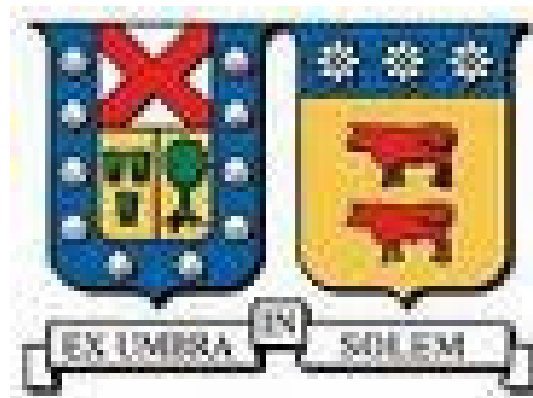


CGC: new b -dependent model in NLO

Eugene Levin

Tel Aviv University/UTFSM



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Carlos Contreras, E.L., Rodrigo Meneses and Irina Potashnikova

Why we are doomed to build a model

- **Short answer:**
- **Because we failed to include in CGC equations the non-perturbative**

large b behavior :

$$A \propto e^{-\mu b}$$

- **Perturbative QCD $\implies A \propto 1/b^4 \implies \sigma \propto s^\lambda \gg \ln^2 s$**

(Kovner & Wiedemann, 2002-2003)

Main building ideas:

1. Non-perturbative large b behavior can be absorbed in $Q_s(Y, b)$

$$Q_s^2 = Q_0^2(Y = Y_0, b) \left(\frac{1}{x}\right)^{\lambda_{\text{pert. QCD}}}$$

with

$$Q_0^2(Y = Y_0, b) = Q_0^2 S(b) \xrightarrow{b \gg 1/\mu} Q_0^2 e^{-\mu b}$$

Theory: $\lambda_{\text{pert. QCD}}$ $e^{-\mu b}$

Fitting parameters: Q_0 and μ

Popular but correct only in semi-classical approach

2. Geometric scaling and matching of two theoretical amplitudes:

- deep inside the saturation domain ($\tau = r^2 Q_S^2(Y, b) \gg 1$)
- in vicinity of the saturation scale ($\tau = r^2 Q_S^2(Y, b) \approx 1$)

(Iancu, Itakura & Munier, 2004)

$$\tau \gg 1 \quad N = 1 - A e^{-\kappa_{\text{pert. QCD}} \ln^2 \tau}$$

(E.L. & Tuchin, 2000)

$$\tau \approx 1 \quad N = N_0 (\tau)^{1 - \gamma_{\text{pert. QCD}}^{cr}}$$

(Iancu, Itakura & McLerran, Mueller & Triantafyllopoulos)

