

Experimental SUSY searches (~LHC)

Lecture 3

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- Lecture 1: the basics
 - what do I mean by SUSY?
 - what happens in an LHC collision?
 - how do the LHC detectors work?
 - which objects can we use to hunt for SUSY?
 - what (roughly) does SUSY look like at the LHC?
- Lecture 2 (mostly for experimentalists)
 - how are the LHC SUSY analyses designed?
 - discriminating variables
 - background measurements
 - detailed worked example (0 lepton search)
 - summary of different approaches
- Lecture 3 (mostly for theorists)
 - what have we learned about SUSY from the LHC (and LEP, etc...)
 - how can a theorist constrain a new model of SUSY (or other BSM physics?)

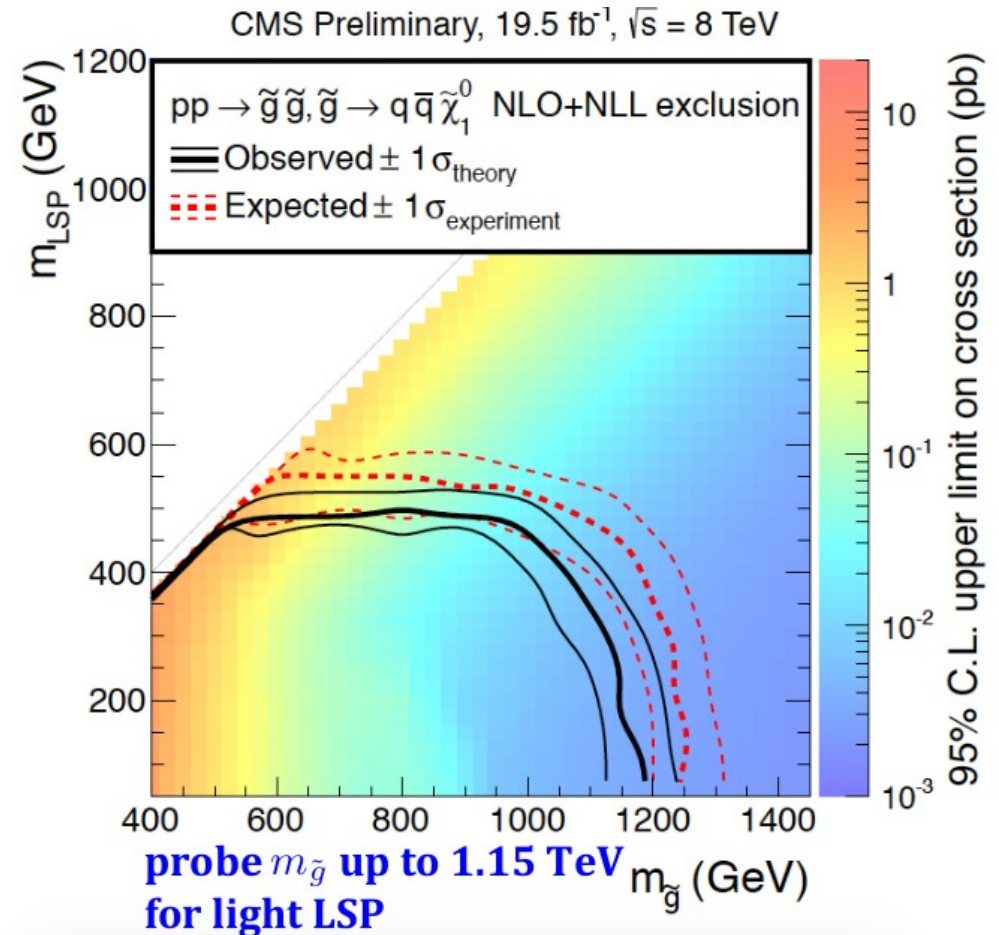
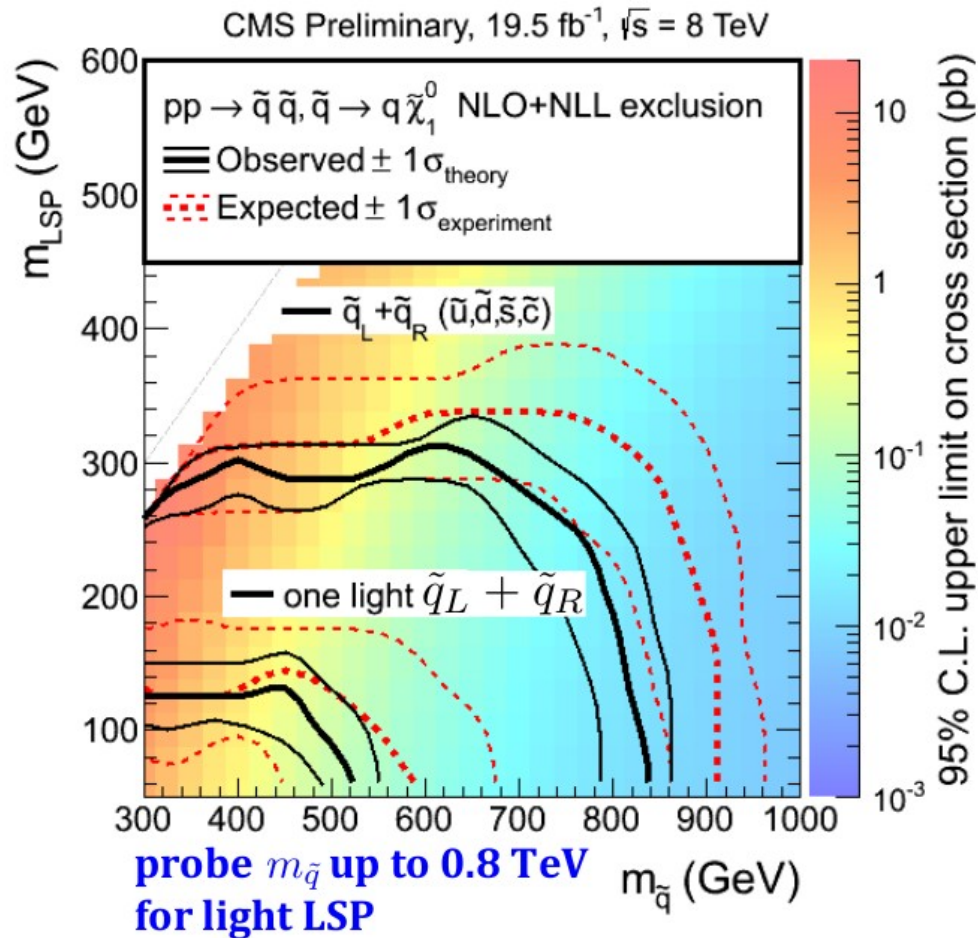
What have we learned about SUSY from the LHC?

- This question is extremely hard to answer
 - we will struggle to answer it in the next hour
- Each search is optimised on a specific scenario
 - no search has seen evidence for SUSY
- Unfortunately this does not mean that SUSY is “ruled out”
 - there are lots of gaps between the limits
 - extreme optimisation means poor coverage of general models
- Will study instead how you should use LHC results to constrain SUSY models
 - some insight can be gained from random explorations of the model space
 - a proper answer requires a global fit

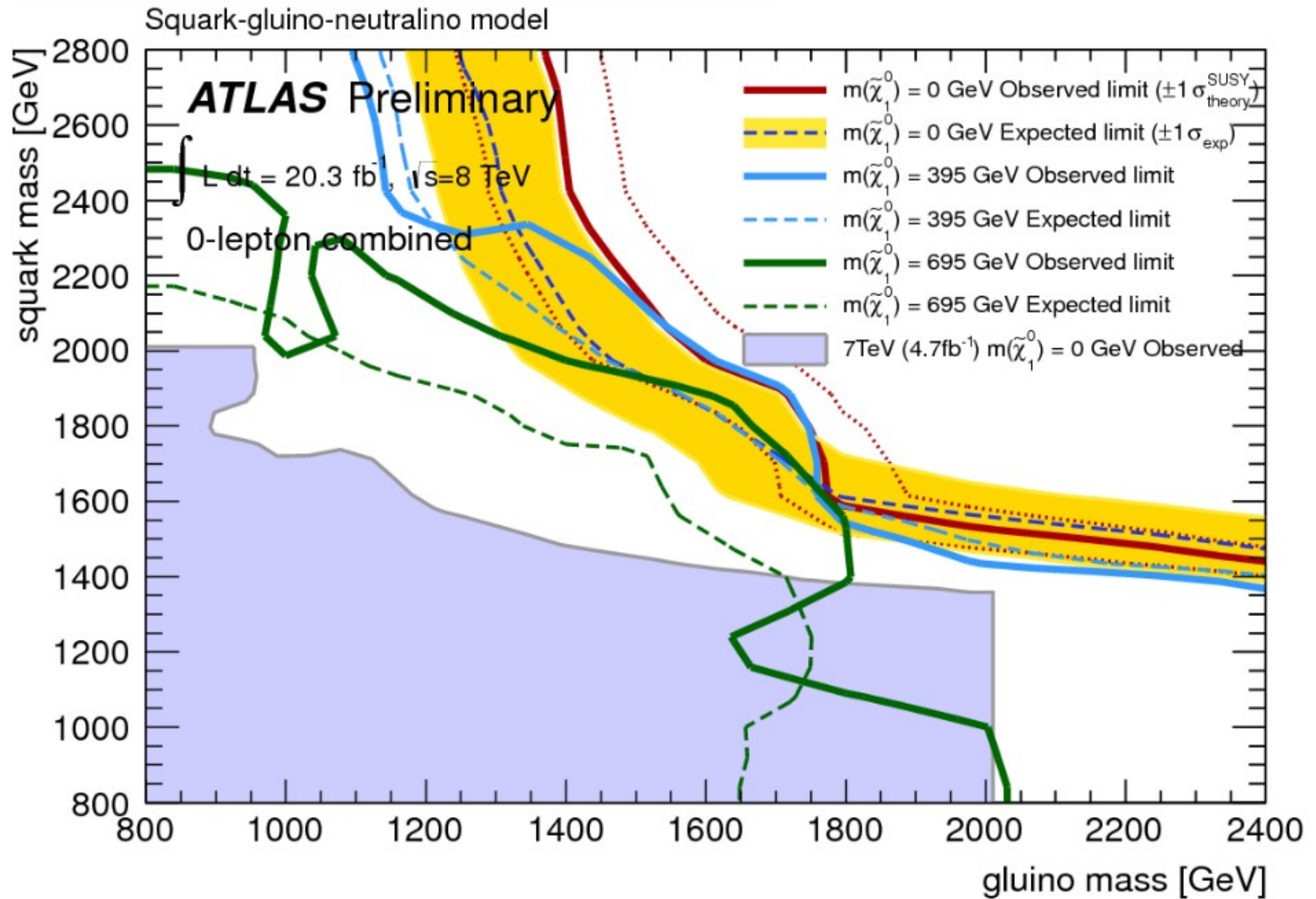
Simplified models

- For several years now, experimental searches have relied on simplified models
 - decouple most sparticles
 - scan two masses per plane
 - fix branching ratio to 100% for a specific process
- Simplified models are used for two specific things
 - optimisation of analysis (based on signal kinematics in the plane)
 - presentation of limits
- Let's quickly review the Run I simplified model limits

CMS multijets (CMS-SUS-13-012)



Squarks + gluinos gives tighter bounds (ATLAS-CONF-2013-047)



Light staus – Jet + MET + at least 1 hadronic tau

GMSB:

Gravitino is LSP (weak coupling to NLSP)

Signature depend on NLSP, e. g. $\tilde{\tau} \rightarrow \tau \tilde{G}$

Common signature:

Strong production (squark & gluinos) dominates \rightarrow jets; light staus $\rightarrow \tau\tau$; LSPs \rightarrow MET

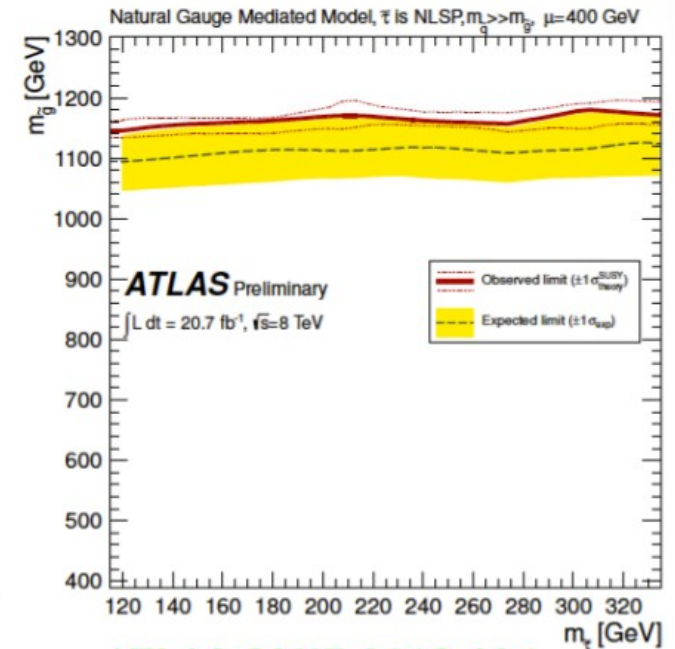
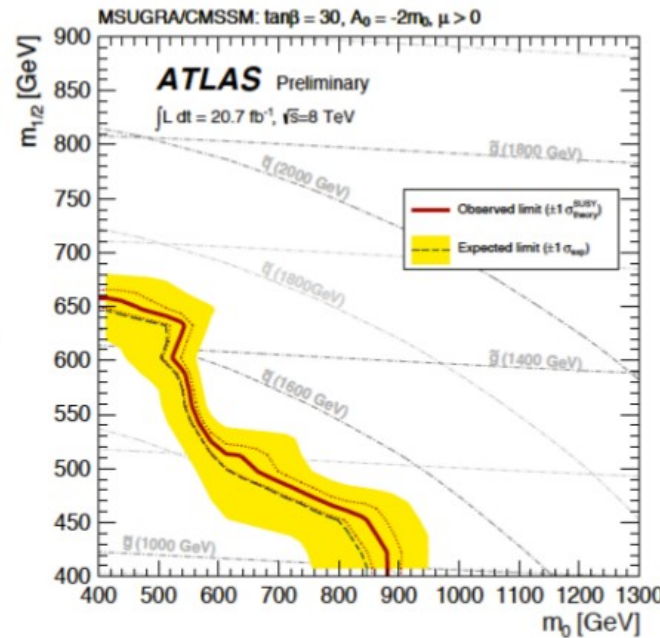
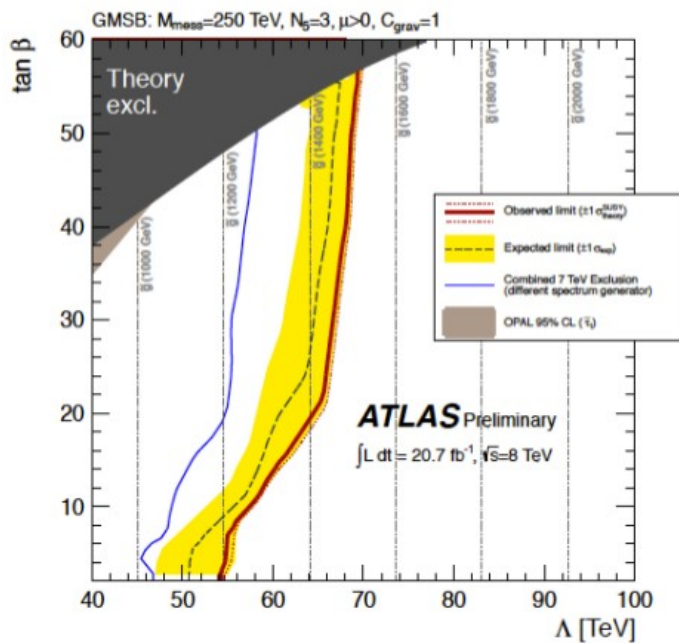
cMSSM:

Fix A_0 , $\tan \beta$ and μ such that $m_H \sim 125$ GeV

For some models NLSP = stau $\rightarrow \tau\tau$ in decay chain

Natural GM:

General Gauge Mediation + set weak scale parameters such to avoid fine tuning \rightarrow Light staus, accessible gluinos, other sparticles heavy

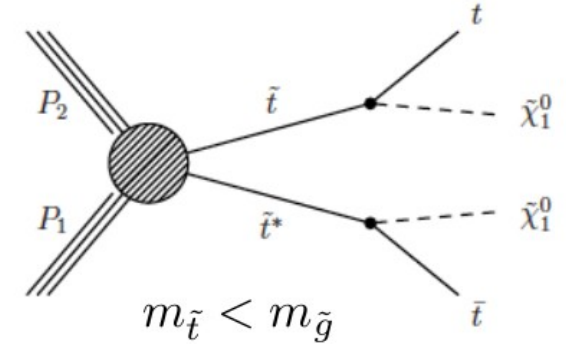
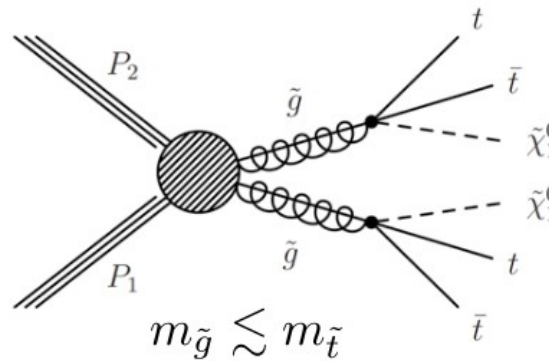
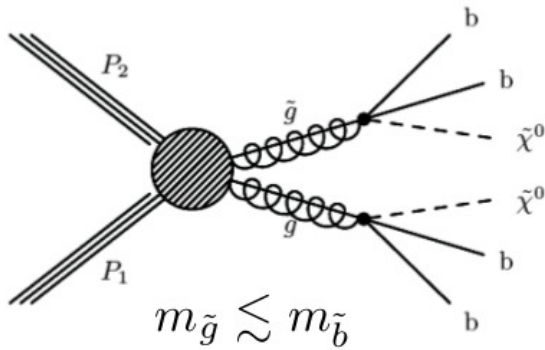
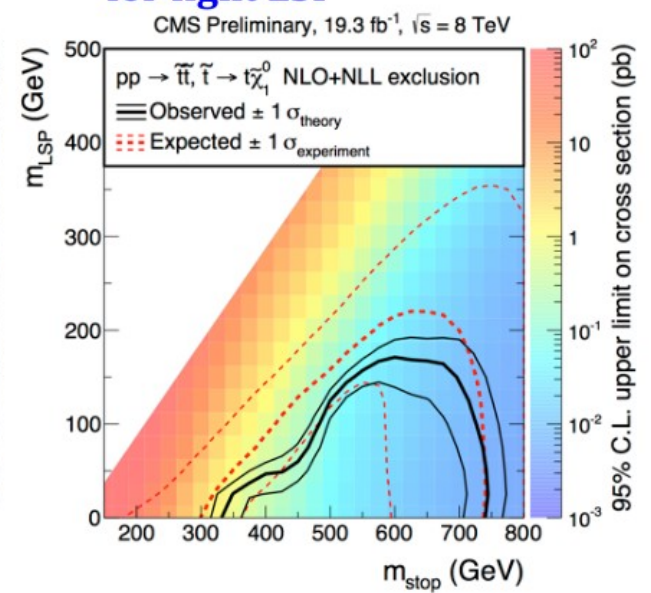
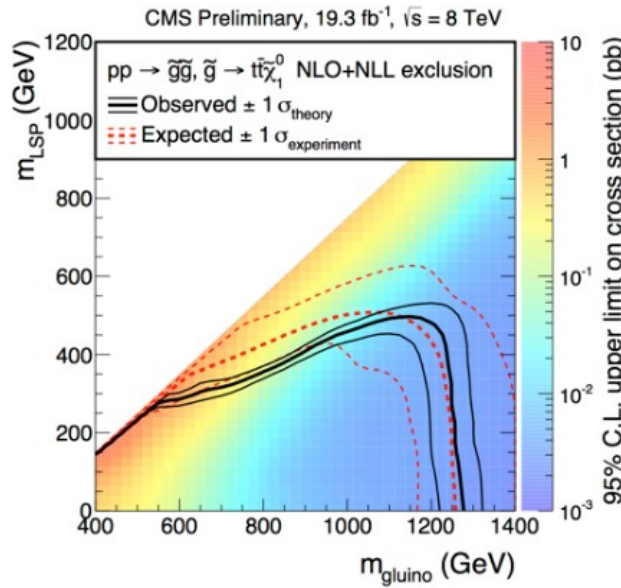
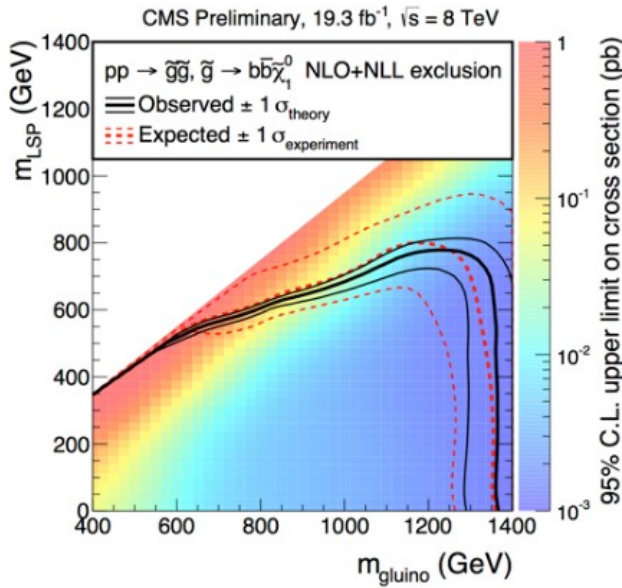


