

Into the multi-TeV scale with $H \rightarrow \gamma\gamma/H \rightarrow ZZ^*$

Abdelhak DJOUADI

(LPT CNRS & U. Paris-Sud)

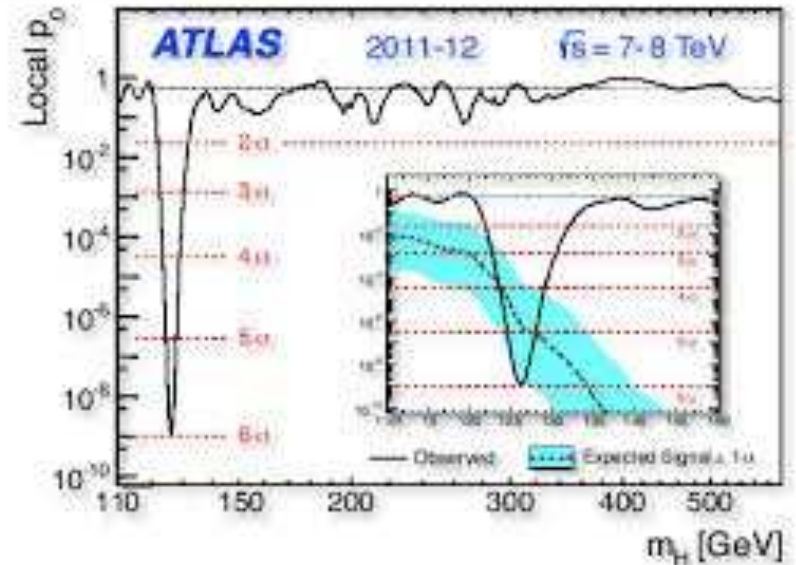
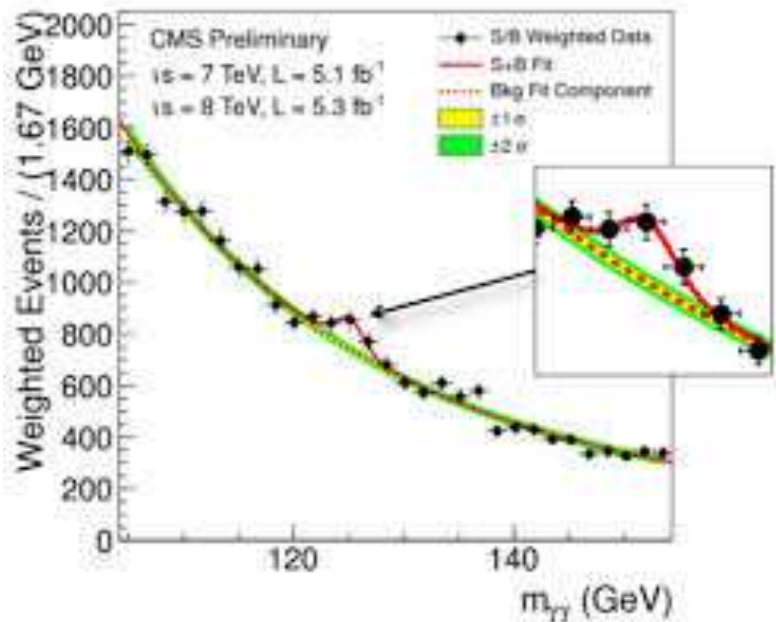
What next after the Higgs discovery?

$D_{\gamma\gamma}$

Search for BSM with $D_{\gamma\gamma}$

What next after the Higgs discovery?

Now that the Higgs is discovered and proved to be approximately SM-like.



Is particle physics closed and we should all go home/multiverse?

What next after the Higgs discovery?

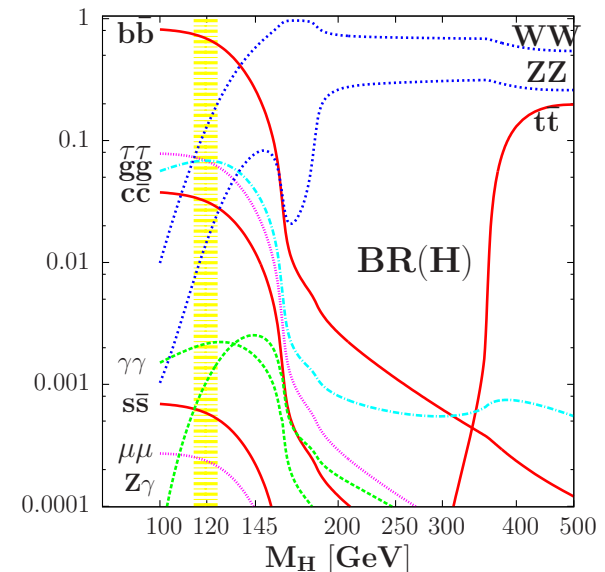
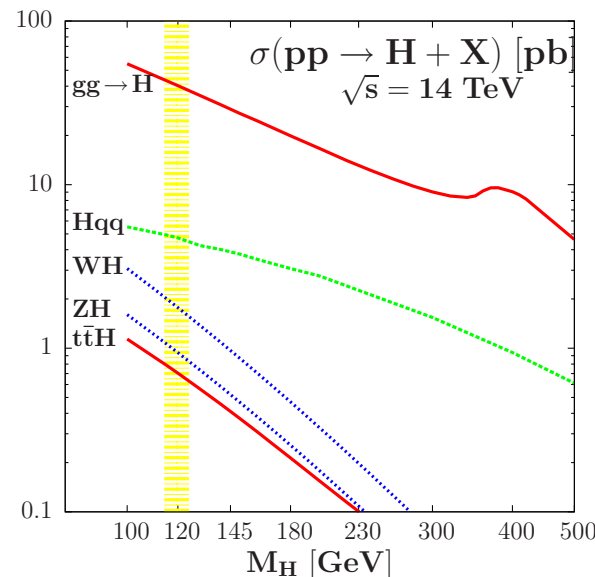
What should we be doing the next 10–30 years in Particle Physics?

Need to check that H is indeed responsible of sEWSB (and SM-like?)

⇒ measure its fundamental properties in the most precise way:

- its mass and total decay width (invisible width due to dark matter?),
- its spin–parity quantum numbers (CP violation for baryogenesis?),
- its couplings to fermions and gauge bosons and check if they are only proportional to particle masses (no new physics contributions?),
- its self-couplings to reconstruct the potential V_S that makes EWSB.

Possible for $M_H \approx 125$ GeV as all production/decay channels useful!



What next after the Higgs discovery?

