Haloscope Searches for DM Axions at the Center for Axion and Precision Physics Research



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Institute for Basic Science, South Korea

ICNFP2016, Kolymbari, Crete

## STRONG CP PROBLEM AND AXIONS

#### THE STRONG CP PROBLEM

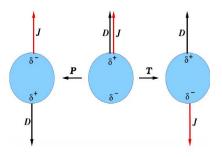
• The QCD lagrangian contains an "absent" term:

$$\mathcal{L}_{QCD} = -\frac{1}{4} G^{a}_{\mu\nu} G^{a\mu\nu} + \sum_{j=1}^{n} \left[ \overline{q_{j}} \gamma^{\mu} i D_{\mu} q_{j} - (m_{j} q^{\dagger}_{Lj} q_{Rj} + \text{H.c.}) \right] + \frac{\theta g^{2}}{32\pi^{2}} G^{a}_{\mu\nu} \tilde{G}^{a\mu\nu}.$$

- Dependence on the theta-term must exist for correct QCD predictions.
- However, it results in CP- and P-violation, unless the *quarks mass matrix* is finely tuned!

At the same time,

- The electric dipole moment of hadrons will imply CP- and P-violation in QCD.
- However, neutron EDM <  $3.0 \times 10^{-26} \text{ e-cm}$  and proton EDM <  $7.9 \times 10^{-25} \text{ e-cm}$ .



- These experimental limits mean that the "theta phase" in the lagrangian is ~10 orders lower than expected.
- (More accurately,  $<10^{-9}$  while expecting  $\sim$ 1. A fine-tuning non-anthropic issue.)

Why is the CP symmetry conserved by QCD? "Strong CP Problem"

#### **ENTER THE AXION**

## Why is the CP symmetry conserved by QCD? "Strong CP Problem"

## Peccei & Quinn ('77):

- > The Strong CP Problem conceals a new symmetry.
- $\rightarrow$  The global " $U_{PQ}(1)$ " quasisymmetry.
- > The potential of theta is modified.
- > Theta is a dynamical variable, with perturbative effects "pulling" it towards zero.
- Axion: the quantum of oscillation of the QCD  $\theta$  parameter. (Weinberg, Wilczek '78)



Searches at CAPP: Two sides of the same coin

The axion was invented to solve the "Strong CP Problem":
Why is the neutron EDM ~10 orders of magnitude smaller than its expected value?

# Microwave cavities DIRECT AXION DETECTION

Today's talk



# Accelerator experiments HADRONIC ELECTRIC DIPOLE



#### **AXION "EVOLUTION"**

Initially the vev was thought to be of the order of EW scale.

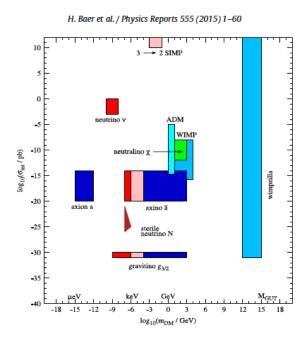
Axion mass -> ~100 keV.

Then a success of the model occured: It was heavily constrained by indirect lab searches and cosmology ^^

Axion mass -> less than 10<sup>-3</sup> eV.

Today, more elaborate models give naturally rise to light axions.

Initially dubbed "invisible axions" for their impossible detection ... until 1983 (see sl.9).

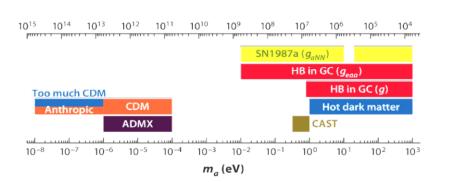


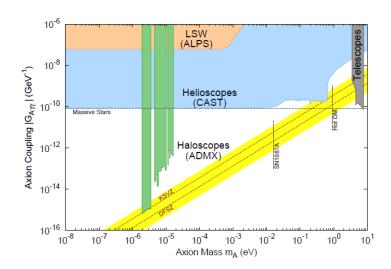
Main invisible axion models:

KSVZ (Kim-Shifman-Vainshtein-Zakharov)

DFSZ (Dine-Fischler-Srednicki-Zhitnitskii)

#### **COSMOLOGICAL AXIONS**



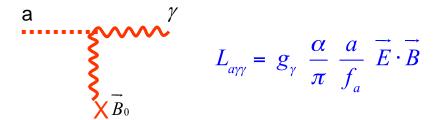


- Good dark matter candidate if mass  $m_a$  is between 1 and 300 $\mu$ eV.
- Kinetic energy ~10<sup>-6</sup> m<sub>a</sub> (cold dark matter).
- Coherent field.
- Cosmological and astrophysical constraints. E.g.:
  - Stellar evolution: Axions emitted in stellar processes would dissipate more energy and stellar evolution would be accelerated.
  - SN1987a: If the collapsing core is also cooled by axion emission, the duration of the neutrino signal will be shorter.

# DIRECT AXION DETECTION / HALOSCOPES

## **Detection scheme by P. Sikivie (PRL 51:1415 1983)**

Axions will convert to photons in a strong magnetic field. ("reverse Primakoff effect")



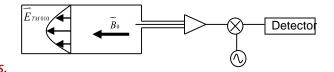
 Process proceeding through loop; both regular SM and new U(1) couplings of axion to fermions.

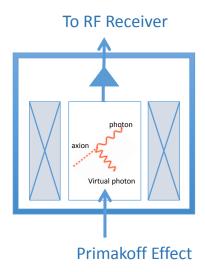
#### **DIRECT AXION DETECTION**

- Powerful magnets.
- Very weak signal (~10-21 W)
  - $\Rightarrow$  Cryogenic temperatures.
  - ⇒ Resonant radiofrequency (RF) cavities for enhancement.
  - $\Rightarrow$  Tunable over a good frequency range.



At the cavity's resonant frequencies the microwaves reinforce to form standing waves in the cavity. Therefore, the cavity functions similarly to an organ pipe or sound box in a musical instrument, oscillating preferentially at a series of frequencies, its resonant frequencies.



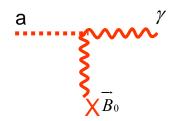


#### **HALOSCOPES AND HELIOSCOPES**

Axions will convert to photons in a strong magnetic field. ("reverse Primakoff effect")

## **Helioscopes**

Pointing to the Sun for axions produced by photons.



## <u>Haloscopes</u>

Looking for ambient axions from the Milky Way's DM halo (~10<sup>14</sup> cm<sup>-3</sup>).







# HALOSCOPE PROGRAMME IN CAPP

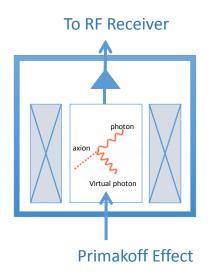
#### **ADVANCING HALOSCOPES**

Axion Conversion Power:

$$P = 1.5 \times 10^{-24} \,\mathrm{W} \left( \frac{g_{a\gamma\gamma}}{10^{15} \,\mathrm{GeV}} \right)^2 \left( \frac{C}{0.69} \right) \left( \frac{B}{8 \,\mathrm{T}} \right)^2 \left( \frac{V}{5l} \right) \left( \frac{\rho_a}{300 \,\mathrm{MeV/cm}^3} \right) \left( \frac{20 \,\mu\mathrm{eV}}{m_a} \right) \left( \frac{Q}{50 \times 10^3} \right)$$

Signal to Noise Ratio:

SNR = 
$$(P_a/P_N)\sqrt{bt} = (P_a/k_BT_S)\sqrt{t/b}$$
.



What can be improved?

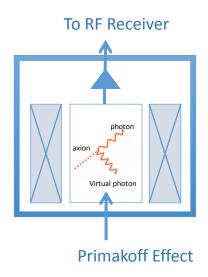
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What can be improved?

