

# Conformal viscous hydrodynamics - *AdS/CFT correspondence at finite temperature*

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## CONTENT

- motivation: viscous effects in RHIC data  
viscosity  $\eta$  - relaxation time  $\tau_\pi, \dots$
- conformal invariance,  
relativistic hydrodynamics -  
second-order gradients
- gauge/gravity duality -  
(AdS/CFT) correspondence

A fluid which has no shear stresses, viscosity or heat conduction is called a

## **PERFECT FLUID**

i.e. it looks isotropic in its rest frame

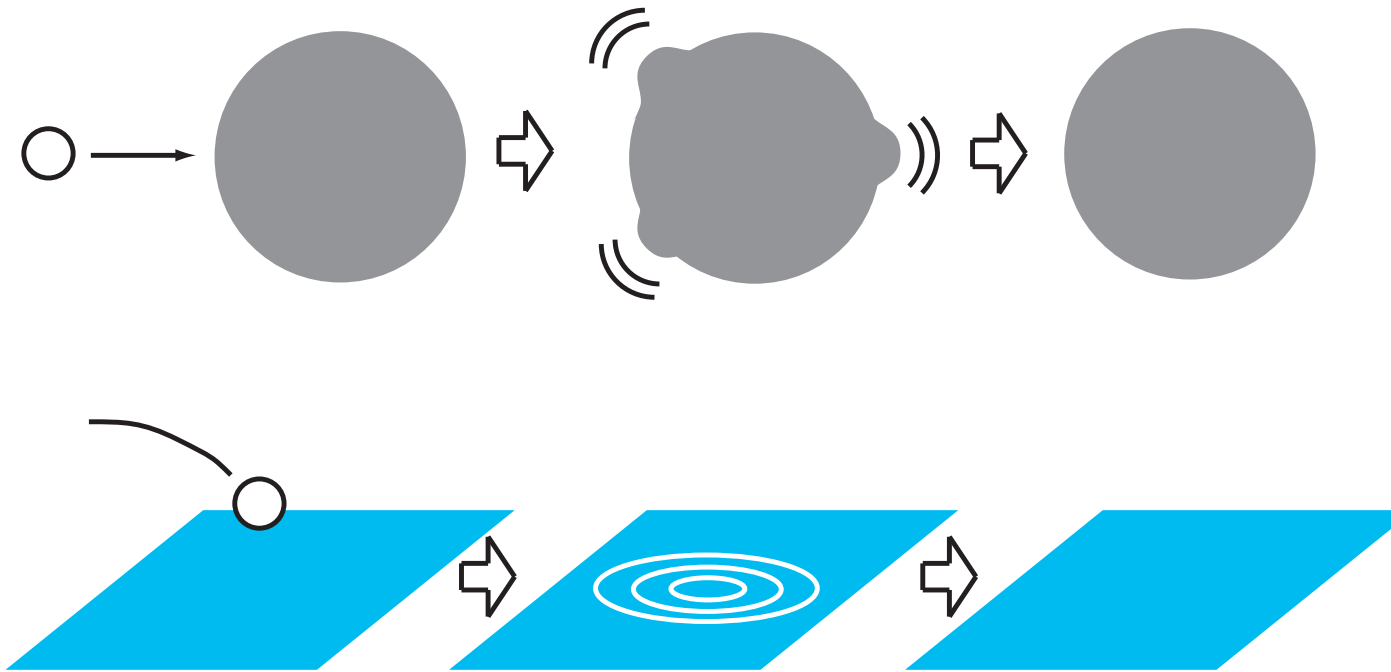
Top physics story of 2005 is the RHIC discovery of the strongly interacting quark-gluon plasma (called **sQGP**), which behaves almost like a perfect fluid, with very low viscosity

[T. D. Lee 06 ]

## **black hole theory**

is now used to explain properties of colliding nuclei

[L. Susskind 08]



[from M. Natsuume]

quasinormal modes:

gravitational perturbation to a black hole  
and to a hydrodynamic system

# hydrodynamics

**energy momentum tensor**:  $\epsilon$  energy density and  $p$  pressure

$$T^{\mu\nu} = (\epsilon + p)u^\mu u^\nu + pg^{\mu\nu} + \Pi^{\mu\nu}$$

$u^\mu$  fluid velocity,  $u_\mu u^\mu = -1$ , and  $\Pi^{\mu\nu}$  shear tensor  
with

$$u_\mu \Pi^{\mu\nu} = 0 \quad (u_\mu T^{\mu\nu} = -\epsilon u^\nu),$$

momentum density is due to energy flow, and

$$\Pi^\mu{}_\mu = 0$$

required from conformal invariance

**local conservation law (covariant derivative  $\nabla_\mu$ ):  $\nabla_\mu T^{\mu\nu} = 0$**

(assume: no net charge in the system)

## a few definitions

conformal fluid in  $d = 4$  (curved) dimensions -

$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu$$

metric  $g_{\mu\nu}$  with signature  $(- + + +)$ ,

$\mu, \nu..$  space-time indices 0, 1, 2, 3

(geometric) covariant derivative:  $\nabla_\mu$

$$\begin{aligned} \nabla_\mu \Phi(x) &= \partial_\mu \Phi(x) , \\ \nabla_\mu V^\nu &= \partial_\mu V^\nu + \Gamma_{\mu\rho}{}^\nu V^\rho \quad \text{etc., e.g. } \nabla_\mu g^{\nu\rho} = 0 \end{aligned}$$

Christoffel symbols:

$$\Gamma_{\mu\rho}{}^\nu = \frac{1}{2} g^{\nu\sigma} (\partial_\mu g_{\rho\sigma} + \partial_\rho g_{\mu\sigma} - \partial_\sigma g_{\mu\rho})$$

Riemann tensor:  $[\nabla_\mu, \nabla_\nu] V^\rho = R_{\mu\nu\sigma}{}^\rho V^\sigma$

# approximation

only retaining shear viscosity terms

- **Navier - Stokes** = first-order theory in gradients with constitutive relation:

$$\Pi^{\mu\nu} = -2\eta \langle \nabla^\mu u^\nu \rangle \equiv -2\eta \sigma^{\mu\nu}$$

with the shear strain tensor

$$\sigma^{\mu\nu} \equiv \frac{1}{2} (\Delta^{\mu\rho} \nabla_\rho u^\nu + \Delta^{\nu\rho} \nabla_\rho u^\mu) - \frac{1}{3} \Delta^{\mu\nu} \Delta^{\rho\sigma} \nabla_\rho u_\sigma$$

and projection

$$\Delta^{\mu\nu} = g^{\mu\nu} + u^\mu u^\nu, \quad \Delta^{\mu\nu} u_\nu = 0$$

## Navier-Stokes

projection  $\Delta_{\alpha}^{\mu} \nabla_{\beta} T^{\alpha\beta} = 0 \rightarrow$   
relativistic Navier-Stokes equation in first-order theory

$$(\epsilon + p)u^{\alpha} \nabla_{\alpha} u^{\mu} + \Delta^{\mu\alpha} \nabla_{\alpha} p + \Delta_{\alpha}^{\mu} \nabla_{\beta} [-2\eta < \nabla^{\alpha} u^{\beta} >] = 0$$

i.e. **parabolic equation:**

time derivative is of first order ( $u^{\alpha} \nabla_{\alpha} \rightarrow \partial/\partial t$ )

**while**

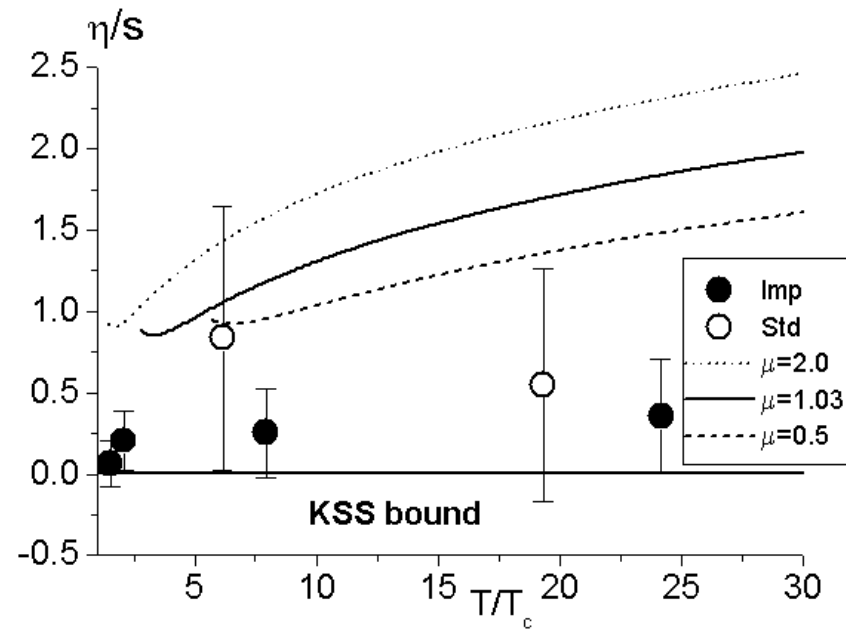
space derivative is of second order ( $\nabla^2$ )

**“relativistic first-order dissipative theory is highly pathological, and therefore should be discarded in favor of the second-order one”**

[Hiscock and Lindblom 1983-1985]



# viscosity

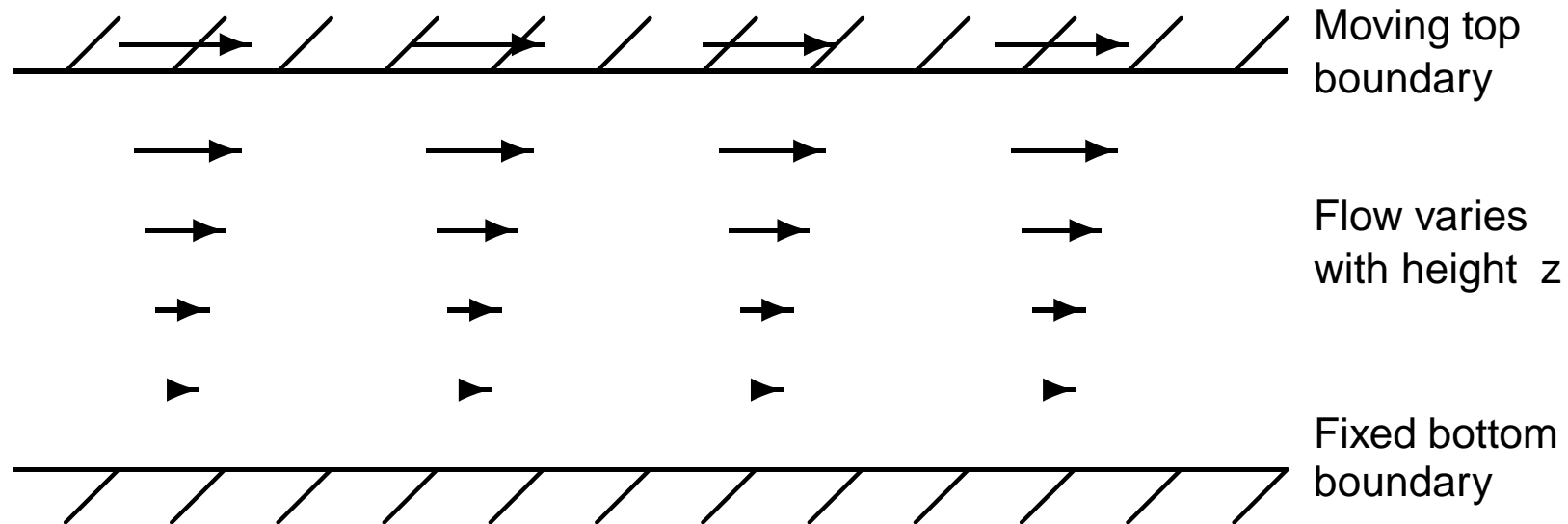


$\eta/s$  from quenched QCD **lattice simulations** [Nakamura and Sakai 05]  
compared with the **perturbative result**

