

Quantum gravity using SUSY as a formal device

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Talk prepared for “Supersymmetry: from M-theory to the LHC”,
University of Kent, Monday 11th January 2016

Introduction

- Einstein's General Relativity (GR) is a **perturbatively** non-renormalizable field theory.
- The field is the spacetime metric.
- The asymptotic safety conjecture suggests that there may exist a **non-perturbative**, non-trivial ultraviolet fixed point for gravity.
- Renormalization Group (RG) flow can be seen intuitively as describing physics at different scales of length by changing the resolution.
- SUSY is not just a theory of BSM particle physics, it can also be used as a more general mathematical device.
- Existing examples are **manifestly gauge invariant regularization** in QFT and **Parisi-Sourlas supersymmetry** in statistical field theory.
- My work develops a **manifestly diffeomorphism invariant** Exact RG, the regularization of which would also use this formal kind of SUSY.

Kadanoff blocking

In the Ising model, Kadanoff blocking is the process of **grouping microscopic spins** together to form **macroscopic “blocked” spins** via a majority rule.

The continuous version integrates out the high-energy modes of a field to give a renormalized field, used in a renormalized action.

The **blocking functional**, b , is defined via the effective Boltzmann factor:

$$e^{-S[\varphi]} = \int \mathcal{D}\varphi_0 \delta[\varphi - b[\varphi_0]] e^{-S_{\text{bare}}[\varphi_0]}.$$

There are an **infinite number** of possible Kadanoff blockings, but a simple linear example is

$$b[\varphi_0](x) = \int_y B(x - y) \varphi_0(y), \quad \text{where the kernel, } B, \text{ contains an } \textbf{infrared cutoff function}.$$

The partition function must be **invariant under change of cutoff scale**, Λ , this will be ensured by construction, i.e.

$$\mathcal{Z} = \int \mathcal{D}\varphi e^{-S[\varphi]} = \int \mathcal{D}\varphi_0 e^{-S_{\text{bare}}[\varphi_0]}.$$

Kadanoff blocking demands a suitable notion of **locality** that requires us to work exclusively in **Euclidean signature**.

Polchinski Flow Equation

Differentiate the effective Boltzmann factor w.r.t. “RG time”:

$$\Lambda \frac{\partial}{\partial \Lambda} e^{-S[\varphi]} = - \int_x \frac{\delta}{\delta \varphi(x)} \int \mathcal{D}\varphi_0 \delta[\varphi - b[\varphi_0]] \Lambda \frac{\partial b[\varphi_0](x)}{\partial \Lambda} e^{-S_{\text{bare}}[\varphi_0]}.$$

Write the “rate of change of blocking functional” as

$$\Psi(x) = \frac{1}{2} \int_y \dot{\Delta}(x, y) \frac{\delta \Sigma}{\delta \varphi(y)}, \quad \text{where } \Sigma = S - 2\hat{S},$$

and $\Delta = c(p^2/\Lambda^2)p^{-2}.$

The RG flow of the effective Boltzmann factor is then

$$\Lambda \frac{\partial}{\partial \Lambda} e^{-S[\varphi]} = \int_x \frac{\delta}{\delta \varphi(x)} \left(\Psi(x) e^{-S[\varphi]} \right).$$

The Polchinski flow equation is then written elegantly as

$$\dot{S} = \frac{1}{2} \frac{\delta S}{\delta \varphi} \cdot \dot{\Delta} \cdot \frac{\delta \Sigma}{\delta \varphi} - \frac{1}{2} \frac{\delta}{\delta \varphi} \cdot \dot{\Delta} \cdot \frac{\delta \Sigma}{\delta \varphi}.$$

where $f \cdot W \cdot g := \int_x f(x) W \left(-\frac{\partial^2}{\Lambda^2} \right) g(x).$

