

Atomic probes of subGeV new physics

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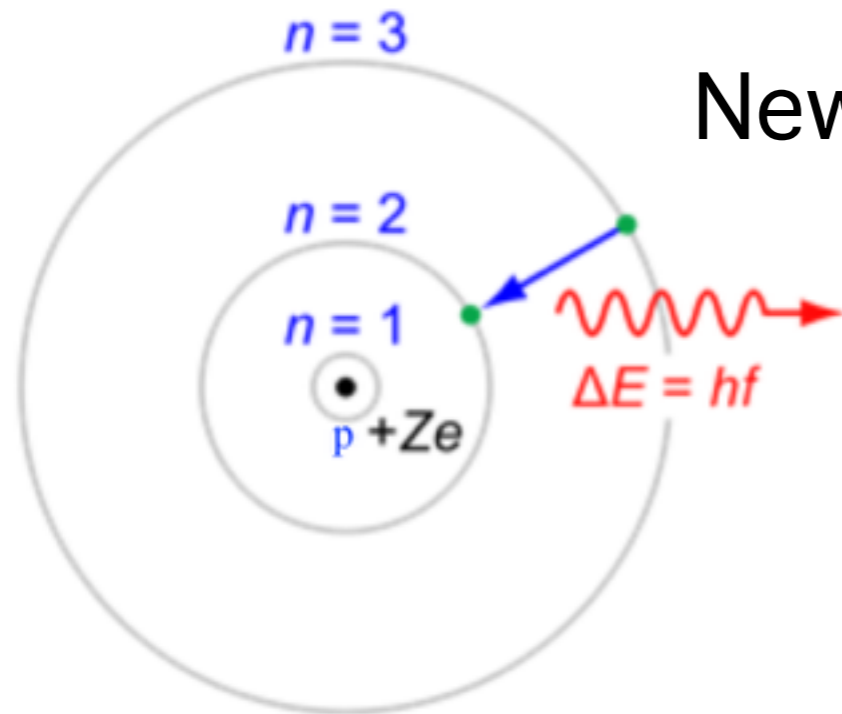
Work in progress with

Delaunay, Fuchs, Grojean, Harnik, Ozeri, Perez, Soreq

NPKI workshop
Korea University

Atomic Probes of new physics

Consider transitions between energy levels



New forces (long or short range)
change the frequency

$$V_{NIP}(r) = \frac{g_e g_A e^{-M r}}{4\pi r}$$

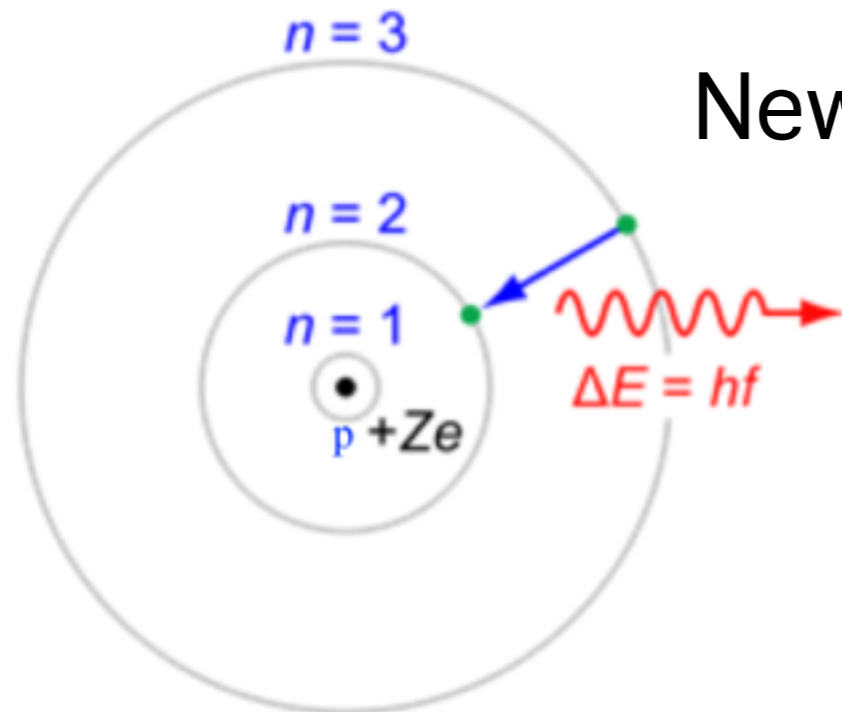
$r > a_0$ $V_{NIP} \propto 1/r$ massless limit

$$\delta E_{nlm} = \langle \psi_{nlm} | V_{NIP} | \psi_{nlm} \rangle = 4\pi \int_0^\infty dr r^2 V_{NIP}(r) |\psi_{nlm}(r)|^2$$

However we can't determine the absolute frequency
shift!

Atomic Probes of new physics

Consider transitions between energy levels



New forces (long or short range)
change the frequency

$$V_{NP}(r) = \frac{g_e g_A e^{-M r}}{4\pi r}$$

δE_n We can measure the difference in isotope frequency shift (IS) $|r)|^2$

However we can't determine the absolute frequency shift!

ISOTOPE SHIFT

consider

- ▶ 2 atomic level transitions with frequencies ν^1, ν^2
- ▶ 2 isotopes A, A' of the same atom/ion that have transitions 1,2

frequency shift of the **same transition** i in A, A' : **factorisation** [King '63]

$$\delta\nu_i^{AA'} \equiv \nu_i^A - \nu_i^{A'} = K_i \mu_{AA'} + F_i \delta\langle r^2 \rangle_{AA'}$$