In fact part of this second chapter has already started. Latest results on

$$\mu_{\mathbf{XX}} = \sigma^{\mathbf{p}}(\mathbf{pp} \rightarrow \mathbf{H}) \times \mathbf{BR}(\mathbf{H} \rightarrow \mathbf{XX})|_{\text{exp/SM}}$$

$\sigma \times$ BRs compatible with those expected in the SM

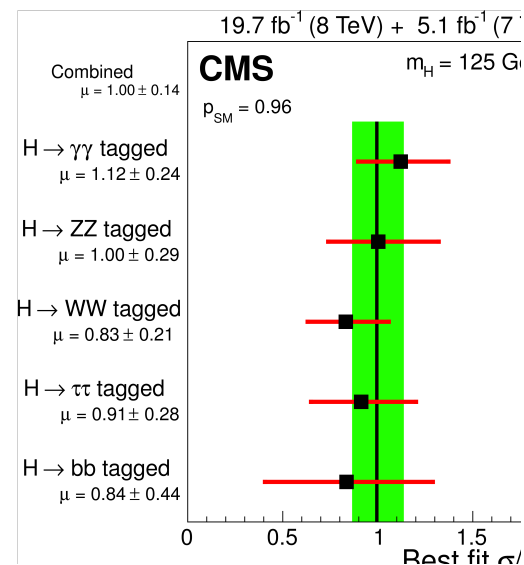
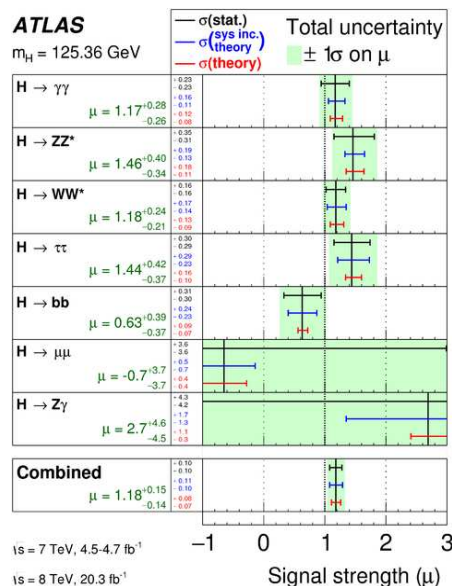
**Fit of all LHC Higgs data \Rightarrow
agreement at 15–30% level**

$$\mu_{\text{tot}}^{\text{ATLAS}} = 1.18 \pm 0.15$$

⇒ Pierre

$$\mu_{\text{tot}}^{\text{CMS}} = 1.00 \pm 0.14$$

⇒ **Guillermo**



Measurement for couplings already precise at the 10–15% level!

Marco St Petersburg: $\mu_{\text{tot}}^{\text{ATLAS+CMS}} = 1.09^{+0.07+0.04+0.07}_{-0.07-0.04-0.06} \approx 1.1 \pm 0.1$

This is particularly the case in the two very clean detection channels

$$\mathbf{H} \rightarrow \gamma\gamma, \mathbf{H} \rightarrow \mathbf{ZZ}^* \rightarrow 4\ell^\pm$$

What next after the Higgs discovery?

channel	ATLAS	CMS
$\mu_{\gamma\gamma}$	$1.17^{+0.23}_{-0.23}^{+0.16}_{-0.11} \left(\begin{smallmatrix} +0.12 \\ -0.08 \end{smallmatrix} \right)$	$1.14^{+0.21}_{-0.21}^{+0.16}_{-0.10} \left(\begin{smallmatrix} +0.09 \\ -0.05 \end{smallmatrix} \right)$
μ_{ZZ}	$1.46^{+0.35}_{-0.31}^{+0.19}_{-0.13} \left(\begin{smallmatrix} +0.18 \\ -0.11 \end{smallmatrix} \right)$	$0.93^{+0.26}_{-0.23}^{+0.13}_{-0.09}$

Is this enough to probe effects of new physics or BSM?

No! Not in the case of weakly interacting theories like 2HDM, SUSY, etc...
effects expected to be at level of $\Delta\mu_{XX} \approx \frac{C_{\text{NEW}}\alpha_W}{\pi} \approx \frac{M_h^2}{M_{\text{NEW}}^2} \approx \text{a few \%}$

Is a 1% accuracy achievable at upgraded LHC with high luminosities?

- **Statistical uncertainty:** $20\% / \sqrt{3 \times 100} \lesssim 1\%$
at least in the clean $H \rightarrow \gamma\gamma, VV$ channels
- **Systematical uncertainties:** can be reduced at the level of a few %
some common to many channels (lumi...).
- **Theoretical uncertainty:** will be by far the limiting issue!

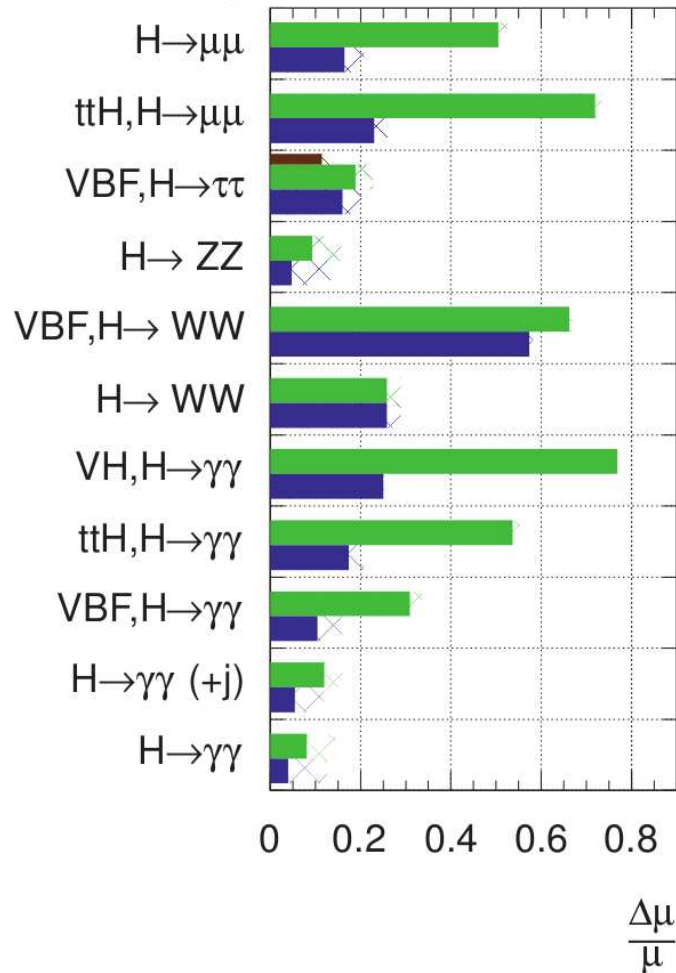
\Rightarrow How big is it? How much can it be reduced? Can it be removed?

What next after the Higgs discovery?