Generalization to Yang-Mills theories

A pure gauge theory is constructed in a manifestly gauge invariant way by building the action from **covariant derivatives** and **field-strength tensors**:

$$D_\mu := \partial_\mu - iA_\mu, \quad F_{\mu\nu} := i[D_\mu, D_\nu].$$

To preserve manifest gauge invariance, wavefunction renormalization must be avoided. This is achieved by writing the coupling, g , as an overall scaling factor:

$$S[A](g) = \frac{1}{4g^2} \text{tr} \int_x F_{\mu\nu} c^{-1} \left(-\frac{D^2}{\Lambda^2} \right) F_{\mu\nu} + \dots$$

The effective action is then written as a **loopwise expansion** that is also an expansion in powers of g :

$$S = \frac{1}{g^2} S_0 + S_1 + g^2 S_2 + \dots$$

The **β -functions** also have a similar expansion: $\beta := \Lambda \partial_\Lambda g = \beta_1 g^3 + \beta_2 g^5 + \dots$

The generalization of the Polchinski flow equation to gauge theories uses a suitably **covariantized kernel**, the easiest way to do this is to replace the partial derivatives with covariant derivatives:

$$\dot{S} = \frac{1}{2} \frac{\delta S}{\delta A_\mu} \cdot \{\dot{\Delta}\} \cdot \frac{\delta \Sigma}{\delta A_\mu} - \frac{1}{2} \frac{\delta}{\delta A_\mu} \cdot \{\dot{\Delta}\} \cdot \frac{\delta \Sigma}{\delta A_\mu}.$$

A refresher on GR

The **Riemann tensor**, representing spacetime curvature, is written in our **sign convention** as

$$R^a_{bcd} = 2\partial_{[c}\Gamma^a_{b]d} + 2\Gamma^a_{[c|f}\Gamma^f_{|b]d}$$

where $A_{[ab]} = \frac{1}{2}(A_{ab} - A_{ba})$

We use the **Levi-Civita connection** (torsion-free metric connection)

$$\Gamma^a_{bc} = \frac{g^{ad}}{2} (\partial_b g_{cd} + \partial_c g_{bd} - \partial_d g_{bc})$$

We have the **Ricci tensor** in our sign convention as $R_{ab} = R^c_{acb}$

Then the **Ricci scalar** by $R = g^{ab} R_{ab}$

Thus the Einstein field equation is

$$R_{ab} - \frac{g_{ab}}{2} R + \Lambda g_{ab} = 8\pi G T_{ab}$$

Diffeomorphism invariance

Consider a general coordinate transformation

$$x'^{\mu} = x^{\mu} - \xi^{\mu}(x)$$

We need our theories to be **diffeomorphism invariant**. This is a surprisingly tough constraint on what we can use.

For some covariant derivative, D , metrics transform as

$$g'_{\mu\nu}(x') = \frac{\partial x^{\rho}}{\partial x'^{\mu}} \frac{\partial x^{\sigma}}{\partial x'^{\nu}} g_{\rho\sigma}(x) = g_{\mu\nu}(x) + 2g_{\lambda(\alpha} D_{\beta)} \xi^{\lambda} + \xi \cdot D g_{\alpha\beta} .$$

So metric perturbations transform as

$$h_{\mu\nu}(x) \rightarrow h_{\mu\nu}(x) + 2g_{\lambda(\alpha} D_{\beta)} \xi^{\lambda} + \xi \cdot D g_{\alpha\beta} .$$

An advantage of manifest diffeomorphism invariance is that we automatically know that our results are not artefacts of some chosen coordinate system.

Generalization to gravity

The generalization of the **Polchinski flow equation** to gravity is

$$\dot{S} = \int_x \frac{\delta S}{\delta g_{\mu\nu}(x)} \int_y K_{\mu\nu\rho\sigma}(x, y) \frac{\delta \Sigma}{\delta g_{\rho\sigma}(y)} - \int_x \frac{\delta}{\delta g_{\mu\nu}(x)} \int_y K_{\mu\nu\rho\sigma}(x, y) \frac{\delta \Sigma}{\delta g_{\rho\sigma}(y)},$$

The **kernel**, which transforms as a two-argument generalization of a tensor, is

$$K_{\mu\nu\rho\sigma}(x, y) = \frac{1}{\sqrt{g}} \delta(x - y) (g_{\mu(\rho} g_{\sigma)\nu} + j g_{\mu\nu} g_{\rho\sigma}) \dot{\Delta}.$$

The **de Witt supermetric** parameter, j , determines the balance of modes propagating in the flow equation. For the “kinetic term” to be a regularized **Einstein-Hilbert** form, $j = -1/2$.