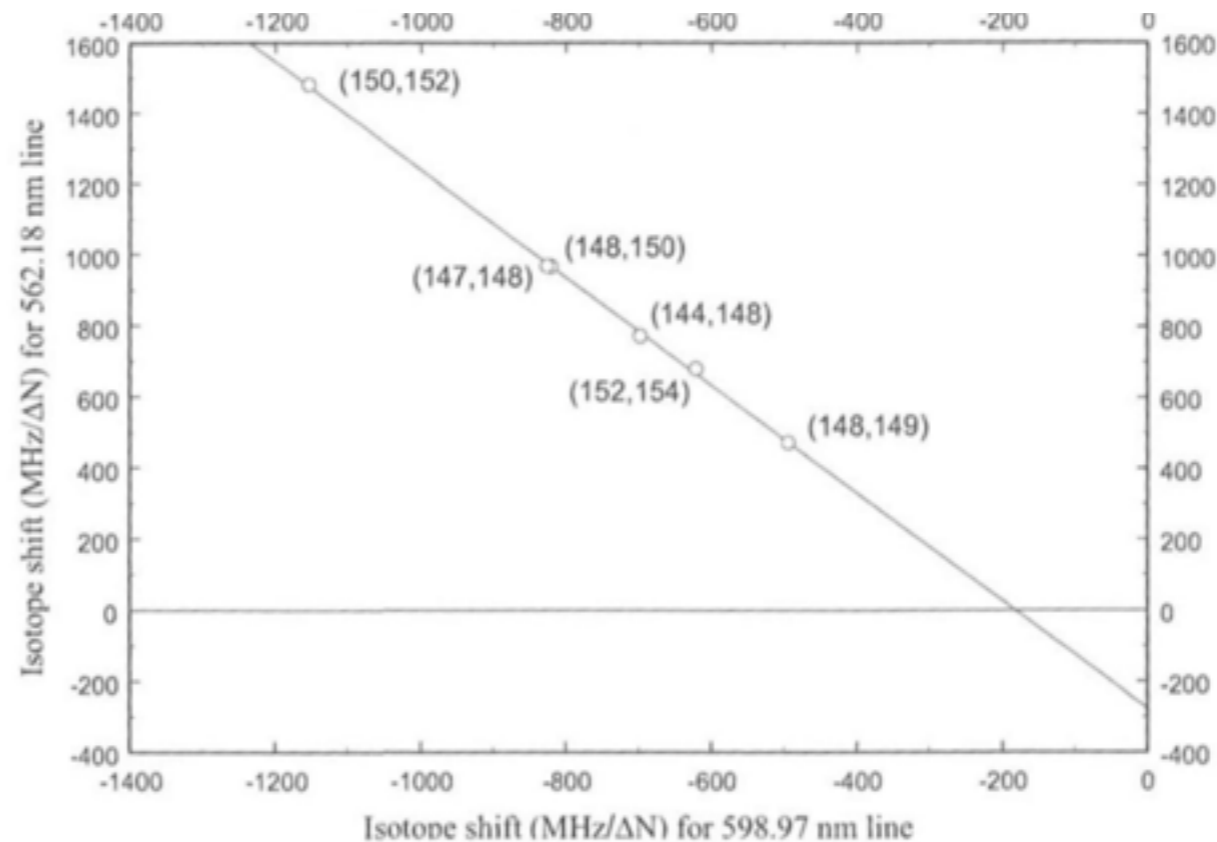
$$\mu_{AA'} = 1/m_A - 1/m_{A'}$$

- ▶ (normal) mass shift: changed electron reduced mass
- ▶ field/volume shift: point charge approximation for nucleus breaks

Absolute frequencies difficult, but frequency shifts more precise

- ▶ eliminate unknown AA' -difference of charge distribution variance $\delta\langle r^2\rangle_{AA'}$, normalise $\delta\nu/\mu \equiv \tilde{\delta\nu}$
- ▶ express frequency shift $\tilde{\delta\nu}_{AA'}^2$ in terms of $\tilde{\delta\nu}_{AA'}^1 \Rightarrow$ linear relation

$$\tilde{\delta\nu}_{AA'}^2 = F \cdot \tilde{\delta\nu}_{AA'}^1 + K$$



classical King plot

- ▶ eliminate unknown AA' -difference of charge distribution variance $\delta\langle r^2 \rangle_{AA'}$, normalise $\delta\nu/\mu \equiv \tilde{\delta\nu}$
- ▶ express frequency shift $\tilde{\delta\nu}_{AA'}^2$ in terms of $\tilde{\delta\nu}_{AA'}^1 \Rightarrow$ linear relation

$$\tilde{\delta\nu}_{AA'}^2 = F \cdot \tilde{\delta\nu}_{AA'}^1 + K$$



Now if we add the NP contribution..

$$\delta\nu_{AA'}^i = K_i \mu_{AA'} + F_i \delta r_{AA'}^2 + H_i (A - A')$$

classical King plot

Isotope shift (MHz/ ΔN) for 598.97 nm line

$$H_{21} = H_2 - F_{21} H_1$$

$$m \delta\nu_{AA'}^2 = F_{21} m \delta\nu_{AA'}^1 + K_{21} - AA' H_{21}$$

➔ new physics can violate King's linearity

Light mediators

$$F_i \text{ contact interaction } F_i \propto |\psi(0)|^2$$

$$\text{For short range interactions } H_i \propto F_i$$

for long range interactions we probe instead a sizeable region of electronic wave function

