Baryogenesis

M.J. Ramsey-Musolf *U Mass Amherst*





http://www.physics.umass.edu/acfi/

SUSY 16, Melbourne July 2016

Cosmic Baryon Asymmetry

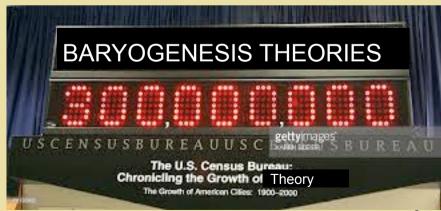
$$Y_B = \frac{n_B}{s} = (8.82 \pm 0.23) \times 10^{-11}$$

One number → BSM Physics

Cosmic Baryon Asymmetry

$$Y_B = \frac{n_B}{s} = (8.82 \pm 0.23) \times 10^{-11}$$

One number → MMM M... Explanations



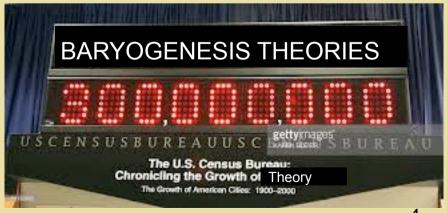
Cosmic Baryon Asymmetry

$$Y_B = \frac{n_B}{s} = (8.82 \pm 0.23) \times 10^{-11}$$

One number → MMM M... Explanations

Experiment can help:

- Discover ingredients
- Falsify candidates



Ingredients for Baryogenesis



- B violation
- C & CP violation
- Out-of-equilibrium or CPT violation

Ingredients for Baryogenesis



Standard Model	BSM
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- B violation (sphalerons)
- C & CP violation
- Out-of-equilibrium or CPT violation

- /
- ×
 - ×

Ingredients for Baryogenesis



Scenarios: leptogenesis, EW baryogenesis, Afflek-Dine, asymmetric DM, cold baryogenesis, postsphaleron baryogenesis...

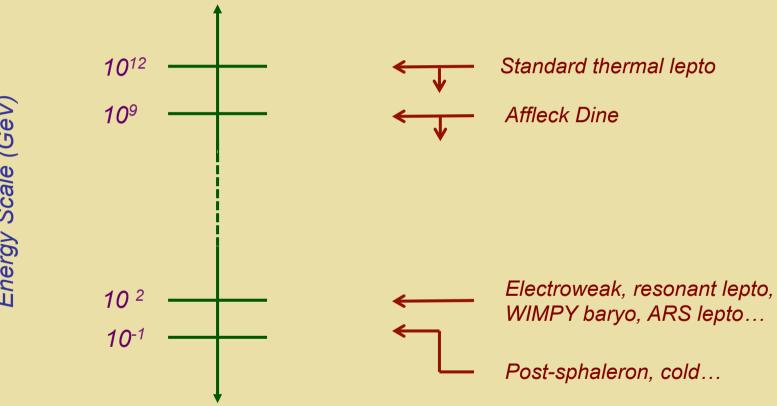
BSM

Standard Model

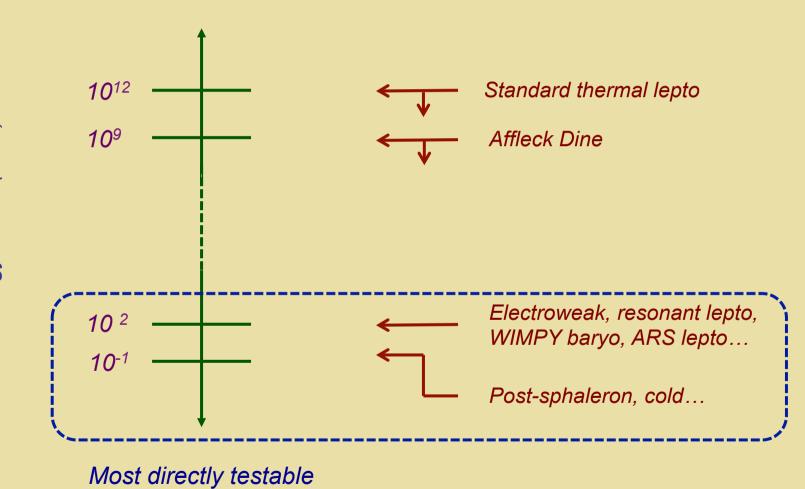
B violation (sphalerons)	✓	✓
C & CP violation	*	V
 Out-of-equilibrium or CPT violation 	*	✓

Energy Scale (GeV)

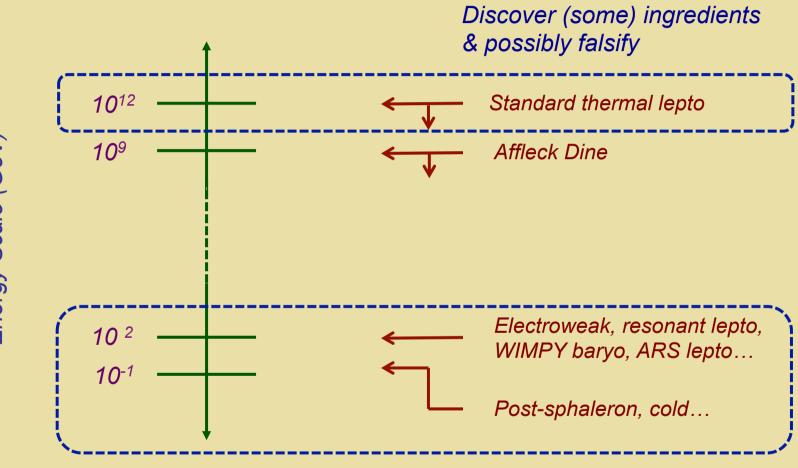
Baryogenesis Scenarios



Baryogenesis Scenarios

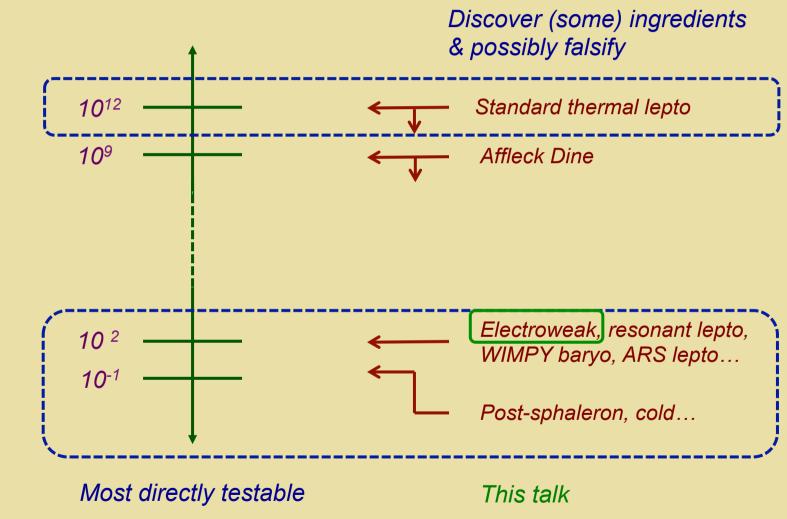


Baryogenesis Scenarios



Most directly testable

Baryogenesis Scenarios



Electroweak Baryogenesis

Was Y_B generated in conjunction with electroweak symmetry-breaking?

Outline

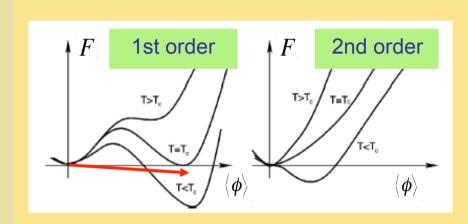
- I. Electroweak Baryogenesis in a Nutshell
- II. Electroweak Phase Transition
- III. CPV: the Baryon Asymmetry & EDMs
- IV. Outlook

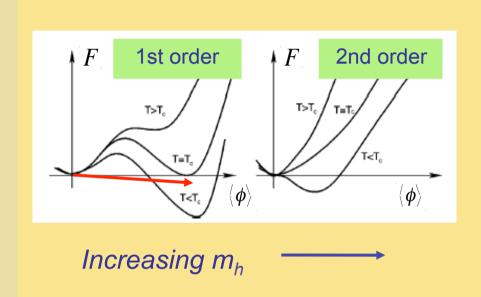
Outline

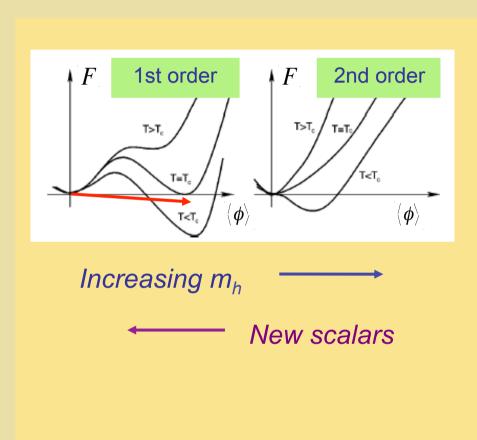
- I. Electroweak Baryogenesis in a Nutshell
- II. Electroweak Phase Transition Time permitting
- III. CPV: the Baryon Asymmetry & EDMs
- IV. Outlook

