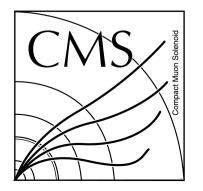
## Dark matter searches: MET+X

# 에너지를 누락

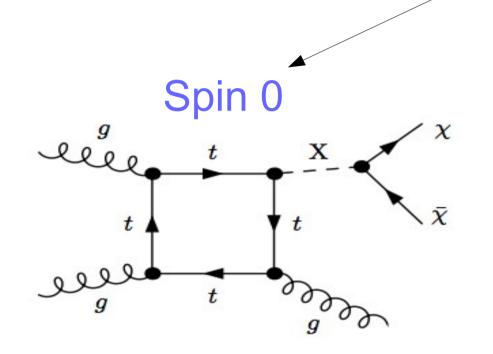


**Phil Harris** 



### Search for Dark Matter at LHC

Can split dark matter into two classes of searches



Yukawa coupling to quarks (At the moment no mixing)

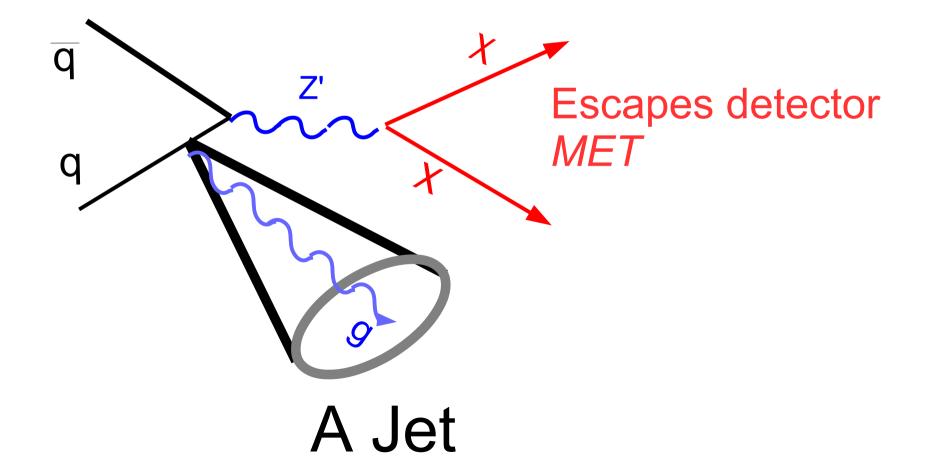
Flavor universal to quarks (At the moment no mixing)

All dark matter searches are really a search for Dark Matter + A mediator

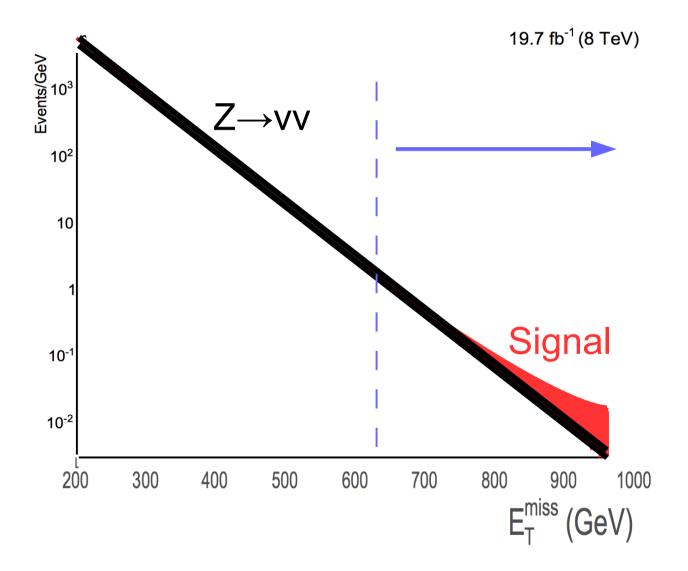
## Experimental Strategy

### The Basic Monojet Search

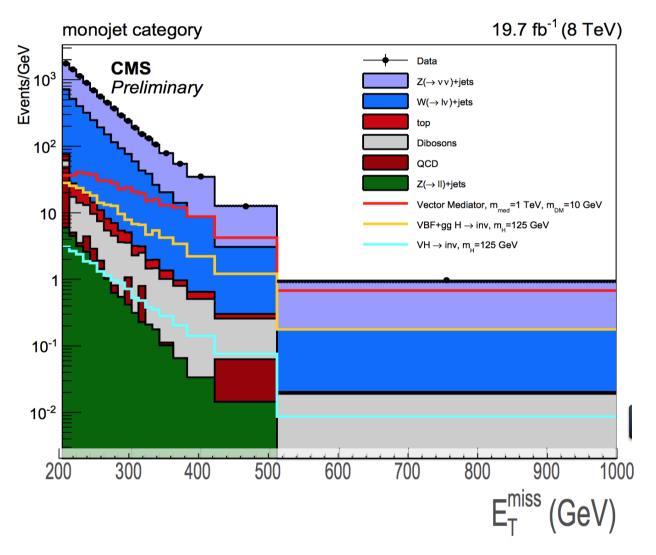
Escaping detector gives us signatures of MET



#### How do we search?



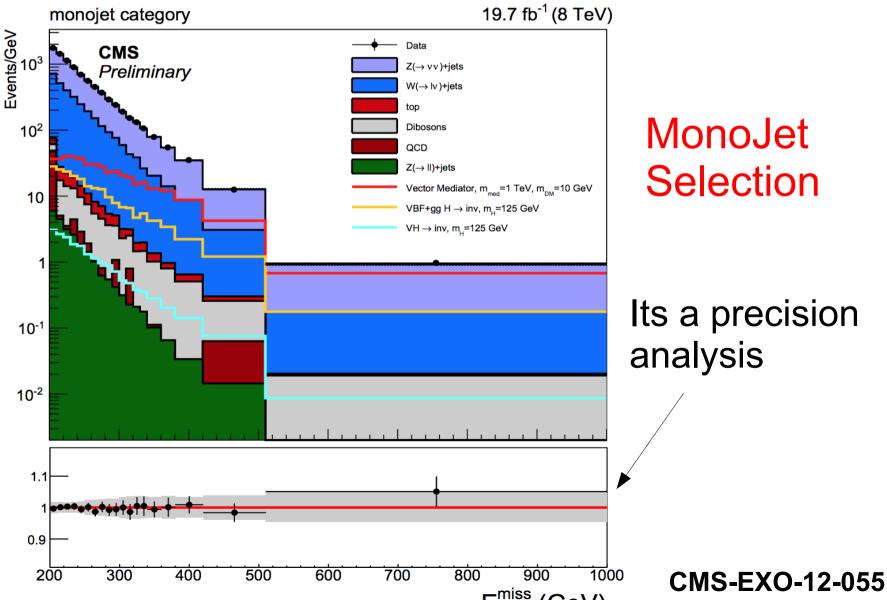
#### How do we search(data)?



#### MonoJet Selection

#### CMS-EXO-12-055

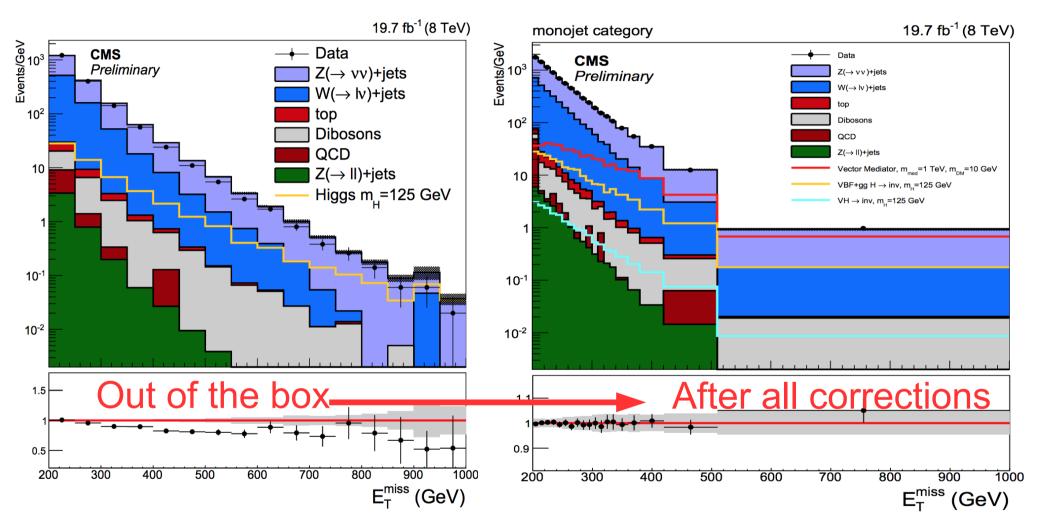
#### How do we search(data)?



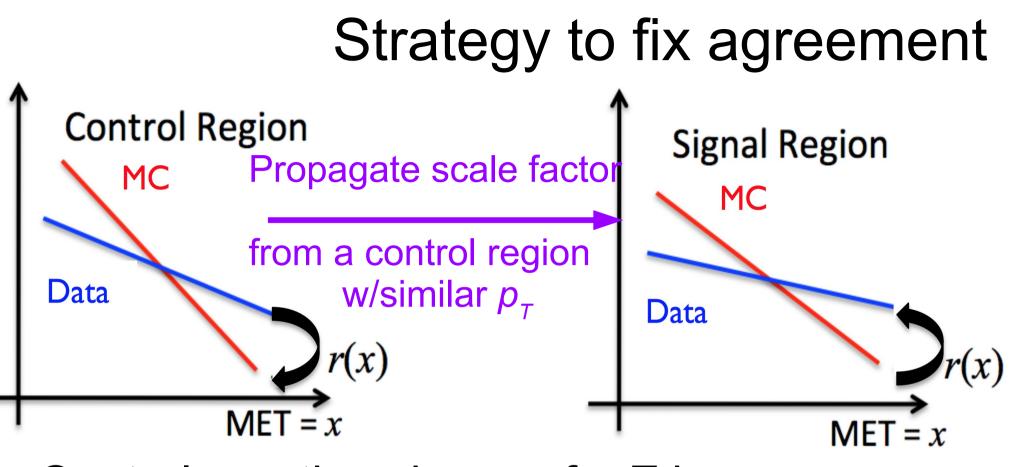
MonoJet Selection 7

Its a precision analysis

### How do we get to this precision?

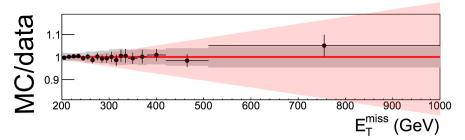


Rely on a series of control regions to correct for the data/MC agreement CMS-EXO-16-010 CMS-EXO-12-055

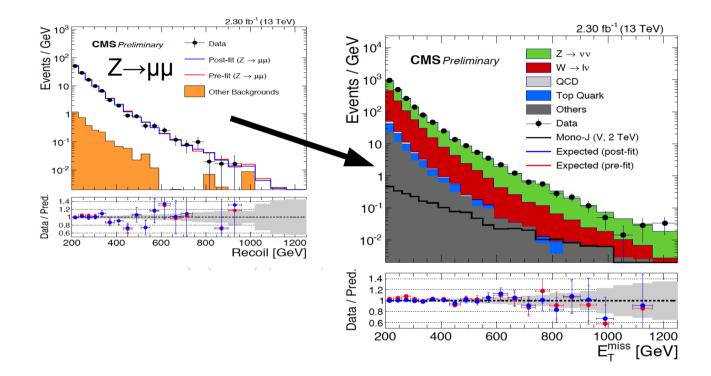


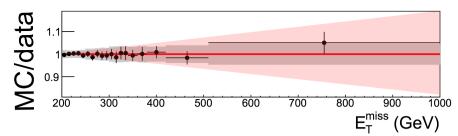
Control: another decay of a Z boson  $Z \rightarrow VV$ Remove

Problem is control regions have less events than signal  $\sigma_{\mu\mu} = 0.1 \sigma_{\nu\nu}$  -Statistical precision is 4x worse CMS-EXO-16-010 CMS-EXO-12-055

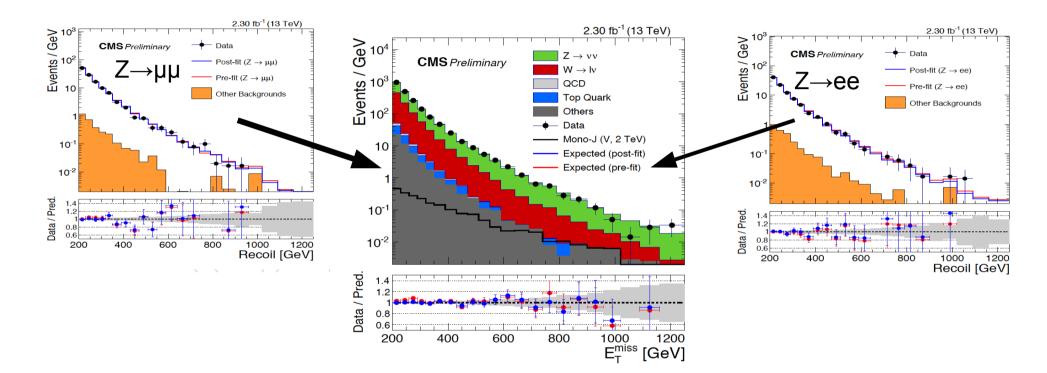


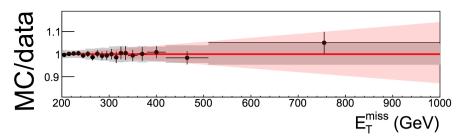
#### 1 Control region 100% uncertainty @ 1 TeV



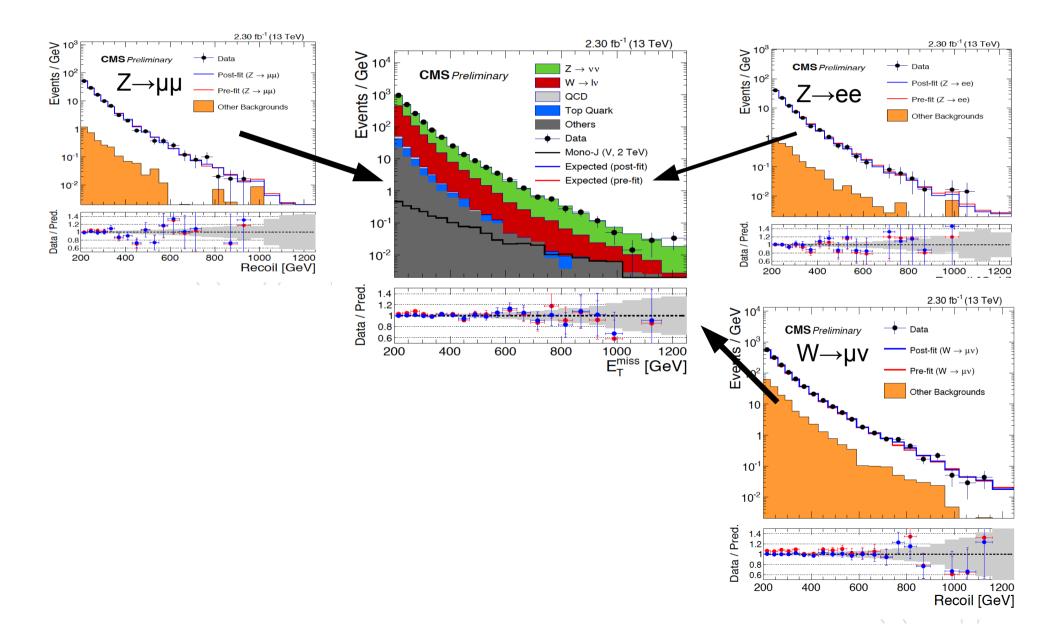


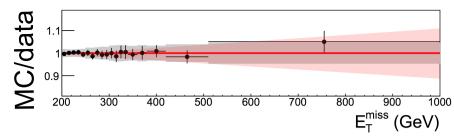
#### 2 Control regions 60% uncertainty @ 1 TeV



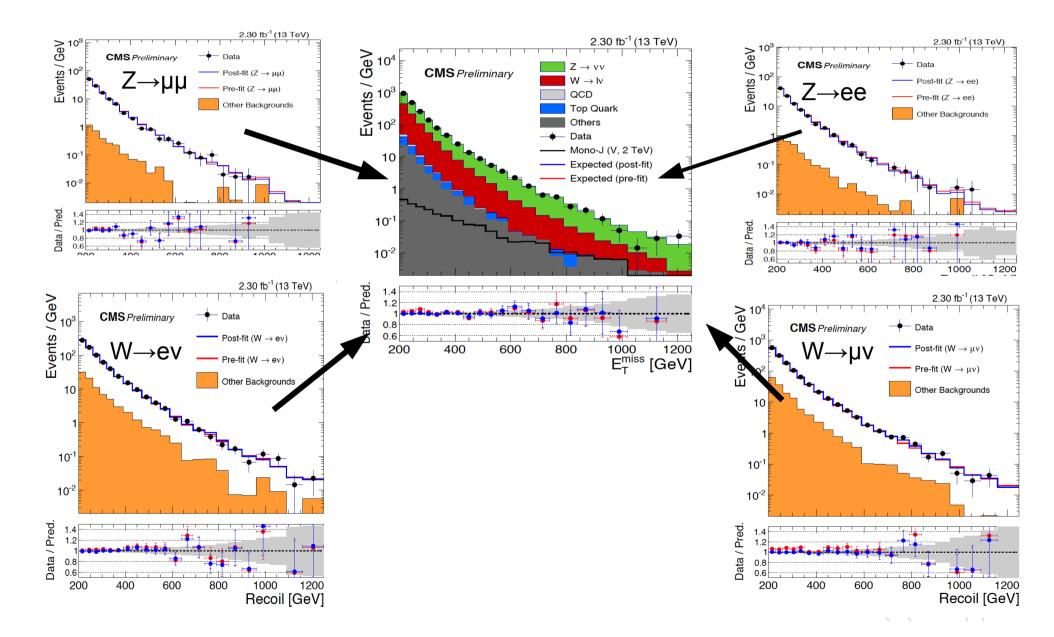


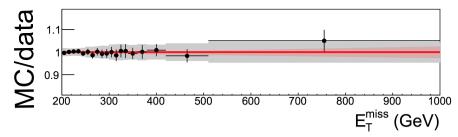
#### 3 Control regions 40% uncertainty @ 1 TeV



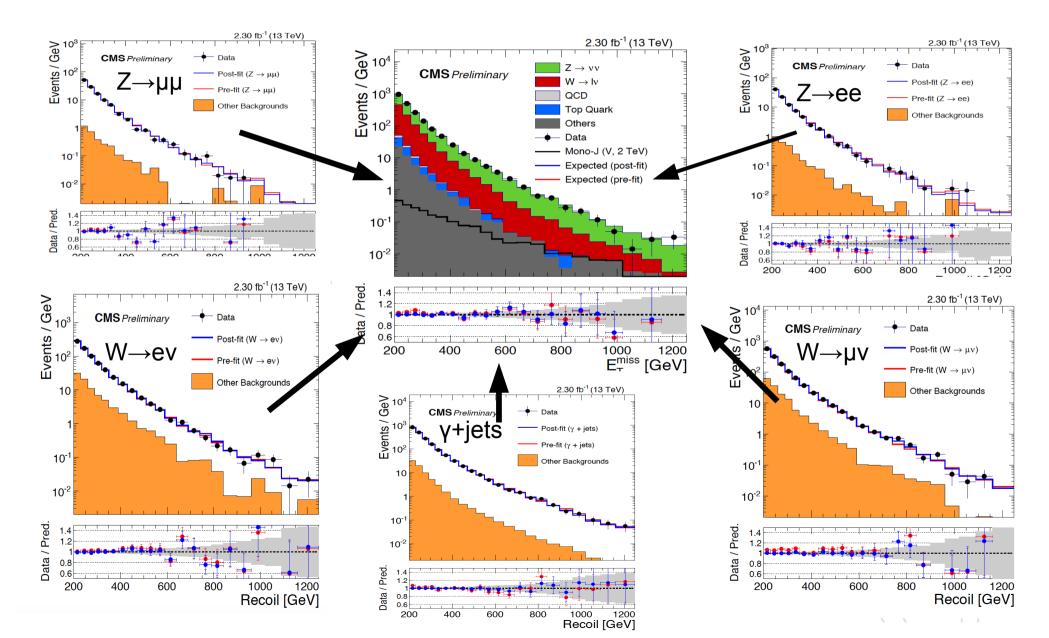


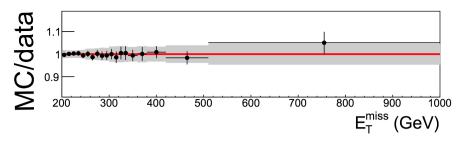
#### 4 Control regions 30% uncertainty @ 1 TeV



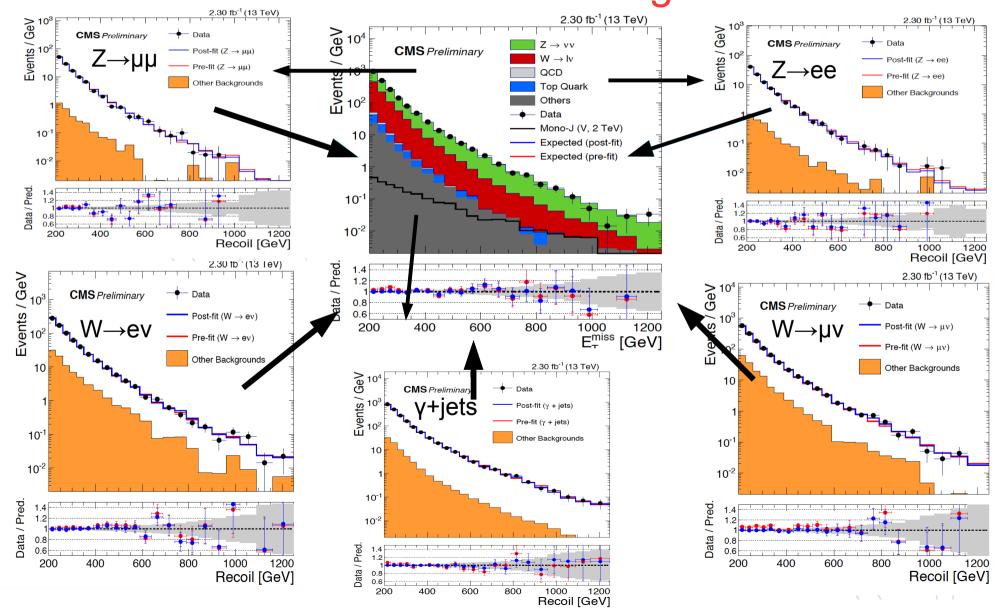


#### 5 Control regions 15% uncertainty @ 1 TeV



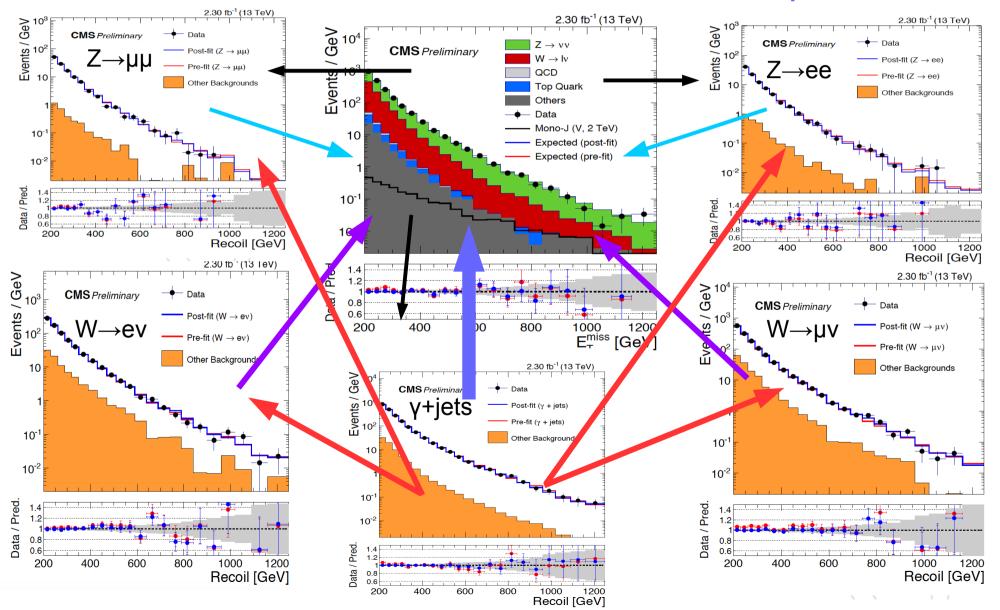


#### 5 Control regions+Signal 15% uncertainty @ 1 TeV All in one big Simultaneous fit



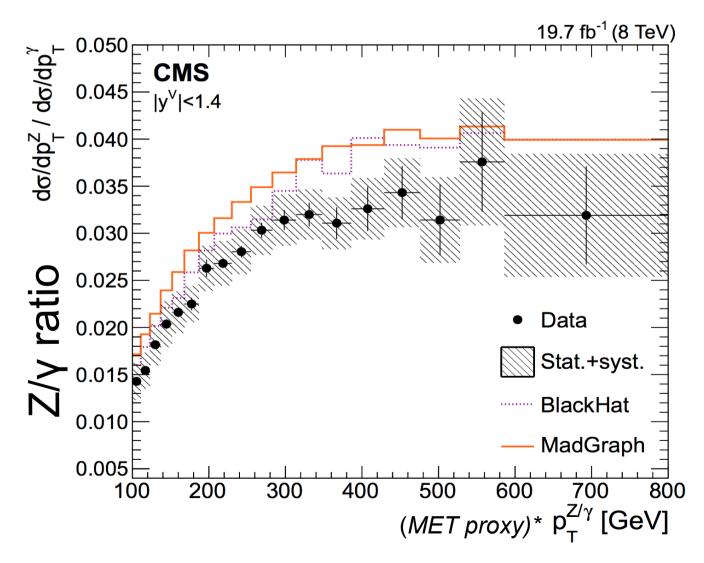
To large extent the γ+jets drives the constraint

- However we need need  $Z \rightarrow II$  to constraint  $\gamma$ 



### A mystery? Understanding $Z/\gamma p_{\tau}$

#### Can we really use Photons to model Zs?

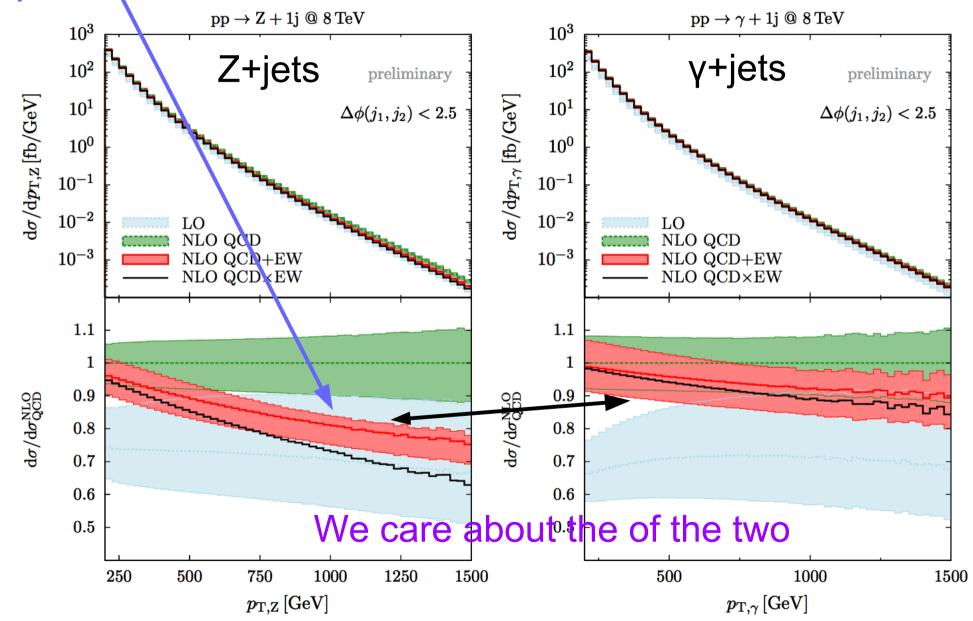


\*See backup

#### CMS-SMP-14-005

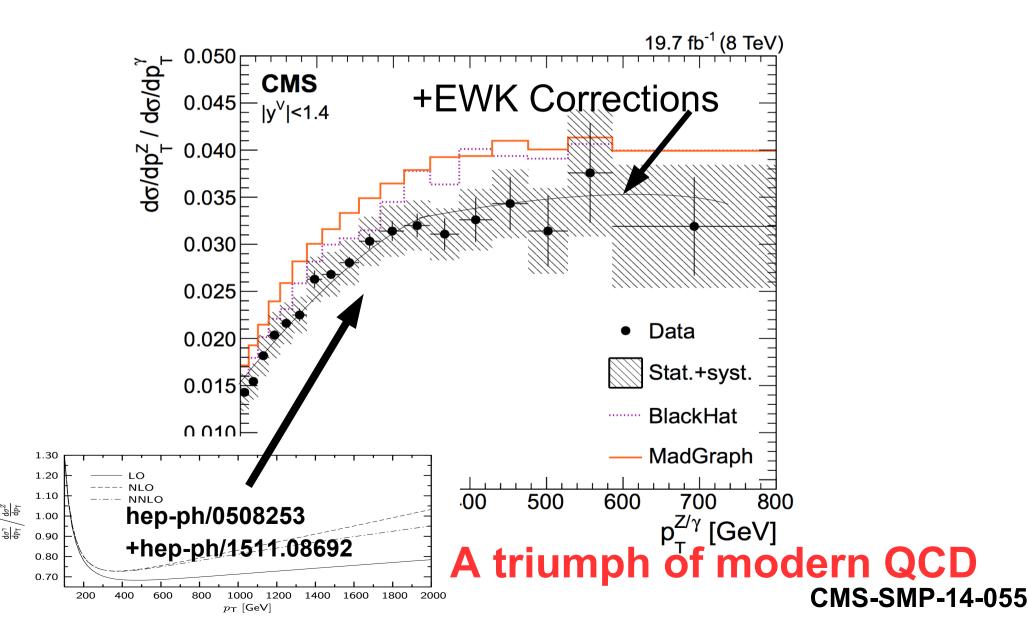
### How do we fix this?

