# **MDI Status and Plans**

Michele Modena and Lau Gatignon on behalf of the MDI working group



# Outline

- MDI (recall of CDR and previous status)
- Studies done and ongoing
  - Studies for longer L\*: impact on the various systems
  - 2. IP feedback
  - 3. Muon scrapers study
- Conclusion and future plans

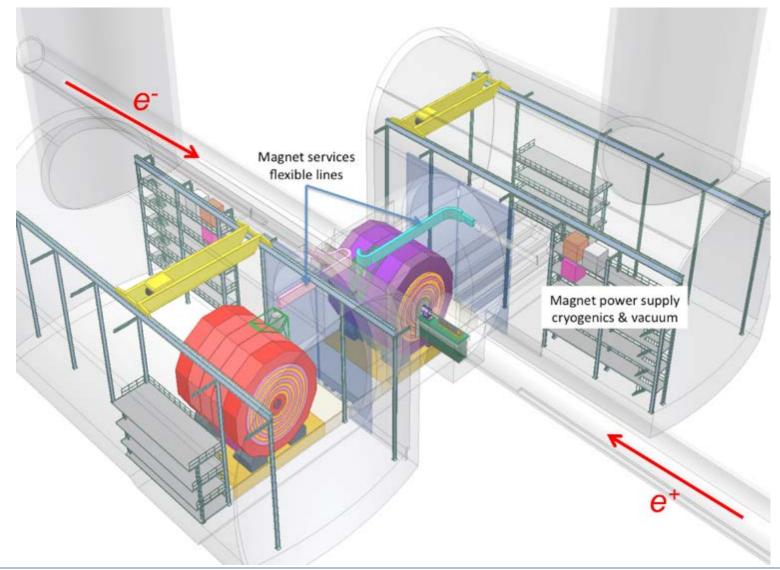


# What is MDI

- The Machine Detector Interface must ensure optimum luminosity for the experiment(s) with minimal backgrounds. It includes the integration of all systems and infrastructure.
- The baseline for the CDR was based on a concept with <u>two</u> <u>detectors</u> operating in <u>push-pull mode</u> and with the final focus quadrupoles QD0 as close as possible to the interaction point ( $L^* = 3.5 \text{ m}$ , i.e. INSIDE the detectors).
- The MDI design and studies include the studies for the QD0 design as well as its stabilisation and pre-alignment, but also IP feedback, BeamCal and Lumical integration, vacuum layout, cavern layout, post-collision line systems etc.