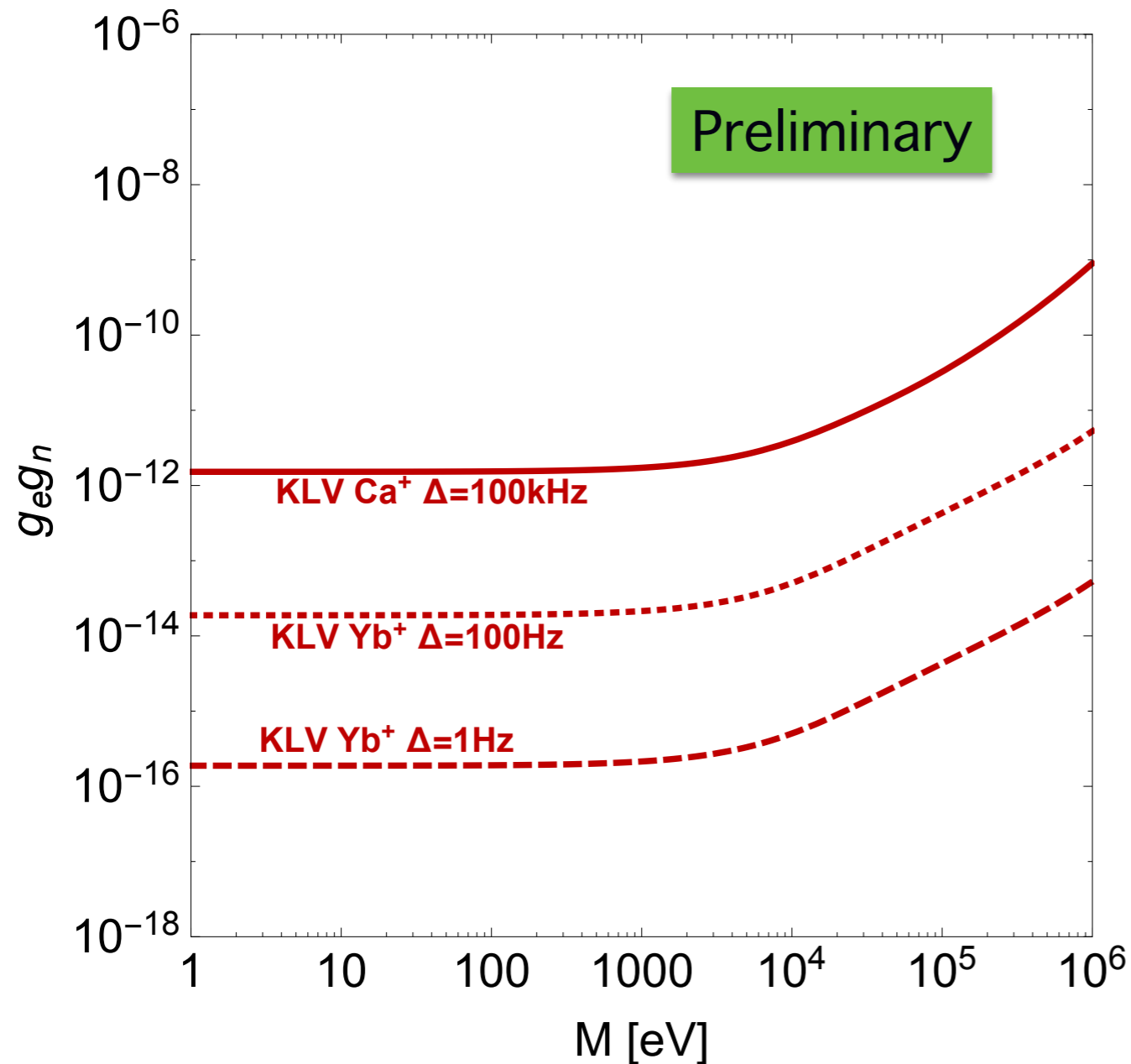


Breaking of Kings linearity

$$m\delta\nu_{AA'}^2 = F_{21}m\delta\nu_{AA'}^1 + K_{21} - AA'H_{21}$$

Probe of new light particles (scalar or vector)
coupled to neutrons and electrons
(spin independent interaction only)

KLV sensitivity to light mediators



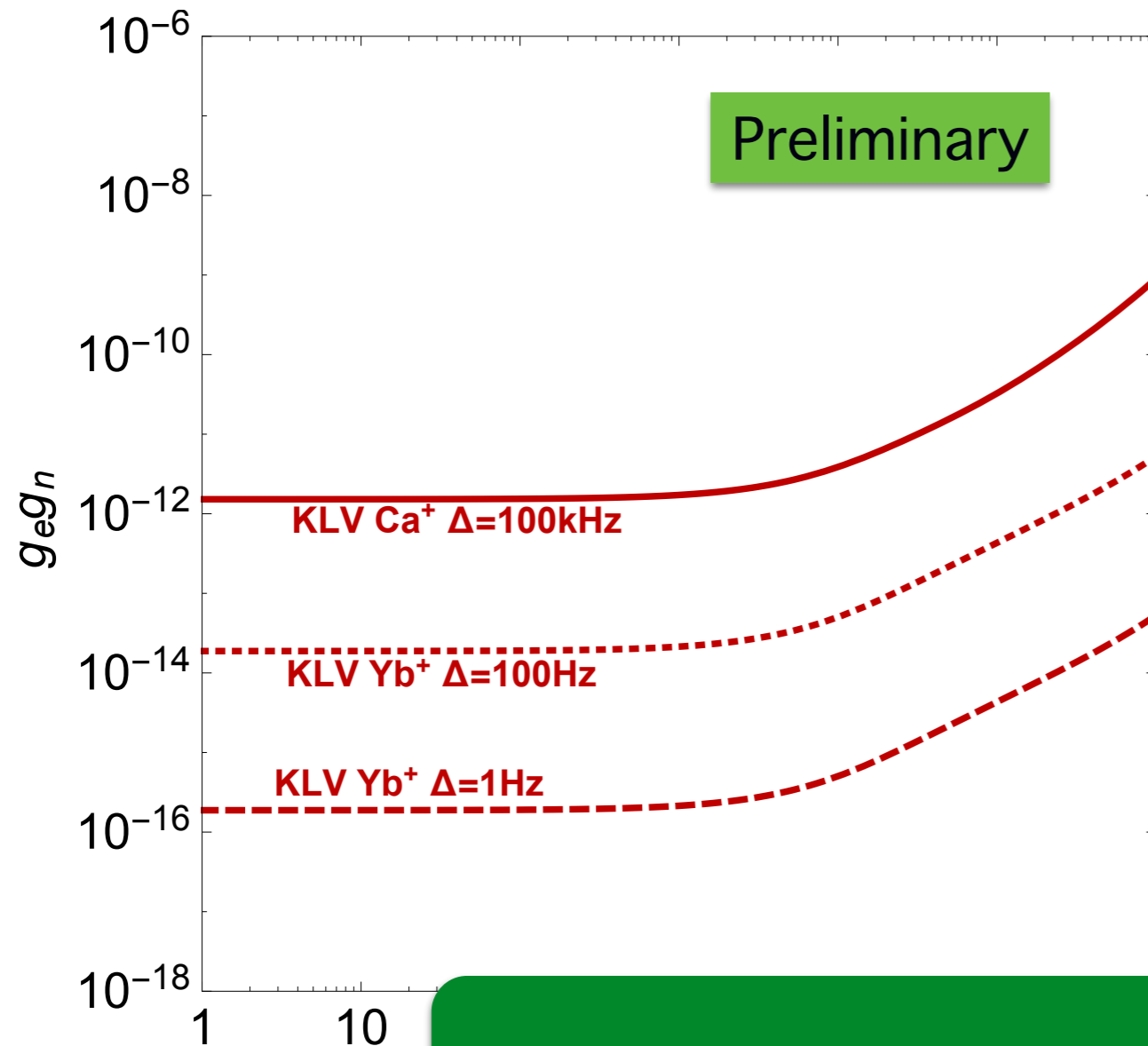
Non linearity supposed
to hold up to 1 Hz