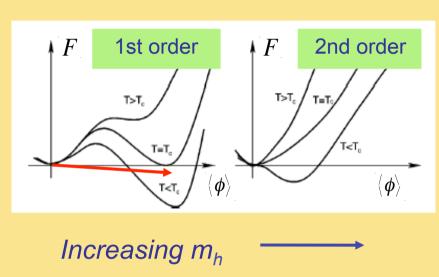


I. EWB in a Nutshell



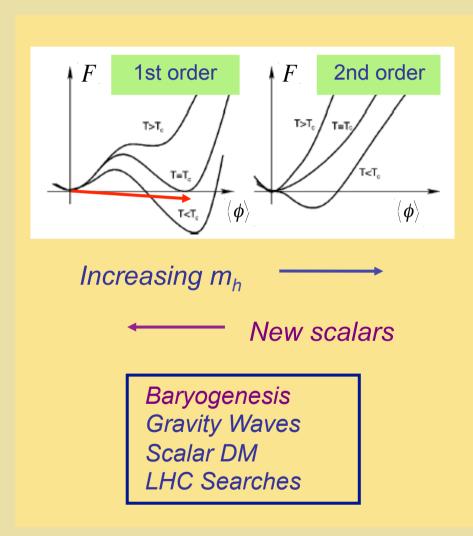




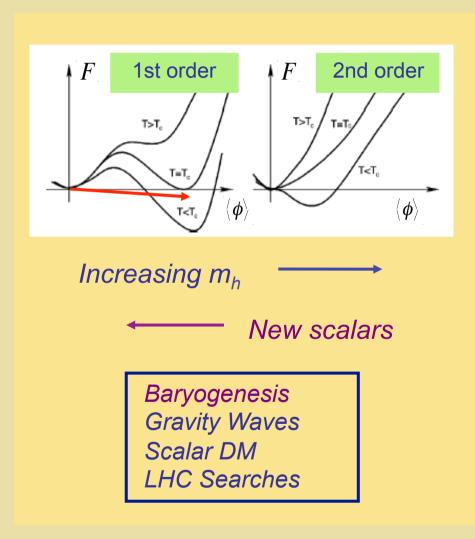


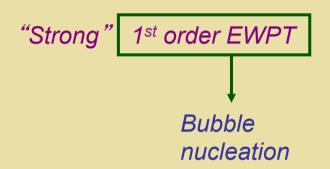
← New scalars

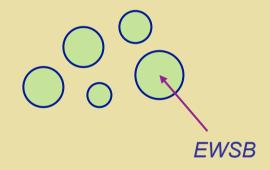
- Loop effects
- Tree-level barrier

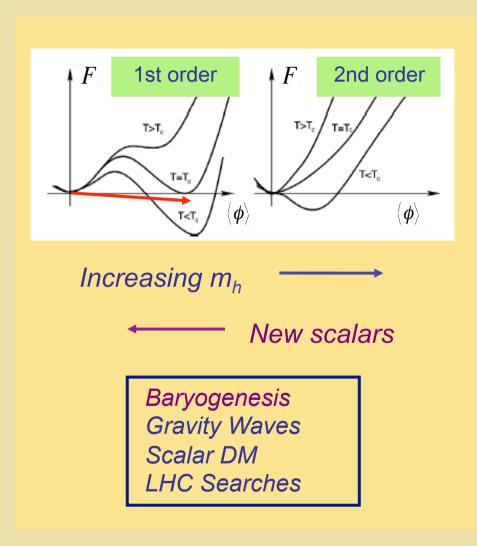


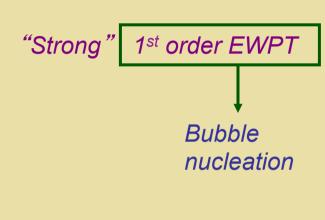
"Strong" 1st order EWPT

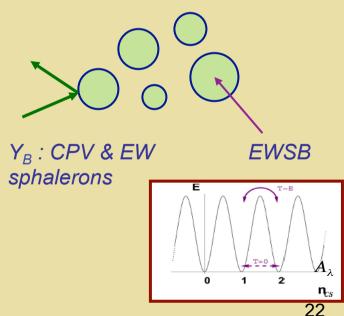


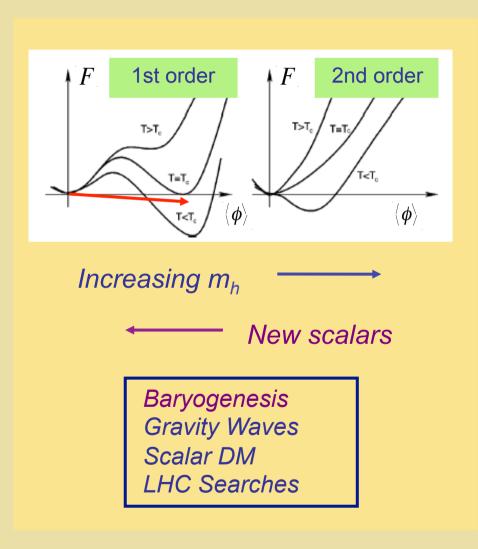


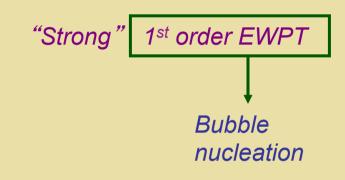


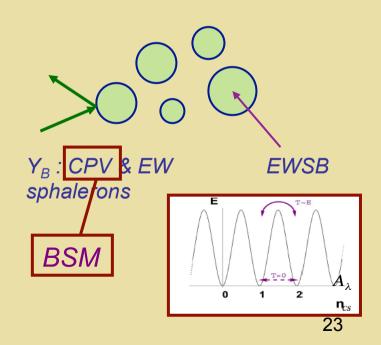


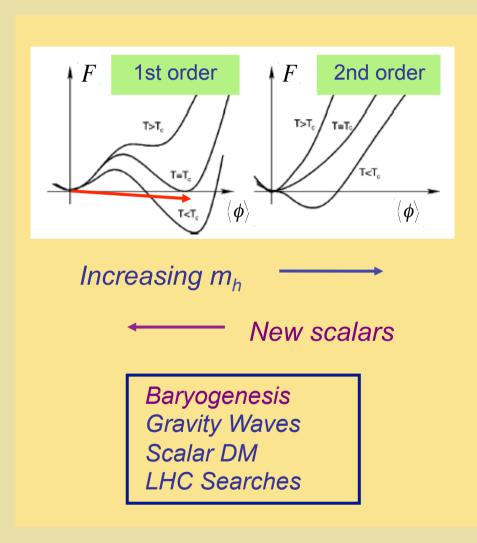


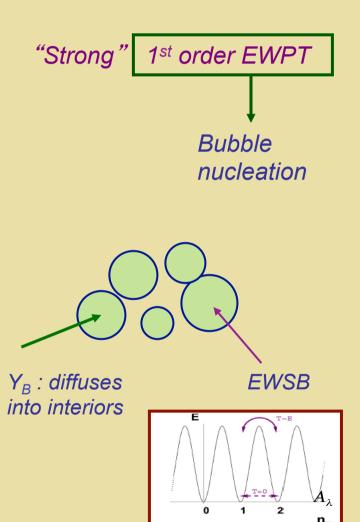


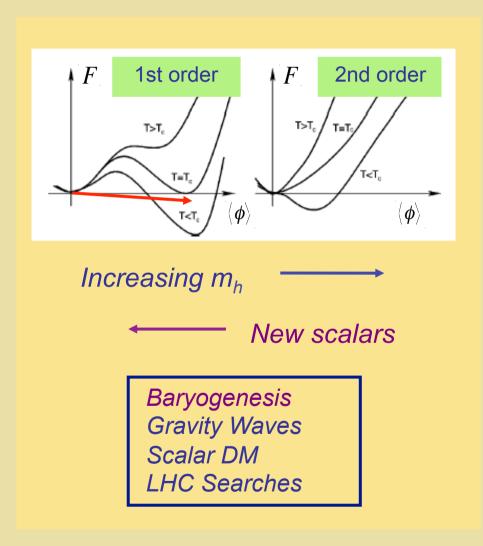


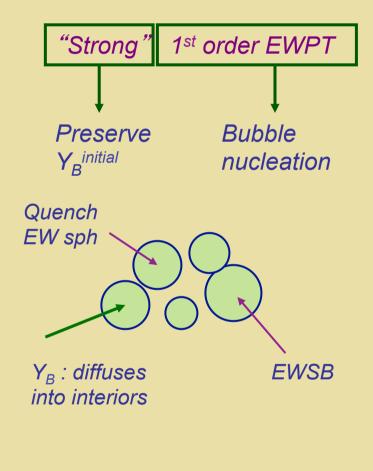








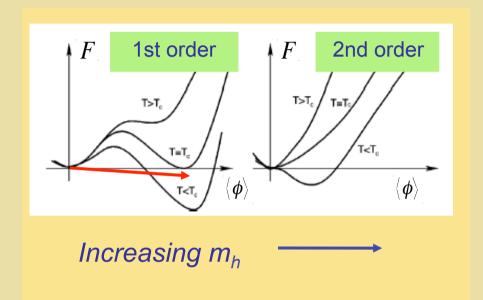


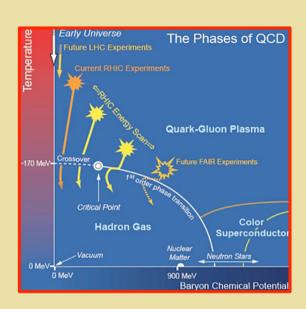


II. Electroweak Phase Transition