**for strong coupling**  $\lambda = g_{YM}^2 N_c$  [Kovtun, Son and Starinets 05]  
**AdS/CFT universal (?) lower bound**

$$\frac{\eta}{s} \geq \frac{1}{4\pi}$$

# shear viscosity $\eta$



shear flow in  $x$ - direction  $\rightarrow$  **force per unit area/momentum transfer**

$$\frac{F_x}{A} = -\Pi_{xz} = \eta \frac{\partial u_x}{\partial z}$$

measure how quickly a system, perturbed from thermal equilibrium, goes back: it takes longer for a weakly coupled system, because particles interact less than for a strongly coupled system

## transverse mode

small linear perturbation and first-order →  
diffusion equation in shear channel:

$$\delta u_{\perp} = \sqrt{\frac{(\epsilon + p)}{4\pi\eta t}} \exp\left[-\frac{(\epsilon + p)x^2}{4\eta t}\right]$$

i.e. propagates outside the light-cone (Gaussian)  
starting from  $\delta u_{\perp}(x, t = 0) = \delta(x)$  !!

minimal modification → second-order:  
relaxation time  $\tau_{\pi} > 0$  → hyperbolic equation

$$\left[\tau_{\pi}\partial_t^2 + \partial_t - \frac{\eta}{(\epsilon + p)}\partial_x^2\right] \delta u_{\perp} = 0$$

# conformal hydrodynamics

mainly based on more recent work by:

R. Baier, P. Romatschke, D. T. Son, A. O. Starinets and  
M. A. Stephanov

“Relativistic viscous hydrodynamics, conformal invariance,  
and holography”

JHEP 0804 (2008) 100 (arXiv: 0712.2451 [hep-th])

and

R. Loganayagam

“Entropy current in conformal hydrodynamics”

JHEP 0805 (2008) 087 (arXiv: 0801.3701 [hep-th])

and references therein

## a new result

all second order terms classified by conformal symmetry,

Weyl transformation:  $g^{\mu\nu} = e^{-2\phi(x)} \tilde{g}^{\mu\nu}$ ,

$$T^{\mu\nu} = e^{-6\phi} \tilde{T}^{\mu\nu}, \quad T_{\mu}^{\mu} = 0, \dots$$

constitutive relation (d= 4):

$$\begin{aligned} \Pi^{\mu\nu} = & -2\eta\sigma^{\mu\nu} + 2\eta\tau_{\pi}u^{\lambda}\mathcal{D}_{\lambda}\sigma^{\mu\nu} \\ & + 4\lambda_1\sigma^{<\mu}_{\lambda}\sigma^{\nu>\lambda} + 2\lambda_2\sigma^{<\mu}_{\lambda}\Omega^{\nu>\lambda} + \lambda_3\Omega^{<\mu}_{\lambda}\Omega^{\nu>\lambda} \\ & + 2\kappa u_{\alpha}C^{\alpha\mu\nu\beta}u_{\beta} \end{aligned}$$

$C^{\alpha\beta\gamma\delta}$  ... Weyl tensor,  $\Omega^{\alpha\beta}$  ... antisymmetric vorticity tensor

[Baier, Romatschke, Son, Starinets, Stephanov 07]

[Bhattacharyya, Hubeny, Minwalla, Rangamani 07, Loganayagam 08]

## Weyl transformation

$$g_{\mu\nu} = e^{2\phi} \tilde{g}_{\mu\nu}; \quad g^{\mu\nu} = e^{-2\phi} \tilde{g}^{\mu\nu}, \quad \phi = \phi(x)$$

leads to:

$$\Gamma_{\lambda\mu}{}^{\nu} = \tilde{\Gamma}_{\lambda\mu}{}^{\nu} + \delta_{\lambda}^{\nu} \partial_{\mu} \phi + \delta_{\mu}^{\nu} \partial_{\lambda} \phi - \tilde{g}_{\lambda\mu} \tilde{g}^{\nu\sigma} \partial_{\sigma} \phi$$

$$u^{\mu} = e^{-\phi} \tilde{u}^{\mu}, \quad \Delta^{\mu\nu} = g^{\mu\nu} + u^{\mu} u^{\nu} = e^{-2\phi} \tilde{\Delta}^{\mu\nu}$$

usefull quantities (d dimensions):

$$\vartheta \equiv \nabla_{\mu} u^{\mu} = e^{-\phi} \left[ \tilde{\vartheta} + (d-1) \tilde{u}^{\sigma} \partial_{\sigma} \phi \right],$$

$$a^{\nu} \equiv u^{\mu} \nabla_{\mu} u^{\nu} = e^{-2\phi} \left[ \tilde{a}^{\nu} + \tilde{\Delta}^{\nu\sigma} \partial_{\sigma} \phi \right],$$

$$\mathcal{A}_{\nu} \equiv a_{\nu} - \frac{\vartheta}{d-1} u_{\nu} = \tilde{\mathcal{A}}_{\nu} + \partial_{\nu} \phi$$

## Weyl derivative

cf. (abelian) gauge transform:

$$\tilde{\psi} = e^{i\Lambda(x)} \psi, \quad \tilde{A}^\mu = A^\mu - \partial^\mu \Lambda,$$

$$\tilde{D}^\mu \tilde{\psi} = e^{i\Lambda} (\partial^\mu + iA^\mu) \psi = e^{i\Lambda} D^\mu \psi$$

Define covariant Weyl derivative for:

$$Q_{\nu\dots}^{\mu\dots} = e^{-w\phi} \tilde{Q}_{\nu\dots}^{\mu\dots}, \quad \mathcal{D}_\lambda Q_{\nu\dots}^{\mu\dots} = e^{-w\phi} \tilde{\mathcal{D}}_\lambda \tilde{Q}_{\nu\dots}^{\mu\dots}$$

where

$$\begin{aligned} \mathcal{D}_\lambda Q_{\nu\dots}^{\mu\dots} &\equiv \nabla_\lambda Q_{\nu\dots}^{\mu\dots} + w A_\lambda Q_{\nu\dots}^{\mu\dots} \\ &+ [g_{\lambda\alpha} \mathcal{A}^\mu - \delta_\lambda^\mu \mathcal{A}_\alpha - \delta_\alpha^\mu \mathcal{A}_\lambda] Q_{\nu\dots}^{\alpha\dots} + \dots \\ &- [g_{\lambda\nu} \mathcal{A}^\alpha - \delta_\lambda^\alpha \mathcal{A}_\nu - \delta_\nu^\alpha \mathcal{A}_\lambda] Q_{\alpha\dots}^{\mu\dots} - \dots \end{aligned}$$



## examples

$$\mathcal{D}_\lambda g_{\mu\nu} = 0, \quad \mathcal{D}_\mu u^\mu = 0, \quad u^\mu \mathcal{D}_\mu u^\nu = 0$$

and

$$\mathcal{D}_\mu u^\nu = \nabla_\mu u^\nu + u_\mu a^\nu - \frac{\vartheta}{d-1} \Delta_\mu{}^\nu = \sigma_\mu{}^\nu + \Omega_\mu{}^\nu = e^{-\phi} \tilde{\mathcal{D}}_\mu \tilde{u}^\nu,$$

$$\begin{aligned} \sigma^{\mu\nu} &\equiv \frac{1}{2} \left( \Delta^{\mu\lambda} \nabla_\lambda u^\nu + \Delta^{\nu\lambda} \nabla_\lambda u^\mu \right) - \frac{1}{d-1} \vartheta \Delta^{\mu\nu} = \frac{1}{2} (\mathcal{D}^\mu u^\nu + \mathcal{D}^\nu u^\mu) \\ &= e^{-3\phi} \tilde{\sigma}^{\mu\nu}, \end{aligned}$$

$$\Omega^{\mu\nu} \equiv \frac{1}{2} \left( \Delta^{\mu\lambda} \nabla_\lambda u^\nu - \Delta^{\nu\lambda} \nabla_\lambda u^\mu \right) = \frac{1}{2} (\mathcal{D}^\mu u^\nu - \mathcal{D}^\nu u^\mu) = e^{-3\phi} \tilde{\Omega}^{\mu\nu},$$

temperature  $T$ :

$$T = e^{-\phi} \tilde{T}, \quad \mathcal{D}_\mu T = e^{-\phi} \tilde{\mathcal{D}}_\mu \tilde{T}$$

## Weyl tensor

$$(d-2)C_{\mu\alpha\nu\beta}u^\alpha u^\beta = \\ \Delta^{\mu\lambda}\Delta^{\nu\sigma}R_{\lambda\sigma} + (d-2)\Delta^{\mu\lambda}\Delta^{\nu\sigma}R_{\lambda\alpha\sigma\beta}u^\alpha u^\beta \\ - \frac{\Delta^{\mu\nu}}{d-1}(\Delta^{\lambda\sigma}R_{\lambda\sigma} + (d-2)\Delta^{\lambda\sigma}R_{\lambda\alpha\sigma\beta}u^\alpha u^\beta)$$

$R_{\mu\nu}\dots$  Ricci tensor

symmetry properties (same as for the Riemann tensor):

$$C_{\mu\nu\lambda\sigma} = -C_{\nu\mu\lambda\sigma} = -C_{\mu\nu\sigma\lambda} = C_{\lambda\sigma\mu\nu} \\ \text{and } C_{\mu\alpha\lambda}{}^\alpha = 0$$

## dispersion relations

quasi-normal modes in the limit of vanishing momentum  
 $k \rightarrow 0$ :