#### TECHNOLOGIES FOR ADVANCING HALOSCOPES

Axion Conversion Power:

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Signal to Noise Ratio:

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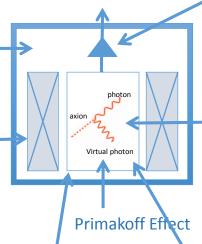




## **Cryogenics**

< 100mK with KAIST (Korea Advanced Institute of Science and Technology)

#### To RF Receiver



## **High Field SC Magnet**

12T -> 25T -> 35T / 40T with Korean company / BNL



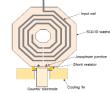
## Multiple cavities readout

increased volume in high frequencies



## **SQUID Amplifier**

quantum limit for noise [50mK·f(GHz)]
with KRISS (Korea Research Institute of Standards and Science)



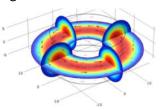
## **Superconducting Cavity**

superconducting coating, high Q-factor with KAIST (Korea Advanced Institute of Science and Technology)



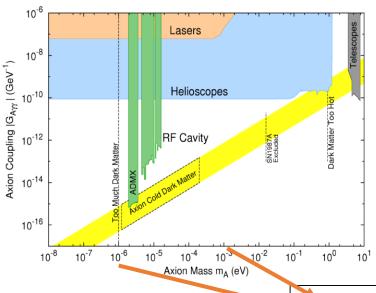
## **Giant toroidal cavity**

gain 10-20x



#### **CAPP ROADMAP**

## **Expected coverage:**

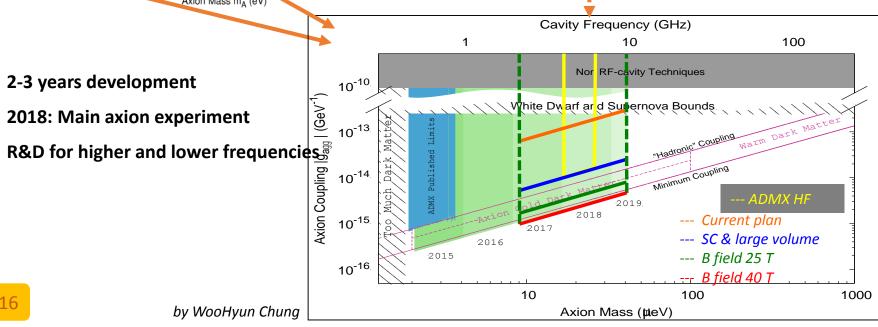


# Major improvements in haloscopes expected within 5 years

## **CAPP Projected Sensitivity**

e.g. 25µeV: 5.8GHz





#### **BEYOND DIRECT DETECTION: ARIADNE EXPERIMENT**

#### **ARIADNE:**



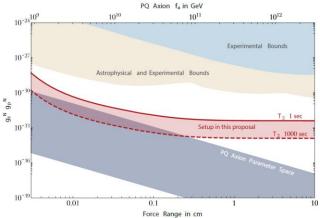
by YunChang Shin

Axion Resonant InterAction DetectioN Experiment

A rotating mass will induce spin precession and magnetization detectable by SQUIDS. Explores monopole-dipole interactions in the axion lagrangian.

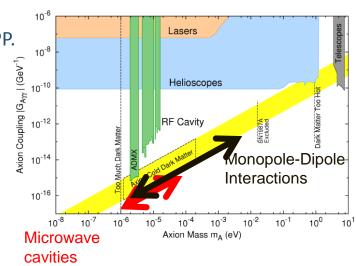
(No dark matter assumed.)