# The CDR MDI concept:





## **CDR Detectors Concepts**

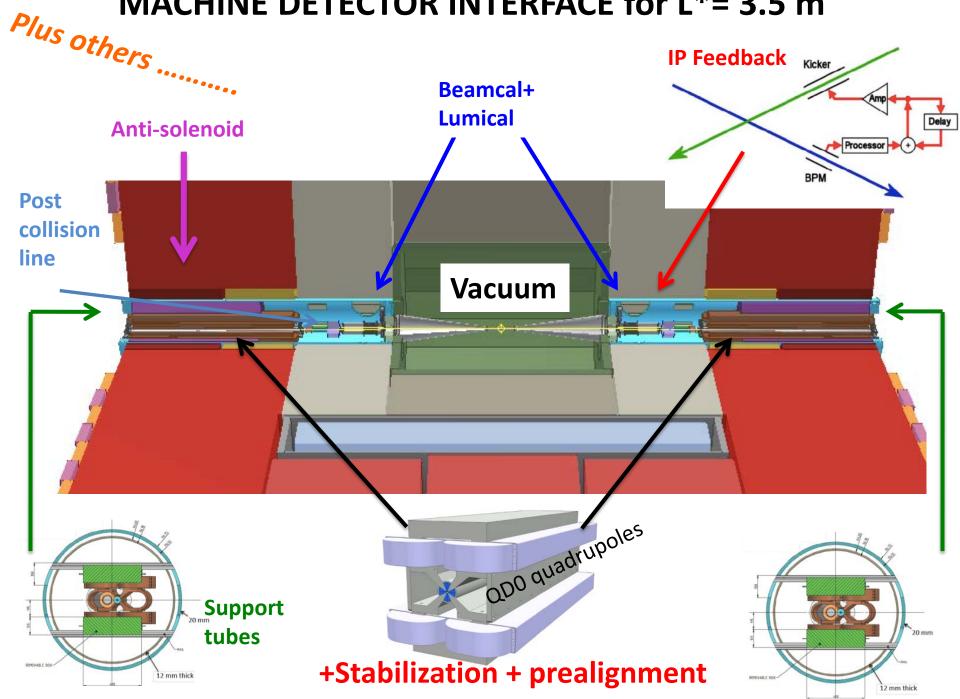


SiD: 5 Tesla field;  $L^* = 4.4 \text{ m}$ 

ILD: 4 Tesla field;  $L^* = 3.5 \text{ m}$ 



#### **MACHINE DETECTOR INTERFACE for L\*= 3.5 m**



### **1. Studies for longer L\*:**

Some justifications for the CDR choice (L\*=3.5 m)

The choice of short L\* was justified by:

- this option would provide the maximum (peak) luminosity
- this layout is the most challenging: if you have a plausible solution for short L\*, the longer L\* should be easier for the stabilisation, radiation, impact of detector solenoid B-field, etc.
- at the time the pre-alignment tolerance for longer L\* was considered unrealistic (2  $\mu$ m for L\*=8 m, 10  $\mu$ m for L\*= 3.5m), but since then significant progress has been made in the BDS optics.



### 1. Studies for longer L\*: PROS & CONS

#### Pros:

- <u>Maximize detecting volume</u> (forward acceptance)
- <u>Less complex integration</u>
   (QD0, stabilization system
   integration, alignment
   concept, vacuum systems,
   etc.)
- <u>No need of an antisolenoid</u> (at least for QD0 operation)

Cons:

- Lower peak luminosity

(see F. Plassard presentation)

- Impact on Beam Delivery System (BDS)

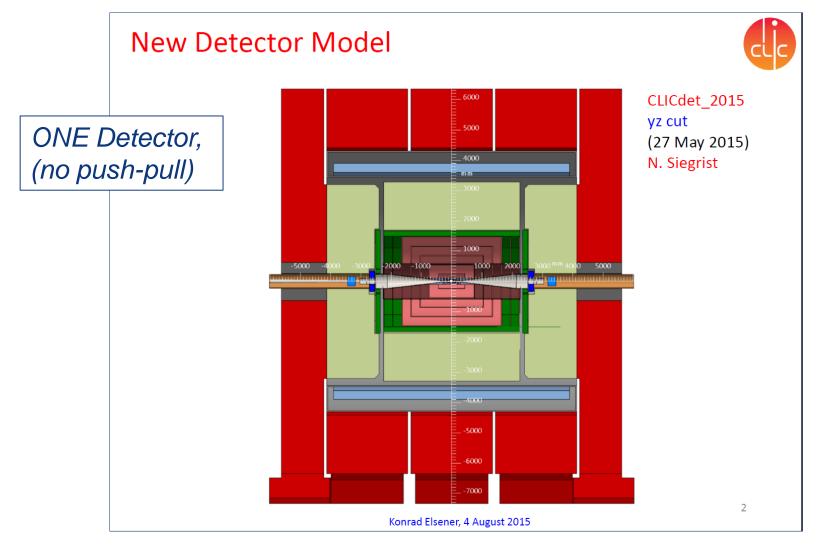
(see F. Plassard presentation)

- Alignment requirements are <u>tighter</u> (more precise evaluations are on-going)



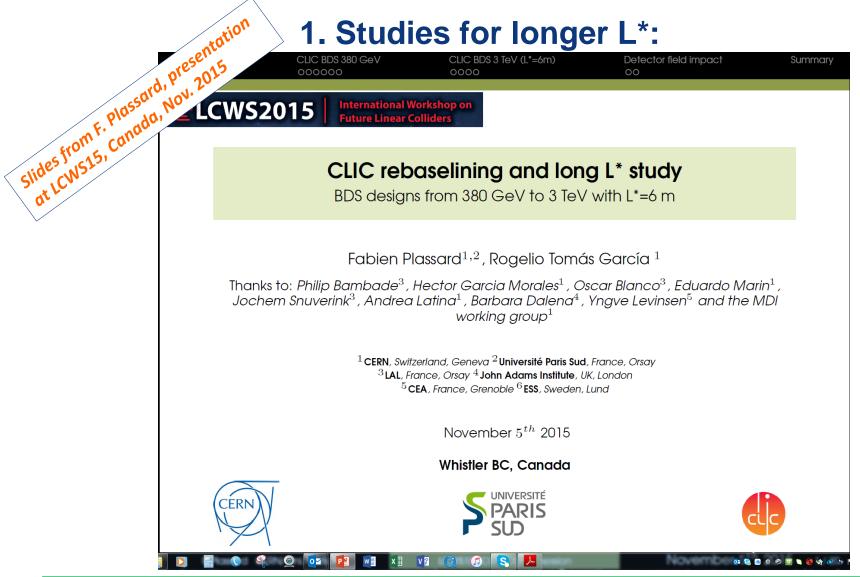
## 1. Studies for longer L\*:

Detector new conceptual design





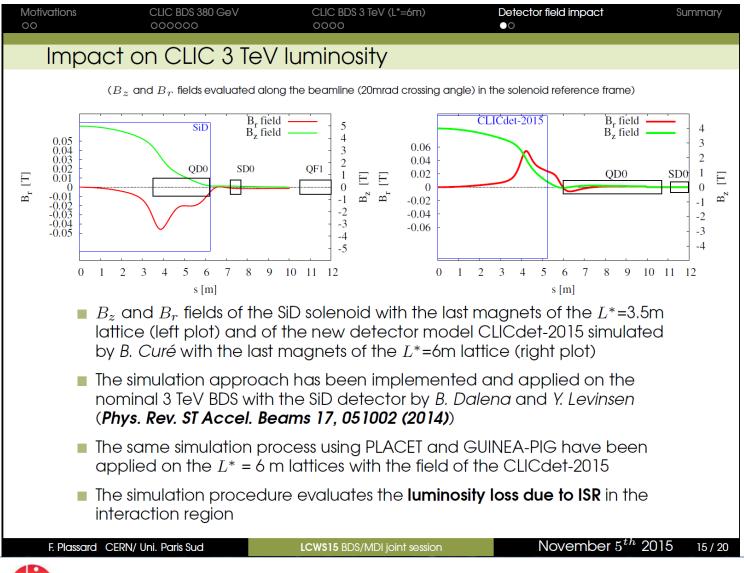




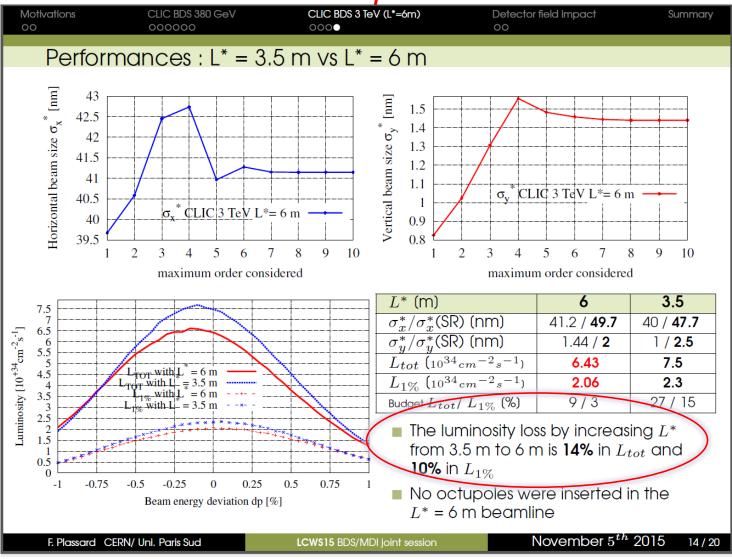
Please refer to F. Plassard presentation: "Rebaselining and longer L\* for CLIC and ATF2" AT THIS WORKSHOP, Accelerator Parallel Sessions, Tuesday 19 at 14h20



