

AdS/CFT correspondence

duality \sim **equality**

is between

Quantum Field Theory (a special one) in $d = 4$

and

Classical Gravity in $d = 5$ (for $N_c \gg 1$, $g_{YM}^2 N_c \gg 1$)

(at finite temperature)

some classical gravity

Einstein-Hilbert with nonzero cosmological constant (d dimensions)



$$R_{\mu\nu} - \frac{R}{2}g_{\mu\nu} = -\Lambda_d g_{\mu\nu}, \quad \mu, \nu = 0, 1, \dots, d-1$$

• ansatz:

$$ds^2 = g_{tt}dt^2 + g_{rr}dr^2 + r^2 d\Omega_{(d-2)}^2, \quad g_{tt} = -\frac{1}{g_{rr}} = -f(r)$$



$$R\left(1 - \frac{d}{2}\right) = -\Lambda_d d, \quad R_{\mu\nu} = \frac{2\Lambda_d}{d-2}g_{\mu\nu}$$

AdS gravity

• with

$$T_{\mu\nu} = (\epsilon + p)u_\mu u_\nu + pg_{\mu\nu} = -\frac{\Lambda_d}{8\pi G_d}g_{\mu\nu}$$

follows

$$\epsilon_{vac} = -p_{vac} = \frac{\Lambda_d}{8\pi G_d} < 0$$

for $\Lambda_d < 0 \dots$

Anti-deSitter

[Hawking and Page 83]

solution

$$R_{tt} = \frac{f}{2} \frac{[r^{d-2} f']'}{r^{d-2}} = -\frac{2\Lambda_d}{d-2} f ,$$

$$f' = -\frac{4\Lambda_d}{(d-1)(d-2)} r \quad (+const) ,$$

$$f = 1 - \frac{2\Lambda_d}{(d-1)(d-2)} r^2 \equiv 1 + \frac{r^2}{L^2}$$

$r \gg L \equiv R$ (radius of AdS space = distance scale)

$$ds^2 = -\frac{r^2}{R^2} dt^2 + \frac{R^2}{r^2} dr^2 + \dots$$

• AdS_5

$$ds^2 = \frac{r^2}{R^2} (-dt^2 + d\vec{x}^2) + \frac{R^2}{r^2} dr^2 = \frac{R^2}{z^2} (-dt^2 + d\vec{x}^2 + dz^2) , \quad z = \frac{R^2}{r}$$

exercise

consider AdS_5 metric, but start from coordinates y^μ , $\mu = 0, 1, \dots, 5$:

$$(y^0)^2 - (y^1)^2 - \dots - (y^4)^2 + (y^5)^2 = R^2$$

and change to

$$x^\mu = \frac{R}{r}(y^0, y^1, y^2, y^3) \quad , \quad r = y^4 + y^5 \quad ,$$

$$y^4 - y^5 = \frac{R^2}{r} + \frac{r}{R^2}((x^0)^2 - (x^1)^2 - (x^2)^2 - (x^3)^2)$$

leads to

$$ds^2 = \frac{r^2}{R^2}(-dt^2 + d\vec{x}^2) + \frac{R^2}{r^2}dr^2$$

AdS_5 has $SO(2, 4)$ symmetry

Hawking temperature

black hole metric:

$$ds^2 = g_{tt}dt^2 + g_{rr}dr^2 + \dots, \quad g_{tt} = -\frac{1}{g_{rr}} = -\left[1 - \left(\frac{r_0}{r}\right)^{(d-3)}\right]$$

expand around the horizon $r \approx r_0$:

$$g_{tt} \approx -\frac{d-3}{r_0}(r-r_0) = -\gamma_t(r-r_0), \quad g_{rr} \approx \frac{\gamma_r}{r-r_0}, \quad \gamma_r = \frac{r_0}{d-3}$$

claim:

$$T_H = \frac{1}{4\pi} \sqrt{\frac{\gamma_t}{\gamma_r}} = \frac{d-3}{4\pi r_0}$$

and AdS_5 : $\gamma_t = 1/\gamma_r = 4r_0/R^2$

$$T_H = \frac{r_0}{\pi R^2}$$

proof: with $\rho^2 = \frac{4r_0(r-r_0)}{d-3}$

$$ds^2 = \rho^2 \left(\frac{d-3}{2r_0} \right)^2 d\tau^2 + d\rho^2 + \dots, \quad \text{euclidean } d\tau^2 = -dt^2$$