- sound mode/pole

$$\omega_{1,2} = \pm c_s k - i\Gamma k^2 \pm \frac{\Gamma}{c_s} \left( c_s^2 \tau_\pi - \frac{\Gamma}{2} \right) k^3 + \mathcal{O}(k^3),$$

with  $\Gamma = \frac{2}{3} \frac{\eta}{\epsilon + p}$

- shear mode/pole

$$\omega = -i \frac{\eta}{\epsilon + p} k^2 + \mathcal{O}(k^4)$$

## references

### NOT a COMPLETE LIST!

- S. Weinberg, *Gravitation and Cosmology* (1972)
- S. Weinberg, *Cosmology*, Appendix B (2008)
- R. M. Wald, *General Relativity* (1984)
- B. Zwiebach, *A first course in string theory* (2007)
- T. Ortin, *Gravity and Strings* (2004)
- C. V. Johnson, *D-branes* (2003)
- E. Kiritsis, *String Theory in a Nutshell* (2007)

## recent papers

- J.-Y. Ollitrault, *Relativistic hydrodynamics for heavy-ion collisions*  
arXiv: 0708.2433
- R. Baier, P. Romatschke, D. T. Son, A. O. Starinets and M. A. Stephanov, *Relativistic viscous hydrodynamics, conformal invariance, and holography*  
JHEP **0804** (2008) 100
- R. Loganayagam, *Entropy current in conformal hydrodynamics*  
JHEP **0805** (2008) 87
- M. Luzum and P. Romatschke, *Conformal relativistic viscous hydrodynamics: Applications to RHIC at  $\sqrt{s_{NN}} = 200$  GeV*  
PRC **78** (2008) 034915
- M. A. York and G. D. Moore, *Second order hydrodynamic coefficients from kinetic theory*  
arXiv: 0811.0729

## reviews

- S. R. Wadia, *String theory: A framework for quantum gravity and various applications*  
arXiv: 0809.1036
- M. Natsuume, *String theory implications on causal hydrodynamics*  
arXiv: 0807.1394
- R. C. Myers and S. E. Vazquez, *Quark soup al dente: Applied superstring theory*  
arXiv: 0804.2423
- H. Nastase, *Introduction to AdS/CFT*  
arXiv: 0712.0689
- D. Mateos, *String theory and quantum chromodynamics*  
arXiv: 0709.1523

## reviews, cont.

- M. Natsuume, *String theory and quark-gluon plasma*  
arXiv: hep-ph/0701201
- D. T. Son and A. O. Starinets, *Viscosity, black holes and quantum field theory*  
Annu. Rev. Nucl. Part. Sci 2007, **57**, p. 95 - 118
- I. R. Klebanov, *TASI lectures: Introduction to the AdS/CFT correspondence*  
arXiv: hep-th/0009139
- J. L. Petersen, *Introduction to the Maldacena conjecture on AdS/CFT*  
arXiv: hep-th/9902131

# AdS/CFT correspondence



**duality**  $\sim$  **equality**

is between

**Quantum** Field Theory (a special one) in  $d = 4$

and

**Classical** Gravity in  $d = 5$  (for  $N_c \gg 1$ ,  $g_{YM}^2 N_c \gg 1$ )

(at finite temperature)

## some classical gravity

### Einstein-Hilbert with nonzero cosmological constant ( $d$ dimensions)



$$R_{\mu\nu} - \frac{R}{2}g_{\mu\nu} = -\Lambda_d g_{\mu\nu}, \quad \mu, \nu = 0, 1, \dots, d-1$$

• ansatz:

$$ds^2 = g_{tt}dt^2 + g_{rr}dr^2 + r^2 d\Omega_{(d-2)}^2, \quad g_{tt} = -\frac{1}{g_{rr}} = -f(r)$$



$$R\left(1 - \frac{d}{2}\right) = -\Lambda_d d, \quad R_{\mu\nu} = \frac{2\Lambda_d}{d-2}g_{\mu\nu}$$

# AdS gravity

• with

$$T_{\mu\nu} = (\epsilon + p)u_\mu u_\nu + pg_{\mu\nu} = -\frac{\Lambda_d}{8\pi G_d}g_{\mu\nu}$$

follows

$$\epsilon_{vac} = -p_{vac} = \frac{\Lambda_d}{8\pi G_d} < 0$$

for  $\Lambda_d < 0 \dots$

**Anti-deSitter**

[Hawking and Page 83]

## solution

$$R_{tt} = \frac{f}{2} \frac{[r^{d-2} f']'}{r^{d-2}} = -\frac{2\Lambda_d}{d-2} f ,$$

$$f' = -\frac{4\Lambda_d}{(d-1)(d-2)} r \quad (+const) ,$$

$$f = 1 - \frac{2\Lambda_d}{(d-1)(d-2)} r^2 \equiv 1 + \frac{r^2}{L^2}$$

$r \gg L \equiv R$  (radius of AdS space = distance scale)

$$ds^2 = -\frac{r^2}{R^2} dt^2 + \frac{R^2}{r^2} dr^2 + \dots$$

•  $AdS_5$

$$ds^2 = \frac{r^2}{R^2} (-dt^2 + d\vec{x}^2) + \frac{R^2}{r^2} dr^2 = \frac{R^2}{z^2} (-dt^2 + d\vec{x}^2 + dz^2) , \quad z = \frac{R^2}{r}$$

## exercise

consider  $AdS_5$  metric, but start from coordinates  $y^\mu, \mu = 0, 1, \dots, 5$  :

$$(y^0)^2 - (y^1)^2 - \dots - (y^4)^2 + (y^5)^2 = R^2$$

and change to

$$x^\mu = \frac{R}{r}(y^0, y^1, y^2, y^3) \quad , \quad r = y^4 + y^5 \quad ,$$

$$y^4 - y^5 = \frac{R^2}{r} + \frac{r}{R^2}((x^0)^2 - (x^1)^2 - (x^2)^2 - (x^3)^2)$$

leads to

$$ds^2 = \frac{r^2}{R^2}(-dt^2 + d\vec{x}^2) + \frac{R^2}{r^2}dr^2$$

**$AdS_5$  has  $SO(2, 4)$  symmetry**

## Hawking temperature

black hole metric:

$$ds^2 = g_{tt}dt^2 + g_{rr}dr^2 + \dots, \quad g_{tt} = -\frac{1}{g_{rr}} = -\left[1 - \left(\frac{r_0}{r}\right)^{(d-3)}\right]$$

expand around the horizon  $r \approx r_0$ :

$$g_{tt} \approx -\frac{d-3}{r_0}(r-r_0) = -\gamma_t(r-r_0), \quad g_{rr} \approx \frac{\gamma_r}{r-r_0}, \quad \gamma_r = \frac{r_0}{d-3}$$

claim:

$$T_H = \frac{1}{4\pi} \sqrt{\frac{\gamma_t}{\gamma_r}} = \frac{d-3}{4\pi r_0}$$

and  $AdS_5$ :  $\gamma_t = 1/\gamma_r = 4r_0/R^2$

$$T_H = \frac{r_0}{\pi R^2}$$

proof: with  $\rho^2 = \frac{4r_0(r-r_0)}{d-3}$

$$ds^2 = \rho^2 \left( \frac{d-3}{2r_0} \right)^2 d\tau^2 + d\rho^2 + \dots, \quad \text{euclidean } d\tau^2 = -dt^2$$

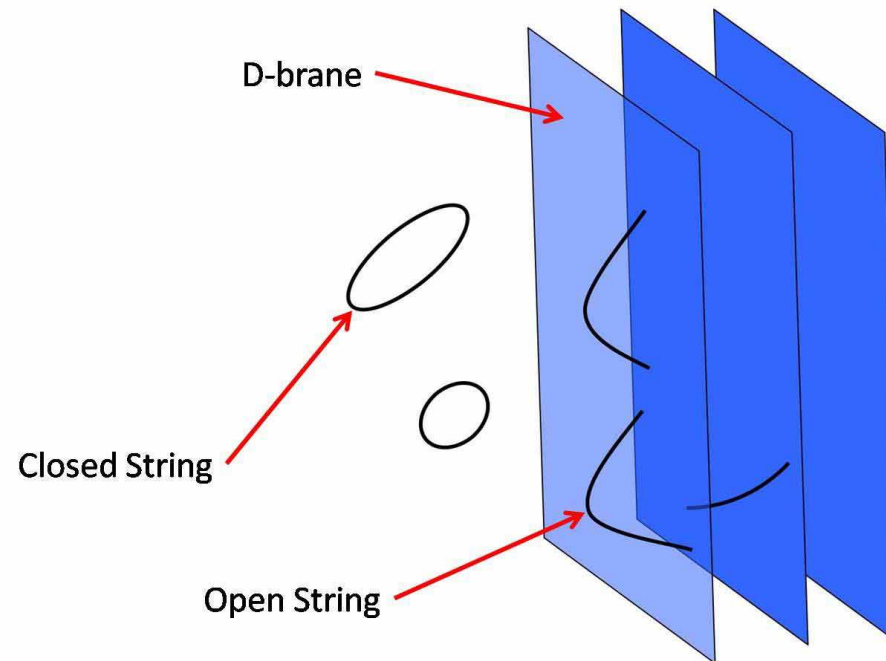
→

$$ds^2 = \rho^2 d\phi^2 + d\rho^2, \quad \phi = \frac{d-3}{2r_0} \tau$$

**requirement:** periodicity  $\phi \rightarrow \phi + 2\pi$ , i.e. **NO** conical singularity

$$2\pi = \frac{(d-3)}{2r_0} \beta, \quad \beta = \frac{1}{T_H}$$

## $D$ -branes



[from Myers and Vazquez 08]

**dynamical walls on which strings can end:**

theory of open strings living on  $N_c - D3$ -branes ( $\mathcal{N} = 4$

SYM,  $d = 4$ )

$\iff$

gravity theory of fields living in the space curved by the  
branes ( $AdS_5$ ,  $d = 5$ )



## *AdS*<sub>5</sub> metric at finite *T*

generalization of the Einstein-Hilbert-Maxwell equation  
- strong analogy to Reissner-Nordström black hole