ATLAS-CONF-2013-026

Extrapolation in various “boxes”: with 0, 1, or 2 leptons; $N_{b\text{-tags}}$ and N_{jets}

No excess \rightarrow Limits in various SMSs

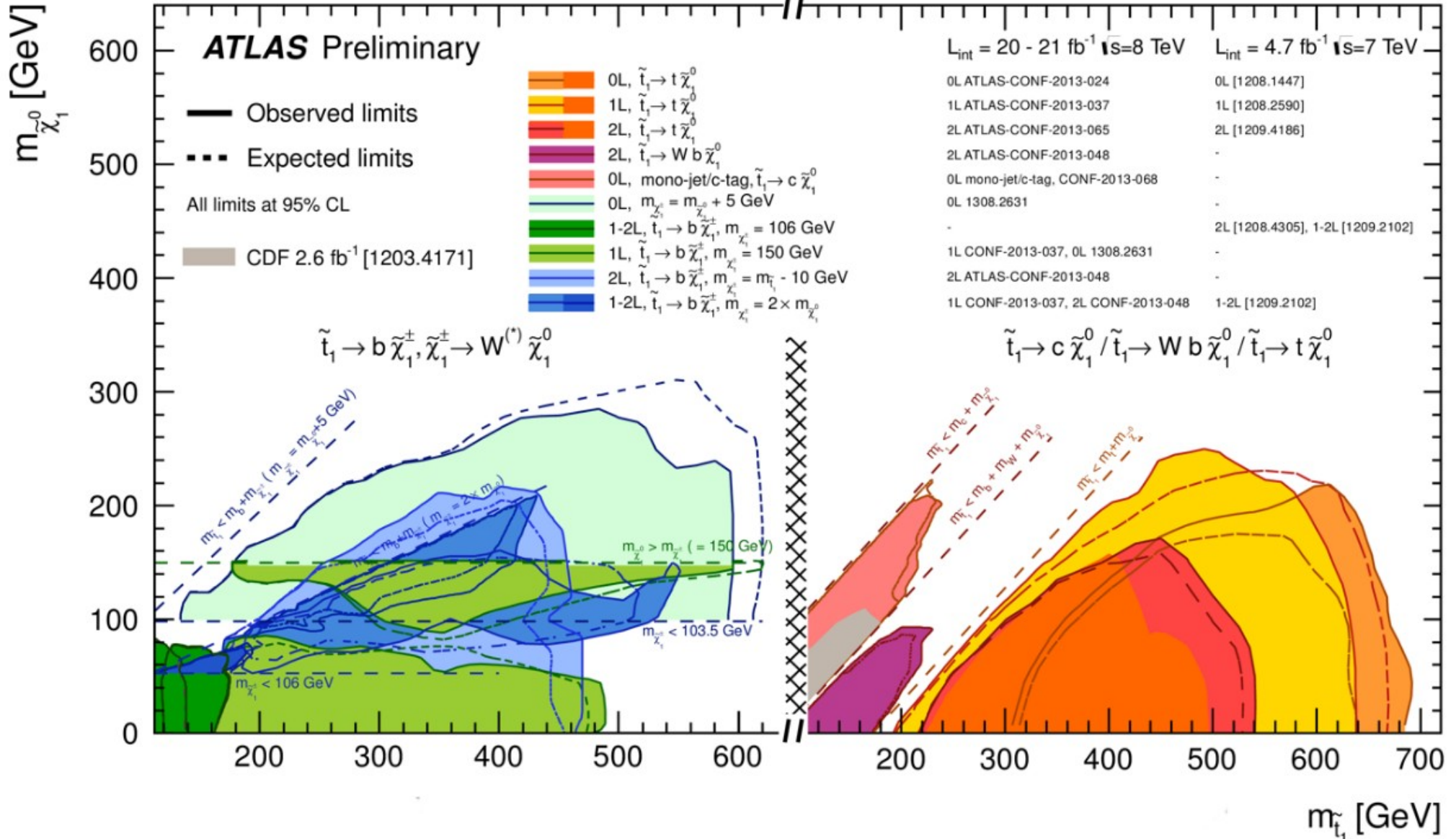
Probe $m_{\tilde{t}}$ up to 0.7 TeV for light LSP



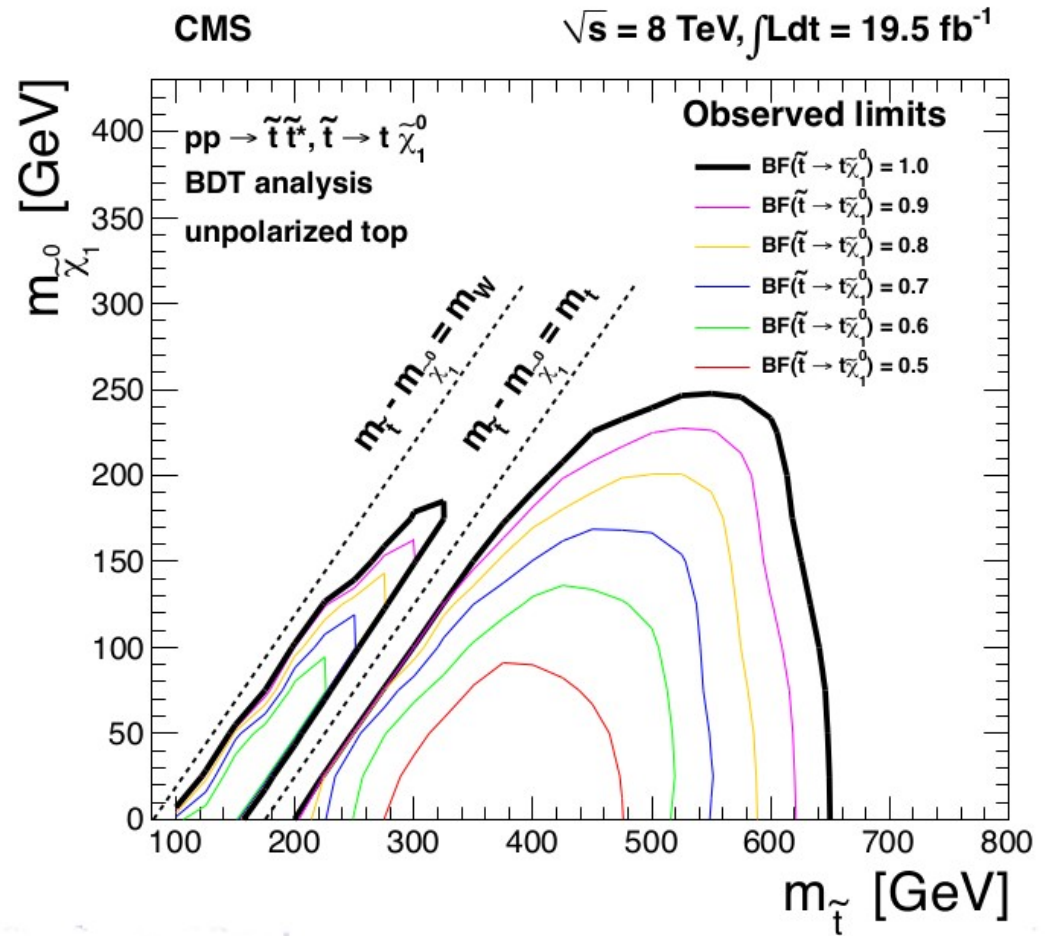
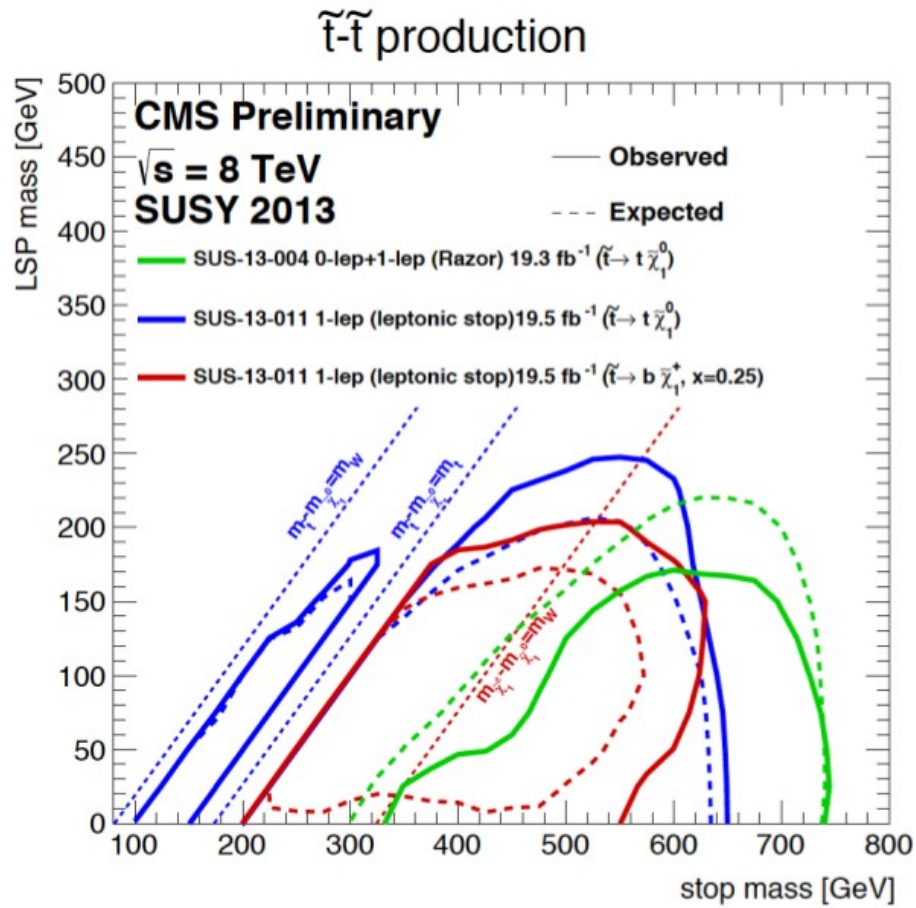
ATLAS stop search summary