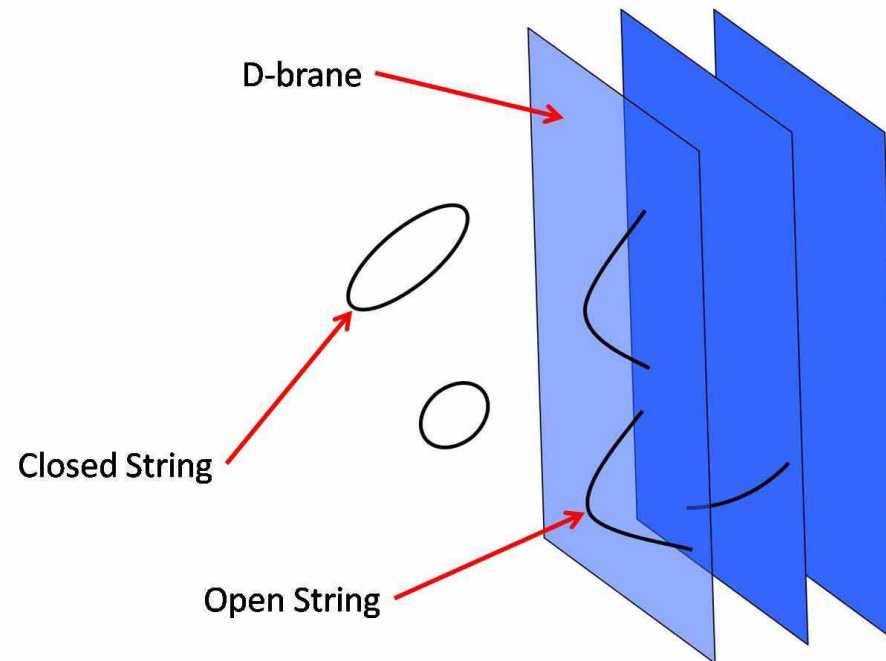
→

$$ds^2 = \rho^2 d\phi^2 + d\rho^2, \quad \phi = \frac{d-3}{2r_0} \tau$$

requirement: periodicity $\phi \rightarrow \phi + 2\pi$, i.e. **NO** conical singularity

$$2\pi = \frac{(d-3)}{2r_0} \beta, \quad \beta = \frac{1}{T_H}$$

D -branes



[from Myers and Vazquez 08]

dynamical walls on which strings can end:

theory of open strings living on $N_c - D3$ -branes ($\mathcal{N} = 4$

SYM, $d = 4$)

\iff

gravity theory of fields living in the space curved by the
branes (AdS_5 , $d = 5$)

*AdS*₅ metric at finite *T*

generalization of the Einstein-Hilbert-Maxwell equation
- strong analogy to Reissner-Nordström black hole

