### QUARKONIUM 2016 Trento, March 1<sup>st</sup> 2016

# $J/\psi$ production in p-A collisions from SPS to LHC

Roberta Arnaldi INFN Torino

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### **Outlook:**

pA J/ $\psi$  results as:

- tool to understand cold nuclear matter effects
- reference for AA

from SPS to LHC experiments

Facility	Experiment	System	√s <sub>NN</sub> (GeV)	Y <sub>cms</sub> range	Data taking
	NA50	p-Be,Al,Cu,Ag,W,Pb	27	-0.4 <y<0.6< td=""><td>1996-</td></y<0.6<>	1996-
SPS	NA30		29	-0.5 <y<0.5< td=""><td>2000</td></y<0.5<>	2000
373	NA60 p-Be,Al,Cu,In,W,Pb,U	17	0.3 <y<0.8< td=""><td>2004</td></y<0.8<>	2004	
	NAGO	p-Be,Al,Cu,In,W,Pb,U	27	-0.1 <y<0.3< td=""><td>2004</td></y<0.3<>	2004
FNAL	E866	p-Be, Fe, W	39	-0.6 <y<2.5*< td=""><td>~1996</td></y<2.5*<>	~1996
HERA	HERA-B	p-C, Ti, W	42	-1.5 <y<0.8*< td=""><td>2002</td></y<0.8*<>	2002
RHIC	PHENIX,	d-Au	200	-2.2 <y<2.4< td=""><td>&gt;2003</td></y<2.4<>	>2003
ппс	STAR	p-Al, Au	200	1.2< y <2.2	2015
	ALICE			-4.46 <y<3.53< td=""><td></td></y<3.53<>	
	ATLAS	p-Pb	5020	-2.87 <y<1.94< td=""><td rowspan="3">2013</td></y<1.94<>	2013
LHC	CMS			-2.87 <y<1.93< td=""></y<1.93<>	
	LHCb			-5.0 <y<-2.5< td=""></y<-2.5<>	
				1.5 <y<4.0< td=""><td></td></y<4.0<>	

Facility	Experiment	System	√s <sub>NN</sub> (GeV)	Y <sub>cms</sub> range	Data taking
	NA50	p-Be,Al,Cu,Ag,W,Pb	27 29	-0.4<γ<0.6 -0.5<γ<0.5	1996- 2000
SPS	NA60	p-Be,Al,Cu,In,W,Pb,U	17 27	0.3 <y<0.8 -0.1<y<0.3< td=""><td>2004</td></y<0.3<></y<0.8 	2004
FNAL	E866	p-Be, Fe, W	39	-0.6 <y<2.5< td=""><td>~1996</td></y<2.5<>	~1996
HERA	HERA-B	p-C, Ti, W	4.2	-1.5 <y<0.8< td=""><td>2002</td></y<0.8<>	2002
RHIC	PHENIX, STAR	d-Au p-Al, Au		-2.2 <v<2.4 ed target exper</v<2.4 	
	ALICE			several A target	S
	ATLAS	p-Pb	5020	-2.87 <y<1.94< td=""><td rowspan="2">2013</td></y<1.94<>	2013
LHC	CMS			-2.87 <y<1.93< td=""></y<1.93<>	
	LHCb			-5.0 <y<-2.5 1.5<y<4.0< td=""><td></td></y<4.0<></y<-2.5 	
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Facility	Experiment	System	√s <sub>NN</sub> (GeV)	Y <sub>cms</sub> range	Data taking
	NA50	p-Be,Al,Cu,Ag,W,Pb	27	-0.4 <y<0.6< td=""><td>1996- 2000</td></y<0.6<>	1996- 2000
SPS			29	-0.5 <y<0.5< td=""><td>2000</td></y<0.5<>	2000
5, 5	NA60	p-Be,Al,Cu,In,W,Pb,U	17	0.3 <y<0.8< td=""><td>2004</td></y<0.8<>	2004
	NACO	p-be,Al,Cu,III,W,Pb,O	27	-0.1 <y<0.3< td=""><td>2004</td></y<0.3<>	2004
FNAL	E866	p-Be, Fe, W	39	-0.23 <y<2.5< td=""><td>~1996</td></y<2.5<>	~1996
HERA	HERA-B	p-C, Ti, W	4 <mark>Co</mark>	llider experimer usually p vs a si	
RHIC	PHENIX,	d-Au	20	beam specie	ingle
RHIC	STAR	p-Al, Au	20	forward and	
	ALICE			backward y rang	ge
	ATLAS			might be covere	ed
LHC	CMS	p-Pb	5020	-2.87 <y<1.93< td=""><td>2013</td></y<1.93<>	2013
	LHCb			-5.0 <y<-2.5 1.5<y<4.0< td=""><td></td></y<4.0<></y<-2.5 	
Roberta Arnaldi Quarkonium 201		6	March 1 <sup>st</sup>	2016	

Facility	Experiment	System	√s <sub>NN</sub> (GeV)	Y <sub>cms</sub> range	Data taking
SPS	NA50	p-Be,Al,Cu,Ag,W,Pb	27 29	Large number target nuclei	of <mark>,</mark>
	NA60	p-Be,Al,Cu,In,W,Pb,U	17 27	0.3 <y<0.8 -0.1<y<0.3< td=""><td>2004</td></y<0.3<></y<0.8 	2004
FNAL	E866	p-Be, Fe, W	39	-0.6 <y<2.5< td=""><td>~1996</td></y<2.5<>	~1996
HERA	HERA-B	p-C, Ti, W	42	-1.5 <y<0.8< td=""><td>2002</td></y<0.8<>	2002
RHIC	PHENIX,	d-Au	200	-2.2 <y<2.4< td=""><td>&gt;2003</td></y<2.4<>	>2003
ппс	STAR	p-Al, Au	200	1.2< y <2.2	2015
	ALICE			-4.46 <y<3.53< td=""><td></td></y<3.53<>	
	ATLAS			-2.87 <y<1.94< td=""><td rowspan="2">2013</td></y<1.94<>	2013
LHC	CMS	p-Pb	5020	-2.87 <y<1.93< td=""></y<1.93<>	
	LHCb			-5.0 <y<-2.5 1.5<y<4.0< td=""><td></td></y<4.0<></y<-2.5 	
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Facility	Experiment	System	√s <sub>NN</sub> (GeV)	Y <sub>cms</sub> range	Data taking
	NA50	p-Be,Al,Cu,Ag,W,Pb	27 29	-0.4 <y<0.6 -0.5<y<0.5< td=""><td>1996- 2000</td></y<0.5<></y<0.6 	1996- 2000
SPS	NA60	p-Be,Al,Cu,In,W,Pb,U	17 27	0.3 <y<0.8 -0.1<y<0.3< td=""><td>2004</td></y<0.3<></y<0.8 	2004
FNAL	E866	p-Be, Fe, W	39	Two energies	in the
HERA	HERA-B	p-C, Ti, W	42	same experim	
RHIC	PHENIX,	d-Au	200	-2.2 <y<2.4< td=""><td>&gt;2003</td></y<2.4<>	>2003
	STAR	p-Al, Au	200	1.2< y <2.2	2015
	ALICE			-4.46 <y<3.53< td=""><td></td></y<3.53<>	
	ATLAS			-2.87 <y<1.94< td=""><td rowspan="2">2013</td></y<1.94<>	2013
LHC	CMS	p-Pb	5020	-2.87 <y<1.93< td=""></y<1.93<>	
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SPS	NASO		29	-0.5 <y<0.5< td=""><td>2000</td></y<0.5<>	2000
JFJ	NA60	p-Be,Al,Cu,In,W,Pb,U	17	0.3 <y<0.8< td=""><td>2004</td></y<0.8<>	2004
	117400	p-be,Al,Cu,III,W,Pb,O	27	-0.1 <y<0.3< td=""><td>2004</td></y<0.3<>	2004
FNAL	E866	p-Be, Fe, W	39	-0.6 <y<2.5< td=""><td>~1996</td></y<2.5<>	~1996
HERA	HERA-B	p-C, Ti, W	42	- <b>1.5<v< b="">&lt;0.8</v<></b>	2002
RHIC	PHENIX,	d-Au	200	Largest x <sub>F</sub> coverage -0.10 <x<sub>F&lt;0.93</x<sub>	
	STAR	p-Al, Au	200		
	ALICE			-4.46 <y<3.53< td=""><td></td></y<3.53<>	
				-4.40 <y<3.55< td=""><td></td></y<3.55<>	
	ATLAS			-4.40 <y<3.53 -2.87<y<1.94< td=""><td></td></y<1.94<></y<3.53 	
LHC		p-Pb	5020	• 	2013
LHC	ATLAS	p-Pb	5020	-2.87 <y<1.94< td=""><td>2013</td></y<1.94<>	2013