$\tilde{t}_1, \tilde{t}_1^*$ production

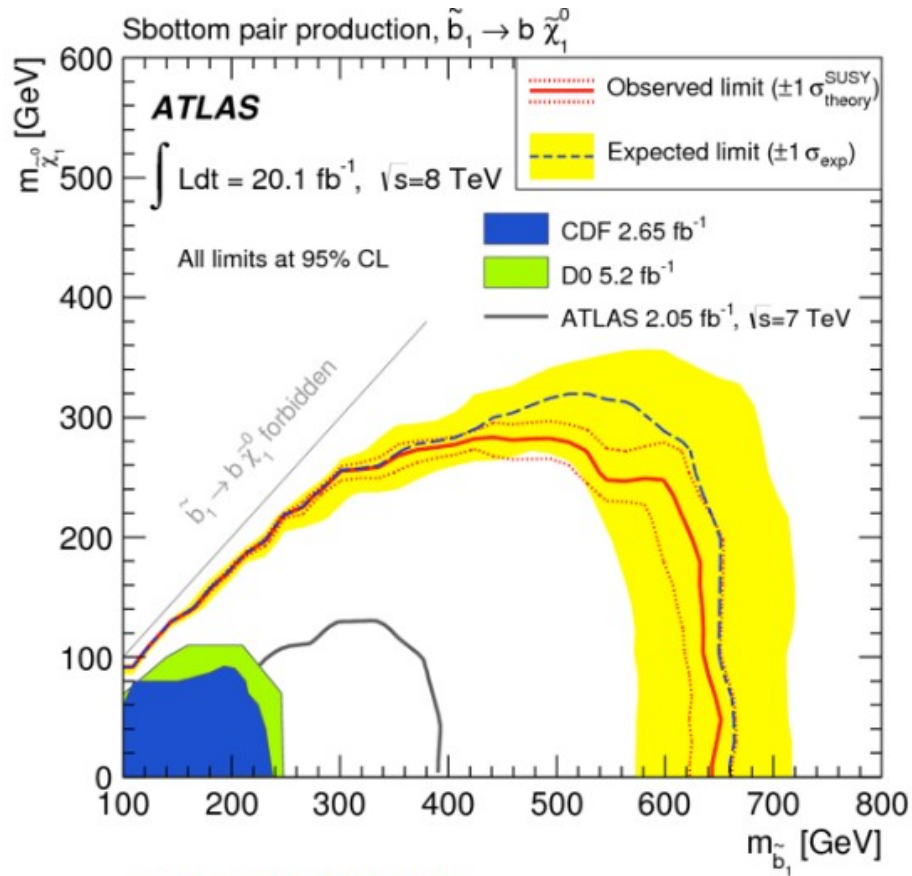
Status: SUSY 2013



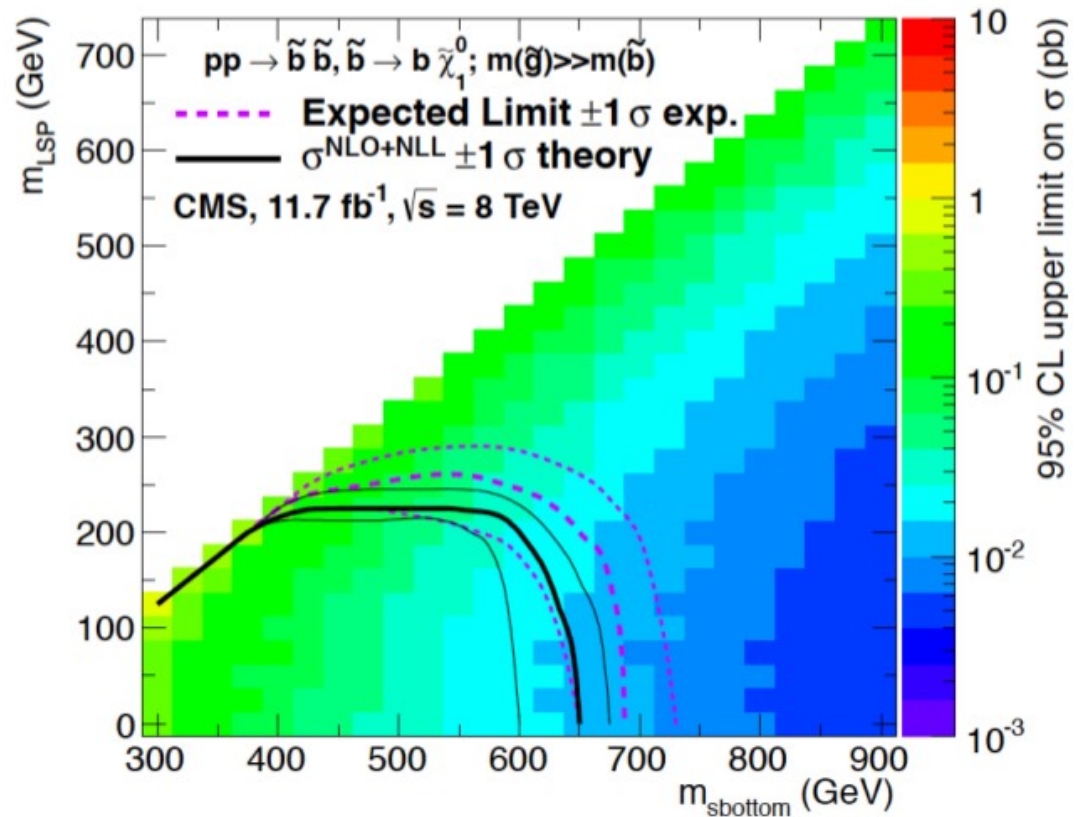
CMS stop summary



Sbottom searches

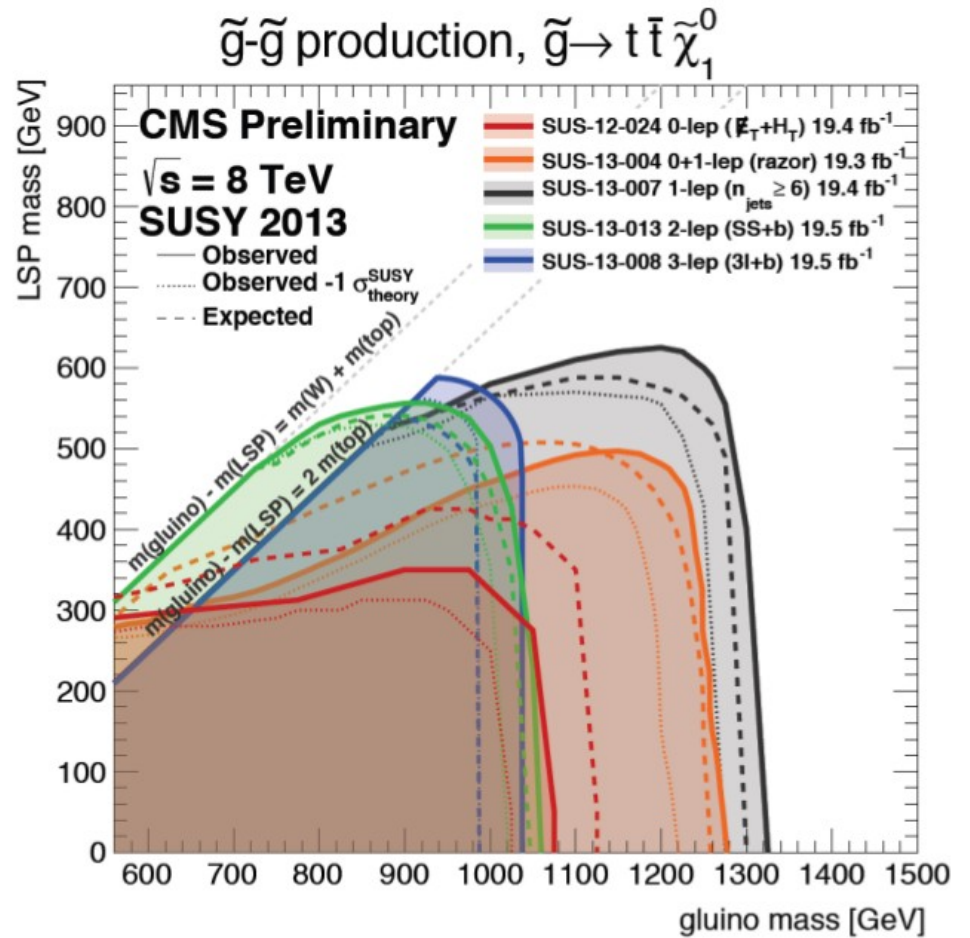
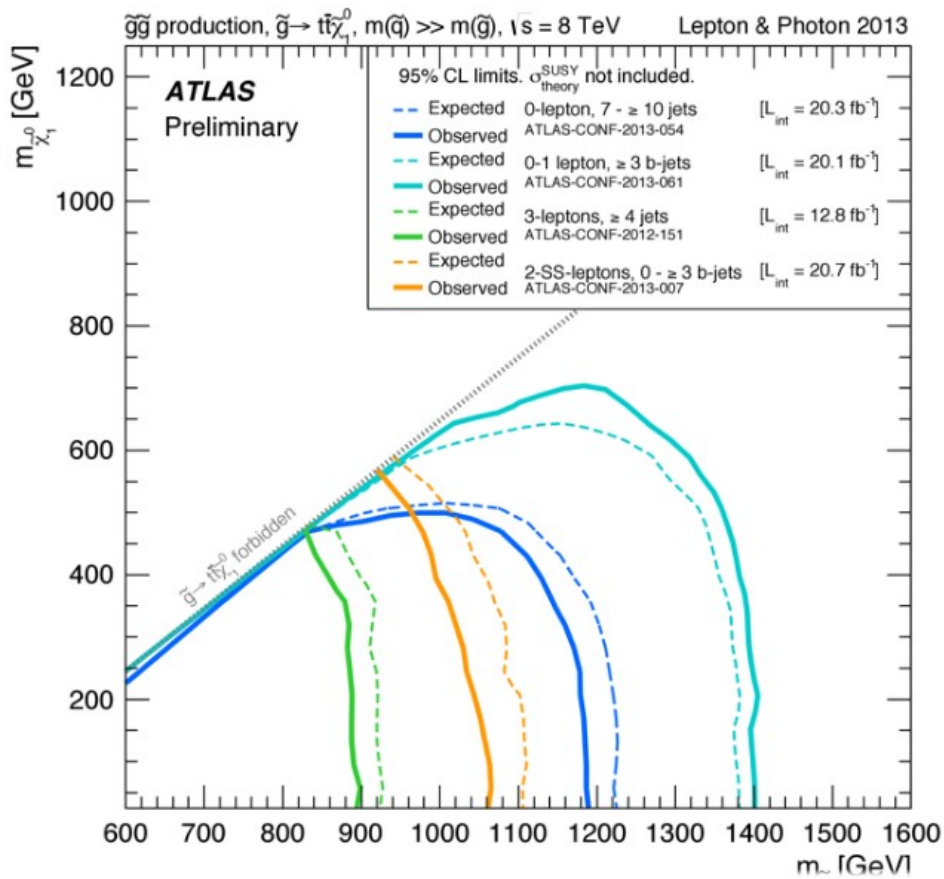
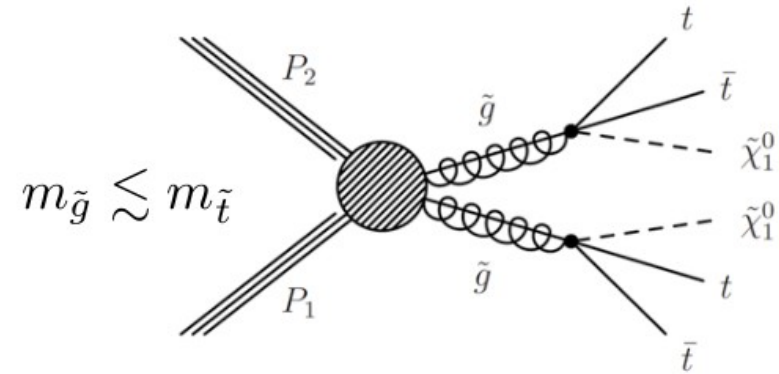


arXiv:1308.2631

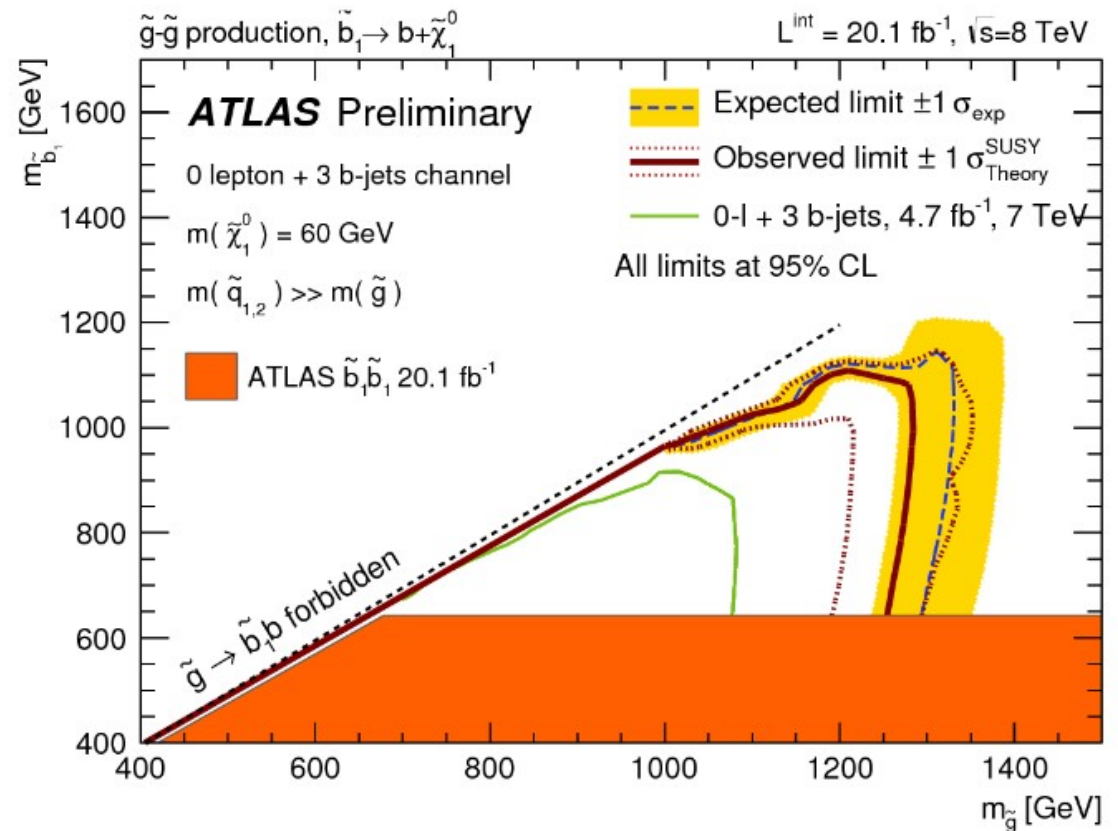
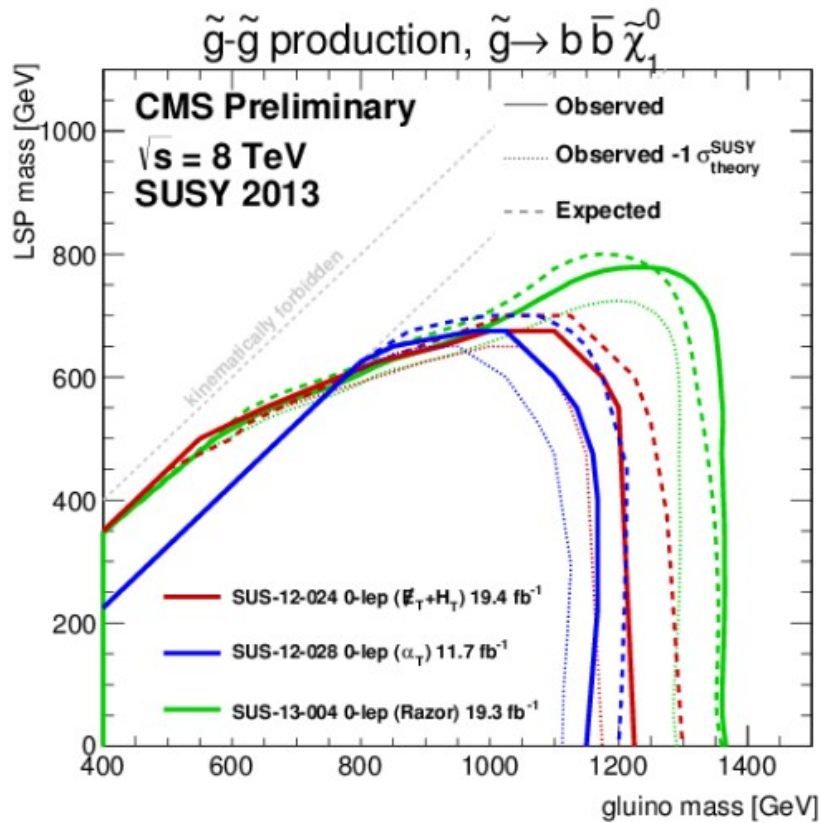
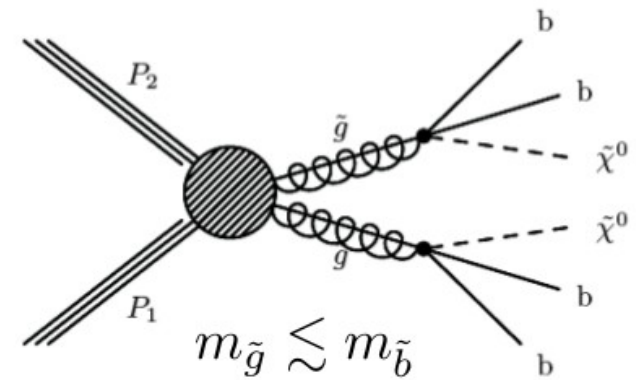


arXiv:1303.2985 (accepted by EPJC)

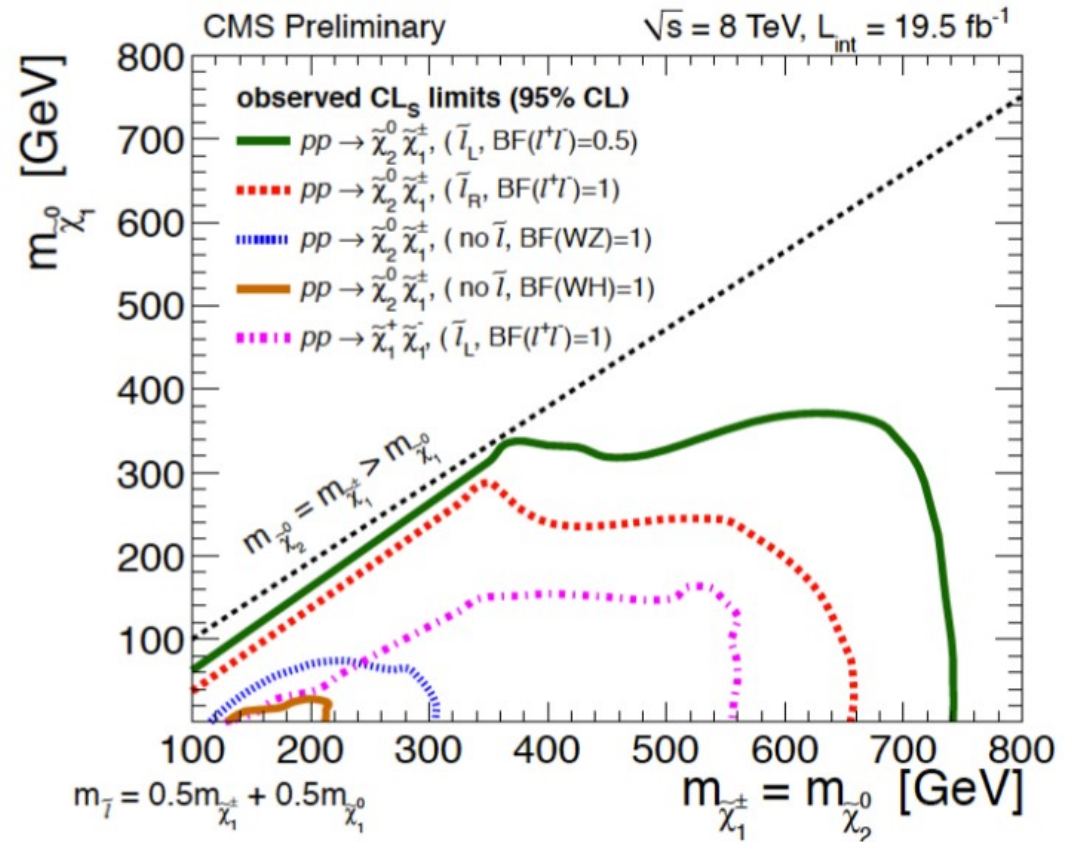
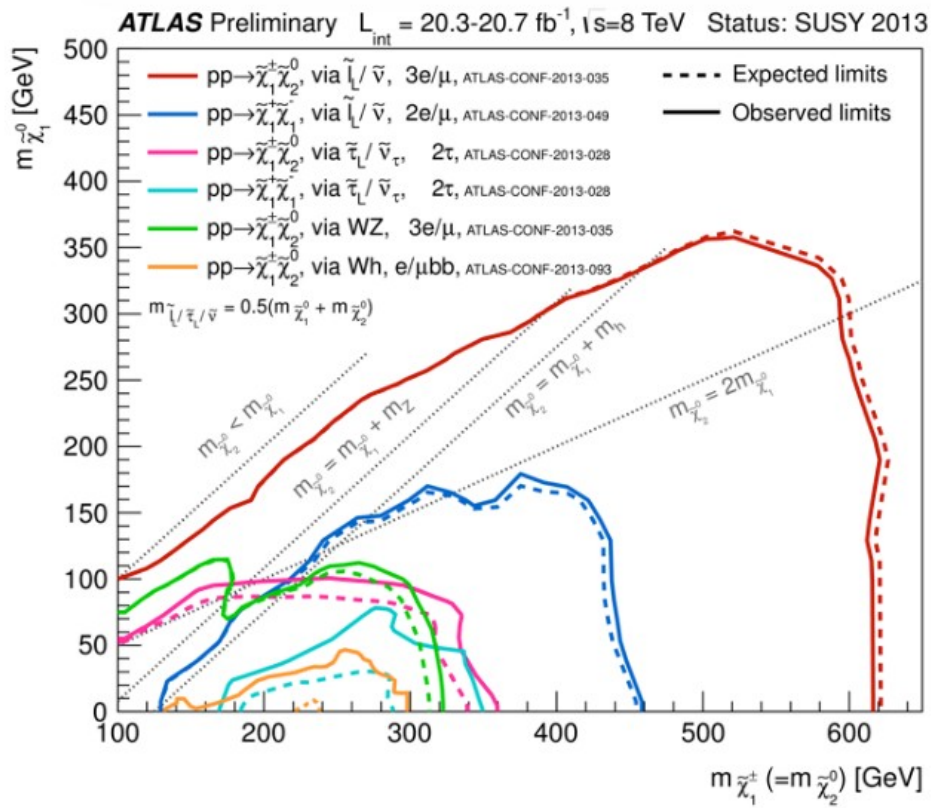
Four top searches



Four bottom searches (ATLAS-CONF-2013-061)

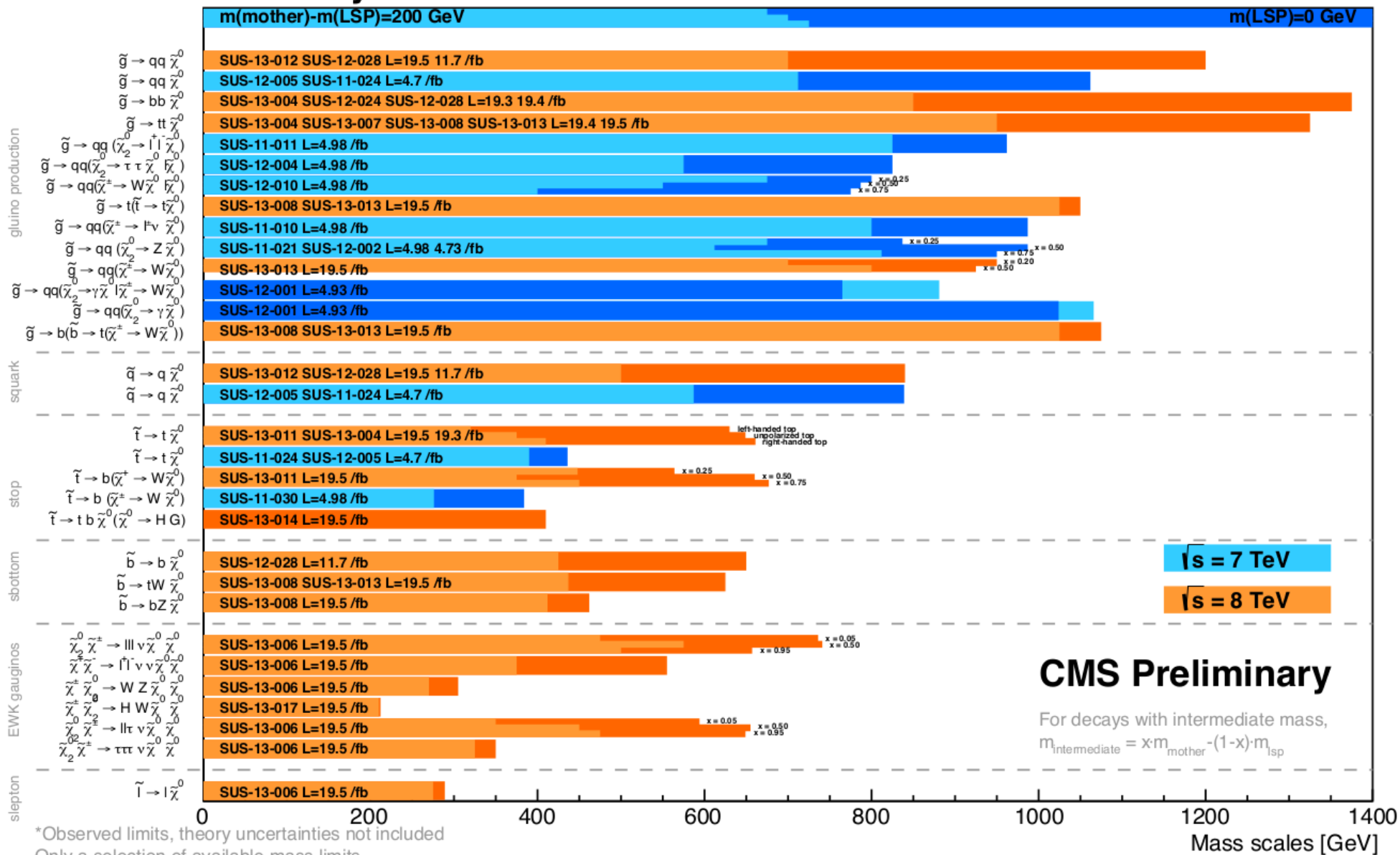


Multilepton searches



Summary of CMS SUSY Results* in SMS framework

SUSY 2013



*Observed limits, theory uncertainties not included
 Only a selection of available mass limits
 Probe *up to* the quoted mass limit