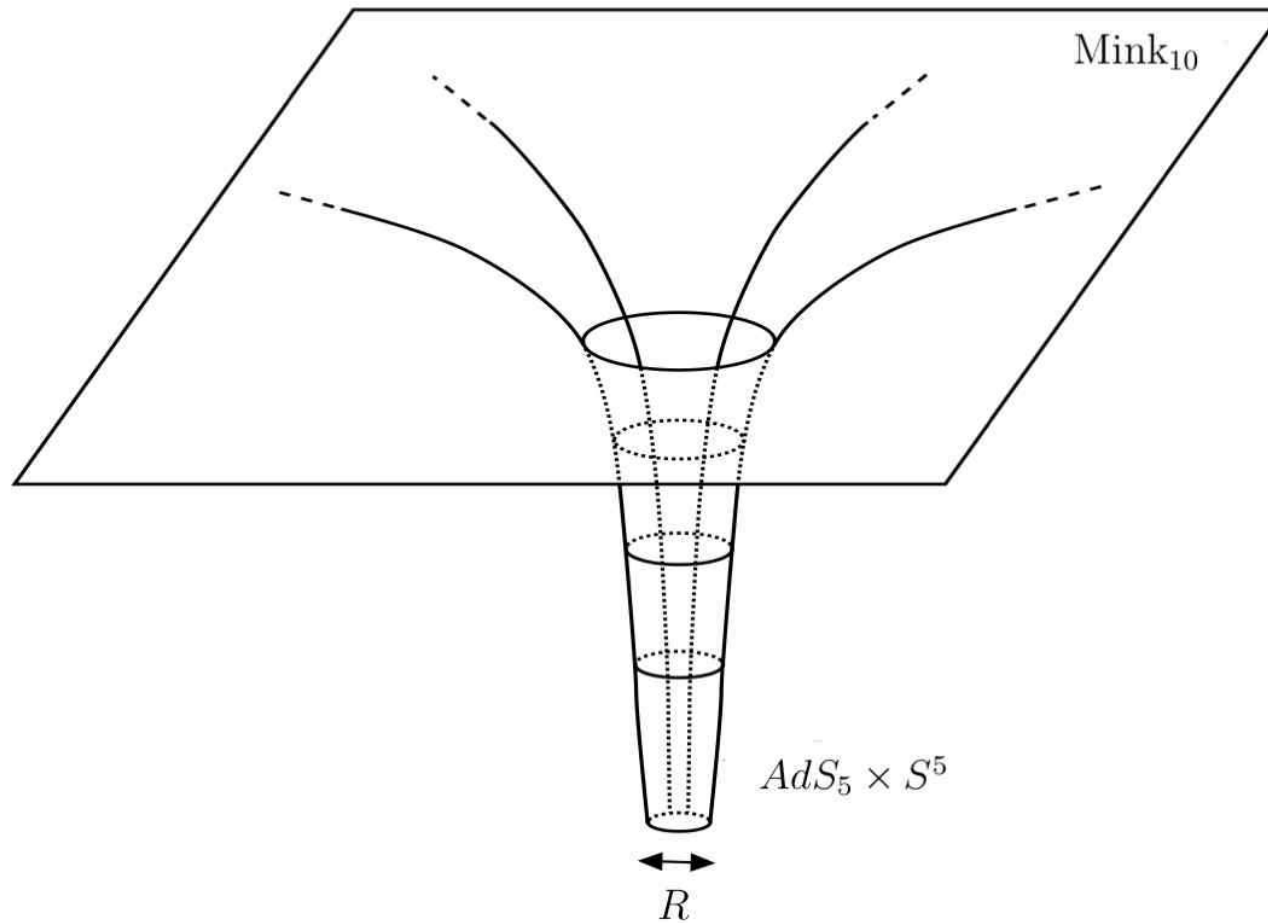
$$S = \frac{1}{16\pi G_{(10)}} \int d^{10}x \sqrt{-g} \left[R - \frac{F_5^2}{2 \cdot 5!} + \dots \right],$$

$$\int_{S^5} *F_5 \propto N_c, \quad G_{(10)} = \frac{\pi^4 L^8}{2N_c^2}, \quad H(r) = 1 + \frac{L^4}{r^4}, \quad f(r) = 1 - \frac{r_0^4}{r^4}$$

$$ds^2 = -\frac{f(r)}{\sqrt{H(r)}} dt^2 + \frac{1}{\sqrt{H(r)}} (dx_1^2 + dx_2^2 + dx_3^2) + \frac{\sqrt{H(r)}}{f(r)} dr^2 + r^2 \sqrt{H} d\Omega_5^2$$

near extremal black *D3*-brane metric with horizon $r = r_0$
for $r_0 < r \ll L = R$, i.e. factorized metric for $AdS_5 \otimes S^5$

$$ds^2 = \frac{r^2}{R^2} (-f(r) dt^2 + d\vec{x}^2) + \frac{R^2}{r^2 f(r)} dr^2 + R^2 d\Omega_5^2$$

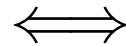


[from D. Mateos]

space-time around $D3$ -branes

AdS/CFT [Maldacena 98]

strongly coupled quantized conformal gauge theory
in $d = 4$ dimensions ($\mathcal{N} = 4$ SYM with $8N_c^2$ (1 gauge and 6
scalar) bosons and $(4N_c^2)$ Weyl fermions) **[[NOT QCD !]]**



weakly coupled classical supergravity (type IIB)
in $d = 10$ dimensions (on $AdS_5 \times S^5$)
via **holographic property**: radial coordinate $r_0 \leq r < \infty$ with
gauge theory on the boundary at ∞
in the limit:

't Hooft coupling $\lambda = g_{YM}^2 N_c$ is large, $N_c \rightarrow \infty$, $g_{YM}^2 \ll 1$

i.e. string coupling $g_s = \frac{g_{YM}^2}{4\pi} \ll 1$ – NO LOOPS

and

small curvature $\frac{l_s^4}{R^4} = \frac{1}{\lambda} \ll 1$ – RADIUS R of CURVATURE
is LARGE compared to the STRING SCALE $l_s = \sqrt{\alpha'}$

symmetries

duality maps the operators of QFT to the boundary values of the SUGRA fields

- QFT ($\mathcal{N} = 4$, $d = 4$ CFT)

$$S = -\frac{1}{g_{YM}^2} \int d^4x \left[\frac{1}{4} F_{\mu\nu}^a F^{a\mu\nu} + \frac{1}{2} D_\mu \Phi_i^a D^\mu \Phi_i^a + \dots \right]$$

.... massless Weyl fermions + interactions , $i = 1, \dots, 6$, $a = 1, \dots, N_c^2$

has Poincaré and conformal symmetry $SO(2, 4)$, e.g.