g_e coupling to the electron
 g_N coupling to the neutron

Focus on region below 1 MeV to be safe-
potential good sensitivity up few tens of MeV

work in
progress

KLV sensitivity to light mediators



Non linearity supposed
to hold up to 1 Hz

g_e coupling to the electron
 g_N coupling to the neutron

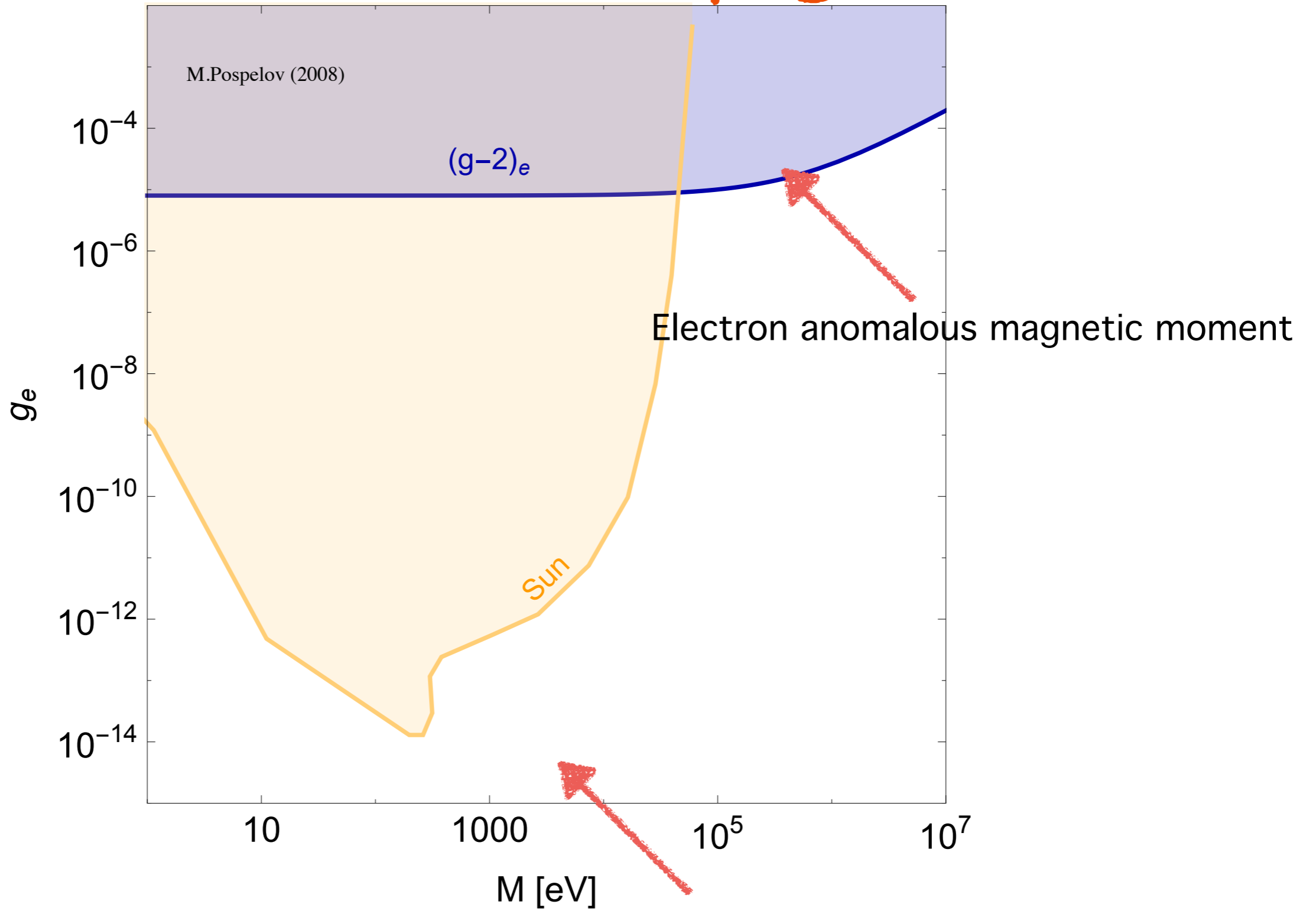
What are the existing constraints?
which frequency is necessary to set the
strongest bounds?