The **effective action** then goes like $S = \frac{1}{16\pi G} \int_x \sqrt{g} \left(-R + R_{\mu\nu} \frac{d}{\Lambda^2} R^{\mu\nu} - \frac{1}{2\Lambda^2} R \frac{d}{\Lambda^2} R + \dots \right),$

where d is a function of **covariant derivatives** that is related to the **inverse cutoff**.

Writing the action as a **loopwise expansion**: $S = \frac{1}{\tilde{\kappa}} S_0 + S_1 + \tilde{\kappa} S_2 + \tilde{\kappa}^2 S_3 \dots$

The β -function also expands as $\beta := \Lambda \partial_\Lambda \tilde{\kappa} = \beta_1 + \beta_2 \tilde{\kappa} + \beta_3 \tilde{\kappa}^2 + \dots$

For this action, we have $\dot{\Delta} = -\frac{2}{\Lambda^2} c' \left(-\frac{\nabla^2}{\Lambda^2} \right),$

which is related to the “**effective propagator**”, Δ , in the fixed-background description.

Fixed-background description

If we fix a **Euclidean** background metric, we can define the **graviton field** as the **perturbation** to that background:

$$h_{\mu\nu}(x) := g_{\mu\nu}(x) - \delta_{\mu\nu}.$$

The position representation is related to a momentum representation via a Fourier transform:

$$h_{\mu\nu}(x) = \int \mathrm{d}p \, e^{-ip \cdot x} h_{\mu\nu}(p), \quad \text{where} \quad \mathrm{d}p := \frac{d^D p}{(2\pi)^D}.$$

The action is defined as a series expansion in the perturbation:

$$\begin{aligned} S = & \int \mathrm{d}p \, \delta(p) \mathcal{S}^{\mu\nu}(p) h_{\mu\nu}(p) + \frac{1}{2} \int \mathrm{d}p \, \mathrm{d}q \, \delta(p+q) \mathcal{S}^{\mu\nu\rho\sigma}(p, q) h_{\mu\nu}(p) h_{\rho\sigma}(q) \\ & + \frac{1}{3!} \int \mathrm{d}p \, \mathrm{d}q \, \mathrm{d}r \, \delta(p+q+r) \mathcal{S}^{\mu\nu\rho\sigma\alpha\beta}(p, q, r) h_{\mu\nu}(p) h_{\rho\sigma}(q) h_{\alpha\beta}(r) + \dots \end{aligned}$$

In this picture, we are able to define an “effective propagator”, $\Delta := \frac{1}{p^2} c \left(\frac{p^2}{\Lambda^2} \right)$.

Gravity does not have a unique transverse 2-point function because there are two linearly independent **transverse projectors**. If we choose the linear combination with **Einstein-Hilbert structure**, the 2-point function is simply that transverse projector times Δ . This then **solves the flow equation** at the 2-point level.

Parisi-Sourlas formalism (SUSY in condensed matter)

Consider a D-dimensional **system of spins** with a **random external field**, **h** . We can write the free energy, using “natural” units of $k_B T=1$, as

$$F[h] = \ln \int \mathcal{D}\phi \exp \left\{ - \int d^D x [\mathcal{L}(x) + h(x)\phi(x)] \right\}$$

where the Lagrangian density is

$$\mathcal{L}(x) = -\frac{1}{2} \phi(x) \partial^2 \phi(x) + V(\phi(x))$$

The **averaged** 2-point function goes like

$$\langle \phi(x) \phi(0) \rangle_R \sim \int \mathcal{D}h \phi_h(x) \phi_h(0) \exp \left(-\frac{1}{2} \int d^D y h^2(y) \right)$$