Conditions for Electroweak Baryogenesis?

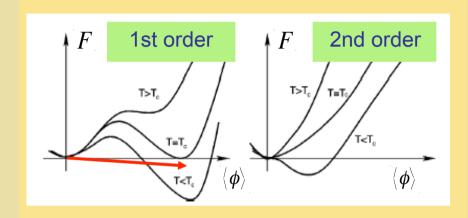
EW Phase Transition: St'd Model





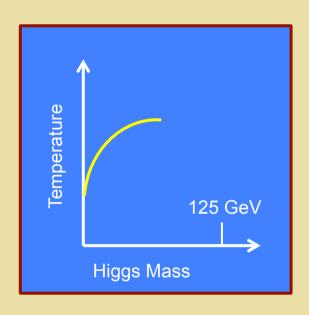
QCD Phase Diagram

EW Phase Transition: St'd Model



Increasing m_h

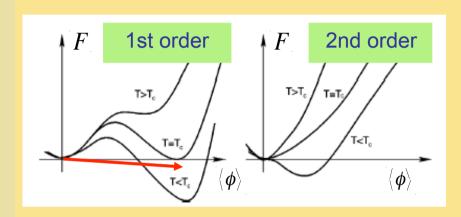
Authors	$M_{\rm h}^C$ (GeV)
[76]	80±7
[74]	72.4 ± 1.7
[72]	72.3 ± 0.7
[70]	72.4 ± 0.9
	[76] [74] [72]



EW Phase Diagram

SM EW: Cross over transition

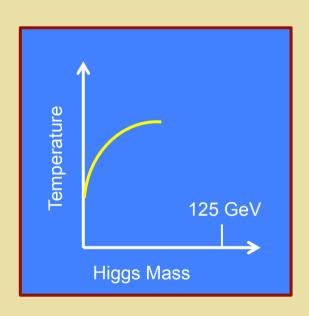
EW Phase Transition: St'd Model



Increasing m_h

Lattice	Authors	$M_{\rm h}^C ({\rm GeV})$
4D Isotropic	[76]	80±7
4D Anisotropic	[74]	72.4 ± 1.7
3D Isotropic	[72]	72.3 ± 0.7
3D Isotropic	[70]	72.4 ± 0.9

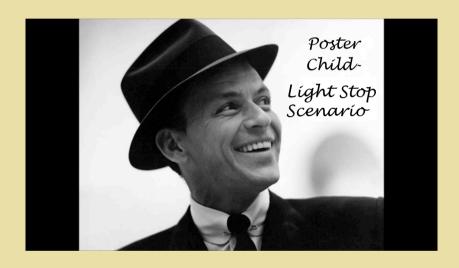
SM EW: Cross over transition



EW Phase Diagram

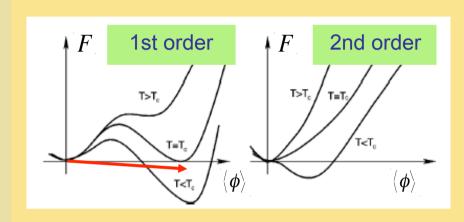
How does this picture change in presence of new TeV scale physics? What is the phase diagram?