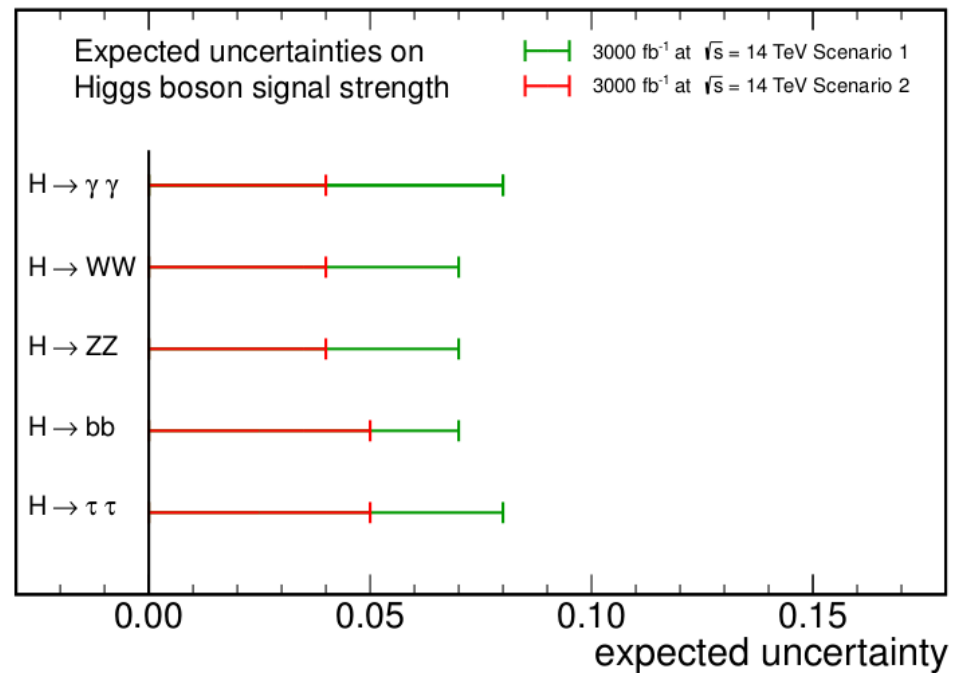
ATLAS Simulation

$\sqrt{s} = 14$ TeV: $\int L dt = 300 \text{ fb}^{-1}$; $\int L dt = 3000 \text{ fb}^{-1}$

$\int L dt = 300 \text{ fb}^{-1}$ extrapolated from 7+8 TeV



CMS Projection



$$D_{\gamma\gamma}$$

Best way to eliminate the theory uncertainty is to use ratios of signal rates

Take for instance $H \rightarrow VV$ with $V = W \rightarrow \ell\nu$ or $Z \rightarrow \ell\ell$ as reference,
and for detection channel $H \rightarrow XX$ with Higgs produced in process p :

$$\begin{aligned} D_{XX} &= \sigma^P(pp \rightarrow H \rightarrow XX) / \sigma^P(pp \rightarrow H \rightarrow VV) \\ &= \sigma^P(pp \rightarrow H) \times \text{BR}(H \rightarrow XX) / \sigma^P(pp \rightarrow H) \times \text{BR}(H \rightarrow VV) \\ &= \text{BR}(H \rightarrow XX) / \text{BR}(H \rightarrow VV) \\ &= \Gamma(H \rightarrow XX) / \Gamma(H \rightarrow VV) \end{aligned}$$

$$D_{XX} = c_X^2 / c_V^2$$

**Works only if one selects exactly the same kinematical configuration
(i.e. same "fiducial cross sections") for the two channels X and V!**

- The theoretical uncertainties from the cross sections drop out
- The parametric uncertainties from the branching ratios drop out
- The theoretical ambiguities in the Higgs total width also drop out

$\Rightarrow D_{XX}$ measures only the ratio of squared couplings!

$D_{\gamma\gamma}$

- Extremely clean theoretically, although some information will be lost.
- And maybe it has also some advantages from the experimental side?
e.g. some common experimental systematical errors also drop out:
 - common uncertainty from the luminosity measurement
 - other common systematics such as errors on efficiencies etc...?

The **decay ratios** that can already be built are the following ones:

$$D_{ww} = \frac{\sigma(pp \rightarrow H \rightarrow WW)}{\sigma(pp \rightarrow H \rightarrow VV)} = \frac{\Gamma(H \rightarrow WW)}{\Gamma(H \rightarrow VV)} = d_{ww} \frac{c_W^2}{c_V^2}$$

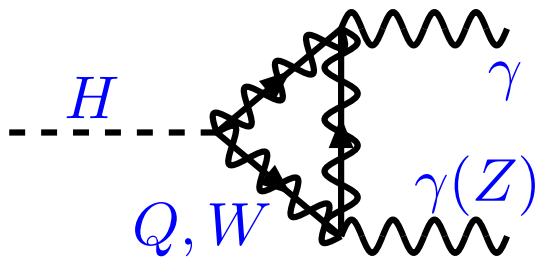
$$D_{\tau\tau} = \frac{\sigma(pp \rightarrow H \rightarrow \tau\tau)}{\sigma(pp \rightarrow H \rightarrow VV)} = \frac{\Gamma(H \rightarrow \tau\tau)}{\Gamma(H \rightarrow VV)} = d_{\tau\tau} \frac{c_\tau^2}{c_V^2}$$

$$D_{bb} = \frac{\sigma(q\bar{q} \rightarrow HV \rightarrow bbV)}{\sigma(q\bar{q} \rightarrow HV \rightarrow VVV)} = \frac{\Gamma(H \rightarrow bb)}{\Gamma(H \rightarrow VV)} = d_{bb} \frac{c_\tau^2}{c_V^2}$$

$$D_{\gamma\gamma} = \frac{\sigma(pp \rightarrow H \rightarrow \gamma\gamma)}{\sigma(pp \rightarrow H \rightarrow VV)} = \frac{\Gamma(H \rightarrow \gamma\gamma)}{\Gamma(H \rightarrow VV)} = d_{\gamma\gamma} \frac{c_\gamma^2}{c_V^2}$$

Best probe by far is $D_{\gamma\gamma}$ which measures the deviation of the $\gamma\gamma$ loop!

AD, Eur.Phys.J. C73 (2013) 2498, arXiv:1208.3436



$$\Gamma = \frac{G_\mu \alpha^2 M_H^3}{128 \sqrt{2} \pi^3} \left| \sum_f N_c e_f^2 A_{\frac{1}{2}}^H(\tau_f) + A_1^H(\tau_W) \right|^2$$

$$A_{1/2}^H(\tau) = 2[\tau + (\tau - 1)f(\tau)] \tau^{-2}$$

$$A_1^H(\tau) = -[2\tau^2 + 3\tau + 3(2\tau - 1)f(\tau)] \tau^{-2}$$

- Photon massless and Higgs has no charge: must be a loop decay.
- In SM: only W-loop and top-loop are relevant (b-loop too small).
- For $m_i \rightarrow \infty \Rightarrow A_{1/2} = \frac{4}{3}$ and $A_1 = -7$: W loop dominating!
(approximation $\tau_W \rightarrow 0$ valid only for $M_H \lesssim 2M_W$: relevant here!).

$\gamma\gamma$ width counts the number of charged particles coupling to Higgs!

