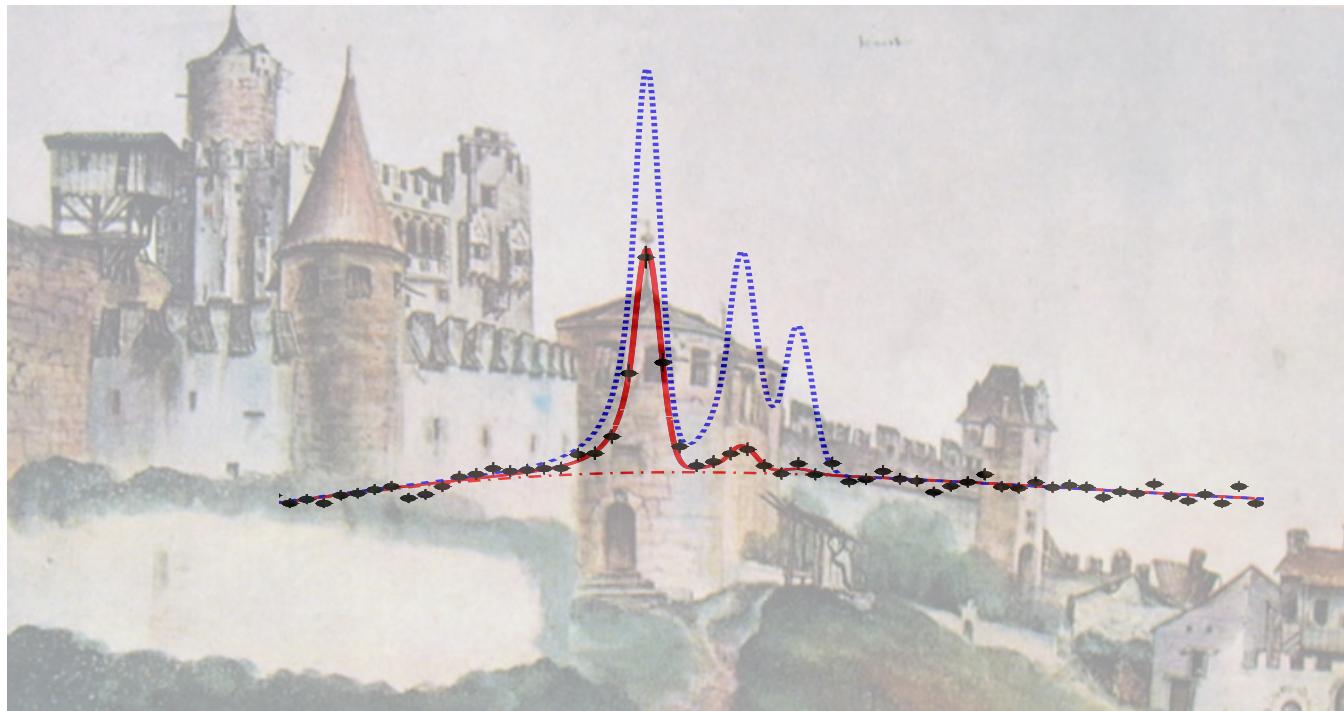


Υ production in pA and AA collisions from Fermilab, RHIC, and LHC



New Observables in Quarkonium Production
ECT* Workshop
1/March/2016

Manuel Calderón de la Barca Sánchez



UCDAVIS

Outline

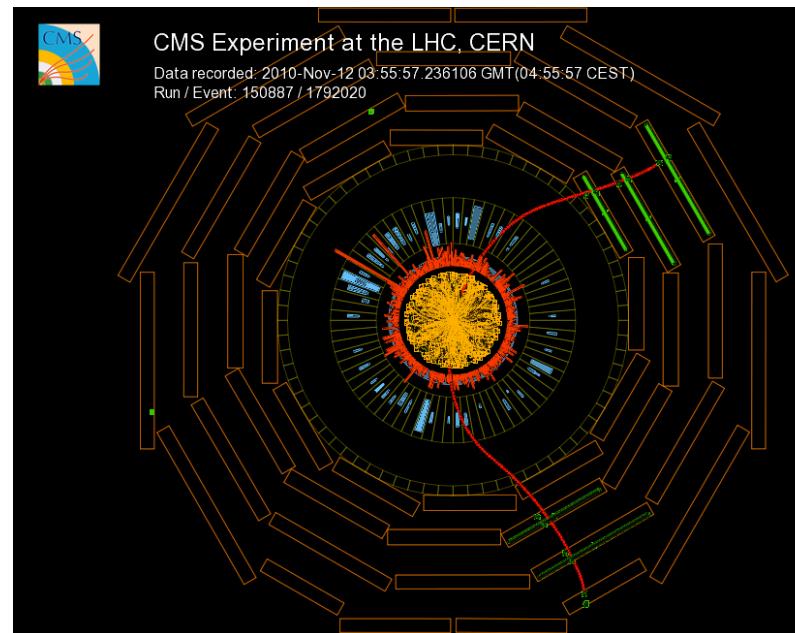


- Bottomonium Measurement in pA and AA collisions.
 - Emphasis on $\Upsilon(1S)$
 - Caveat: Many results include ground states.
 - Will be mentioned, but leave emphasis for session on excited states.
 - Fermilab pA
 - $\sqrt{s}=40 \text{ GeV}$
 - RHIC dA, AA
 - $\sqrt{s}=200 \text{ GeV}$
 - LHC pA, AA
 - $\sqrt{s}=2.76, 5.02 \text{ TeV}$
- Summary

Bottomonium in pA and AA

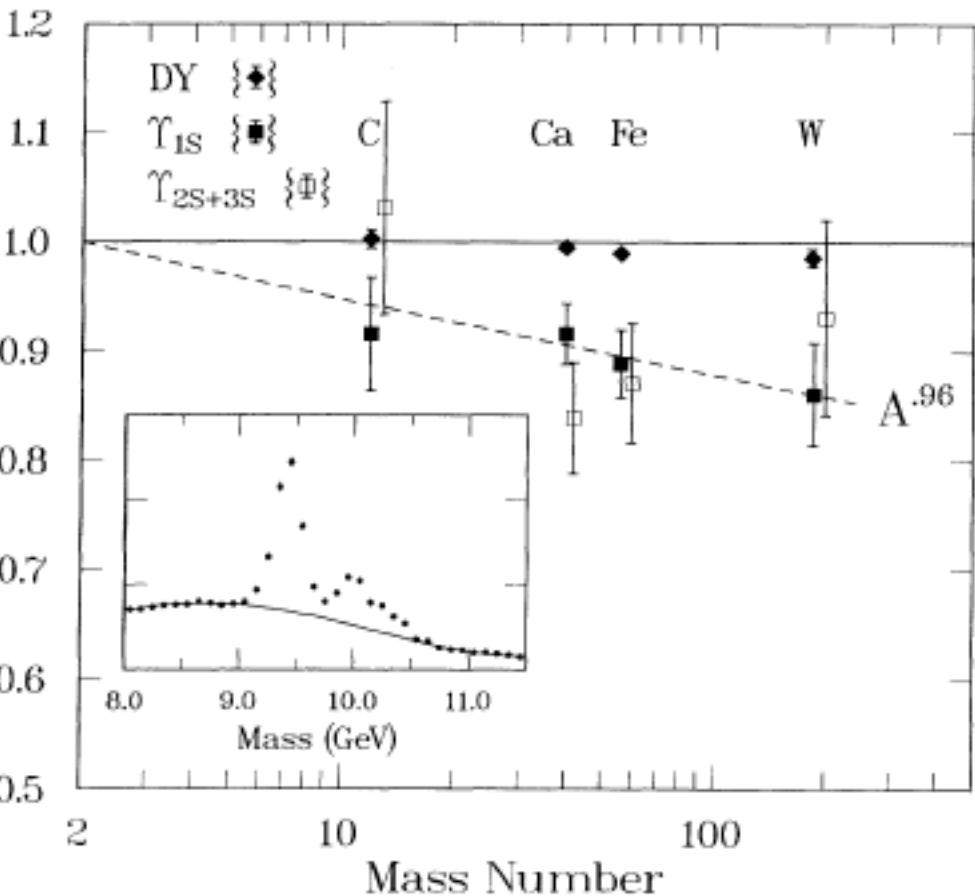


- Υ in AA: a probe of color deconfinement
 - Suppression of yield: states do not form in a color deconfined medium
 - Color screening, gluodissociation, partonic breakup/Landau damping.
- Υ in pA: Production in Cold nuclear matter
 - Initial state effects: e.g. nPDF
 - Final state: energy loss, absorption
- Bottomonium: a cleaner probe than charmonium...
 - 3 states are accessible experimentally
 - Expect smaller recombination contribution compared to charmonium
 - $m_b \sim 3 m_c$





E772 pA Υ measurement



- Ratio of nuclear targets normalized to deuterium
- Suppression seen with increasing A .
- A dependence:

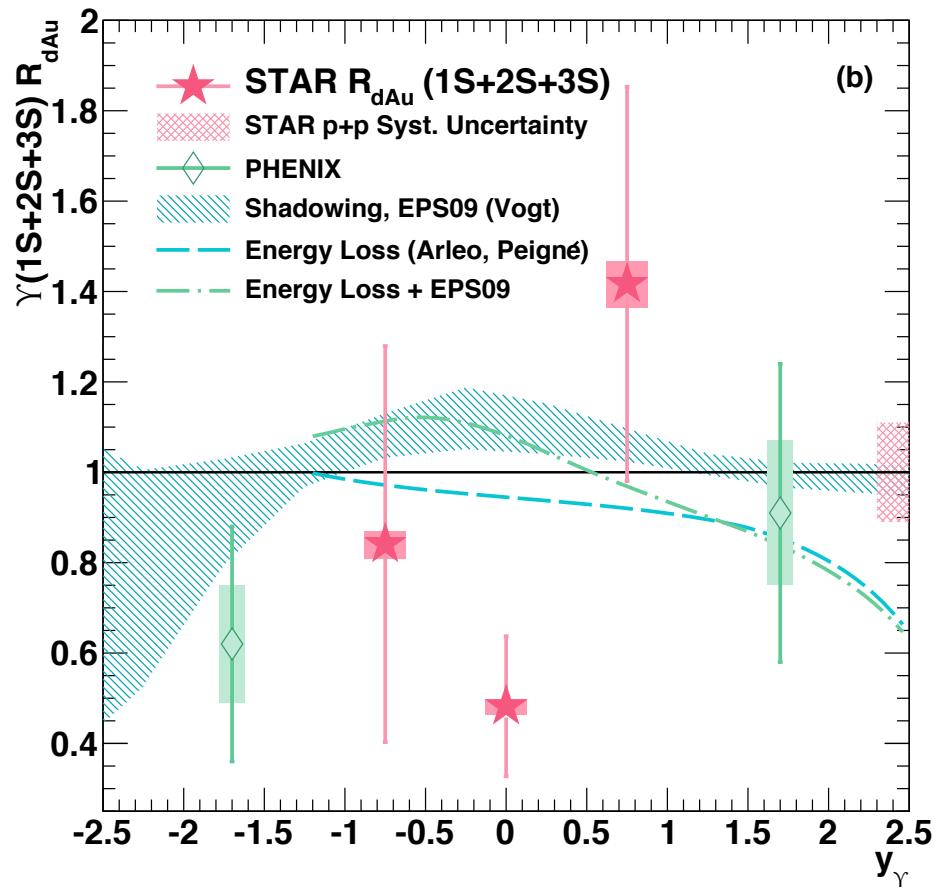
$$\alpha_{1S} = 0.962 \pm 0.006$$

$$\alpha_{2S+3S} = 0.948 \pm 0.012$$

PRL 66 (1991) 2285



R_{dAu} at RHIC

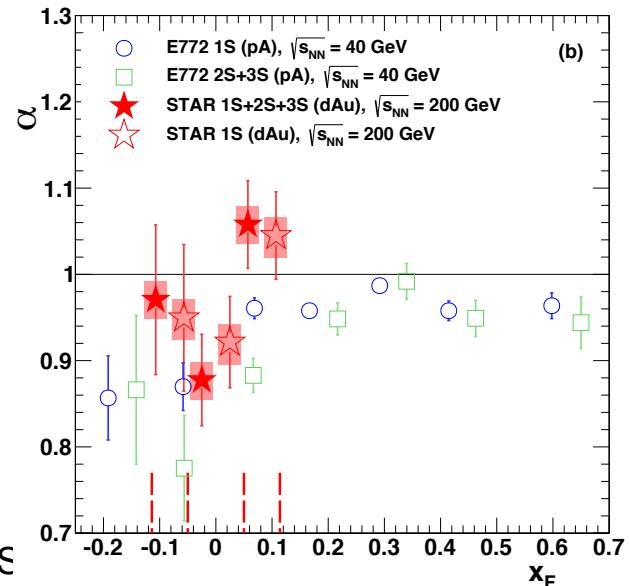
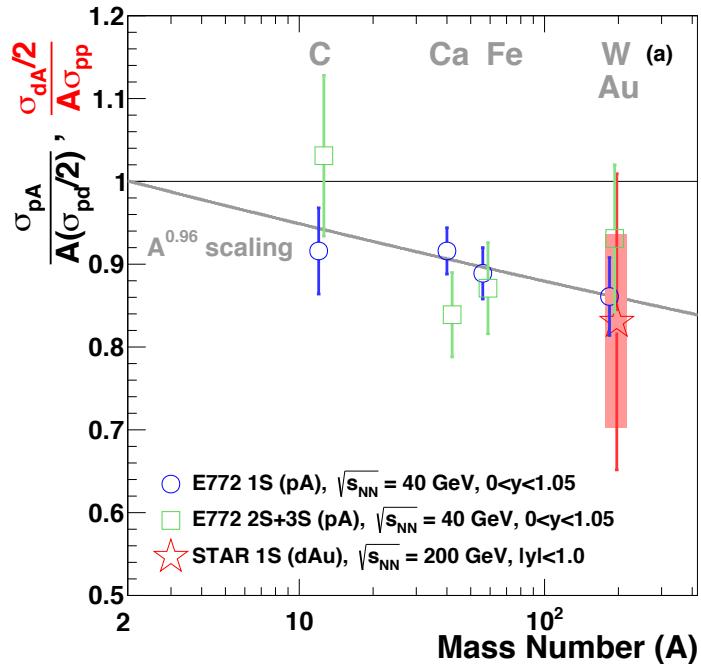


