

# The Spectra of Type IIB Flux Compactifications at Large Complex Structure

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# Introduction

- ▶ Type IIB string theory is 10d
  - ⇒ (Usually) say 6d are 'compactified'
  - Compactification has parameters ('moduli')
- ▶ Low energy limit is supergravity
  - ⇒ Compactification affects content of supergravity
  - ⇒ Different compactifications  $\leftrightarrow$  different low energy theories
- ▶ Flux associates energy with parameters of compactification
  - ⇒ Important for constructing metastable vacua
  - ⇒ Flux compactification useful to get to particle physics
  - (But: we *will not* be talking about moduli stabilisation)

# Motivation: Flux landscape, computational difficulty

Many possible flux compactification choices ('flux landscape')

( $\Rightarrow$  Many low-energy theories)

And too difficult to compute in generic compactifications

Computations possible only when there are few moduli

$\Rightarrow$  Large literature on statistical approach, e.g. random matrix theory (RMT)

But: difficult to verify applicability of RMT techniques, as used precisely where explicit compactifications not available

Here we present a way to test these statistical methods

Will focus on statistical models of two important matrices:  
the Hessian  $\mathcal{H}$  and the matrix  $\mathcal{M}$

# Type IIB flux compactifications, I

Low-energy degrees of freedom include axio-dilaton  $\tau$  and complex structure moduli  $u^i$  ('large complex structure'  $\leftrightarrow u^i \gtrsim 1$ )

Integrally quantised flux wraps cycles of the compactification  
 $\Rightarrow$  Gives superpotential  $W$  (Gukov-Vafa-Witten),

$$W = \vec{N} \cdot \vec{\Pi} = (\vec{f} - \tau \vec{h}) \cdot (1, u^i, 2F - u^j F_j, F_i)$$

$\vec{N}$  is the 'flux vector',  $\vec{\Pi}$  is the 'period vector',  $F$  is cubic in  $u^i$

Kähler potential (in LCS expansion):

$$K = -\ln \left( \frac{i}{6} \chi_{ijk} (u^i - \bar{u}^i)(u^j - \bar{u}^j)(u^k - \bar{u}^k) - 2\text{Im}\chi_0 \right) - \ln(-i(\tau - \bar{\tau}))$$

F-term scalar potential:

$$V = e^K \left( K^{a\bar{b}} D_a W \bar{D}_{\bar{b}} \bar{W} - 3|W|^2 \right),$$

$K$ : Kähler potential,  $K_{a\bar{b}}$ : Kähler metric,  $D_a$ : Kähler covariant derivative ( $D_a W = \partial_a W + \partial_a K W$ ),  $a, b \in \{\tau, u^i\}$

# Type IIB flux compactifications, II

Hessian matrix  $\mathcal{H}$ ,

$$\mathcal{H} = \begin{pmatrix} \partial_a \bar{\partial}_{\bar{b}} V & \partial_a \partial_b V \\ \bar{\partial}_{\bar{a}} \bar{\partial}_{\bar{b}} V & \bar{\partial}_{\bar{a}} \partial_b V \end{pmatrix}$$

$\mathcal{H}$  important: must be positive definite for metastable vacuum

Matrix  $\mathcal{M}$ ,

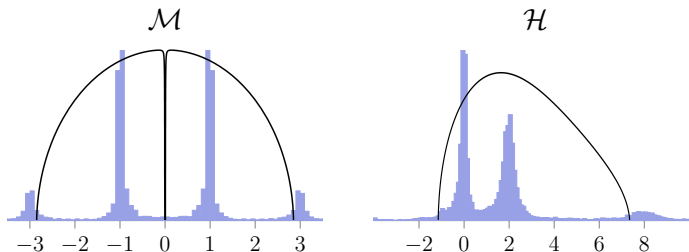
$$\mathcal{M} = \begin{pmatrix} 0 & Z_{ab} e^{-i\vartheta} \\ \bar{Z}_{\bar{a}\bar{b}} e^{i\vartheta} & 0 \end{pmatrix}$$

$\mathcal{M}$  important:  $\partial_a V = 0 \Leftrightarrow$  eig'val equation for  $\mathcal{M}$  (eig'val  $2|W|$ )

$$Z_{ab} \equiv D_a D_b W, \quad \vartheta \equiv \arg(W)$$

RMT spectra for  $\mathcal{H}$  and  $\mathcal{M}$  proposed in literature

# Results: Numeric spectra at large complex structure



**Figure 1:** In blue: Sampled numerical spectra of  $\mathcal{M}$  (on left) and  $\mathcal{H}$  (on right) in explicit compactification, in units of  $|W|$  and  $m_{3/2}^2$ . RMT expectations shown in black. Here  $\tau \sim 5i$  and  $u^i \sim 10i$ .