CMS Preliminary

For decays with intermediate mass,
 $m_{\text{intermediate}} = x m_{\text{mother}} - (1-x) m_{\text{LSP}}$

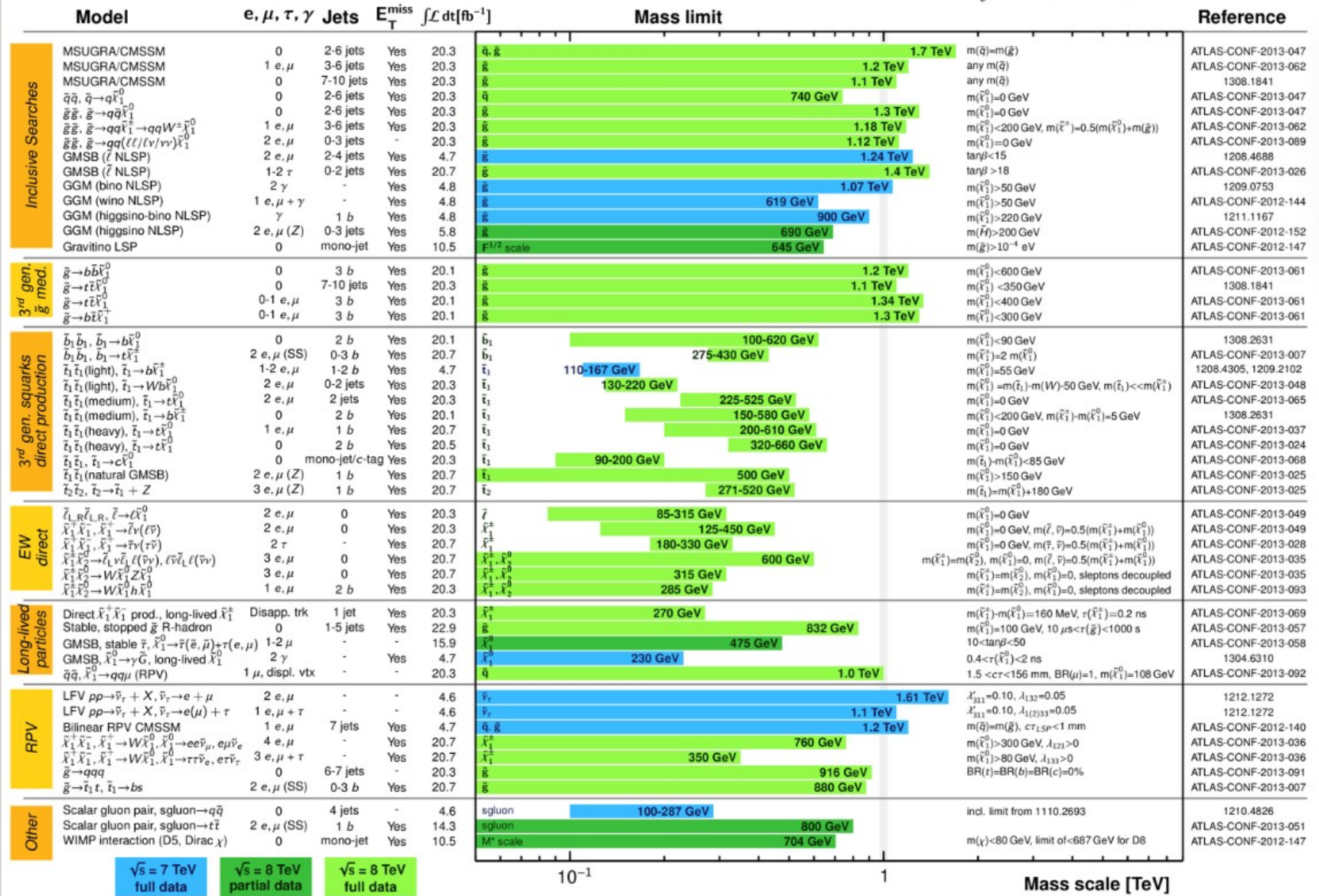
Total overview (ATLAS)

ATLAS SUSY Searches* - 95% CL Lower Limits

Status: SUSY 2013

ATLAS Preliminary

$\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1}$ $\sqrt{s} = 7, 8 \text{ TeV}$



*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1σ theoretical signal cross section uncertainty.

So SUSY is excluded then?

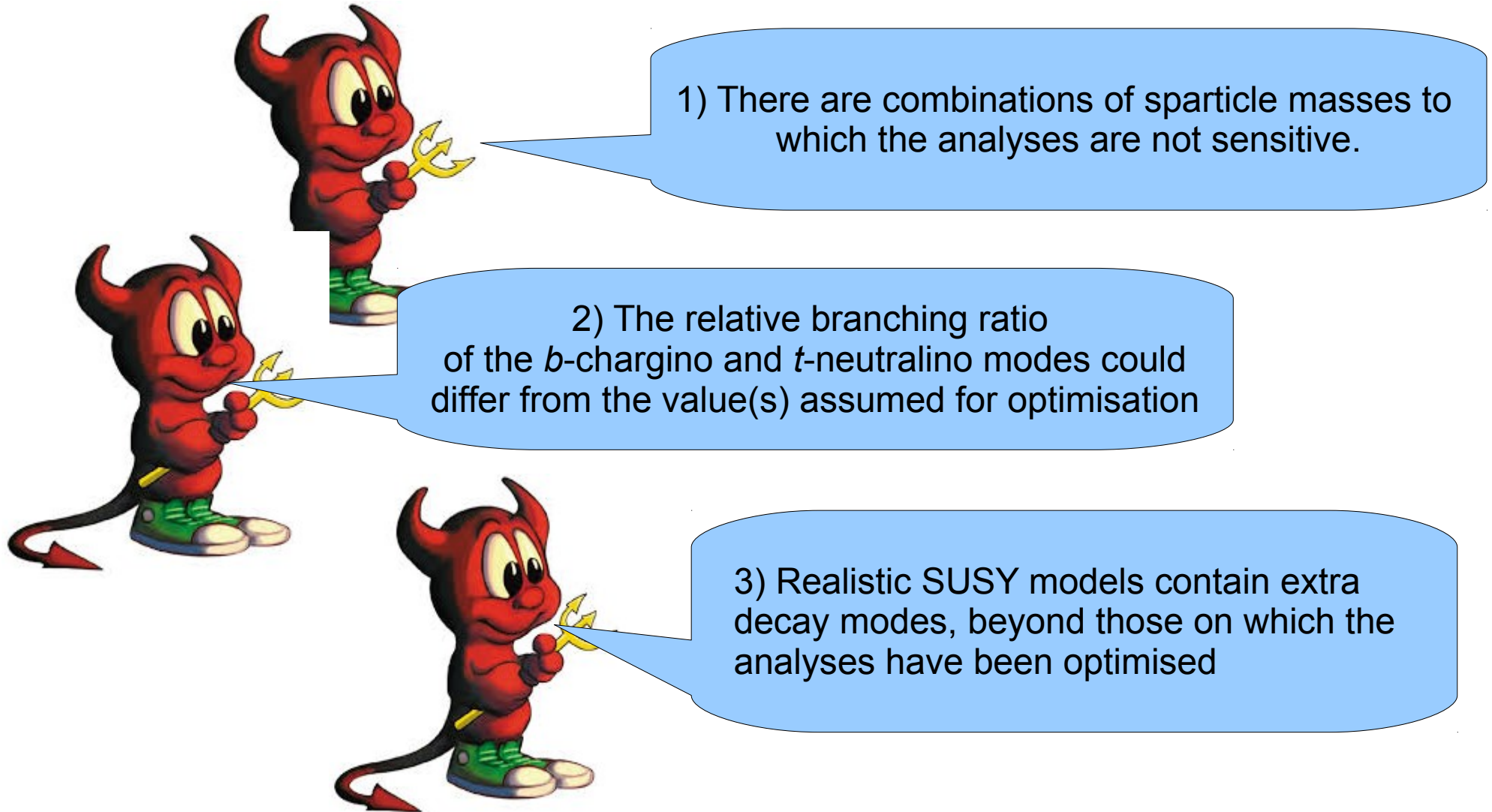
NOT SO FAST!

DISCLAIMER: The following work is taken from a quick exercise on my laptop for illustration. It is not endorsed by the ATLAS collaboration.

Stop searches: What's next?

- Naively the ATLAS results remove room for natural SUSY

- things are not as they first appear



A quick scan

1) Have taken the ATLAS simplified model space and explored *all four parameters*

$$m_{\tilde{t}}, m_{\tilde{\chi}_1^\pm}, m_{\tilde{\chi}_1^0}, BR(\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0)$$

2) Use Multinest Bayesian sampling algorithm, using the following analyses:

- **ATLAS-CONF-2013-024:** The ATLAS 0 lepton direct stop analysis using 20 fb⁻¹ of data [1].
- **ATLAS-CONF-2013-037:** The ATLAS 1 lepton direct stop analysis using 20 fb⁻¹ of data [2].
- **ATLAS-CONF-2013-048:** The ATLAS 2 lepton direct stop analysis using 20 fb⁻¹ of data, based on the use of the leptonic M_{T2} variable [3].
- **arXiv:1308.2631:** The ATLAS 2 b jets plus missing E_T search for direct stop and/or sbottom production, using 20 fb⁻¹ of data [4].
- The “mixed” signal regions from the 2014 paper update to the ATLAS 0 lepton direct stop analysis, using 20 fb⁻¹ of data [5].

• I do not include compressed spectra in the scan

- thus have not looked at monojet or c -neutralino searches
- this is due to limited time rather than sound physical reasoning...

Summary of scan details

- Run Multinest scanning algorithm with parameters in the following range

Parameter	Range
$m_{\tilde{t}}$	100 – 700 GeV
$m_{\tilde{\chi}_1^0}$	1 – 700 GeV
$m_{\tilde{\chi}_1^\pm}$	103 – 700 GeV
$BR(\tilde{t}_1 \rightarrow t\tilde{\chi}_1^0)$	0 – 1

- $(m_{\tilde{t}} - m_{\tilde{\chi}_1^\pm}) > 10 \text{ GeV}$.
- $(m_{\tilde{\chi}_1^\pm} - m_{\tilde{\chi}_1^0}) > 1 \text{ GeV}$
- **Top quarks must be on shell!**

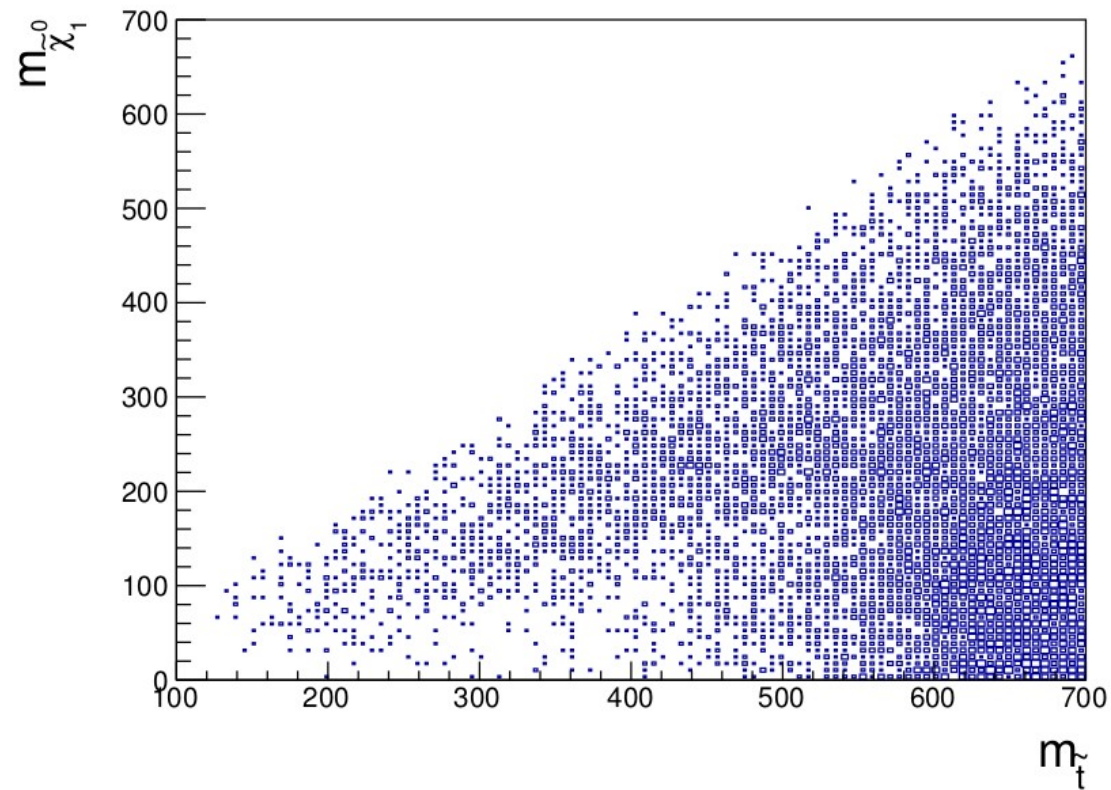
- At each point, generate 10,000 stop pair events

- have parallelised Pythia 8 using Open MP
- use NLLFAST for the production cross-section
- use DELPHES with custom hacks for the detector simulation
- apply ATLAS analysis cuts and derive approximate likelihood
- parallelisation gives 10,000 events in ~3.5 seconds on 8 cores
- get a convergent fit with 2,000 live points within a few hours
- 10,000 live points \Rightarrow ~ 3 days of running on 24 cores

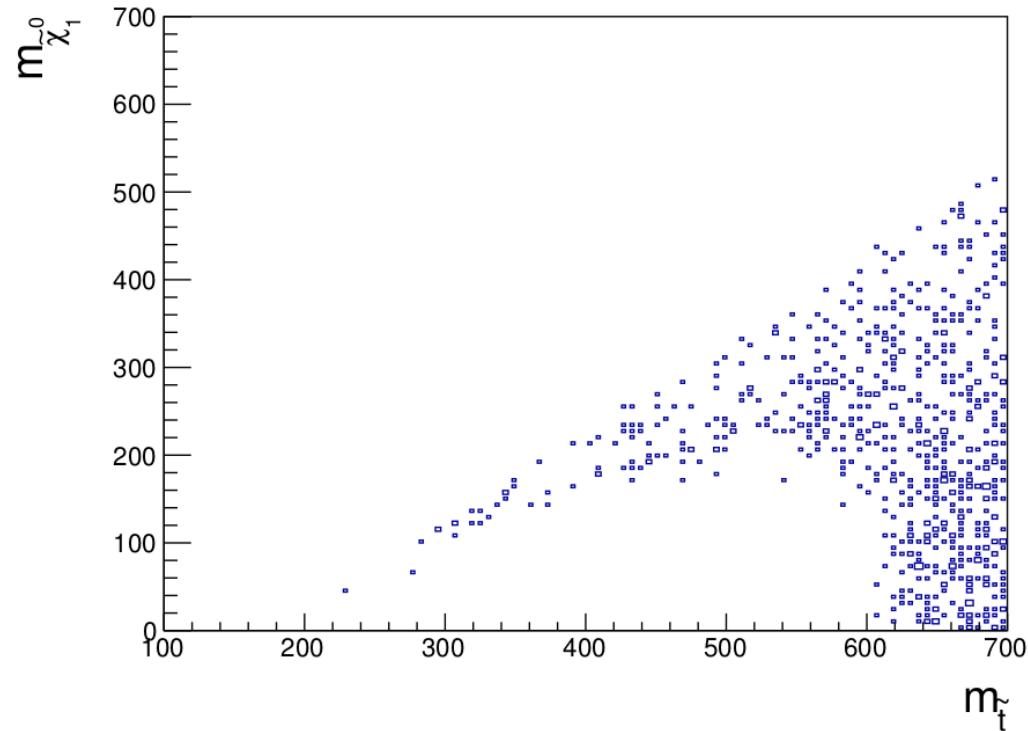
RESULTS

- A set of posterior samples in the 4D stop simplified model space
- All have likelihood 1 by construction \Rightarrow we will use them as a set of equally likely benchmark points

Results: Stop-Neutralino mass plane

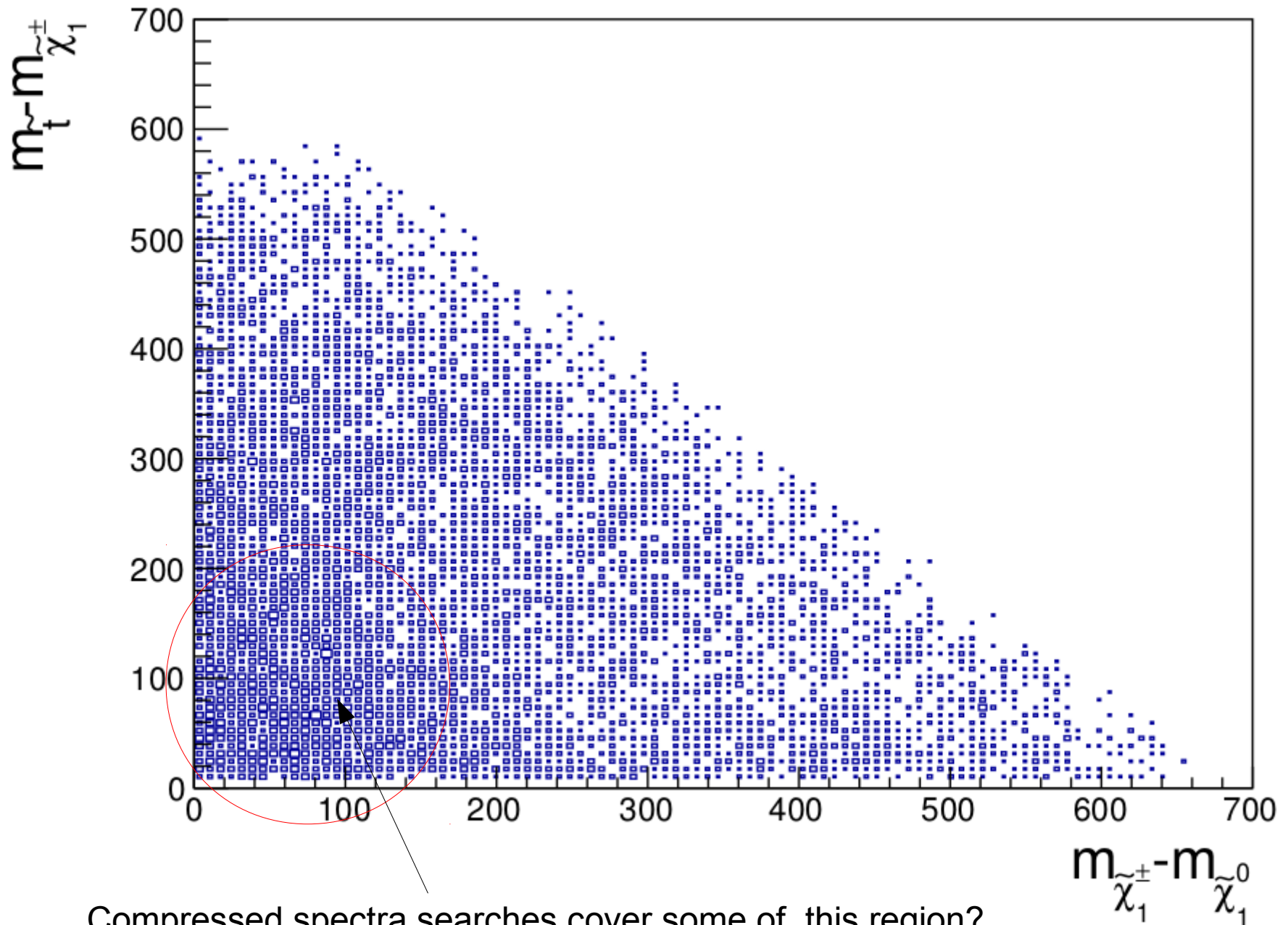


Results: Stop-Neutralino mass plane for t -neutralino decays



- This plot shows models with $\text{BR}(\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0) > 0.9$
 - we recover the ATLAS limit
 - provides a nice sanity check