- R_{dAu} vs. y
 - Model comparison:
 - Shadowing, EPS09
 - R. Vogt
 - Energy loss
 - Energy loss + shadowing
 - $y \sim 0$ is right in the middle of the antishadowing region
 - Expect $R_{dAu} > 1$ (small effect)
 - Observe $R_{dAu} < 1$



Comparison: STAR & E772

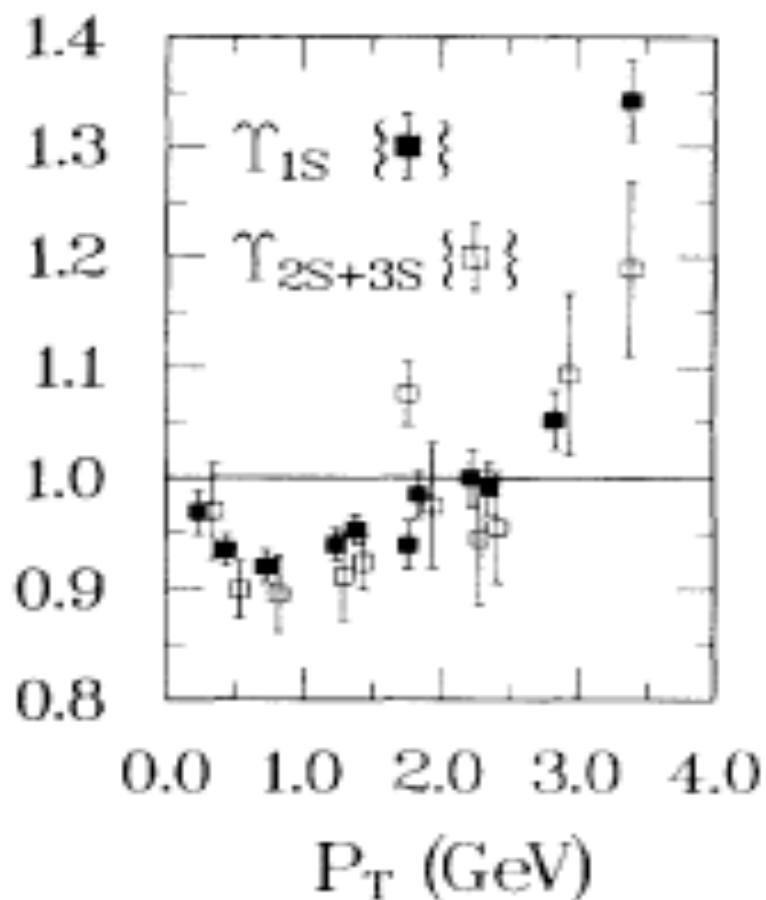
- For a similar comparison, separate $\Upsilon(1S)$ stat
 - Use $|y| < 1$, check A dependence.
- STAR result: consistent with trend vs. A from E772.
- Large suppression seen near $x_F \sim 0$ by E772, $\alpha \sim 0.9$.
 - Similar to STAR $|y| < 0.5$ points.



Note: p_T dependence of α in E772



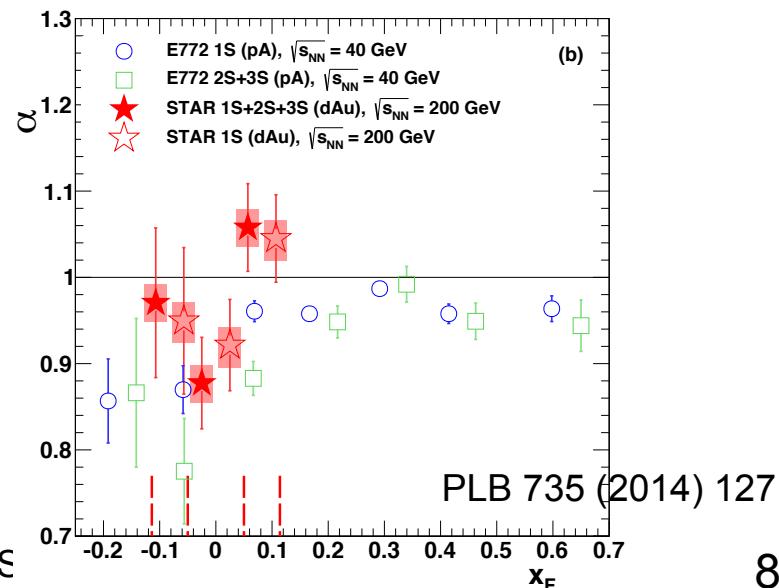
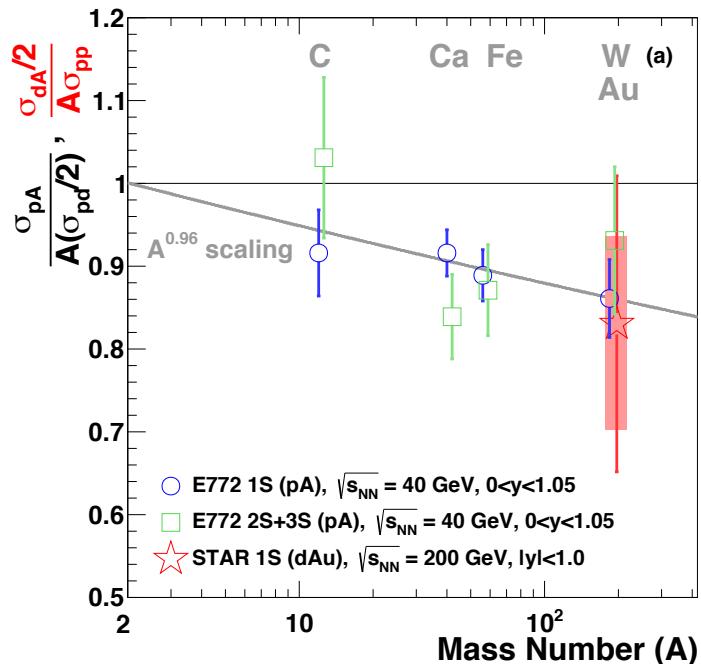
- Not trivial either.
 - Suppression largest at ~ 1 GeV
 - Large enhancement above 3 GeV
- No LHC measurements vs. p_T for pPb, yet.
 - pp reference at 5.02 TeV just taken



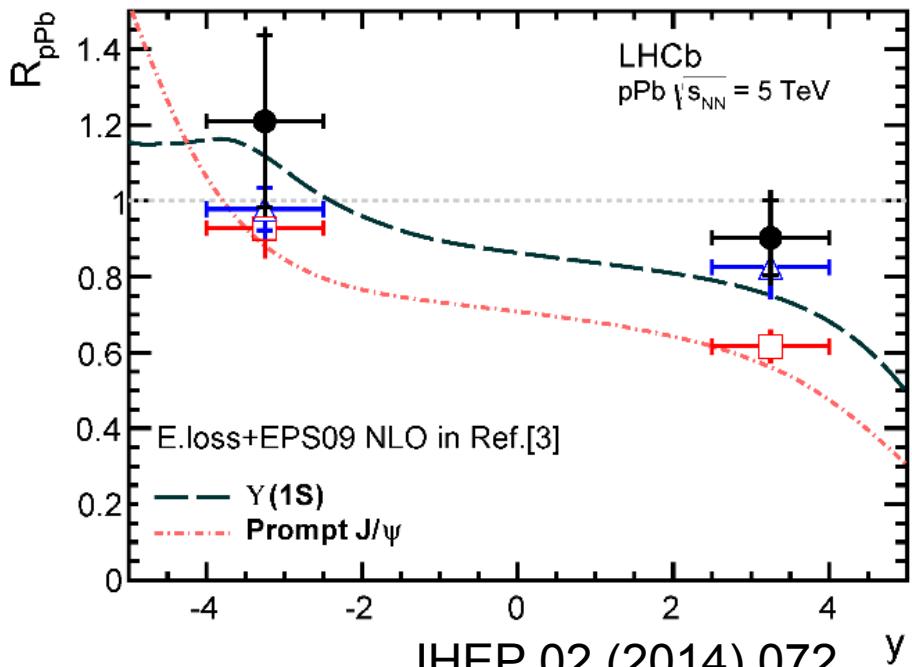
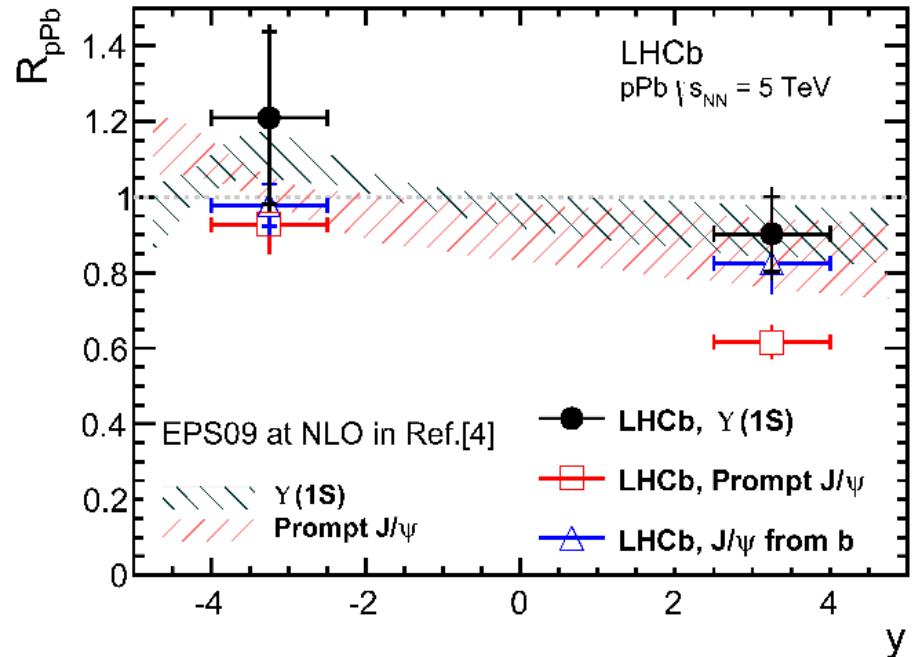
Comparison: STAR & E772



- For a similar comparison, separate $\Upsilon(1S)$ stat
 - Increases the statistical uncertainty compared to sum $\Upsilon(1S+2S+3S)$
 - Use $|y| < 1$, check A dependence.
 - Also compare y or x_F dependence.
- STAR result: consistent with A trend from E772.
- Large suppression seen near $x_F \sim 0$ by E772, $\alpha \sim 0.9$.
 - Same as STAR $|y| < 0.5$ points.
- Shadowing, or shadowing+E. Loss cannot explain suppression at $y=0$.
- Effect goes away in the forward y bins.
- A higher-statistics d+Au run would help.
 - Note: dAu 2008 run was first attempt at measuring bottomonium in cold nuclear matter, can revisit with higher statistics



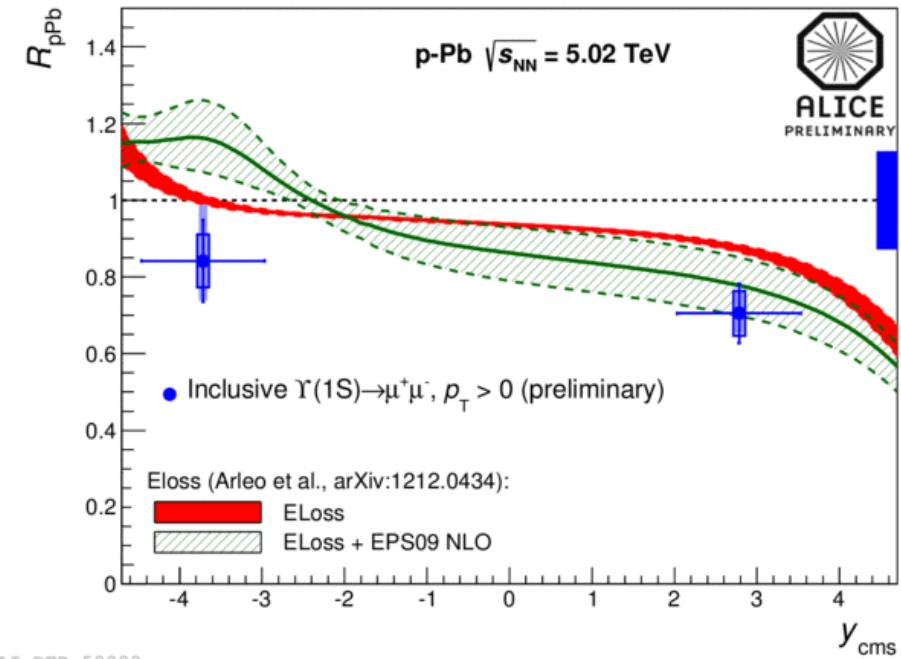
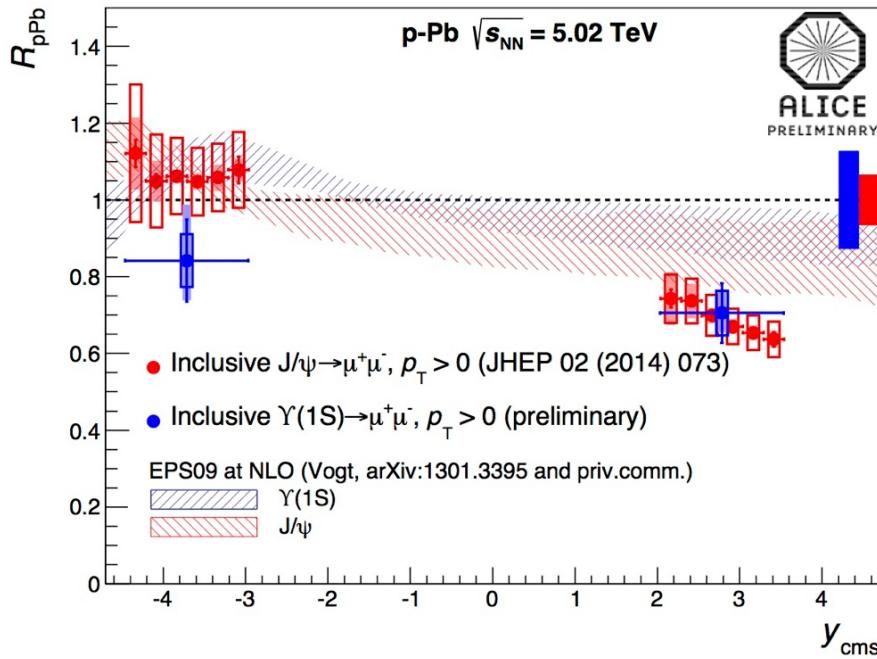
LHCb Υ results for $R_{p\text{Pb}}$