Impact of the electroweak corrections



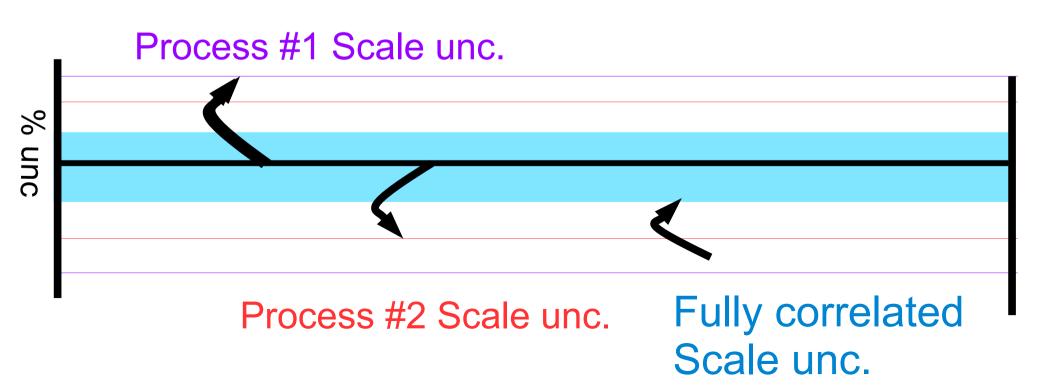
### A mystery? The Z $p_{\tau}$ spectrum

• These results are missing NLO EWK corrections!



#### However we still have a problem!

Unc. 
$$- \frac{d\sigma^{\gamma}}{dp_{\tau}} / \frac{d\sigma^{z}}{dp_{\tau}} = d\sigma^{\gamma}/d\sigma^{z}(\mu)$$



#### Uncertainty on ratio? How is it done? Scale uncertainty on process #1 Scale uncertainty on process #2 Uncertainty on process #1/process #2 (fully correlated) % unc $\frac{d\sigma^{\gamma}}{dp_{T}} / \frac{d\sigma^{Z}}{dp_{T}} = d\sigma^{\gamma}/d\sigma^{Z}(\mu)$ Unc. $\begin{pmatrix} d\sigma^{\gamma}(+\sigma) \\ d\sigma^{Z}(+\sigma) \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \begin{pmatrix} d\sigma^{\gamma}(\mu^{up})/d\sigma^{i}(\mu_{0}) \\ d\sigma^{Z}(\mu^{up})/d\sigma^{i}(\mu_{0}) \end{pmatrix}$

## Uncertainty on ratio? How is it done?

Scale uncertainty on process #1 Scale uncertainty on process #2

%

UN

Uncertainty on process #1/process #2 (fully correlated)

Jnc. 
$$\rightarrow \frac{d\sigma^{\gamma}}{dp_{T}} / \frac{d\sigma^{Z}}{dp_{T}} = d\sigma^{\gamma}/d\sigma^{Z}(\mu)$$

 $\begin{pmatrix} d\sigma^{\gamma}(+\sigma) \\ d\sigma^{Z}(+\sigma) \end{pmatrix} = \begin{pmatrix} 1 & C \\ C & 1 \end{pmatrix} \begin{pmatrix} d\sigma^{\gamma}(\mu^{up})/d\sigma^{i}(\mu_{0}) \\ d\sigma^{Z}(\mu^{up})/d\sigma^{i}(\mu_{0}) \end{pmatrix}$  Adjust C until uncertainty is

### Uncertainty on ratio? How is it done?

Scale uncertainty on process #1 Scale uncertainty on process #2 Uncertainty on process #1/process #2 (fully correlated)

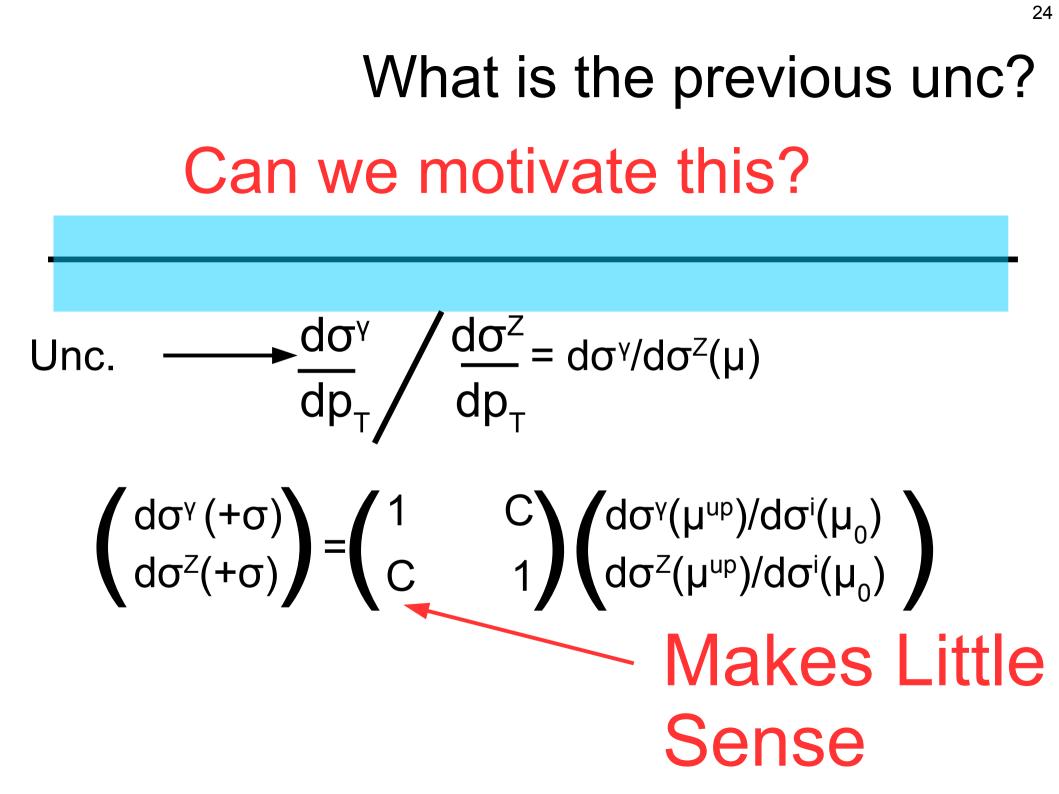
%

Unc.  $\frac{d\sigma^{\gamma}}{dp_{T}} / \frac{d\sigma^{Z}}{dp_{T}} = d\sigma^{\gamma}/d\sigma^{Z}(\mu)$  $d\sigma^{\gamma}(+\sigma) / 1 C / d\sigma^{\gamma}(\mu^{up})/d\sigma^{i}(\mu_{0})$ 

 $\begin{pmatrix} d\sigma^{\gamma}(+\sigma) \\ d\sigma^{Z}(+\sigma) \end{pmatrix} = \begin{pmatrix} 1 & C \\ C & 1 \end{pmatrix} \begin{pmatrix} d\sigma^{\gamma}(\mu^{up})/d\sigma^{i}(\mu_{0}) \\ d\sigma^{Z}(\mu^{up})/d\sigma^{i}(\mu_{0}) \end{pmatrix}$ 

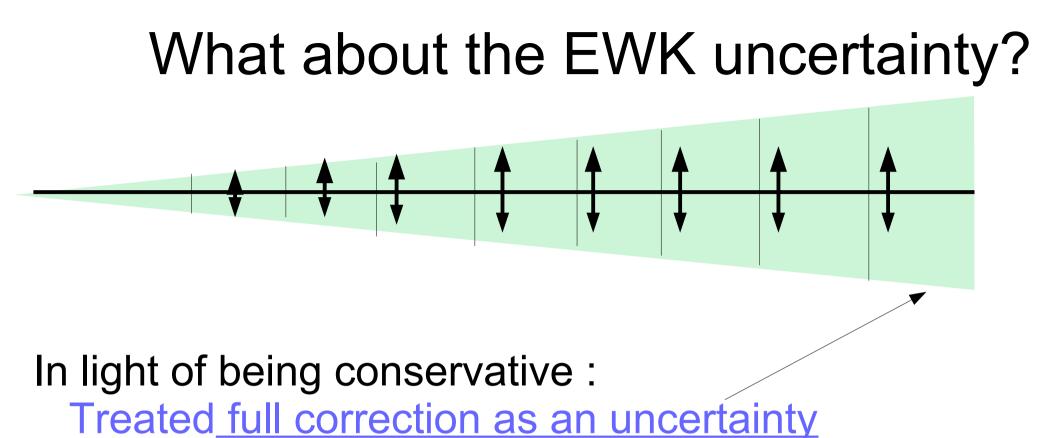
Decorrelate scale unc. until its max of either process

 $d\sigma^{\gamma}/d\sigma^{z}(+\sigma) < \max_{i} (d\sigma^{i}(\mu^{up})/d\sigma^{i}(\mu_{0}))$ 



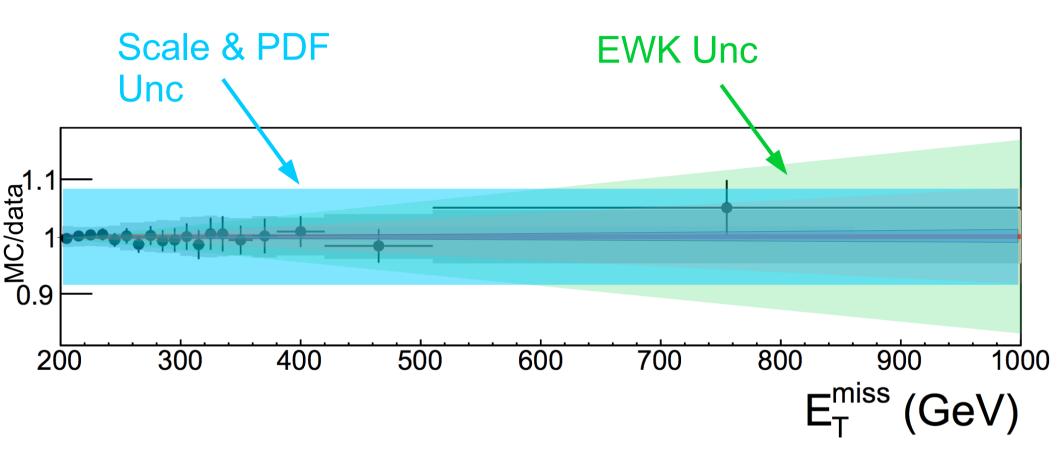
#### What about the EWK uncertainty?

#### In light of being conservative : Treated <u>full correction as an uncertainty</u> More formal way could be with scale



Additionally <u>de-correlated this per bin</u> Avoids low *MET* to high *MET* constraints Not very logical Other (better) schemes exist

#### What do the uncertainties look like?

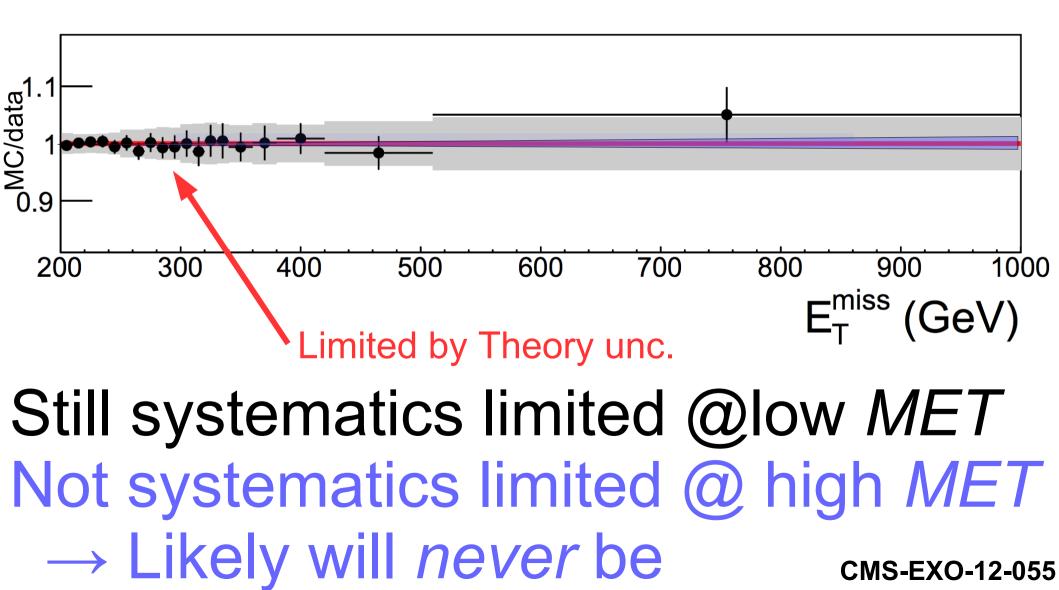


#### Updated unc still too large

CMS-EXO-12-055

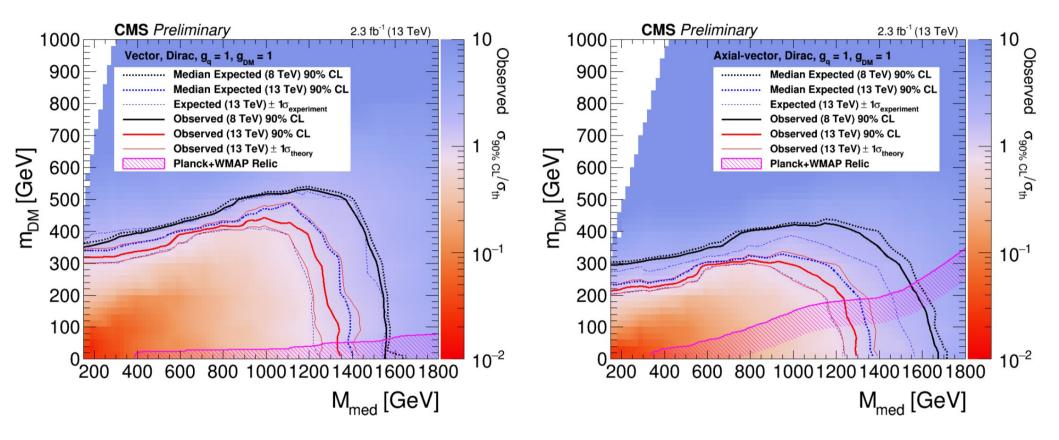
### Profiling them in the fit

#### Constraints after the fit



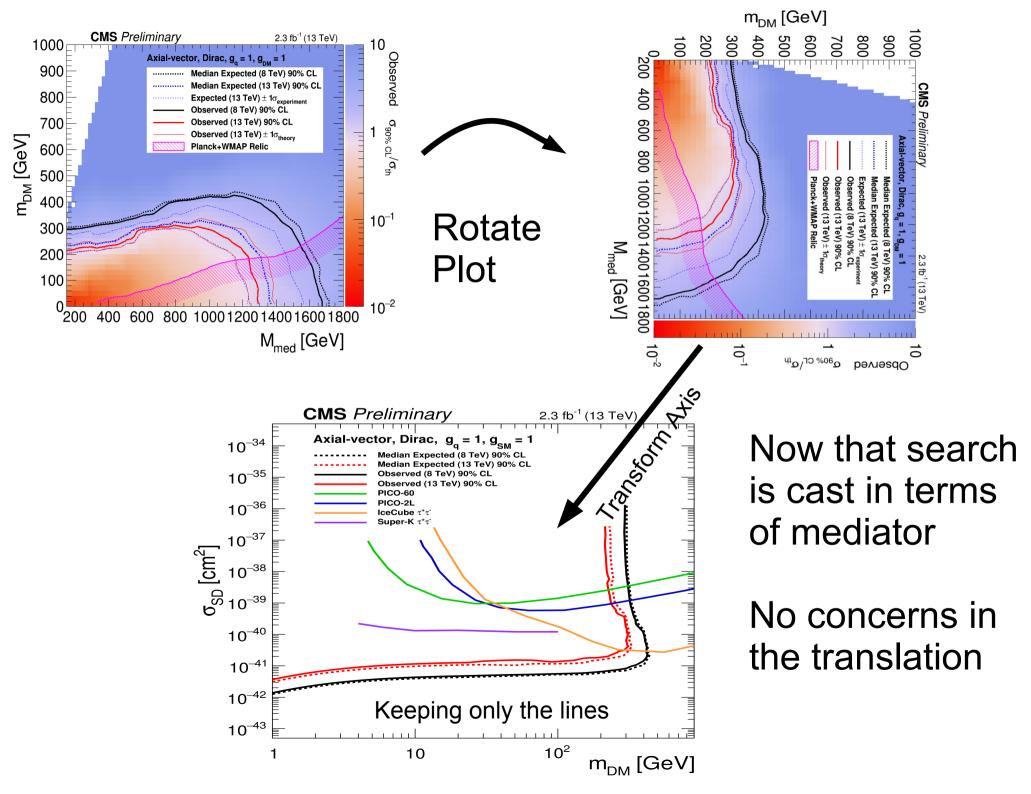
Spin 1

#### Results

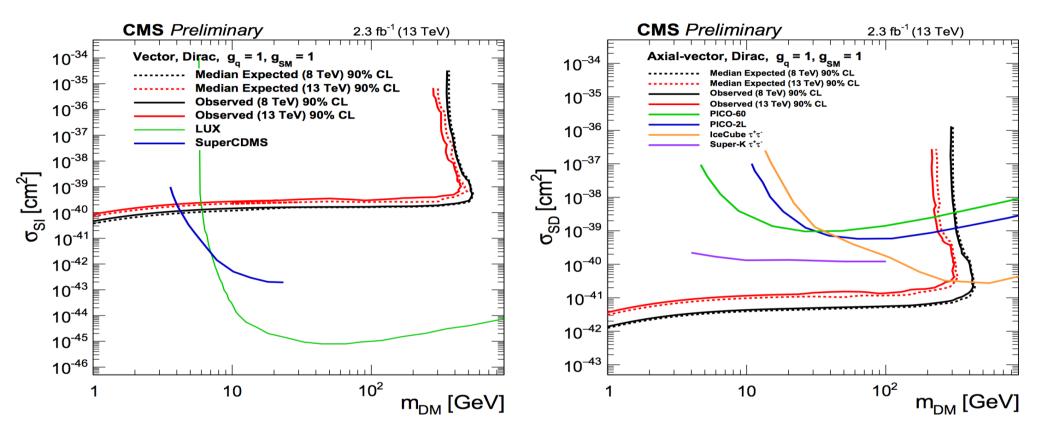


#### Both 13 TeV and 8 TeV analysis treat: mono-V and monojet on equal footing

An 1-2 excess is present in both data sets in tail



### **Our Current Public Results**

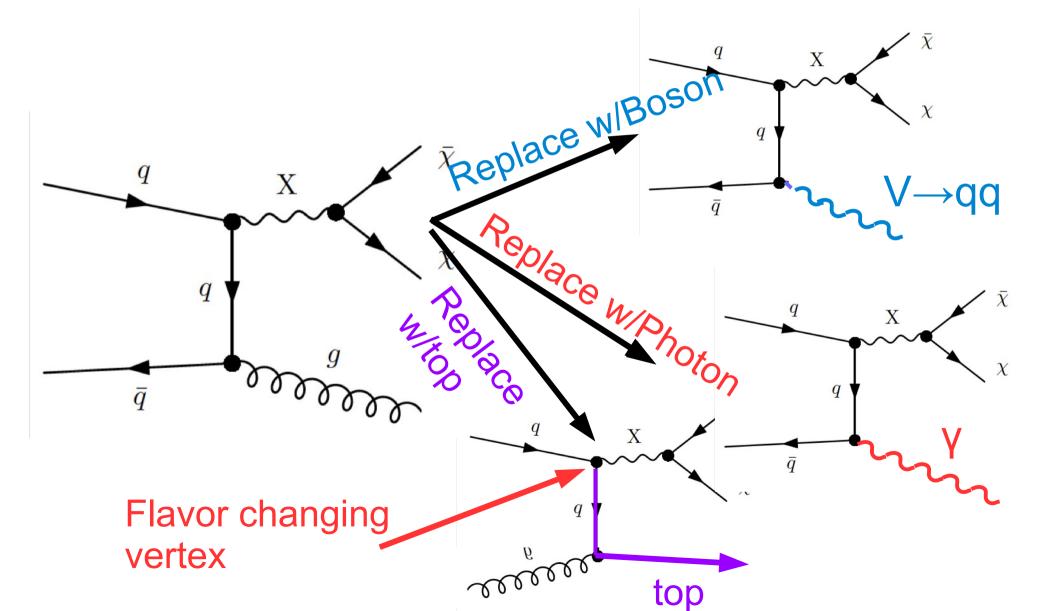


Translation to direct detection now standardized

An 1-2 $\sigma$  excess is present in both data sets in tail

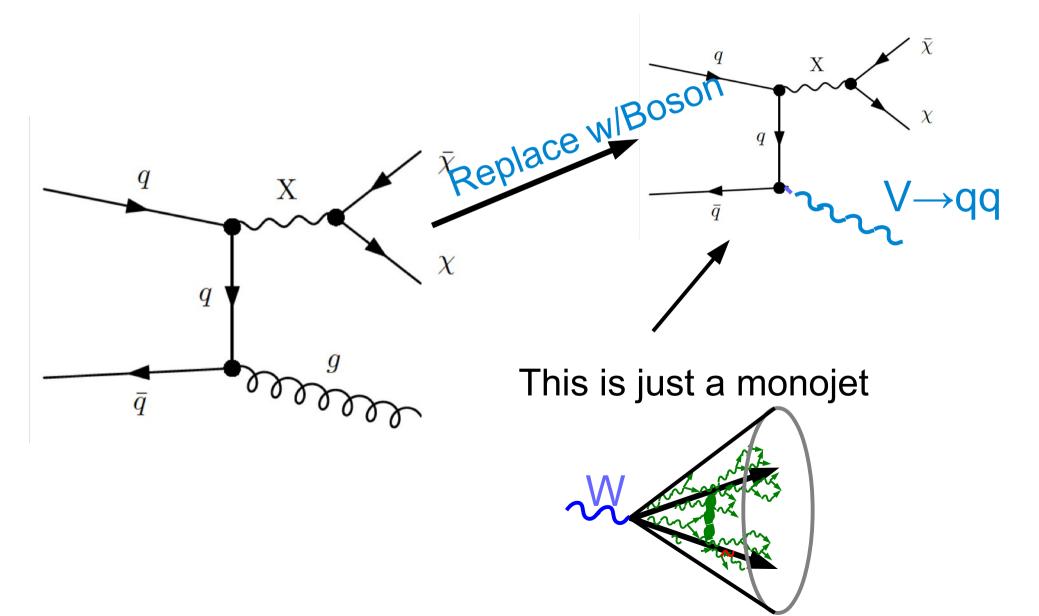
### The split in simplified model terms

• With spin 1 can generate other final states :