A. Arvanitaki, A. A. Geraci, arXiv:1403:1290

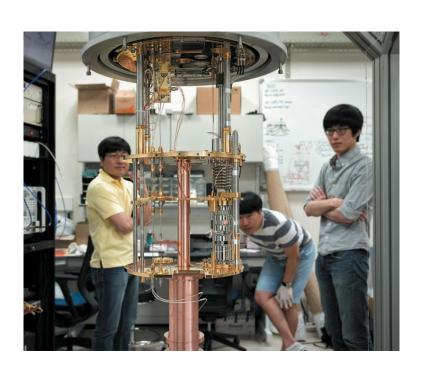


Projected reach for monopole-dipole axion mediated interaction

- Collaboration: Indiana, Nevada, Stanford Universities, CAPP.
- R&D in progress.
- Data taking in 2019.
- Poster at Patras2016



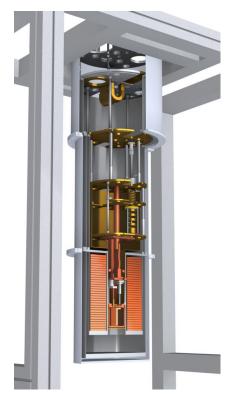
# HALOSCOPE PROGRAMME IN CAPP: STATUS



#### **CULTASK: CAPP ULTRA-LOW TEMPERATURE AXION SEARCH IN KOREA**

- New lab space February -> June 2016
- Two Bluefors dilution refrigerators LD400
- Nominal temperature 10mK (~50mK with load)
- 8T magnet
- Cu cavity. 9cm  $\phi$  -> ~2GHz -> ~10 $\mu$ eV axion mass
- Sapphire tuning rod





by WooHyun Chung



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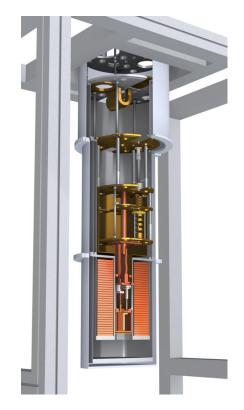




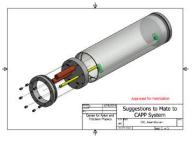


- Cavity R&D
- Reasonable sensitivity data expected in 2016.

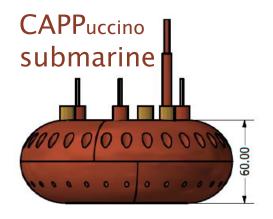




by WooHyun Chung



#### **CAPPuccino SUBMARINE**







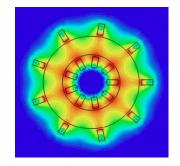


- Major radius 1.4cm, minor radius 0.2cm.
- Precursor to small toroid
  - Magnetic field 12T. Major radius 50cm, minor radius 10cm, volume ~80 liters.
  - 1.3-1.8 GHz.
- Precursor to giant toroidal cavity
  - Magnetic field 5T. Major radius 2m, minor radius 0.5m, volume ~10,000 liters.
  - Less than 1GHz (axion mass ~4μeV).

## Why toroid:

- Better form factor (EM field coverage)
- · Less fringe magnetic field
- No losses from mode crossing (between different EM modes)

Talk at Patras2016





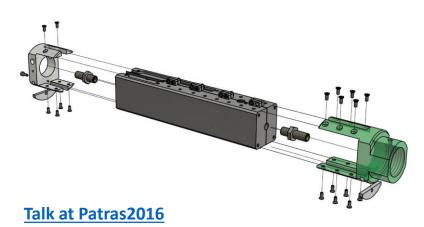
#### **CAPP-CAST EXPERIMENT**

Collaboration with the CAST experiment at CERN.



- Rectangular cavity inside the dipole magnet bore.
- Frequency 5-6 GHz. Axion mass 20-24 μeV.
- Cavity installed in June '16.