Theory: $\kappa_{\text{pert. QCD}}$ and $\gamma_{\text{pert. QCD}}^{cr}$

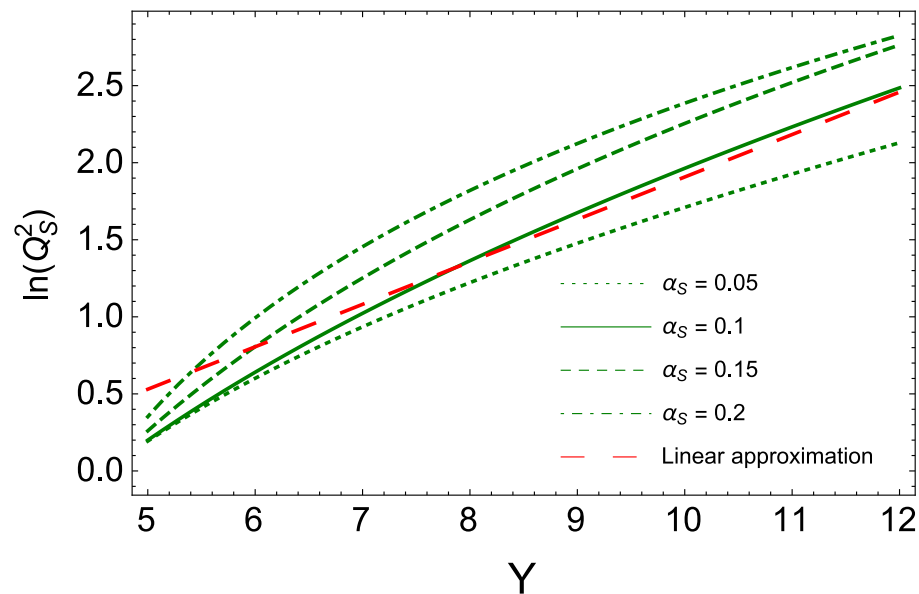
Fitting parameters: A and N_0

3. Saturation momentum $Q_s(Y, b)$:

- $$\ln \left(Q_s^2(Y, b) / Q_s^2(Y = Y_0, b) \right) =$$

$$d_1^{\text{pert.QCD}}(\bar{\alpha}_S) (Y - Y_0) - d_2^{\text{pert.QCD}}(\bar{\alpha}_S) \ln(Y/Y_0)$$

$$+ d_3^{\text{pert.QCD}}(\bar{\alpha}_S) \left(\frac{1}{\sqrt{Y_0}} - \frac{1}{\sqrt{Y}} \right)$$



Linear approximation \longrightarrow $0.7 d_1^{\text{pert.QCD}}(\bar{\alpha}_S) (Y - Y_0)$ at $\bar{\alpha}_S = 0.1$.

4. Phenomenological input: impact parameter dependence

- Theoretical restrictions:

- At large b $A \propto e^{-\mu b}$ (Froissart ,1961);

- At large Q_T (conjugated variable to b) $A \propto 1/Q_T^4$
(Brodsky & Lepage, ,1979);

- ♠ Our choice:

- $$A(Q_T) = \frac{1}{(1 + Q_T^2/m^2)^2} \implies A(b) = mb K_1(mb)$$

- $$S(b) = \left(mb K_1(mb) \right)^{1/(1-\gamma^{cr})}$$

Theoretical results for $\tau \gg 1$

Non-linear equation: (Balisky & Chirilli, 2008)

$$\begin{aligned}
 \frac{\partial S_{12}}{\partial Y} &= \frac{\bar{\alpha}_S}{2\pi} \int d^2 x_3 \frac{x_{12}^2}{x_{13}^2 x_{23}^2} \left\{ 1 + \bar{\alpha}_S b \left(\ln x_{12}^2 \mu^2 - \frac{x_{13}^2 - x_{23}^2}{x_{12}^2} \ln \frac{x_{13}^2}{x_{23}^2} \right) \right. \\
 &+ \bar{\alpha}_S \left(\frac{67}{36} - \frac{\pi^2}{12} - \frac{5}{18} \frac{N_f}{N_c} - \frac{1}{2} \ln \frac{x_{13}^2}{x_{12}^2} \ln \frac{x_{23}^2}{x_{12}^2} \right) \left. \right\} (S_{13} S_{32} - S_{12}) \\
 &+ \frac{\bar{\alpha}_S^2}{8\pi^2} \int \frac{d^2 x_3 d^2 x_4}{x_{34}^4} \left\{ -2 + \frac{x_{13}^2 x_{24}^2 + x_{14}^2 x_{23}^2 - 4x_{12}^2 x_{34}^2}{x_{13}^2 x_{24}^2 - x_{14}^2 x_{23}^2} \right. \\
 &\left. \ln \frac{x_{13}^2 x_{24}^2}{x_{14}^2 x_{23}^2} + \frac{x_{12}^2 x_{34}^2}{x_{13}^2 x_{24}^2} \left(1 + \frac{x_{12}^2 x_{34}^2}{x_{13}^2 x_{24}^2 - x_{14}^2 x_{23}^2} \right) \ln \frac{x_{13}^2 x_{24}^2}{x_{14}^2 x_{23}^2} \right\} \\
 &\times (S_{13} S_{34} S_{42} - S_{13} S_{32})
 \end{aligned}$$

For $S_{ij} \ll 1$ the equation degenerates to (E.L. & Tuchin , 2000)