where ϕ_h is the solution to the **field equation**

$$-\partial^2 \phi + V'(\phi) + h = 0$$

Equivalence to SUSY

A procedure similar to Fadeev-Popov ghosts rewrites the 2-point function to

$$\langle \phi(x)\phi(0) \rangle \sim \int \mathcal{D}\phi \mathcal{D}\omega \mathcal{D}\psi \mathcal{D}\bar{\psi} \exp \left[- \int d^D y \mathcal{L}_R(y) \right] \phi(x)\phi(0)$$

where ψ is an **anti-commuting (ghost)** scalar field and the **averaged Lagrangian** is

$$\mathcal{L}_R = -\frac{1}{2}\omega^2 + \omega[-\partial^2 \phi + V'(\phi)] + \bar{\psi}[-\partial^2 + V''(\phi)]\psi$$

which is **supersymmetric** under

$$\begin{aligned} \delta\phi &= -\bar{a}\epsilon_\mu x_\mu \psi, & \delta\omega &= 2\bar{a}\epsilon_\mu \partial_\mu \psi, \\ \delta\psi &= 0, & \delta\bar{\psi} &= \bar{a}(\epsilon_\mu x_\mu \omega + 2\epsilon_\mu \partial_\mu \phi), \end{aligned}$$

where \bar{a} is an **infinitesimal, anti-commuting** number and ϵ_μ is an **arbitrary vector**.

Dimensional reduction

The action, written in a manifestly supersymmetric form, is

$$S = \int d^D x d\bar{\theta} d\theta \left(-\frac{1}{2} \Phi \left[\partial^2 + \frac{\partial^2}{\partial \bar{\theta} \partial \theta} \right] \Phi + V(\Phi) \right)$$

where θ is a **Grassmann number** and Φ is the **superfield**

$$\Phi(x, \theta) = \phi(x) + \bar{\theta} \psi(x) + \bar{\psi}(x) \theta + \theta \bar{\theta} \omega(x)$$

This system is equivalent to a non-supersymmetric system with dimension D-2 via the relation

$$\int d^{D-2} x f(x^2) = \int d^D x d\theta d\bar{\theta} f(x^2 + \bar{\theta} \theta)$$

Green's functions at **all orders** of perturbation theory for the **non-SUSY D-2** system are the same as for the **SUSY D-dimensional** system. Thus a **random external field** effectively **reduces the dimension** by 2.

SUSY in gauge and gravity ERG

The manifestly gauge invariant ERG for $SU(N)$ requires [additional regularization](#) at 1-loop level using covariant Pauli-Villars fields. The elegant way to do this is with [SU\(N|N\) regularization](#).

The field is promoted to a [supermatrix](#) of bosonic components, A , and fermionic components, B :

$$\mathcal{A}_\mu = \begin{pmatrix} A_\mu^1 & B_\mu \\ \bar{B}_\mu & A_\mu^2 \end{pmatrix} + \mathcal{A}_\mu^0 \mathcal{I}, \text{ and } D_\mu = \partial_\mu - i\mathcal{A}_\mu.$$

The action is built in a similar way:

$$S = \frac{1}{4g^2} \text{str} \int \mathcal{F}_{\mu\nu} c^{-1} \left(-\frac{D^2}{\Lambda^2} \right) \mathcal{F}_{\mu\nu} + \dots$$

$$\text{where } \text{str} \begin{pmatrix} X^{11} & X^{12} \\ X^{21} & X^{22} \end{pmatrix} = \text{tr}_1 X^{11} - \text{tr}_2 X^{22}.$$

This supersymmetry is spontaneously broken by a [super-Higgs](#) mechanism with a mass at order Λ , so that the physical $SU(N)$ can be recovered at [low energy](#).

A similar procedure is required for the 1-loop manifestly diffeomorphism invariant ERG for gravity.

Summary

- Supersymmetry is broader than just a property of BSM physics.
- SUSY appears in statistical field theory in the [Parisi-Sourlas formalism](#) for spin-glasses, where it is related to dimensional reduction by 2.
- A related form of SUSY appears as a regularization mechanism in the manifestly [gauge](#) invariant ERG and is expected to be useful in the [diffeomorphism](#) invariant case also.