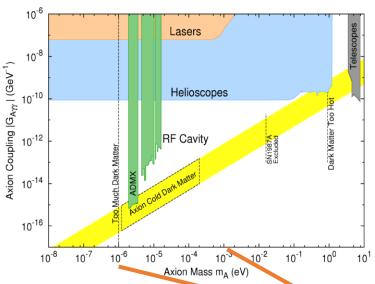






## **SUMMARY (ROADMAP, AGAIN)**

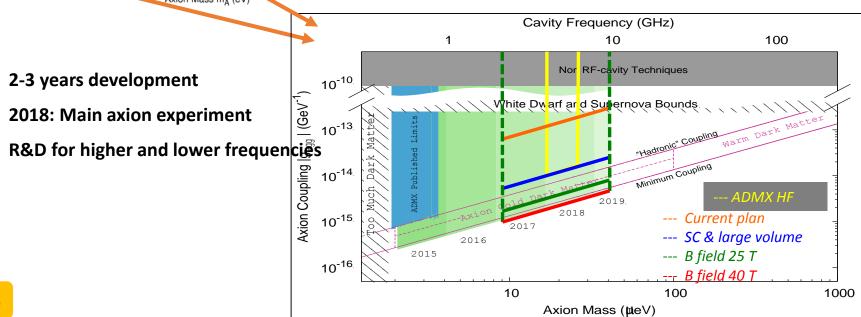
## **Expected coverage:**



# Major improvements in haloscopes expected within 5 years

## **CAPP Projected Sensitivity**





# CAPP/IBS



## The IBS Center for Axion and Precision Physics Research (KAIST campus, Daejeon)



## Institute for Basic Science (IBS):

- Founded in November 2011.
- Currently 26 Centers.
- https://www.ibs.re.kr/



http://capp.ibs.re.kr/

- Director: Prof. Yannis K. Semertzidis
- Currently:

~15 research fellows + visitor program,

~10 students, ~5 administrators.

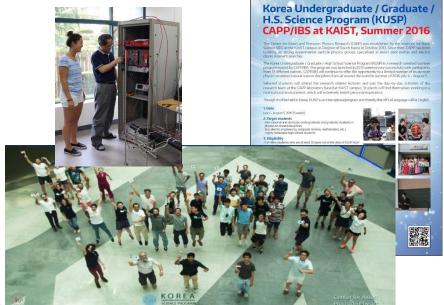




## The IBS Center for Axion and Precision Physics Research (KAIST campus, Daejeon)



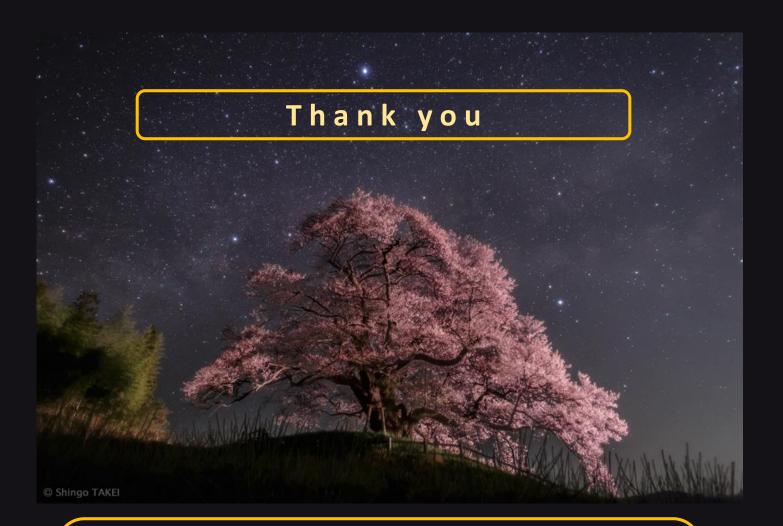
**Patras 2016** 



KUSP

- Green-field center. Motivation to play leading role in axion searches.
- Cutting edge research (cryogenics, SC, quantum electronics, precision accelerator physics and more).
- Next generation of direct axion detection with RF cavities.
- Closing in on theoretical limit of hadronic EDMs.
- Scientific potential, education & outreach.

The two sides of axionic CP-violation and DM searches.



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