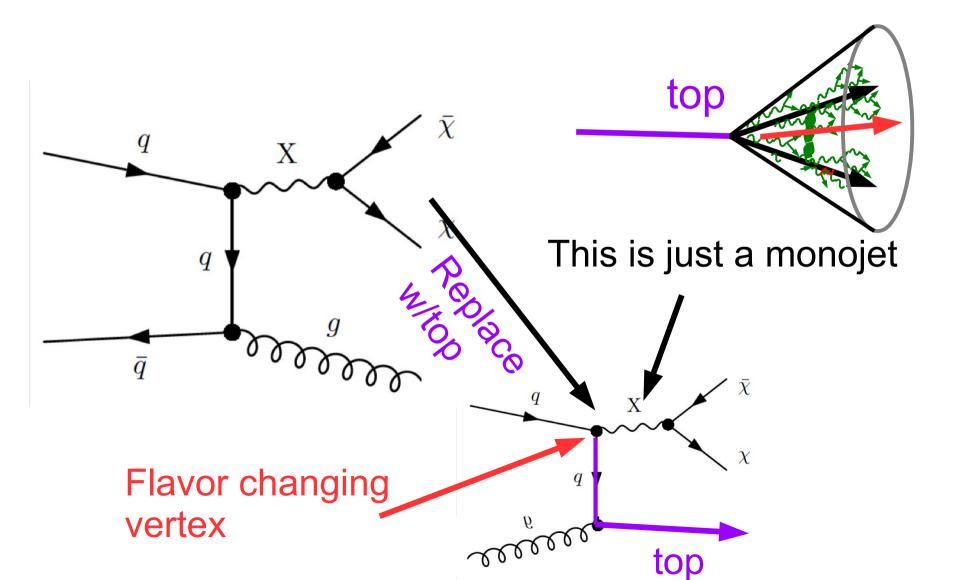
### The split in simplified model terms

• With spin 1 can generate other final states :



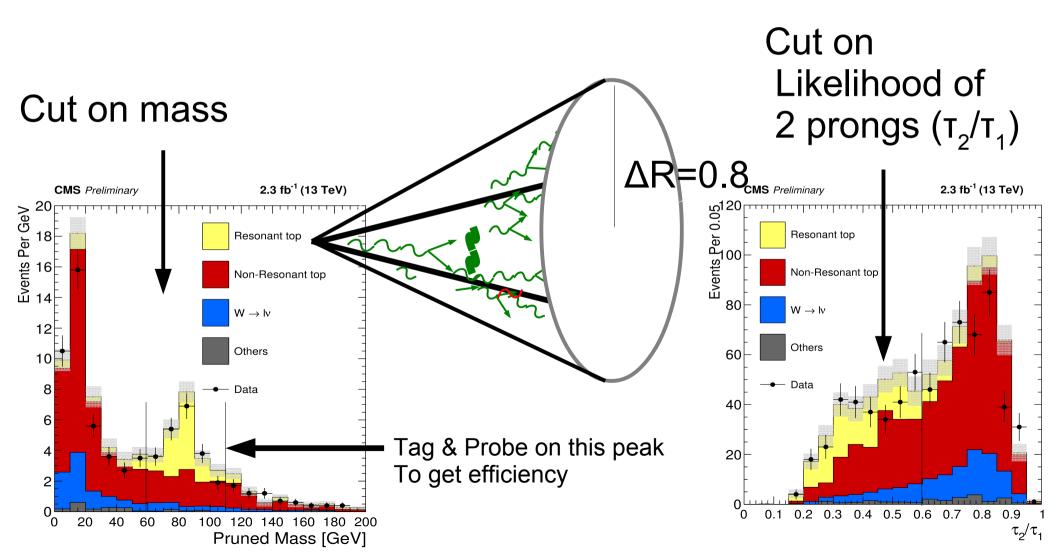
### The split in simplified model terms

• With spin 1 can generate other final states :



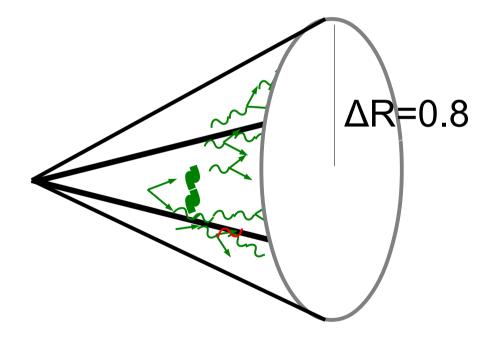
### What differentiates them?

• For both the mono-top and mono-V we tag



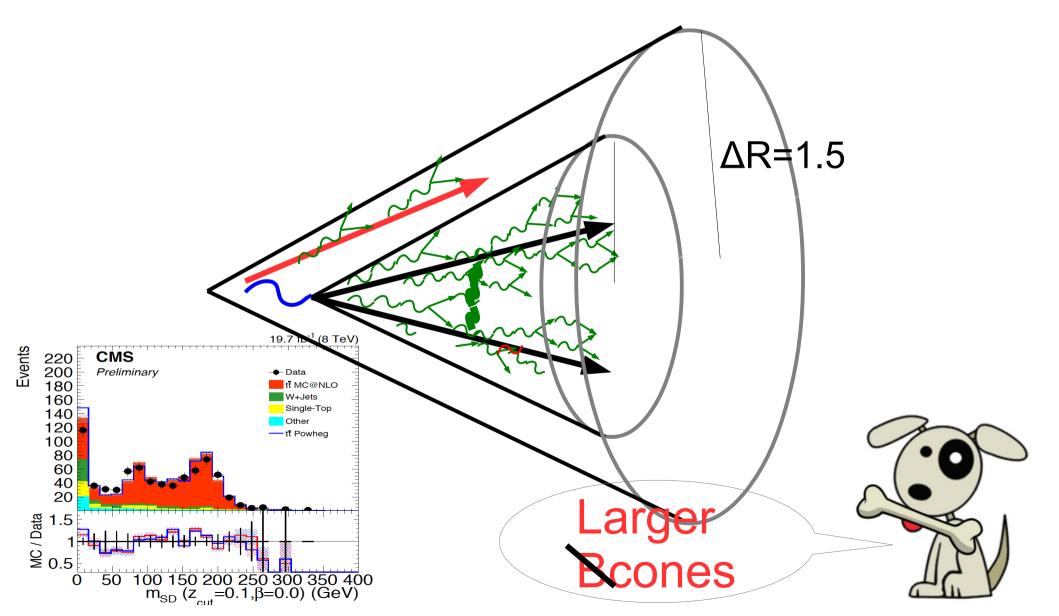
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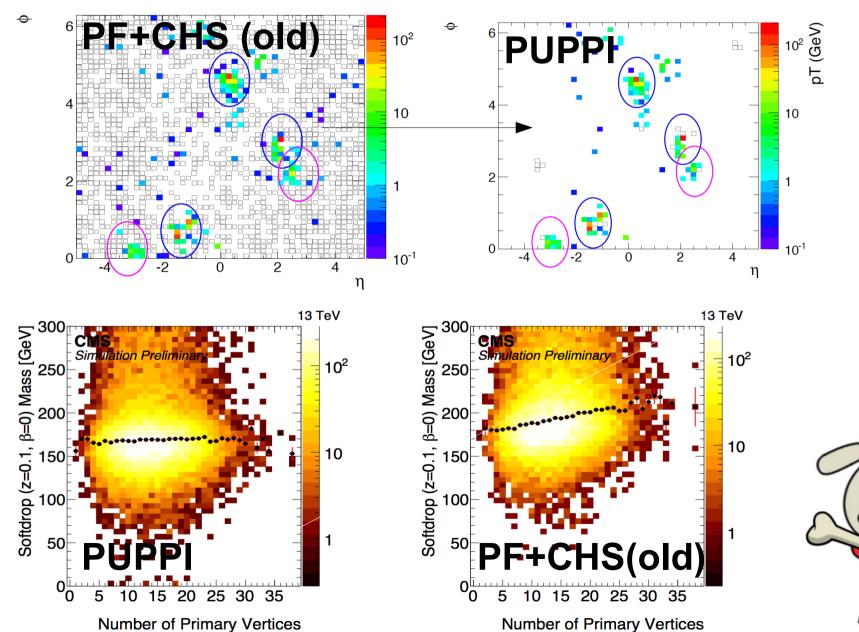
## What differentiates them?

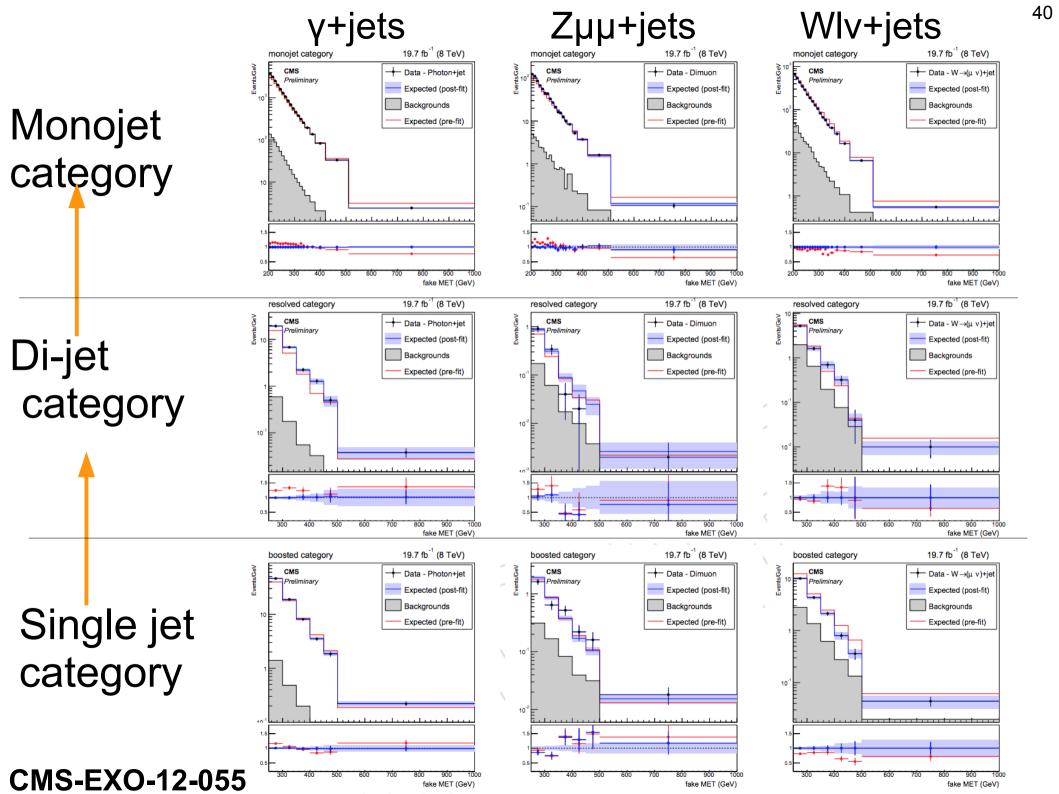
For both the mono-top and mono-V we tag



## PUPPI-ganda

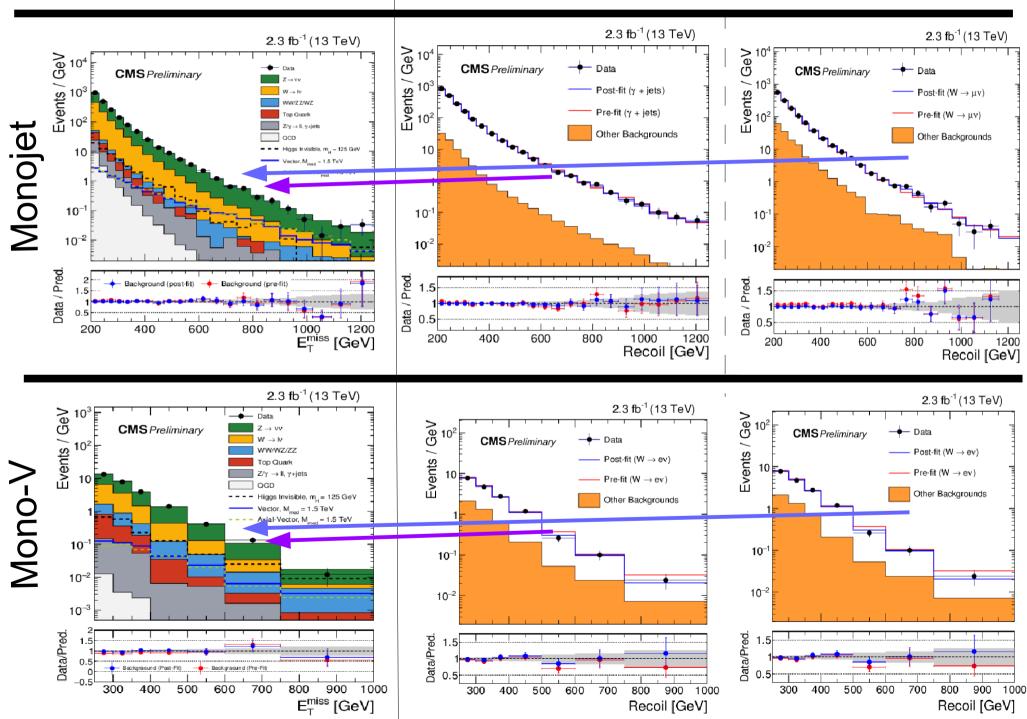
#### Key to Large Cones is PUPPI





Signal

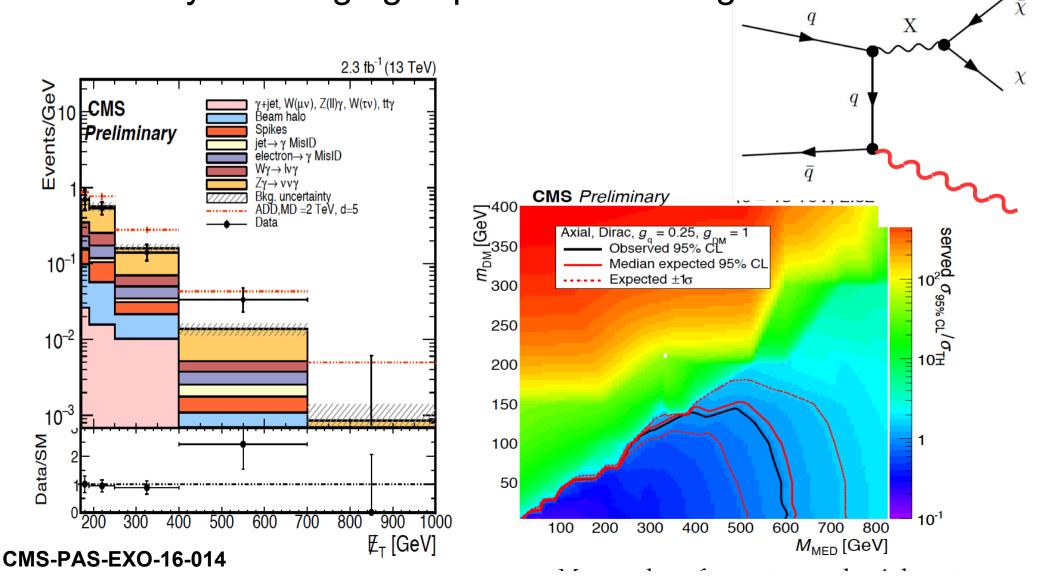
CRs: γ+jets



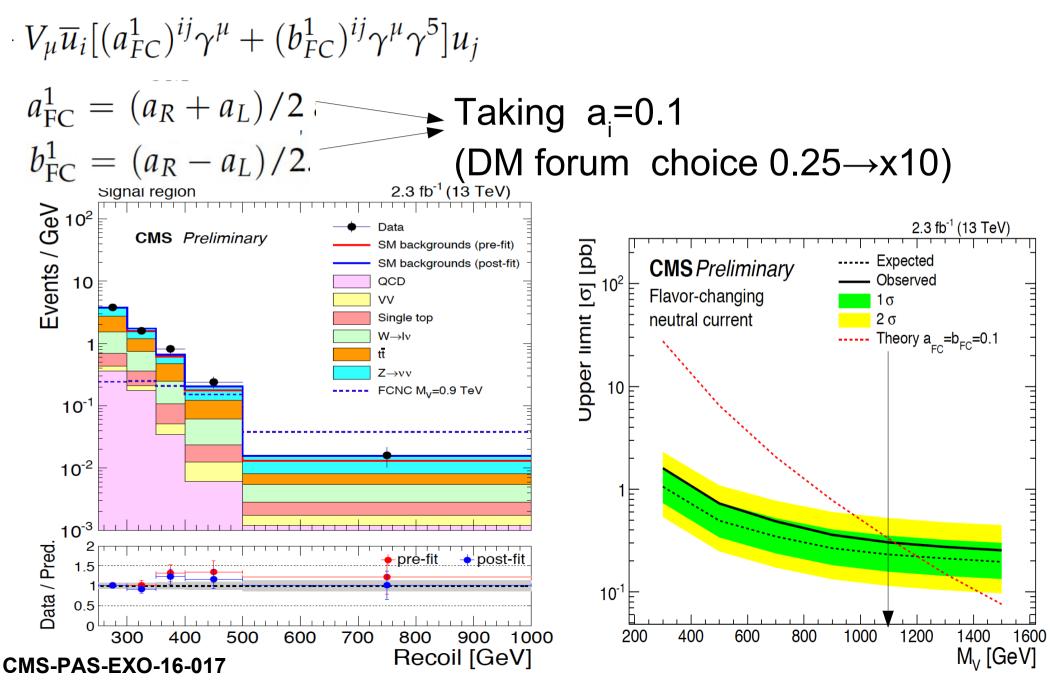
W +

## Monophoton

- Tag a photon and look for MET
  - Many challenging experimental backgrounds

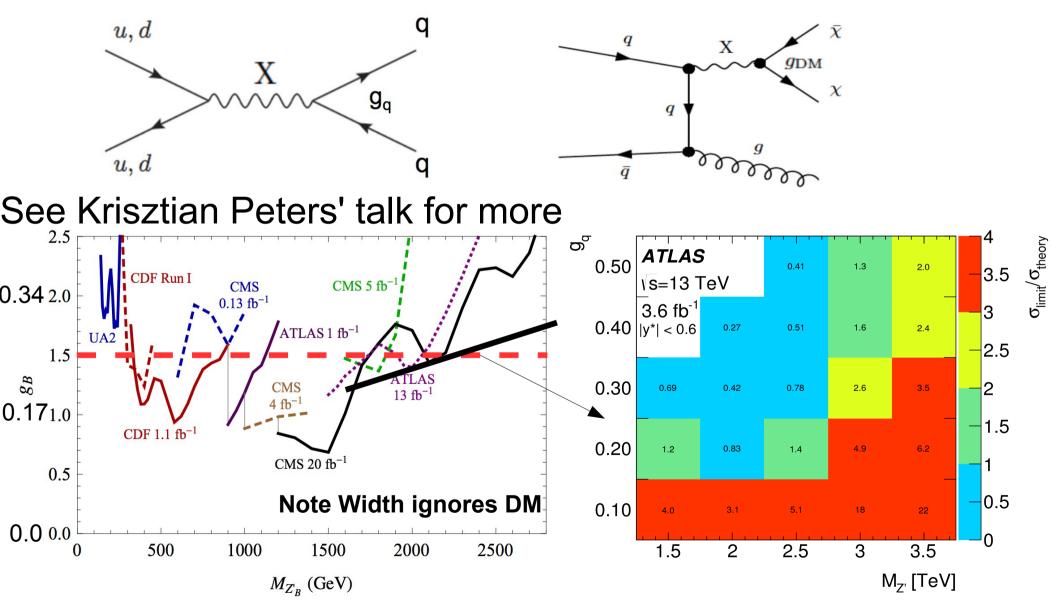


### Monotop

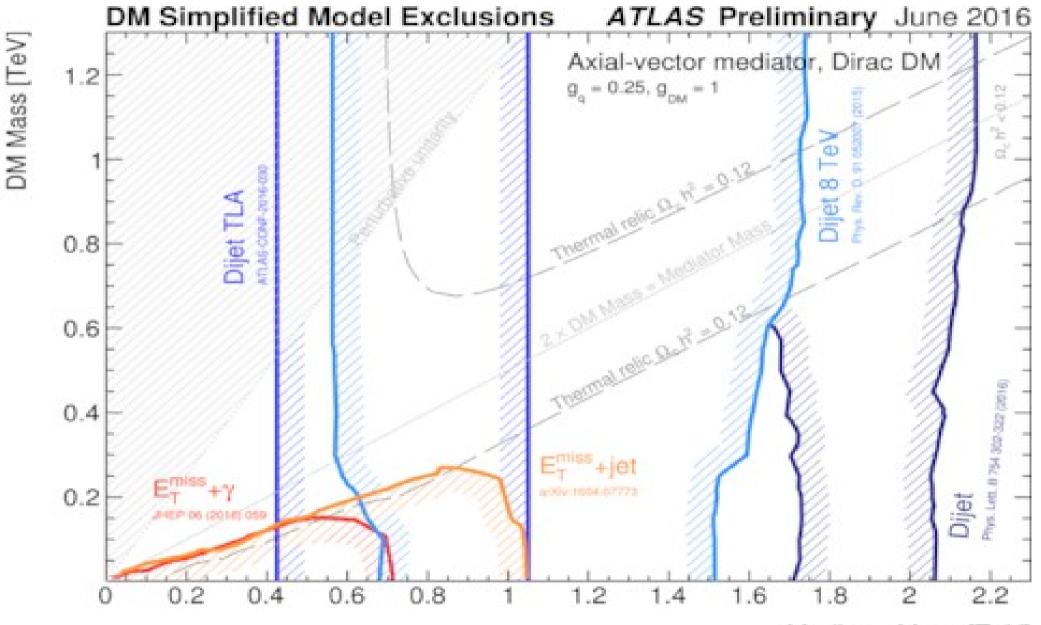


## **Mediator Search**

In addition we can just look for the mediator

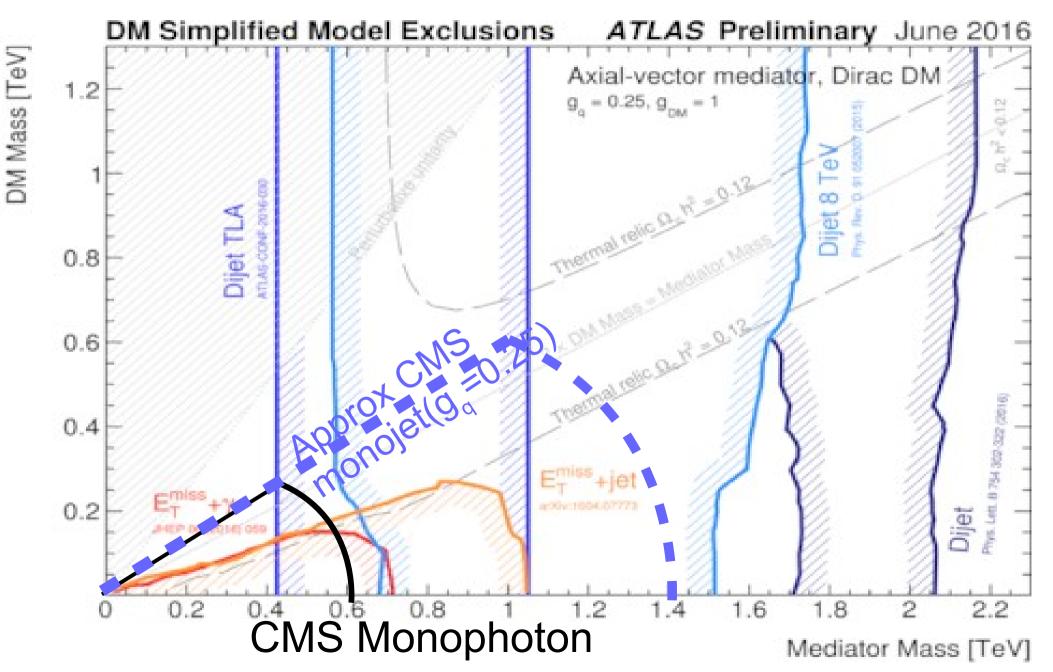


## Putting it all together

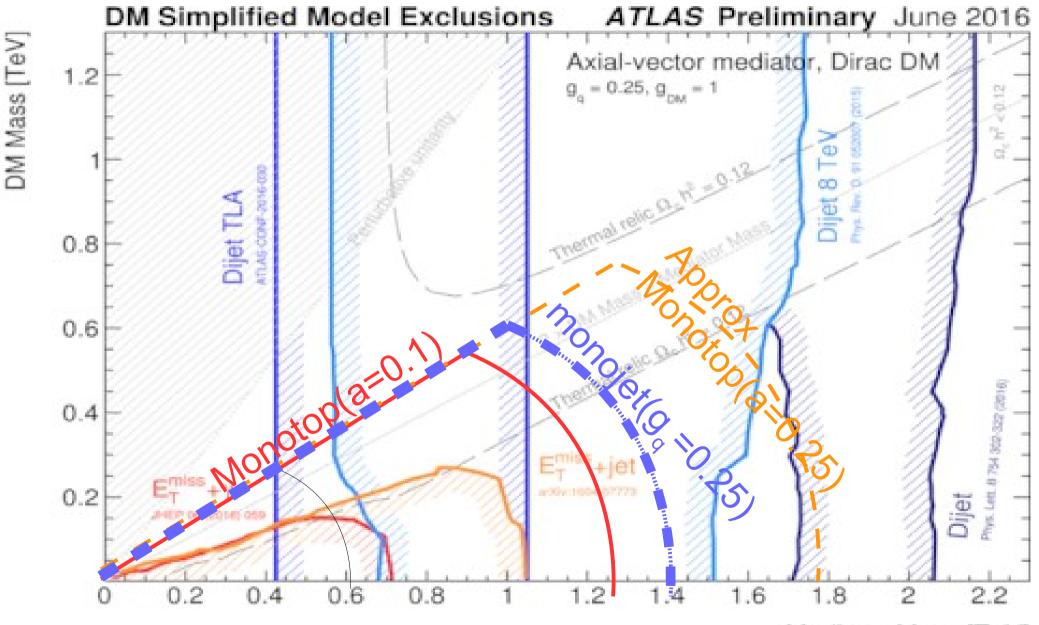


Mediator Mass [TeV]