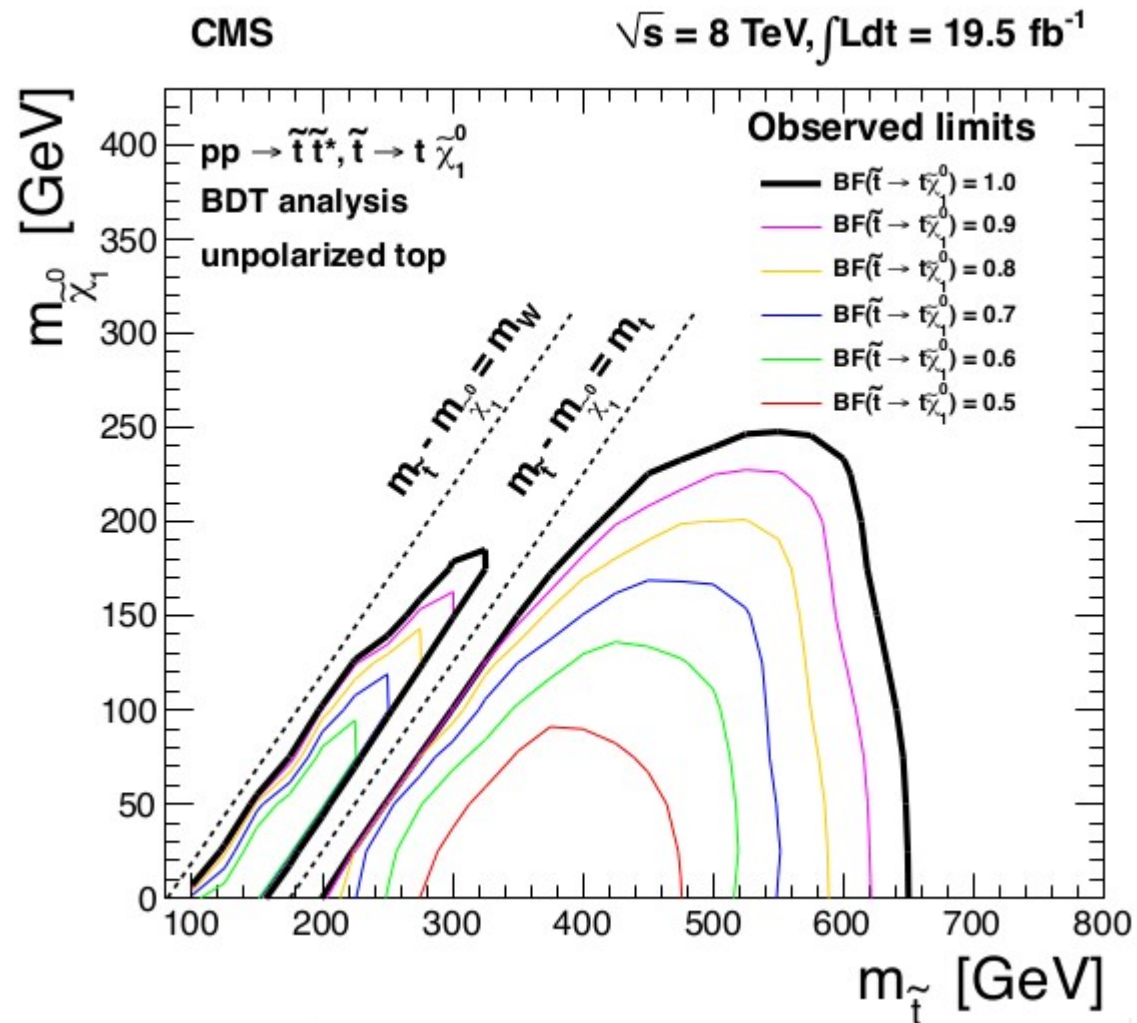
Results: Mass difference plane

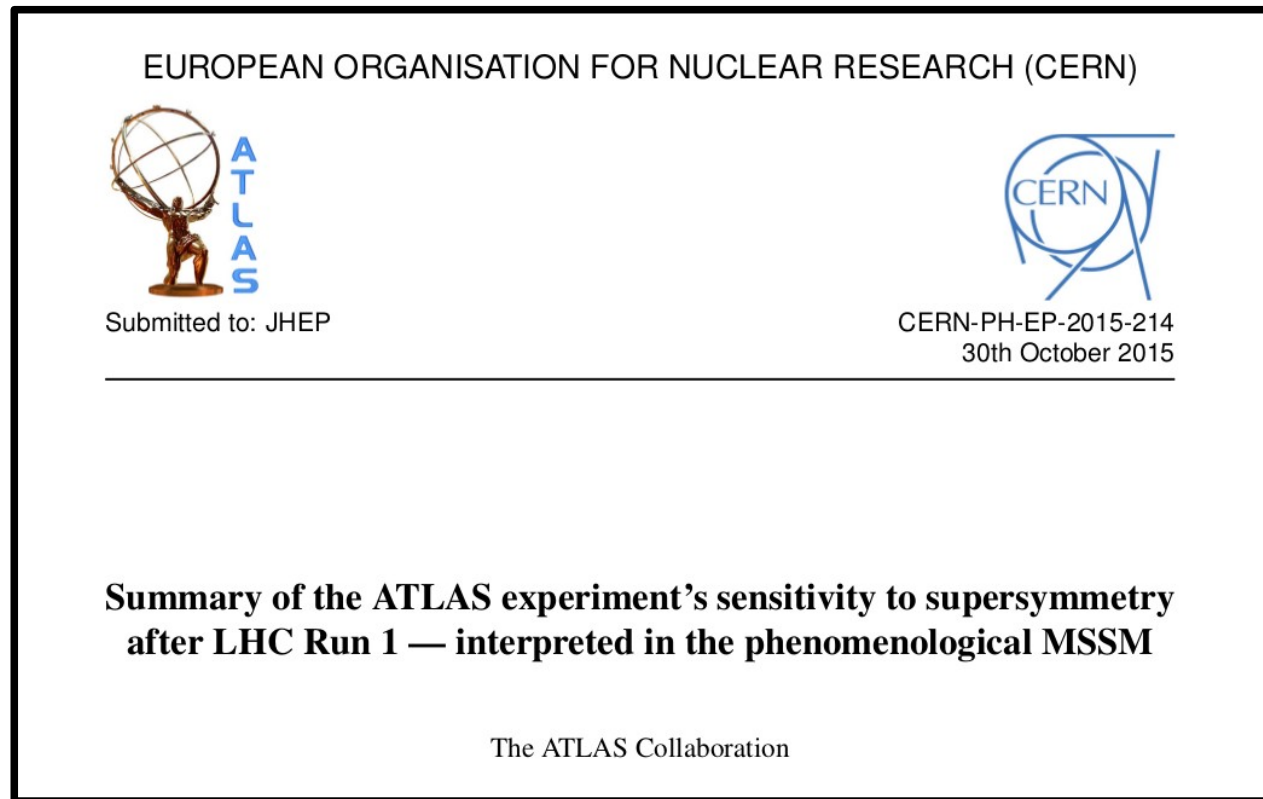


Compressed spectra searches cover some of this region?

A case study: ATLAS stop searches

- There is apparently *no* limit on the stop mass from ATLAS
- You actually see this in the official limits (for both CMS and ATLAS)





- Most complete summary of ATLAS limits: lots of searches included
 - full ATLAS likelihood included
 - no statistical combination of analyses however
- Based on a large number of points in the pMSSM
 - 19 parameter version of MSSM (unify 1st and 2nd gen squark masses)

List of included searches

Analysis	Ref.	Category
0-lepton + 2–6 jets + E_T^{miss}	[58]	Inclusive
0-lepton + 7–10 jets + E_T^{miss}	[59]	
1-lepton + jets + E_T^{miss}	[60]	
$\tau(\tau/\ell)$ + jets + E_T^{miss}	[61]	
SS/3-leptons + jets + E_T^{miss}	[62]	
0/1-lepton + 3b-jets + E_T^{miss}	[63]	
Monojet	[64]	
0-lepton stop	[65]	Third generation
1-lepton stop	[56]	
2-leptons stop	[66]	
Monojet stop	[67]	
Stop with Z boson	[68]	
2b-jets + E_T^{miss}	[69]	
$tb + E_T^{\text{miss}}$, stop	[57]	
ℓh	[70]	Electroweak
2-leptons	[54]	
2- τ	[55]	
3-leptons	[53]	
4-leptons	[71]	
Disappearing Track	[72]	
Long-lived particle	[73, 74]	Other
$H/A \rightarrow \tau^+ \tau^-$	[75]	

- Includes 7 TeV and 8 TeV results
- Most signal regions of each analysis are included

→ over 200 distinct signal regions!

pMSSM parameters and ranges

Parameter	Min value	Max value	Note
$m_{\tilde{L}_1} (= m_{\tilde{L}_2})$	90 GeV	4 TeV	Left-handed slepton (first two gens.) mass
$m_{\tilde{e}_1} (= m_{\tilde{e}_2})$	90 GeV	4 TeV	Right-handed slepton (first two gens.) mass
$m_{\tilde{L}_3}$	90 GeV	4 TeV	Left-handed stau doublet mass
$m_{\tilde{e}_3}$	90 GeV	4 TeV	Right-handed stau mass
$m_{\tilde{Q}_1} (= m_{\tilde{Q}_2})$	200 GeV	4 TeV	Left-handed squark (first two gens.) mass
$m_{\tilde{u}_1} (= m_{\tilde{u}_2})$	200 GeV	4 TeV	Right-handed up-type squark (first two gens.) mass
$m_{\tilde{d}_1} (= m_{\tilde{d}_2})$	200 GeV	4 TeV	Right-handed down-type squark (first two gens.) mass
$m_{\tilde{Q}_3}$	100 GeV	4 TeV	Left-handed squark (third gen.) mass
$m_{\tilde{u}_3}$	100 GeV	4 TeV	Right-handed top squark mass
$m_{\tilde{d}_3}$	100 GeV	4 TeV	Right-handed bottom squark mass
$ M_1 $	0 GeV	4 TeV	Bino mass parameter
$ M_2 $	70 GeV	4 TeV	Wino mass parameter
$ \mu $	80 GeV	4 TeV	Bilinear Higgs mass parameter
M_3	200 GeV	4 TeV	Gluino mass parameter
$ A_t $	0 GeV	8 TeV	Trilinear top coupling
$ A_b $	0 GeV	4 TeV	Trilinear bottom coupling
$ A_\tau $	0 GeV	4 TeV	Trilinear τ lepton coupling
M_A	100 GeV	4 TeV	Pseudoscalar Higgs boson mass
$\tan\beta$	1	60	Ratio of the Higgs vacuum expectation values

- Parameters are chosen at random, then constraints are applied
 - note: this is not statistically meaningful (we will return to this point)
 - allows at least some insight however

Non-LHC constraints

Parameter	Minimum value	Maximum value
$\Delta\rho$	-0.0005	0.0017
$\Delta(g-2)_\mu$	-17.7×10^{-10}	43.8×10^{-10}
$\text{BR}(b \rightarrow s\gamma)$	2.69×10^{-4}	3.87×10^{-4}
$\text{BR}(B_s \rightarrow \mu^+\mu^-)$	1.6×10^{-9}	4.2×10^{-9}
$\text{BR}(B^+ \rightarrow \tau^+\nu_\tau)$	66×10^{-6}	161×10^{-6}
$\Omega_{\tilde{\chi}_1^0} h^2$	—	0.1208
$\Gamma_{\text{invisible(SUSY)}}(Z)$	—	2 MeV
Masses of charged sparticles	100 GeV	—
$m(\tilde{\chi}_1^\pm)$	103 GeV	—
$m(\tilde{u}_{1,2}, \tilde{d}_{1,2}, \tilde{c}_{1,2}, \tilde{s}_{1,2})$	200 GeV	—
$m(h)$	124 GeV	128 GeV

- Direct detection also considered

- spin-independent: LUX
- spin-dependent: COUPP and Xenon 100

Neutralino type

LSP type	Definition	Sampled	Simulated		Weight
			Number	Fraction	
'Bino-like'	$N_{11}^2 > \max(N_{12}^2, N_{13}^2 + N_{14}^2)$	480×10^6	103,410	35%	1/24
'Wino-like'	$N_{12}^2 > \max(N_{11}^2, N_{13}^2 + N_{14}^2)$	} 20×10^6 {	80,233	26%	1
'Higgsino-like'	$(N_{13}^2 + N_{14}^2) > \max(N_{11}^2, N_{12}^2)$		126,684	39%	1
Total		500×10^6	310,327		

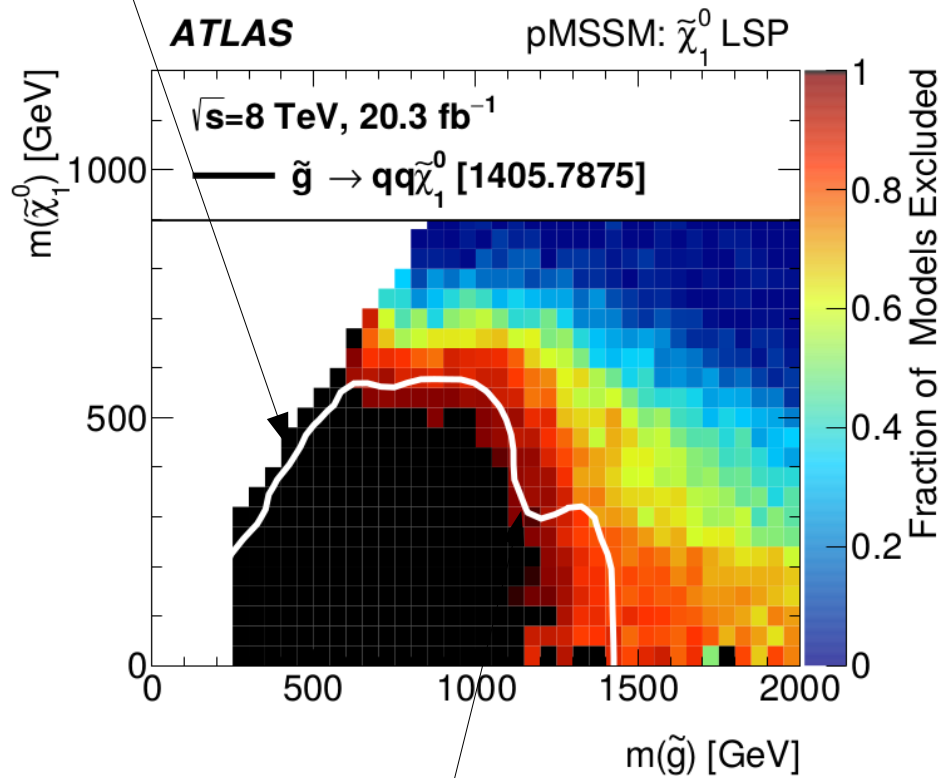
- Samples are chosen carefully to give \sim equal numbers of different neutralino types
- Most samples had pure neutralinos of a particular type
 - this is generically expected for their sampling method
 - does not mean in general that models with mixtures are not favoured
 - allows reasonable exploration of the extremes however

Applying ATLAS constraints

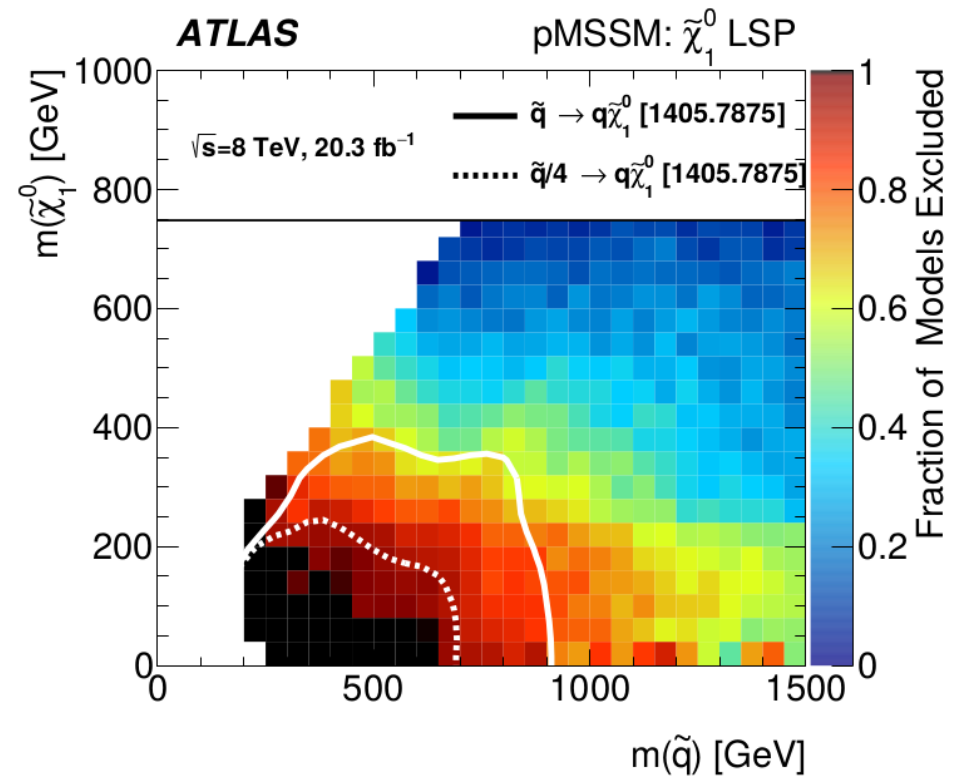
- 500, 000, 000 points were chosen at random
 - 310,327 points pass the non-ATLAS constraints
- For each of these points, the following is performed
 - check to see if cross-sections are high enough
 - if yes: truth level (+ basic efficiency) MC to see if acceptance values are high enough
 - separate into: a) obviously excluded, b) obviously not excluded, c) close to exclusion
 - run proper ATLAS simulation on category (c) to clarify status
- Results are presented in mass planes, with fraction of excluded models
 - the paper is honest: this tells you **nothing** in general about exclusions
 - can give a sort of qualitative, hand-waving insight however
 - BONUS: get to check how accurate simplified model limits are