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SPS	NA60		17	0.3 <y<0.8< td=""><td></td></y<0.8<>	
	NAOU	p-Be,Al,Cu,In,W,Pb,U	27	-0.1 <y<0.3< td=""><td>2004</td></y<0.3<>	2004
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		d-Au			
ршс	PHENIX,	d-Au	200	-2-2 <y<2.4< td=""><td>&gt;2003</td></y<2.4<>	>2003
RHIC	PHENIX, STAR	p-Al, Au 🚽	200	Telvier 2	2015
RHIC		p-Al, Au 🚽	Coverag	e up to negative	2015
	STAR	p-Al, Au 🚽	Coverag	Telvier 2	2015
RHIC	STAR ALICE	p-Al, Au 🚽	Coverag	e up to negative	2015
	STAR ALICE ATLAS	p-Al, Au	Coverag -(	e up to negative 0.34 <x<sub>F&lt;0.14</x<sub>	2015 X <sub>F</sub>

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Facility	Experiment	System	√s <sub>NN</sub> (GeV)	Y <sub>cms</sub> range	Data taking	
	NA50	p-Be,Al,Cu,Ag,W,Pb	27 <u>29</u>	-0.4 <y<0.6 -0.5<y<0.5< td=""><td>1996- 2000</td></y<0.5<></y<0.6 	1996- 2000	
SPS	NA60	p-Be,Al,Cu,In,W,Pb,U	<ul> <li>very high energies</li> <li>complementary</li> </ul>			
FNAL	E866	p-Be, Fe, W	KI	nematic ranges		
HERA	HERA-B	p-C, Ti, W	92	-1.5 <y<0.8< td=""><td>2002</td></y<0.8<>	2002	
RHIC	PHENIX,	d-Au	200	-2.2 <y<2.4< td=""><td>&gt;2003</td></y<2.4<>	>2003	
ппіс	STAR	p-Al, Au	200	1.2< y <2.2	2015	
	ALICE			-4.46 <y<3.53< td=""><td></td></y<3.53<>		
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### Quarkonium production in pA<sub>1</sub>

The study of the interaction of the cc pair with the nuclei provides:

# Constraints to production models The strength of this interaction may depend on the cc̄ quantum states and kinematics (R.Vogt, Nucl.Phys. A700,539 (2002), B.Z. Kopeliovich et al, Phys. Rev.D44, 3466 (1991))

Tool to investigate cold nuclear matter effects
 → complicated issue, interplay of many competing mechanisms as shadowing, energy loss, break-up in the medium...

Reference to disentangle genuine QGP effect in AA collisions  $\rightarrow$  Approach followed at SPS, RHIC and LHC

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### How is $J/\psi$ studied in pA? 12

Varying the amount of nuclear matter crossed by cc pair (studying  $J/\psi$  production as a function of A or centrality)

Selecting the kinematics of the quarkonium states e.g. selecting events where resonance is formed inside or outside the nucleus

Comparing the behavior of different resonances

"Effective" quantities are defined to evaluate the size of CNM effects

1) 
$$\sigma_{J/\psi}^{pA} = \sigma_{J/\psi}^{pp} \cdot A \cdot e^{-\langle \rho L \rangle \sigma} abs$$
 the larger  $\sigma_{abs}$ , the more important the nuclear effects  
2)  $\sigma_{J/\psi}^{pA} = \sigma_{J/\psi}^{pp} \cdot A^{\alpha} \leftarrow a=1 \Rightarrow \text{ no nuclear effects}$   
3)  $R_{J/\psi}^{pA} = \frac{\sigma_{J/\psi}^{pA}}{A \cdot \sigma_{J/\psi}^{pp}} \leftarrow R_{pA}=1 \Rightarrow \text{ no nuclear effects}$   
 $R_{pA}=1 \Rightarrow \text{ no nuclear effects}$ 

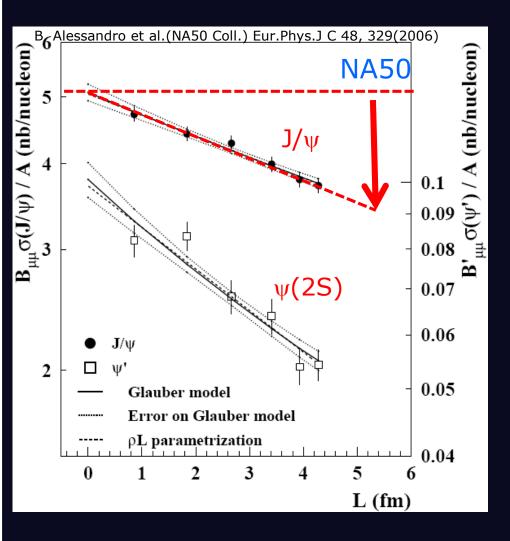
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### $J/\psi$ in pA at SPS

13

A significant reduction of the yield per NN collision is observed



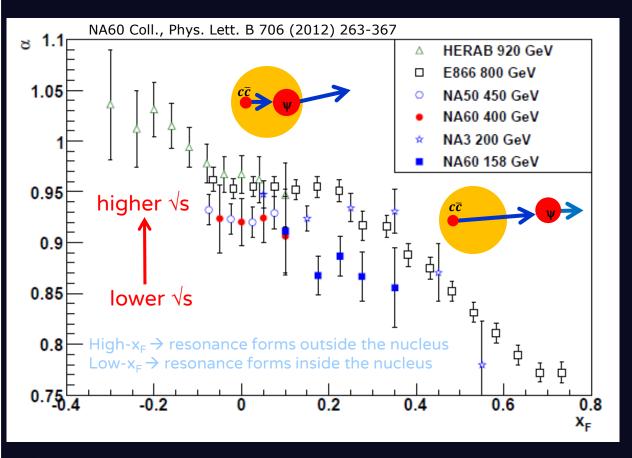
- Early studies interpreted this reduction as due to "nuclear absorption"
- Stronger absorption for the less bound state  $\psi(2S)$  at mid-y
  - → Nucleus crossing time (τ~0.3 fm/c) comparable or larger than charmonium formation time:
    - → fully formed resonances traversing the nucleus

 $\sigma_{abs} J/\psi = 4.5 \pm 0.5 \text{ mb}$  $\sigma_{abs} \psi(2S) = 8.3 \pm 0.9 \text{ mb}$ 

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### $J/\psi$ production vs x<sub>F</sub>

Compilation of fixed target results, collected at different  $\sqrt{s}$  and kinematical regions



 $J/\psi$  yield in pA is modified with respect to pp collisions

 $\alpha$  strongly decreases with  $x_{\rm F}$ 

for a fixed x<sub>F</sub>, CNM are stronger at lower √s

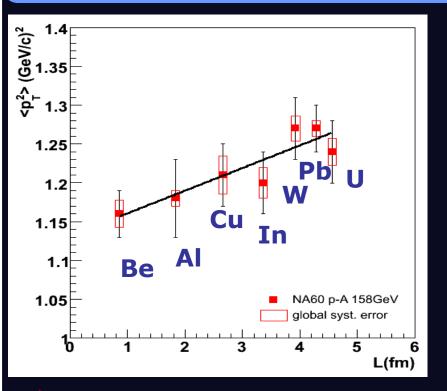
Theoretical description over the full x<sub>F</sub> range very complicate!