Contribution A_s^P of particle p of spin s with Higgs coupling g_{Hpp} :

$$A_0^P = -\frac{1}{3}g_{Hpp}^2/m_P^2, A_{1/2}^P = +\frac{4}{3}g_{Hpp}^2/m_P^2, A_1^P = -7g_{Hpp}^2/m_P^2,$$

$$\text{If } g_{Hpp} \propto m_p \Rightarrow A_0^P \rightarrow -\frac{4}{3}, A_{1/2}^P \rightarrow +\frac{1}{3}, A_1^P \rightarrow +7.$$

Small/calculated QCD and EW corrections: only of order few percent.

$$D_{\gamma\gamma}$$

In the SM, the top and W loop contributions to the $H \rightarrow \gamma\gamma$ amplitude is

$$c_\gamma \approx 1.26 \times |c_W - 0.21 c_t|$$

Assuming the custodial symmetry relation $g_{HZZ} = g_{HWW} = c_V$ (which is well checked experimentally and hard to violate in theory)

The SM value of the ratio $D_{\gamma\gamma} = c_\gamma^2 / c_V^2$ is then simply given by

$$c_\gamma^2 / c_V^2 \approx 6.5 \times |1 - \frac{1}{5} c_t / c_V|^2$$

with $c_V = c_t = 1$ in SM. Any new physics effects will alter this value.

Big question: how well this observable can be experimentally measured?

If it is $\mathcal{O}(1\%)$, then best possible probe of new physics at the LHC:

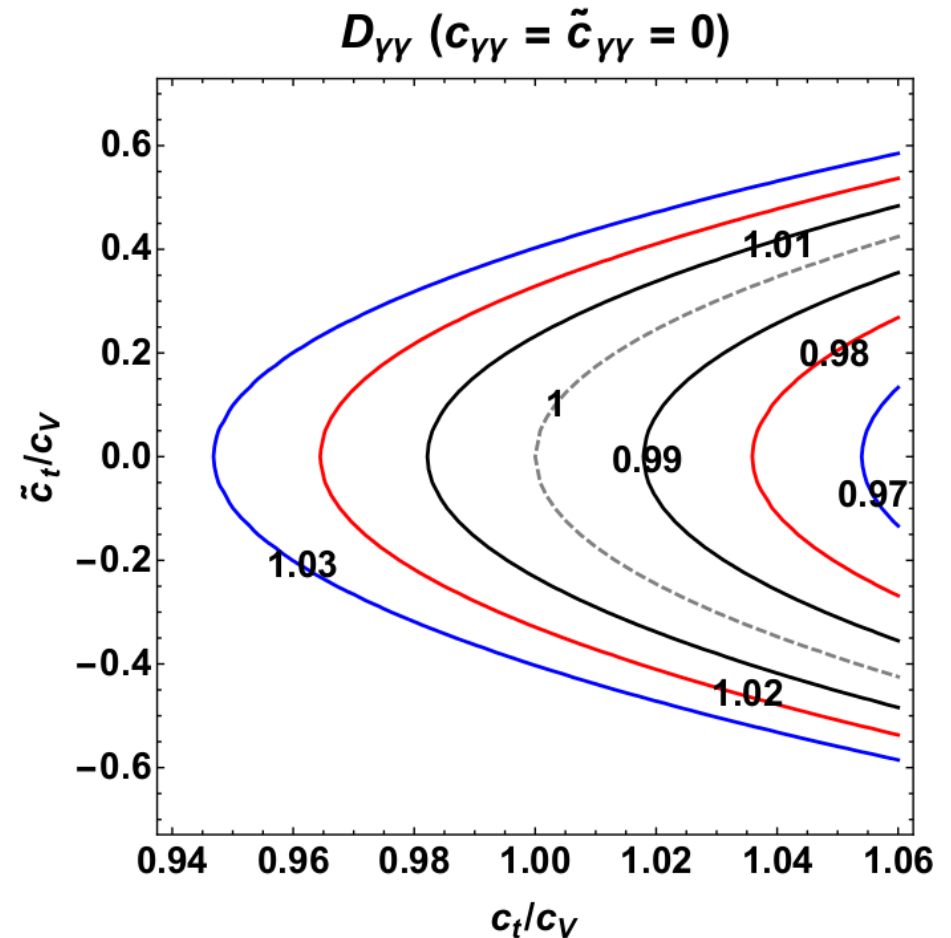
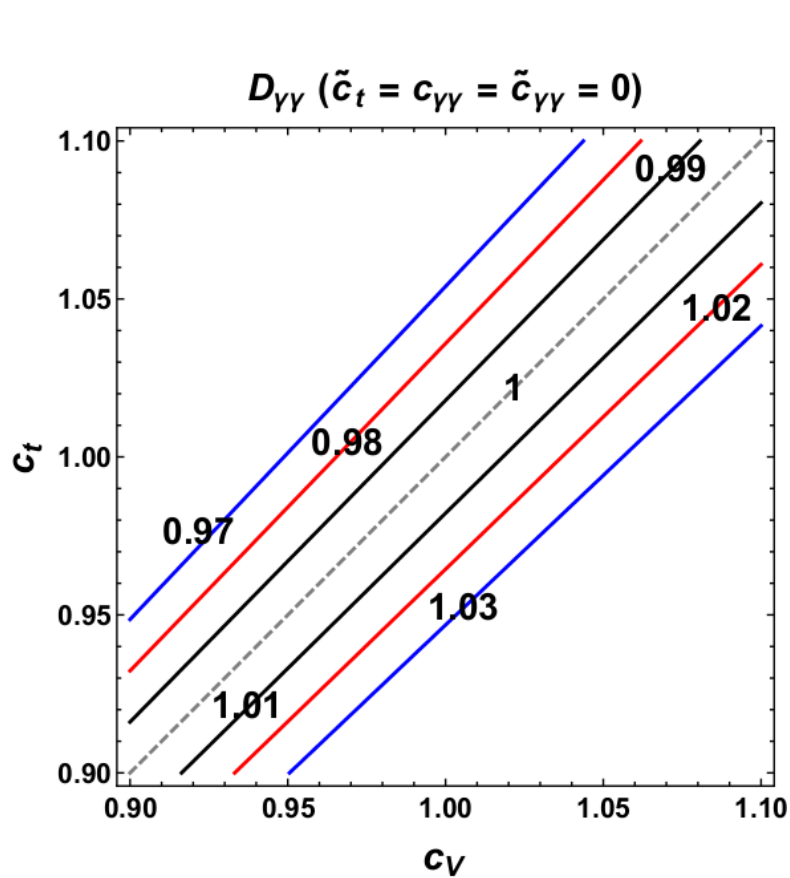
- such accuracy was envisaged only at the "clean" e^+e^- machines..
- impact comparable to $\sin^2\theta_W$ at LEP and M_W at Tevatron/LHC..
- the g-2 of the LHC?

Examples of BSM searches that can be done with the observable follow.