Glauino-LSP and squark-LSP planes

monojet search is good here



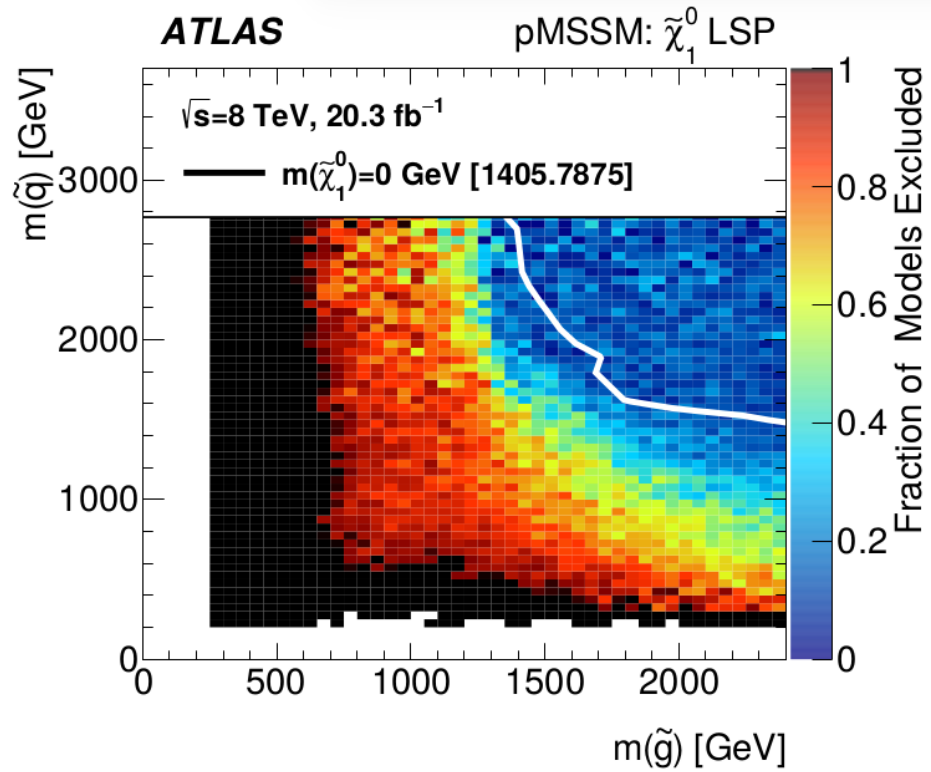
(a) Gluino / LSP



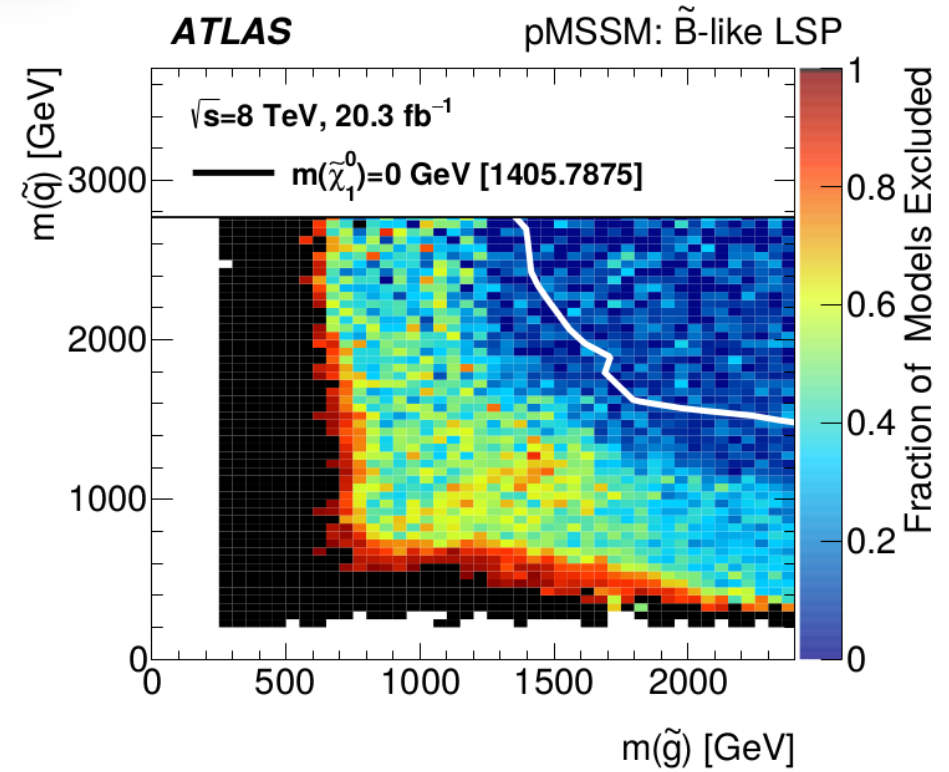
(b) Lightest (1st/2nd gen) squark / LSP

simplified gluino model

Glino-squark planes

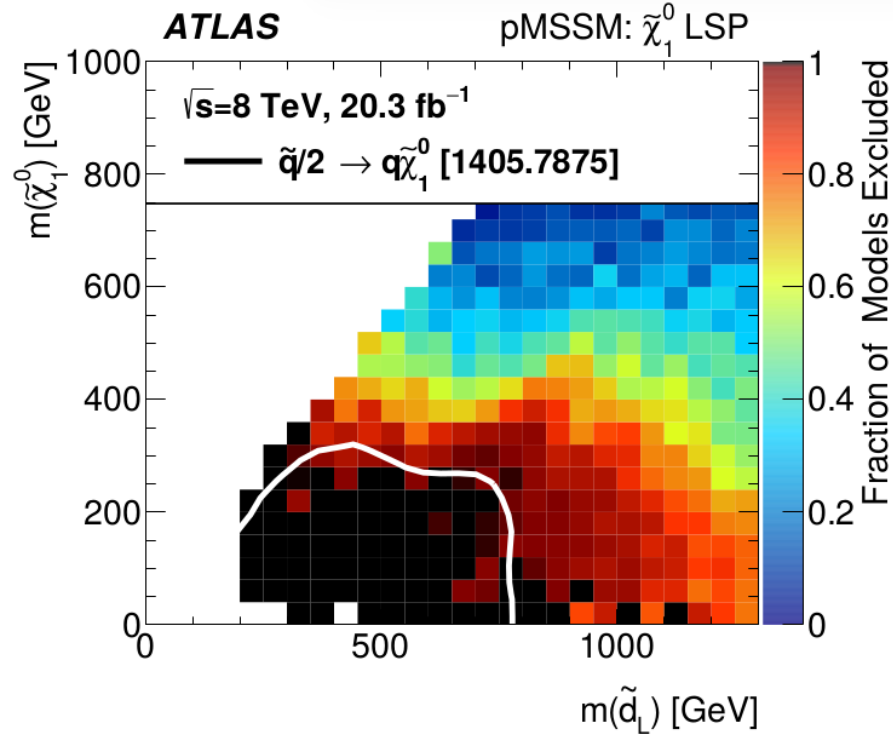


(a) All LSP types

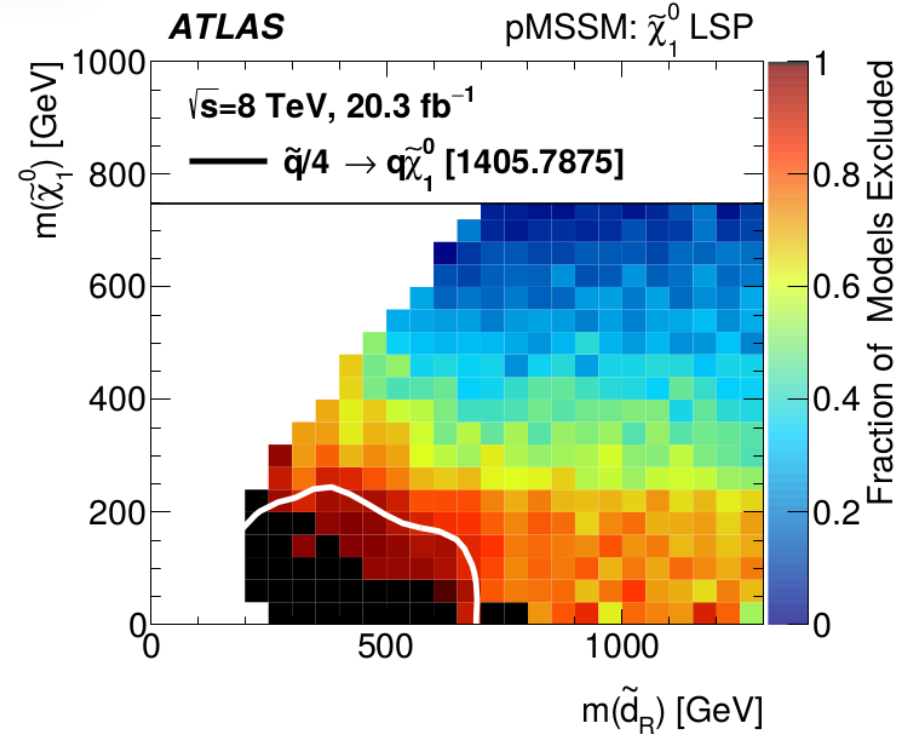


(b) Bino-like LSPs

Limits depend on squark L or R



(c) Left down squark



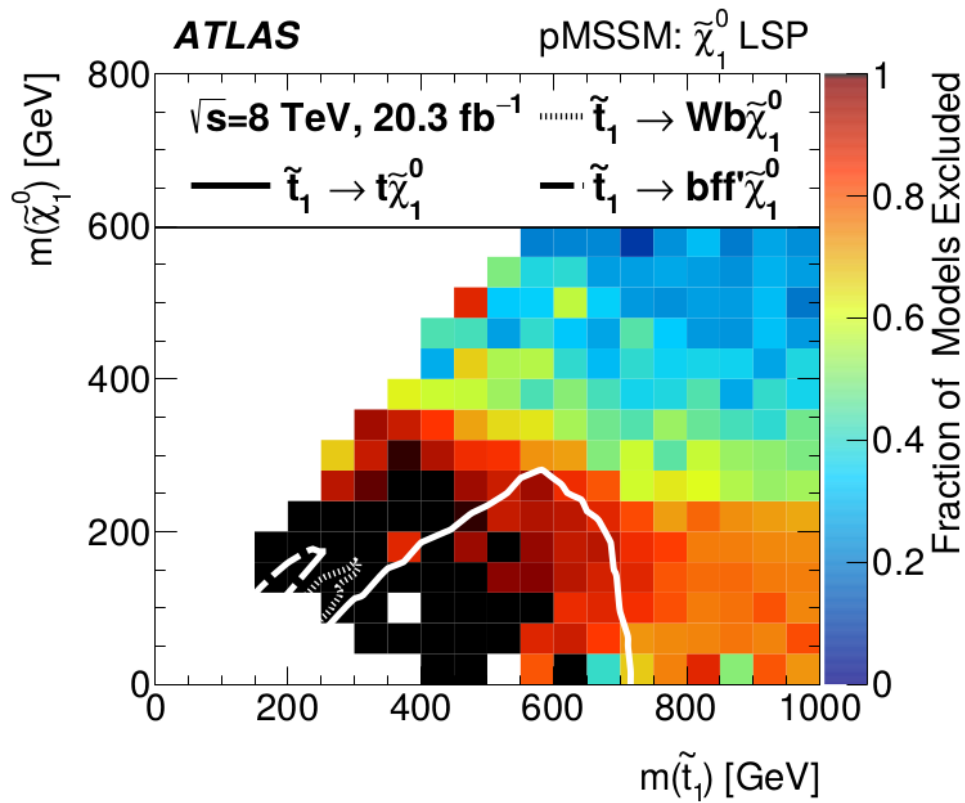
(d) Right down squark

- LH squarks form an EW doublet \rightarrow roughly mass degenerate \rightarrow higher cross-section

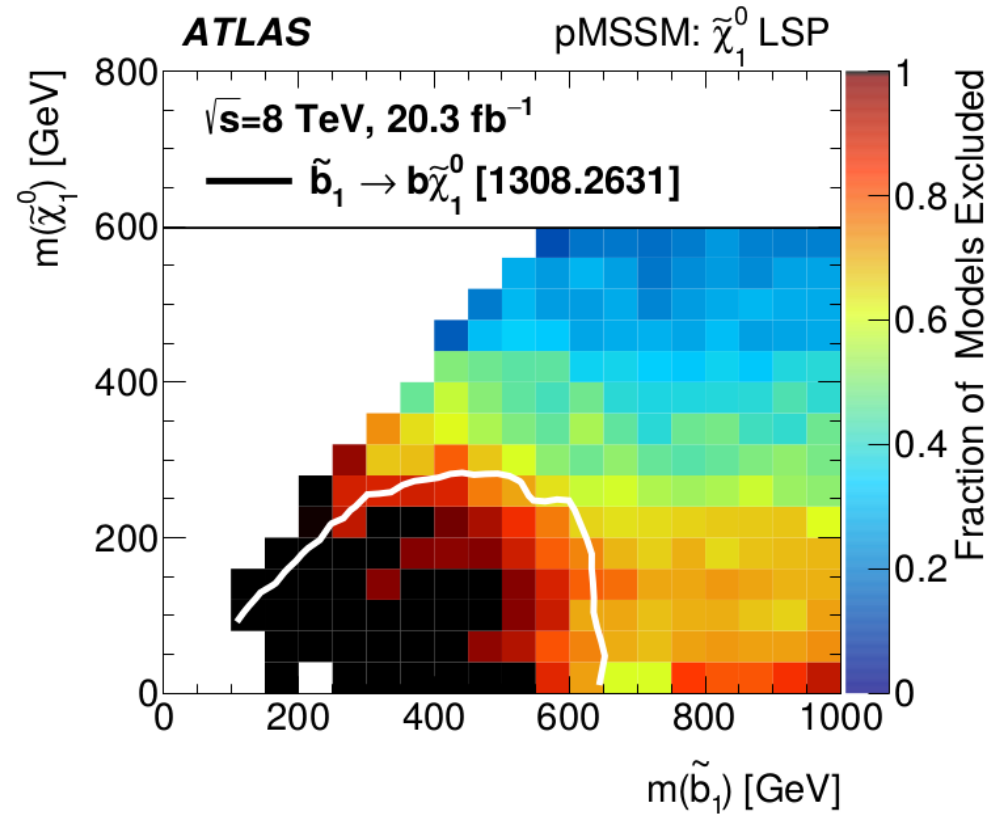
Also, get more cascade decays for RH squarks

- can more easily fall in the cracks between searches

3rd generation squarks



(a) All pMSSM points, all searches

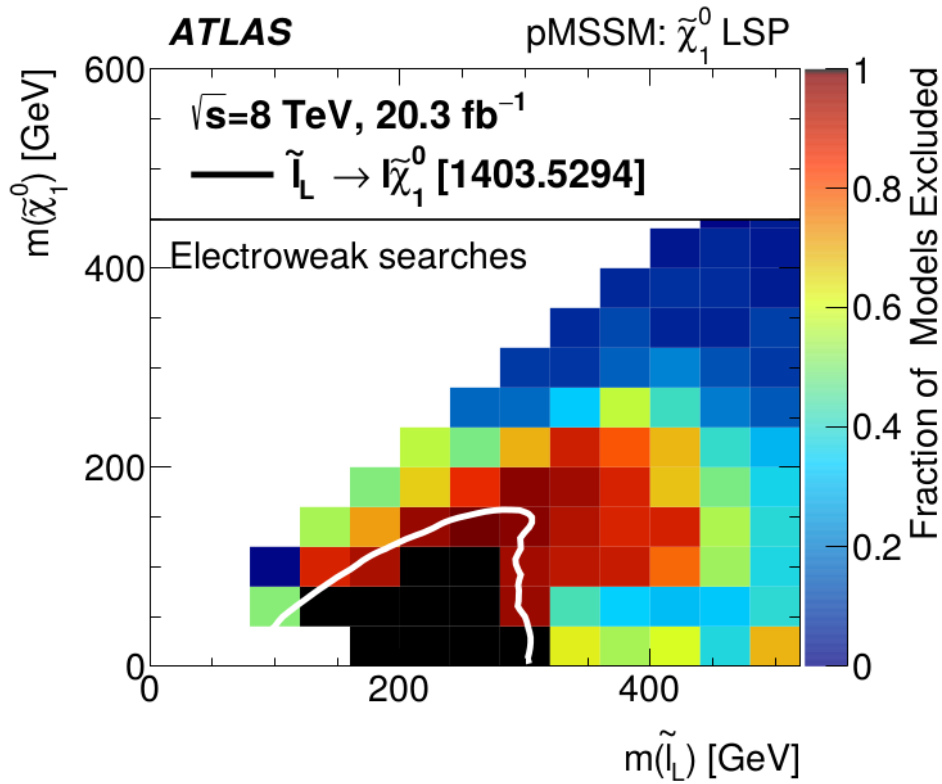


(a) All searches

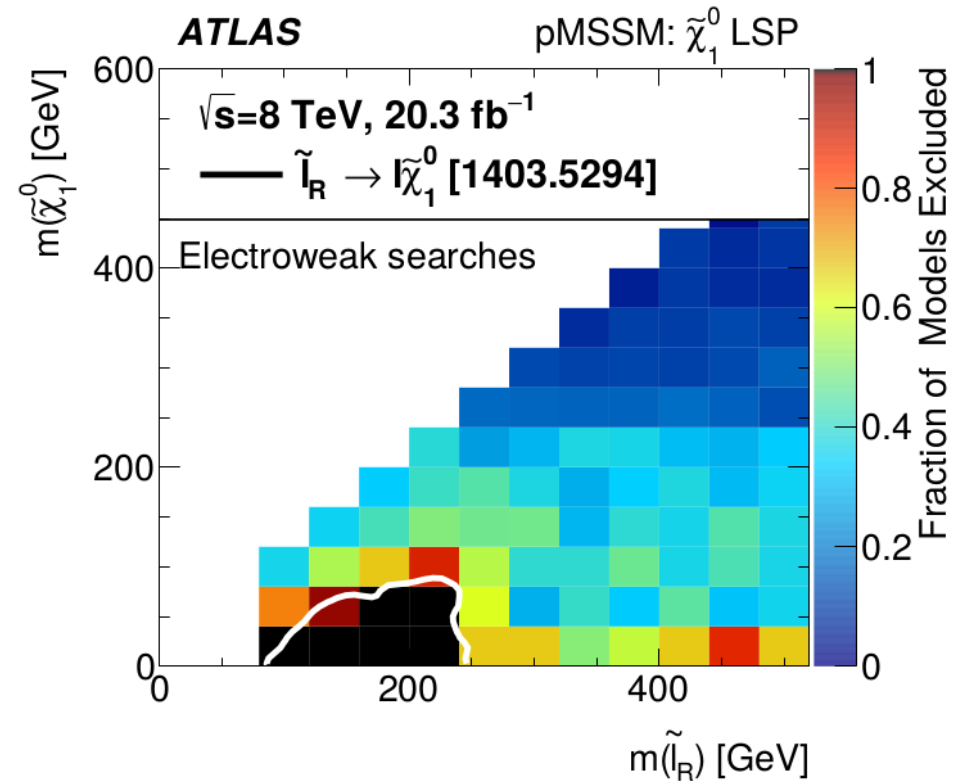
- Note: scan has few points with low stop mass