JHEP 02 (2014) 072
JHEP 07 (2014) 094

- Caveat emptor:
 - All LHC $R_{p\text{Pb}}$ results to date done without pp Reference at 5.02 TeV:
 - LHCb: Power-law interpolation: $\text{BR} \times s = 1.12 \pm 0.11 \text{ nb}$
- Slight enhancement at negative rapidity, indication of antishadowing
- Slight suppression at forward rapidity
- Different theoretical models are consistent with data, within uncertainties
 - EPS09 NLO: IJMP E22 (2013) 1330007
 - E. loss : JHEP 03 (2013) 122

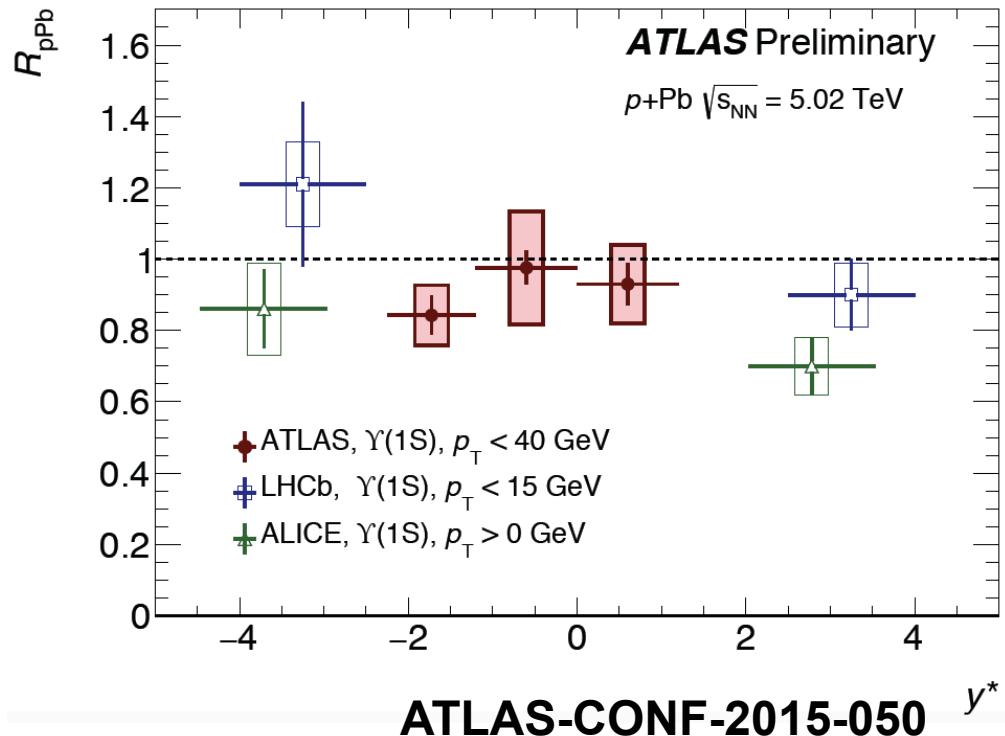
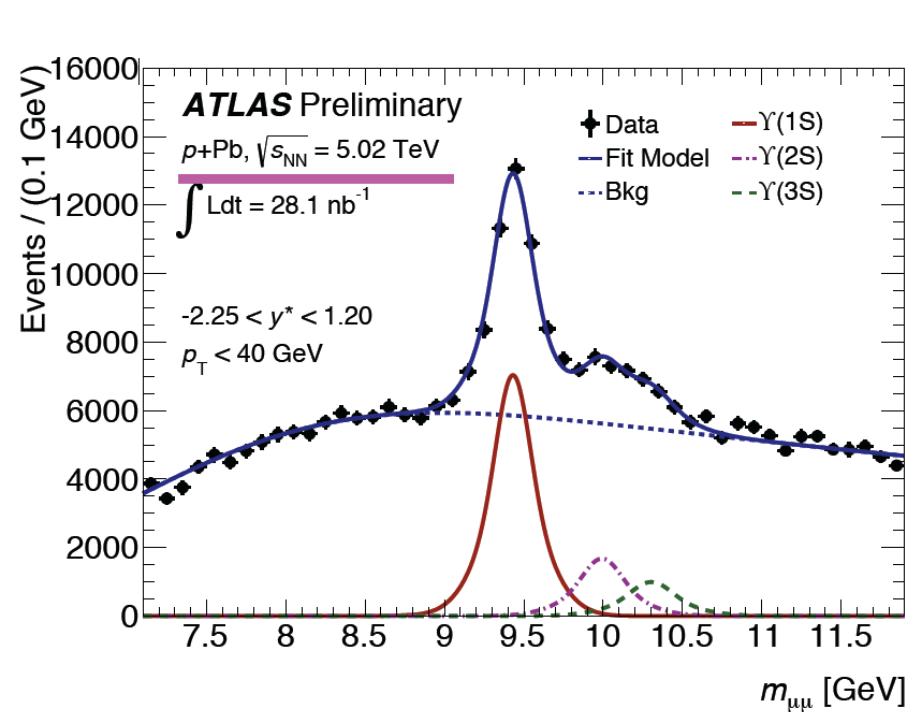
ALICE Υ results for R_{pPb}



ALI-DER-58999

- Backward rapidity: ALICE sees no enhancement, slight suppression.
 - EPS09 NLO: antishadowing in forward region. Enhancement remains when including ELoss + EPS09.
- Forward rapidity data:
 - EPS09 NLO expects only modest suppression.
 - Including E. Loss lowers R_{pPb} , data near lower end of prediction
- Note: **ALICE data in both cases lower than LHCb data.**
- ALICE pp interpolation:
 - Used LHCb data (not in the same rapidity range).
 - Fit with CEM (15 different choices of scale, pdf), FONLL bare quark energy dependence; power-law, linear, exponential.
 - $1451 \pm 114(\text{syst}) \text{ pb}$ and $770 \pm 87(\text{syst}) \text{ pb}$,

ATLAS LHC R_{pPb} Υ(1S)



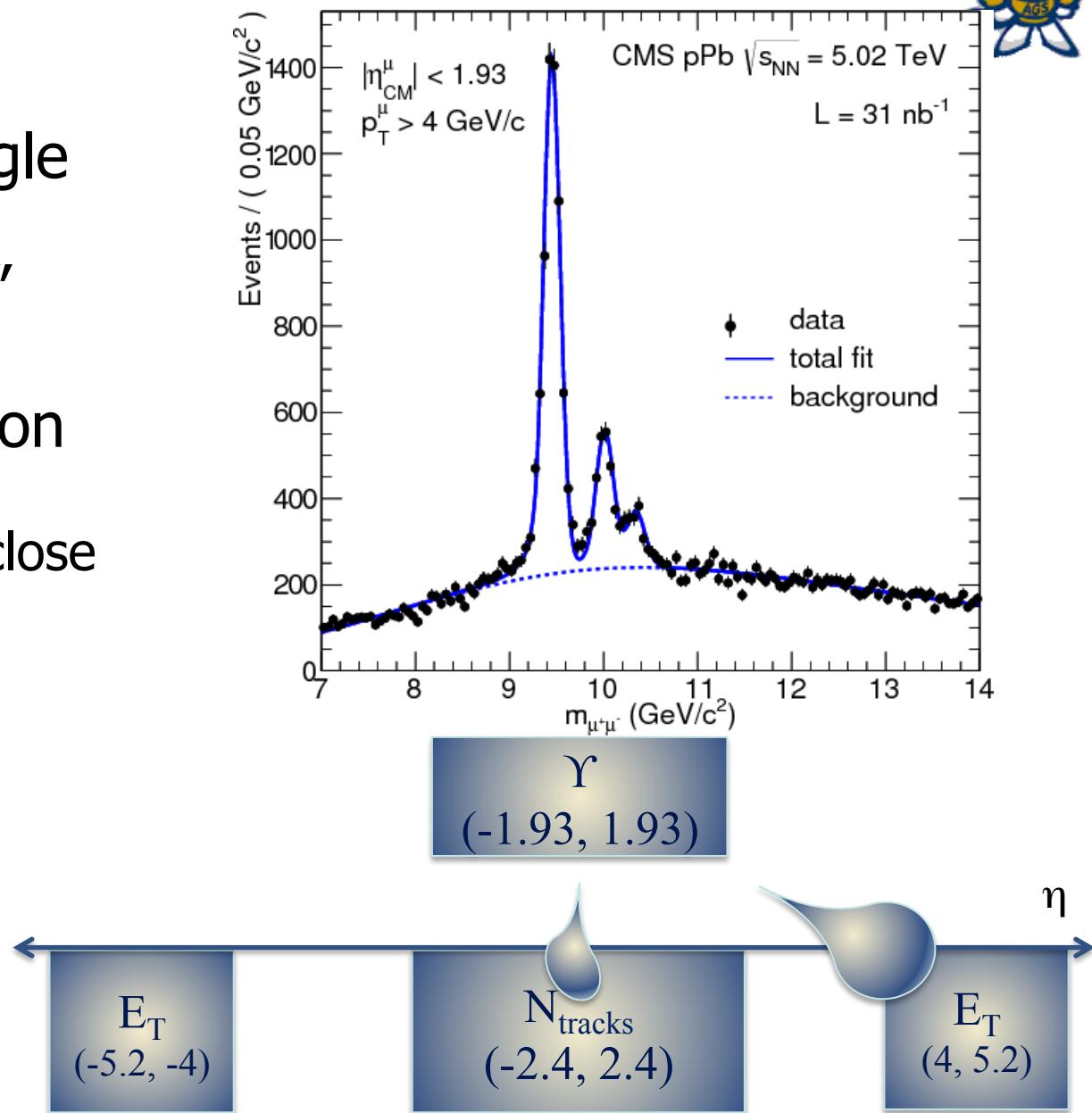
- 1S in pPb, Little to no modification seen at midrapidity.
 - Contrast to E772 & STAR lower energy data for $y \sim 0$.

Υ in pPb at LHC with CMS

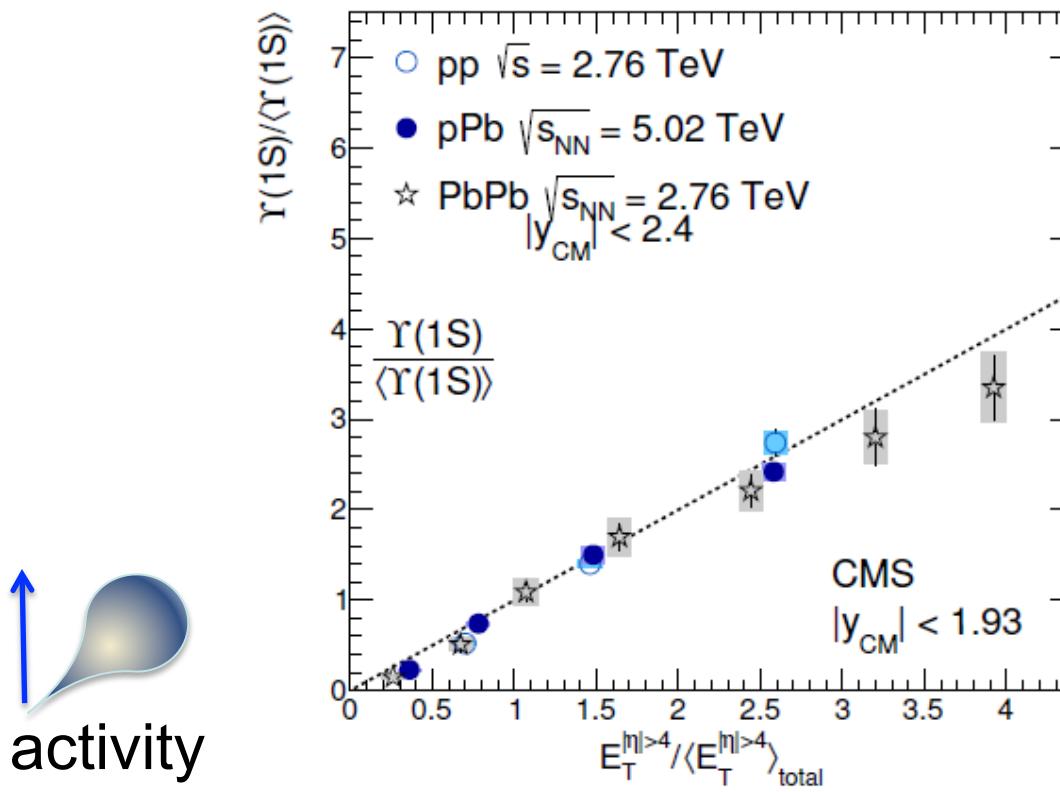


- Observables:

- Double ratio, single ratio
- “Self-normalized” yields
- Study as a function of event activity
 - Look at activity close to or far from Υ meson.
 - Close: N_{tracks}
 - Far: E_T

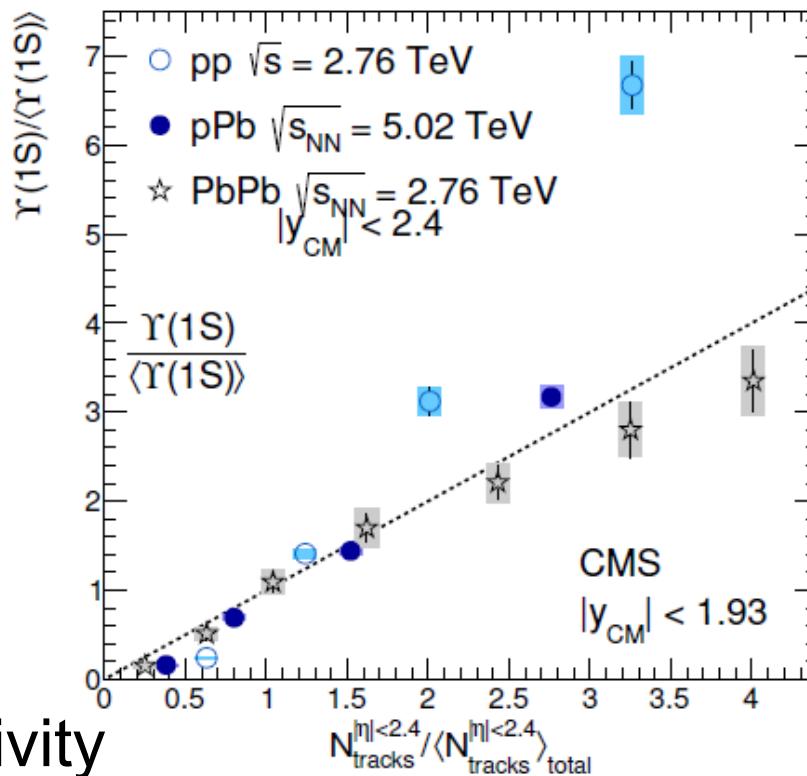


Υ Self-normalized yields: Far activity



- E_T , Far activity
- Close to linear scaling is observed for all systems.
 - Suppression in PbPb for high E_T : central events.
- pPb, pp follow very closely line with slope 1 (dashed line)
 - Fit gives slope consistent with 1 within errors.
 - All systems, all Υ states.

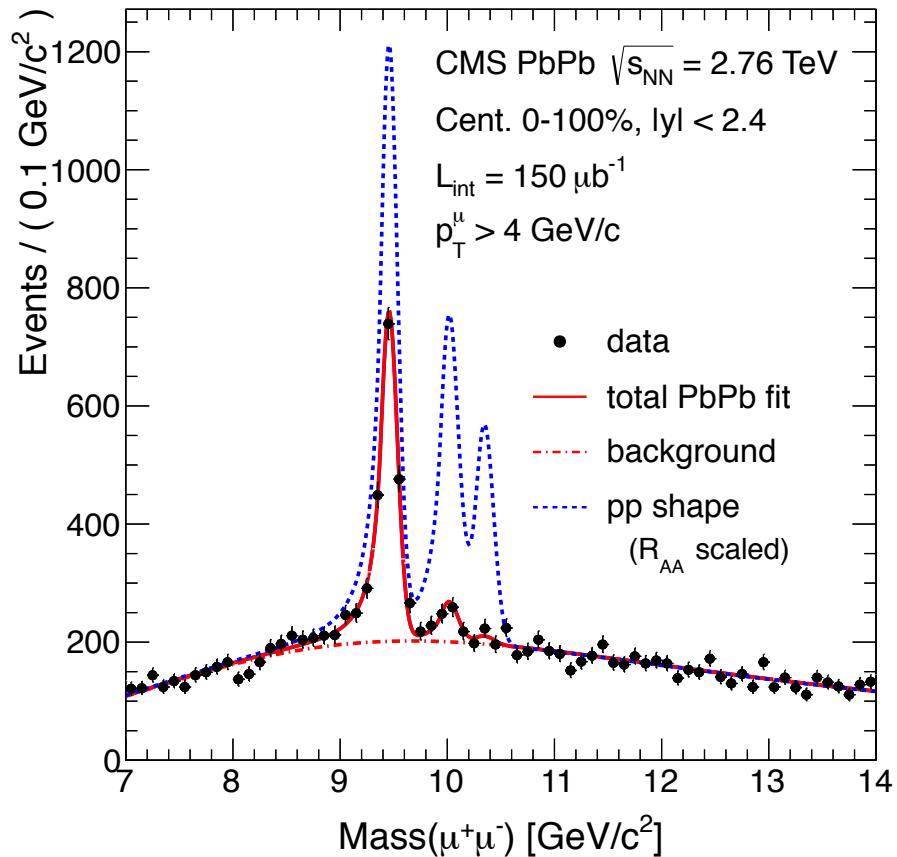
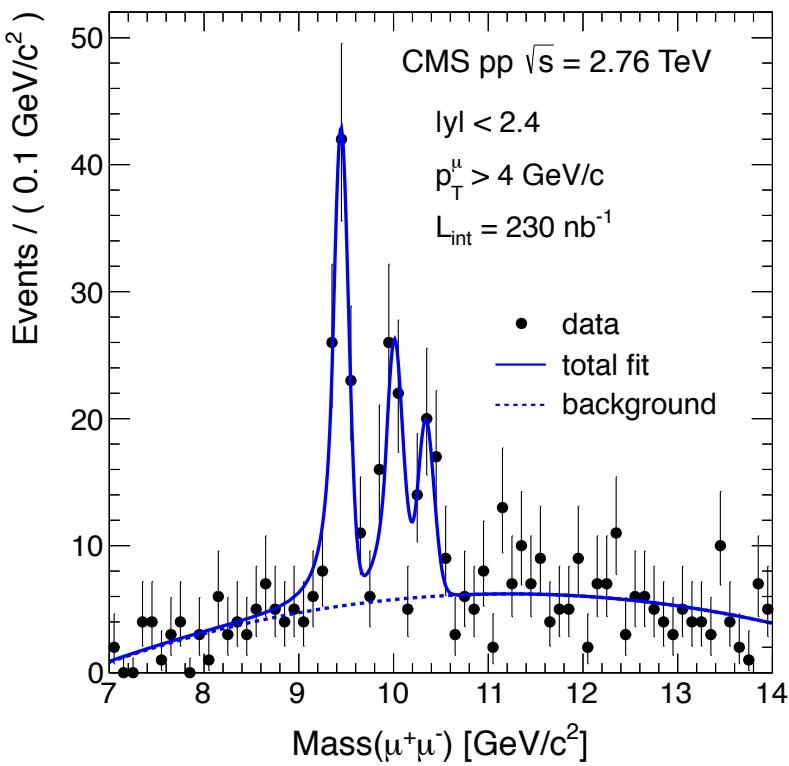
Υ Self-normalized yields: Near activity



- N_{tracks} , Near activity
- Significant differences among systems and among states!
 - $\Upsilon(1S)$ production scaling: stronger than linear in pp.
- All states, even in pp, regardless of whether activity is far or near, show increasing relative yield in higher activity events.
 - Multi-parton interactions in pp?



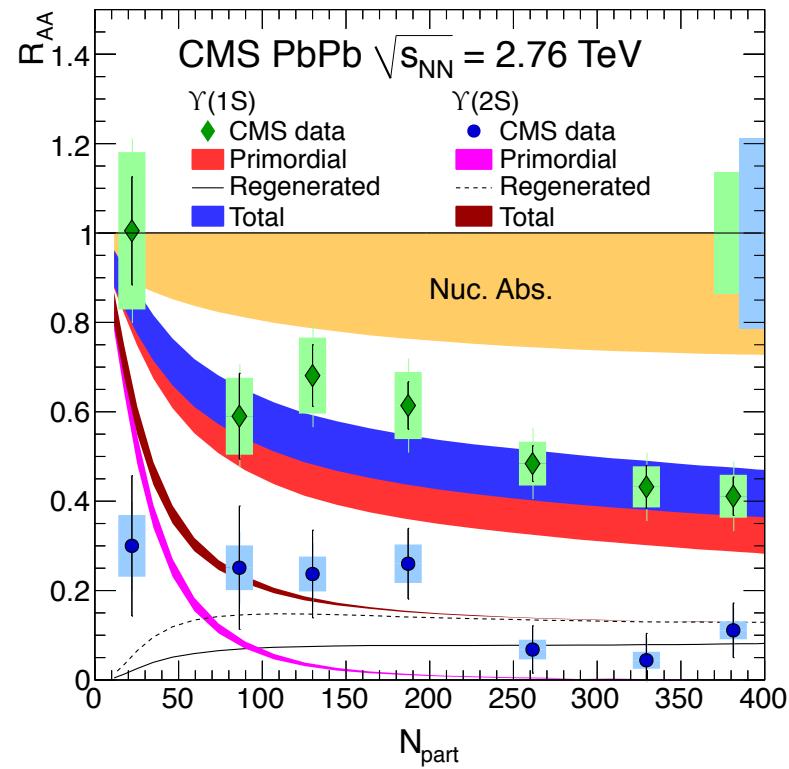
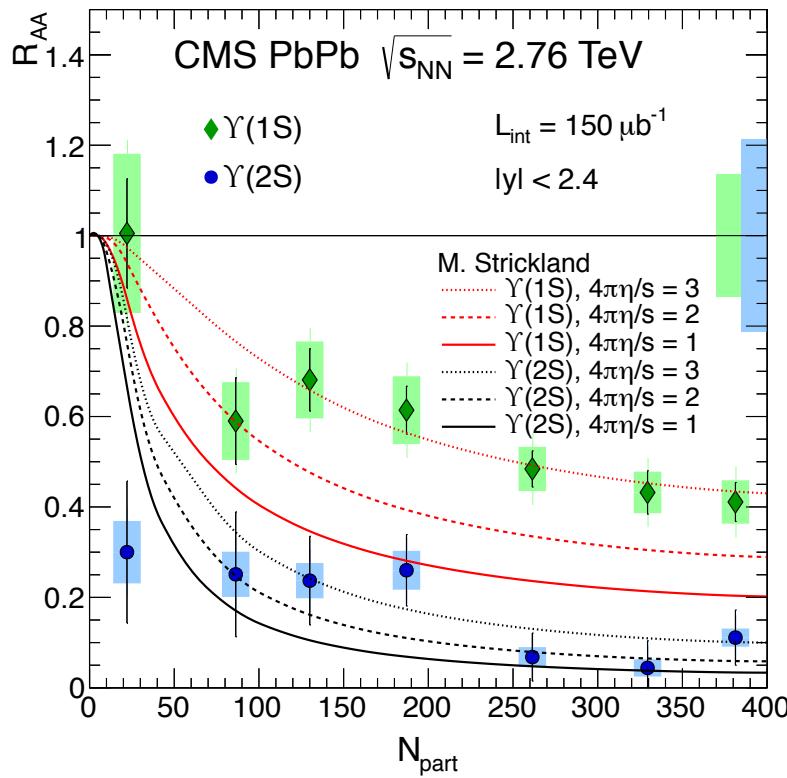
Υ in CMS PbPb



CMS: PRL 109 222301 (2012)