AD, J. Quevillon and R. Vega-Morales, arXiv:1509.03913

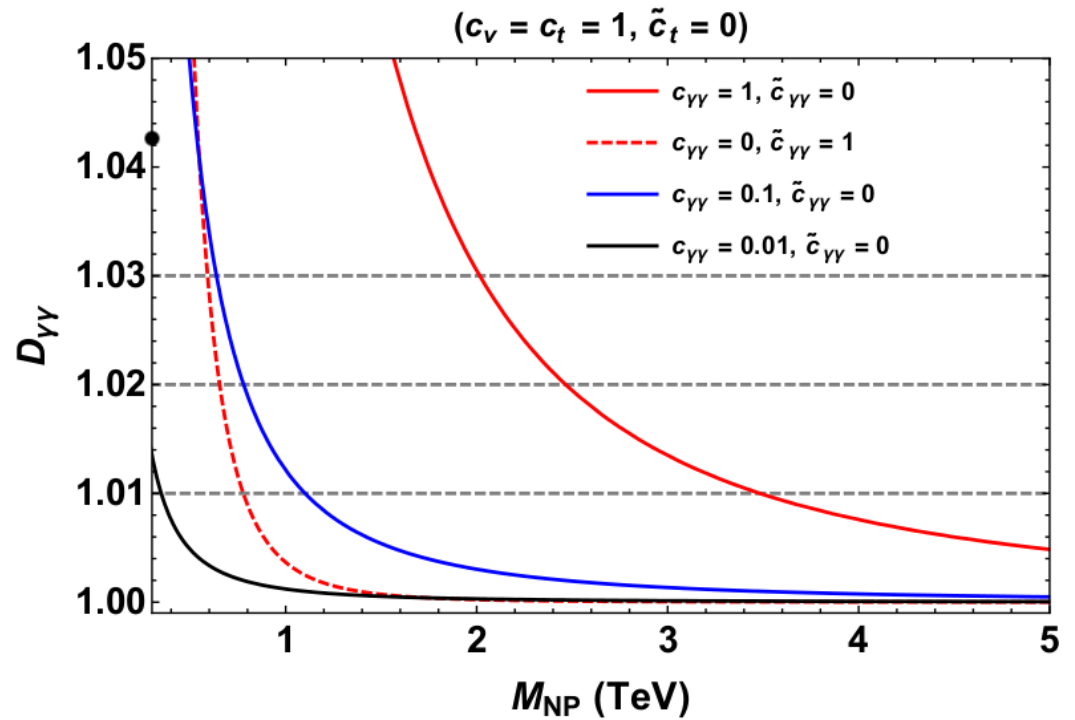
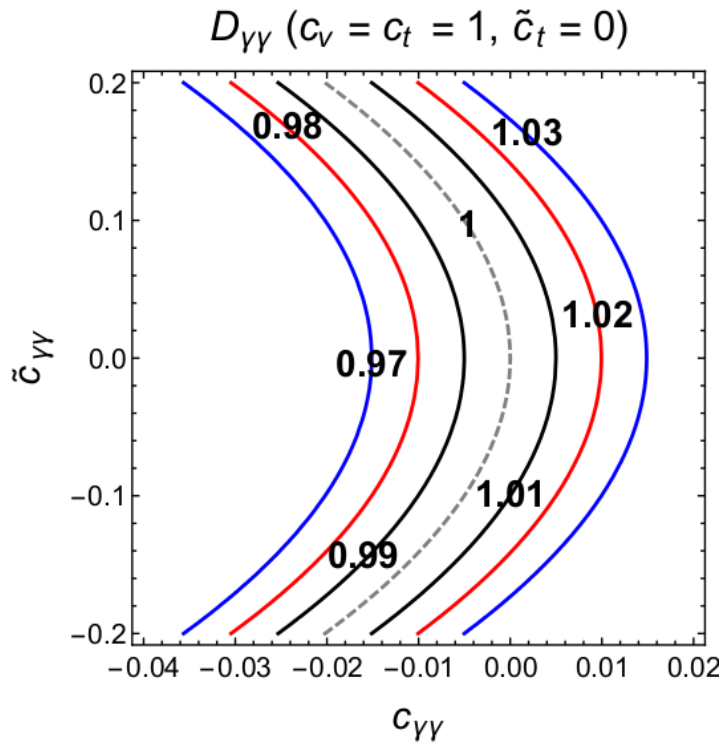
Search for BSM with $D_{\gamma\gamma}$

$$\mathcal{L} = \frac{H}{v} \left(c_V (2M_W^2 W_\mu^+ W^{-\mu} + M_Z^2 Z_\mu Z^\mu) - m_t \bar{t} (c_t + i\tilde{c}_t \gamma^5) t \right. \\ \left. + \frac{c_{\gamma\gamma}}{4} F^{\mu\nu} F_{\mu\nu} + \frac{\tilde{c}_{\gamma\gamma}}{4} \tilde{F}^{\mu\nu} F_{\mu\nu} \right)$$