- $$\frac{\partial S_{12}}{\partial Y} = -\frac{\bar{\alpha}_S}{2\pi} \int d^2 x_3 \frac{x_{12}^2}{x_{13}^2 x_{23}^2} \left\{ \dots \right\} S_{12}$$

- $$\frac{d \ln S_{12}}{dY} = -\bar{\alpha}_S \left[1 + \bar{\alpha}_S b \ln(\mu^2 x_{12}^2) + \bar{\alpha}_S \left(\frac{67}{36} - \frac{\pi^2}{12} - 5 \frac{N_f}{N_c} \right) \right] \ln(Q_s^2 x_{12}^2)$$

$$+ \frac{\bar{\alpha}_S^2 b}{2} \ln^2 Q_s^2 x_{12}^2 + \frac{\bar{\alpha}_S \zeta(3)}{16}$$

$$\mu = Q_s = -\bar{\alpha}_S(Q_s) \left[1 + \frac{3}{2} \bar{\alpha}_S(Q_s) b \ln(Q_s^2 x_{12}^2) \right.$$

$$\left. + \bar{\alpha}_S(Q_s) \left(\frac{67}{36} - \frac{\pi^2}{12} - 5 \frac{N_f}{N_c} \right) \right] \ln(Q_s^2 x_{12}^2) + \frac{\bar{\alpha}_S(Q_s) \zeta(3)}{16}$$

- $$\ln S_{12} = -\mathcal{Z} =$$

$$-\frac{1}{2 \varrho} \left(\left[z + \bar{\alpha}_S(Q_s) b z^2 + \bar{\alpha}_S(Q_s) z \left(\frac{67}{36} - \frac{\pi^2}{12} - 5 \frac{N_f}{N_c} \right) \right] z + \frac{\zeta(3)}{8} z \right)$$

$$z = \bar{\alpha}_S(Q_s) \varrho (Y - Y_0) + \ln(Q_0^2 (Y = Y_0) x_{12}^2)$$

- **Results:**
- Geometric scaling behavior with $\bar{\alpha}_S(Q_s)$;
- Particular form $\tau = x_{12}^2 Q_s^2$ dependence for $\tau \gg 1$;

$$N(z) = 1 - A \exp(-\mathcal{Z})$$

(Contreras, E.L. & Meneses, 2014)

- $N^{z \gg 1}(z) = 1 - 2Ae^{-\mathcal{Z}} - A^2 \frac{1}{\mathcal{Z}} e^{-2\mathcal{Z}} + \mathcal{O}(e^{-3\mathcal{Z}})$
- $\mathcal{Z} = \mathcal{Z} \left(z \rightarrow z - \frac{1}{2} A \sqrt{\varrho \pi / 2 - 2\psi(1)} \right)$