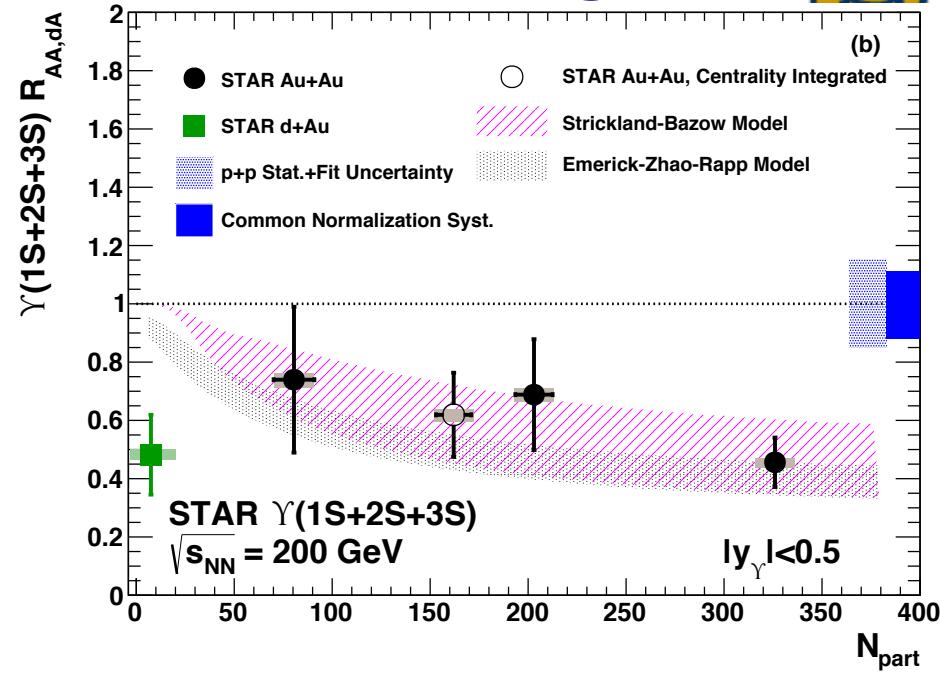
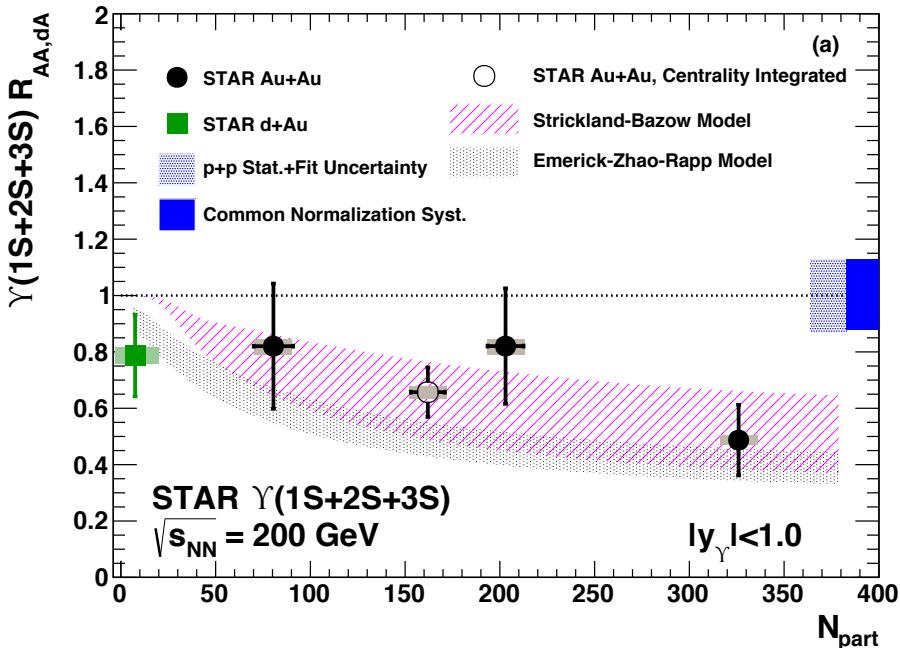
- CMS: Clear suppression of all states in PbPb compared to pp.

CMS Υ R_{AA} in PbPb and models



- Strickland et al. : Nucl. Phys. A 879 (2012) 25
 - Incorporate Re and Im Potential, insert into hydrodynamic expansion, include feed-down contributions.
- Emerick et al. : EPJ A (2012) 48:72
 - Attempt to model binding behavior, test strong/weak binding, kinetic transport model, includes production of Upsilon via “regeneration” of uncorrelated bottom quarks, cold-nuclear-matter absorption.
- Both models consistent with data.
 - But now we have more information, e.g. feeddown contribution is not 50%. What does this do to R_{AA} in models?

STAR γ R_{AA} vs. Centrality



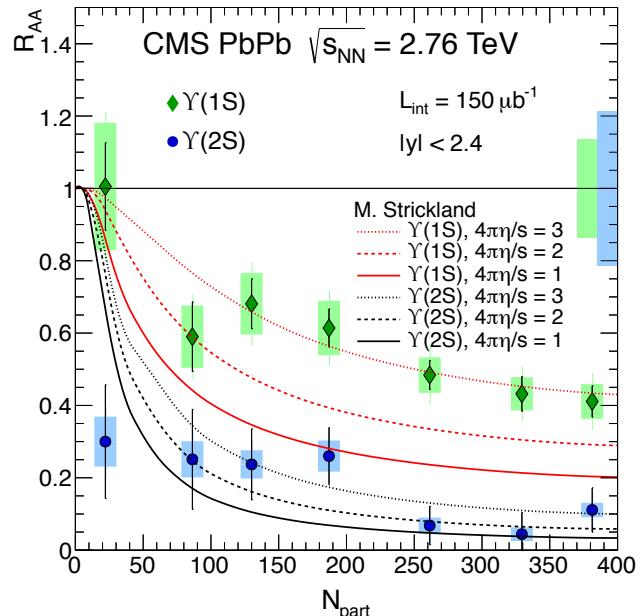
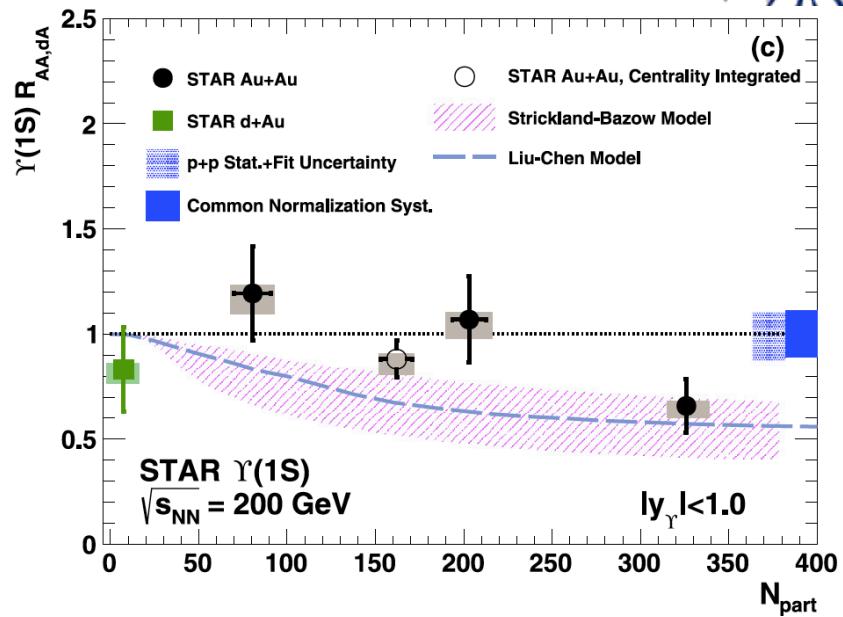
PLB 735 (2014) 127

- Left: all data in STAR acceptance $|y| < 1$
 - dAu, and two most peripheral bins: consistent with no suppression
 - Suppression most central Au+Au: Consistent with expectations for hot & cold nuclear matter, however...
- Right: bin closest to midrapidity, $|y| < 0.5$
 - dAu suppression is of the same magnitude as central AuAu: Important to understand dAu system
- Calculations:
 - Strickland & Bazow: Includes estimate of heavy quarkonium potential, Re and Im. Models evolution through anisotropic hydro.
 - Nucl. Phys. A 879 (2012) 25
 - Emerick, Zhao & Rapp: attempt to include both Hot & Cold nuclear effects.
 - Eur. Phys. J. A (2012) 48: 72

STAR & CMS data and model parameters



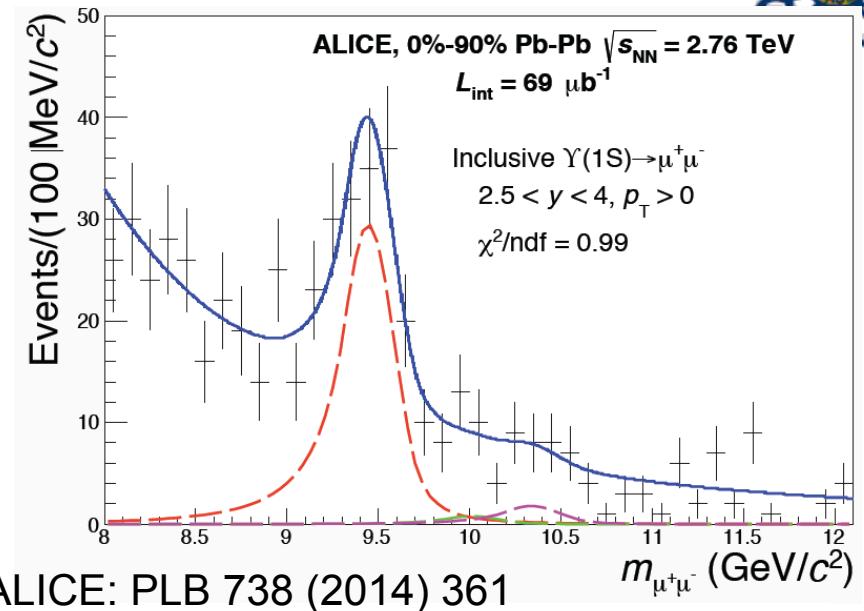
- Strickland & Bazow:
 - Incorporating lattice-based potentials,
 - including real and imaginary parts
 - Includes sequential melting and feed-down contributions
 - Used: $\sim 50\%$ feed-down (what was known at the time)
 - Dynamical expansion, variations in initial conditions (T_0 , η/S)
 - Data indicate:
 - $428 < T_0 < 442$ MeV at RHIC
 - $552 < T_0 < 580$ MeV at LHC
 - for $3 > 4\pi\eta/S > 1$
- Emerick, Zhao, Rapp:
 - Kinetic Theory Model
 - Rate Equation: dissociation + regeneration
 - Fireball model: T evolution.
 - $T_0 \sim 300$ MeV at RHIC
 - $T_0 \sim 600$ MeV at LHC
- All models have T_0 well above $T_c \sim 150\text{-}160$ MeV



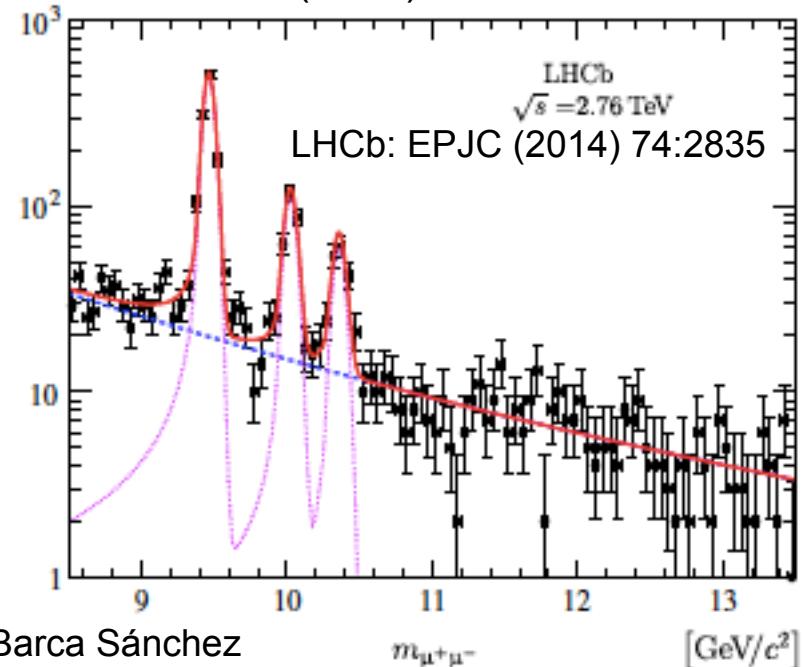


ALICE, LHCb: Υ R_{AA} forward

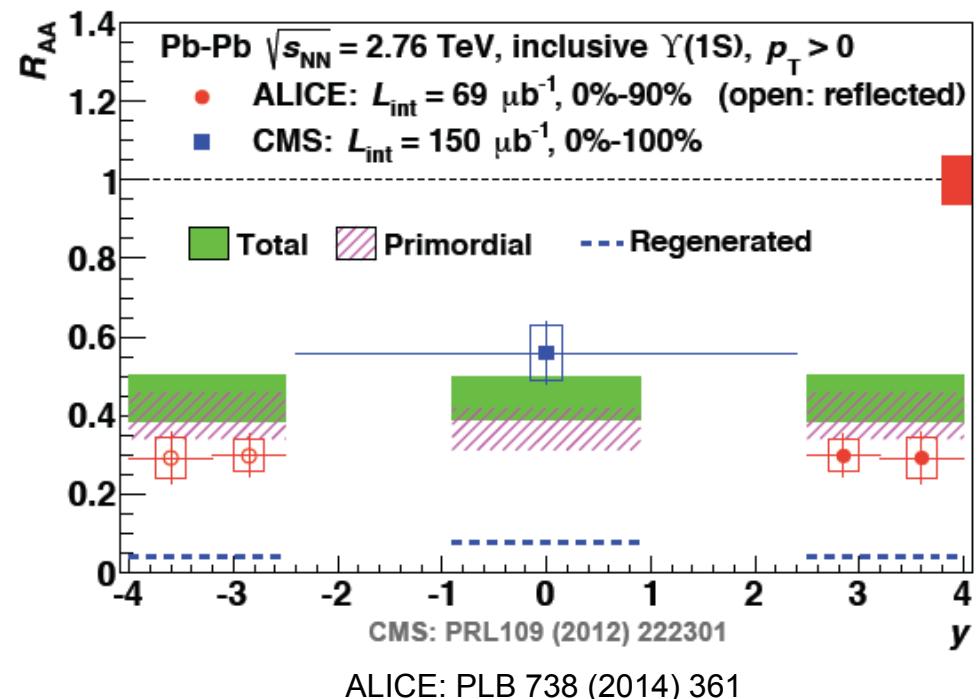
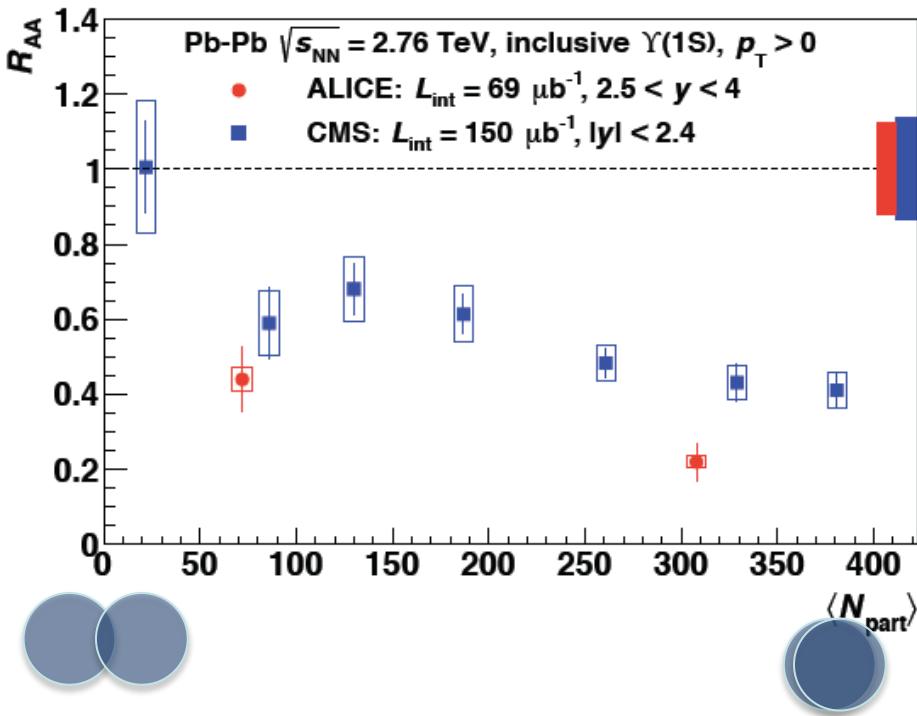
- ALICE Measures Υ in PbPb
 - Forward rapidity region
 - $2.5 < y < 4$
 - Note: CMS, $|y| < 2.4$
 - Fit to 1S to extract yield in PbPb
- pp Reference: LHCb
 - $2.0 < y < 4.5$