- due to the Higgs mass constraint (see lecture 1)

Sleptons

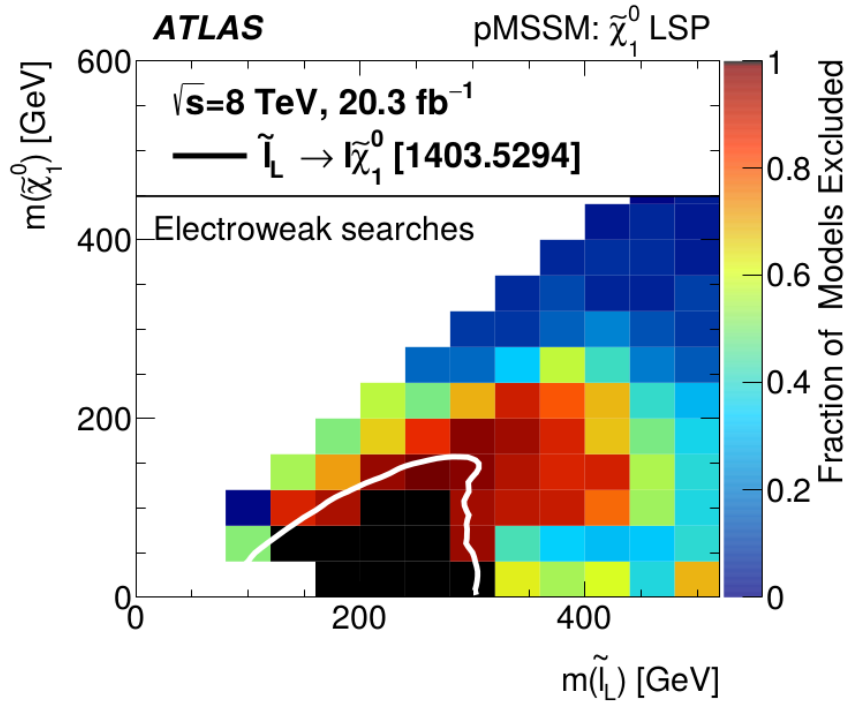


(a) All LSPs, $\tilde{\ell}_L$

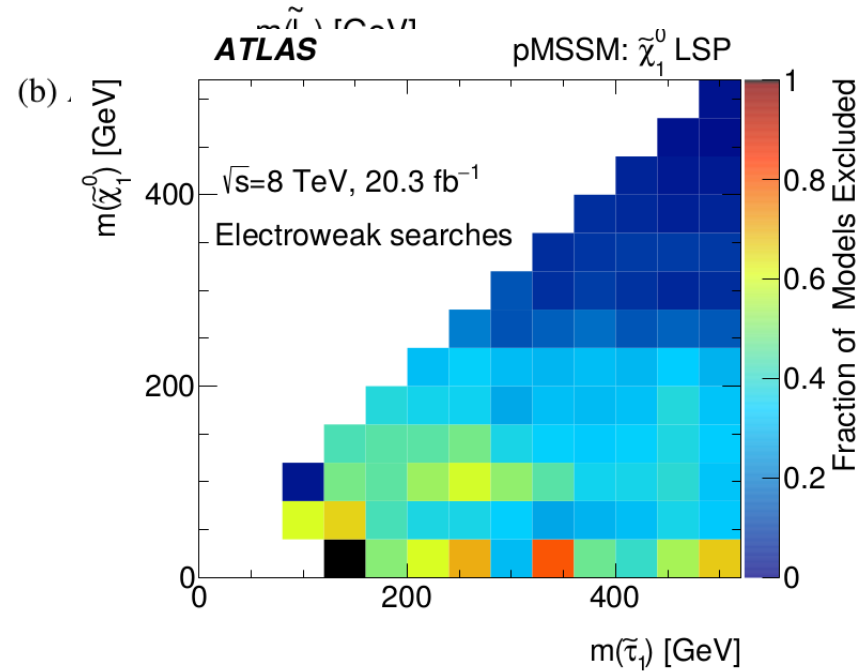
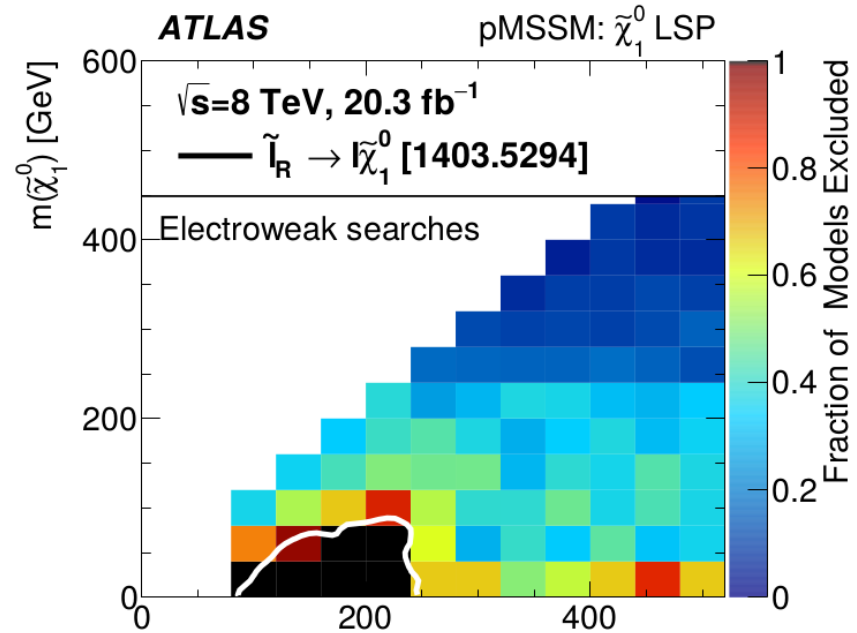


(b) All LSPs, $\tilde{\ell}_R$

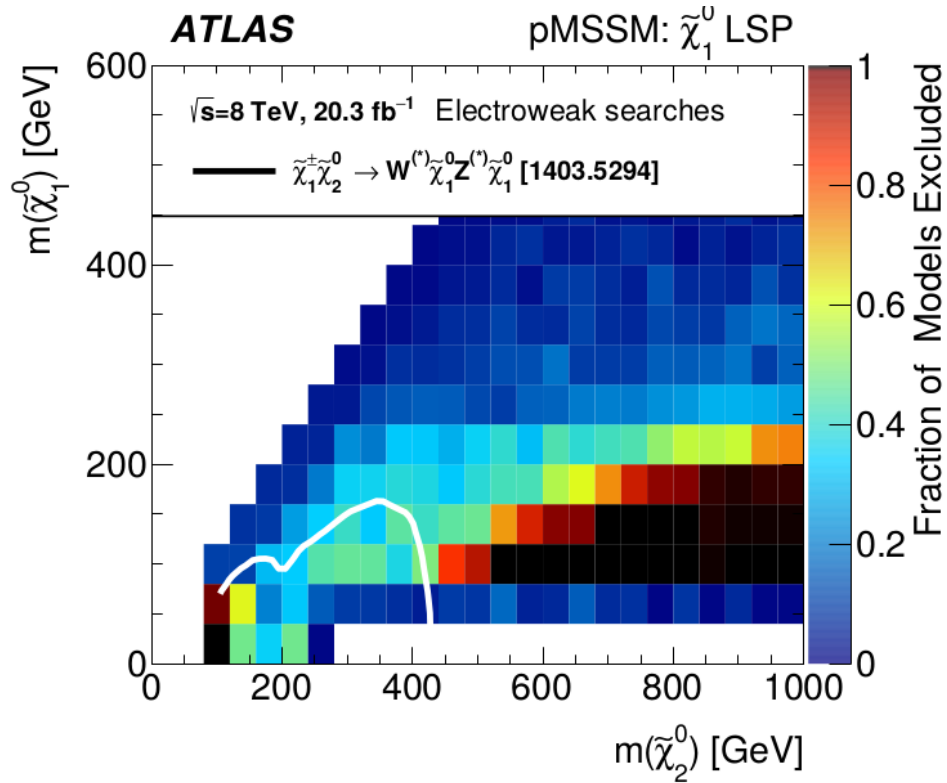
Sleptons



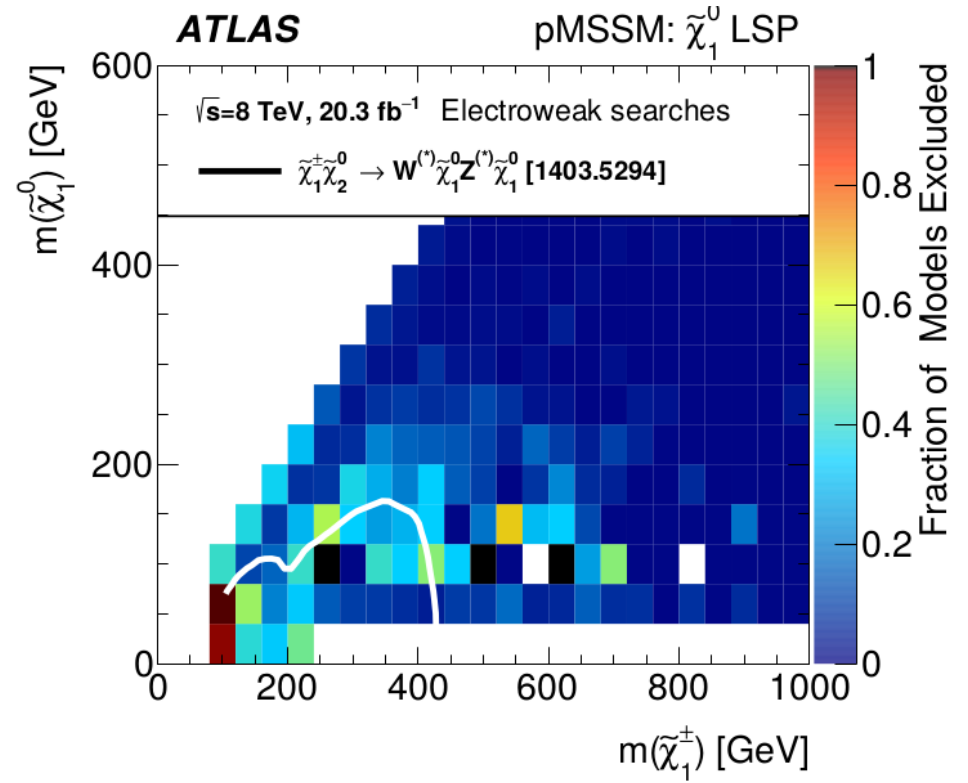
(a) All LSPs, \tilde{l}_L



Weak gauginos



(a) Neutralinos



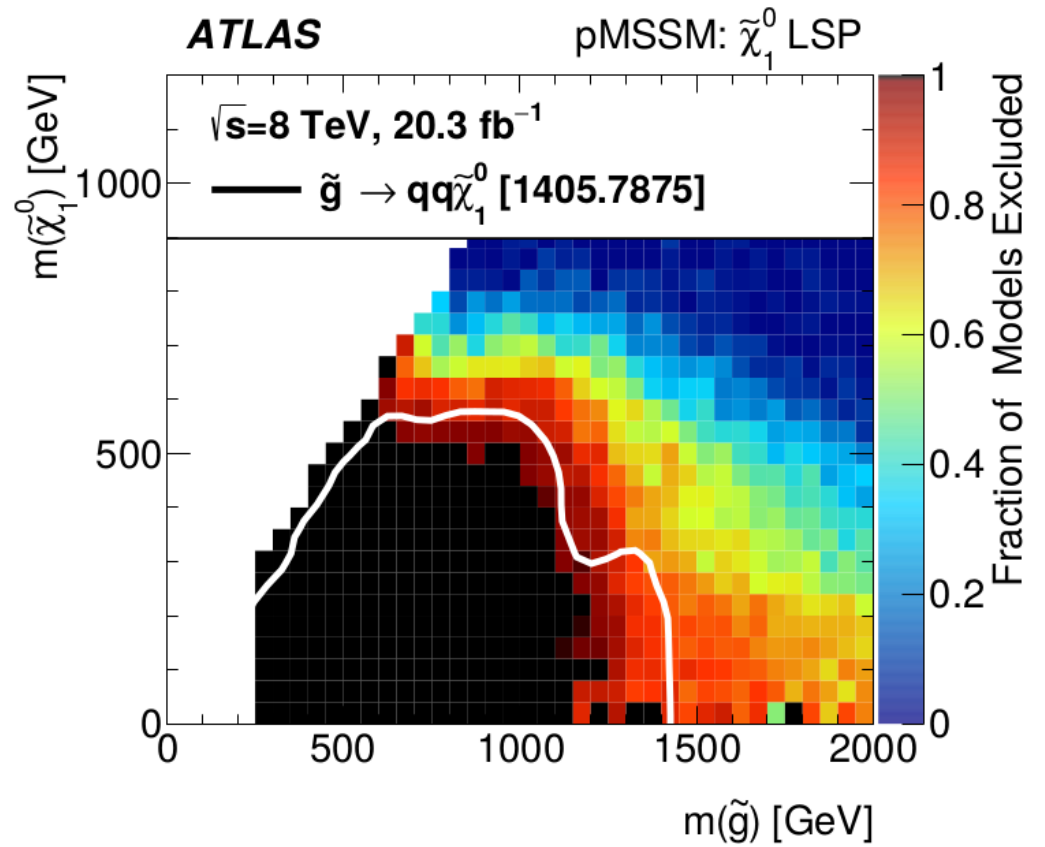
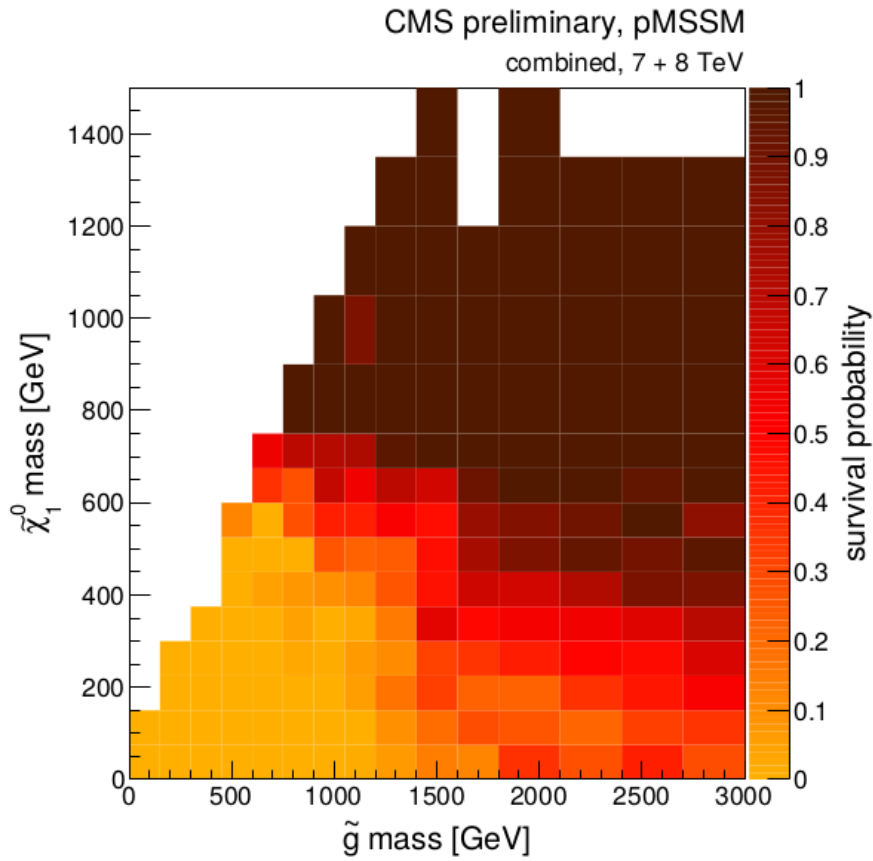
(b) Chargino–neutralino