work in
progress

Foc
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V

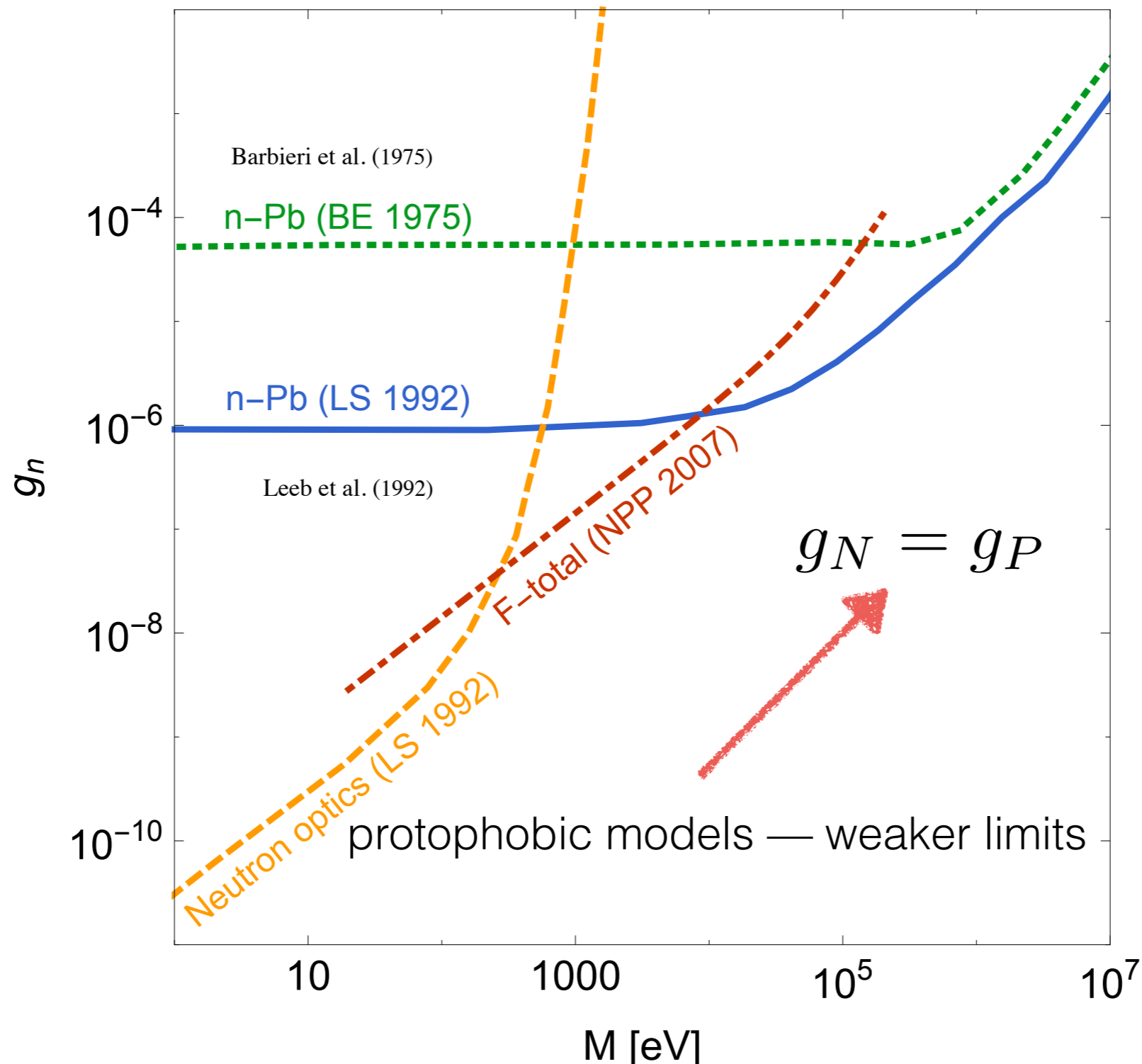
Probes of g_e



Astro bounds coming from the production via thermal radiation in the sun of a new scalar or vector boson G. G. Raffelt and G. D. Starkman (1989)