EWPT "Poster Child": MSSM Light Stop Scenario



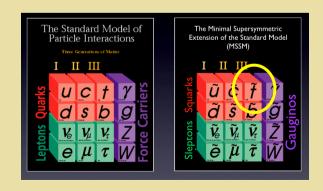
Thermal loops

EW Phase Transition: SUSY



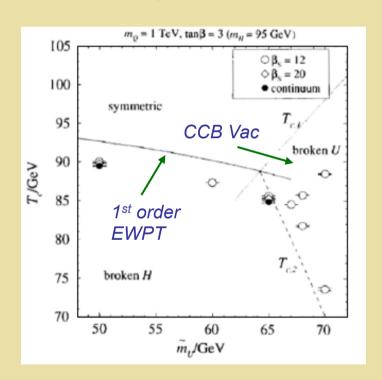
Increasing m_h

New scalars



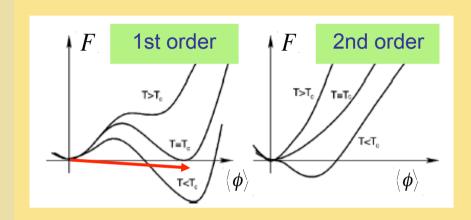
MSSM: Light Stop Scenario

Lattice: Laine, Rummukainen



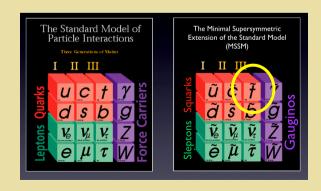
Decreasing RH stop mass

EW Phase Transition: SUSY



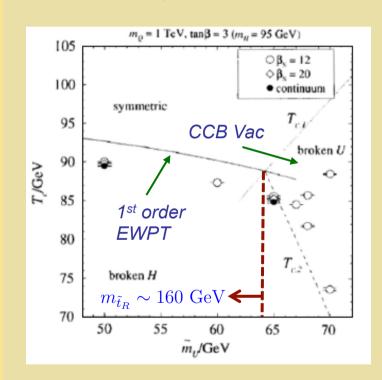
Increasing m_h

New scalars



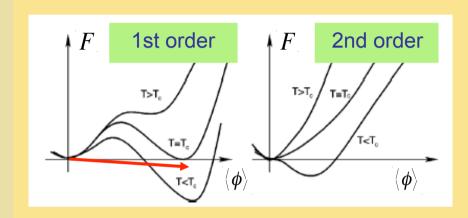
MSSM: Light Stop Scenario

Lattice: Laine, Rummukainen



Decreasing RH stop mass

EW Phase Transition: MSSM

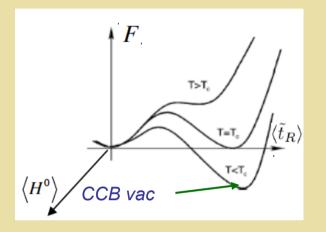


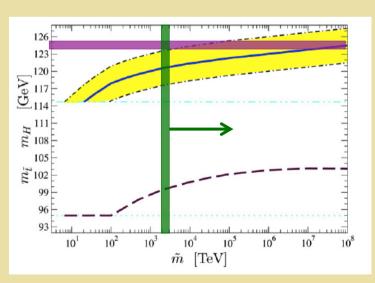
Increasing m_h

—— New scalars

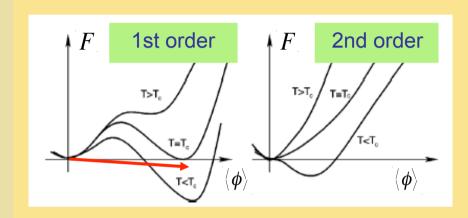
MSSM: Light RH stops

Carena et al 2008: MSSM strong 1st order EWPT: RH stop mass < 105 GeV





EW Phase Transition: MSSM

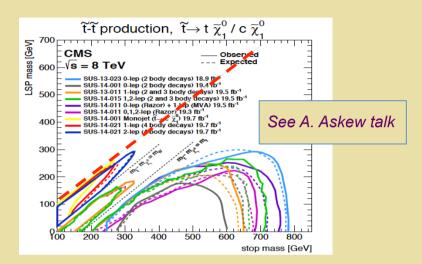


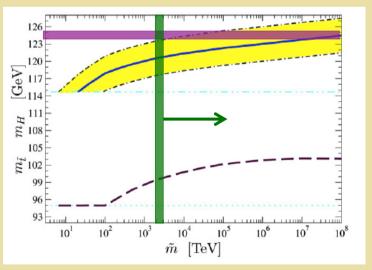
Increasing m_h

New scalars

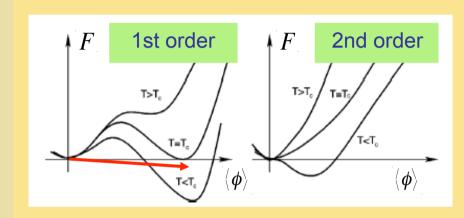
MSSM: Light RH stops

Carena et al 2008: MSSM strong 1st order EWPT: RH stop mass < 105 GeV





EW Phase Transition: MSSM

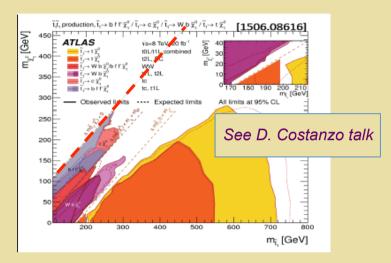


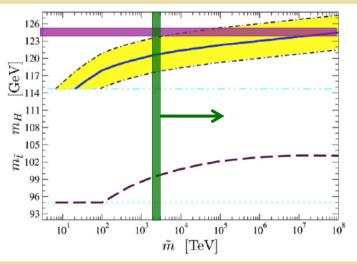
Increasing m_h

New scalars

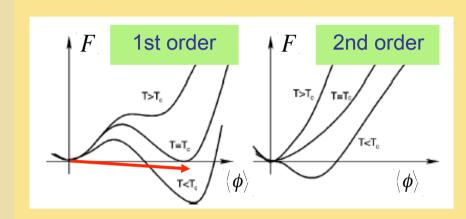
MSSM: Light RH stops

Carena et al 2008: MSSM strong 1st order EWPT: RH stop mass < 105 GeV





EW Phase Transition: SUSY



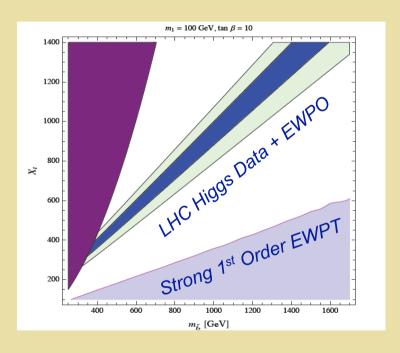
Increasing m_h

← New scalars

Light RH stops also affect Higgs properties

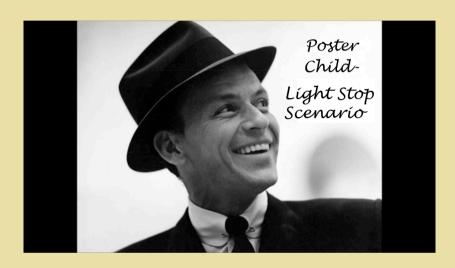
Curtin, Jaiswal, Meade 1203.2932

$MSSM + \delta \lambda_4 (H_u^{\dagger} H_u)^2$



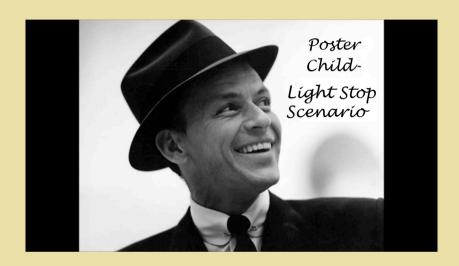
Katz, Perelstein, R-M, Winslow 1509.02934

Beyond the Poster Child



- Gauge singlets (tree-level)
- EW multiplets (tree + loops)

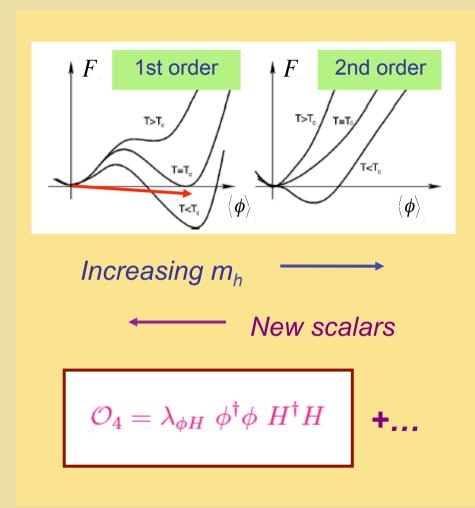
Beyond the Poster Child



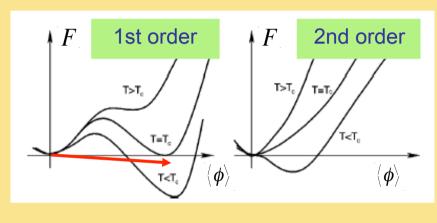
- Gauge singlets (tree-level)
- EW multiplets (tree + loops)

Higgs portal: SUSY or otherwise

EW Phase Transition: Higgs Portal



EW Phase Transition: Higgs Portal

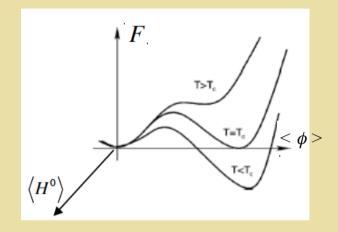


Increasing m_h

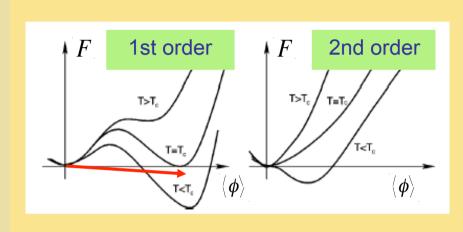
New scalars

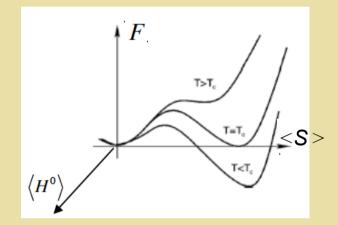
$$\mathcal{O}_4 = \lambda_{\phi H} \; \phi^\dagger \phi \; H^\dagger H$$





- Renormalizable
- φ : singlet or charged under SU(2)_L x U(1)_Y
- Generic features of full theory (NMSSM, GUTS...)
- More robust vacuum stability
- Novel patterns of SSB