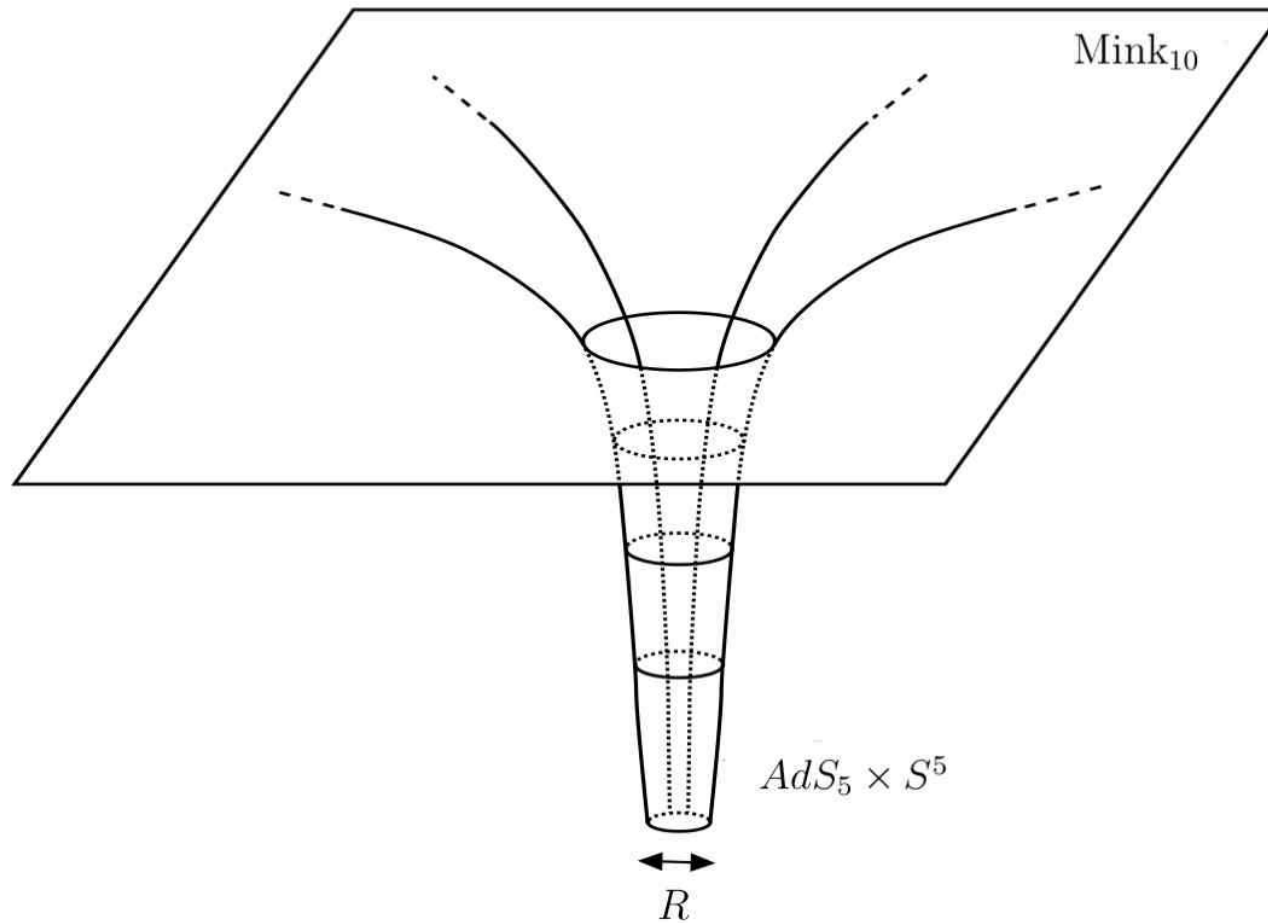
$$S = \frac{1}{16\pi G_{(10)}} \int d^{10}x \sqrt{-g} \left[ R - \frac{F_5^2}{2 \cdot 5!} + \dots \right],$$

$$\int_{S^5} *F_5 \propto N_c, \quad G_{(10)} = \frac{\pi^4 L^8}{2N_c^2}, \quad H(r) = 1 + \frac{L^4}{r^4}, \quad f(r) = 1 - \frac{r_0^4}{r^4}$$

$$ds^2 = -\frac{f(r)}{\sqrt{H(r)}} dt^2 + \frac{1}{\sqrt{H(r)}} (dx_1^2 + dx_2^2 + dx_3^2) + \frac{\sqrt{H(r)}}{f(r)} dr^2 + r^2 \sqrt{H} d\Omega_5^2$$

near extremal black *D3*– brane metric with horizon  $r = r_0$   
for  $r_0 < r \ll L = R$ , i.e. factorized metric for  $AdS_5 \otimes S^5$

$$ds^2 = \frac{r^2}{R^2} (-f(r) dt^2 + d\vec{x}^2) + \frac{R^2}{r^2 f(r)} dr^2 + R^2 d\Omega_5^2$$

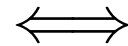


[from D. Mateos]

space-time around  $D3$ -branes

## AdS/CFT [Maldacena 98]

strongly coupled quantized conformal gauge theory  
in  $d = 4$  dimensions ( $\mathcal{N} = 4$  SYM with  $8N_c^2$  (1 gauge and 6  
scalar) bosons and  $(4N_c^2)$  Weyl fermions) **[[ NOT QCD !]]**



weakly coupled classical supergravity (type IIB)  
in  $d = 10$  dimensions (on  $AdS_5 \times S^5$ )  
via **holographic property**: radial coordinate  $r_0 \leq r < \infty$  with  
gauge theory on the boundary at  $\infty$   
in the limit:

't Hooft coupling  $\lambda = g_{YM}^2 N_c$  is large,  $N_c \rightarrow \infty$ ,  $g_{YM}^2 \ll 1$

i.e. string coupling  $g_s = \frac{g_{YM}^2}{4\pi} \ll 1$  – NO LOOPS

and

small curvature  $\frac{l_s^4}{R^4} = \frac{1}{\lambda} \ll 1$  – RADIUS  $R$  of CURVATURE  
is LARGE compared to the STRING SCALE  $l_s = \sqrt{\alpha'}$

## symmetries

duality maps the operators of QFT to the boundary values of the SUGRA fields

- QFT ( $\mathcal{N} = 4, d = 4$  CFT)

$$S = -\frac{1}{g_{YM}^2} \int d^4x \left[ \frac{1}{4} F_{\mu\nu}^a F^{a\mu\nu} + \frac{1}{2} D_\mu \Phi_i^a D^\mu \Phi_i^a + \dots \right]$$

.... massless Weyl fermions + interactions ,  $i = 1, \dots, 6, a = 1, \dots, N_c^2$

has Poincaré and conformal symmetry  $SO(2, 4)$ , e.g.

$x^\mu \rightarrow \lambda x^\mu$ ,  $\Phi_i^a(x) \rightarrow \Phi_i^a(\lambda x)$ , etc,  $\beta(\mu) = 0$ , and  $SU(\mathcal{N} = 4)$  symmetry and  $SU(N_c)$

- $AdS_5 \otimes S^5$ :

$$SO(2, 4) \otimes SO(6) = SO(2, 4) \otimes SU(4)$$

besides  $SU(N_c)$  due to  $N_c$  parallel  $D - 3$  branes

## relaxation phenomena/strong coupling limit

hydrodynamic transport coefficients by gauge/gravity duality:

compare with quite involved AdS/CFT-gravity calculations

at Hawking temperature  $T_H = \frac{r_0}{\pi R^2}$ ,

and for momentum  $\omega, k \ll T_H$ ,

e.g. from sound channel dispersion up to  $O(k^3)$ , etc :

$$\frac{\eta}{s} = \frac{1}{4\pi}, \quad \tau_\pi = \frac{2 - \ln 2}{2\pi T}, \quad \kappa = \frac{\eta}{\pi T} = 2\lambda_1,$$

$$\lambda_2 = -\frac{\ln 2}{2\pi T}\eta, \quad \lambda_3 = 0$$

## conformal $\Pi^{\mu\nu}$

all 5 second order terms classified by conformal symmetry

constitutive relation (d= 4):

$$\begin{aligned}\Pi^{\mu\nu} = & -2\eta\sigma^{\mu\nu} + 2\eta\tau_\pi u^\lambda \mathcal{D}_\lambda \sigma^{\mu\nu} \\ & + 4\lambda_1 \sigma^{\langle\mu}{}_\lambda \sigma^{\nu\rangle\lambda} + 2\lambda_2 \sigma^{\langle\mu}{}_\lambda \Omega^{\nu\rangle\lambda} + \lambda_3 \Omega^{\langle\mu}{}_\lambda \Omega^{\nu\rangle\lambda} \\ & + 2\kappa u_\alpha C^{\alpha\mu\nu\beta} u_\beta\end{aligned}$$

$C^{\alpha\beta\gamma\delta}$  ... Weyl tensor,  $\Omega^{\alpha\beta}$  ... antisymmetric vorticity tensor

[Baier, Romatschke, Son, Starinets, Stephanov 07]

[Bhattacharyya, Hubeny, Minwalla, Rangamani 07, Loganayagam 08]

## Bekenstein- Hawking entropy

thermodynamics:  $dS = \frac{dE}{T}$

- Schwarzschild BH ( $G = G_{(4)}, T_H = \frac{1}{4\pi r_0}$ )

$$E = M = \frac{r_0}{2G}, \quad dE = \frac{dr_0}{2G}, \quad dS = \frac{4\pi}{2G} r_0 dr_0,$$

entropy

$$S = \frac{\pi r_0^2}{G} = \frac{A}{4G}$$

a universal relation with  $A = 4\pi r_0^2$ ... area of the horizon

## Bekenstein-Hawking entropy, cont.

●  $AdS_5 \otimes S^5$ :  $T_H = \frac{r_0}{\pi R^2}$  ,  $S = \frac{A_{(8)}}{4G_{(10)}}$

$$A_{(8)} = \left(\frac{r_0}{R}\right)^3 V_3 R^5 \Omega_{(5)} = \pi^6 T_H^3 R^8 V_3 , \quad G_{(10)} = \frac{\pi^4 R^8}{2N_c^2}$$