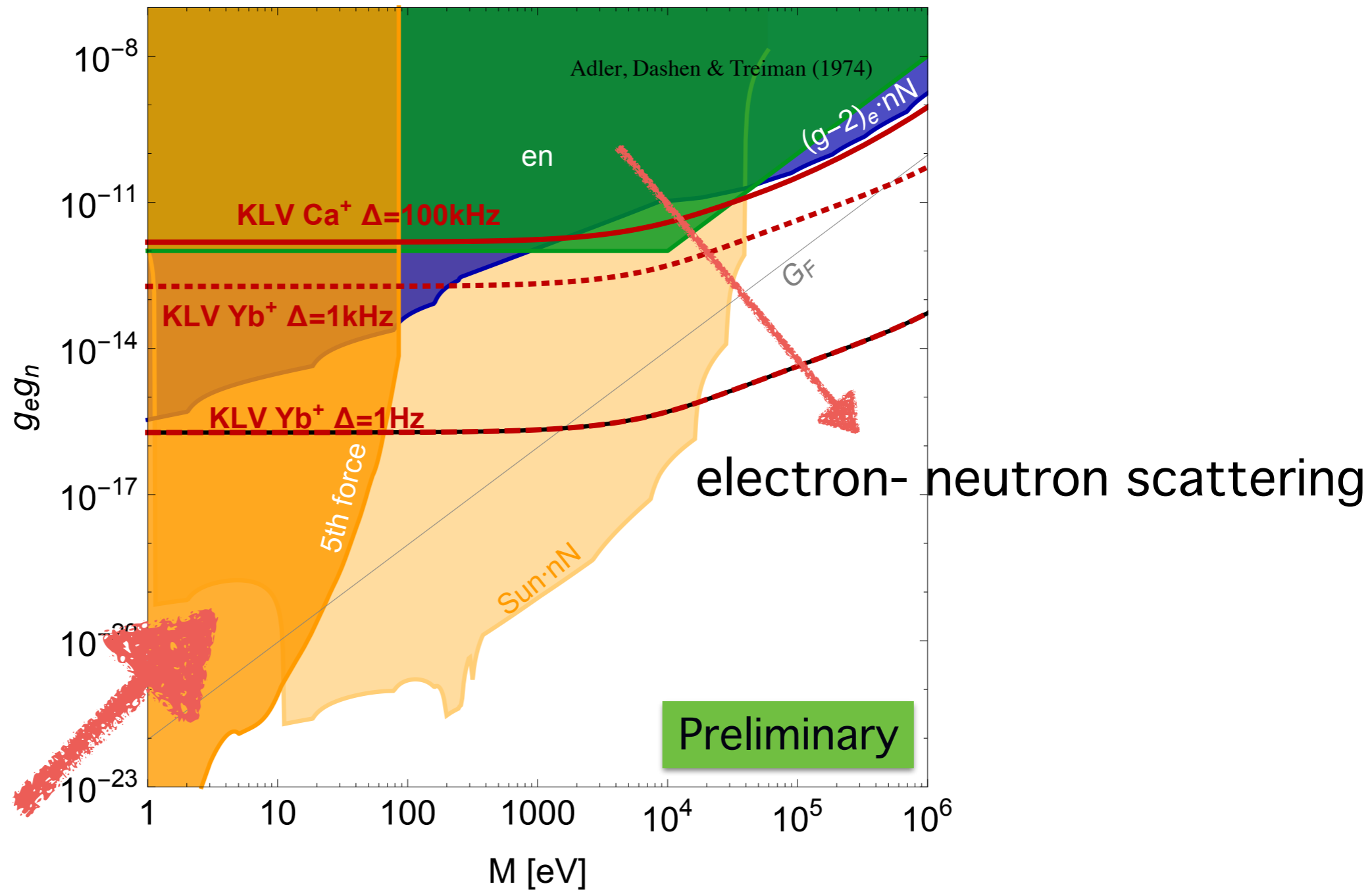
Probes of g_N

Low energy neutron-nucleus scattering



ex: angular distribution measurements in low energy neutron-nucleus scattering.

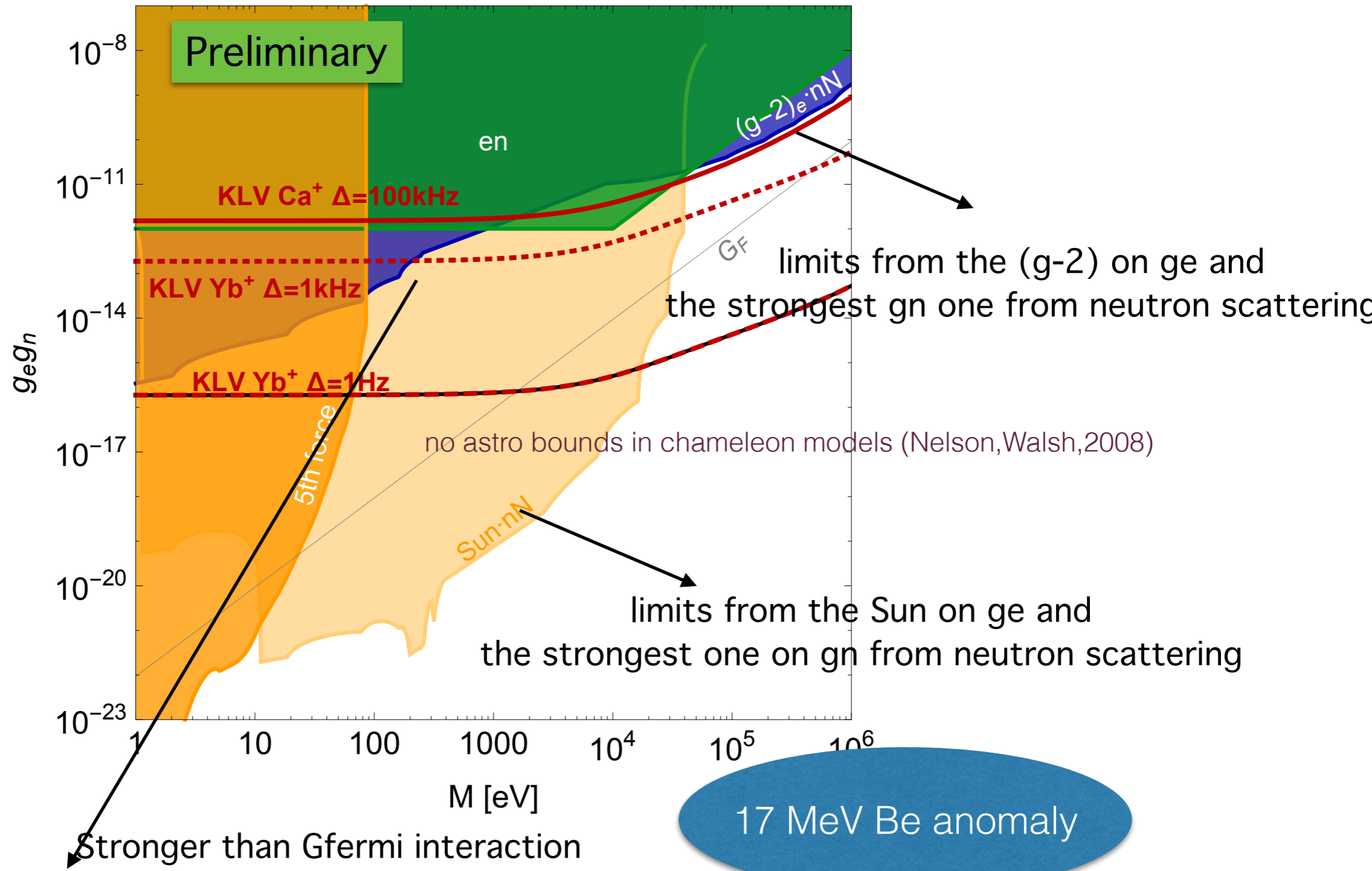
Probes of $g_N g_e$



Fifth force experiments

too strong to compete- we focus on > 50 eV mass range

Probes of $g_N g_e$



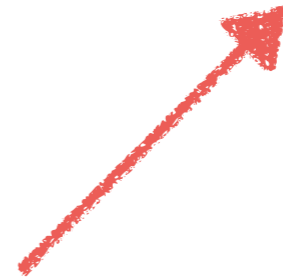
Heavier mediators: work in progress

Models?