### 1. Studies for longer L\*: BDS Implication



#### 1. Studies for longer L\*: BDS Implication





### 1. Studies for longer L\*: BDS Implication

ations CLIC BDS 380 GeV 000000	study	COVERING a		ctor field impact	Summary
ummary	energy staging case (380 GeV)				
CLIC		380 GeV	380 GeV		3 TeV
L* (m)		4.3	6	3.5	6
$\sigma_x^*$ (SR) (nm)		150	160	47.7	49.7
$\sigma_y^*$ (SR) (nm)		2.7	3.5	2.5	2
$L_{tot}$ (design) / $L_{tot}$ (10 <sup>34</sup> cm <sup>-2</sup> s <sup>-1</sup> )		1.5 / <b>1.86</b>	1.5 / <b>1.5</b>	2 5.9 / 7.5	5.9 / <b>6.43</b>
$L_{1\%}$ (design) / $L_{1\%}$ (10 $^{34}cm^{-2}s^{-1}$ )		0.9 / <b>1.09</b>	0.9 / 0.94	2 / 2.3	2 / <b>2.06</b>
Chromaticity $\xi_y$ (computed)		68464	95697	82637	93017
Budget $L_{tot}/L_{1\%}$ (%)		24 / 21	1.5 / 4.5	27 / 15	9/3
Impact of solenoid on $L_{tot}/L_{1\%}$ (%)		-	-	7.8 / 8.2	3.7 / 4.6
Tuning performances		-	-	-	-
<ul> <li>All lattices fulfill now the a</li> <li>For L*= 6m option for ea dynamic imperfections is</li> <li>The impact of the solence and should not require a</li> </ul>	ch stag s low pid on th	e, the lumino	osity budge	et for static a	
The tuning is still on progr FFS (Tradition or Local sch Plassard CERN/ Uni. Paris Sud	neme?		L* ?)		"Work in pro
				perjormane	



IMP!: "Work in progress"; performances could be probably even improved but FEEDBACKS and MOTIVATIONS from Detector Community are NEEDED!

### **1. Studies for longer L\*:** *Magnet system implication*

• The QD0 requirement for  $E= 3\text{TeV} / L^* = 3.5 \text{ m}$  are:

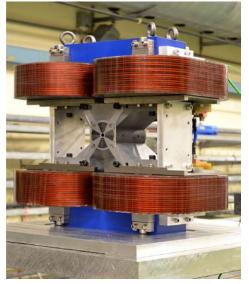
Gradient: 575 T/m; Aperture Ø: 8.25 mm, Length: 2.73 m, tunability: 20%

• The (*preliminary*) requirements for <u>E=3TeV / L\*=6 m</u> are:

Gradient: 197 T/m; Aperture Ø: 10 mm, Length: 4.7 m (eventually split in 2-3 elements)

The QD0 parameters for L\*=6 *m* are evidently <u>more relaxed</u>. Furthermore, the magnet would be positioned OUTSIDE the detector.

The hybrid design developed for the  $L^*=3.5$  m case is a possible but maybe not necessary solution. To be reminded that the magnet still need to be nanometer stabilized and has to be compatible with the passage of the post-collision line (chamber at ~ 60 mm in transverse direction of QD0 axis)

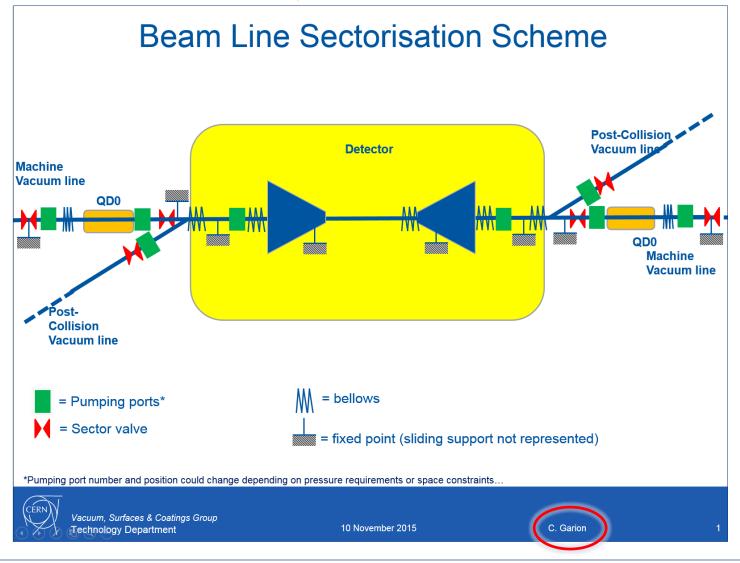


Hybrid QD0 prototype developed with the L\*=3.5 m main parameters



#### **1. Studies for longer L\*:**

Vacuum system implication





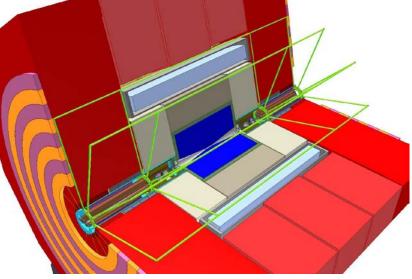
#### **1. Studies for longer L\*:** *Alignment system implication*

• Tolerances will be tighter: for CDR requirements were evaluated at  $10\mu m$  for L\* = 3.5 m.