Search for BSM with $D_{\gamma\gamma}$

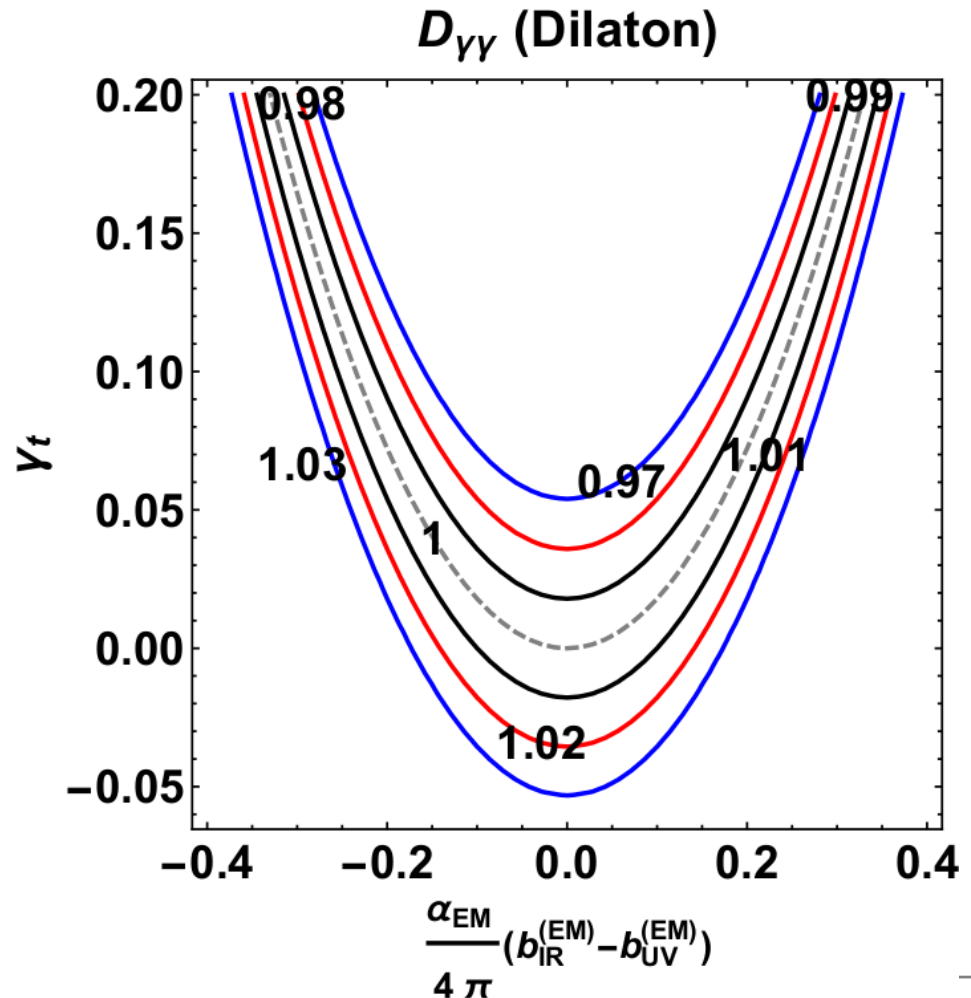
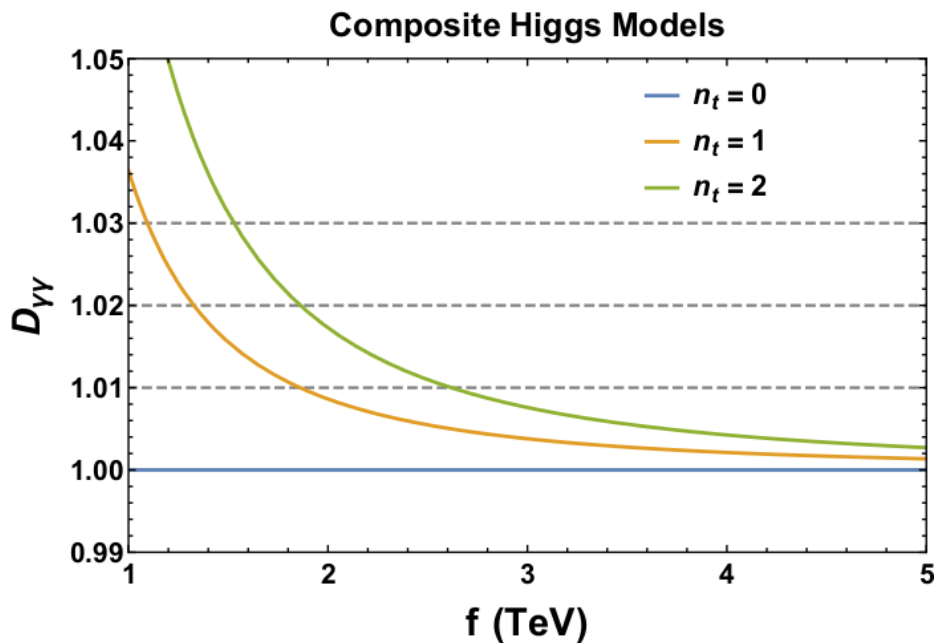
$$\mathcal{L} = \frac{H}{v} \left(c_V (2M_W^2 W_\mu^+ W^{-\mu} + M_Z^2 Z_\mu Z^\mu) - m_t \bar{t} (c_t + i\tilde{c}_t \gamma^5) t \right. \\ \left. + \frac{c_{\gamma\gamma}}{4} F^{\mu\nu} F_{\mu\nu} + \frac{\tilde{c}_{\gamma\gamma}}{4} \tilde{F}^{\mu\nu} F_{\mu\nu} \right)$$



Search for BSM with $D_{\gamma\gamma}$

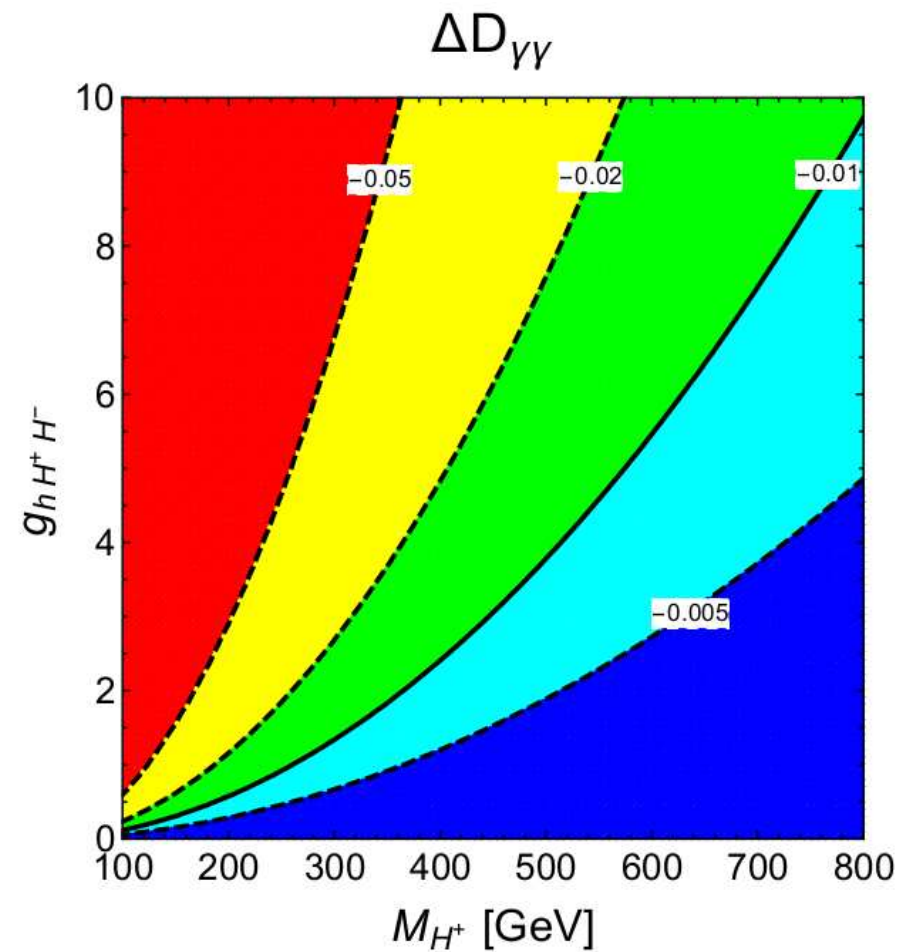
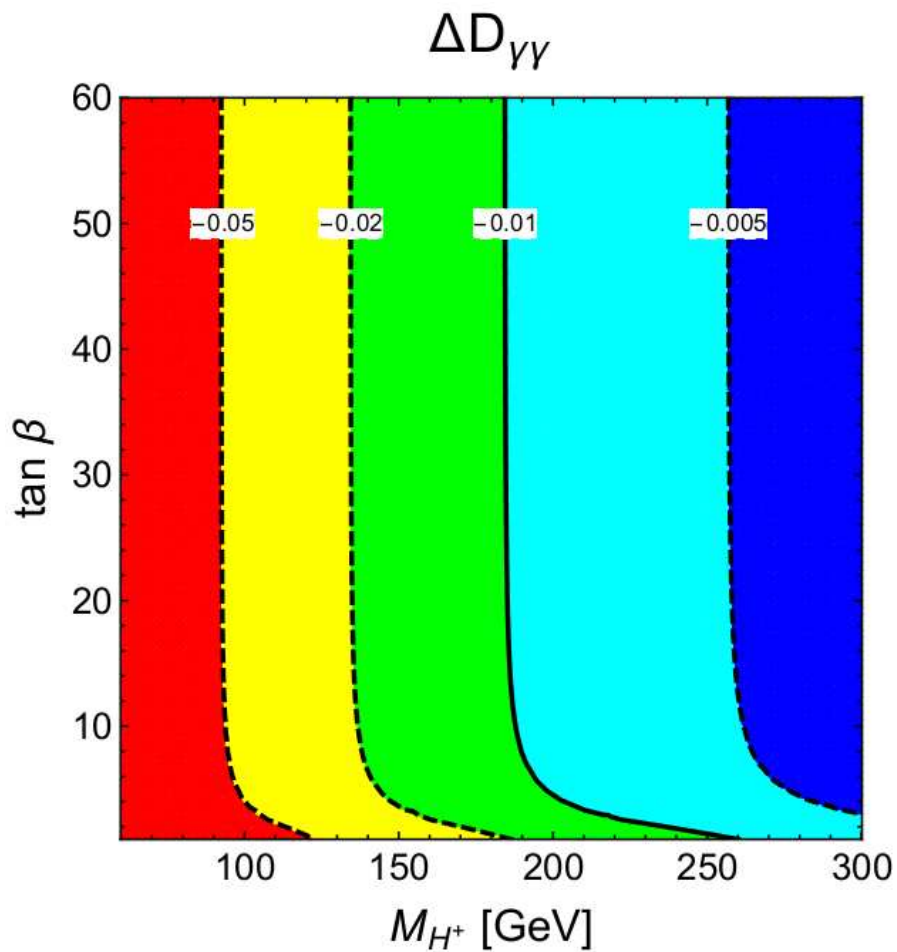
$$c_t/c_V = [1 - (1 + n)\xi]/((1 - \xi)), \quad \tilde{c}_t = c_{\gamma\gamma} = \tilde{c}_{\gamma\gamma} = 0$$