gives for the entropy density

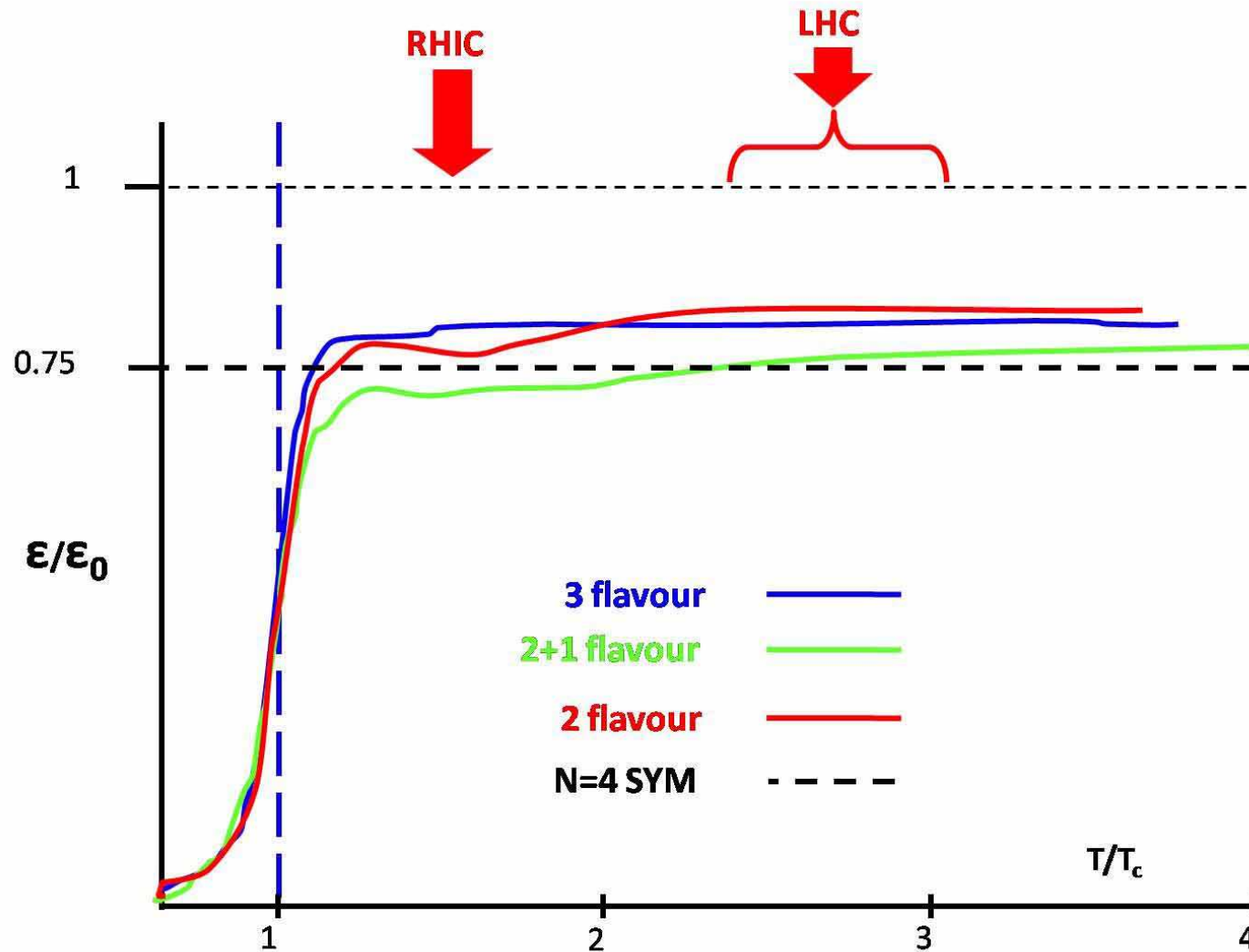
$$s_{BH} = \frac{S}{V_3} = \frac{\pi^2 N_c^2}{2} T_H^3$$

NOTE:

$$s_{Boltzmann} = \frac{4p}{T} = \frac{4\pi^2 T^3}{90} \left[ \frac{7}{4} \cdot 4 + 2 + 6 \right] N_c^2 = \frac{2\pi^2 N_c^2}{3} T^3$$

$$s_{BH} = \frac{3}{4} s_{Boltzmann}$$





energy density of QCD and SYM - via BH entropy

[from Myers and Vazquez 08]

## master formula for AdS/CFT

schematically in terms of coinciding partition functions:

$$\int e^{iS_{4d}^{gauge} + i\Phi_0 O} = \int e^{iS_{5d}[\Phi]} \simeq e^{iS^{classical}[\Phi_0]}$$

$S_{5d}$  is computed with non-trivial boundary condition  
(holography)

$$\Phi(t, \vec{x}, r) \stackrel{r \rightarrow \infty}{\simeq} \Phi_0(t, \vec{x})$$

$\implies$

quantum correlation = classical two-point function

$$G_R(x, y) = -i \langle O(x)O(y) \rangle = - \frac{\delta^2 S^{classical}}{\delta\Phi_0(x)\delta\Phi_0(y)} \Big|_{r=\infty}$$

gauge:  $O = T_{\mu\nu}$  .. energy-momentum tensor

gravity:  $\phi = g_{\mu\nu}$  .. graviton

## correlators from gravity, $d = 4$

response of the fluid to small metric perturbations

**example:**  $g_{xy} = h_{xy}(t, z, r) \neq 0$ ,  $u^0 = 1$ ,  $T = \text{const}$

e.g. it leads from Christoffel symbols in the covariant derivatives to linear approximation:

$$\sigma^{xy} = \sigma_{xy} \approx \frac{1}{2}(\Gamma_{x0}^y + \Gamma_{y0}^x) \approx \frac{1}{2}\partial_t h_{xy}$$

and for the stress tensor

$$\delta\Pi^{xy} \approx -\eta\partial_t h_{xy} + \eta\tau_\pi\partial_t^2 h_{xy} - \frac{\kappa}{2}[\partial_t^2 h_{xy} + \partial_z^2 h_{xy}]$$

Fourier transform  $h(t, z) = \exp(-i\omega t + ikz) h(\omega, k)$ , etc.

$$\delta\Pi^{xy}(\omega, k) = -G_R^{xy,xy}(\omega, k)h_{xy}(\omega, k)$$

via **linear response**

## correlators, cont.

retarded Green function:

$$G_R^{xy,xy}(\omega, k) = -i\eta\omega + \eta\tau_\pi\omega^2 - \frac{\kappa}{2}(\omega^2 + k^2) + \dots$$

i.e. **Kubo formula**

$$\begin{aligned}\eta &= \lim_{\omega} \frac{1}{\omega} iG_R^{xy,xy}(\omega, \vec{0})|_{\omega=0} \\ &= \lim_{\omega} \frac{1}{\omega} \int dt d^3x \exp(i\omega t) \langle [T_{xy}(t, \vec{x}), T_{xy}(0, \vec{0})] \rangle |_{\omega=0}\end{aligned}$$

## quasinormal mode

$G_R$  from gravity action with  $AdS_5$  metric:

$$S_{5d} = \frac{N_c^2}{8\pi^2 R^3} \int d^5x [\sqrt{-g}R_{5d} + \dots],$$

factor from  $\frac{1}{16\pi G_{(10)}} R^5 \Omega_{(5)}$ ,  $\Omega_{(5)} = \pi^3$

and “dictionary”  $G_{(10)} = 8\pi^6 g_s^2 l_s^8$ ,  $g_{YM}^2 N_c = 4\pi g_s N_c = R^4/l_s^4$   
scalar mode and EoM (of Heun’s type)

$$\sqrt{-g}R_{(5d)} \rightarrow -\frac{1}{2}\sqrt{-g}g^{\mu\nu} \partial_\mu \Phi \partial_\nu \Phi, \quad \Phi \equiv h_{xy} :$$

$$\Phi(t, z, u = \frac{r_0^2}{r^2}) = \exp(-i\omega t + ikz) \Phi_k(u), \quad \Phi'_k = \frac{d}{du} \Phi_k,$$

$$\Phi_k'' - \frac{1+u^2}{u(1-u^2)} \Phi_k' + \frac{1}{4\pi^2 T^2} \frac{\omega^2 - k^2(1-u^2)}{u(1-u^2)^2} \Phi_k(u) = 0$$

## quasinormal mode, cont.

solution with ansatz  $\Phi_k(u) = f_k(u)\Phi_0(k)$  ,  $f_k(0) = 1$  and  
boundary condition: only the incoming wave, the one which  
moves toward the horizon at  $r- > r_0 \equiv u = 1$ , i.e. nothing  
comes out the horizon

$$f_k(u) = (1 - u^2)^{-i\omega/(4\pi T)} F_k(u) \approx 1 + i\frac{\omega}{4\pi T}u^2 + \dots$$

note, near  $u \sim 1$ :

$$\exp(-i\omega t)f_k(u) \rightarrow \exp[-i\omega(t + r^*)] , \quad r^* = \frac{\ln(1 - u)}{4\pi T}$$

moves from  $u = 0$  at  $t = 0$  to  $u = 1$  at  $t- > \infty$

## classical gravity action

inserting the solution into  $S_{(5d)}$

$$S_{(5d)} \approx \frac{N_c^2}{8\pi^2 R^3} \int d^4x \int_0^1 du \sqrt{-g} \left(-\frac{1}{2}\right) g^{uu} \partial_u \Phi \partial_u \Phi$$

and keeping only the surface contribution at  $u = 0$ ,  $\Rightarrow$

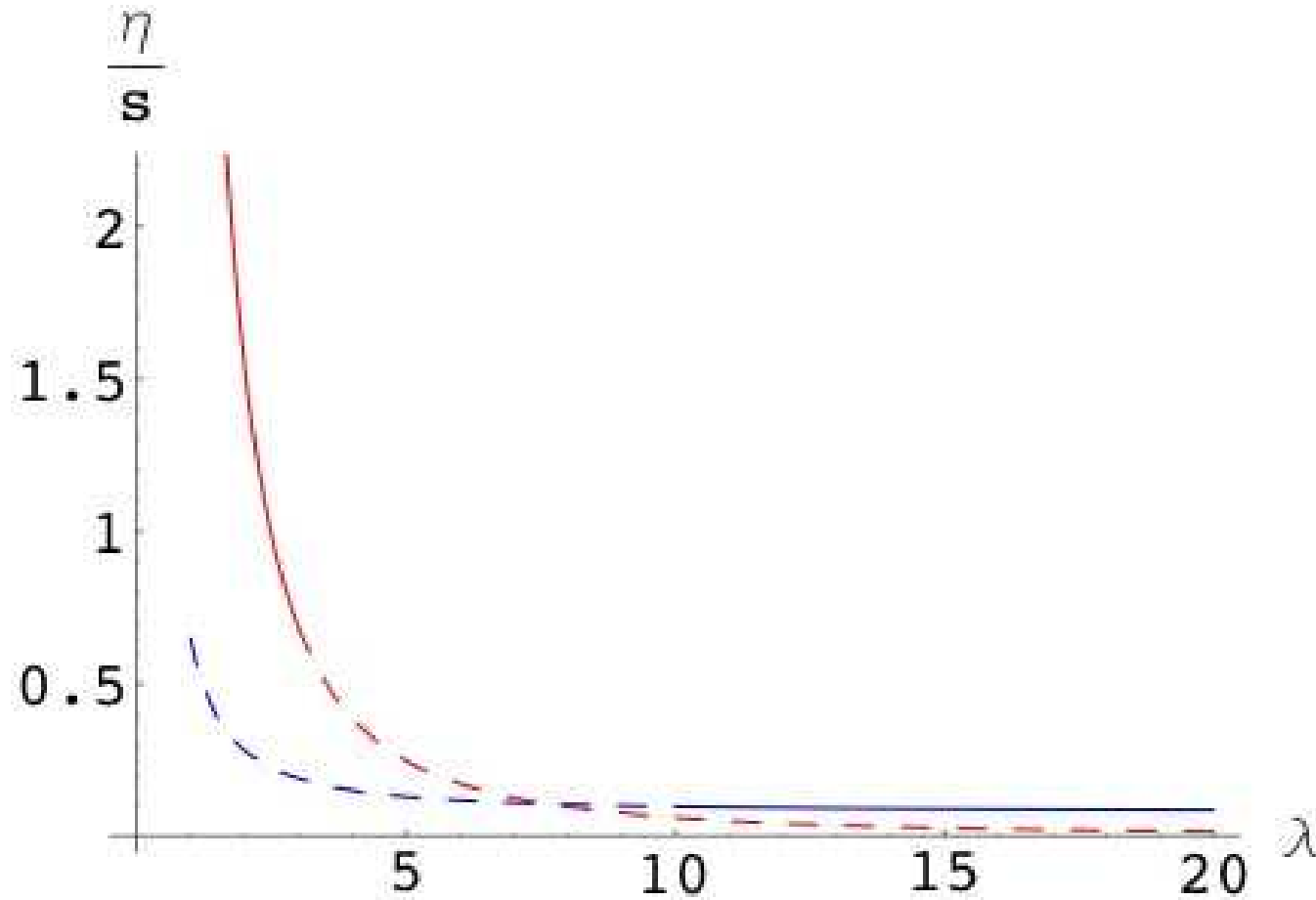
$$S^{classical} = -\frac{\pi^2 N_c^2 T^4}{8} \int \frac{d^4k}{(2\pi)^4} \Phi_0(-k) \left[ \frac{(1-u^2)}{u} f'_k(u) f_k(u) \right] \Big|_{u=0} \Phi_0(k)$$