Given the strong  $x_F$  and  $\sqrt{s}$  dependence, pA data used as reference for AA collisions should be collected in the same kinematical domain

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### $J/\psi$ production vs $p_T$



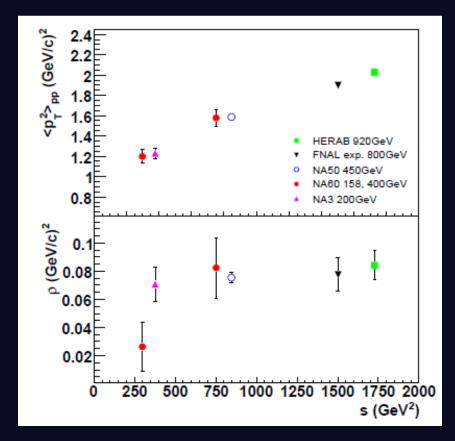
 $\boldsymbol{p}_{T}$  broadening can be parametrized as

 $< p_T^2 > = < p_T^2 >_{pp} + \rho (A^{1/3}-1)$ 

slope  $\rho$  is almost energy independent (apart from very low  $\sqrt{s}$ )

 $< p_T^2 >_{pp}$  increases with  $\sqrt{s}$ 

## <p<sup>2</sup>> increasing with the A of the target nuclei → interpreted as Cronin effect



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### Disentangling CNM effects 16

Assume dominant effects are shadowing and cc breakup at mid-y

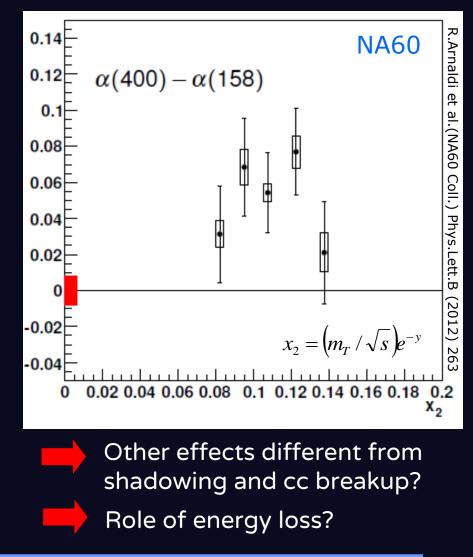
Shadowing in the target nucleus depends only on  $x_2$  (2 $\rightarrow$ 1approach)

J/ $\psi$  break-up depends on  $\sqrt{s_{J/\psi-N}}$ which is a function of  $x_2$ 

$$\sqrt{s_{J/\psi N}} \sim m_{J/\psi} \sqrt{\frac{1+x_2}{x_2}}$$

If parton shadowing and final state absorption were the only relevant mechanisms

α should not depend on √s at constant x<sub>2</sub> ... and this is clearly not the case

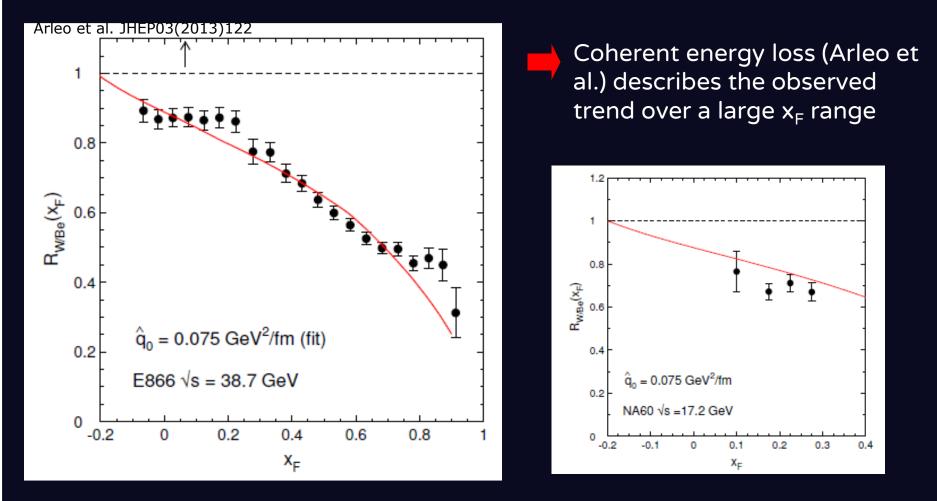


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### Disentangling CNM effects 17

The increase of the J/ $\psi$  suppression towards high x<sub>F</sub> might be interpreted as due to energy loss

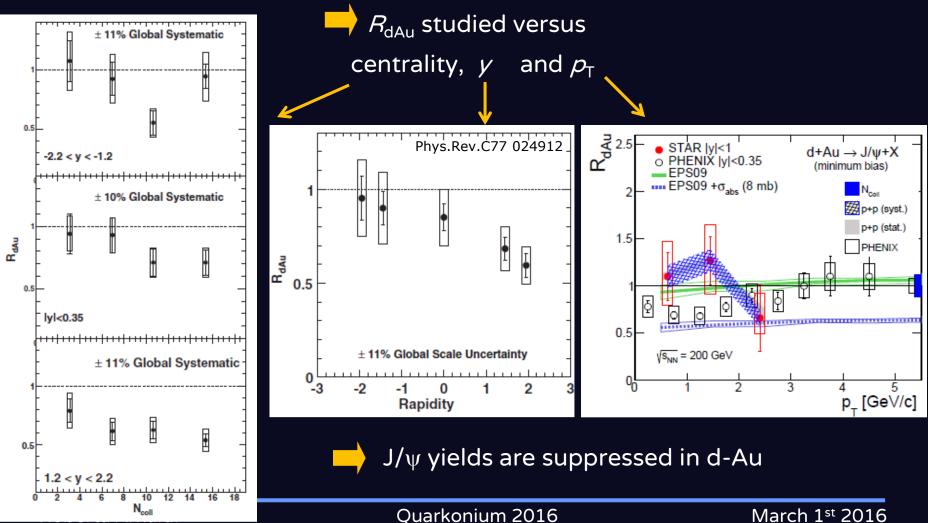


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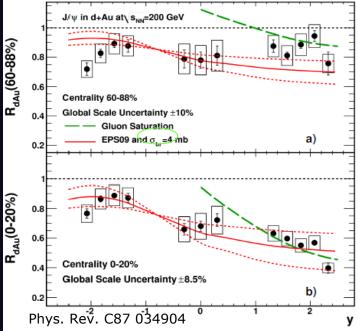
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### Moving to higher energies: RHIG8

Different approach wrt to fixed target experiments → Proton/deuteron on a single nucleus species and events selected on impact parameter