## Putting it all together



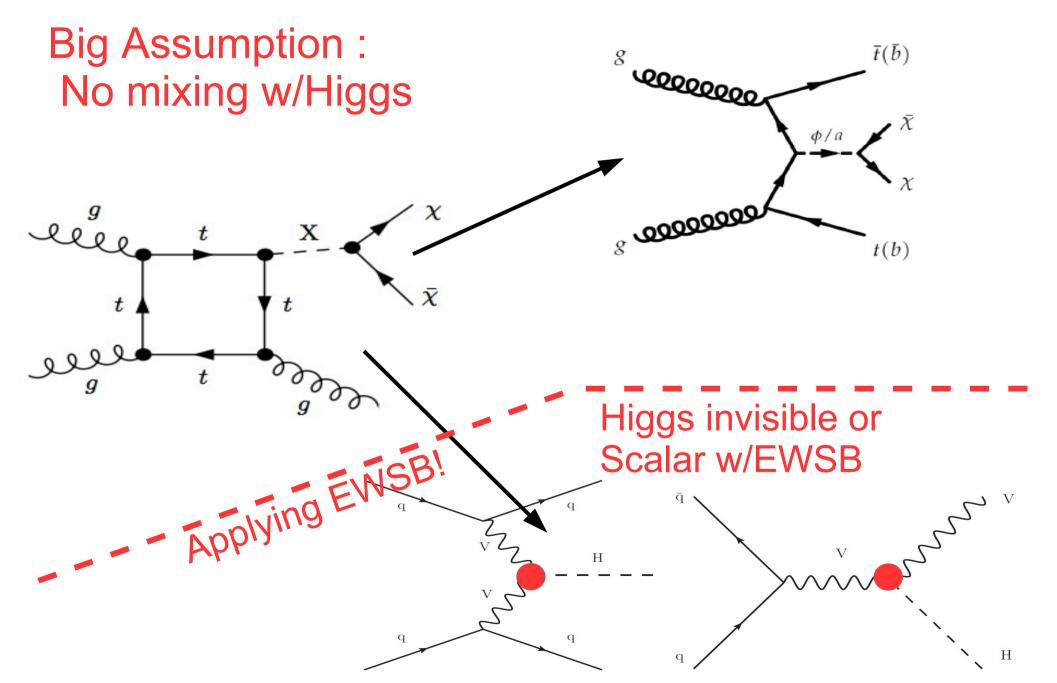
## Putting it all together



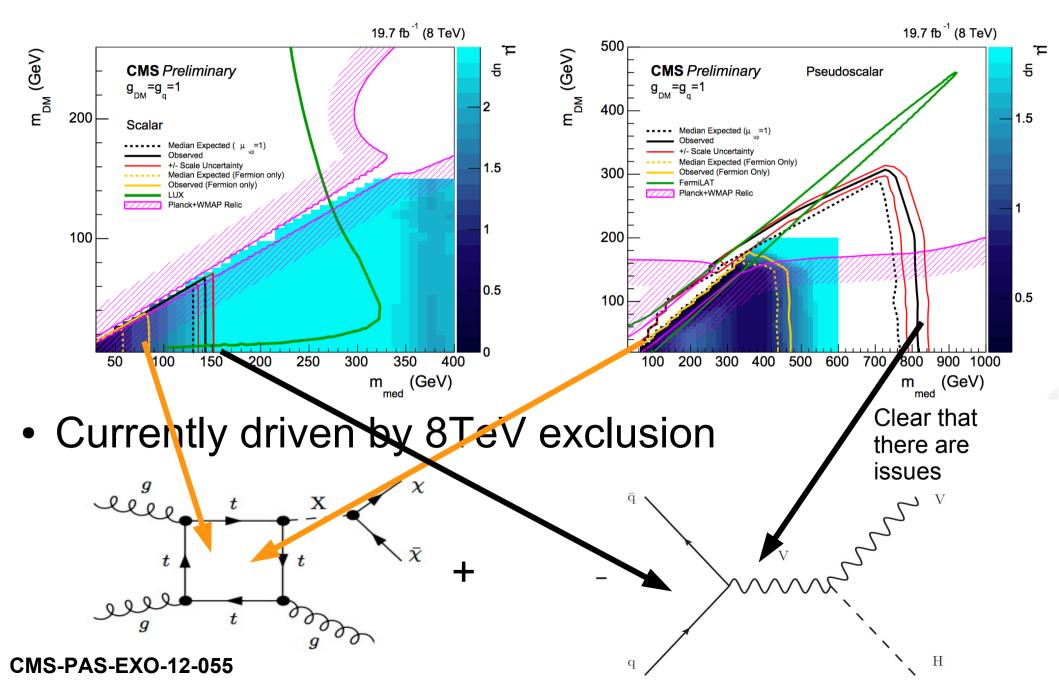
Mediator Mass [TeV]

Spin 0

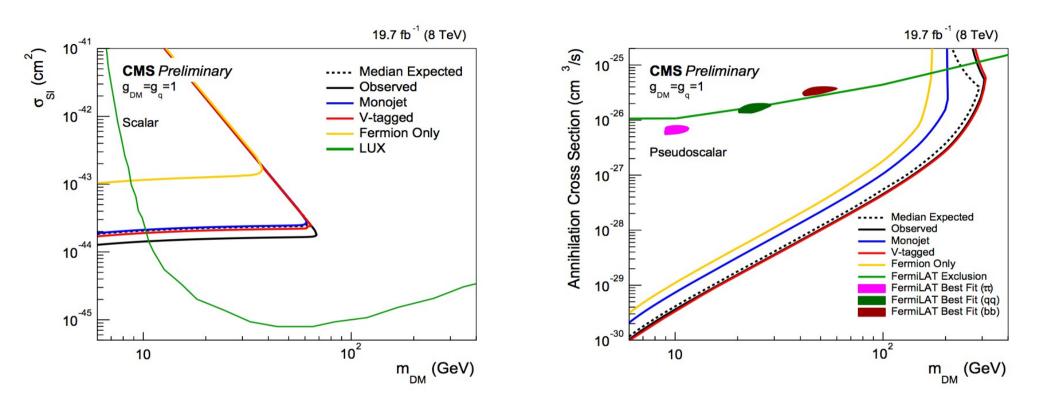
## What can you do with Spin 0?



## **Our Current Scalar & Psuedo results**



## **Our Current Scalar & Psuedo results**

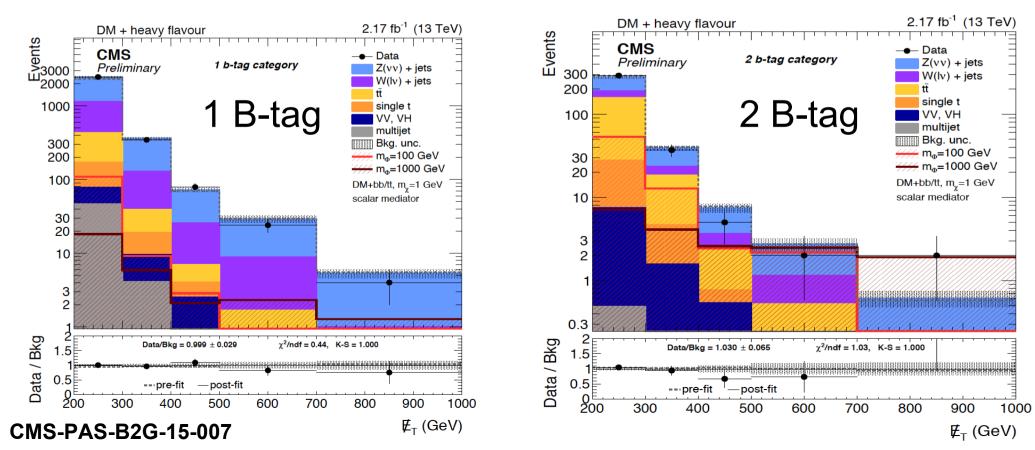


- When the dark matter is not onshell
  - Strong exclusion of pseudoscalar interpreation of LAT
  - Scalar and Direct detection are in close comopetition
    - Expect LHC to pass LUX this summer

CMS-PAS-EXO-12-055

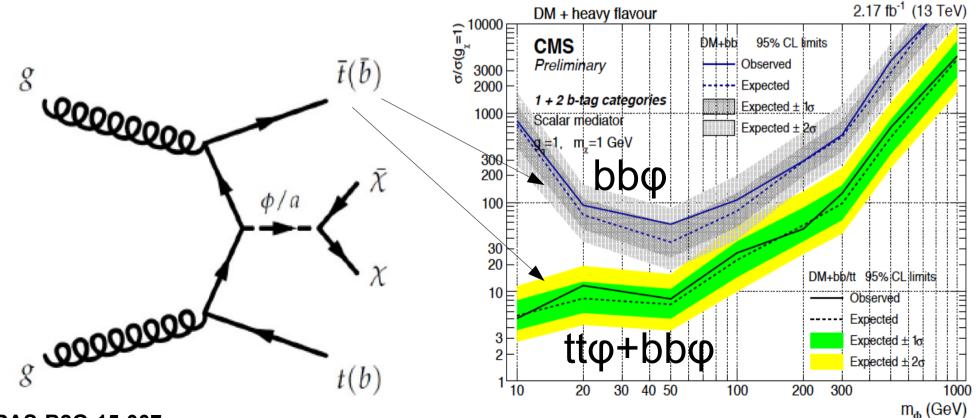
## Heavy Flavor

- Mono-B or B(s)
  - Require less than 4 jets
  - Basically the monojet analysis with either 1 or 2 bs
  - Inject both tt+DM and bb+DM into the analysis



## Heavy Flavor Results

- Mono-B or B(s)
  - Note that this is only < 4 jets</li>
  - Inject both tt+DM and bb+DM into the analysis



CMS-PAS-B2G-15-007

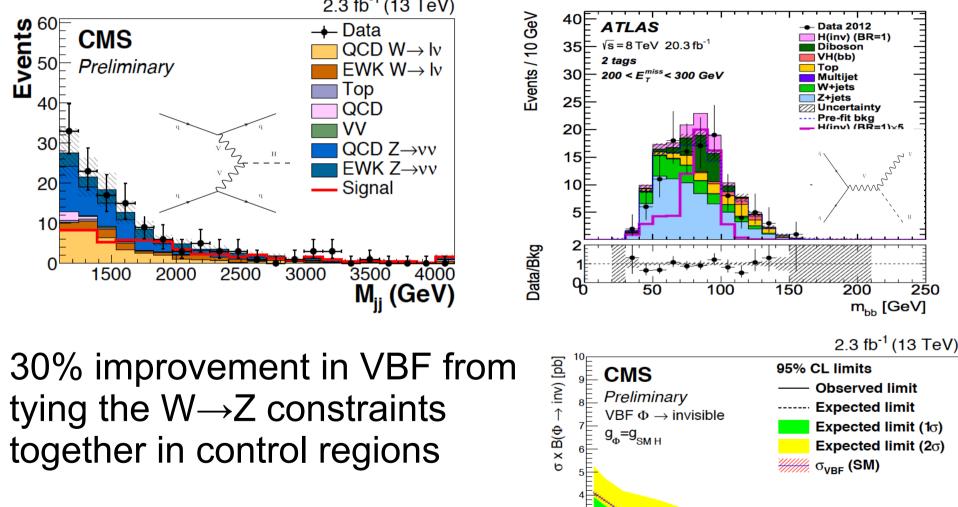
## Higgs Invisible Interpretation

200

300

400

• Higgs Invisible is scalar model with EWSB



Starting to scan mass-

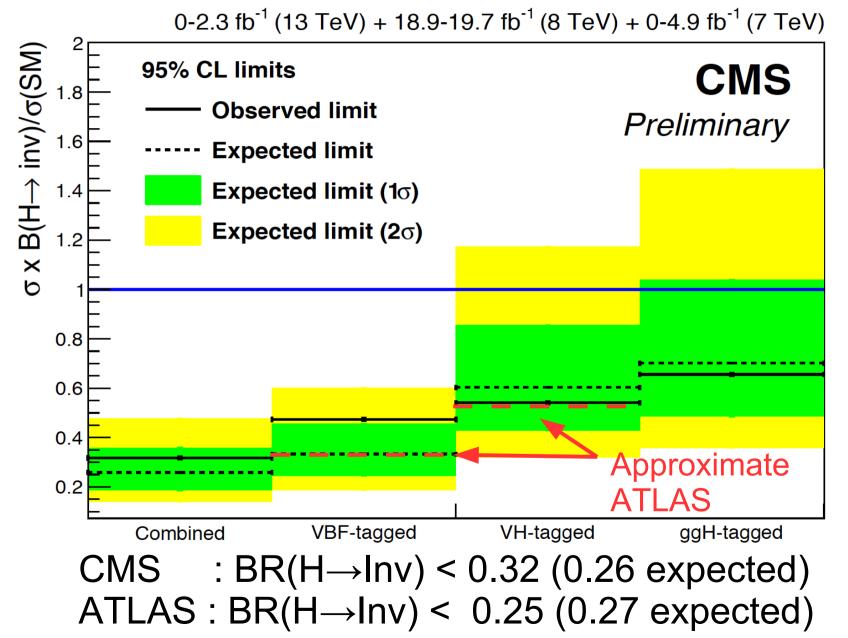
CMS-PAS-HIG-16-009/ATLAS-HIGG-2015-03

600

m<sub>o</sub> [GeV]

500

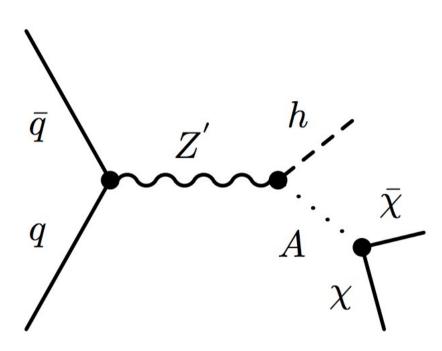
## Results



CMS-PAS-HIG-16-009/ATLAS-HIGG-2015-03

## Mono-Higgs

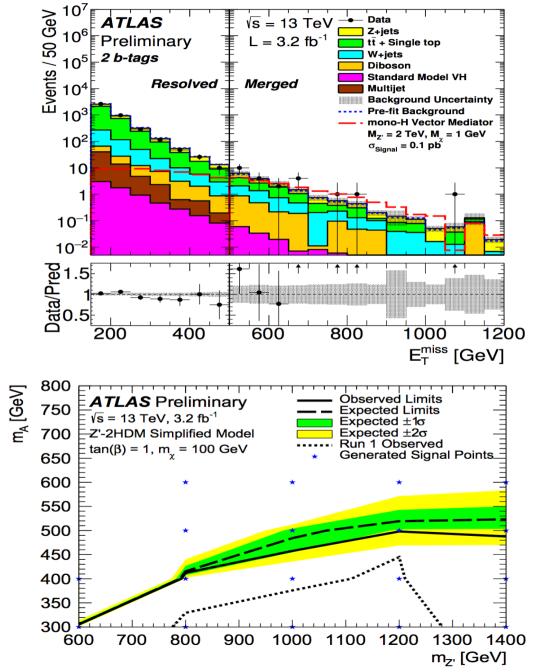
#### Adding Spin 1 and Spin 0 mediators



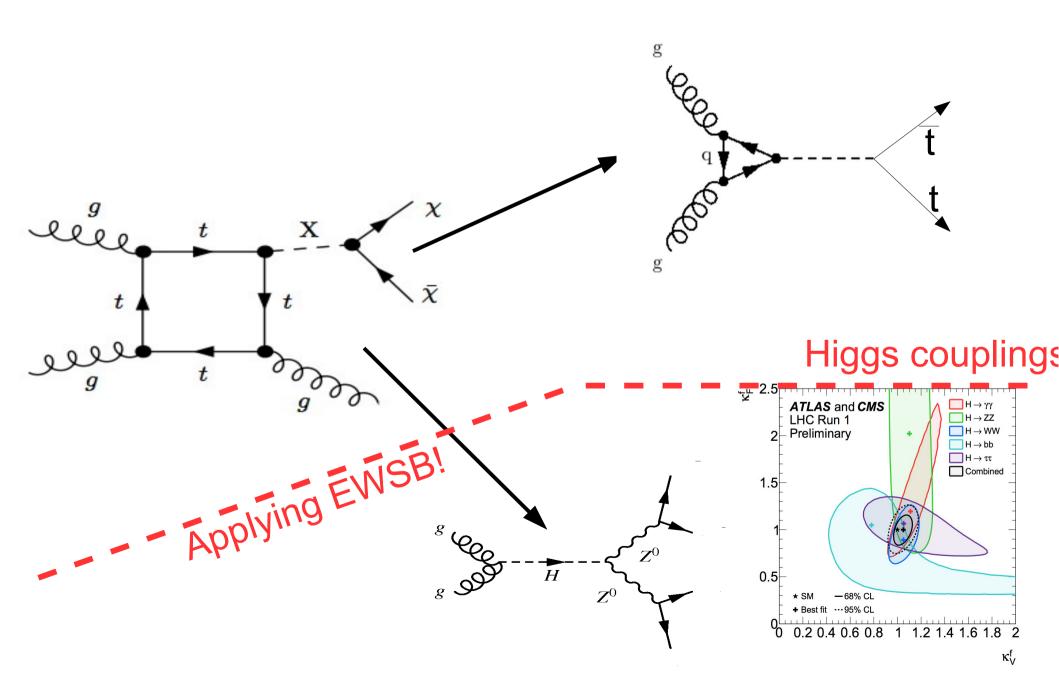
h→bb bounds drive mono-Higgs

B-tagging forces ttbar background to drive analysis

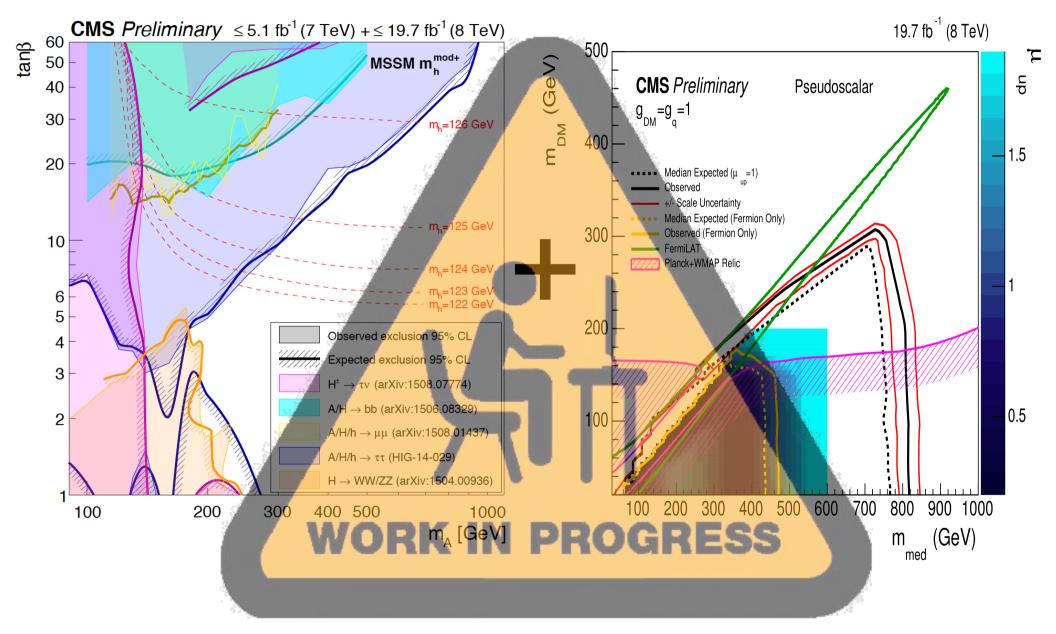
ATLAS-CONF-2016-019



## What about the visible?



## Not yet available



Spin 0 Di-photon ( 미안해요 )

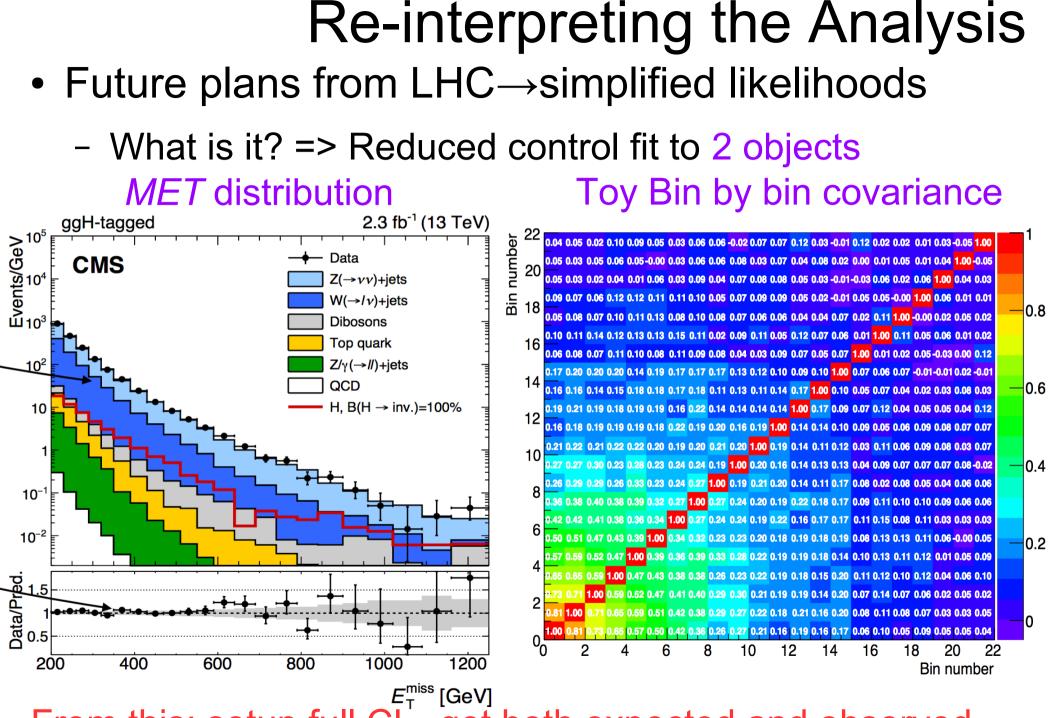
Preliminary Work with U. Haish, O. Buchmuller, K. Hahn, N. Wardle, T. Du Pree

## The simplest DM model

- Lets try to make something super basic
  - Basic model

$$\mathcal{L}_{S} = g_{\chi} S \bar{\chi} \chi + \frac{\alpha_{s}}{4\pi} \frac{c_{g}}{\Lambda} S G^{a}_{\mu\nu} G^{a \,\mu\nu} + \frac{\alpha}{4\pi} \frac{c_{\gamma}}{\Lambda} S F_{\mu\nu} F^{\mu\nu}$$

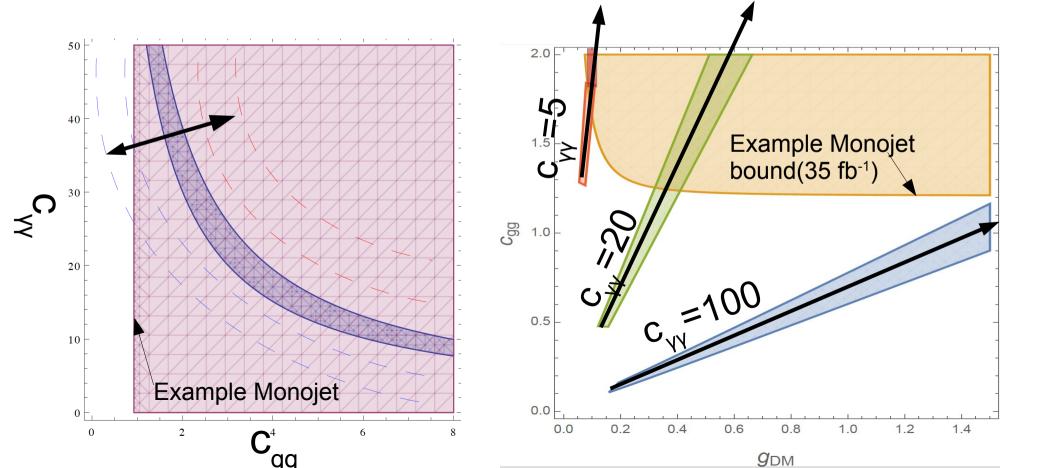
$$??$$

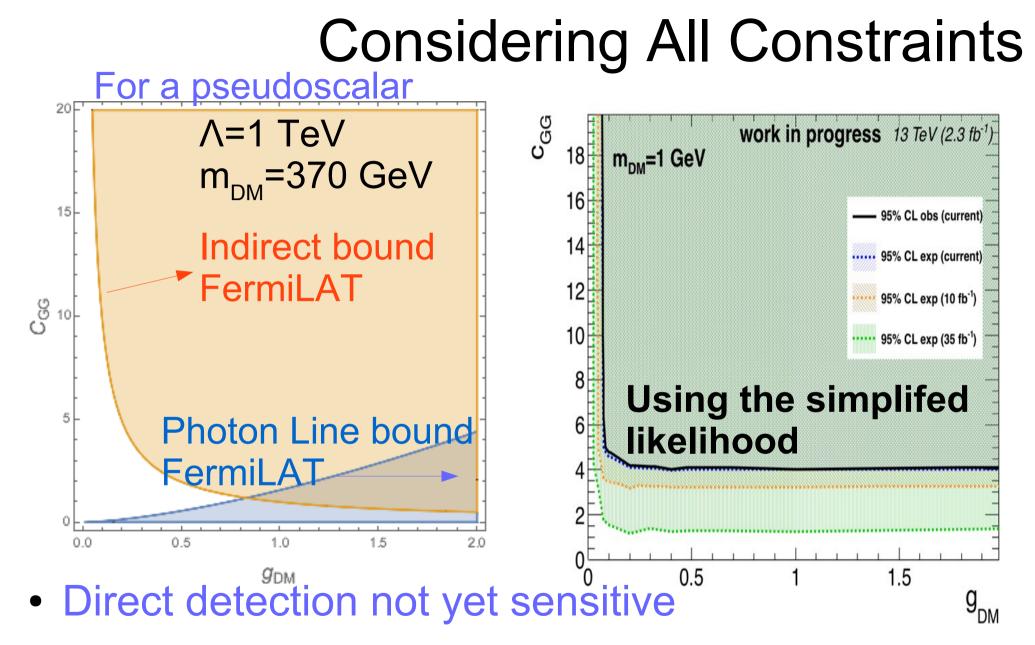


From this: setup full CL<sub>s</sub> get both expected and observed

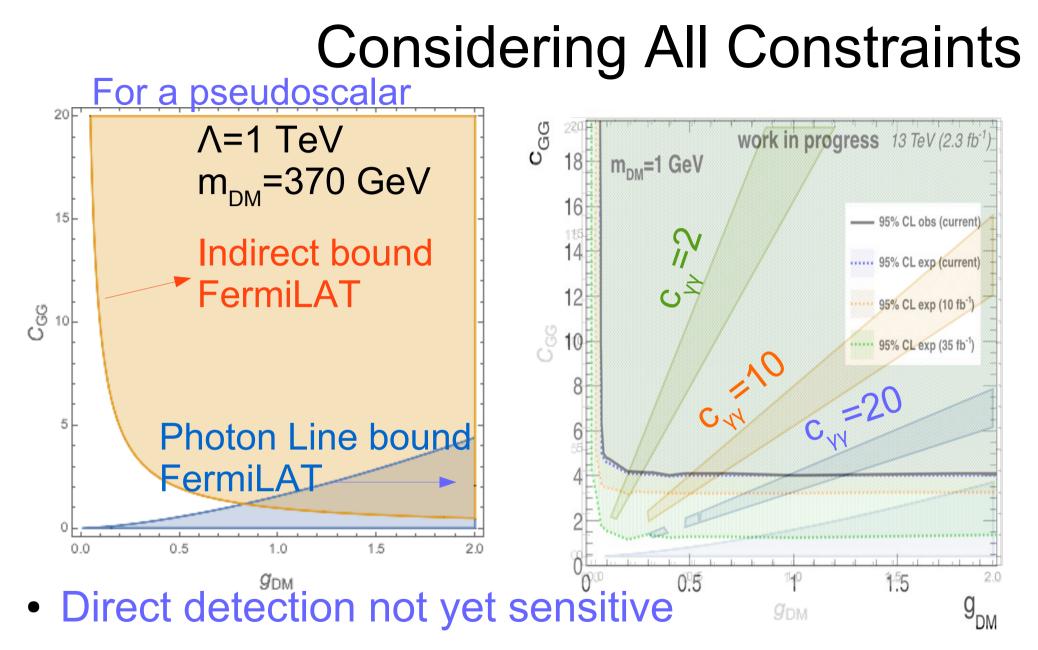
## Using the best fit cross section

- We have 3 free couplings :
  - $g_{DM}, C_{GG}, C_{\gamma\gamma}$
- Taking the photon best fit can constrain one