$x^\mu \rightarrow \lambda x^\mu$, $\Phi_i^a(x) \rightarrow \Phi_i^a(\lambda x)$, etc, $\beta(\mu) = 0$, and

$SU(\mathcal{N} = 4)$ symmetry and $SU(N_c)$

- $AdS_5 \otimes S^5$:

$$SO(2, 4) \otimes SO(6) = SO(2, 4) \otimes SU(4)$$

besides $SU(N_c)$ due to N_c parallel $D - 3$ branes

relaxation phenomena/strong coupling limit

hydrodynamic transport coefficients by gauge/gravity duality:

compare with quite involved AdS/CFT-gravity calculations

at Hawking temperature $T_H = \frac{r_0}{\pi R^2}$,

and for momentum $\omega, k \ll T_H$,

e.g. from sound channel dispersion up to $O(k^3)$, etc :

$$\frac{\eta}{s} = \frac{1}{4\pi}, \quad \tau_\pi = \frac{2 - \ln 2}{2\pi T}, \quad \kappa = \frac{\eta}{\pi T} = 2\lambda_1,$$

$$\lambda_2 = -\frac{\ln 2}{2\pi T}\eta, \quad \lambda_3 = 0$$

conformal $\Pi^{\mu\nu}$

all 5 second order terms classified by conformal symmetry

constitutive relation (d= 4):

$$\begin{aligned}\Pi^{\mu\nu} = & -2\eta\sigma^{\mu\nu} + 2\eta\tau_\pi u^\lambda \mathcal{D}_\lambda \sigma^{\mu\nu} \\ & + 4\lambda_1 \sigma^{\langle\mu}{}_\lambda \sigma^{\nu\rangle\lambda} + 2\lambda_2 \sigma^{\langle\mu}{}_\lambda \Omega^{\nu\rangle\lambda} + \lambda_3 \Omega^{\langle\mu}{}_\lambda \Omega^{\nu\rangle\lambda} \\ & + 2\kappa u_\alpha C^{\alpha\mu\nu\beta} u_\beta\end{aligned}$$

$C^{\alpha\beta\gamma\delta}$... Weyl tensor, $\Omega^{\alpha\beta}$... antisymmetric vorticity tensor

[Baier, Romatschke, Son, Starinets, Stephanov 07]

[Bhattacharyya, Hubeny, Minwalla, Rangamani 07, Loganayagam 08]

Bekenstein- Hawking entropy

thermodynamics: $dS = \frac{dE}{T}$

- Schwarzschild BH ($G = G_{(4)}, T_H = \frac{1}{4\pi r_0}$)

$$E = M = \frac{r_0}{2G}, \quad dE = \frac{dr_0}{2G}, \quad dS = \frac{4\pi}{2G} r_0 dr_0,$$

entropy

$$S = \frac{\pi r_0^2}{G} = \frac{A}{4G}$$

a universal relation with $A = 4\pi r_0^2$... area of the horizon

Bekenstein-Hawking entropy, cont.

● $AdS_5 \otimes S^5$: $T_H = \frac{r_0}{\pi R^2}$, $S = \frac{A_{(8)}}{4G_{(10)}}$

$$A_{(8)} = \left(\frac{r_0}{R}\right)^3 V_3 R^5 \Omega_{(5)} = \pi^6 T_H^3 R^8 V_3 , \quad G_{(10)} = \frac{\pi^4 R^8}{2N_c^2}$$