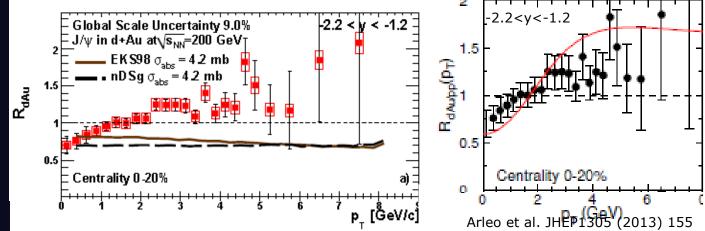
### **CNM effects at RHIC**



Disentangling CNM mechanisms is challenging

shadowing + cc break-up describe  $R_{dAu}$ vs y, but meets some difficulties for  $R_{dAu}$ vs  $p_T$ 

coherent energy loss contribution induces a less flat  $R_{dAu}$  dependence on  $p_T$ 



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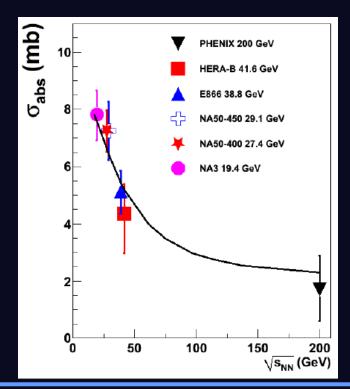
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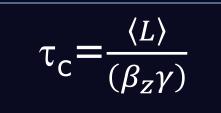
### Which CNM effects at LHC?20

Large Lorentz- $\gamma$  factor  $\rightarrow$  short crossing time of the cc in the nuclear matter

 $\rightarrow c\bar{c}$  pair almost point-like after crossing the nuclear matter

→ final state effects (as cc break-up) might be negligible





forward-y:  $\tau_c \sim 10^{-4}$  fm/c backward-y:  $\tau_c \sim 7.10^{-2}$  fm/c

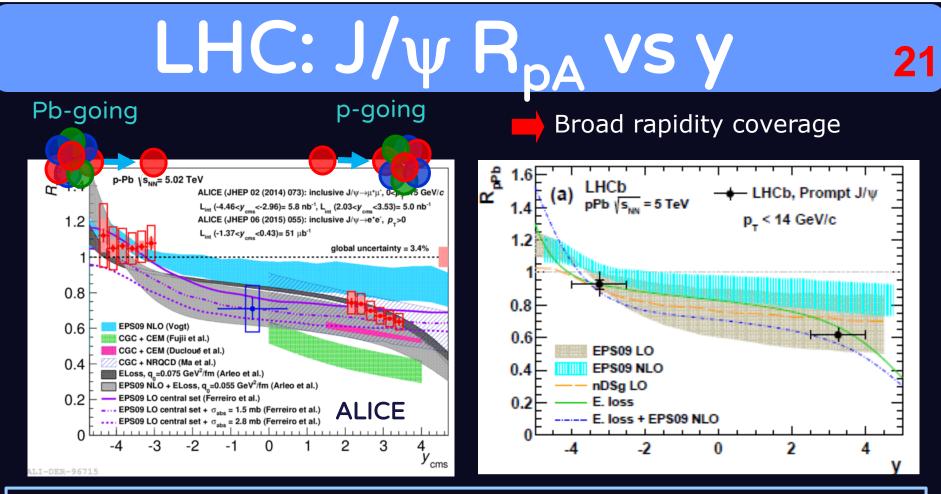
> D. McGlinchey, A. Frawley and R.Vogt, PRC 87,054910 (2013)

shadowing and/or energy loss might be the dominant effects

parton saturation effects can also be investigated at low-x

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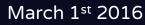
#### J/ $\psi$ production modified by CNM effects $\rightarrow R_{pA}$ decreases at forward y

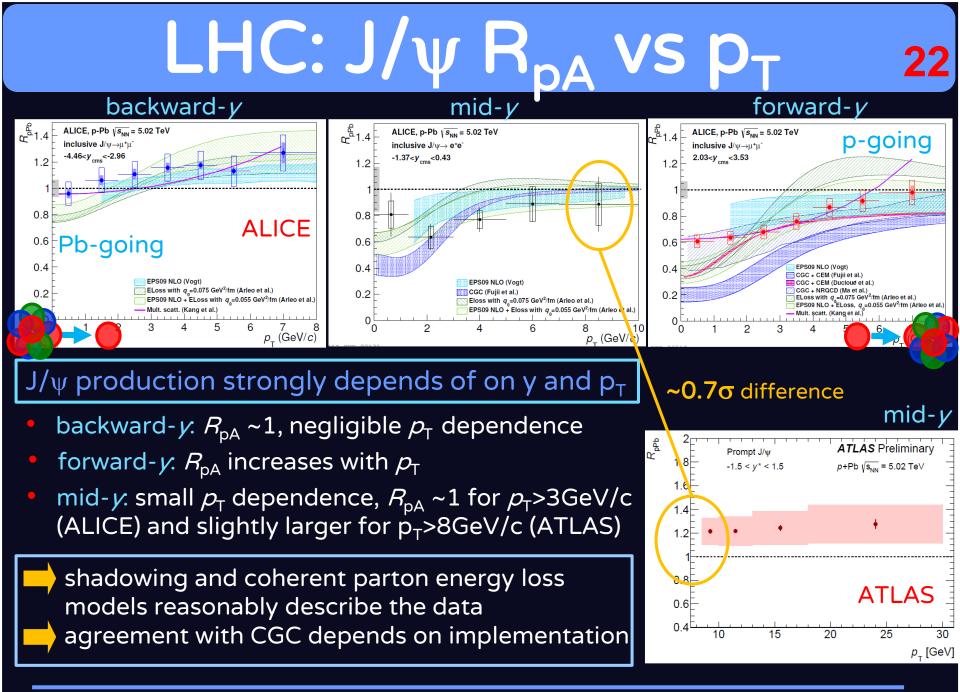
#### Theoretical predictions:

- shadowing calculations and models including coherent parton energy loss reasonably describe the data
- agreement with CGC depends on the implementation

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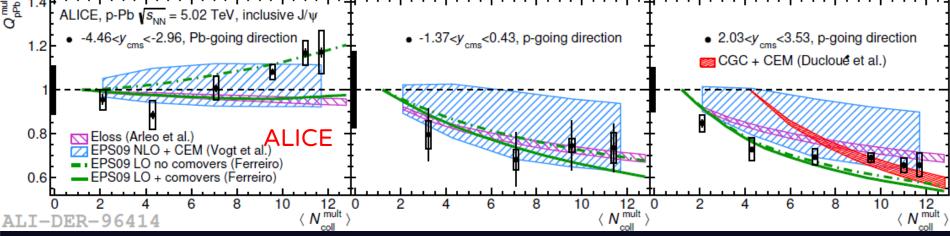
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## LHC: $J/\psi R_{pA}$ vs centrality<sub>23</sub>

#### backward-y

#### mid-y

#### forward-y



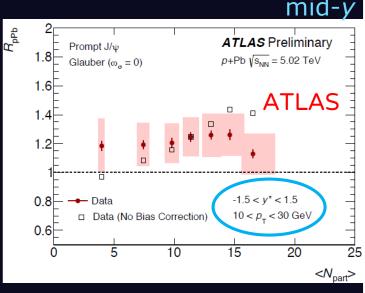
#### ALICE:

mid & fw-y: suppression increases with centrality backward-y: hint for increasing  $Q_{pA}$  with centrality

Shadowing and coherent energy loss models in fair agreement with data No strong comovers effect expected for J/ $\psi$ 

#### ATLAS

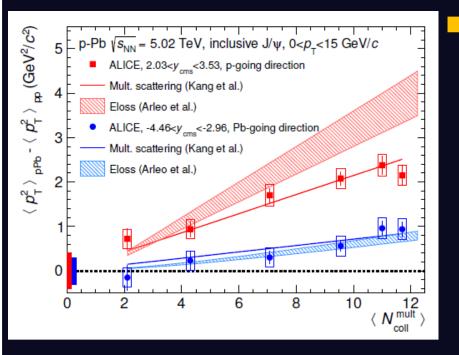
Flat centrality dependence in the high  $p_T$  range