Theoretical results for $\tau \approx 1$ in NLO

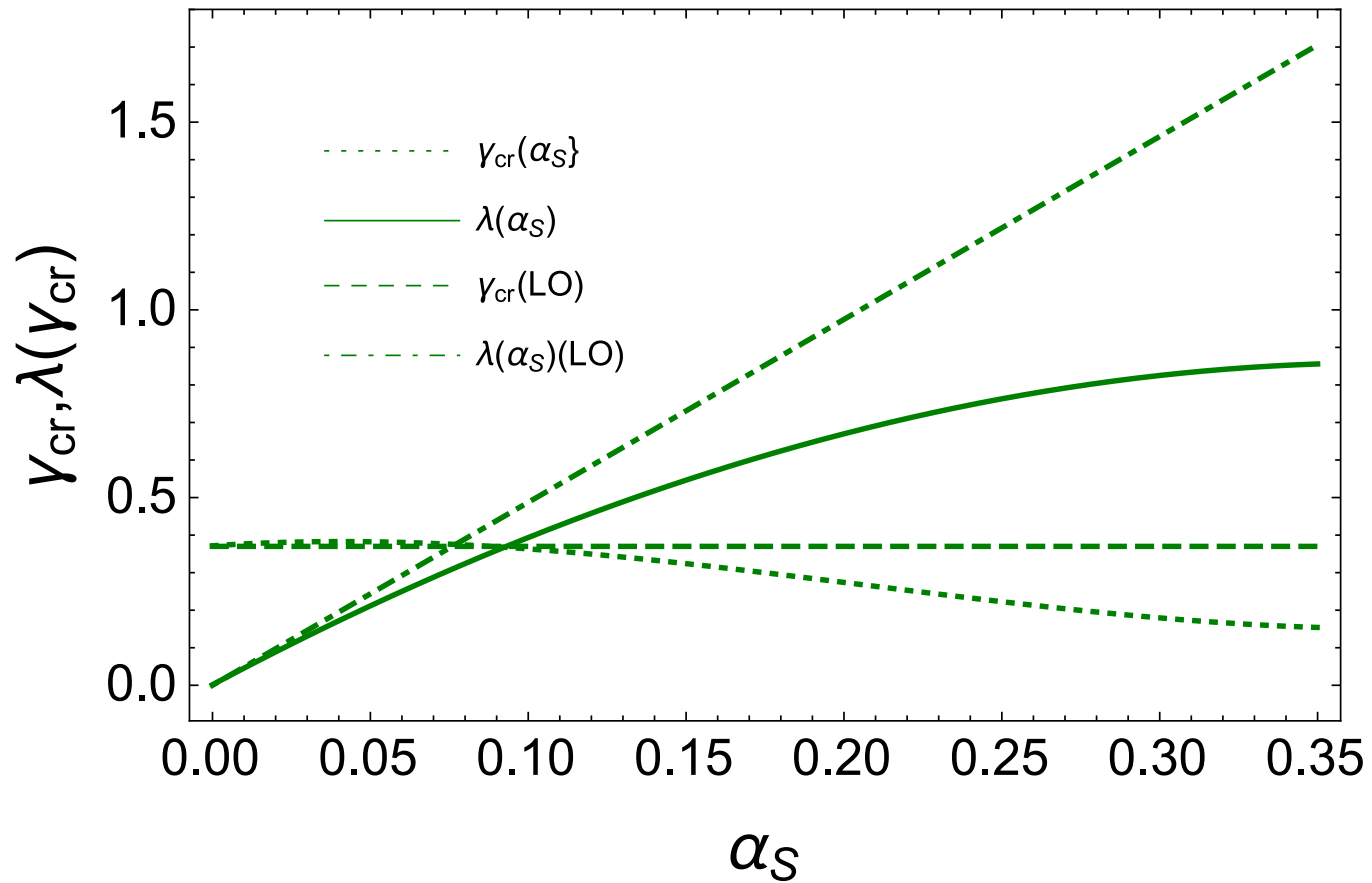
We follow the well known procedure:

- (Lipatov & Fadin, 1998) $\omega = \bar{\alpha}_S (\chi_{LO}(\gamma) + \bar{\alpha}_S \chi_{NLO}(\gamma))$

$\xrightarrow{\text{Salam, 1998}}$ $\omega = \bar{\alpha}_S \left(\chi_0(\omega, \gamma) + \omega \frac{\chi_1(\omega, \gamma)}{\chi_0(\omega, \gamma)} \right)$

- $\frac{\omega^{NLO}(\gamma_{cr})}{1 - \gamma_{cr}} = \left| \frac{d\omega^{NLO}(\gamma_{cr})}{d\gamma_{cr}} \right|$

(Triantafyllopoulos, 2003; Khoze, Martin, Ryskin & Stirling, 2004)