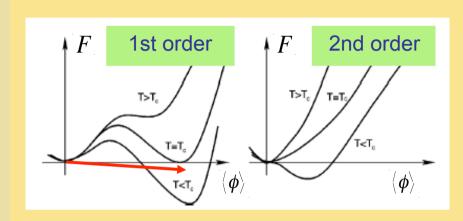


Increasing m_h

New scalars

Real Singlet: $\phi \rightarrow S$

Simplest Extension: two states h_1 & h_2

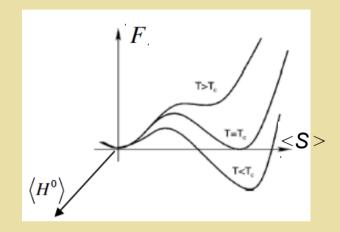


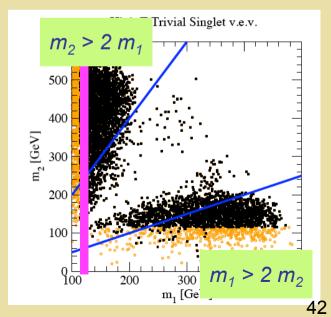
Increasing m_h

New scalars

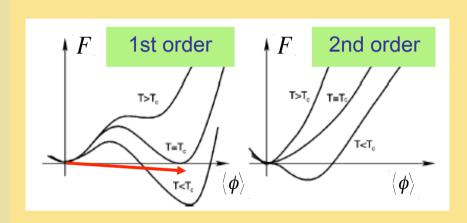
Real Singlet: $\phi \rightarrow S$

Simplest Extension: two states h_1 & h_2





Profumo, R-M, Shaugnessy JHEP 0708 (2007) 010

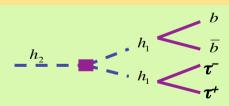


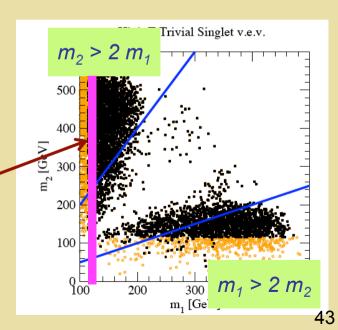
F $T > T_c$ $T = T_c$ $S > T < T_c$

Increasing m_h

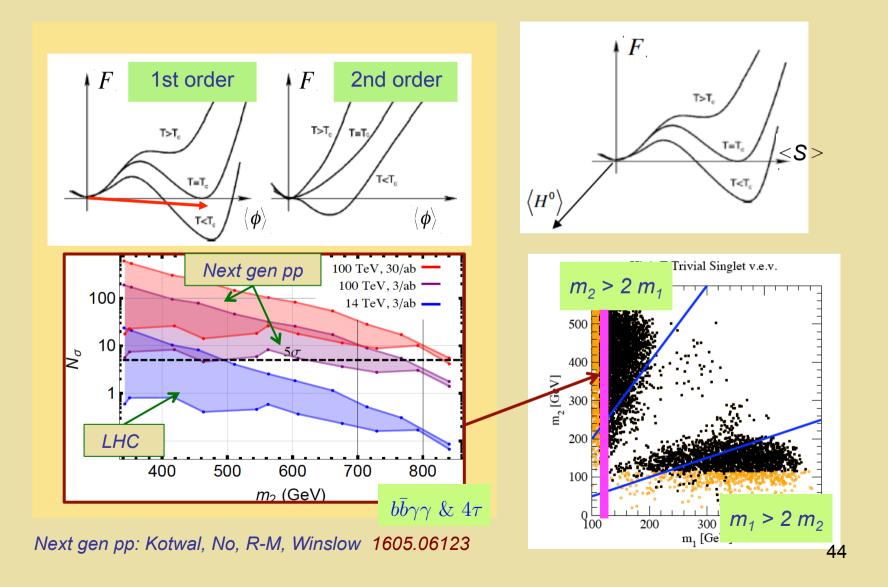
— New scalars

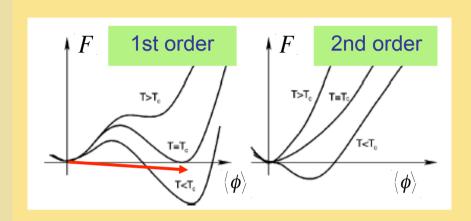
Resonant di-Higgs production:

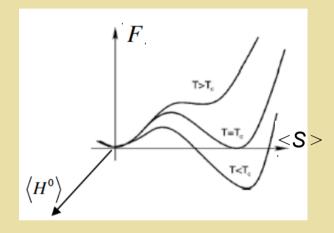


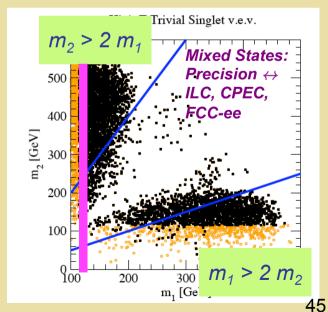


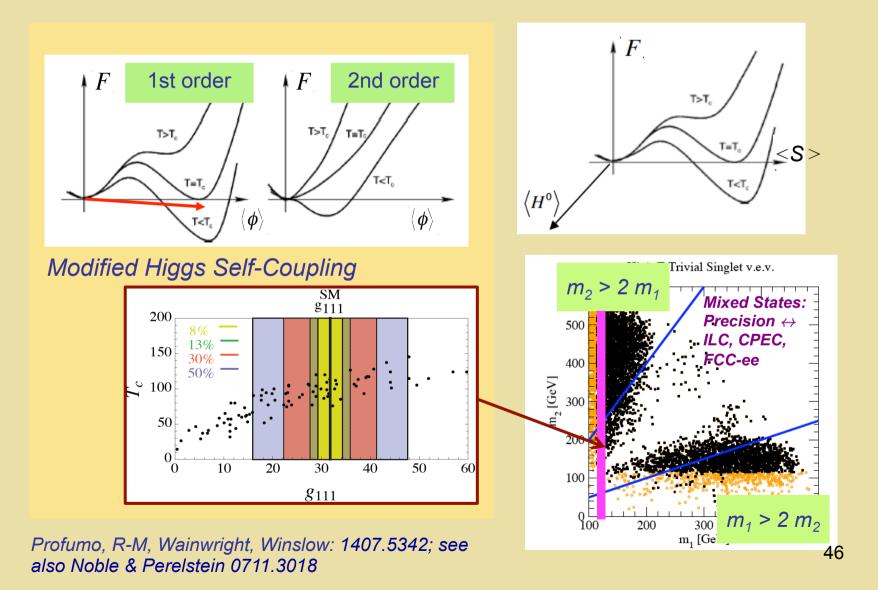
No & RM, arXiv:1310.6035 : LHC Discovery w/ 100 fb⁻¹

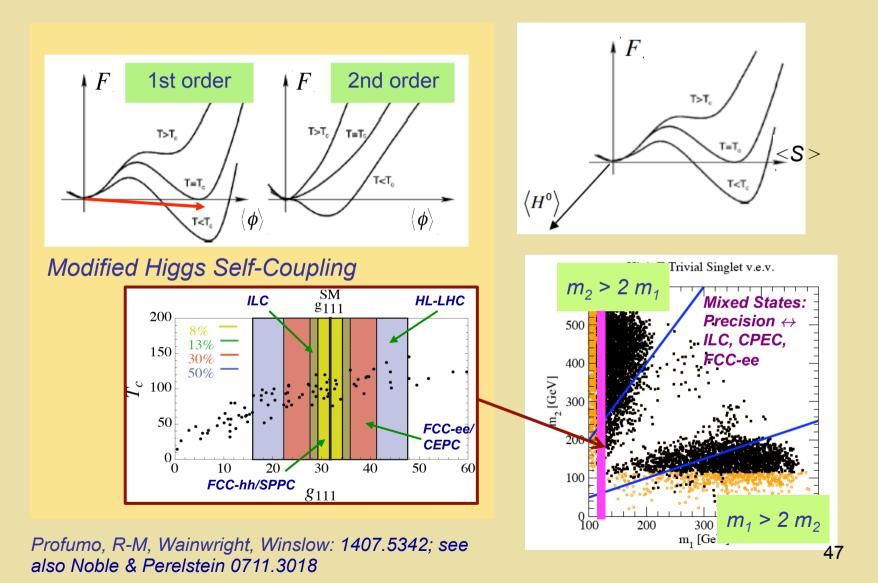


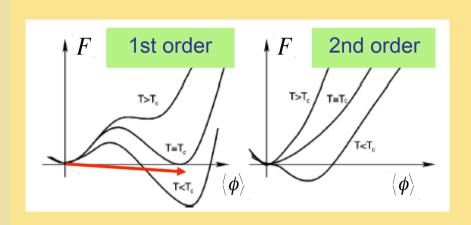


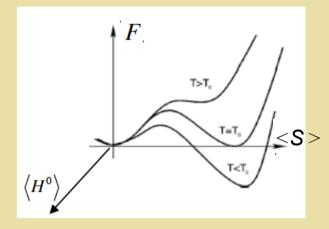






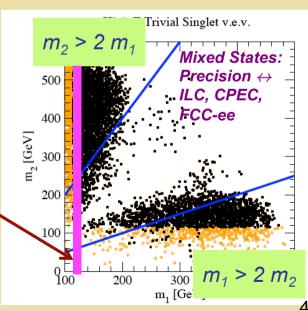


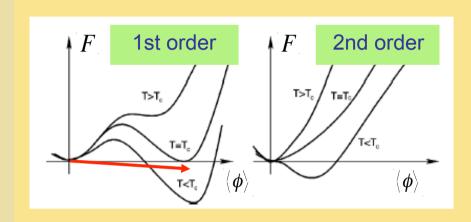




Exotic Higgs Decays

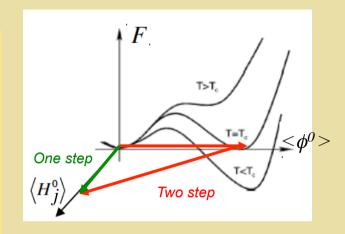
?