- Different methodology

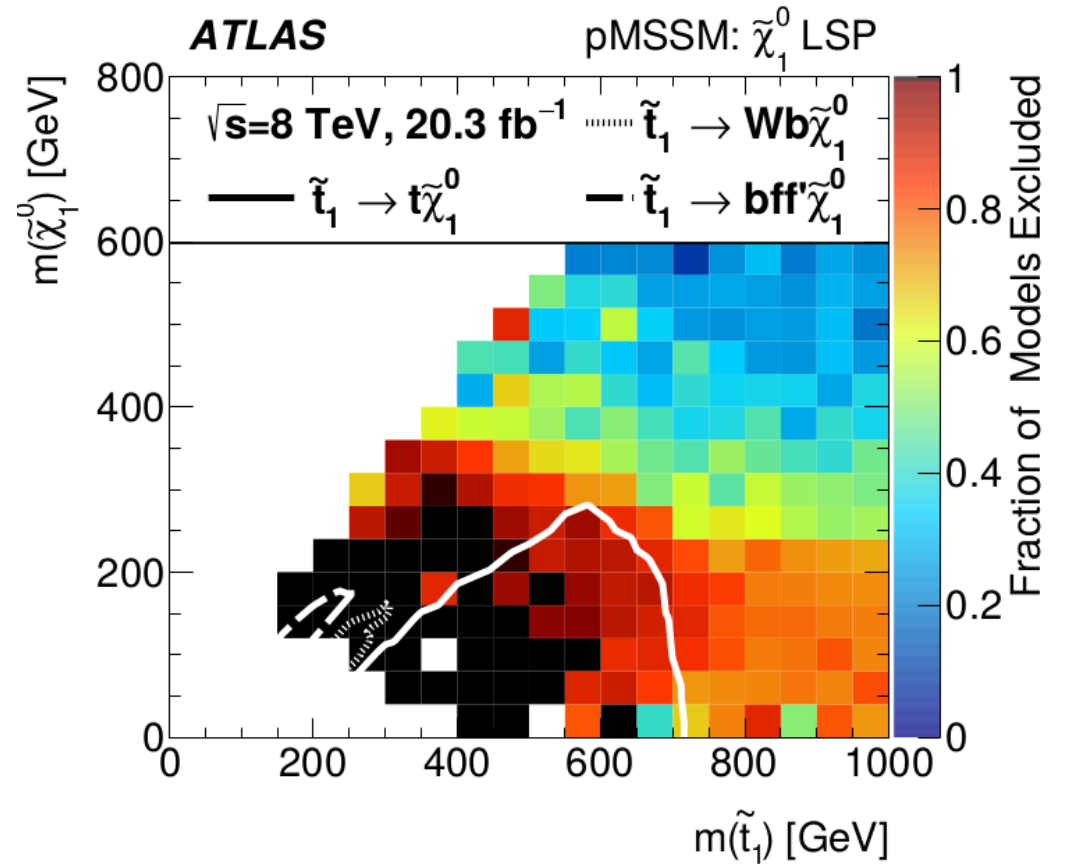
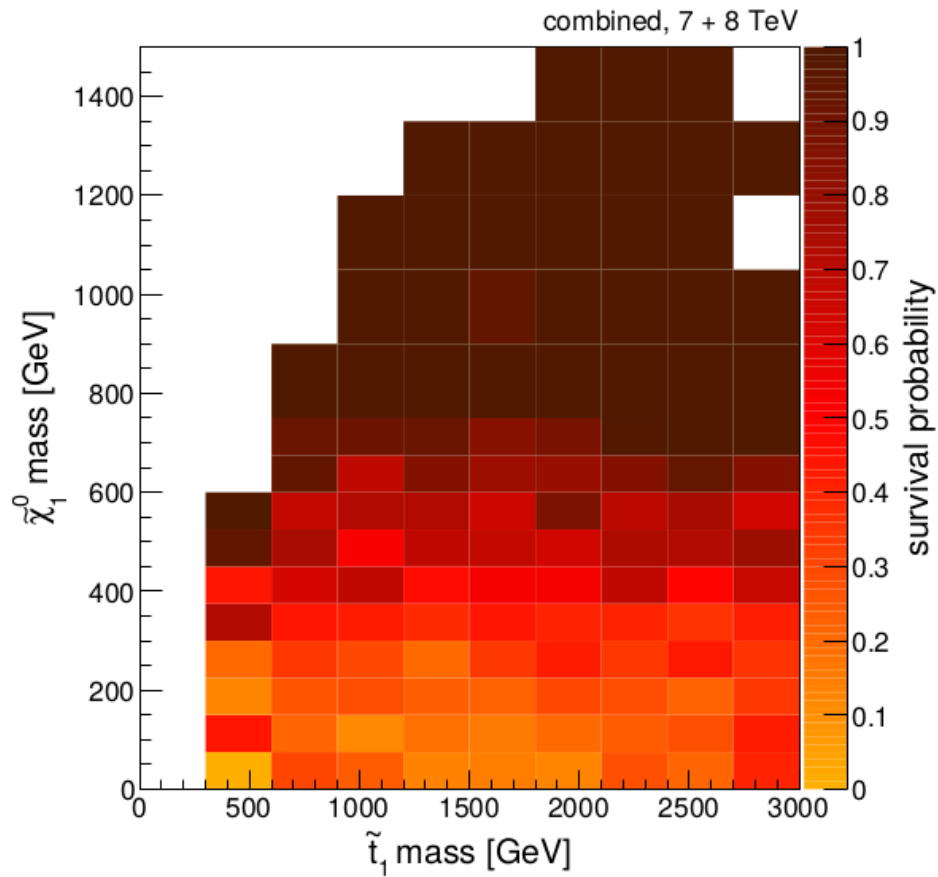
- sampled points come from an MCMC Bayesian scan of the pMSSM
- but only 7200 points are selected for simulation studies!
- 10,000 MC events per point

i	Observable $\mu_i(\theta)$	Constraint $D_i^{\text{non-DCS}}$	Likelihood function $L(D_i^{\text{non-DCS}} \mu_i(\theta))$	comment
1	$BR(b \rightarrow s\gamma)$ [40]	$(3.43 \pm 0.21^{\text{stat}} \pm 0.24^{\text{th}} \pm 0.07^{\text{sys}}) \times 10^{-4}$	Gaussian	reweight
2	$BR(B_s \rightarrow \mu\mu)$ [41]	$(2.9 \pm 0.7 \pm 0.29^{\text{th}}) \times 10^{-9}$	Gaussian	reweight
3	$R(B_u \rightarrow \tau\nu)$ [40]	1.04 ± 0.34	Gaussian	reweight
4	Δa_μ [42]	$(26.1 \pm 6.3^{\text{exp}} \pm 4.9^{\text{SM}} \pm 10.0^{\text{SUSY}}) \times 10^{-10}$	Gaussian	
5	m_t [43]	$173.20 \pm 0.87^{\text{stat}} \pm 1.3^{\text{sys}}$ GeV	Gaussian	reweight
6	$m_b(m_b)$ [44]	$4.19_{-0.06}^{+0.18}$ GeV	Two-sided Gaussian	
7	$\alpha_s(M_Z)$ [44]	0.1184 ± 0.0007	Gaussian	
8	m_h	LHC: $m_h^{\text{low}} = 120, m_h^{\text{up}} = 130$	1 if $m_h^{\text{low}} \leq m_h \leq m_h^{\text{up}}$ 0 if $m_h < m_h^{\text{low}}$ or $m_h > m_h^{\text{up}}$	reweight
9	μ_h	CMS and ATLAS in LHC RunI, Tevatron	Lilith1.01 [36, 37]	post-MCMC
10	sparticle masses	LEP [45] (via micrOMEGAs [29–31])	1 if allowed 0 if excluded	

CMS v ATLAS: gluino mass



CMS v ATLAS: stop mass



How *should* we decide if SUSY is ruled out

- Both of the previous studies had statistical flaws:
 - ATLAS: ignore fitting entirely, density of points has no meaning
 - CMS: too few points simulated to ensure conclusions are meaningful (in both cases this was necessary to reduce computational complexity)
- In general, these studies need a more rigorous approach
- For simple models with only two parameters and a few constraints:
 - determining exclusion regions by overlapping limits is fine
- For lots of constraints, ought to use a **combined likelihood**:
 - can trade off constraints and still get a good fit to the data
- If there are lots of constraints *and* lots of parameters:
 - things are much harder

What if there are many parameters?

- Much harder in principle
- Need to:
 - scan the space (need very smart methods for a large number of parameters)
 - interpret the results (Bayesian/frequentist)
 - project down to parameters of interest (marginalise/profile)

i.e. need a global statistical fit

Reminder: LHC cut and count analyses

- Given s expected signal events, b background events and o observed events:

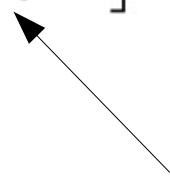
$$L = e^{-(s+b)} (s+b)^o \div o!$$

- Systematics on signal and background can be handled via:

$$\mathcal{L}(n|s, b) = \int_0^\infty \frac{[\xi(s+b)]^n e^{-\xi(b+s)}}{n!} P(\xi) d\xi$$

- With, e.g. , $P(\xi|\sigma_\xi) = \frac{1}{\sqrt{2\pi\sigma_\xi}} \frac{1}{\xi} \exp \left[-\frac{1}{2} \left(\frac{\ln \xi}{\sigma_\xi} \right)^2 \right]$

quadrature sum of signal and background systematics



How to calculate s

$$S = \sigma \times \epsilon \times L$$

Read this from the title of the paper

Need to calculate SUSY production cross-sections. Best tool is probably Prospino (general NLO cross-sections). Fast NLO+NLL cross-sections exist in NLLFAST for some models.

Need to run an MC generator (see Skands lectures) and pass the events through a detector simulation. Also need to apply LHC analysis cuts.

Current tools for applying LHC constraints to SUSY models

- Several tools for the application of collider limits have appeared in recent years:

CHECKMATE: Contains a custom version of the DELPHES event simulation plus analysis code for lots of LHC searches. Does not do any MC event generation or cross-section calculation (but version 2 is coming soon).

S MODELS: Looks at LHC “simplified model” limits and checks to see if a given model is excluded by those limits (by scaling each topology by the appropriate cross-section times branching ratio). Gives very conservative limits and is only applicable to SUSY.

FASTLIM: Uses simplified model limits plus efficiency tables to approximate the efficiency for general SUSY models. Only applicable to SUSY.

- Although some of the approaches taken above may generalise, there is no fully general package capable of quickly applying LHC limits to new models out of the box

Coming soon: SUSY-AI, GAMBIT

Global analysis of the pMSSM in light of the Fermi GeV excess: prospects for the LHC Run-II and astroparticle experiments

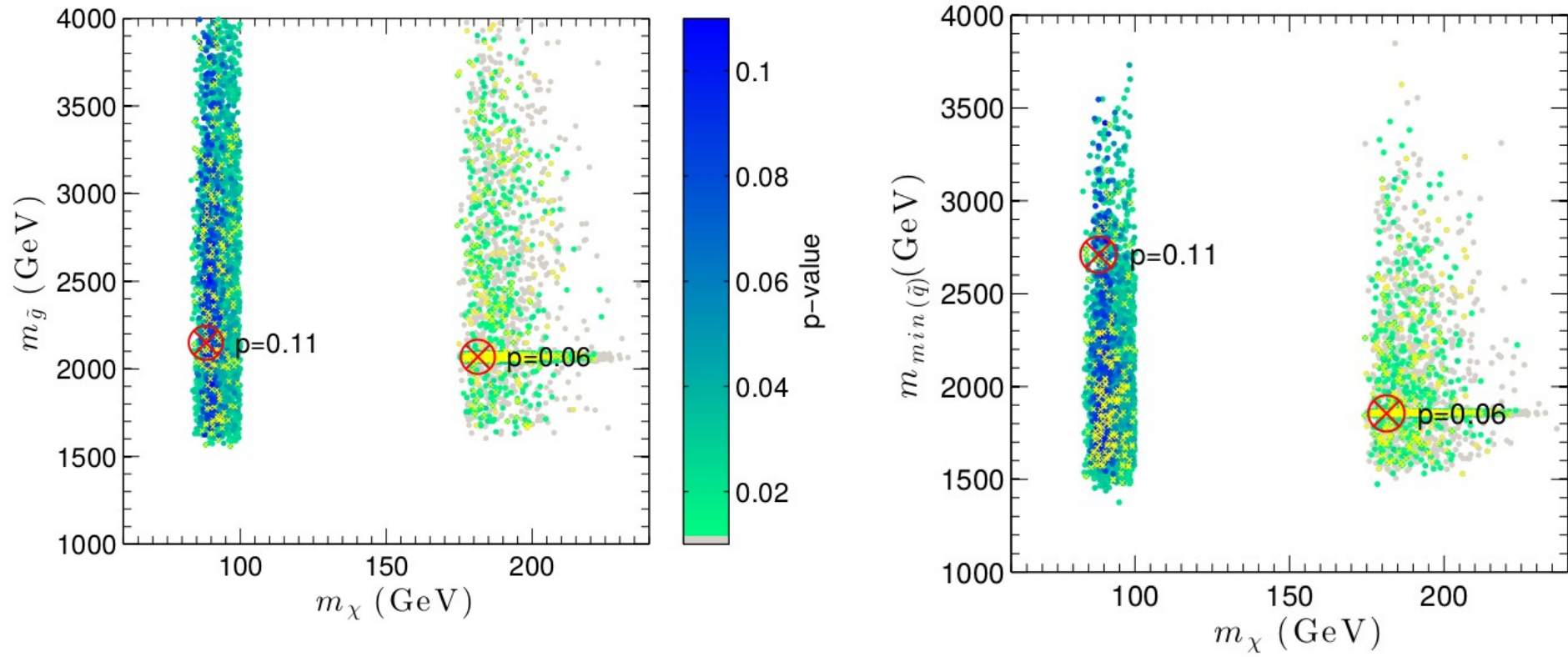
Gianfranco Bertone,^a Francesca Calore,^a Sascha Caron,^b Roberto Ruiz,^c Jong Soo Kim,^d Roberto Trotta,^e Christoph Weniger.^a

- A proper global fit of the pMSSM

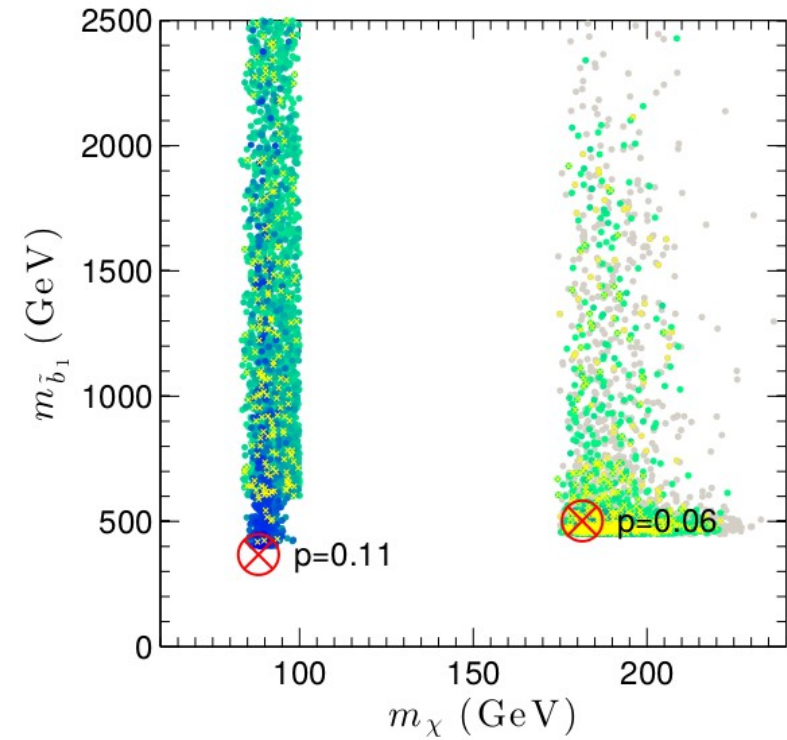
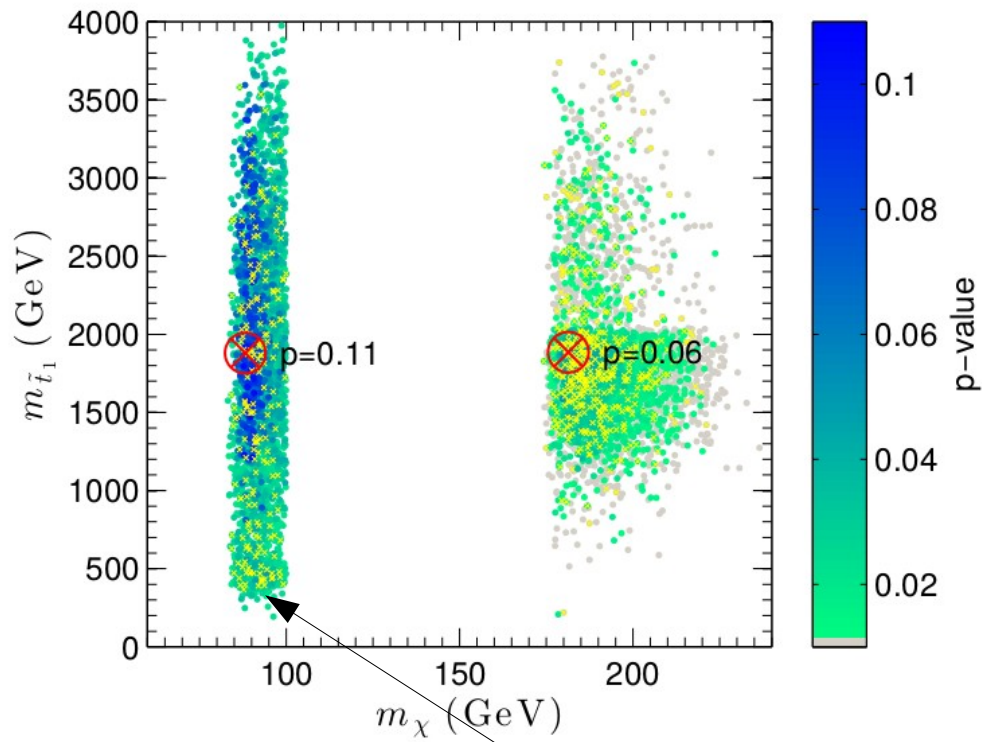
- LHC constraints applied posthoc, but to 100,000 points
- uses Fermi-LAT data, EW precision constraints, flavour observables, relic density, LUX constraints, IceCube constraints, Higgs constraints + LHC sparticle searches (CheckMATE)

$$\begin{aligned} \ln \mathcal{L}_{\text{Joint}} = & \ln \mathcal{L}_{\text{GCE}} + \ln \mathcal{L}_{\text{EW}} + \ln \mathcal{L}_{\text{B(D)}} + \ln \mathcal{L}_{\Omega_{\chi} h^2} \\ & + \ln \mathcal{L}_{\text{LUX}} + \ln \mathcal{L}_{\text{IC}} + \ln \mathcal{L}_{\text{Higgs}} + \ln \mathcal{L}_{\text{SUSY}}, \end{aligned}$$

Glino and minimum squark mass



Stop and sbottom masses



Light stop points not excluded
by LHC searches

GAMBIT: The Global And Modular BSM Inference Tool

- Fast definition of new datasets and theoretical models
- Plug and play scanning, physics and likelihood packages
- Extensive model database – not just SUSY
- Extensive observable/data libraries
- Many statistical and scanning options (Bayesian & frequentist)
- *Fast* LHC likelihood calculator
- Massively parallel
- Fully open-source

ATLAS

LHCb

Belle-II

Fermi-LAT

CTA

HESS

IceCube

XENON/DARWIN

Theory

A. Buckley, P. Jackson, C. Rogan, M. White,

M. Chrzęszcz, N. Serra

F. Bernlochner, P. Jackson

J. Conrad, J. Edsjö, G. Martinez, P. Scott

C. Balázs, T. Bringmann, J. Conrad, M. White

J. Conrad

J. Edsjö, P. Scott

J. Conrad, R. Trotta

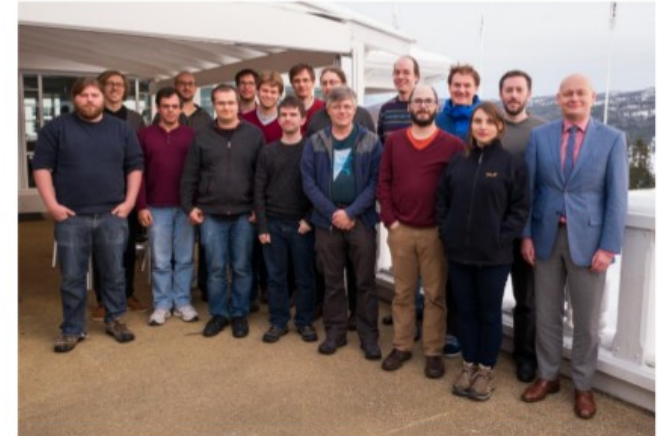
P. Athron, C. Balázs, T. Bringmann,

J. Cornell, J. Edsjö, B. Farmer, T. Gonzalo, S. Hoof,

F. Kahlhoefer, A. Krislock, A. Kvellestad, M. Pato,

F. Mahmoudi, J. McKay, A. Raklev, R. Ruiz, P. Scott,

R. Trotta, C. Weniger, M. White



27 Members, 9 Experiments, 4 major theory codes, 10 countries

Summary

- This lecture, we have learnt that SUSY limits are complicated in general
 - LHC gluino limit is relatively stringent
 - much evidence suggests that light stops are not ruled out
 - light squark mass limits are not watertight either
- SUSY can apparently be perfectly healthy, except:
 - Higgs mass makes things interesting
 - naturalness is complicated
 - SUSY is **fine** if you don't care at all about fine tuning
- Lots still to do
 - this year is an extremely important year
 - could get as much as 30 fb^{-1} of LHC data at 13 TeV

Pay close attention to the new results presented at this conference!