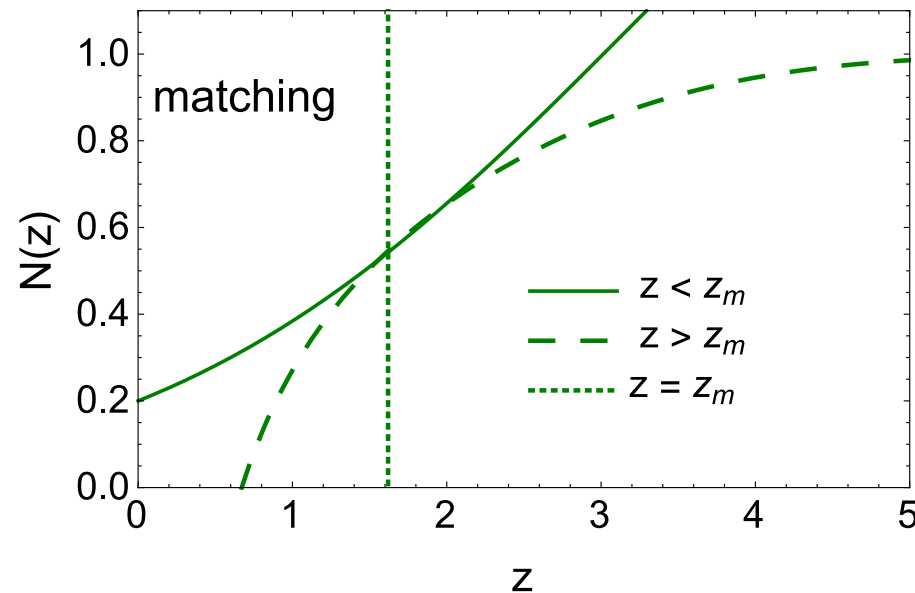


● $\ln \left(Q_s^2(Y) / Q_s^2(Y_0) \right) = \lambda(\gamma_{cr})(Y - Y_0)$

Matching

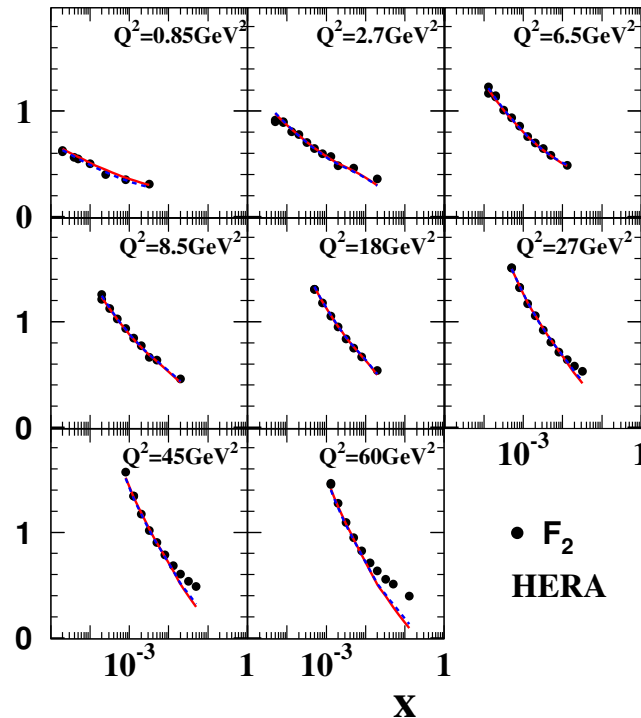
- $N^{0 < z \ll 1} (z = z_m) = N^{z \gg 1} (z = z_m) ;$

- $$\frac{dN^{0 < z \ll 1} (z = z_m)}{dz_m} = \frac{dN^{z \gg 1} (z = z_m)}{dz_m} ;$$



Descriptions of HERA data

$\bar{\alpha}_S$	N_0	Y_0	m (GeV)	Q_0^2 (GeV ²)	m_u (MeV)	m_d (MeV)	m_s (MeV)	m_c (GeV)
0.133	0.1075	3.77	0.83	3.0	2.3	4.8	95	1.4



$$\chi^2/d.o.f. = 183/153 = 1.2$$

Difficulties of the model

- $\bar{\alpha}_S(Q_0) = 0.133$ is too small for $\Lambda_{\text{QCD}} \approx 200 \text{ MeV}$;
- $Q_s^2(Y = Y_0) = 3 \text{ GeV}^2$ is too large ;
- $Q_s^2(Y)$ tends to be extremely large at small x ;