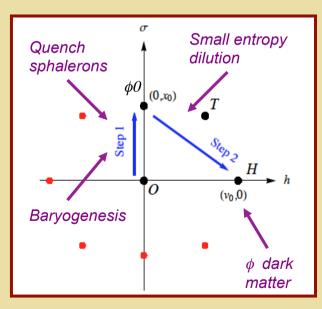


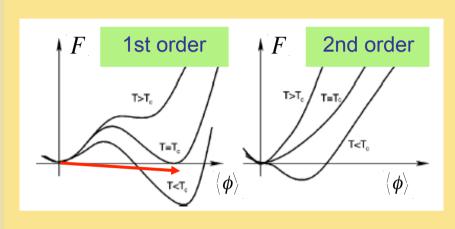


Increasing m_h New scalars

- Step 1: thermal loops
- Step 2: tree-level barrier





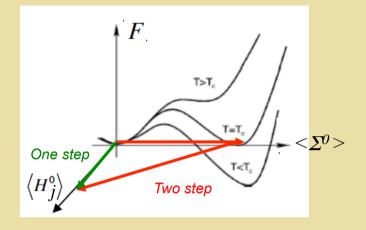


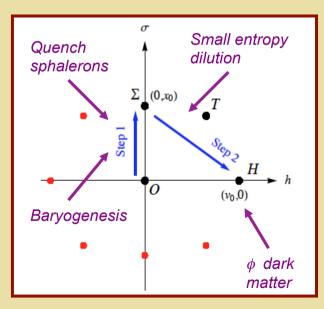
Increasing m_h

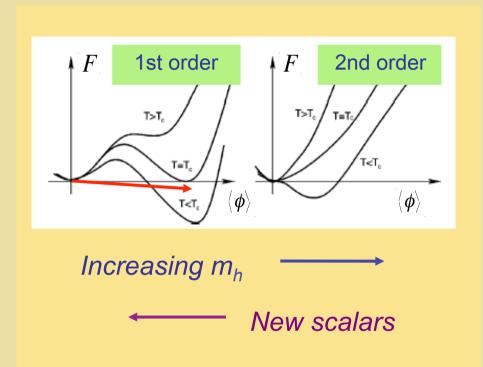
── New scalars

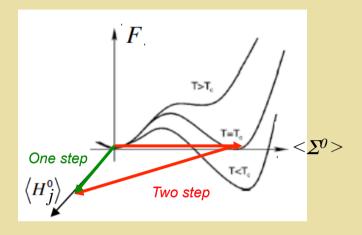
Real Triplet $\Sigma \sim (1,3,0)$

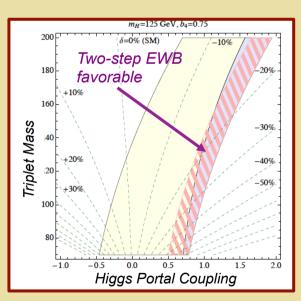
Two-step EWPT & dark matter

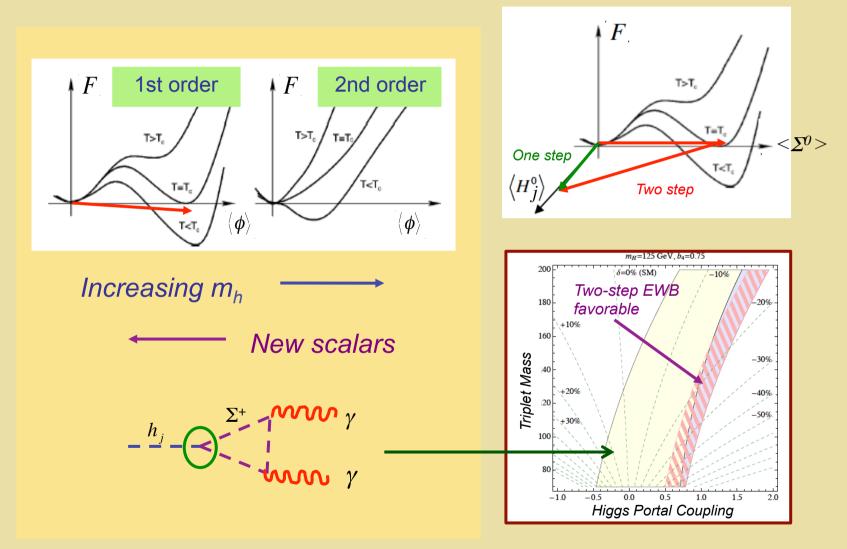












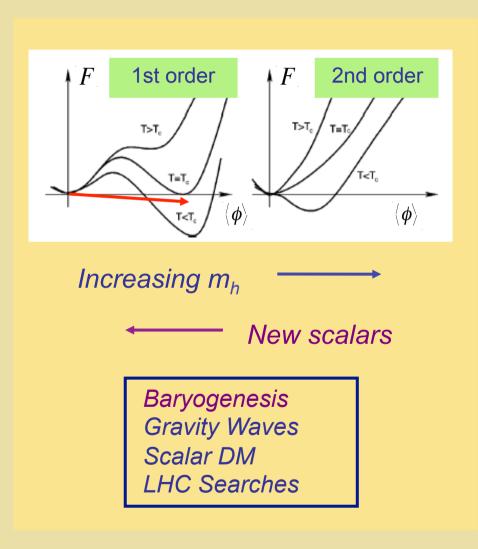
Strong 1st Order EWPT

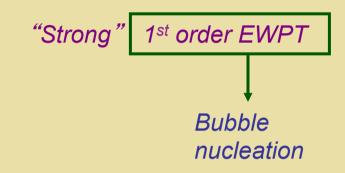


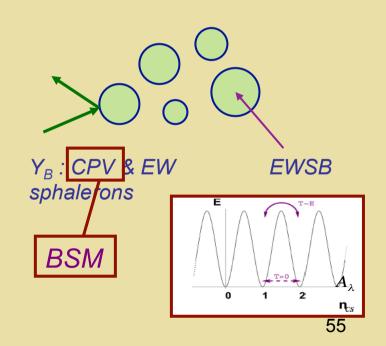


Definitive probe of the possibilities \rightarrow LHC + next generation colliders

III. CPV: Baryon Asymmetry & EDMs







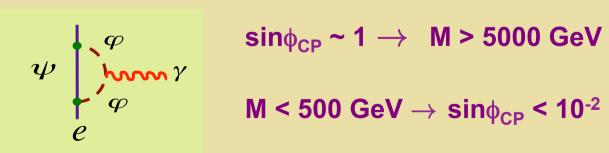
System	Limit (e cm)*	SM CKM CPV	BSM CPV
¹⁹⁹ Hg	7.4 x 10 ⁻³⁰	10 ⁻³³	10 ⁻²⁹
ThO	8.7 x 10 ⁻²⁹ **	10 ⁻³⁸	10 ⁻²⁸
n	3.3 x 10 ⁻²⁶	10 - ³¹	10 ⁻²⁶

^{* 95%} CL ** e- equivalent

System	Limit (e cm)*	SM CKM CPV	BSM CPV
¹⁹⁹ Hg	7.4 x 10 ⁻³⁰	10 ⁻³³	10 ⁻²⁹
ThO	8.7 x 10 ⁻²⁹ **	10 ⁻³⁸	10 ⁻²⁸
n	3.3 x 10 ⁻²⁶	10 - ³¹	10 ⁻²⁶

* 95% CL ** e-equivalent

Mass Scale Sensitivity

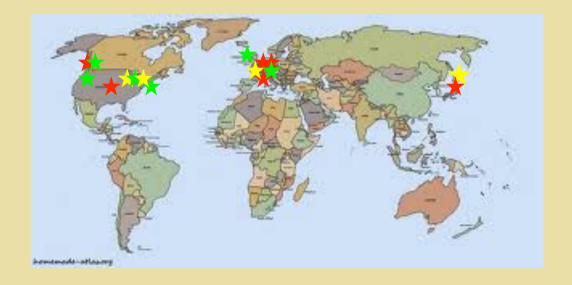


$$\sin\phi_{CP} \sim 1 \rightarrow M > 5000 \text{ GeV}$$

M < 500 GeV
$$ightarrow$$
 sin $\phi_{ extsf{CP}}$ < 10-2

System	Limit (e cm)*	SM CKM CPV	BSM CPV
¹⁹⁹ Hg	7.4 x 10 ⁻³⁰	10 ⁻³³	10 ⁻²⁹
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n	3.3 x 10 ⁻²⁶	10 - ³¹	10 ⁻²⁶