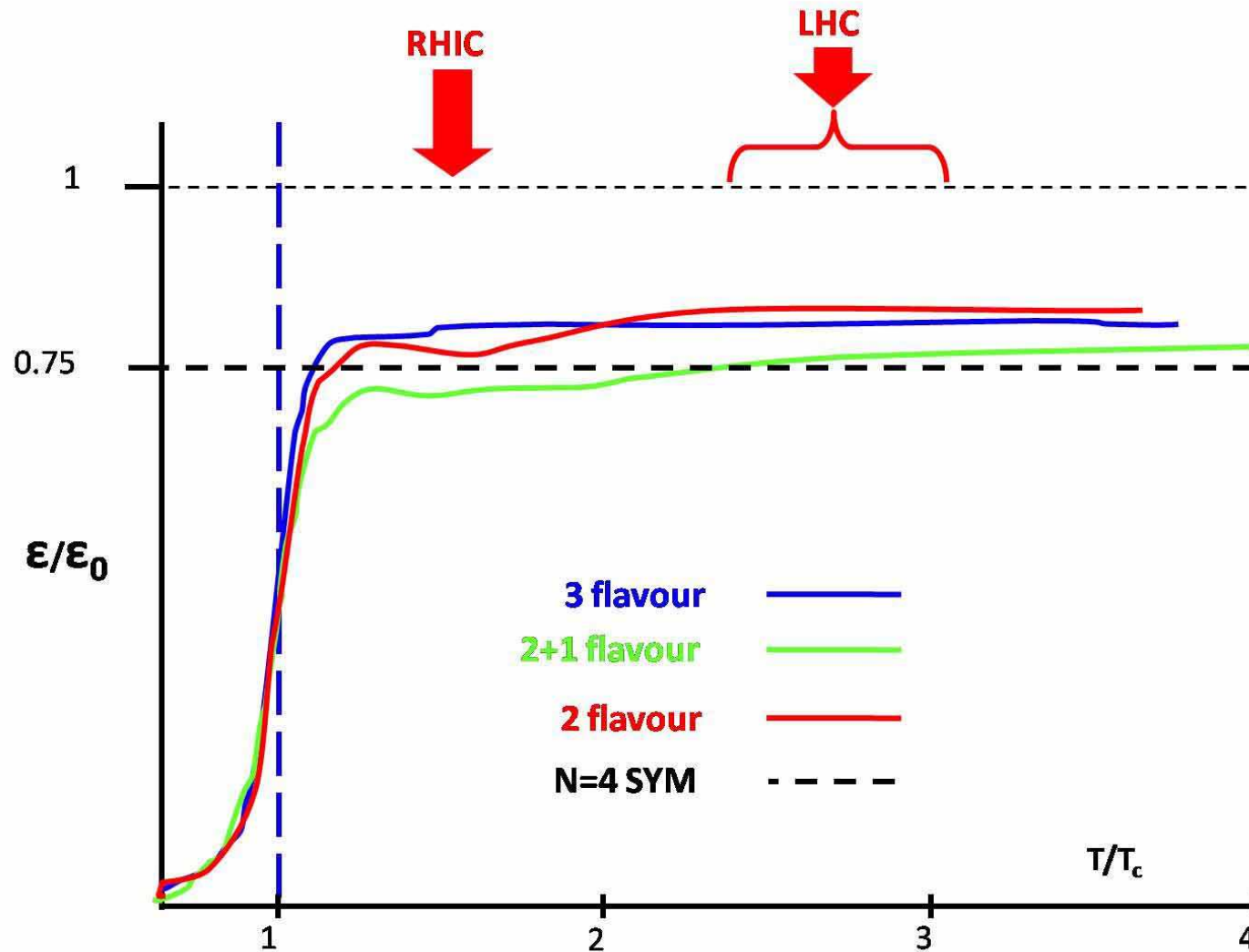
gives for the entropy density

$$s_{BH} = \frac{S}{V_3} = \frac{\pi^2 N_c^2}{2} T_H^3$$

NOTE:

$$s_{Boltzmann} = \frac{4p}{T} = \frac{4\pi^2 T^3}{90} \left[\frac{7}{4} \cdot 4 + 2 + 6 \right] N_c^2 = \frac{2\pi^2 N_c^2}{3} T^3$$

$$s_{BH} = \frac{3}{4} s_{Boltzmann}$$



energy density of QCD and SYM - via BH entropy

[from Myers and Vazquez 08]

master formula for AdS/CFT

schematically in terms of coinciding partition functions:

$$\int e^{iS_{4d}^{gauge} + i\Phi_0 O} = \int e^{iS_{5d}[\Phi]} \simeq e^{iS^{classical}[\Phi_0]}$$

S_{5d} is computed with non-trivial boundary condition
(holography)

$$\Phi(t, \vec{x}, r) \stackrel{r \rightarrow \infty}{\simeq} \Phi_0(t, \vec{x})$$