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## $J/\psi < p_T^2 >$

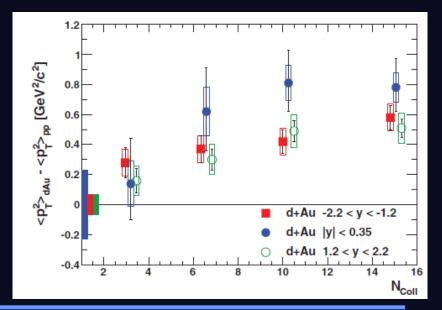


#### PHENIX:

- $p_{\rm T}$  broadening similar as the one observed by ALICE at backward-y
- large uncertainties prevent conclusions on the y dependence

#### ALICE:

- the  $p_T$  broadening  $\Delta < p_T^2 >$  increases from peripheral to central collisions
- effect is stronger at forward y
- initial/final state parton multiple scattering model describe the results energy loss describes the bck-y results, but predicts a steeper trend at forward y



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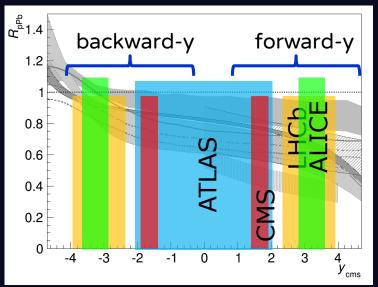
## $J/\psi R_{FB}$

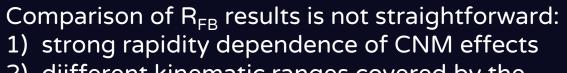


#### Forw. to backw. ratio in a common y range

$$R_{FB} = \frac{Y_{J/\Psi}^{forward}}{Y_{J/\Psi}^{backward}}$$

→ no pp reference is needed
→ but less straightforward to interpret

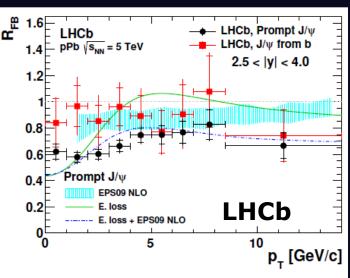


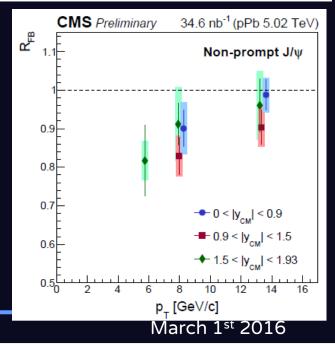


2) diifferent kinematic ranges covered by the experiments

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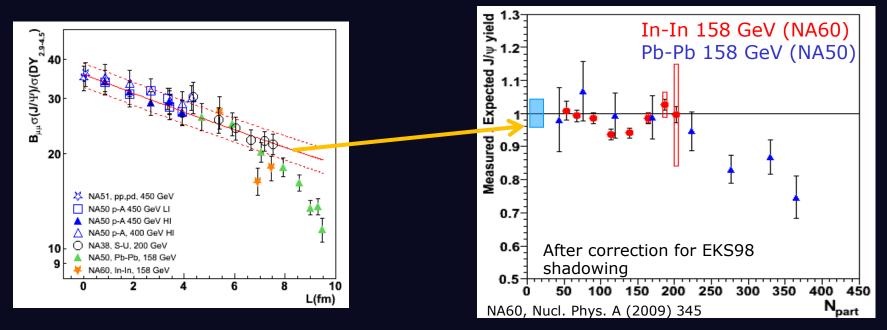
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### From pA to AA: SPS

Once CNM effects are measured in pA, how can they be extrapolated to AA?



#### SPS $\rightarrow$ the reference is built

- evaluating  $\sigma$  break-up in pA collisions (in the same kinematic range as AA)
- including project/target (anti)shadowing
- determining the reference centrality dependence through Glauber approach



### From pA to AA: RHIC

Same energy for AA and d-Au collisions

Reference evaluated with several approaches as:



 $R_{dAu}$  vs centrality (y) is described with various shadowing + break-up  $\sigma$ 

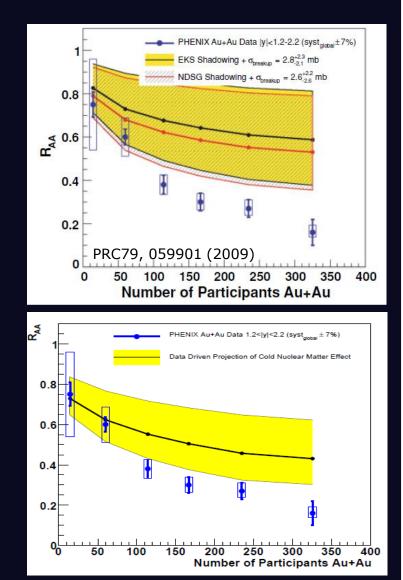
Shadowing scenario + break-up  $\sigma$  (evaluated in pA) are then compared to AA result



Data-driven approach:

All CNM effects (not disentangled) depend on the radial position in the nucleus

 $R_{AA} \sim R_{dAu}(-y) \times R_{dAu}(y)$ 



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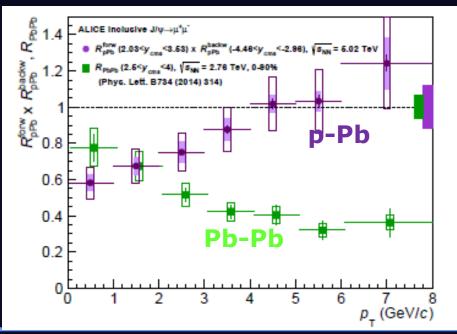
### From pA to AA: LHC 28

#### Different pA and AA $\sqrt{s}$ and y range

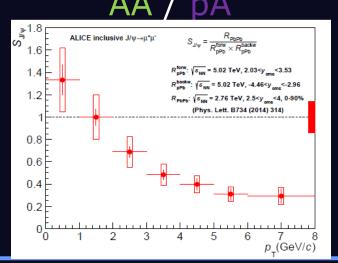
#### Hypothesis: $2 \rightarrow 1$ kinematics for J/ $\psi$ production

- CNM effects (dominated by shadowing) factorize in p-A
- CNM obtained as  $R_{pA} \times R_{Ap} (R_{pA}^2)$ , similar x-coverage as PbPb





CNM effects are "removed" via



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### Conclusions

The production of quarkonia in nuclear matter has been studied since a long time, both at fixed target and at colliders

The J/ $\psi$  production is modified in pA (d-Au) with respect to pp, with a strong kinematic dependence

Interplay of many cold nuclear matter effects as shadowing, energy loss and, at low  $\sqrt{s}$ , also cc break-up in the nucleus