- Indirect detection limits on-shell production
- Photon Line bounds limit photon coupling < 100



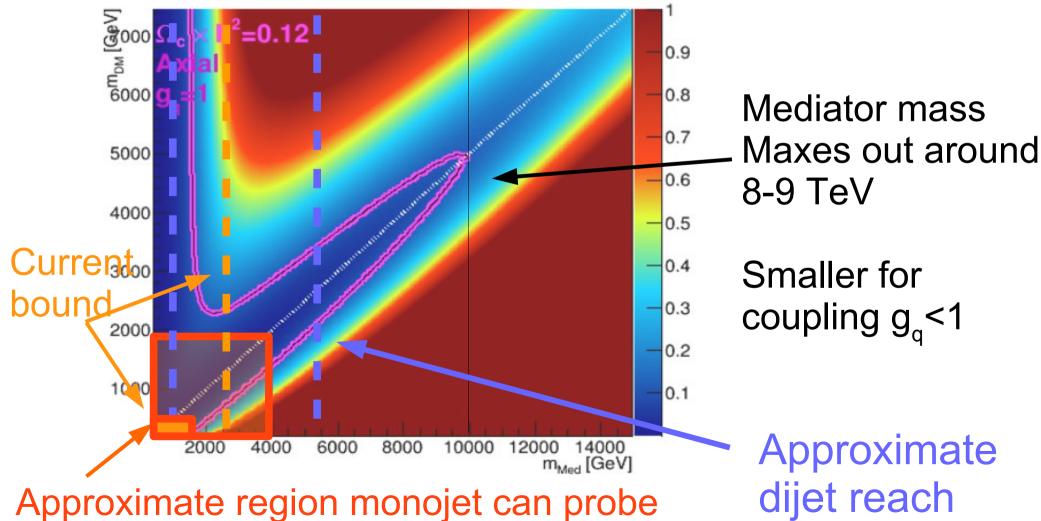
- Indirect detection limits on-shell production
- Photon Line bounds limit photon coupling < 100

# Outlook

Hep-ph/1603.08525 Hep-ph/1509.02904 ATL-PHYS-PUB-2015-004

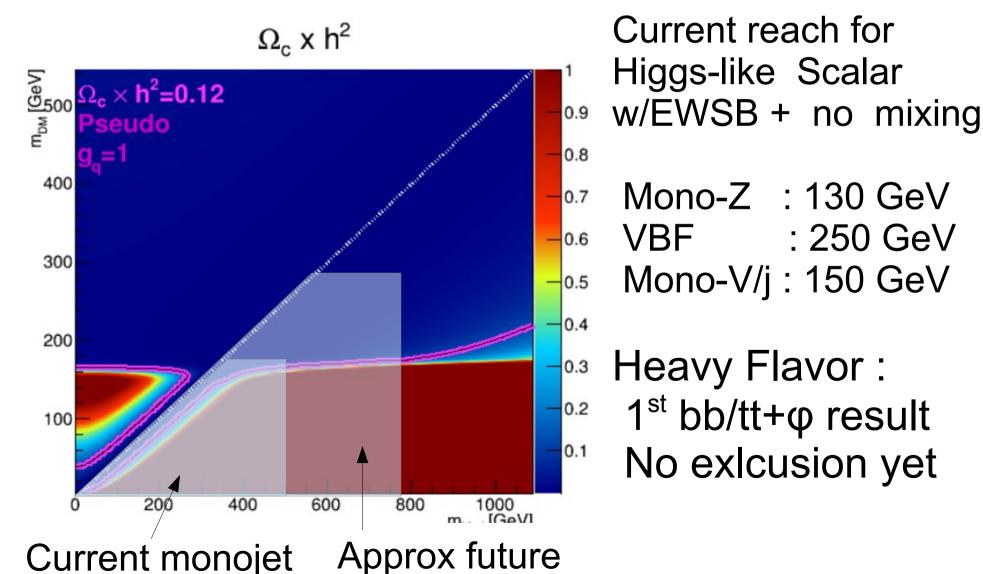
Outlook

- Spin 1 :
  - Dijet and monojet will continue to push out the bounds  $\Omega_c \ge h^2$



Hep-ph/1603.08525 Hep-ph/1509.02904 EXO-12-055 HIG-16-012

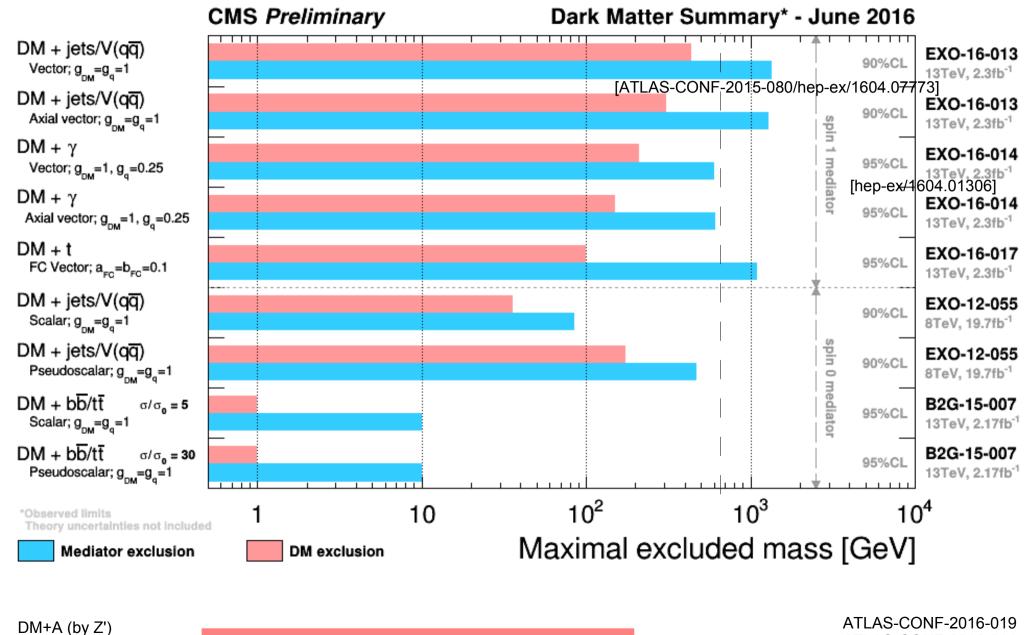
- Spin 0 :
  - Yet to truly coalesce in 13 TeV



67

Outlook

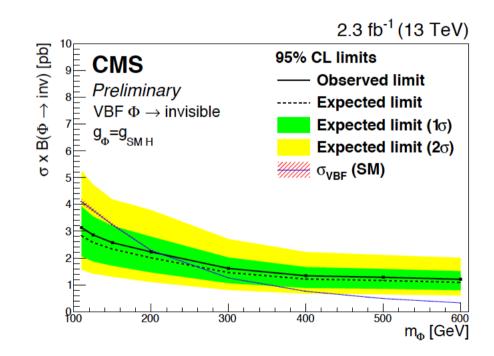
## Summary



Pseudoscalar ( $g_{DM} = g_{a} = 1$ )

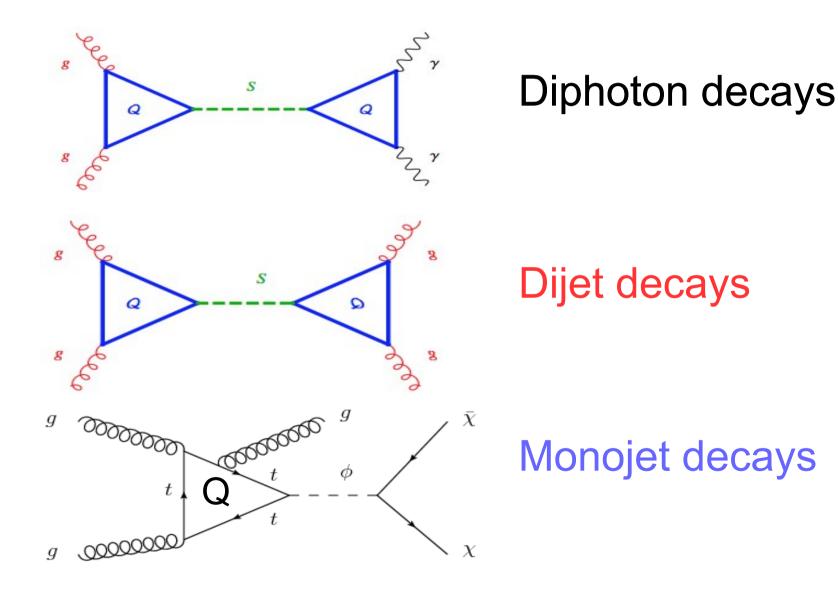
ATLAS-CONF-2016-011

# Thanks 강사합니다



## What are the decays

• We only really have a few decays:



## What goes into this?

• To find a signal we look for high MET :

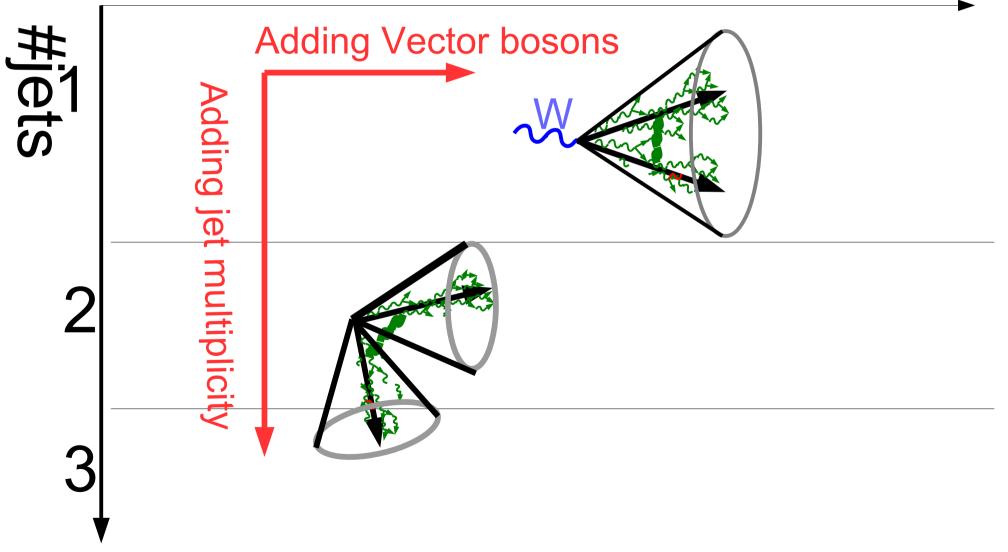
Modeling of production mode is needed (HO corrections)

 $MET = -\Sigma_{All \text{ particles}} p_T$  $^{\mu}MET(Z \rightarrow vv) = - Z \text{ recoil} + p_T(vv)$ 

Modelling of the calorimeter response and resoltuion

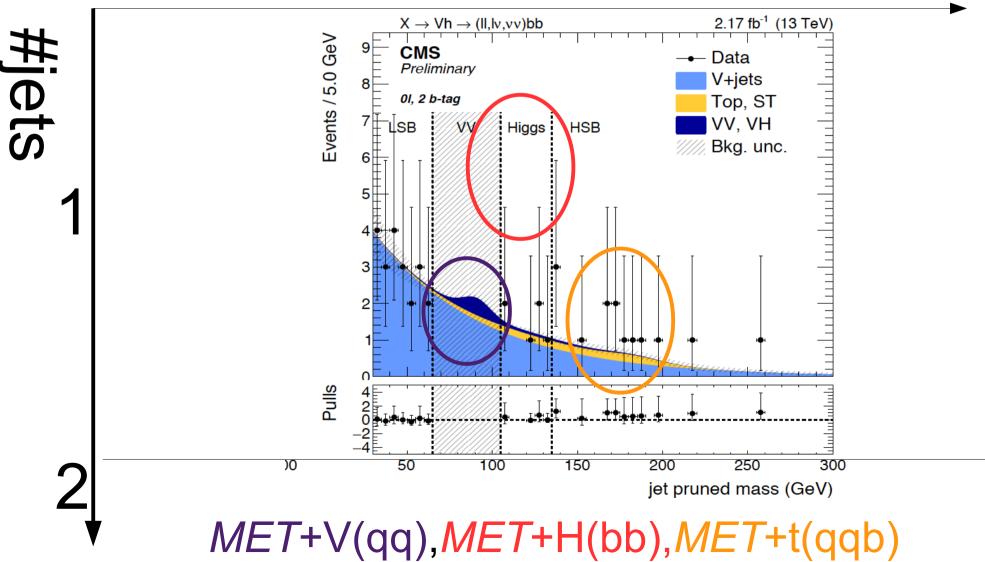
#### **Monojet Extensiion Plane**

#### Jet Mass



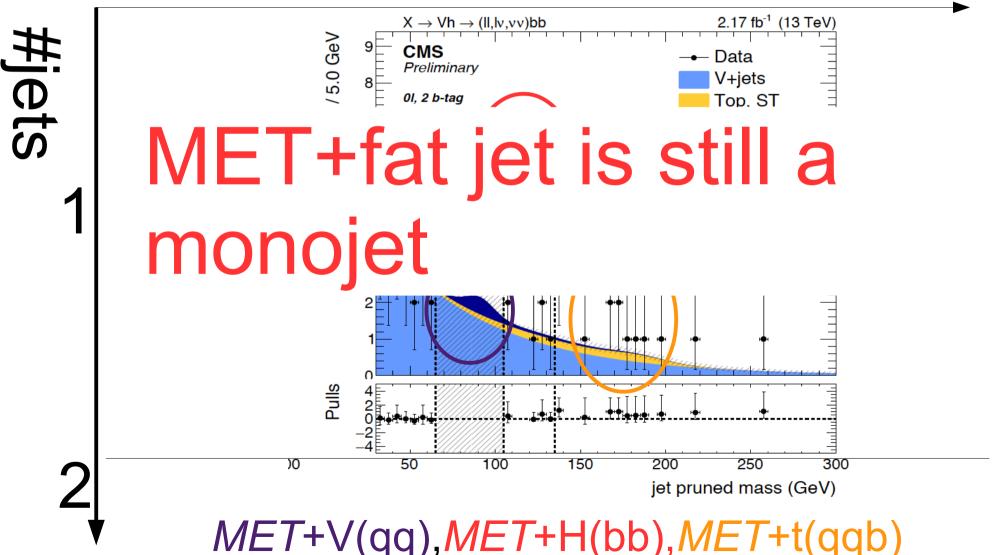
#### Monojet Extension #1 ( $V \rightarrow qq$ )

#### Jet Mass



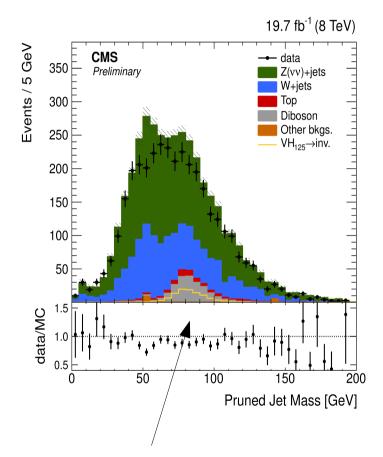
#### **Beyond Monojet**

#### Jet Mass



#### MET + fat jet

• There is no clean way to separate fat jets form jets



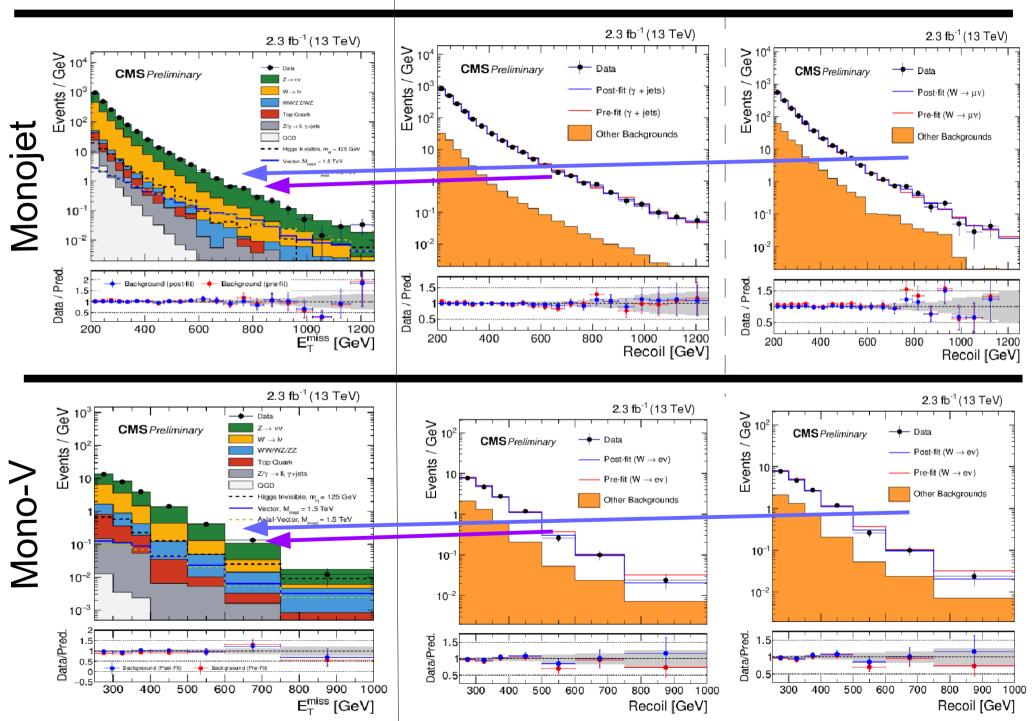
Is there room for improvement?

Yes

Currently require a simple : jet mass cut + τ2/τ1

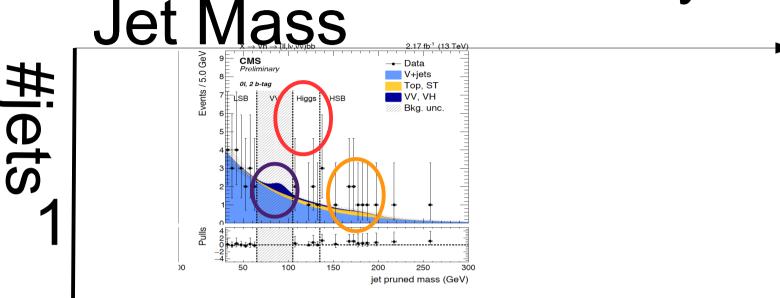
Signal

CRs: γ+jets



W +

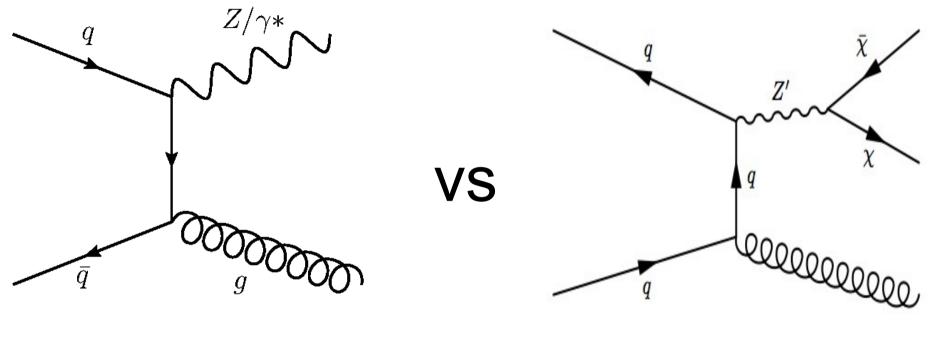
#### **Beyond Monojet**



2

# Using the 2<sup>nd</sup> jet or more can add to discrimination

#### Where do we gain from 2<sup>nd</sup> Jet?

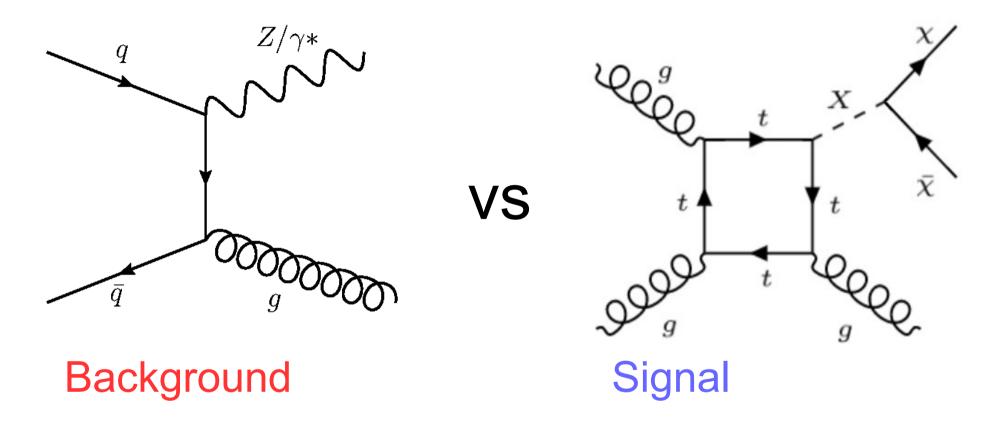


Background

Signal

For Vector and Axial mediators not much Only real difference is mediator mass

#### Where do we gain from 2<sup>nd</sup> Jet?



For Scalar and Pseudoscalar mediators more Now the production modes are different

In addition to 2<sup>nd</sup> jet can consider a quarkgluon discriminator

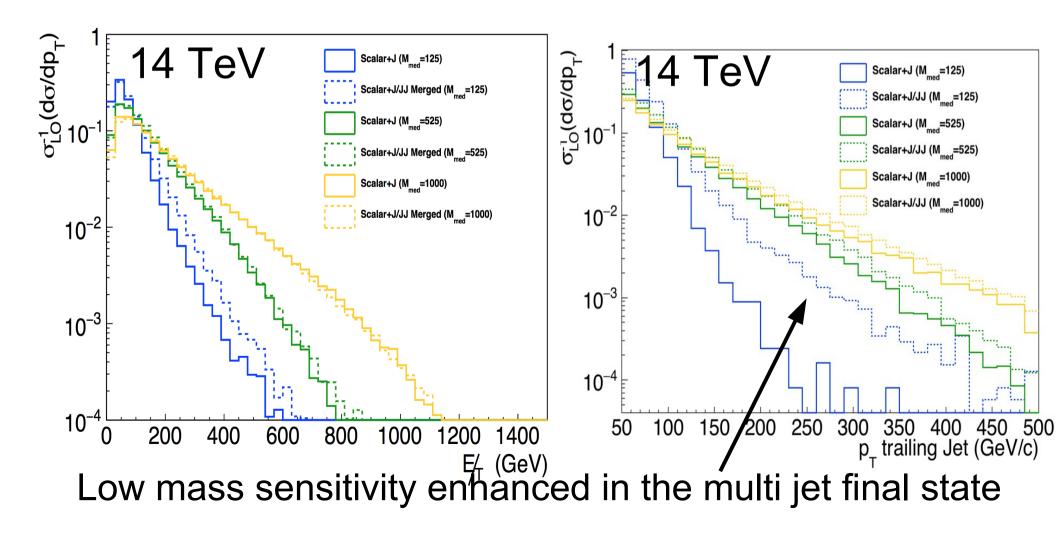
#### Spectrum of Signal MCs

Sample	LO/ Leading Loop	LO in 2j	NLO,1,2j
Vector/ Axial	Madgraph MCFM	Powheg	aMC@NLO
Scalar/ Pseudoscalar	Powheg MCFM aMC@NLO	VBF@NLO aMC@NLO	

aMC@NLO+MG get highest order 1/2 jets merged

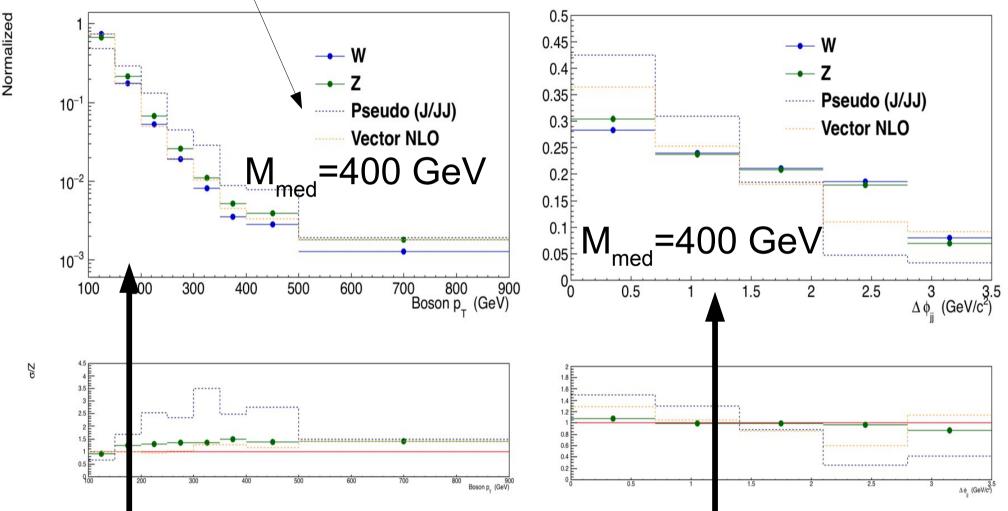
#### Advantage of merged MC

- Taking advantage of the new technology
  - Can consider exploring new regions of phase space