For L\*=6 m will be ~6-8  $\mu$ m (?)  $\rightarrow$  study ongoing by beam dynamic team.

 The system would be <u>simpler</u> (no needs of the "ZERODUR" spokes system as for L\*=3.5 m

(ZERODUR® has a thermal expansion coefficient of  $0\pm0.007\times10^{-6}/K$ in the range 0°to 50°C)

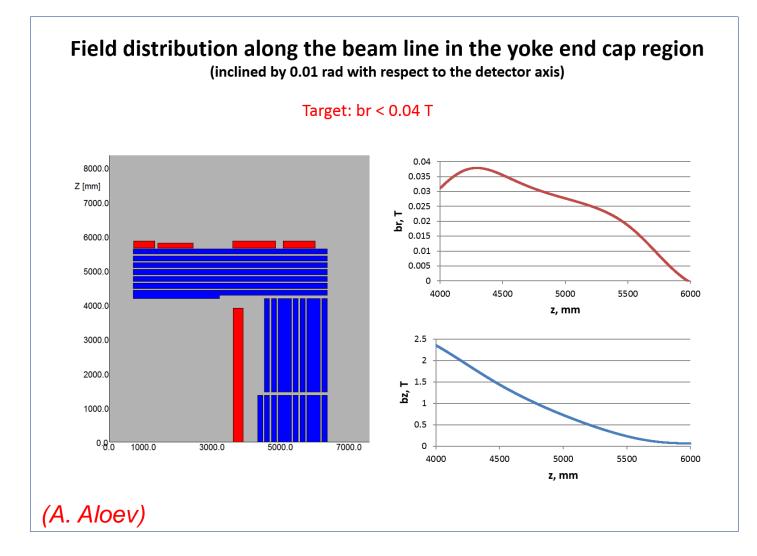


Needs of <u>"survey mini-galleries</u>" bypassing the cavern ?



### 1. Studies for longer L\*:

#### Antisolenoid system implication

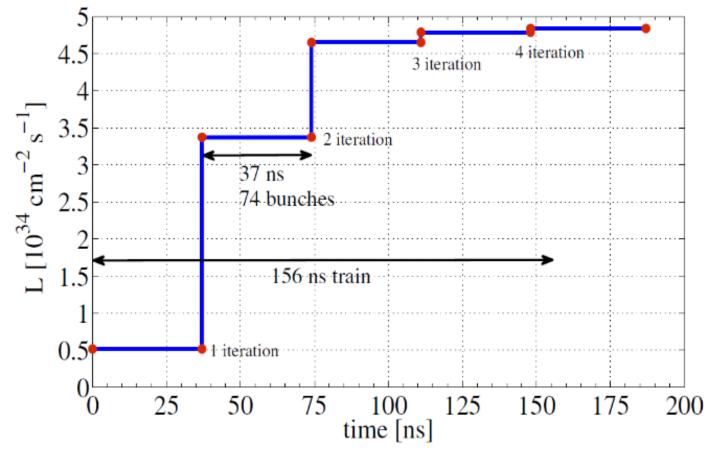




#### 2. IP Feedback study

### CLIC IP FB Performance (CDR)

Single random seed of GM C, CDR implementation



(Ph. Burrows, Resta-Lopez)



#### 2. IP Feedback study

### IP FB with long L\*

- Current CDR geometry: time of flight IP → BPM → kicker → IP ~ 24 ns
- Demonstrated FONT3 electronics latency = 13ns
- Estimated IPFB latency = 37ns
- In principle, change of L\* need not affect IPFB position and latency, but needs to be engineered carefully, considering other beam line components