* 95% CL ** e-equivalent



- * neutron
- proton& nuclei
- ★ atoms

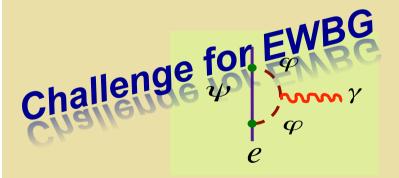
~ 100 x better sensitivity

Not shown: muon

System	Limit (e cm)*	SM CKM CPV	BSM CPV
¹⁹⁹ Hg	7.4 x 10 ⁻³⁰	10 ⁻³³	10 ⁻²⁹
ThO	8.7 x 10 ⁻²⁹ **	10 ⁻³⁸	10 ⁻²⁸
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^{* 95%} CL ** e-equivalent

Mass Scale Sensitivity



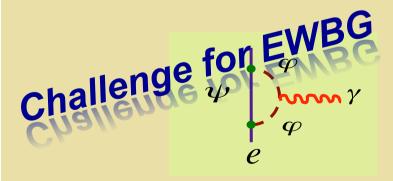
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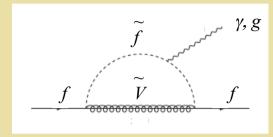
* 95% CL ** e-equivalent

Mass Scale Sensitivity

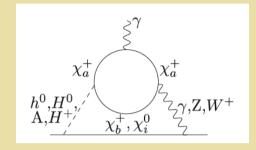


- EDMs arise at > 1 loop
- CPV is flavor non-diagonal
 - CPV is "partially secluded"

EDMs & EWBG: MSSM & Beyond

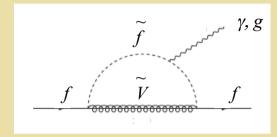


Heavy sfermions: LHC consistent & suppress 1-loop EDMs

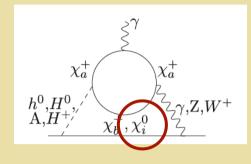


Sub-TeV EW-inos: LHC & EWB - viable but non-universal phases

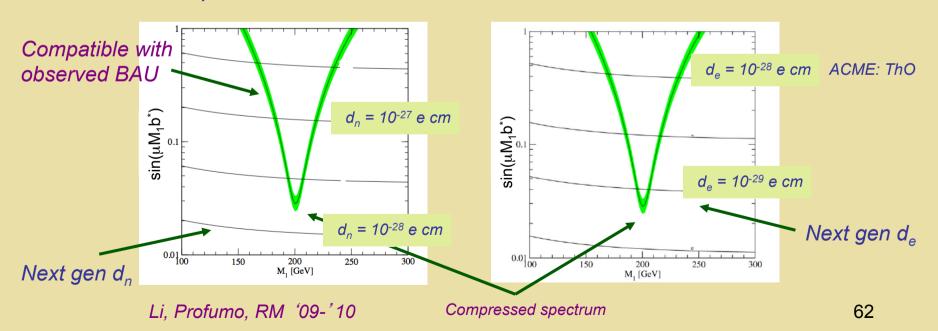
EDMs & EWBG: MSSM & Beyond



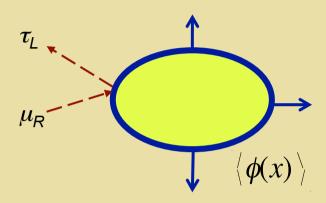
Heavy sfermions: LHC consistent & suppress 1-loop EDMs



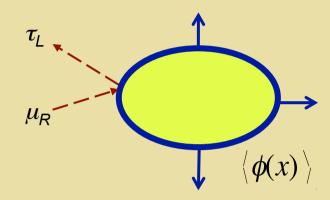
Sub-TeV EW-inos: LHC & EWB - viable but non-universal phases











Flavor basis (high T)

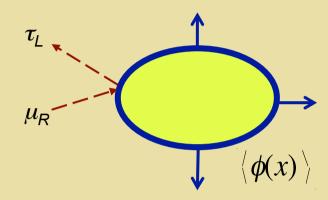
$$\mathcal{L}_{\text{Yukawa}}^{\text{Lepton}} = -\overline{E_L^i} \left[(Y_1^E)_{ij} \Phi_1 + (Y_2^E)_{ij} \Phi_2 \right] e_R^j + h.c.$$

Mass basis (T=0)

$$\frac{m_f}{v}\kappa_{\tau}(\cos\phi_{\tau}\bar{\tau}\tau + \sin\phi_{\tau}\bar{\tau}i\gamma_{5}\tau)h$$

Guo, Li, Liu, R-M, Shu 1607.XXXX





Flavor basis (high T)

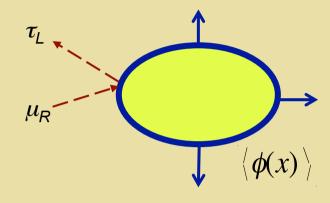
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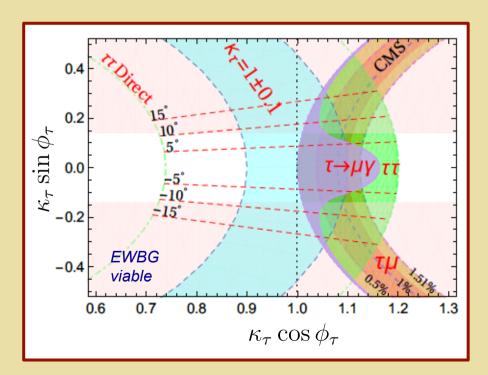
Mass basis (T=0)

$$rac{\mathit{CPV}\,\mathsf{h} o au au}{v} \kappa_{ au} (\cos\phi_{ au}ar{ au} au + \sin\phi_{ au}ar{ au}i\gamma_5 au) h$$

Guo, Li, Liu, R-M, Shu 1607.XXXX







Flavor basis (high T)

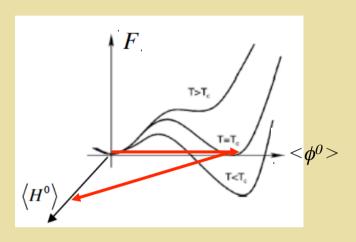
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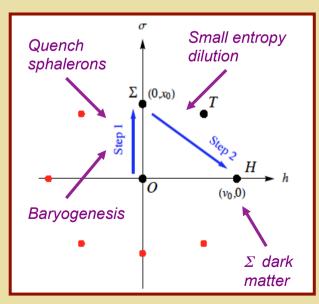
Mass basis (T=0)

$$rac{\mathit{CPV}\,\mathsf{h} o au au}{v} \kappa_{ au} (\cos\phi_{ au}ar{ au} au + \sin\phi_{ au}ar{ au}i\gamma_{5} au) h$$

Guo, Li, Liu, R-M, Shu 1607.XXXX

Two-Step EW Baryogenesis





Illustrative Model:

New sector: "Real Triplet" Σ Gauge singlet S

H → Set of "SM" fields: 2 HDM

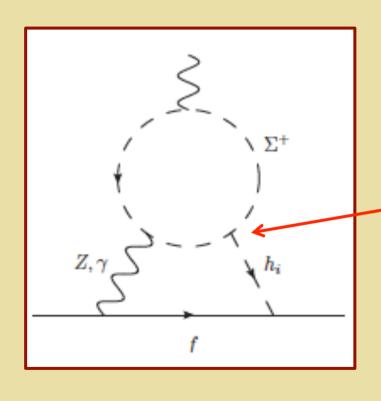
(SUSY: "TNMSSM", Coriano...)

Two CPV Phases:

 δ_{Σ} : Triplet phase

 $\delta_{\mathcal{S}}$: Singlet phase

Two-Step EW Baryogenesis & EDMs



EDMs are Two Loop

Two CPV Phases:

 δ_{Σ} :

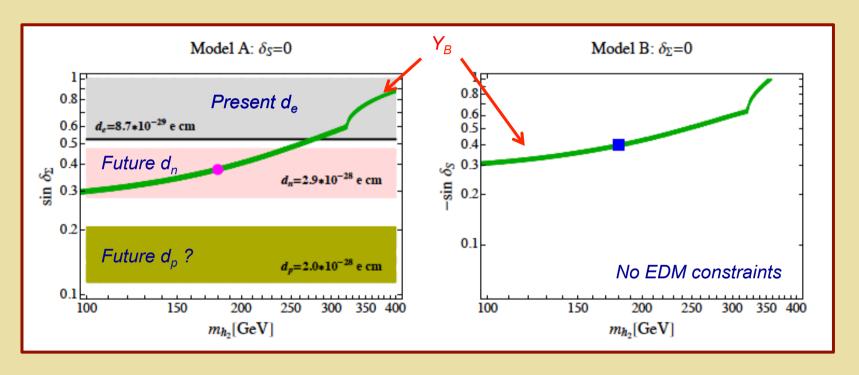
Triplet phase

Singlet phase

Insensitive to δ_S : electrically neutral o "partially secluded"