#### FYI aMC@NLO merged 0,1,2jet pseudoscalar

#### **Basic Concept of Gains**

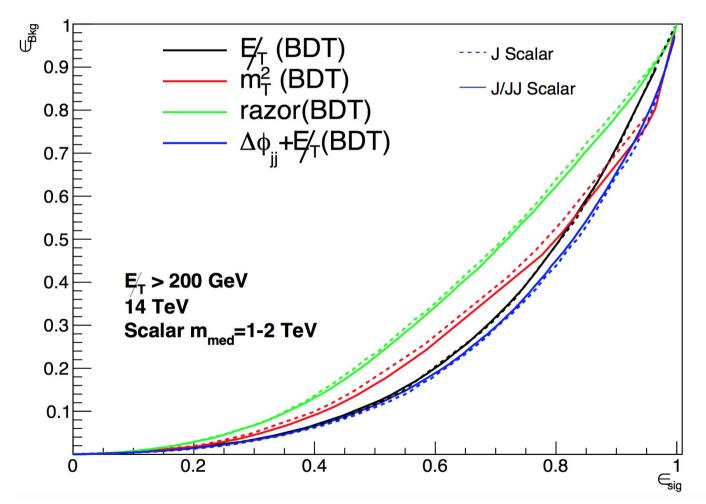


Gluon fusion induces Higher pT spectrum

Heavier mediator forces jets to be closer

#### How do the single variables perform?

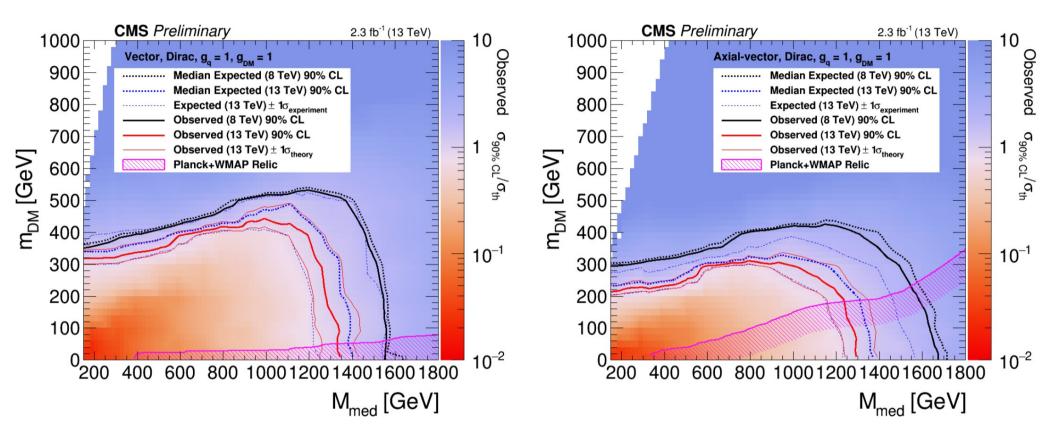
Comparison of single variables



Gain comes from fact that light mass objects have collinear jets Using  $\Delta \phi_{ii}$  can bring as much as 20% gain

# Results

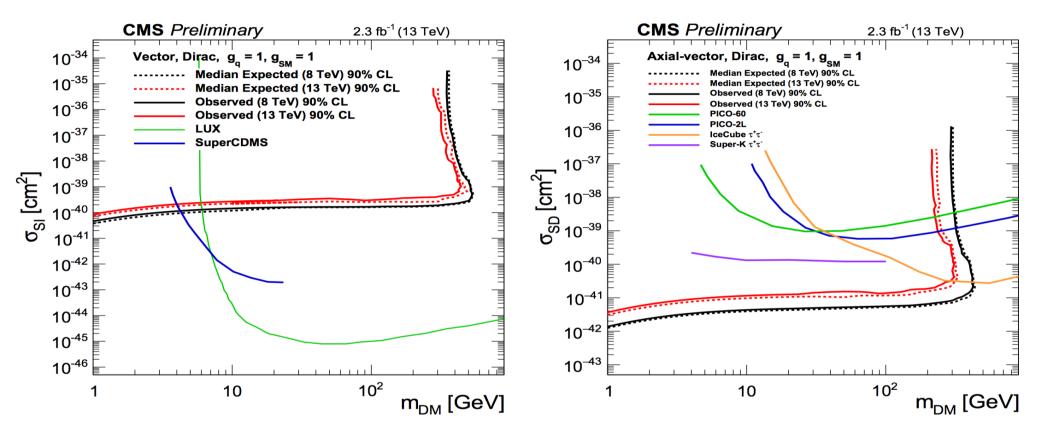
#### **Our Current Public Results**



#### Both 13 TeV and 8 TeV analysis treat: mono-V and monojet on equal footing

An 1-2 excess is present in both data sets in tail

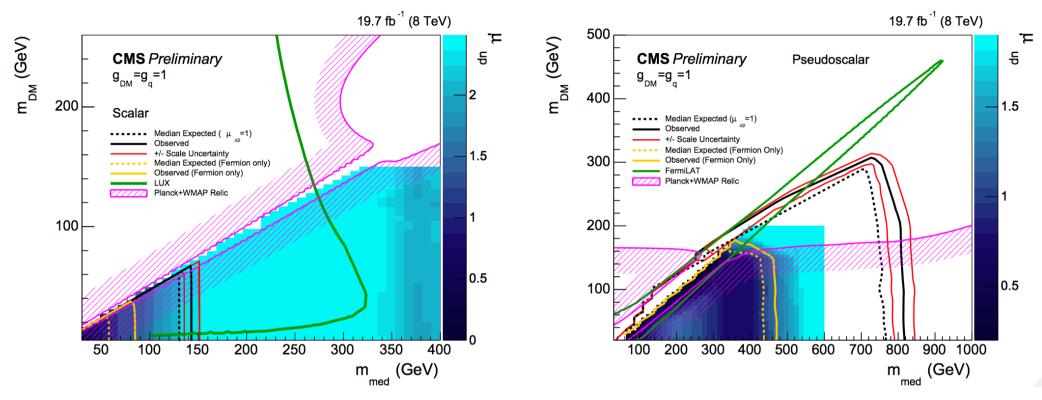
#### **Our Current Public Results**



Translation to direct detection now standardized

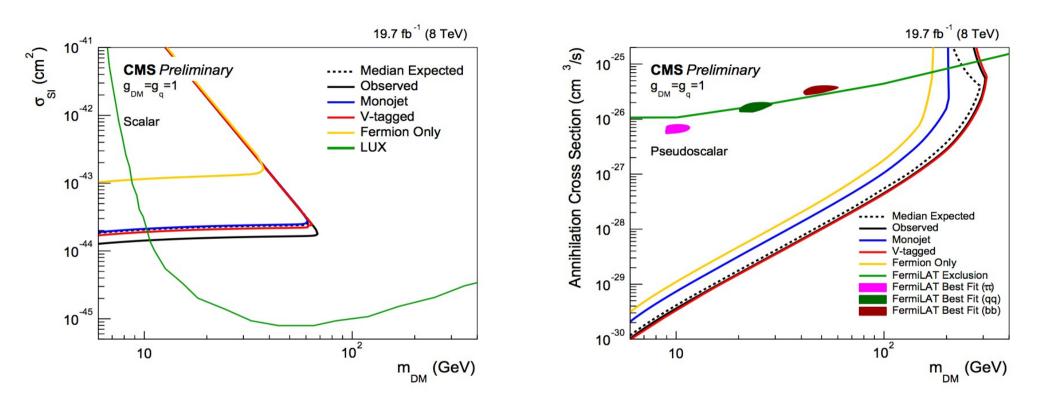
An 1-2 $\sigma$  excess is present in both data sets in tail

#### **Our Current Scalar & Psuedo results**



- Currently only have 8TeV exclusion
  - Yellow line : Official simplified models
  - Black/Red (controversial) : Simplified + EWSB
    - Allows us to add Higgsstrahlung

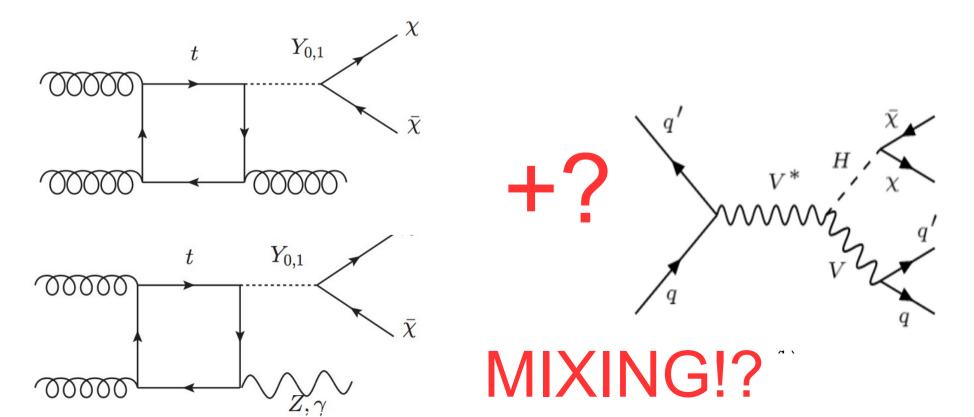
#### **Our Current Scalar & Psuedo results**



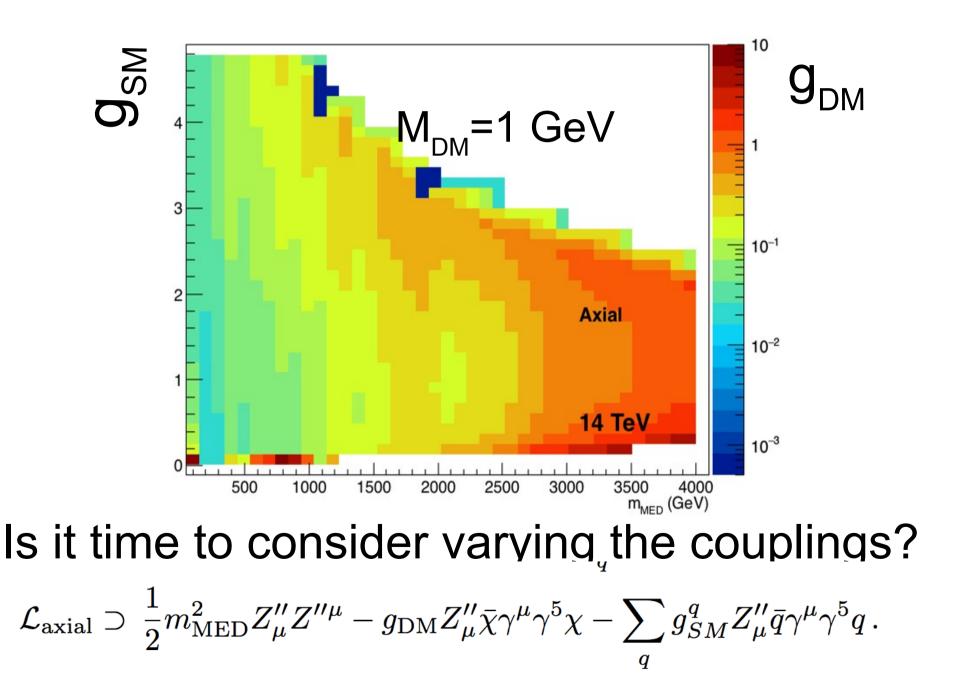
- When the dark matter is not onshell
  - Strong exclusion of pseudoscalar interpreation of LAT
  - Scalar and Direct detection are in close comopetition
    - Expect LHC to pass LUX this summer winter!

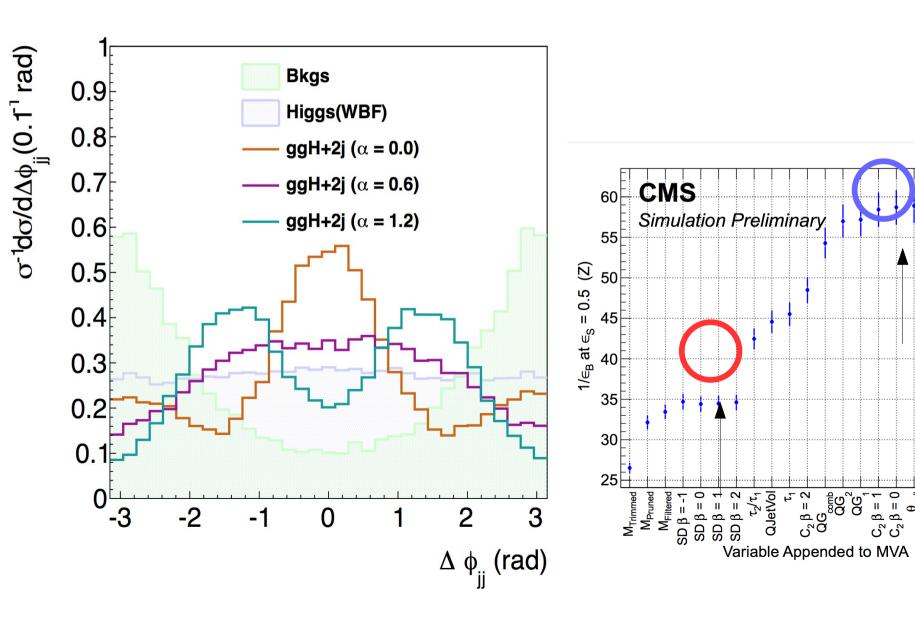
#### To break or not to break?

- EWK symmetry breaking adds lots of mono-V
  - Contribution can be very significant if pseudoscalar
- There are models that do that (e.g. 2HDM...)
  - Need physics at a higher scale (dim-7 operator)



#### Extending Our results





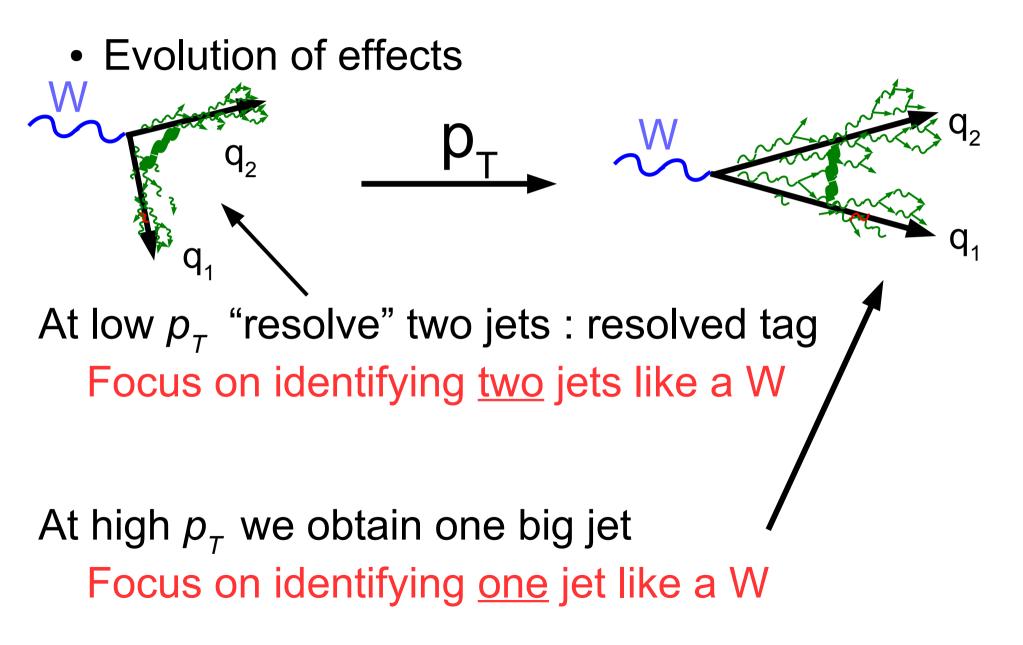
8 TeV

 $C_2 \beta = 0.5$  $C_2 \beta = 0.2$ 

Q/G |Pull

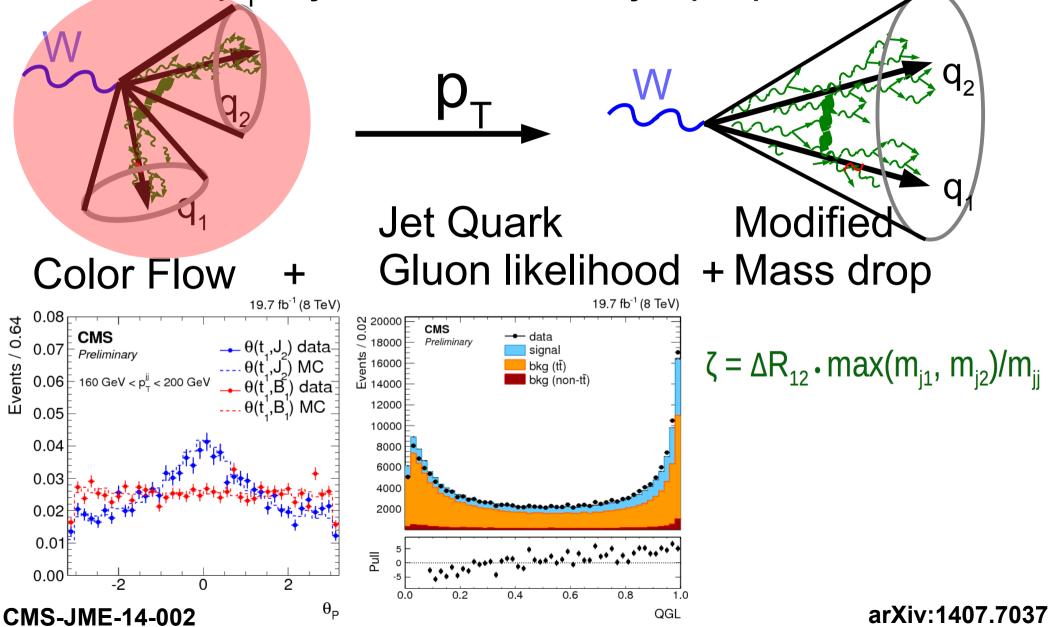
 $\theta_{pull}$ 

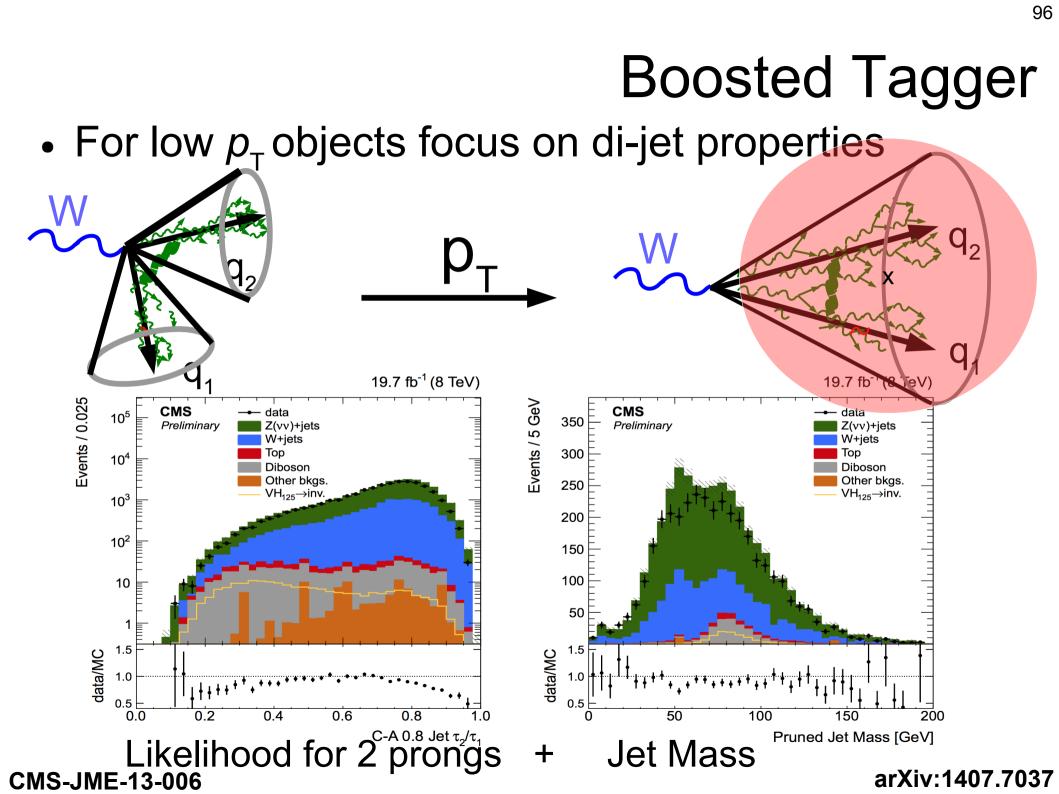
#### **Building a V-tagger**



#### **Resolved Tagger**

• For low  $p_{\tau}$  objects focus on di-jet properties





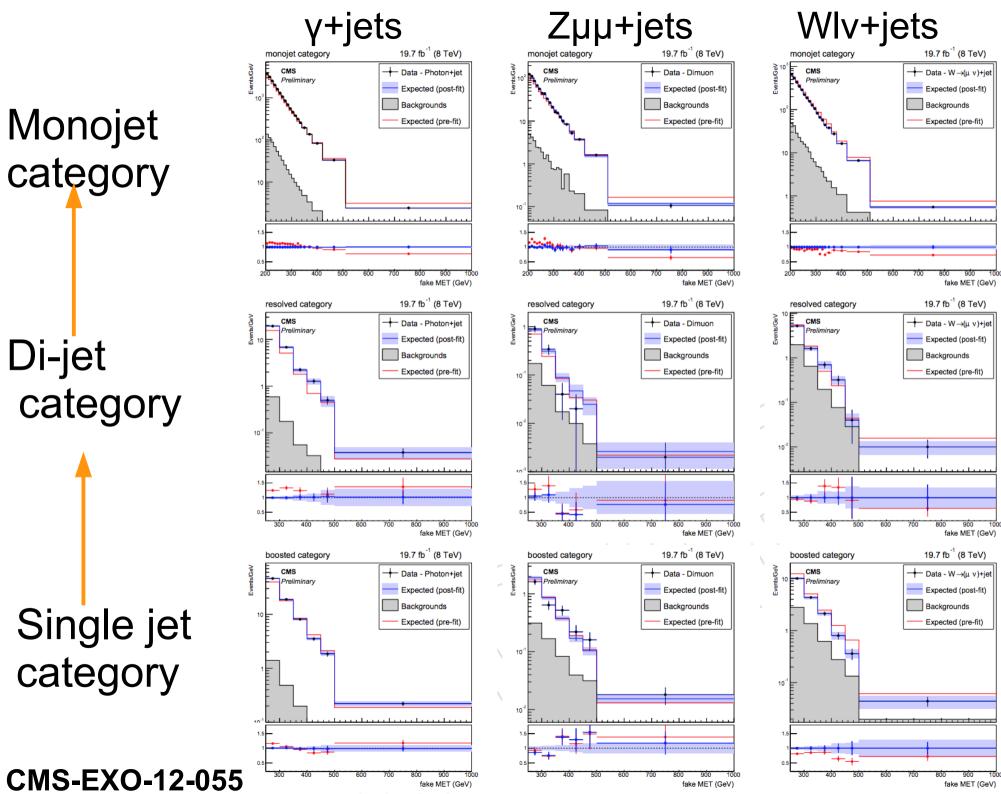
#### **One Big Analysis**

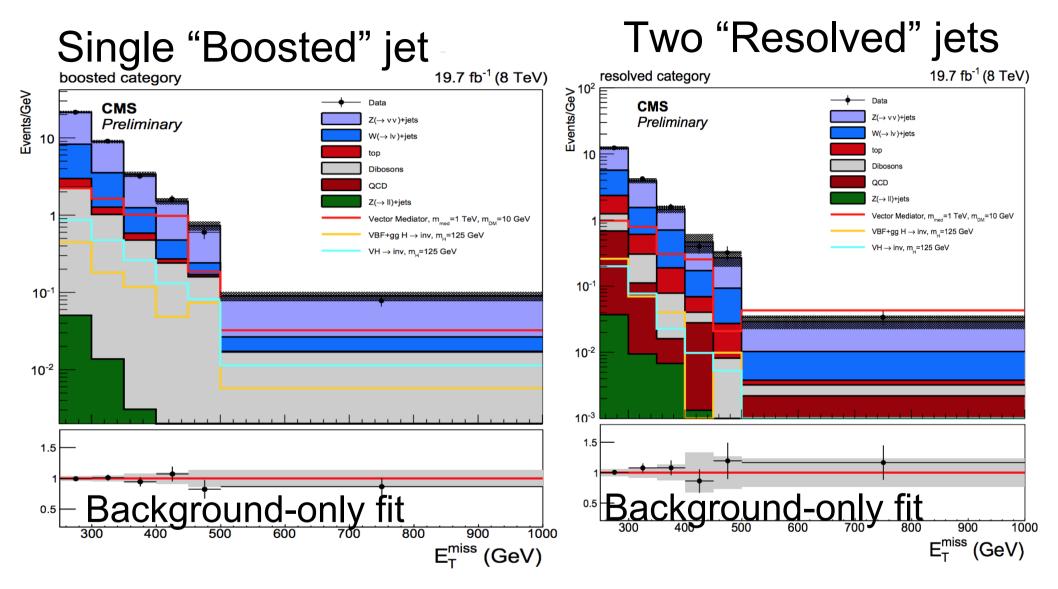
"Its all just jets and MET"