and finally

$$G_R^{xy,xy}(\omega, k) = \frac{\pi N_c^2 T^3}{8} \left[ -i\omega + \frac{1 - 2 \ln 2}{2\pi T} \omega^2 - \frac{1}{2\pi T} k^2 \pm \dots \right],$$

$$\eta = \frac{\pi N_c^2 T^3}{8}, \quad \tau_\pi = \frac{2 - \ln 2}{2\pi T}, \quad \kappa = \frac{\eta}{\pi T}$$

$$\eta/s$$



behaviour of  $\eta/s$  as a function of the 't Hooft coupling



## shear viscosity $\eta$

near equilibrium:  $\eta \simeq \epsilon \bar{v} \lambda_f$ , *entropy density*  $s \simeq \epsilon/m \rightarrow$

$$\frac{\eta}{s} \simeq m \bar{v} \lambda_f \simeq \hbar \frac{\text{mean free path}}{\text{deBroglie wavelength}}$$

- dilute system (QFT  $\rightarrow$  kinetic theory  $\rightarrow$  hydro):  
**scale**  $\lambda_f \rightarrow \frac{\eta}{s} \gg \hbar$ , e.g. **pQCD** ( $N_f = 0$ )

$$\frac{\eta}{s} \simeq 3.8 \frac{1}{g^4 \ln(2.8/g)} \simeq O(1) \text{ for } g = 2.5$$

**BUT** with  $\ln(2.8/g) \simeq O(1) : \frac{\eta}{s} \simeq 0.1 \rightarrow$  **sensitive to constant under the log !**

[Arnold, Moore and Yaffe 03]

- strongly coupled system (QFT  $\rightarrow$  hydro):  
**scale**  $1/T \rightarrow \frac{\eta}{s} = \frac{\hbar}{4\pi} \simeq 0.08$

[Policastro, Son and Starinets 01]

## sound mode

consistency requirement from sound channel:

$$\omega_{1,2} = \pm c_s k - i\Gamma k^2 \pm \frac{\Gamma}{c_s} \left( c_s^2 \tau_\pi - \frac{\Gamma}{2} \right) k^3 + \mathcal{O}(k^3),$$

with  $\Gamma = \frac{2}{3} \frac{\eta}{\epsilon + p}$

from gravity perturbation  $g_{tz} = h_{tz} \ll 1$

$$\omega_{1,2} \pm \frac{1}{\sqrt{3}} k - i \frac{1}{6\pi T} k^2 \pm \frac{3 - 2 \ln 2}{24\pi^2 T^2 \sqrt{3}} k^3$$

gives sound velocity  $c_s = \frac{1}{\sqrt{3}}$  and

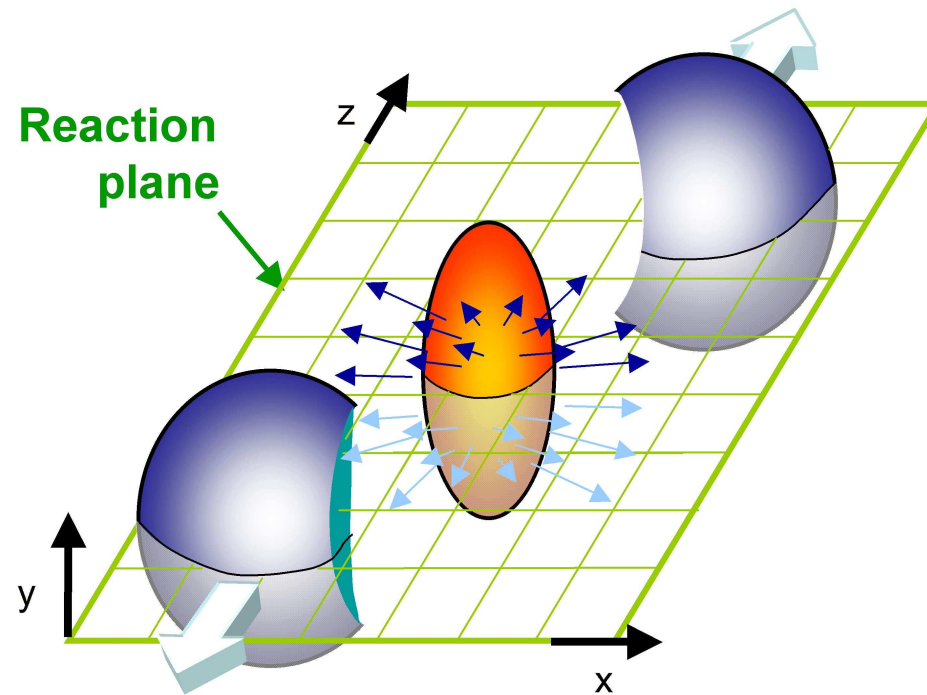
$$\frac{\eta}{s} = \frac{1}{4\pi}, \quad \tau_\pi = \frac{2 - \ln 2}{2\pi T}$$

# hydro + RHIC

heavy-ion collisions require **beyond hydrodynamics**:

- hydrodynamics = differential equations  
**initial conditions !**
- initial = equilibration time
- distribution of energy density (Glauber? CGC ?)
- QCD equation of state
- hadronisation prescription (Cooper-Frye?)

## as important example



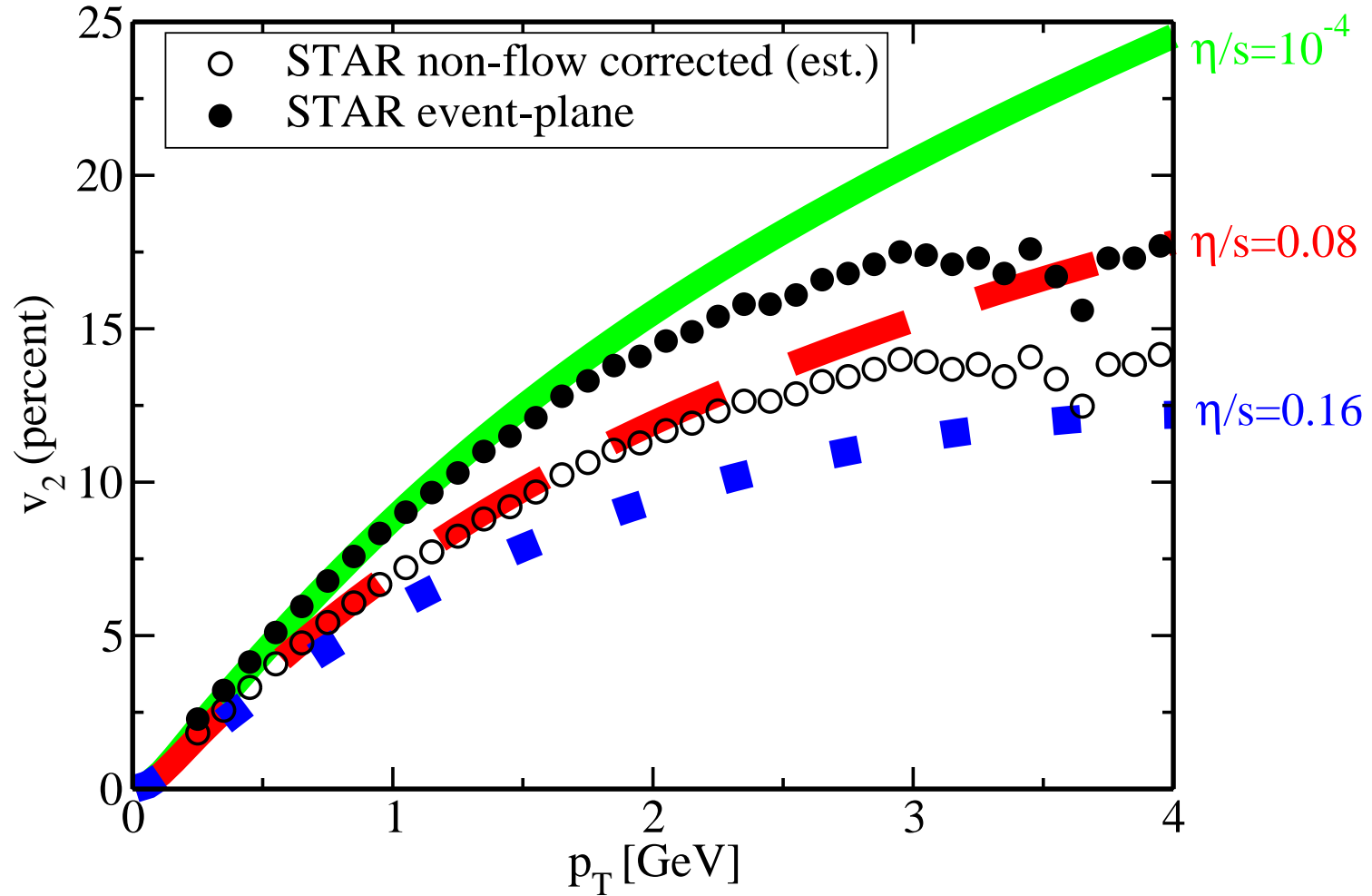
spatial asymmetry



$$\frac{dN}{dy dp_{\perp} d\phi} = \left\langle \frac{dN}{dy dp_{\perp} d\phi} \right\rangle_{\phi} (1 + 2v_2(p_{\perp}) \cos(2\phi) + \dots)$$