$$c_t/c_V = (1 + \gamma_t), \quad c_{\gamma\gamma}/c_V = \alpha/(4\pi)(b_{\text{IR}}^{\text{EM}} - b_{\text{UV}}^{\text{EM}}), \quad \tilde{c}_t = \tilde{c}_{\gamma\gamma} = 0,$$



Search for BSM with $D_{\gamma\gamma}$

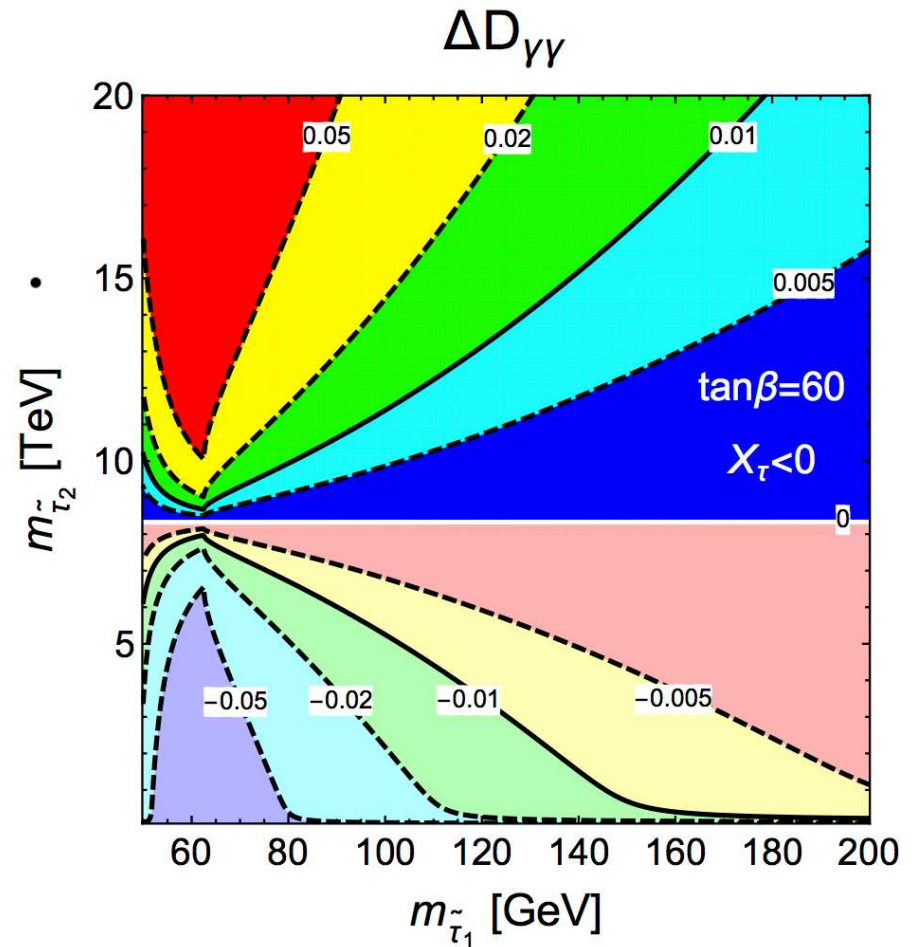
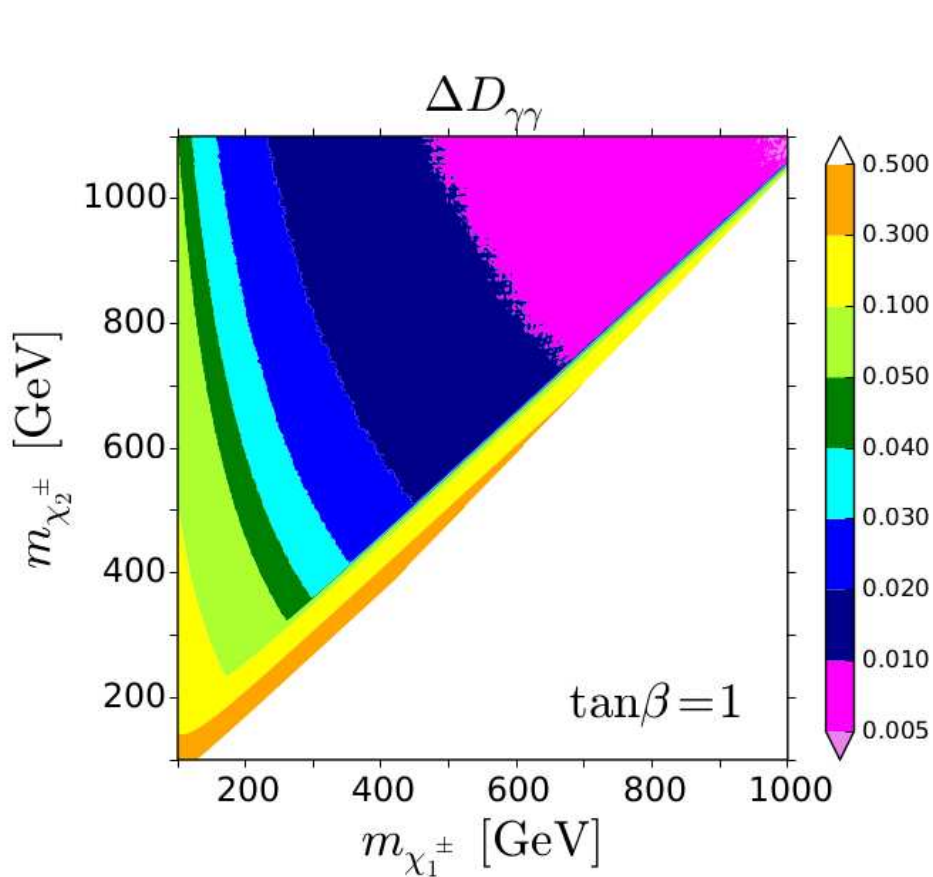
(h)MSSM and 2HDM: charged Higgs contributions