▲ We can't set limits on dark photon kinetic mixing since it is aligned with QED

▲ Benchmark model B-L or the Higgs portal

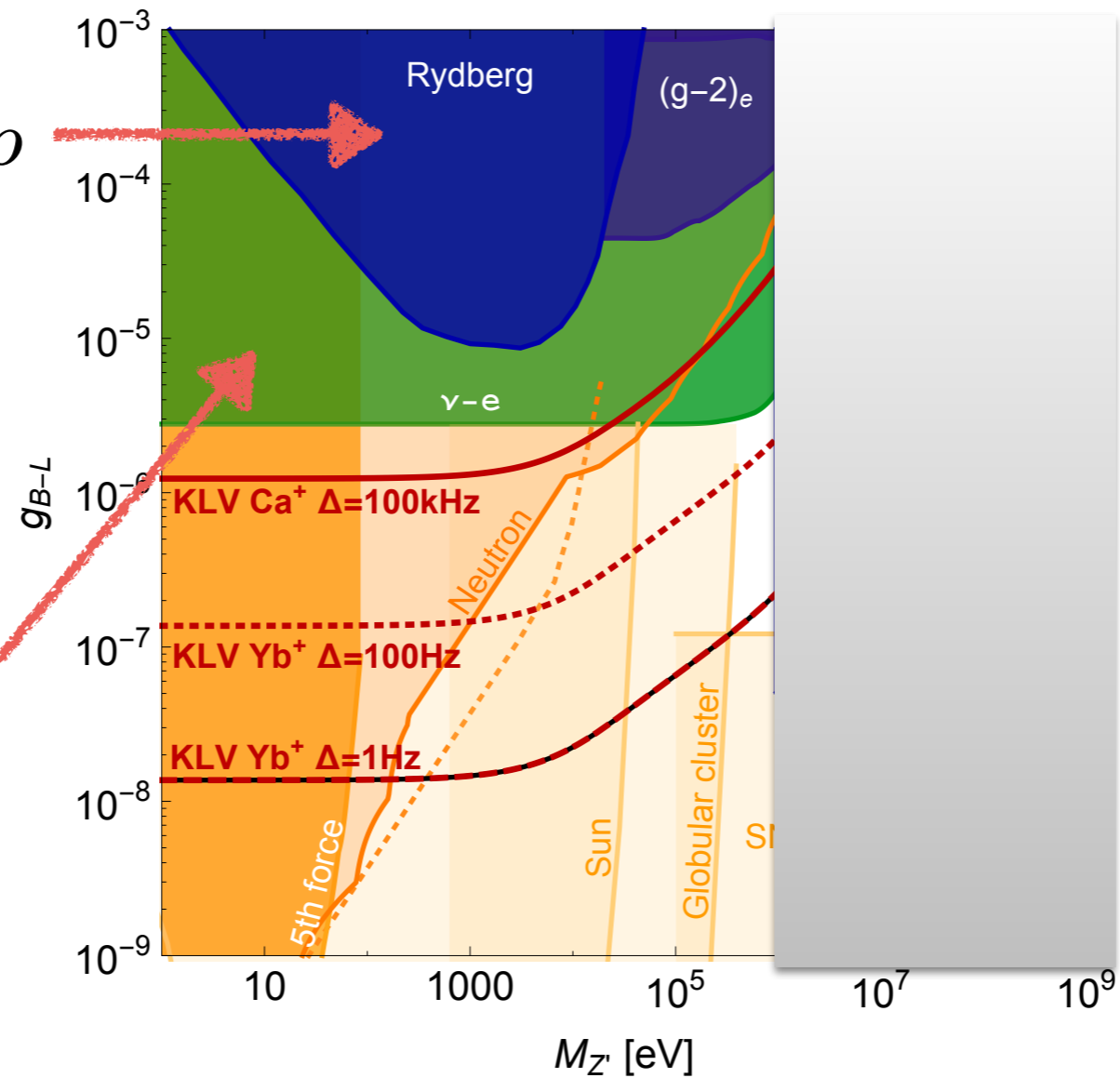
limits weakened by the smallness of electron yukawa coupling



Constraints on B-L

Preliminary

Probe of $g_e \times g_p$



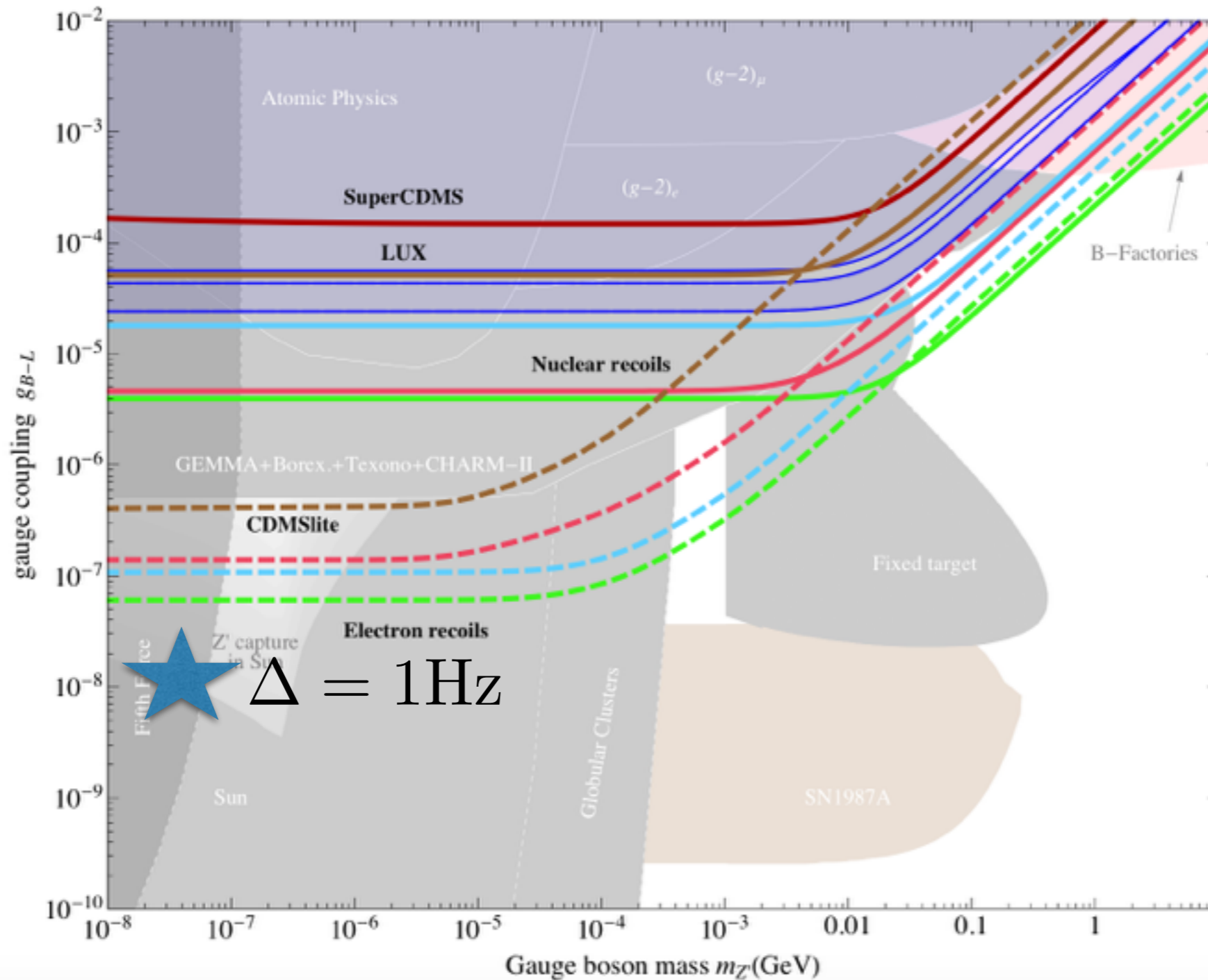
bounds on (solar or reactor) neutrinos- electron scattering