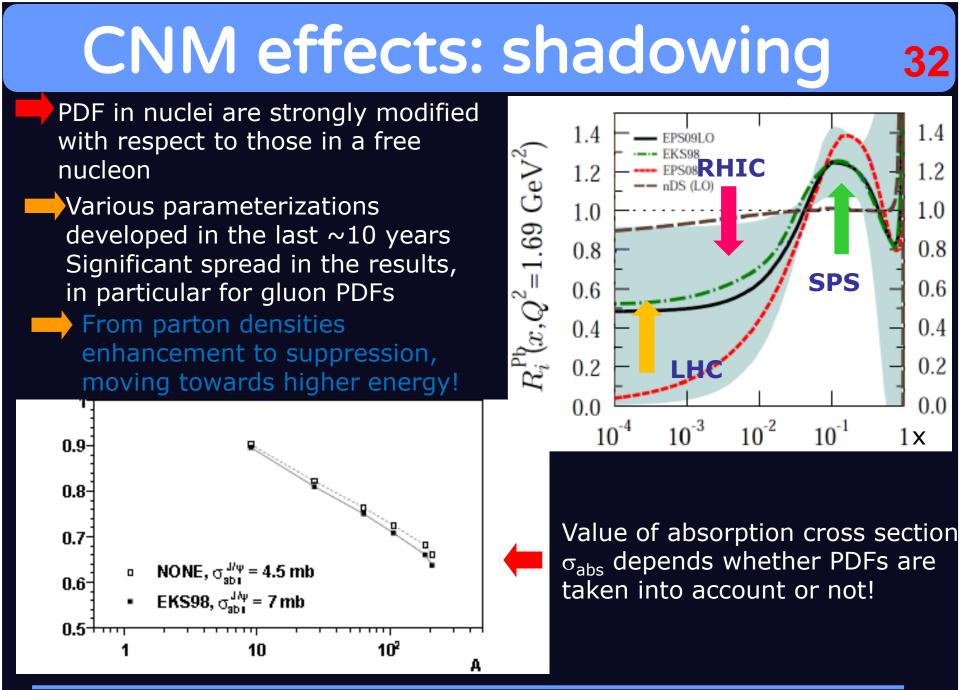
- $\rightarrow$  Modeling is complicate, but progresses have been done!
- → However, size of uncertainties prevents a clear assessment of the role of the various contributions



## J/ψ in pA

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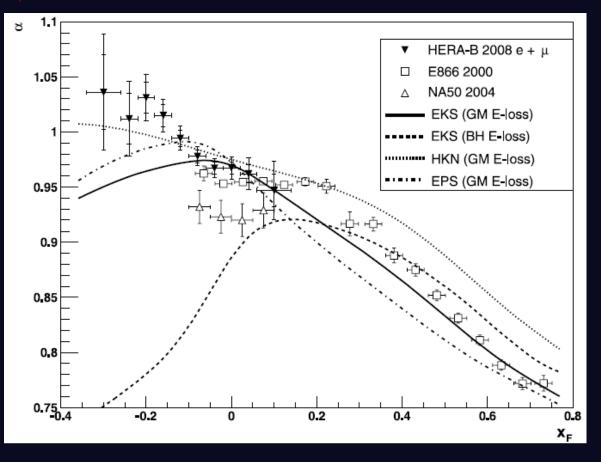
Facility	Experiment	System	√s <sub>NN</sub> (GeV)	x <sub>F</sub> range	Data taking
	NA50	p-Be,Al,Cu,Ag,W,Pb	27	-0.14 <x<sub>F&lt;0.10</x<sub>	1996-
CDC		₽ = -; <i>,,</i>	29	-0.10 <x<sub>F&lt;0.10</x<sub>	2000
SPS	NA60	p-Be,Al,Cu,In,W,Pb,U	17	0.05 <x<sub>F&lt;0.40</x<sub>	2004
	NAGO	p-be,Al,Cu,III,W,Pb,O	27	-0.07 <x<sub>F&lt;0.12</x<sub>	2004
FNAL	E866	p-Be, Fe, W	39	-0.10 <x<sub>F&lt;0.93</x<sub>	~1996
HERA	HERA-B	p-C, W	42	-0.34 <x<sub>F&lt;0.14</x<sub>	2002
	PHENIX,	d-Au		-0.1 <x<sub>F&lt;0.2</x<sub>	>2003
RHIC	STAR	p-Al, Au	200	0.05 < x <sub>F</sub>  <0.14	2015
	ALICE			-0.05 <x<sub>F&lt;0.02</x<sub>	
	ATLAS			-0.01 <x<sub>F&lt;-0.004</x<sub>	2013
LHC	CMS	p-Pb	5020	-0.01 <x<sub>F&lt;-0.004</x<sub>	
	LHCb			-0.09 <x<sub>F&lt;-0.007 0.003<x<sub>F&lt;-0.03</x<sub></x<sub>	
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### $J/\psi$ production vs xF

#### Compilation of fixed target results:



High- $x_F$  → resonance forms outside the nucleus Low- $x_F$  → resonance forms inside the nucleus

$$\sigma_{J/\psi}^{pA} = \sigma_{J/\psi}^{pp} \cdot A^{\alpha}$$

 $J/\psi$  production is modified by the medium already in pA collisions

 $\alpha$  strongly decreases with xF

for a given  $x_F$ , CNM are stronger at lower  $\sqrt{s}$ 

Satisfactory theoretical description still unavailable!



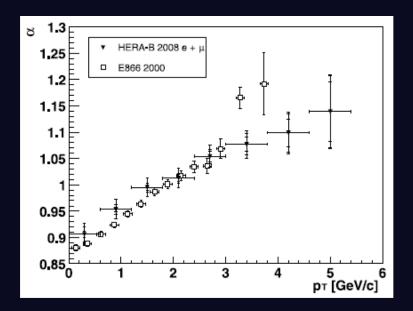
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### $J/\psi$ production vs $p_T$

The J/ $\psi$  suppression is stronger at low pT

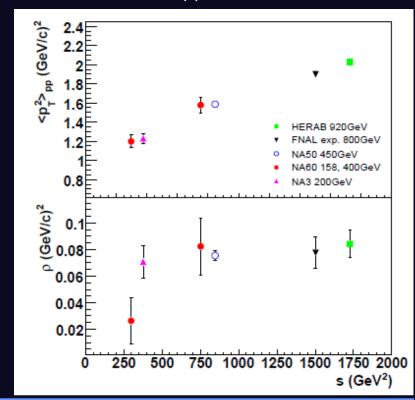
Increase of  $\alpha$  with  $p_T$  interpreted in terms of Cronin effect



Slope  $\rho$  is almost energy independent (apart from very low  $\sqrt{s}$ ) while  $< p_T^2 > pp$  increases with  $\sqrt{s}$ 

A broadening of pT as a function of A is observed:

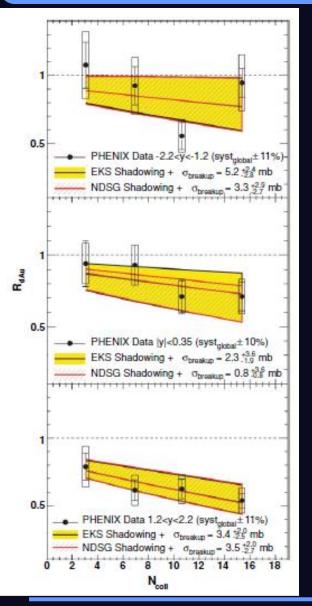
$$< p_T^2 > = < p_T^2 >_{pp} + \rho (A^{1/3} - 1)$$



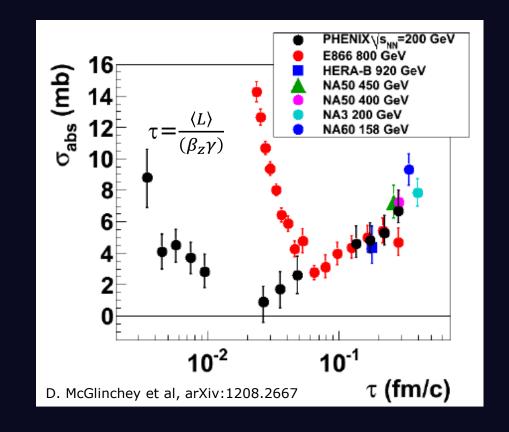
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### Moving to higher energies: RHIG5