ALICE: PLB 738 (2014) 361



Υ R_{AA} at forward y

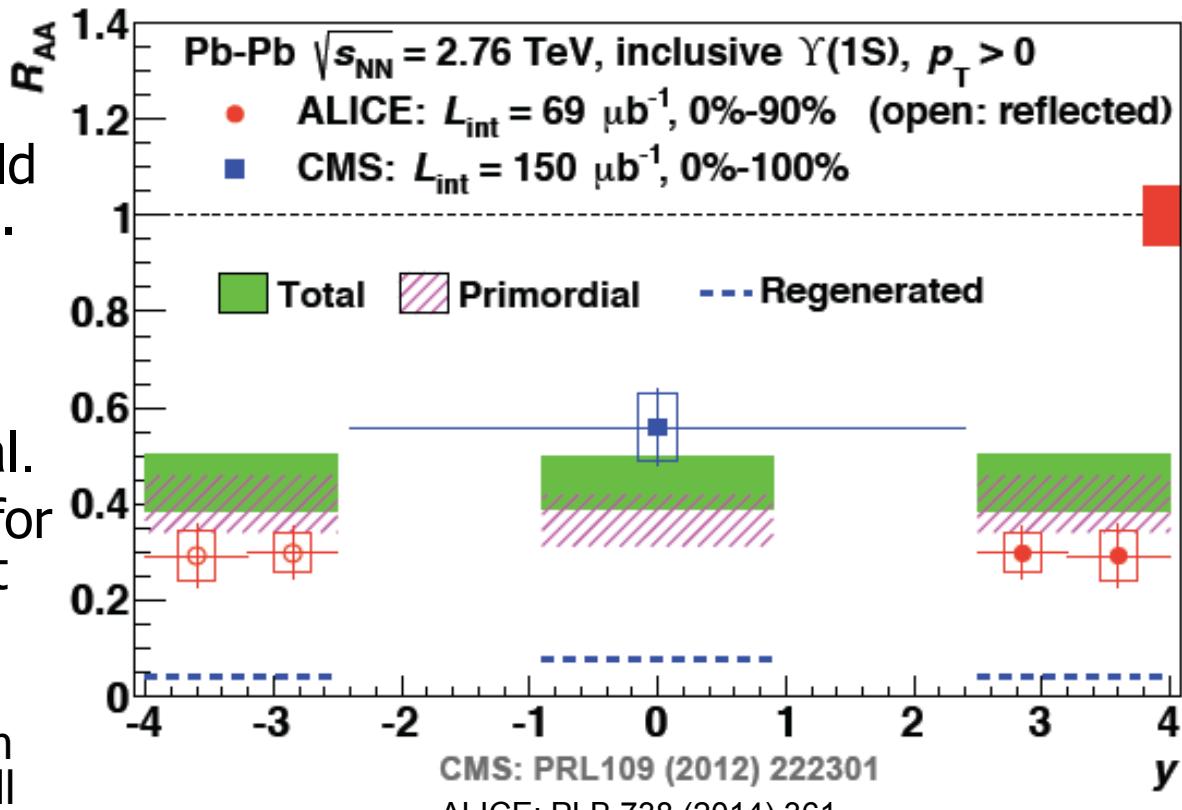


- Comparison between CMS and ALICE/LHCb
 - Υ R_{AA}: more suppression at forward rapidities!
 - Energy density, T should be smaller at forward y.
 - Argues for less suppression, not more.

Comparison to transport/regeneration



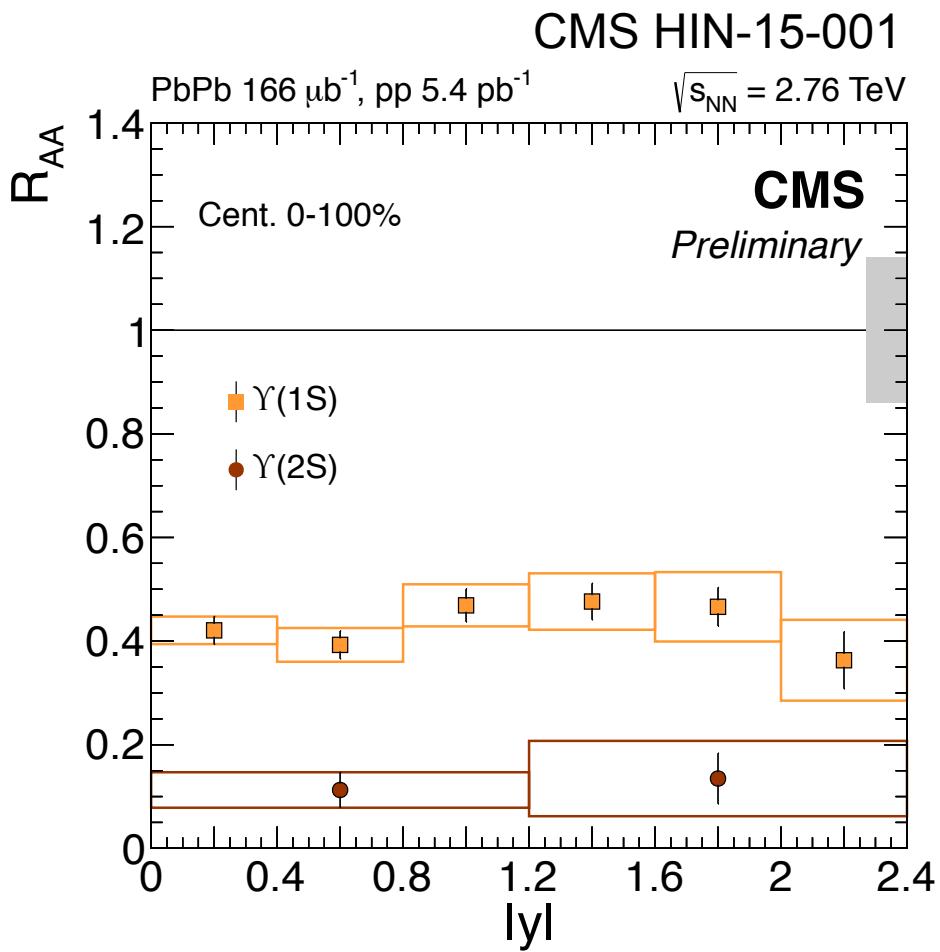
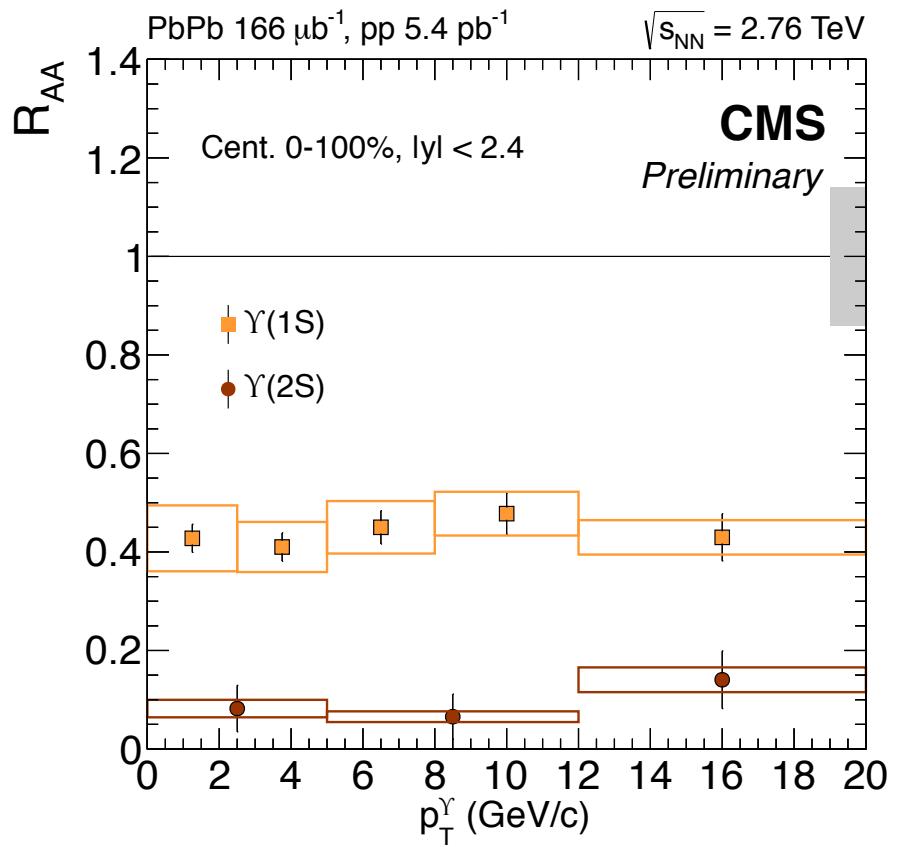
- More suppression at forward rapidity.
- Energy density, T should be smaller at forward y.
 - Argues for less suppression, not more.
- Model from Emerick et al.
 - Model cannot account for stronger suppression at forward rapidity
 - Includes a regeneration component, albeit small
 - Includes absorption component



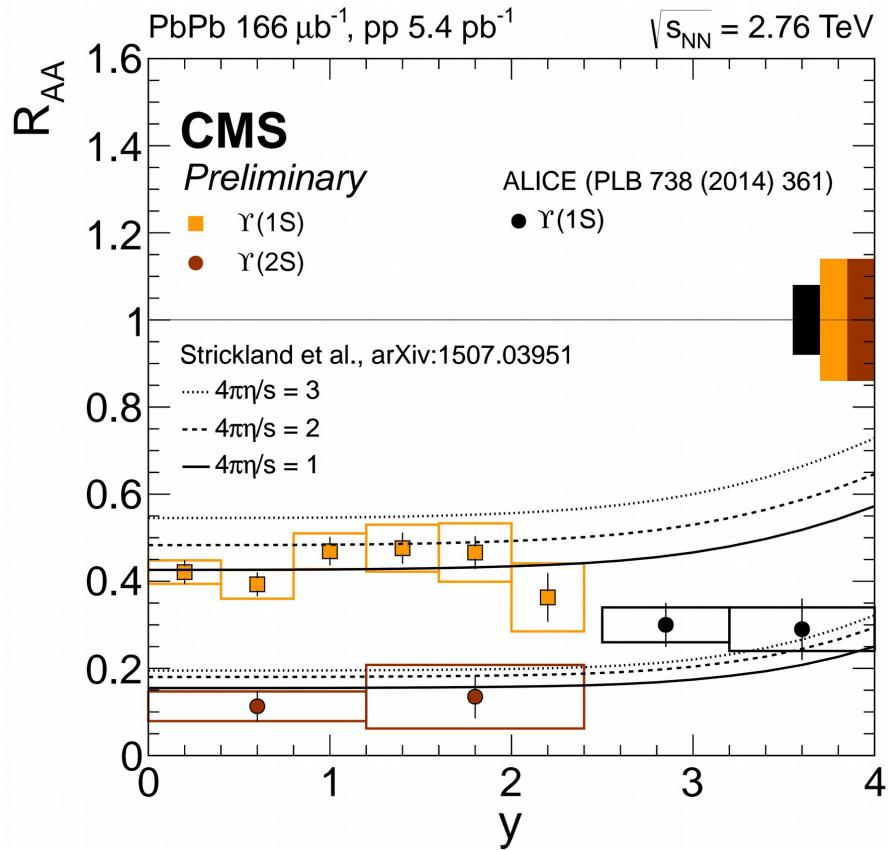
CMS: Kinematic dependence of R_{AA}



- Higher stats. for pp @ 2.76 allow finer bins in p_T and y



CMS and ALICE: Υ rapidity dependence

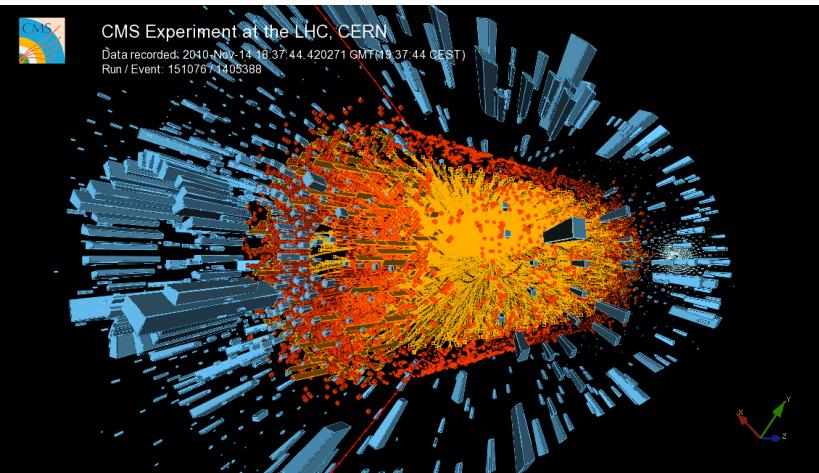
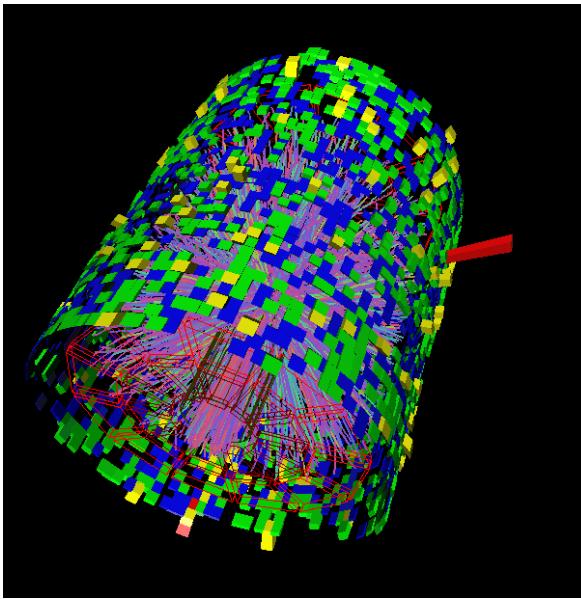