Single Jet Vector boson

**Di-Jet Vector boson** 

One or two jets + *MET* 



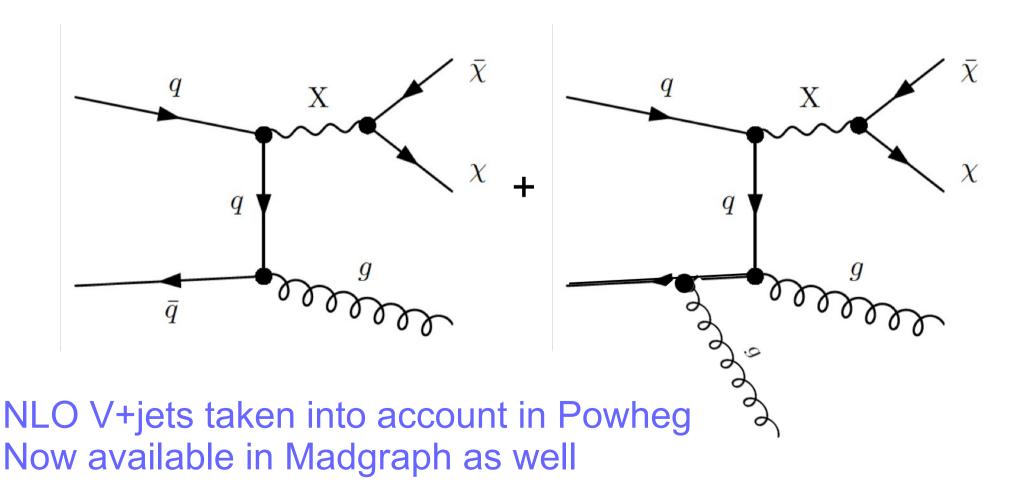


Observe a small excess in resolved *MET* tail (1 $\sigma$ ) Observe a small deficit in the mono-V *MET* tail (1.5 $\sigma$ )

**CMS-EXO-12-055** 

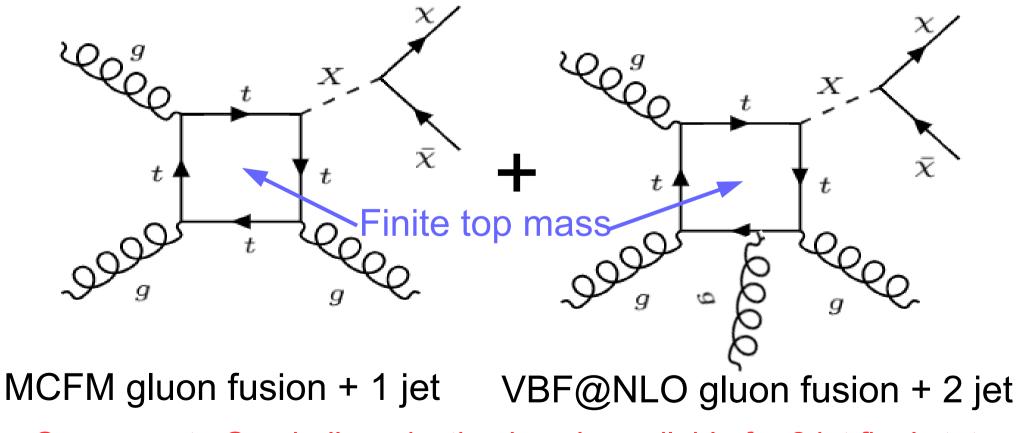
## 2 Jets @ Vector/Axial Simplified Model

- At higher  $\sqrt{s}$  multi-jet final states predominant
- In light of building on new ideas



### (Pseudo)Scalar Simplified Model

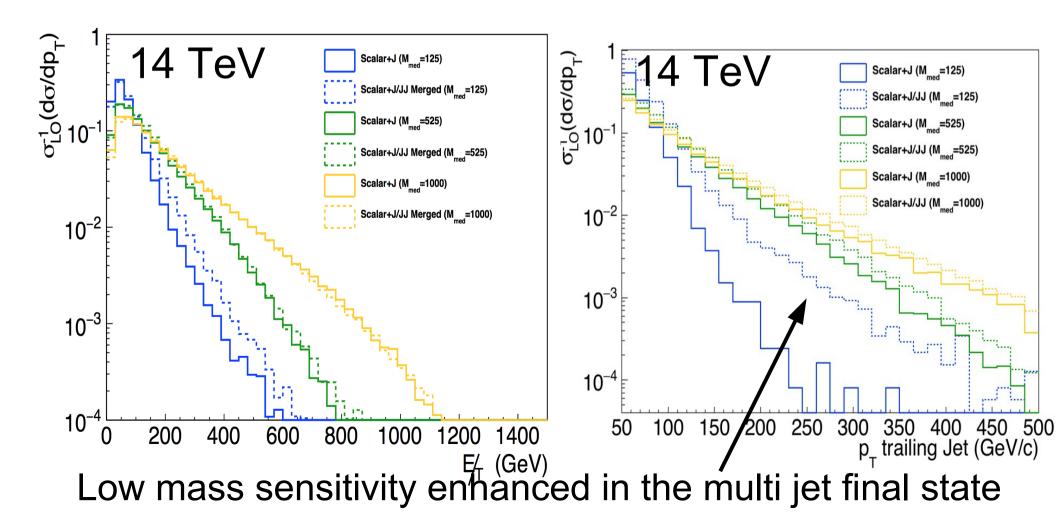
- Requires finite top mass at 2 jets order
  - Available now in now with Madgraph
  - Also can do some hacky procedure



One caveat : On-shell production is only available for 2 jet final state

#### Advantage of merged MC

- Taking advantage of the new technology
  - Can consider exploring new regions of phase space

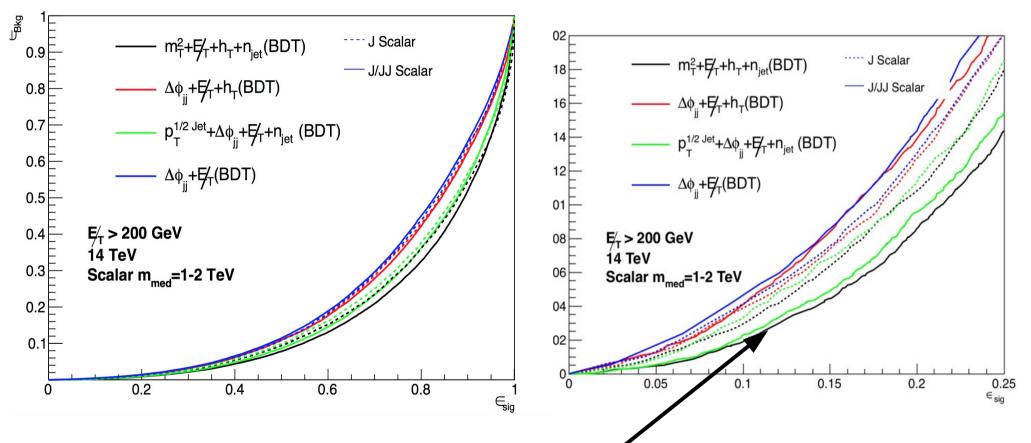


#### Question #2 : Advantage of MC

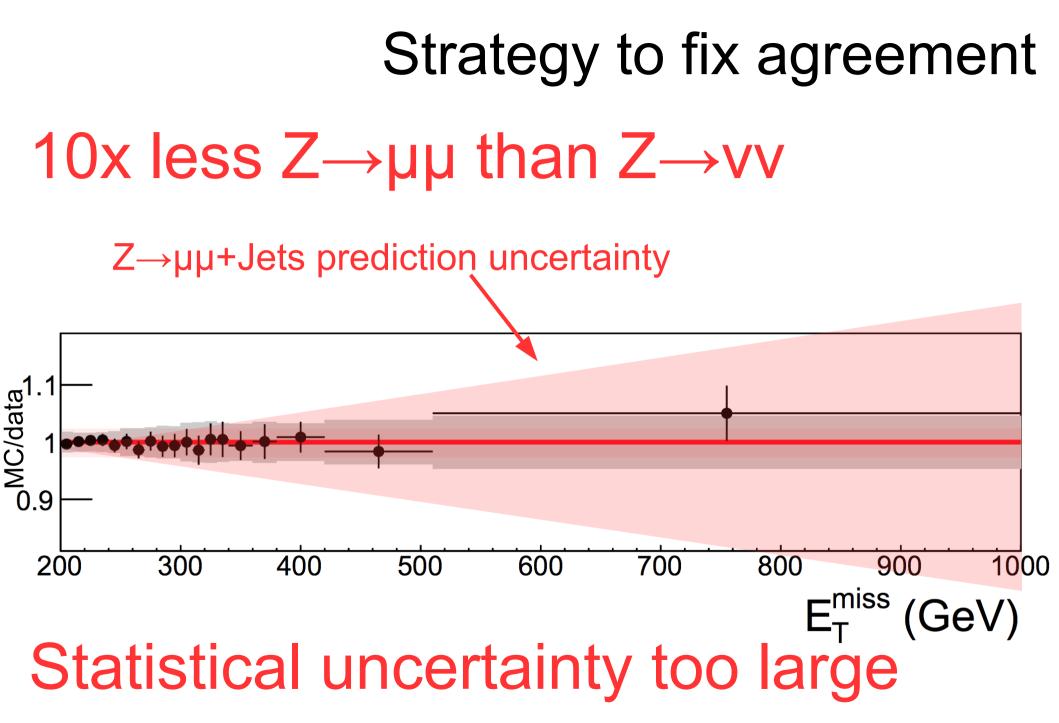
- With 2 jet MC : can now probe multijet final states
- Two questions can be answered :
  - Which variables are most sensitive with 2 jet MC?
  - Which variables are sensitive at 100 TeV?
- Considered a number of multi-jet variables :
  - $M_T^2$ : SUSY like variable obtained for pairwise sparticles
  - Razor variables :  $M_R$  , R : Related SUSY variables
  - MET : standard
  - $\Delta \phi_{jj}$  : angle between the two jets

# Whats the maximum gain?

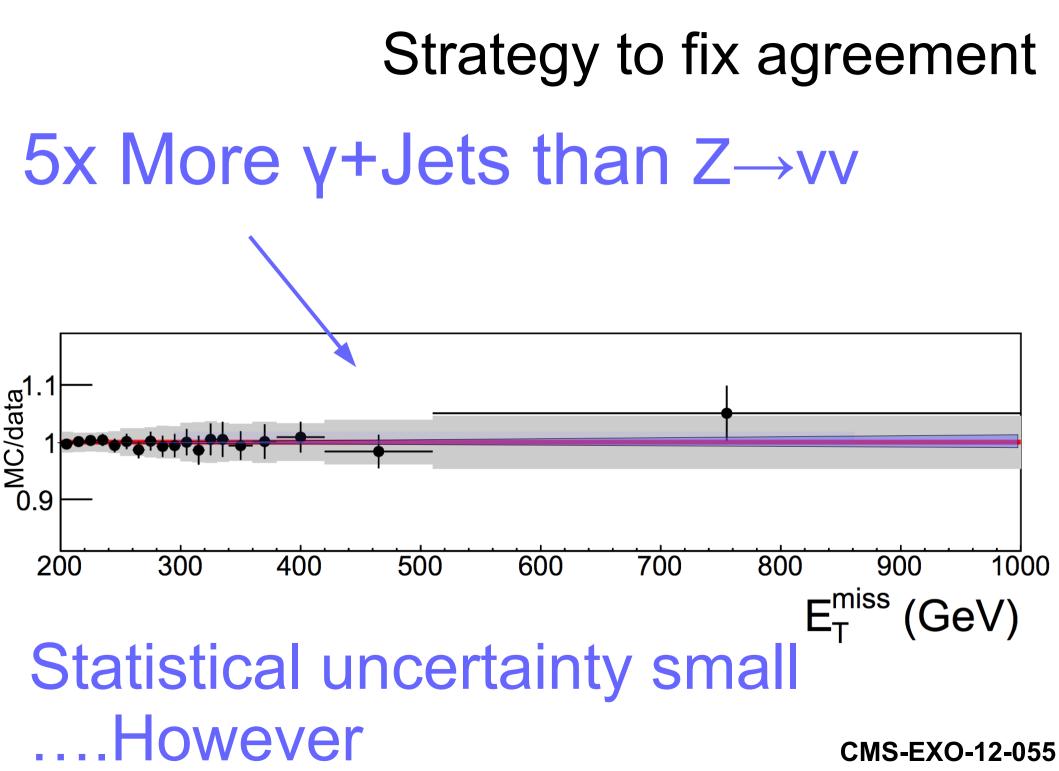
• Making an MVA combining all information



Background drops by a factor of 2 Can maximize sensitivity by an additional sqrt(2) The only other way to gain is to reduce the systematics

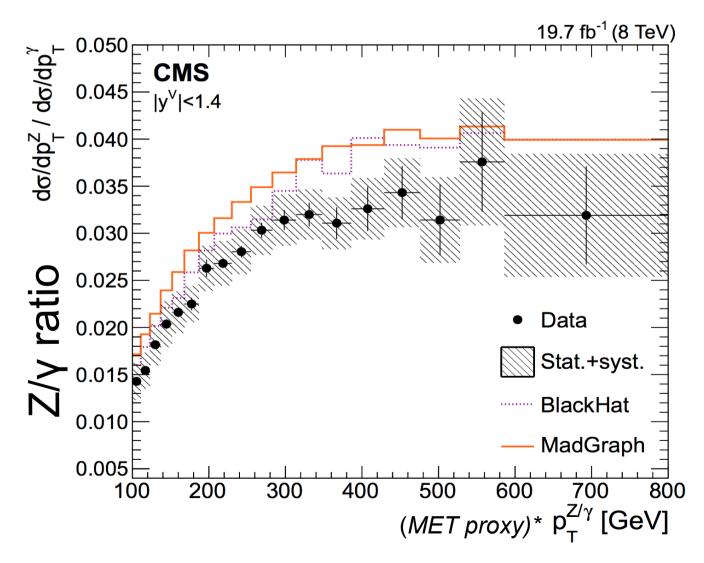


CMS-EXO-12-055



# A mystery? Understanding $Z/\gamma p_{\tau}$

#### Can we really use Photons to model Zs?



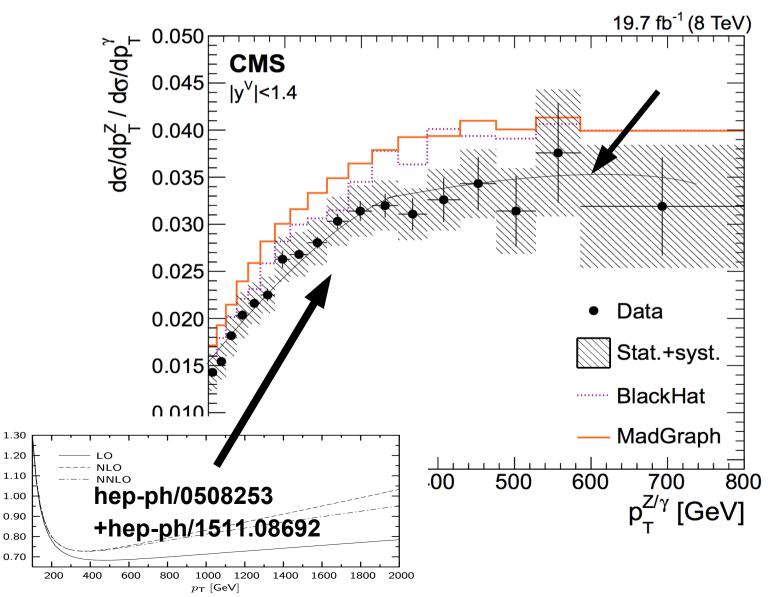
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CMS-SMP-14-005

#### \*See backup

#### A mystery? The Z $p_{\tau}$ spectrum

• These results are missing NLO EWK corrections!



dp1 dp1

빌븉

CMS-SMP-14-055

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#### How do we fix this?

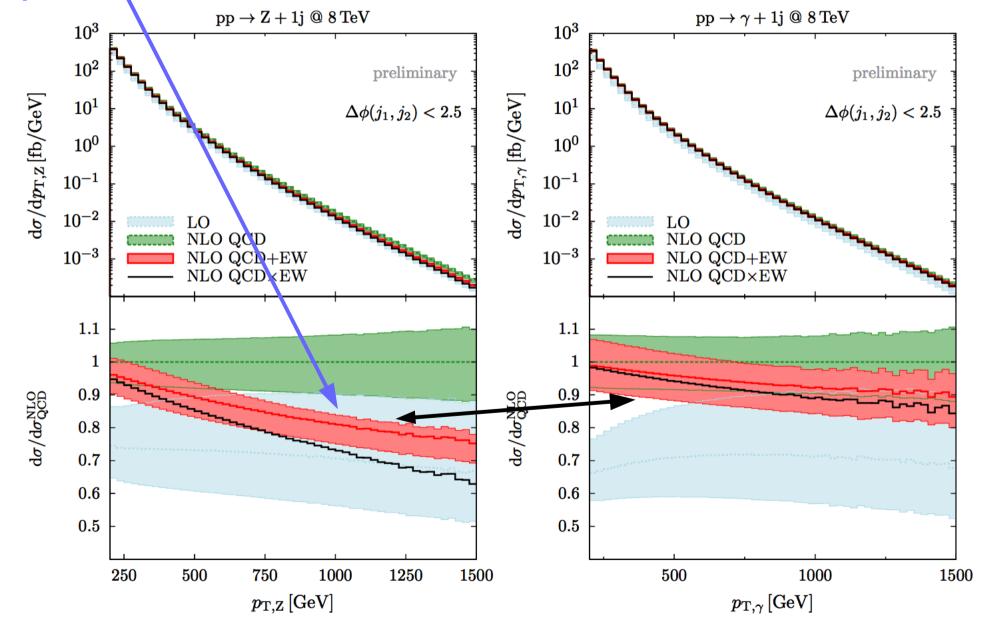
# $\begin{array}{cc} d\sigma^{\gamma} & and & d\sigma^{z} \\ dp_{\tau} & & dp_{\tau} \end{array}$

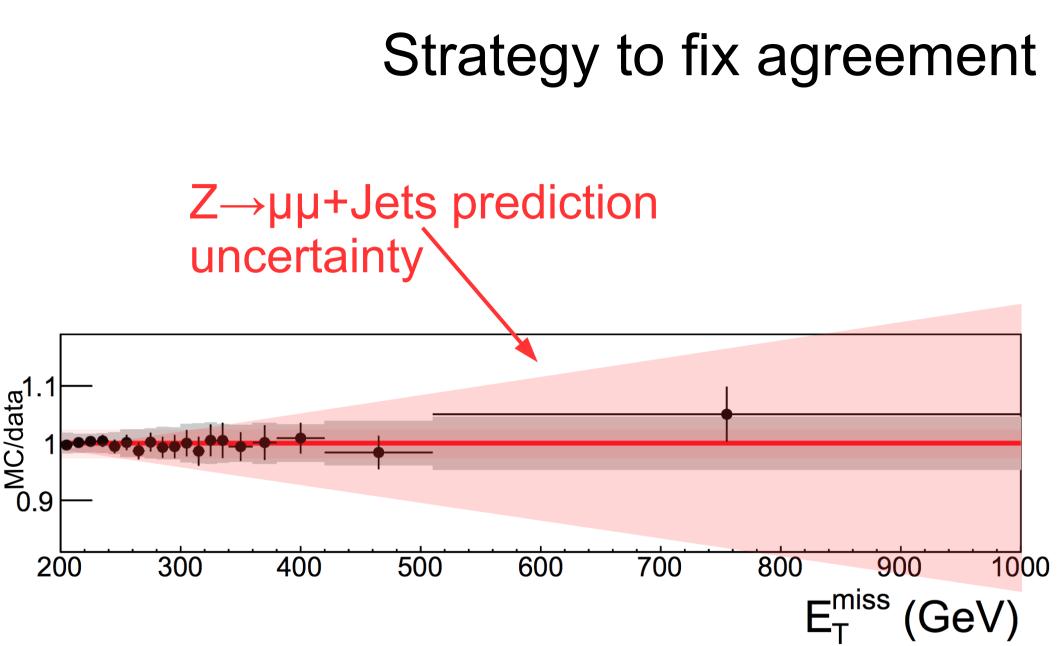
## Before : $\sigma_{tot} = \sigma_{NLO}(0, 1jet)$ After : $\sigma_{tot} = \sigma_{NLO}(0, 1, 2j)(1 + \sigma_{EWK})$ (added)

Energy leakage outside of photon which biases *MET* This was the harder one

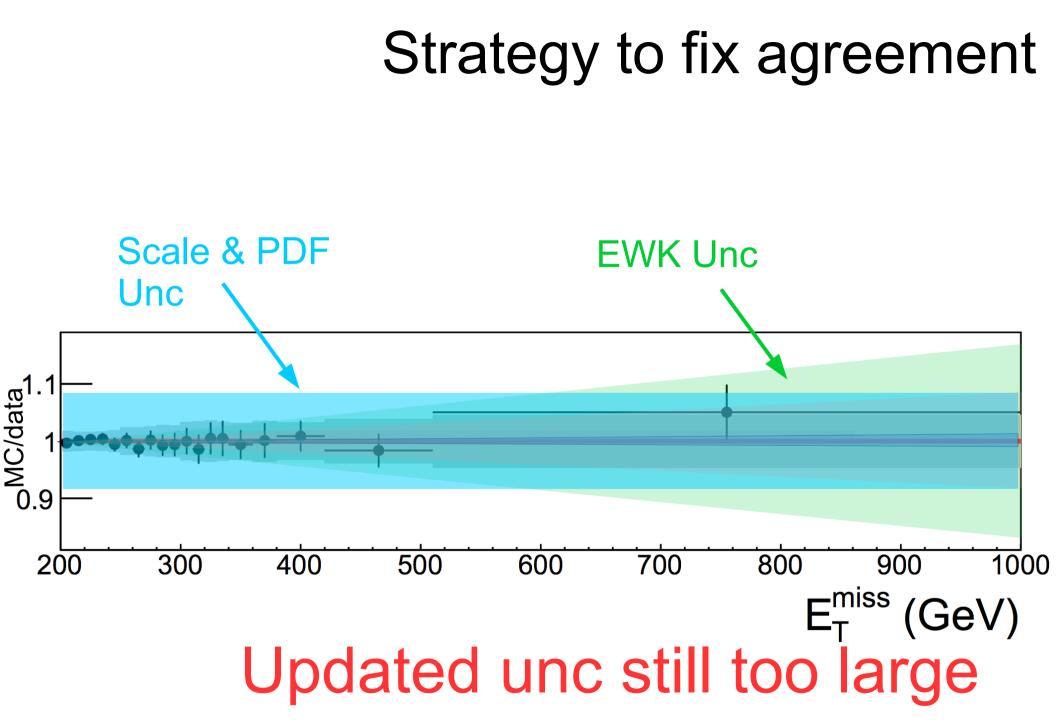
## How do we fix this?

#### Impact of the electroweak corrections





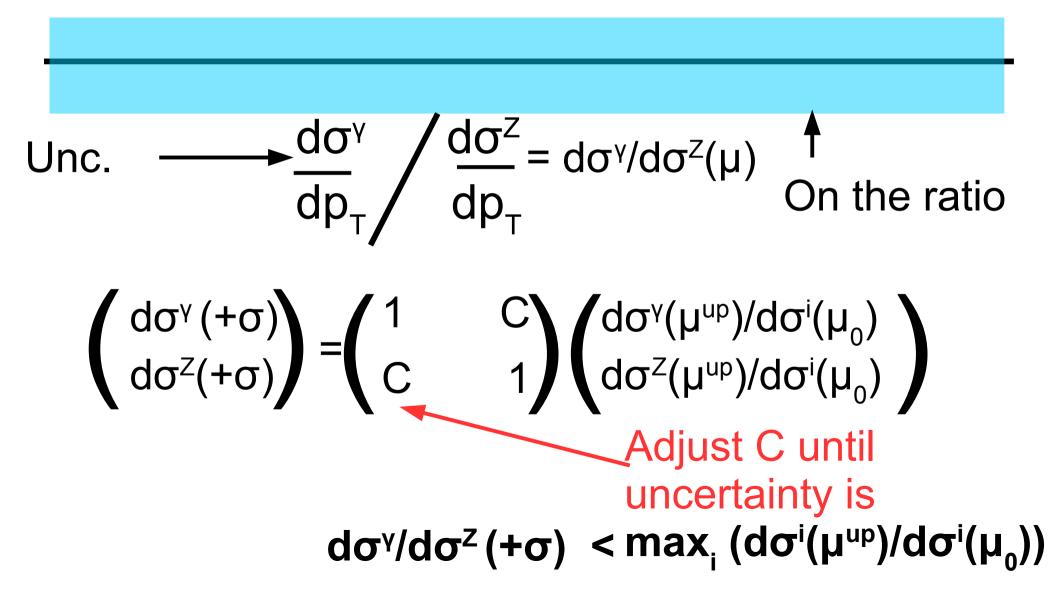
**CMS-EXO-12-055** 



**CMS-EXO-12-055** 

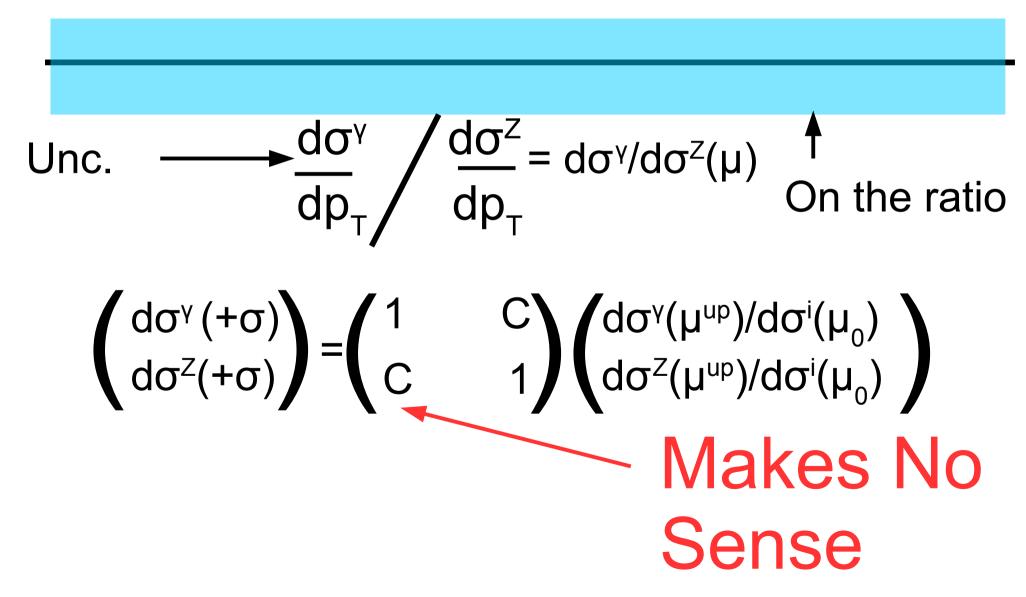
## What is the previous unc?