- ▶ Same peaks seen in many compactifications, so **universal**
- ▶ **Very different** from expectations of random matrix theory
- ▶ Peaks are sharper for larger moduli (not shown)
- ▶ Peaks visible only for canonically normalised fields and in units of  $|W|$  and  $m_{3/2}^2 = e^{K/2} W$

# Results: Analytic spectra at large complex structure, I

Restrict to LCS <sup>1</sup>:  $(u^i)^3$  dominates  $W$ ,  $(u^i - \bar{u}^i)^3$  dominates  $K$   
⇒ Can derive **analytic spectra**:

$$\text{Spectrum}(\mathcal{M}) = \begin{cases} |W| & \text{multiplicity } h_-^{1,2}, \\ 3|W| & \text{multiplicity } 1, \\ -|W| & \text{multiplicity } h_-^{1,2}, \\ -3|W| & \text{multiplicity } 1. \end{cases}$$

$$\text{Spectrum}(\mathcal{H}) = \begin{cases} 0 & \text{multiplicity } h_-^{1,2}, \\ 2m_{3/2}^2 & \text{multiplicity } h_-^{1,2} + 1, \\ 8m_{3/2}^2 & \text{multiplicity } 1. \end{cases}$$

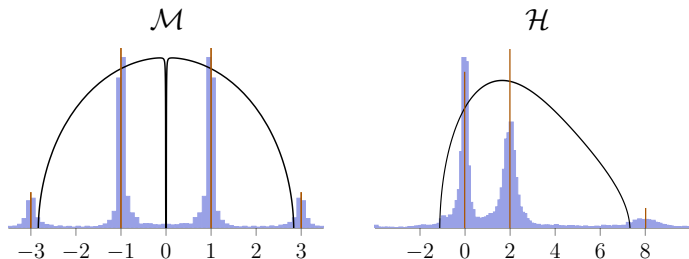
These are **'universal'**: true for any non-vanishing choice of flux, any not-all-vanishing triple intersections  $\chi_{ijk}$ , and any number of complex structure moduli  $u^i$

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<sup>1</sup>See [Marsh, Sousa, JHEP 1603 (2016) 064] for extension to this full subspace



# Results: Analytic spectra at large complex structure, II



Blue: Randomly sampled numerical spectra of  $\mathcal{M}$  and  $\mathcal{H}$  in explicit compactification, in units of  $|W|$  and  $m_{3/2}^2$  respectively

Black: Random matrix theory expectations

Brown: Analytic spectra, with schematic degeneracies

# Things to note

- ▶ Spectra remain **highly peaked at not-so-large complex structure** in explicit compactifications
- ▶ Can show  $F^2 = 4|W|^2 \Rightarrow$  **no supersymmetric points** here
- ▶ No eig'val of  $\mathcal{M}$  at  $2|W| \Rightarrow$  **no critical points** here
- ▶ Slow-roll parameters take universal **large values**:  $\epsilon = 4$ ,  $\eta_{||} = 8$
- ▶ Analytic results can also be **extended to F-theory**:  
see [Marsh, Sousa, JHEP 1603 (2016) 064]
- ▶ Popular continuous flux approximation breaks down quickly at large complex structure (as a result of the dominance of one term in the superpotential)
- ▶ Strong linear correlation between  $|W|$  and the supersymmetric mass at large complex structure

# Summary

- ▶ Many flux compactifications, and difficult to compute  
⇒ Statistical perspective, including RMT
- ▶ In subspace of moduli space, can attempt to verify RMT
- ▶ Explicit compactifications: spectra show peaks, not well-described by RMT
- ▶ At LCS, can find analytic spectra of Hessian matrix  $\mathcal{H}$ , and matrix  $\mathcal{M}$  that governs the critical point equation - find highly degenerate eigenvalues - these results are 'universal'
- ▶ ⇒ Statistical models not as applicable as one might have believed (or hoped)
- ▶ But we can find other generic or universal predictions in some regions of parameter space, via analytic results