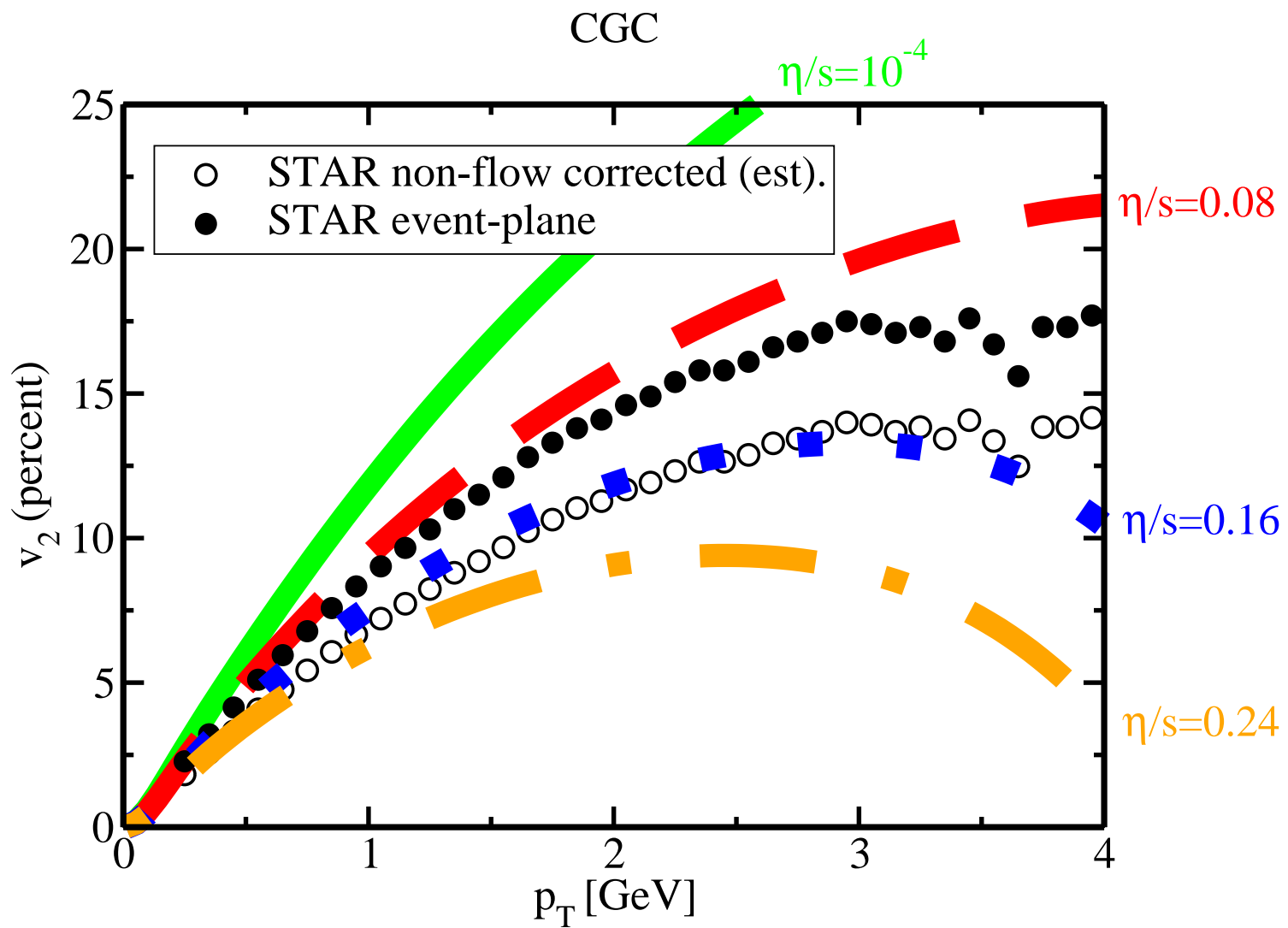
● elliptic flow:  $v_2(p_{\perp})$

# Glauber



elliptic flow

[Luzum and Romatschke 08]



elliptic flow

[Luzum and Romatschke 08]

## main results [Luzum and Romatschke 08]

- viscous hydrodynamics simulation give a good description of RHIC data for

$$\frac{\eta}{s} = 0.1 \pm 0.1(\text{theory}) \pm 0.08(\text{experiment})$$

- modest estimate:

$$\frac{\eta}{s} < 0.5$$

- weak dependence on the values of the second-order parameters  $\tau_\pi, \lambda_1, \dots$
- early thermalisation time is questioned, but

$$\tau_0 < 2 \text{ fm}$$

## transport coefficients

kinetic QCD theory (weak coupling):

$$\frac{\eta/s}{\tau_\pi T} \simeq 0.17-0.2, \quad \lambda_1 \simeq -(2.0-2.2) \frac{\eta^2}{sT}, \quad \lambda_2 = -2\tau_\pi\eta, \quad \lambda_3 = \kappa = 0$$

[York and Moore 08]

**finite 't Hooft coupling**  $\lambda \equiv g_{YM}^2 N_c$ ,  $\lambda \gg 1$  corrections to coefficients by gauge/gravity duality, e.g.:

[Buchel and Paulos 08]

$$\frac{\eta/s}{\tau_\pi T} = 0.383 (1 - 3.52 \lambda^{-3/2} + \dots),$$

$$\kappa = \frac{\eta}{\pi T} \left( 1 - \frac{145\zeta(3)}{8} \lambda^{-3/2} + \dots \right)$$



excitement about gauge/gravity correspondence:  
mainly to gain intuition into **STRONG COUPLING**

ASKING FOR MORE:

Is there an experiment whose outcome could cast strong doubts on the relevance of AdS/CFT to understand QCD ?

[P. Jacobs 08]

**JET PHYSICS ?**

[Hatta, Iancu and Mueller 07 - 08]

**MOST CHALLENGING TASK** of the theory is to find the microscopic mechanism for the rather **RAPID EQUILIBRATION** of matter in RHIC collisions

EXTRAS

## COMPARISON

	QCD	$\mathcal{N}=4$ SYM
$T=0$	$N_c=3=N_f$ , confinement, discrete spectrum, scattering, . . . .	$N_c$ large, $N_f/N_c$ small, deconfined, conformal, supersymmetric, . . . .
	<b>very different !!</b>	
$T > T_c$	strongly-coupled plasma of gluons & <b>fundamental</b> matter  deconfined, screening, finite corr. lengths, . . .	strongly-coupled plasma of gluons & <b>adjoint and fundamental</b> matter  deconfined, screening, finite corr. lengths, . . .
	<b>very similar !!</b>	
$T \gg T_c$	runs to weak coupling	remains strongly-coupled
	<b>very different !!</b>	

QCD and  $\mathcal{N} = 4$  SYM as a function of temperature

[from Myers and Vazquez 08]

## entropy current

Israel - Stewart 79:  $s^\mu = (s - \frac{\tau_\pi}{4\eta T} \Pi_{\alpha\beta} \Pi^{\alpha\beta}) u^\mu$

$\Rightarrow$  **second law:**

$$\nabla_\mu s^\mu = \frac{\Pi_{\alpha\beta} \Pi^{\alpha\beta}}{2\eta T} \geq 0$$

instead in **causal viscous and conformal hydrodynamics**  
a more general current in terms of  $u^\mu$  and its derivatives:

$$s^\mu = s u^\mu + (\# \sigma^2 + \# \Omega^2) u^\mu + O(u \nabla^2 u)$$

with

$$\nabla_\mu s^\mu = \frac{\eta}{2T} \sigma^{\mu\nu} \sigma_{\mu\nu} + \frac{1}{4T} (\kappa - 2\lambda_1) \sigma_\nu^\mu \sigma_\lambda^\nu \sigma_\mu^\lambda$$

**in  $\mathcal{N} = 4$  SYM:  $\kappa = 2\lambda_1$**

[Loganayagam 08]

## Bjorken flow

boost-invariant (irrotational) 1 + 1 flow [Bjorken 83]  
**second-order equations** (proper time  $\tau$ ,  $\Phi$ ... viscous flow):

$$\partial_\tau \epsilon = -\frac{4\epsilon}{3\tau} + \frac{\Phi}{\tau}$$

$$\tau_\pi \partial_\tau \Phi = \frac{4\eta}{3\tau} - \Phi - \frac{4\tau_\pi}{3\tau} \Phi - \frac{\lambda_1}{2\eta^2} \Phi^2$$

**non-linear term NOT in MIS theory!**

[BRSSS 07]

compare with AdS/CFT calculation:

$$\frac{\lambda_1 T}{\eta} = \frac{1}{2\pi} \left[ 1 + \frac{215 \zeta(3)}{8} \lambda^{-3/2} + \dots \right]$$

[Janik, Peschanski, Heller 06; Buchel, Paulos 08]

## Müller - Israel - Stewart theory

keeping essentially **one term** in the derivative expansion  
up to **second order**

$$\Pi^{\mu\nu} = -2\eta\sigma^{\mu\nu} + 2\eta \tau_\pi \langle D\sigma^{\mu\nu} \rangle, \quad D = u \cdot \nabla$$

remark: does not match with AdS/CFT  $\mathcal{N} = 4$  SYM

$$\text{sound} : \tau_\pi = \frac{2 - \ln 2}{2\pi T}$$

$$\text{Bjorken flow} : \tau_\pi = \frac{1 - \ln 2}{2\pi T}$$

$\Rightarrow$  **all second-order terms** consistent with  
conformal symmetry have to be included !

# equilibration time

REMINDER:

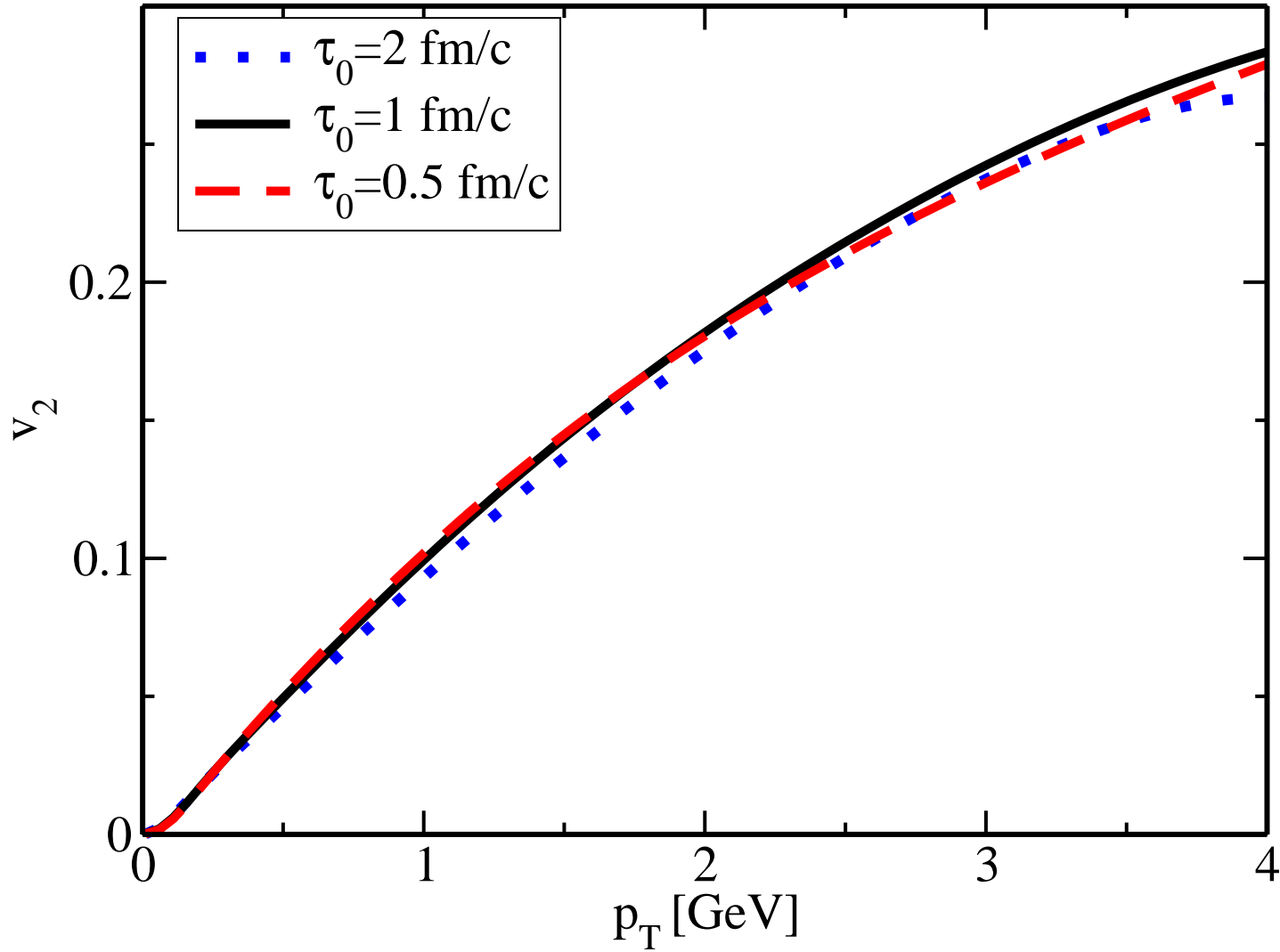
claim of short  $\tau_0 \equiv \tau_{eq} \leq 0.5$  fm at RHIC

OPEN QUESTION:

*CGC* ( $\alpha_s \ll 1$ )  $\rightarrow$  *sQGP* ( $\alpha_s > O(1)$ )

within a very short time  $< 0.5$  fm ?

(b)



early thermalisation ?

[Luzum and Romatschke 08]



# parametric pQCD estimate

for thermalisation in an expanding gluonic medium

near equilibrium at  $T(\tau)$ : Knudsen number  $Kn$

$$\frac{1}{Kn} = \frac{\text{longitudinal expansion time}}{\text{mean free path}} = \frac{\tau}{\lambda_f} \gg 1$$

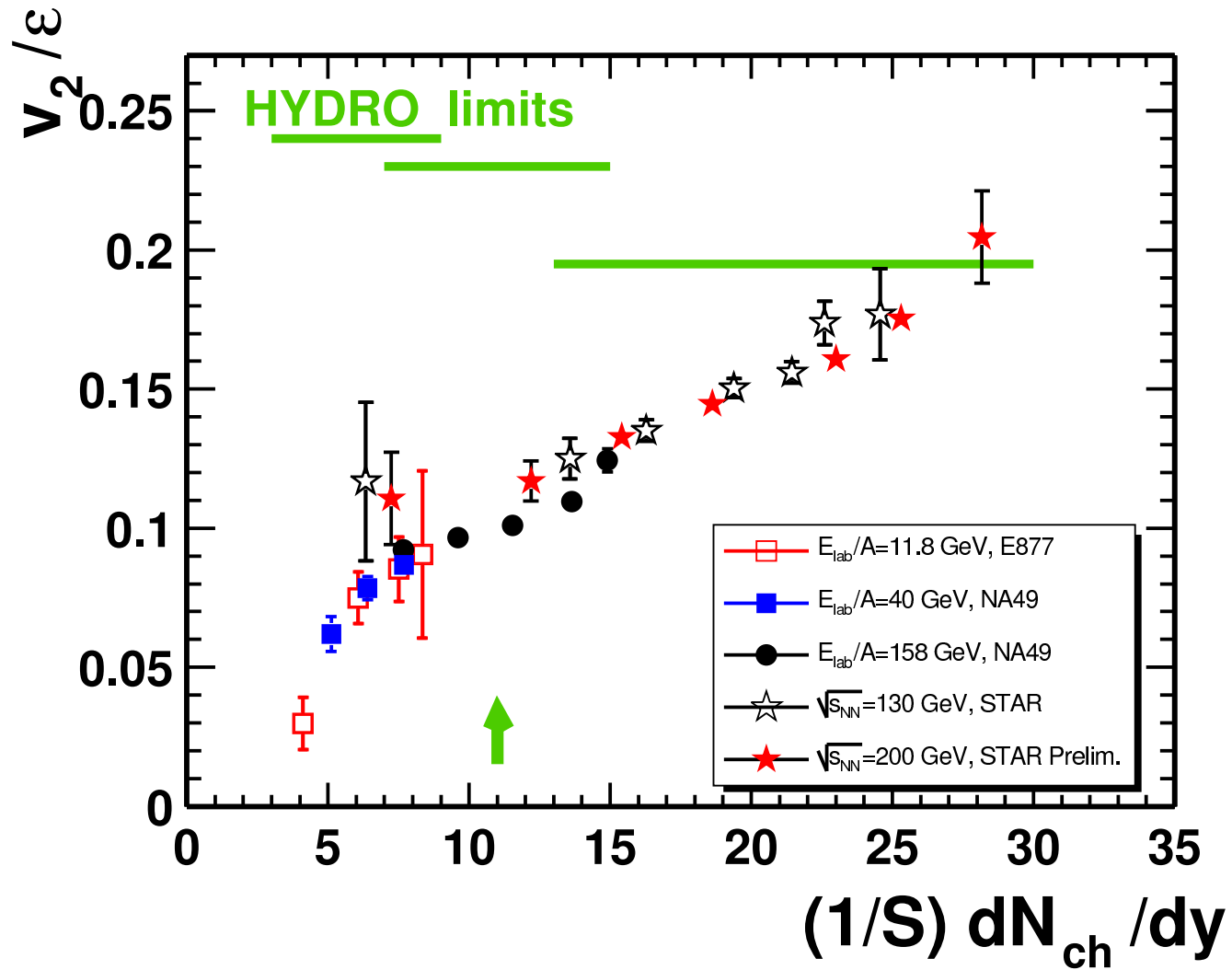
from many gluon interactions (including saturation):

Arnold et al.:  $\tau_{eq} Q_s \geq \alpha_s^{-7/3}$

‘bottom-up’ [Baier, Son, Mueller and Schiff 01]

$$\tau_{eq} Q_s \geq \alpha_s^{-13/5}$$

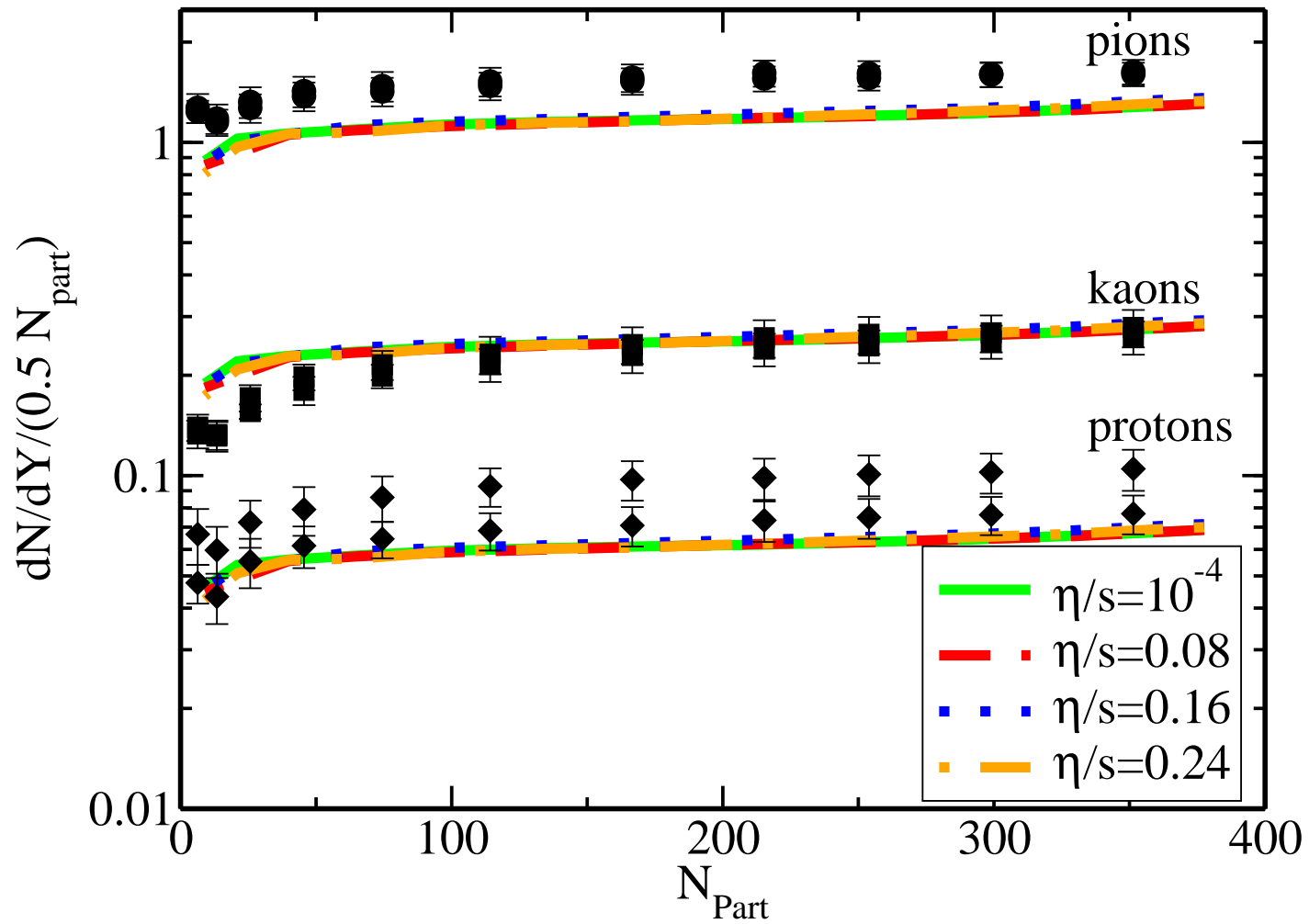
$$\text{RHIC: } \tau_{eq} \geq 2 - 3 \text{ fm}$$



$v_2$  : experiment vs. perfect hydro

[Heinz 04]

# CGC



multiplicity

[Luzum and Romatschke 08]