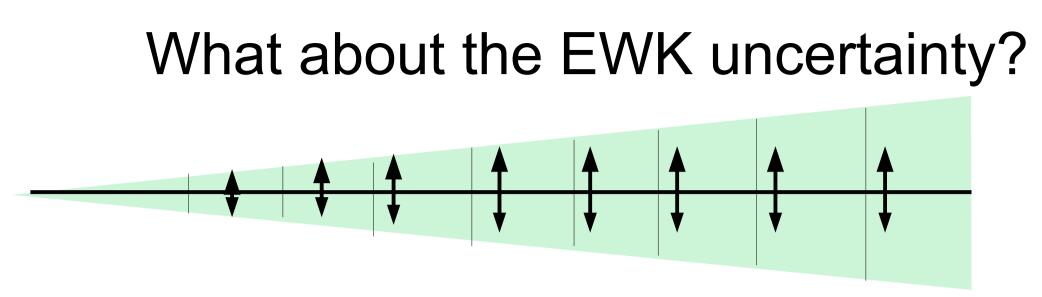
#### Scale & PDF Unc



#### What is the previous unc?

#### Scale & PDF Unc

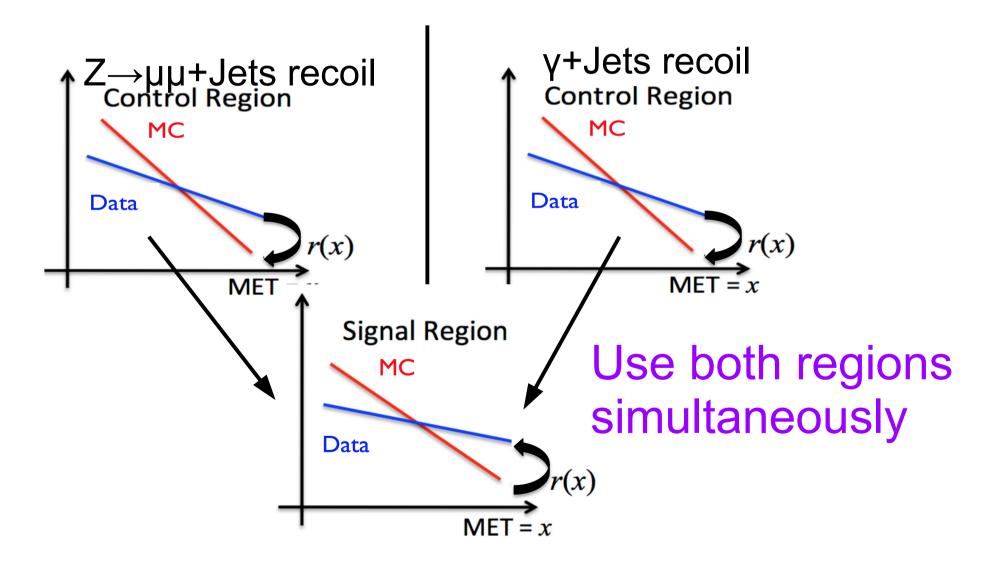




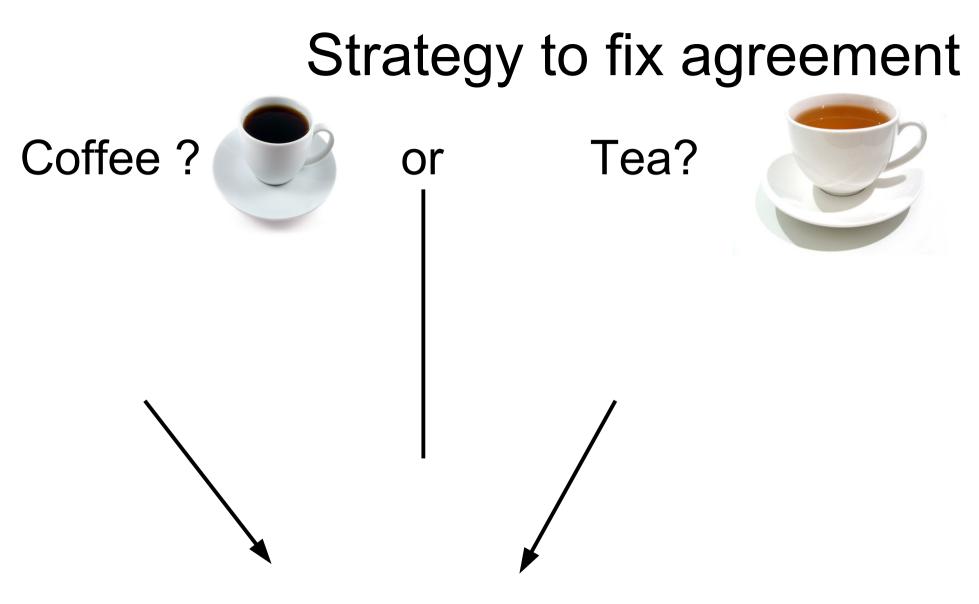
#### In light of being conservative : Treated full correction as an uncertainty

Additionally de-correlated this per bin Avoids low *MET* to high *MET* constraints Not very logical Other (better) schemes exist

## Strategy to fix agreement



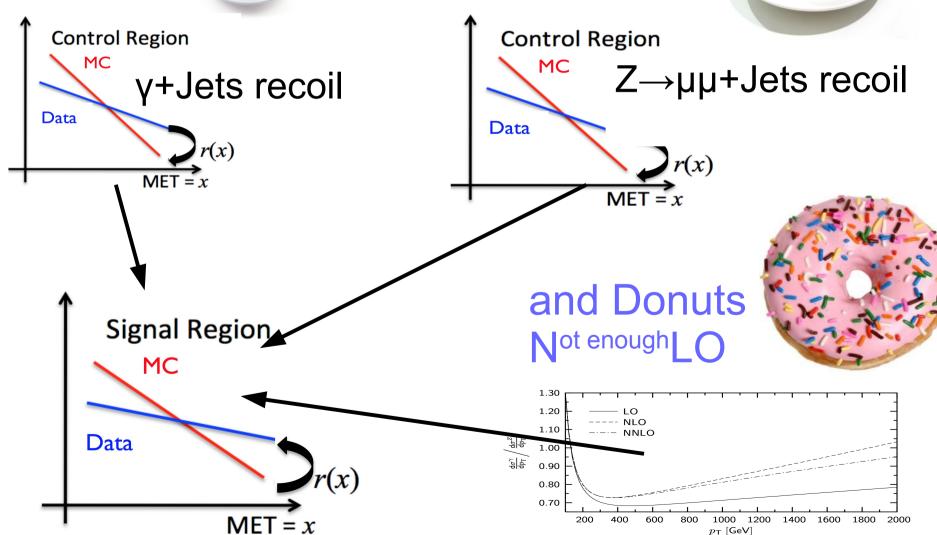
116

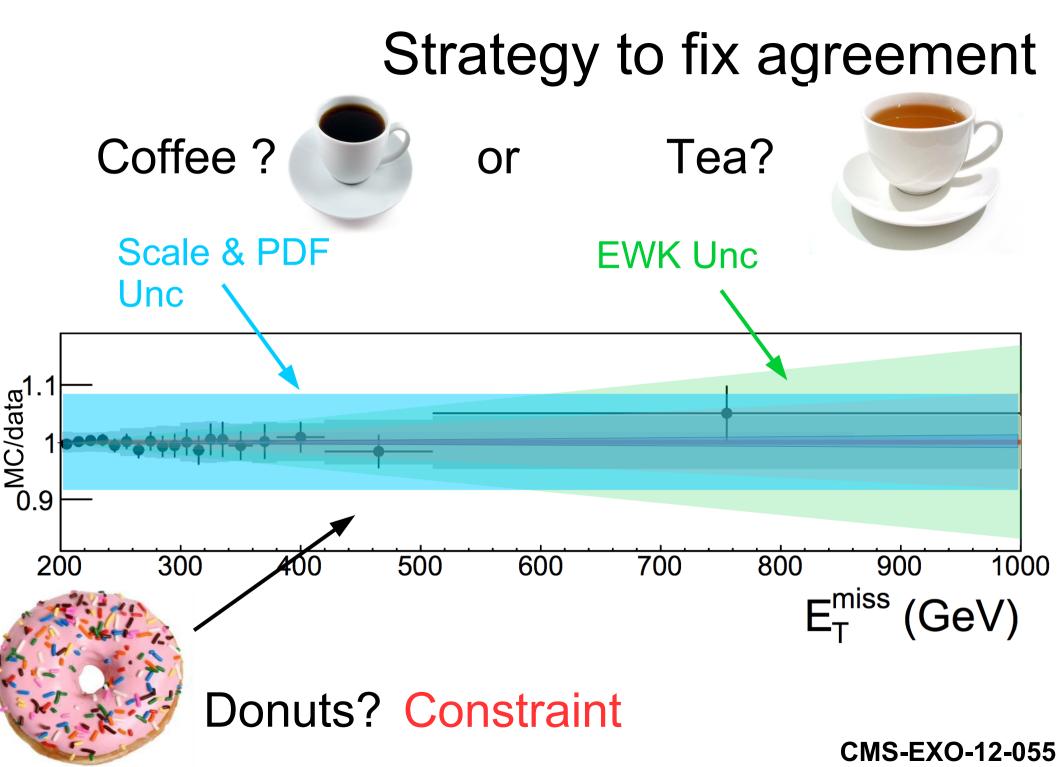


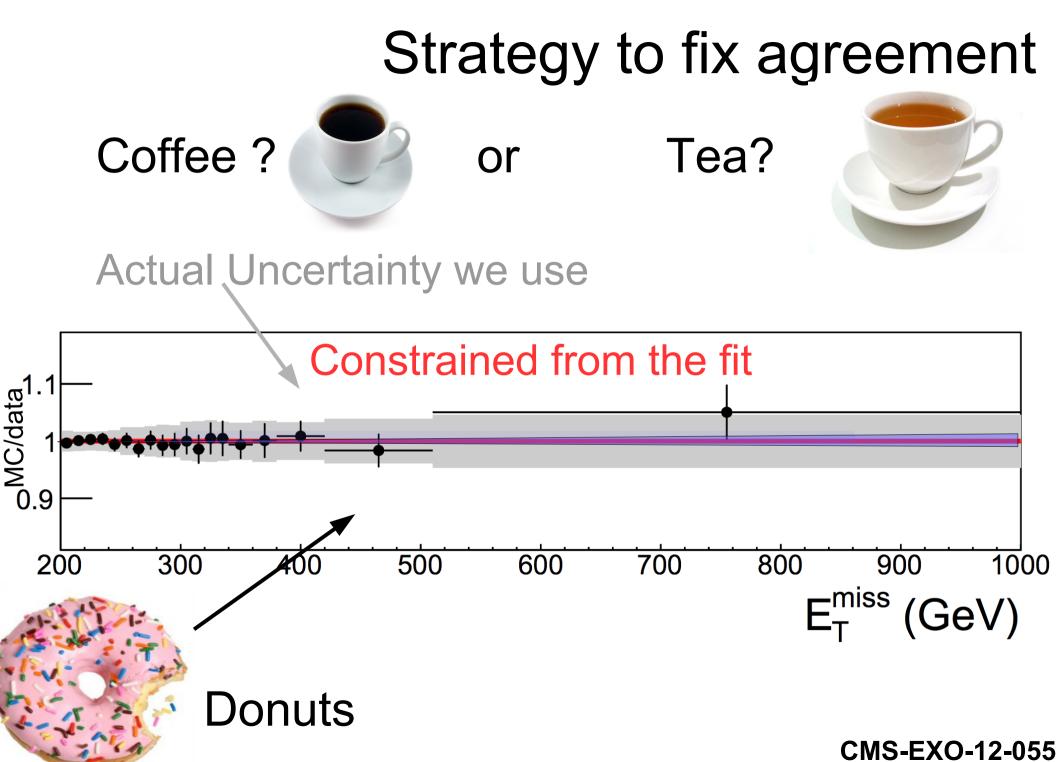
Answer : Yes Please

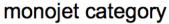
**CMS-EXO-12-055** 

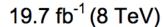
# Coffee ? or Tea?









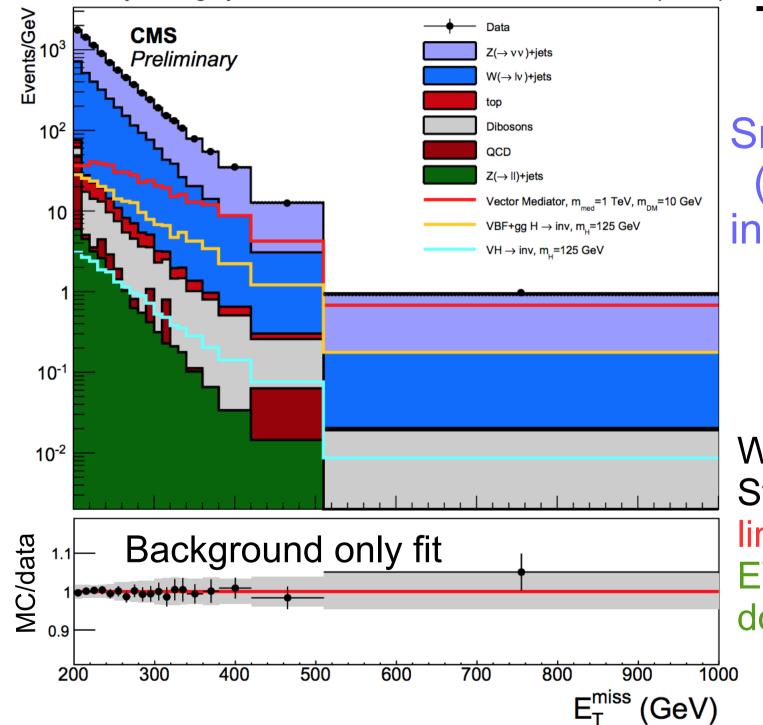


## The Result

Small excess  $(1-2\sigma)$  in *MET* tail

With new method Still systematics limited EWK uncertainty dominates

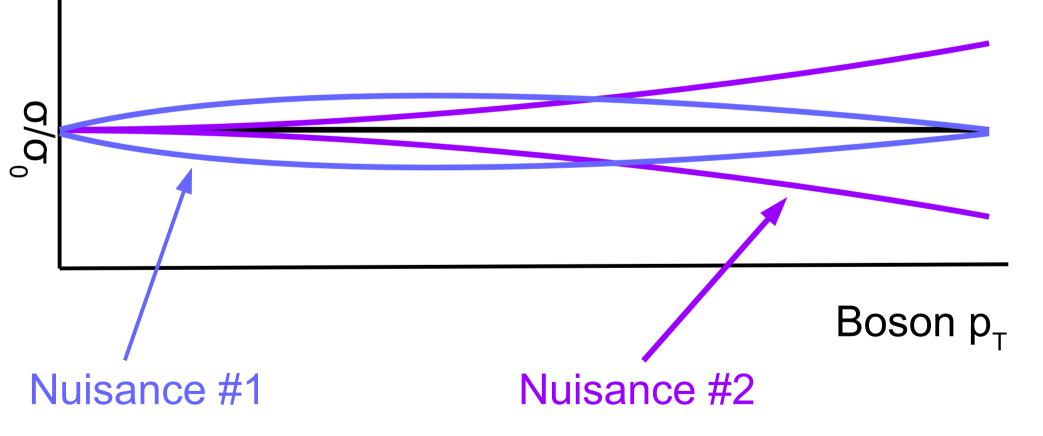
CMS-EXO-12-055



## How do we do the fit?

The updated version of this search fits the  $W,Z,\gamma p_{\tau}$  simultaneously

Simultaneously profile the shapes of the  $p_{\tau}$  spectra



Can we bound our uncertainties into a class of shapes?

## We are Stuck!

We are relying on

 $d\sigma^{\gamma}/d\sigma^{Z}(+\sigma) < \max_{i} (d\sigma^{i}(\mu^{up})/d\sigma^{i}(\mu_{0}))$ 

For the uncertainty on

$$\frac{d\sigma^{\gamma}}{dp_{T}} / \frac{d\sigma^{z}}{dp_{T}}$$

We need a better a approach

Ideally one that we can embed to the likelihood(L)

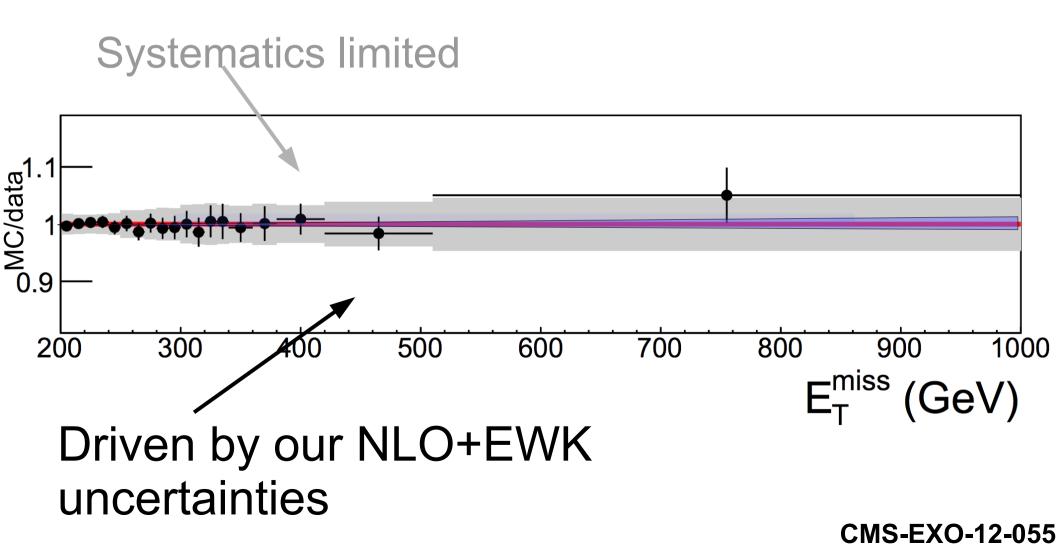
 $Log(L)=Log(L_0)+(d\sigma^{\gamma}/d\sigma^{Z}(\theta)-d\sigma^{\gamma}/d\sigma^{Z}(\mu_0))/\sigma^{2}$ 

#### Profiled nuisance

Improved knowledge of high  $p_{\tau}$  spectrum drives search

#### Can we improve?

Monojet search will not improve quickly in the future



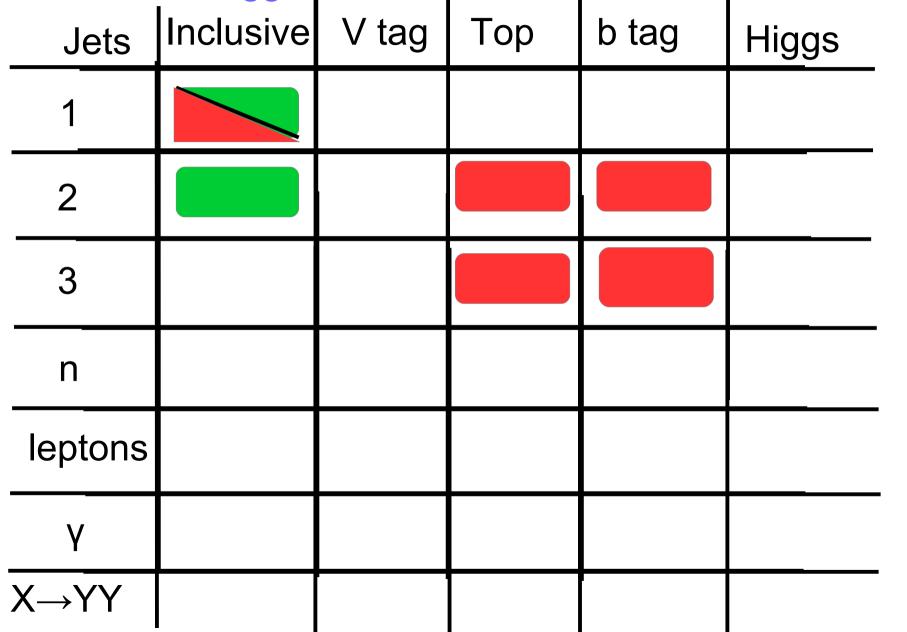
## Towards a complete statement on Dark Matter

#### Analyses presented

• Mono Jet:

Scalar Axial Higgs modified vector modified scalar mixture Inclusive Тор V tag Jets b tag Higgs 2 3 n leptons γ X→YY

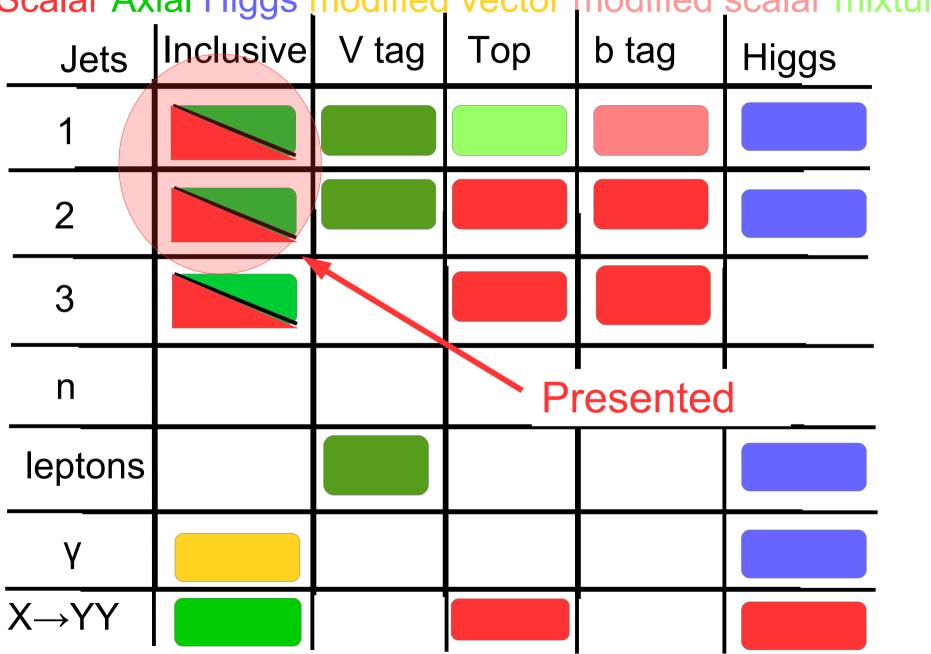
## Mono Everything: at 100 TeV



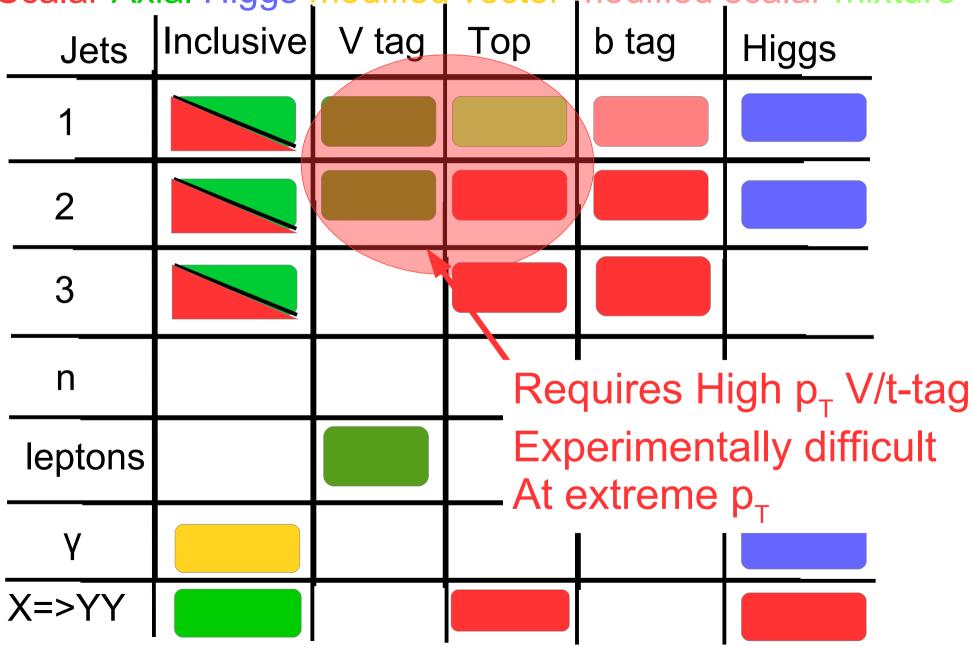
Mono Everything: Extending models to cover modified simplified models



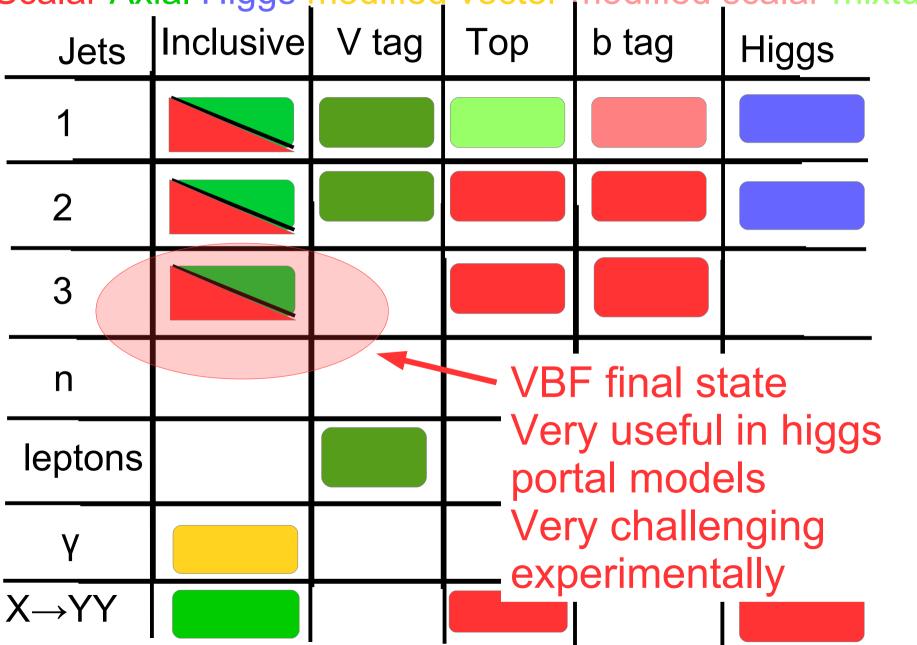
Mono Everything:
 Extending models to cover modified simplified models



Mono Everything: Extending models to cover modified simplified models



Mono Everything:
 Extending models to cover modified simplified models



Mono Everything:
 Extending models to cover modified simplified models