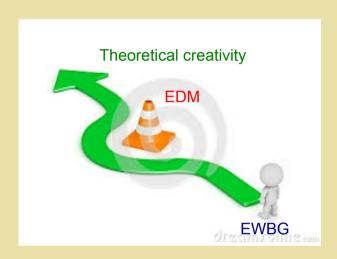
Two-Step EW Baryogenesis & EDMs

Two cases: (A) $\delta_S = 0$ (B) $\delta_\Sigma = 0$



CPV for **EWBG**





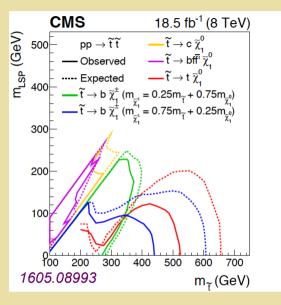
IV. Outlook

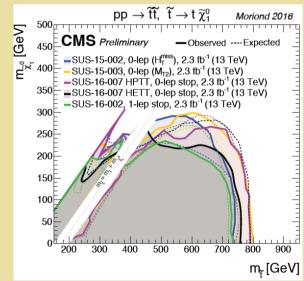
- Explaining the origin of the matter-antimatter asymmetry is a forefront challenge for BSM physics
- Electroweak baryogenesis remains one of the most theoretically rich & experimentally accessible scenarios: "Was the baryon asymmetry produced in conjunction with electroweak symmetry-breaking?"
- EDMs & collider studies (LHC & beyond) provide powerful probes of the ingredients & results to date challenge theoretical creativity
- Exciting array of possibilities to be explored

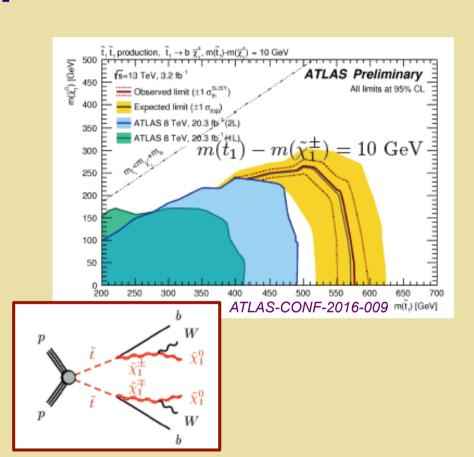
Refs: 1604.05324, 1206.2942

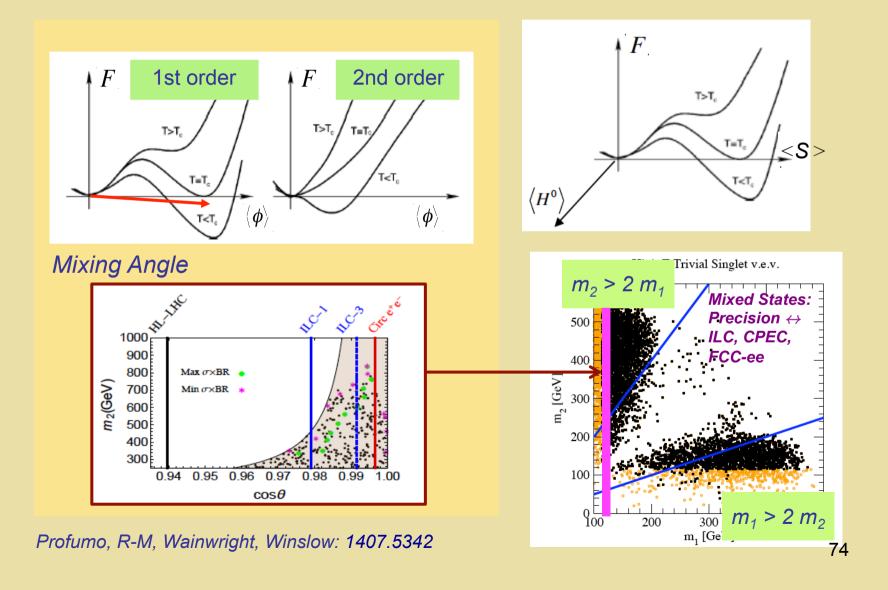
Back Up Slides

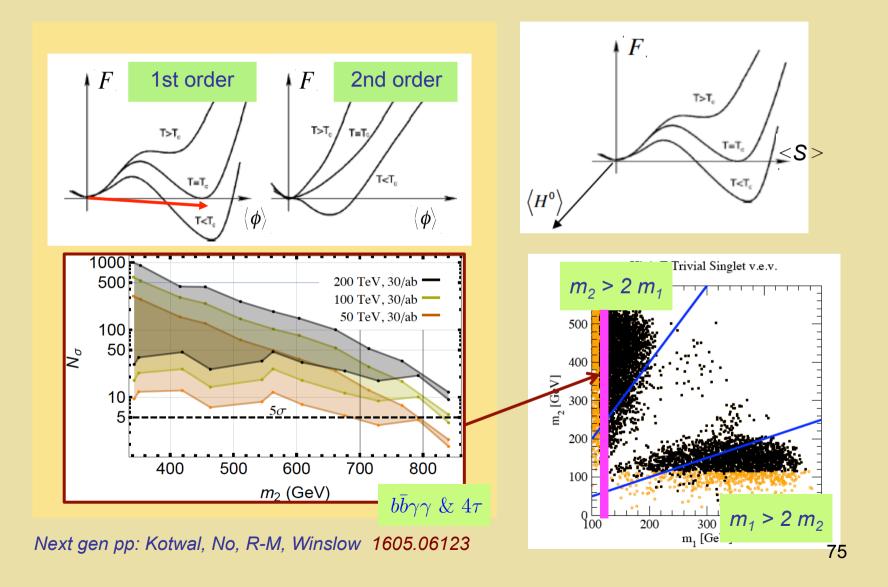
LHC Stop Searches



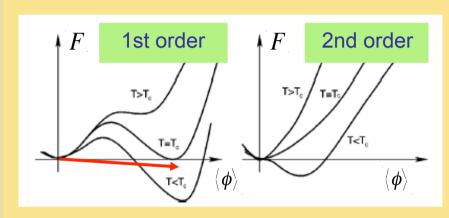








EW Phase Transition: 100 TeV pp



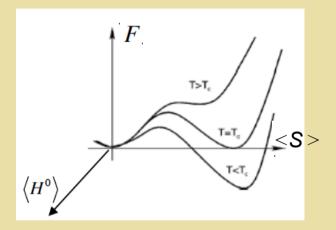
Curtain, Meade, Yu: arXiv: 1409.0005

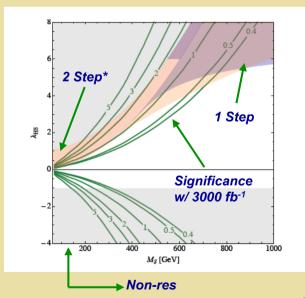
Z₂ symmetric real singlet extension

- Loop-induced 1-step transition
- 2-step transition for $\mu_S^2 < 0$

VBF @ 100 TeV pp:

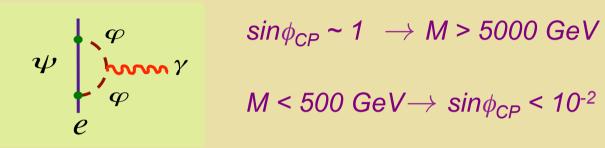
$$pp \rightarrow h jj$$
, $h \rightarrow invis$





^{*} Singlet two step: see also Profumo, R-M, Shaugnessy 2007

EDMs & EWBG: MSSM & Beyond



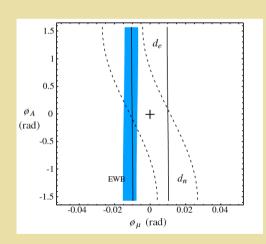
$$sin\phi_{CP} \sim 1 \rightarrow M > 5000 \text{ GeV}$$

$$M < 500 \; \text{GeV} \rightarrow \sin \phi_{CP} < 10^{-2}$$

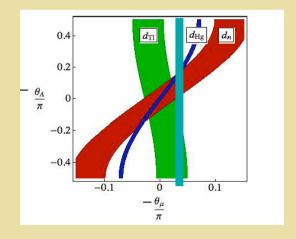
Universal gaugino phases

 $Arg(\mu M_i b^*) =$

 $Arg(\mu M_i b^*)$



Cirigliano, R-M, Tulin, Lee '06



Ritz CIPANP 09 + Cirigliano, R-M, Tulin, Lee '06 77