Results might be described including shadowing and a rapidity-dependent σ<sub>abs</sub>



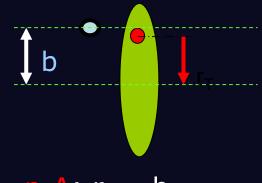
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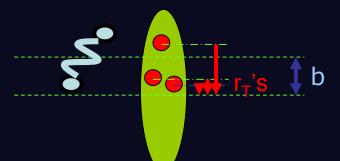
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### Moving to higher energies: RHIG6

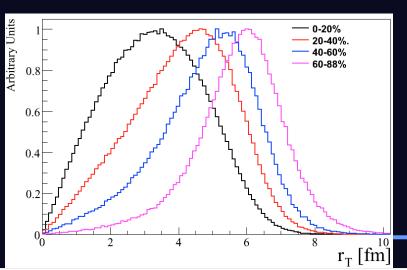
Different approach wrt to fixed target experiments:

Instead of accelerating several different nuclei  $\rightarrow$  Use one single nucleus species and select on impact parameter





p-A:  $r_T \sim b$ 



d-Au: due to the size of the deuteron (<r>~2.5fm) the distribution of transverse positions of the collisions are not very well represented by impact parameter

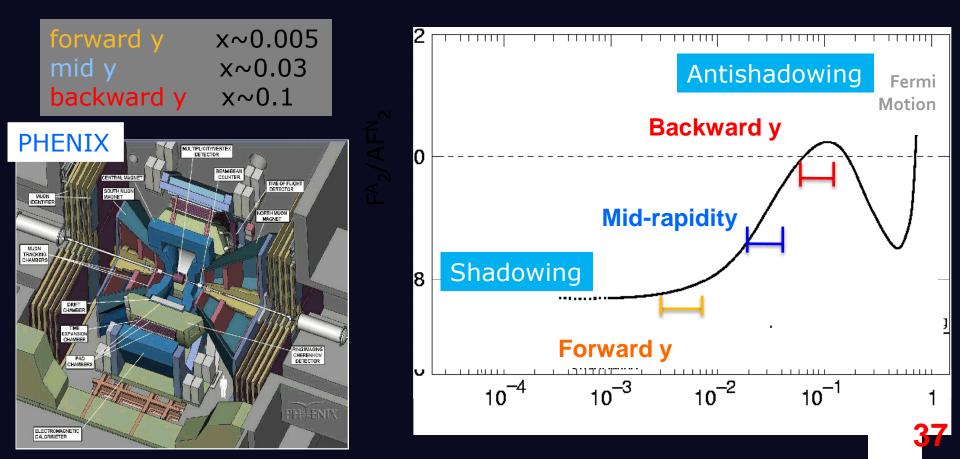
 $\rightarrow$  overlap of the centrality classes

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### d-Au rapidity range

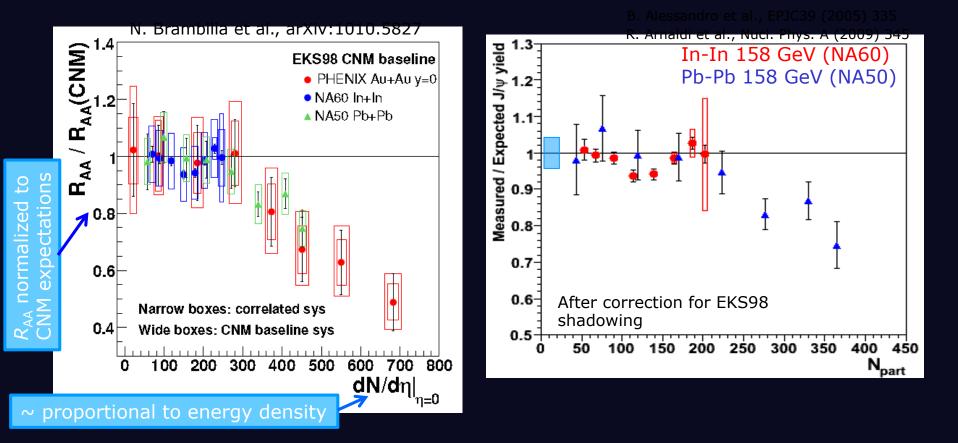
Regions corresponding to very different strength of shadowing effects have been studied: -2.2 < y < -1.2, |y| < 0.35, 1.2 < y < 2.2

 $\rightarrow$  good test of our understanding of the physics!



### From p-A to A-A...

Even if disentangling the different CNM mechanisms is a complicate issue... ...CNM, evaluated in p-A, can be extrapolated to A-A to build a reference for the J/ $\psi$  behaviour in hadronic matter!

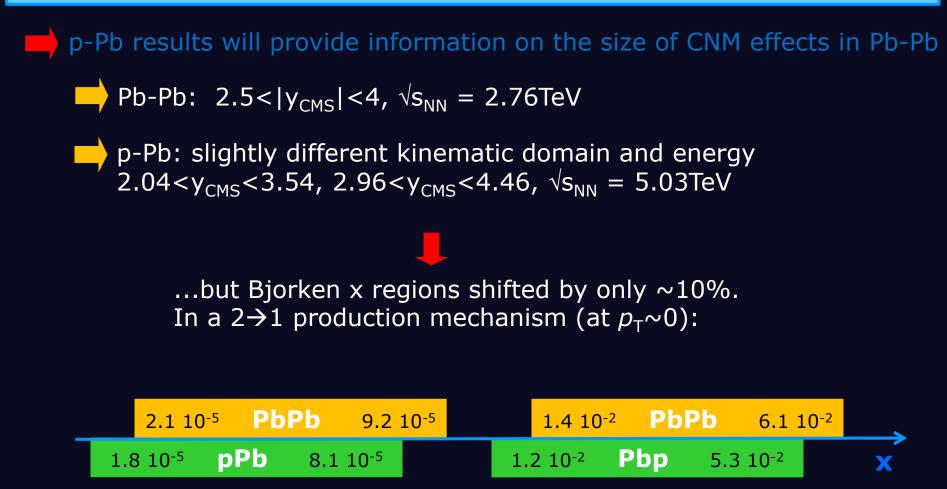


Clear suppression is indeed observed on top of CNM effects!

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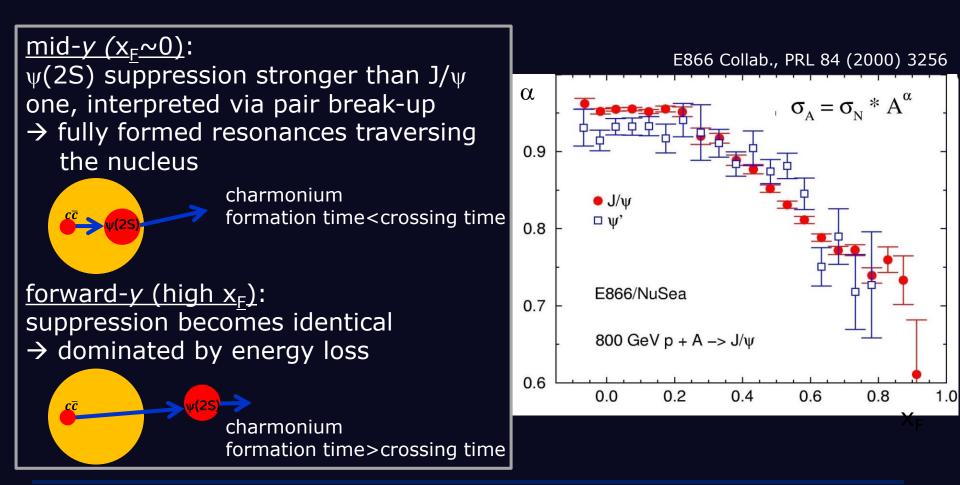
### From p-Pb to Pb-Pb...



