A 4d Domain Wall is D5 on B 4d instanton is D(-1) (cosine modulation)

Conclusions

- Fig. It is amazing that something close to the SM can be realized in string theory
- Branes provide a tractable setup for SM model building in string theory

 Not more fundamental, but more manifest geometric intuition
- Program is not closed: Continuous progress
 - Moduli stabilization
 - Neutrino masses, non-perturbative operators
 - Discrete symmetries. Fluxed inflation models.
- Expect continuous progress and new results and useful input from LHC & cosmo