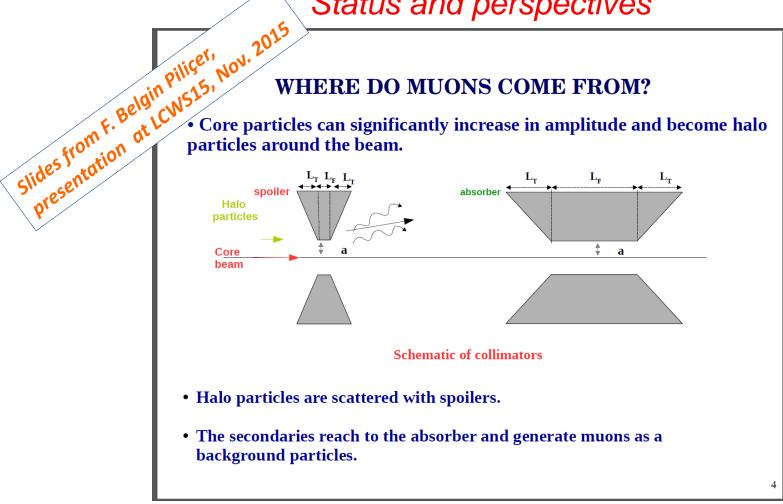


### 3. Muons scrapers studies:

Status and perspectives

#### WHERE DO MUONS COME FROM?

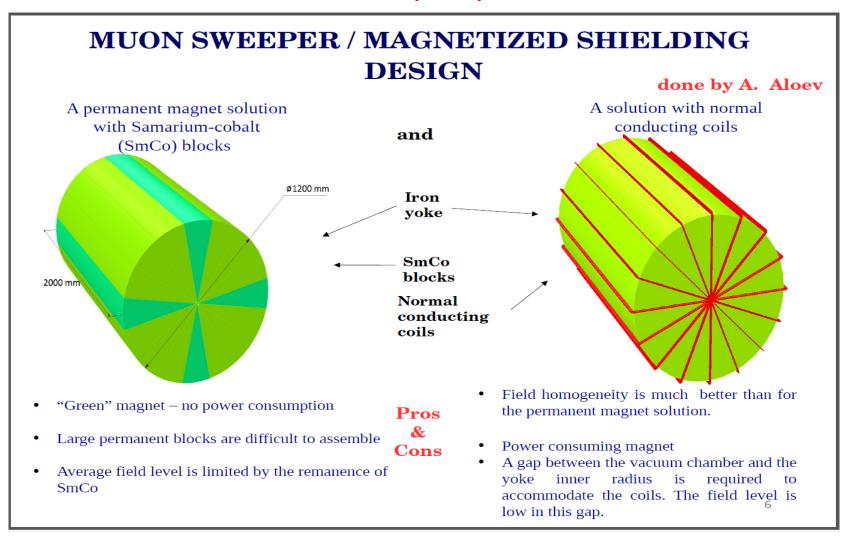
Core particles can significantly increase in amplitude and become halo particles around the beam.



Please refer to A. Aloev presentation: "Magnetized muon absorbers" AT THIS WORKSHOP, Accelerator Parallel Sessions, Tuesday 19 at 15h40

3. Muons scrapers studies:

Status and perspectives





### 3. Muons scrapers studies:

Status and perspectives

#### **SUMMARY**

 $\rightarrow$  Permanent magnet solution has been compared with normal conducting coils.

 $\rightarrow$  Field intensity has been simulated for different # of SmCo blocks.

 $\rightarrow$  The muon sweeper parameters have been updated for BDS.

 $\rightarrow$  0.7 T (min) and 1.2 T (optimum) as permanent magnet option have been simulated with BDSIM.

 $\rightarrow$  The simulation results showed roughly factor of ~10 reduction at 1.2 T for muons at the end of the BDS.

 $\rightarrow$  The remaining muons comes dominantly from last dipole section.

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#### 4. Conclusion and Future Plans:

- <u>COMPLETION of BDS OPTIMIZATION and PERFORMANCES evaluation</u>: work on-going with the beam dynamic team. We need operative requirements for stabilization and prealignment tolerances.
- <u>MAGNET STUDY</u>: The required gradient for QD0 in L\*=6 m permit to envisage different solutions, driving aspects link with the following points:
  - <u>QD0 STABILIZATION</u>: one of the most critical aspect of the new layout. QD0 will be longer (4.7 m) but can be split in 2-3 elements (each length ~ 1.5 m; → always try to minimize the QD0 mass; → correlation/matching of the 2-3 stabilizing systems)
  - REQUIREMENTS FOR QD0 PRE-ALIGNMENT: the requirements are tighter respect to L\*=3.5 m, but QD0 is outside of the detector → a new approach must be study (survey mini-galleries?)
- FEEDBACK and INPUT from the Detector Community are needed in order to advance with the study and for eventual improvement of final performances



# The MDI working group

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### Thanks for your attention