Work in progress to quantify size of CNM effects in Pb-Pb results!



Being more weakly bound than the J/ $\psi$ , the  $\psi$ (2S) is an interesting probe to have further insight on the charmonium behaviour in pA Low energy  $\psi$ (2S) p-A results from NA50, E866 and HERA-B:



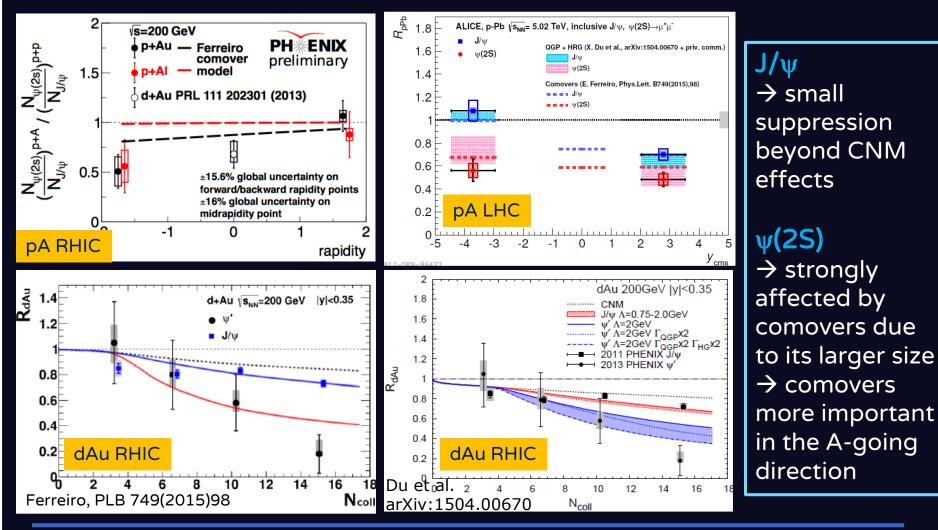
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October 2<sup>nd</sup> 2015

### comparison to theoretical models,

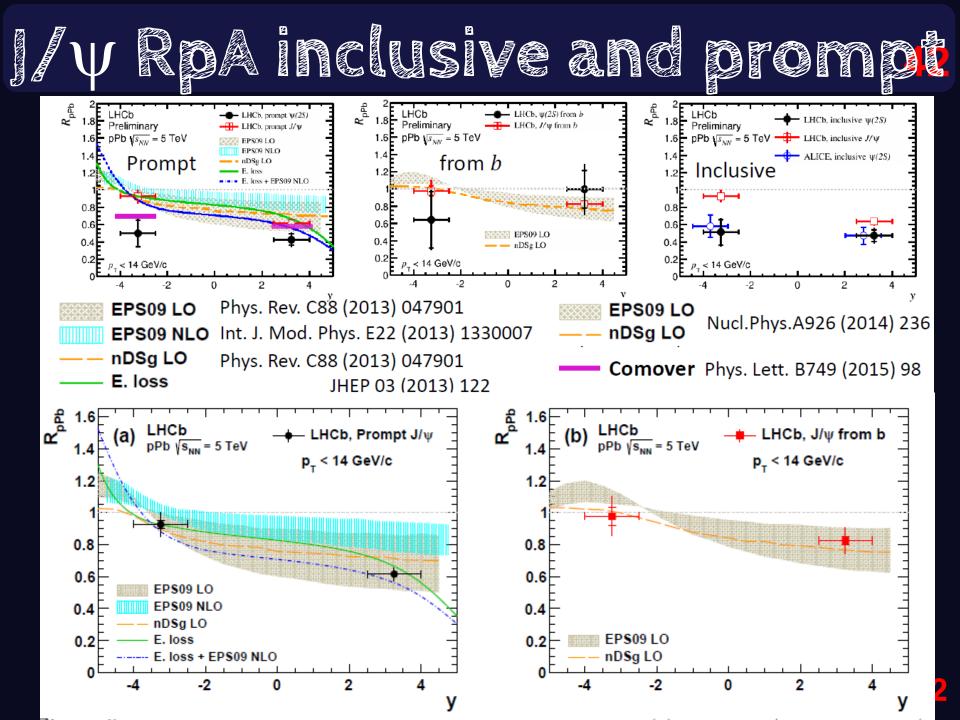
QGP+hadron resonance gas (Rapp) or comovers models (Ferreiro) reasonably describe both J/ $\psi$  and  $\psi$ (2S) suppression at RHIC and LHC



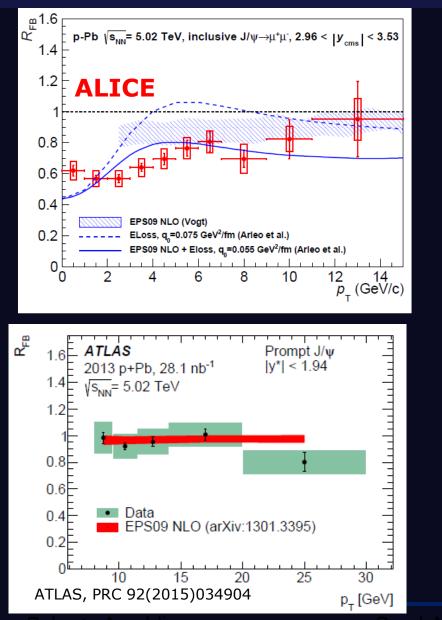
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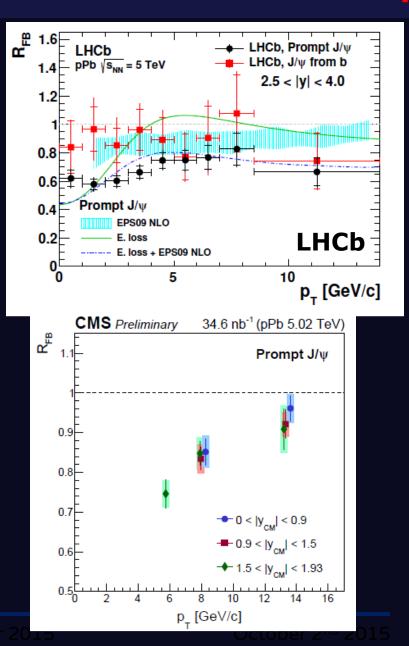
Quark Matter 2015

October 2<sup>nd</sup> 2015

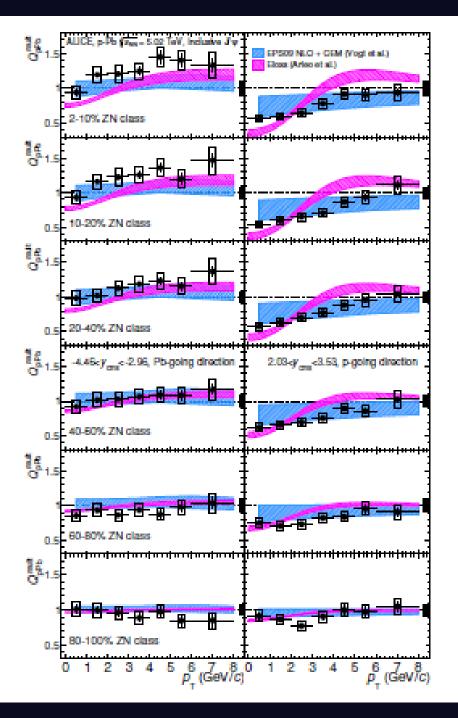


### $//\psi$ forward to backward ratio 43





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Differential results might provide constraints to theoretical calculations

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