Bilmis et al. (2015)

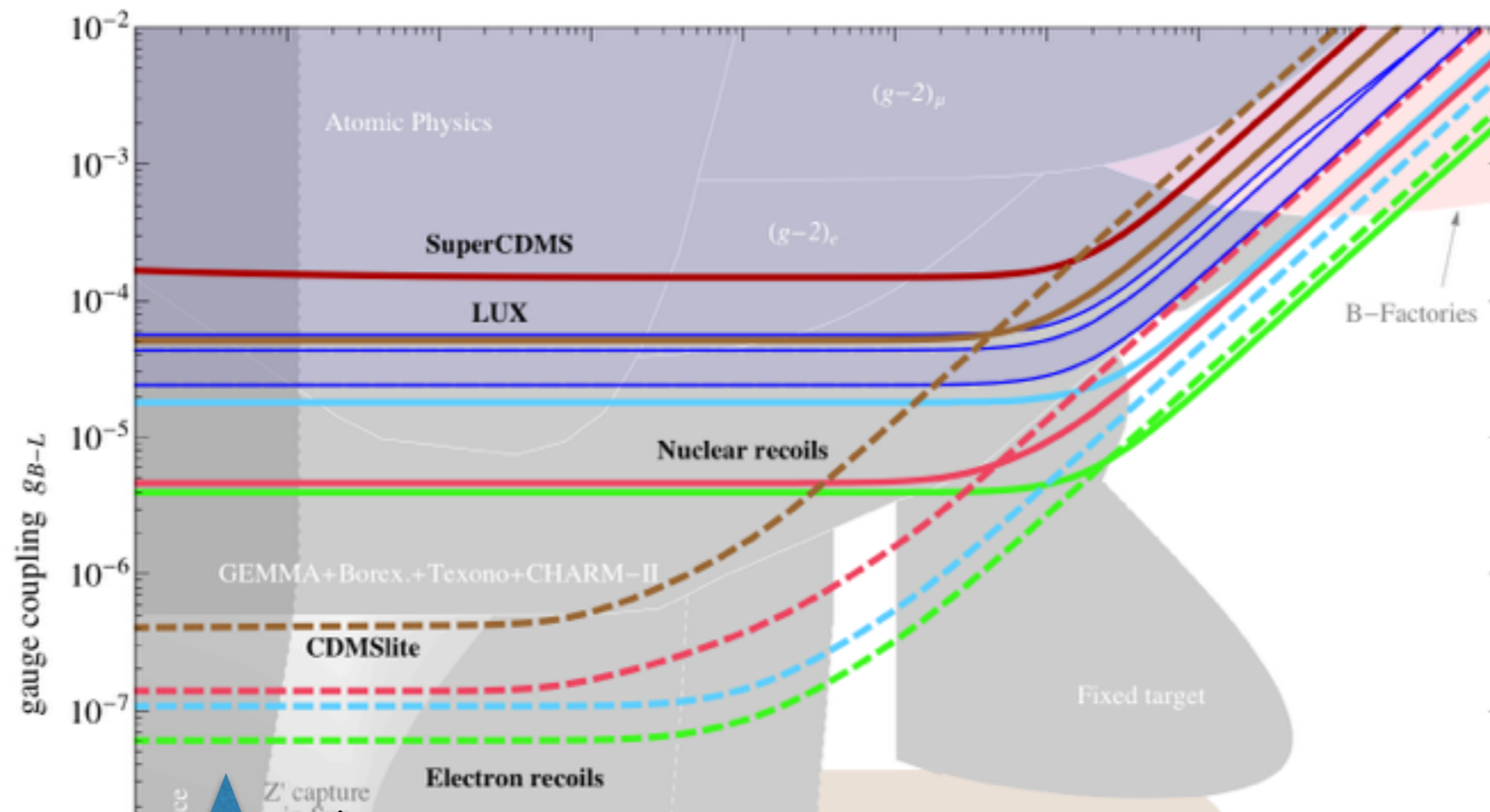
Kopp et al (2013)

Pospelov (2013)

Neutrino signals at dark matter direct detection experiments?



Neutrino signals at dark matter direct detection experiments?



If a signal is detected at future DM experiments our measurements will rule out some neutrino related interpretations

Conclusions

★ We can probe long range interactions with isotope shift measurements via King's linearity

★ We can set the strongest lab limits in the $50 \text{ eV} < m < 1 \text{ MeV}$ region and in some region also compete with astrophysics