ALICE:
Phys. Lett. B 738 (2014) 361
CMS Preliminary:
CMS-PAS-HIN-15-001

- CMS and ALICE data are consistent near their acceptance limits.
 - Observe a flat or falling rapidity dependence of Upsilon suppression.
- Comparison to Strickland & Bazow model, with Re & Im V(T) + hydro:
 - R_{AA} increases at forward rapidities. (Lower energy density/Temperature)



The bottom line...



- pA collisions:
 - Modifications of production compared to pp observed.
 - Shadowing effects, E.Loss.
 - But not consistent with all data
 - dAu RHIC, LHCb vs. ALICE pPb.
- AA collisions:
 - Large suppression of Upsilon production seen.
 - Need quantitative comparisons pA vs. AA to shed disentangle hot vs. cold nuclear matter.
 - Rapidity dependence: not expected.
 - Goal: Pinning down the medium properties, e.g. Temperature.



Results/References

- Fermilab
 - E772, pA: PRL 66 (1991) 2285
- RHIC
 - STAR, pp, dAu, AuAu: PLB 735 (2014) 127
 - PHENIX: PRC 87, 044909 (2013)
- LHC
 - LHCb, pPb: JHEP 07 (2014) 094
 - ATLAS pPb: ATLAS-CONF-2015-050
 - ALICE pPb: PLB 740 (2015) 105
 - CMS pPb: JHEP 04 103 (2014)
- ALICE PbPb: PLB 738 (2014) 361
- CMS PbPb: PRL 109 222301 (2012)

Backup

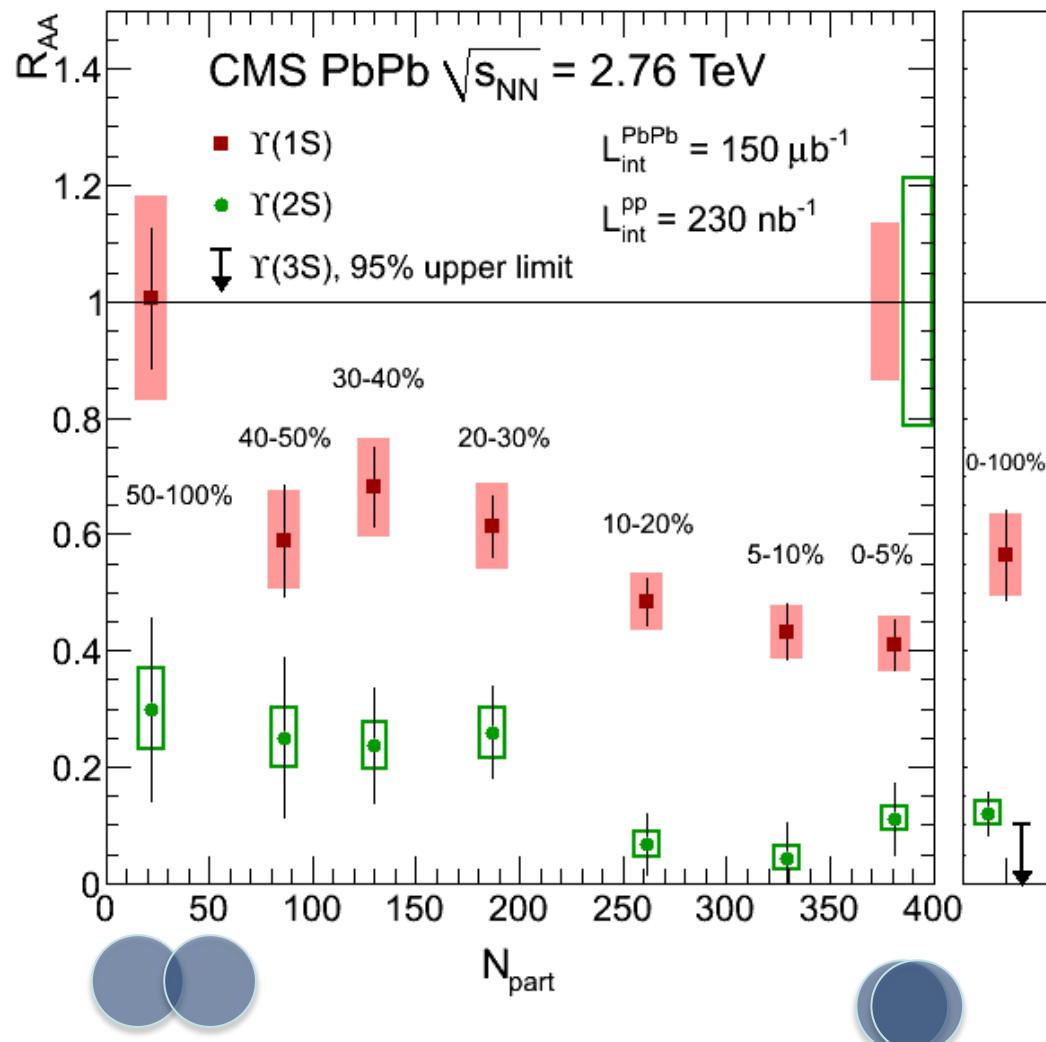


CMS Υ R_{AA} vs. Centrality



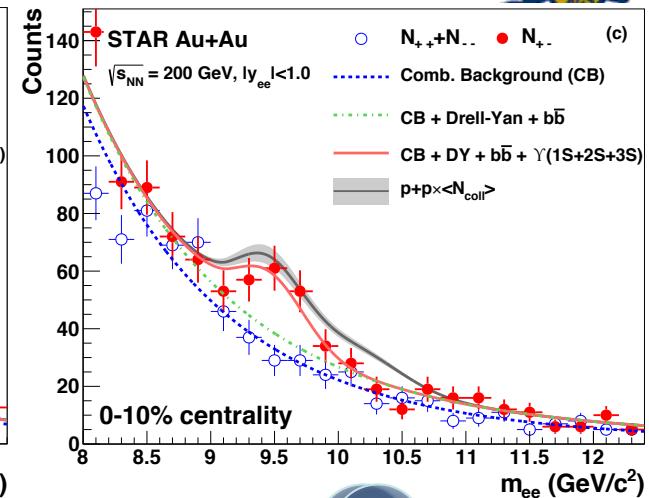
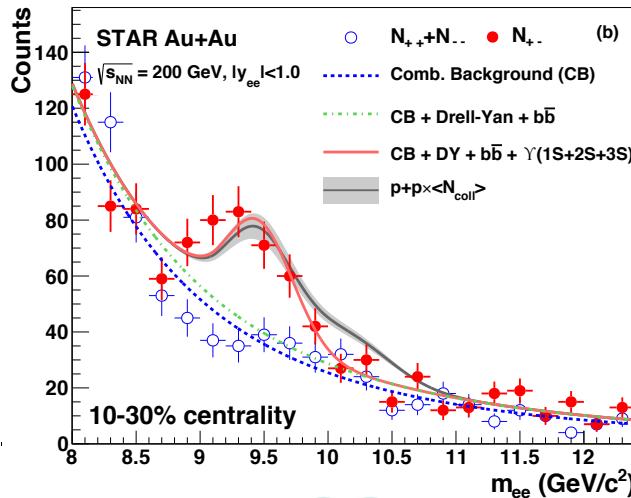
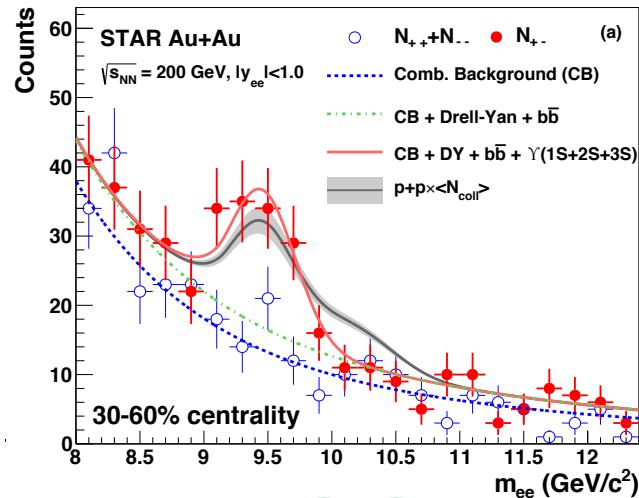
- Centrality integrated:
 - $\Upsilon(1S) : 0.56 \pm 0.08 \pm 0.07$
 - $\Upsilon(2S) : 0.12 \pm 0.04 \pm 0.02$
 - $\Upsilon(3S) : < 0.10 @ 95\% \text{ CL}$
- Observation of sequential suppression.
- Comparison to STAR R_{AA}
 $\Upsilon(1S), |y|<1 :$
 - $0.88 \pm 0.09 \pm 0.13^{+0.03}_{-0.07} \pm 0.11$
 - More suppression at LHC compared to RHIC

PRL 109 222301 (2012)





Υ in STAR, AuAu



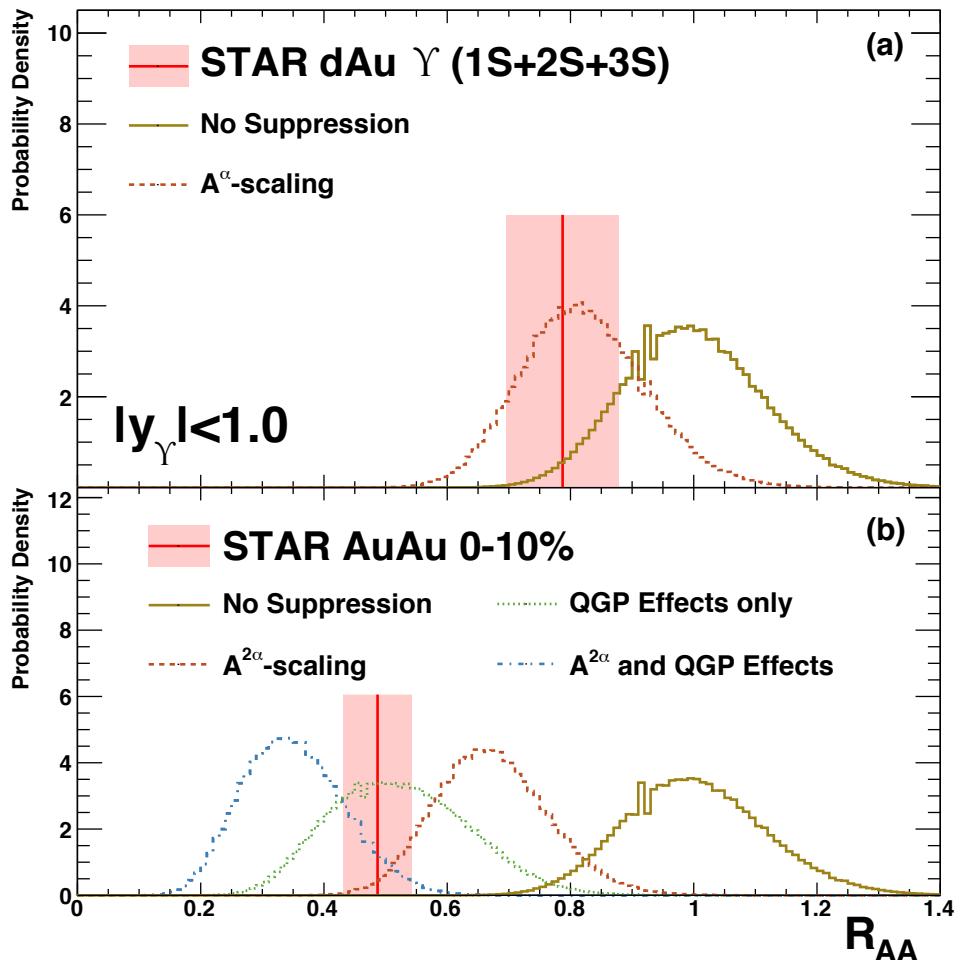
STAR: PLB 735 (2014) 127

- Invariant mass distributions in 3 centrality bins
- Comparison to N_{coll} -scaled pp reference:
 - Clear suppression of excited states.
 - Report results for sum of 3 states, and for ground state.