Search for BSM with $D_{\gamma\gamma}$

MSSM: chargino and stau contributions

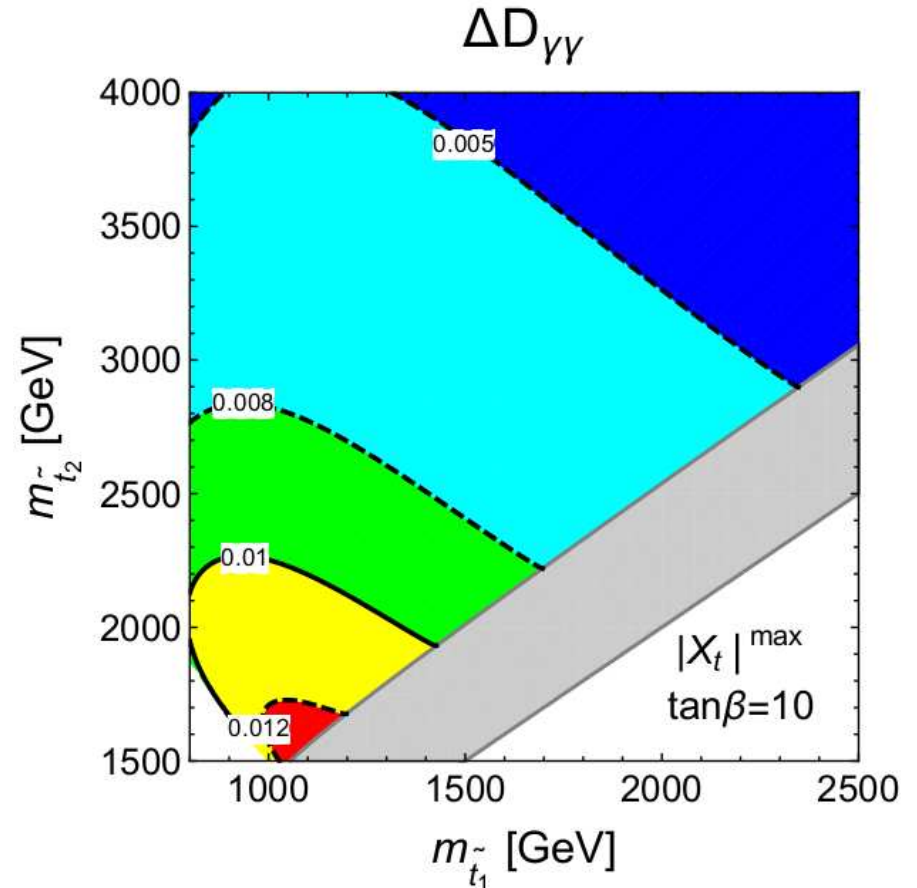
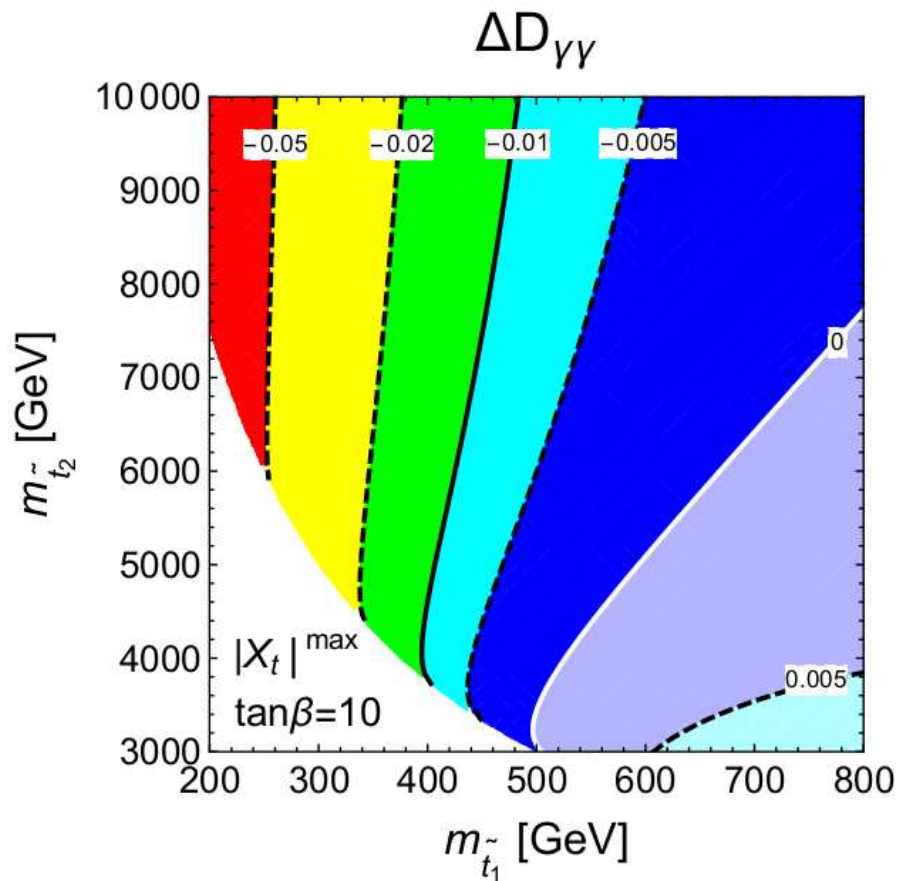
(see also AD, Driesen, Hollik, Illana (Karlsruhe U.), hep-ph/9612362)



Search for BSM with $D_{\gamma\gamma}$

(h)MSSM: stop contributions

(see also AD, Driesen, Hollik, Illana (Karlsruhe U.), hep-ph/9612362)



Search for BSM with $D_{\gamma\gamma}$

Vector-like quarks: $Q_{\text{VLQ}} = +2/3, -4/3, +5/3$

Angelescu, AD, Moreau, to appear.