STAR Hypothesis testing, $|y|<1$



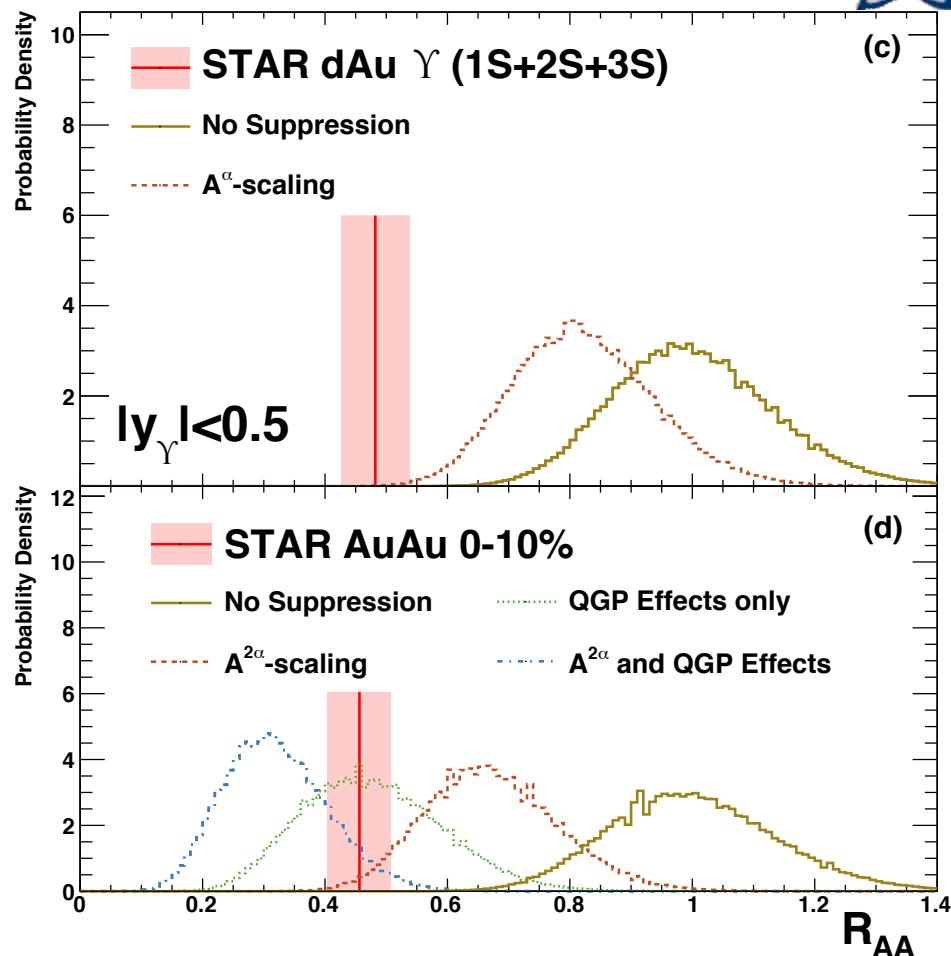
- Measurements: vertical line
 - R_{dAu}
 - R_{AA} , 0-10% most central
 - pink band: syst. unc.
- Hypothesis test:
 - Run pseudoexperiments for various scenarios
 - Stat. unc.: width of distributions
 - No suppression: $RAA=1$
 - A^α scaling for dAu (CNM effect)
 - As seen by E776 at Fermilab
 - $A^{2\alpha}$ for AuAu
 - QGP effects only
 - Based on Strickland et al.
 - QGP effects + A^α scaling
- A^α scaling: consistent with dAu data
- QGP+ A^α scaling: consistent with AuAu data
- Other scenarios are disfavored.



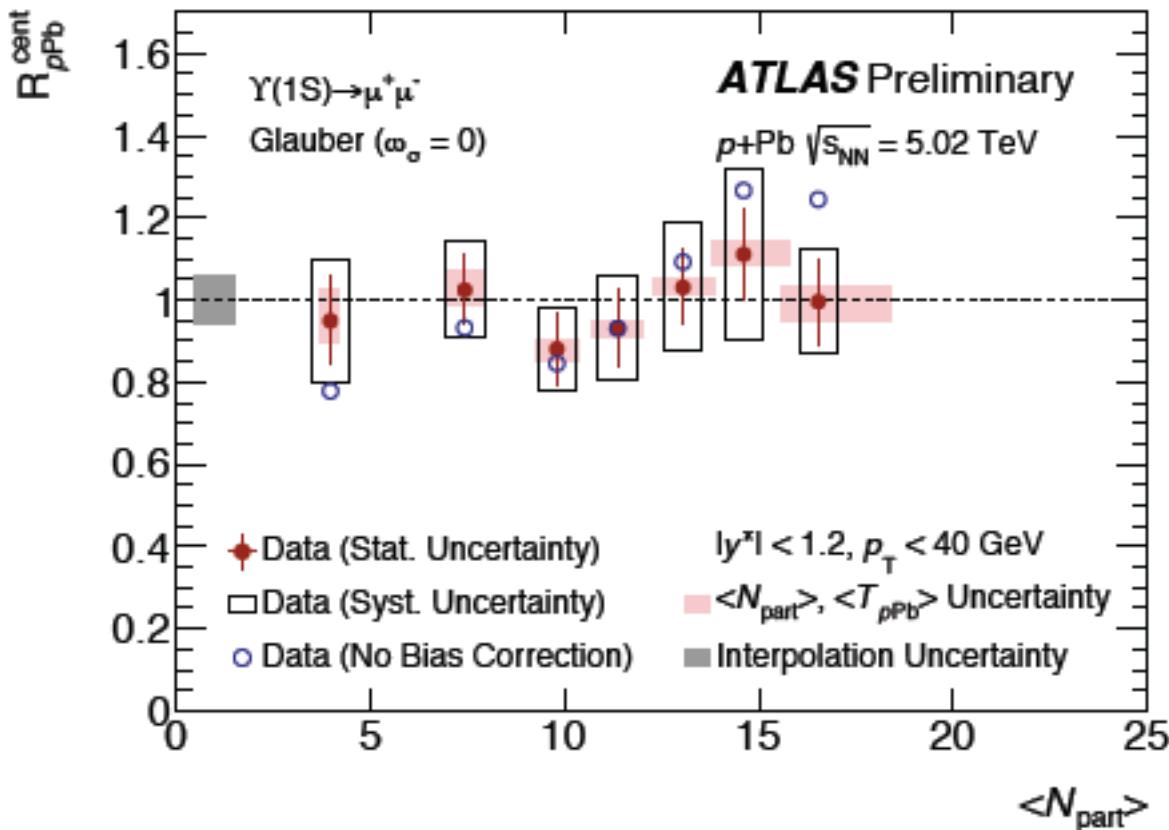
STAR Hypothesis testing, $|y|<0.5$



- Hypothesis tests:
 - No suppression: $RAA=1$
 - A^α scaling for dAu (CNM effect)
 - $A^{2\alpha}$ for AuAu
 - QGP effects only
 - Based on Strickland et al.
 - QGP effects + A^α scaling
- Clear: $|y|<0.5$ shows large suppression in dAu.
 - Comparable to central AuAu!
 - No particular scenario is favored.
 - pAu run at RHIC from 2015, stay tuned!



ATLAS, RpPb $\Upsilon(1S)$ vs. N_{part}



- $\Upsilon(1S)$: No modification seen vs. N_{part} in pPb.
- ATLAS and ALICE do a “bias correction”.

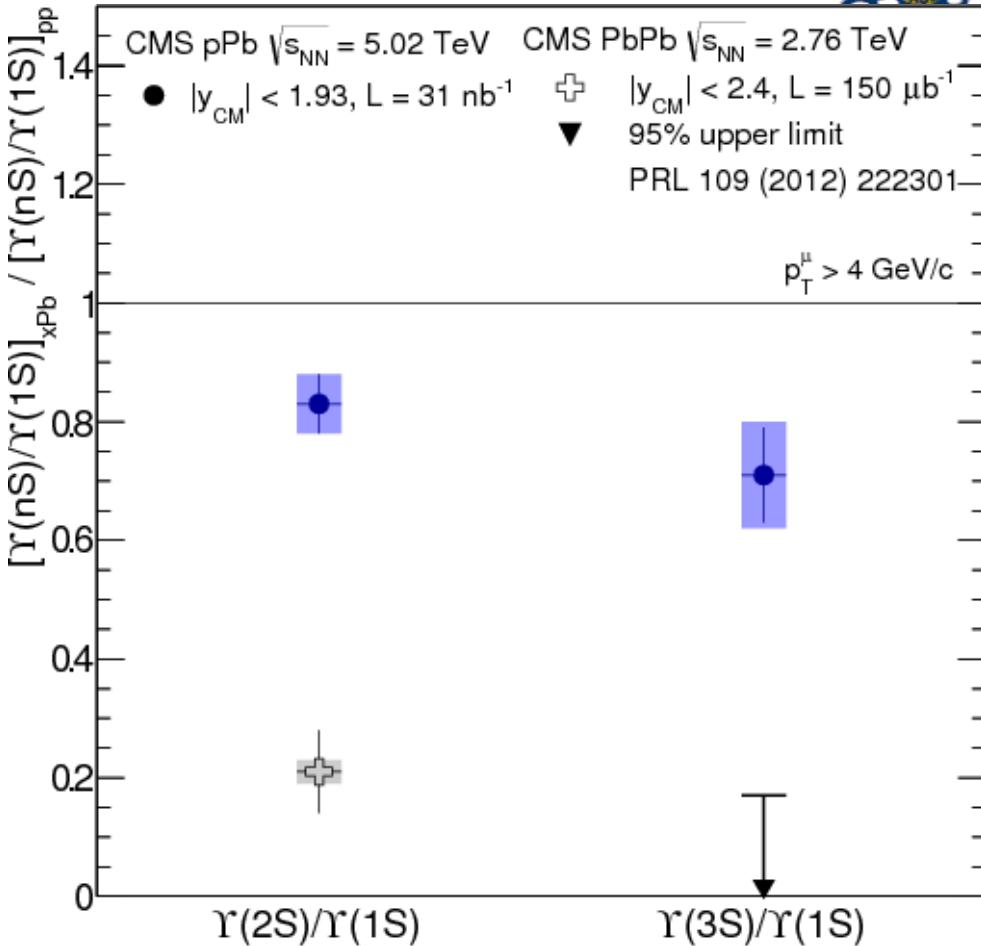


Υ Double ratio, pPb and pp

$$\left[\frac{\Upsilon(nS)}{\Upsilon(1S)} \right]_{pPb} = \frac{R_{pPb}(\Upsilon(nS))}{R_{pPb}(\Upsilon(1S))}$$

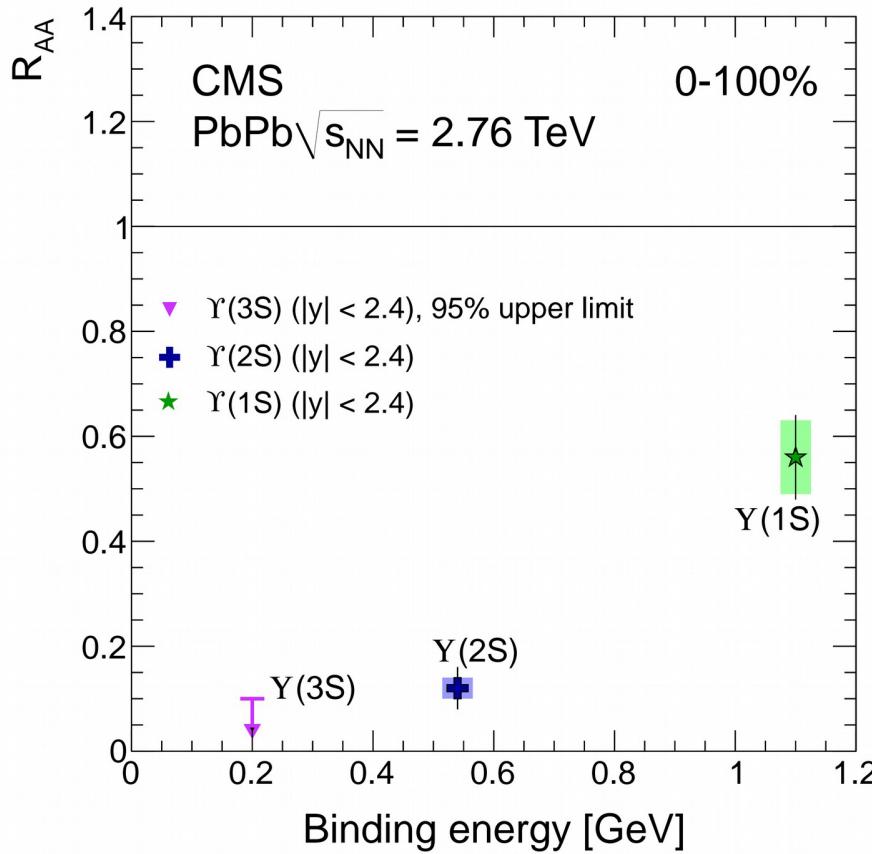
$$\left[\frac{\Upsilon(nS)}{\Upsilon(1S)} \right]_{pp}$$

- Key feature of double ratio:
 - All initial-state effects should cancel
 - e.g. shadowing modifies excited and ground state in the same way
- Observation:
 - Double ratio < 1 in pPb
 - Double ratio in pPb higher than in PbPb
 - Similar for 2S than for 3S
- Implication: possible presence of final state effects in pPb which affect excited states more than ground state



Expect small Υ absorption in hadronic matter
 Lin & Ko, PLB 503 (2001) 104
 But absorption depends on radius:
 Larger absorption for 2S and 3S

Υ Summary plot vs. binding energy



- Overall pattern of sequential suppression is observed.
 - But there are important details that indicate additional physics...