\implies

quantum correlation = classical two-point function

$$G_R(x, y) = -i \langle O(x)O(y) \rangle = - \frac{\delta^2 S^{classical}}{\delta\Phi_0(x)\delta\Phi_0(y)} \Big|_{r=\infty}$$

gauge: $O = T_{\mu\nu}$.. energy-momentum tensor

gravity: $\phi = g_{\mu\nu}$.. graviton

correlators from gravity, $d = 4$

response of the fluid to small metric perturbations

example: $g_{xy} = h_{xy}(t, z, r) \neq 0$, $u^0 = 1$, $T = \text{const}$

e.g. it leads from Christoffel symbols in the covariant derivatives to linear approximation:

$$\sigma^{xy} = \sigma_{xy} \approx \frac{1}{2}(\Gamma_{x0}^y + \Gamma_{y0}^x) \approx \frac{1}{2}\partial_t h_{xy}$$

and for the stress tensor

$$\delta\Pi^{xy} \approx -\eta\partial_t h_{xy} + \eta\tau_\pi\partial_t^2 h_{xy} - \frac{\kappa}{2}[\partial_t^2 h_{xy} + \partial_z^2 h_{xy}]$$

Fourier transform $h(t, z) = \exp(-i\omega t + ikz) h(\omega, k)$, etc.

$$\delta\Pi^{xy}(\omega, k) = -G_R^{xy,xy}(\omega, k)h_{xy}(\omega, k)$$

via **linear response**

correlators, cont.

retarded Green function:

$$G_R^{xy,xy}(\omega, k) = -i\eta\omega + \eta\tau_\pi\omega^2 - \frac{\kappa}{2}(\omega^2 + k^2) + \dots$$

i.e. **Kubo formula**

$$\begin{aligned}\eta &= \lim_{\omega} \frac{1}{\omega} iG_R^{xy,xy}(\omega, \vec{0})|_{\omega=0} \\ &= \lim_{\omega} \frac{1}{\omega} \int dt d^3x \exp(i\omega t) \langle [T_{xy}(t, \vec{x}), T_{xy}(0, \vec{0})] \rangle |_{\omega=0}\end{aligned}$$

quasinormal mode

G_R from gravity action with AdS_5 metric:

$$S_{5d} = \frac{N_c^2}{8\pi^2 R^3} \int d^5x [\sqrt{-g}R_{5d} + \dots],$$

factor from $\frac{1}{16\pi G_{(10)}}R^5\Omega_{(5)}$, $\Omega_{(5)} = \pi^3$

and “dictionary” $G_{(10)} = 8\pi^6 g_s^2 l_s^8$, $g_{YM}^2 N_c = 4\pi g_s N_c = R^4/l_s^4$
scalar mode and EoM (of Heun’s type)

$$\sqrt{-g}R_{(5d)} \rightarrow -\frac{1}{2}\sqrt{-g}g^{\mu\nu} \partial_\mu \Phi \partial_\nu \Phi, \quad \Phi \equiv h_{xy} :$$

$$\Phi(t, z, u = \frac{r_0^2}{r^2}) = \exp(-i\omega t + ikz)\Phi_k(u), \quad \Phi'_k = \frac{d}{du}\Phi_k,$$

$$\Phi_k'' - \frac{1+u^2}{u(1-u^2)}\Phi_k' + \frac{1}{4\pi^2 T^2} \frac{\omega^2 - k^2(1-u^2)}{u(1-u^2)^2}\Phi_k(u) = 0$$

quasinormal mode, cont.

solution with ansatz $\Phi_k(u) = f_k(u)\Phi_0(k)$, $f_k(0) = 1$ and
boundary condition: only the incoming wave, the one which
moves toward the horizon at $r- > r_0 \equiv u = 1$, i.e. nothing
comes out the horizon

$$f_k(u) = (1 - u^2)^{-i\omega/(4\pi T)} F_k(u) \approx 1 + i\frac{\omega}{4\pi T}u^2 + \dots$$

note, near $u \sim 1$:

$$\exp(-i\omega t)f_k(u) \rightarrow \exp[-i\omega(t + r^*)] , \quad r^* = \frac{\ln(1 - u)}{4\pi T}$$

moves from $u = 0$ at $t = 0$ to $u = 1$ at $t- > \infty$

classical gravity action

inserting the solution into $S_{(5d)}$

$$S_{(5d)} \approx \frac{N_c^2}{8\pi^2 R^3} \int d^4x \int_0^1 du \sqrt{-g} \left(-\frac{1}{2}\right) g^{uu} \partial_u \Phi \partial_u \Phi$$

and keeping only the surface contribution at $u = 0$, \Rightarrow

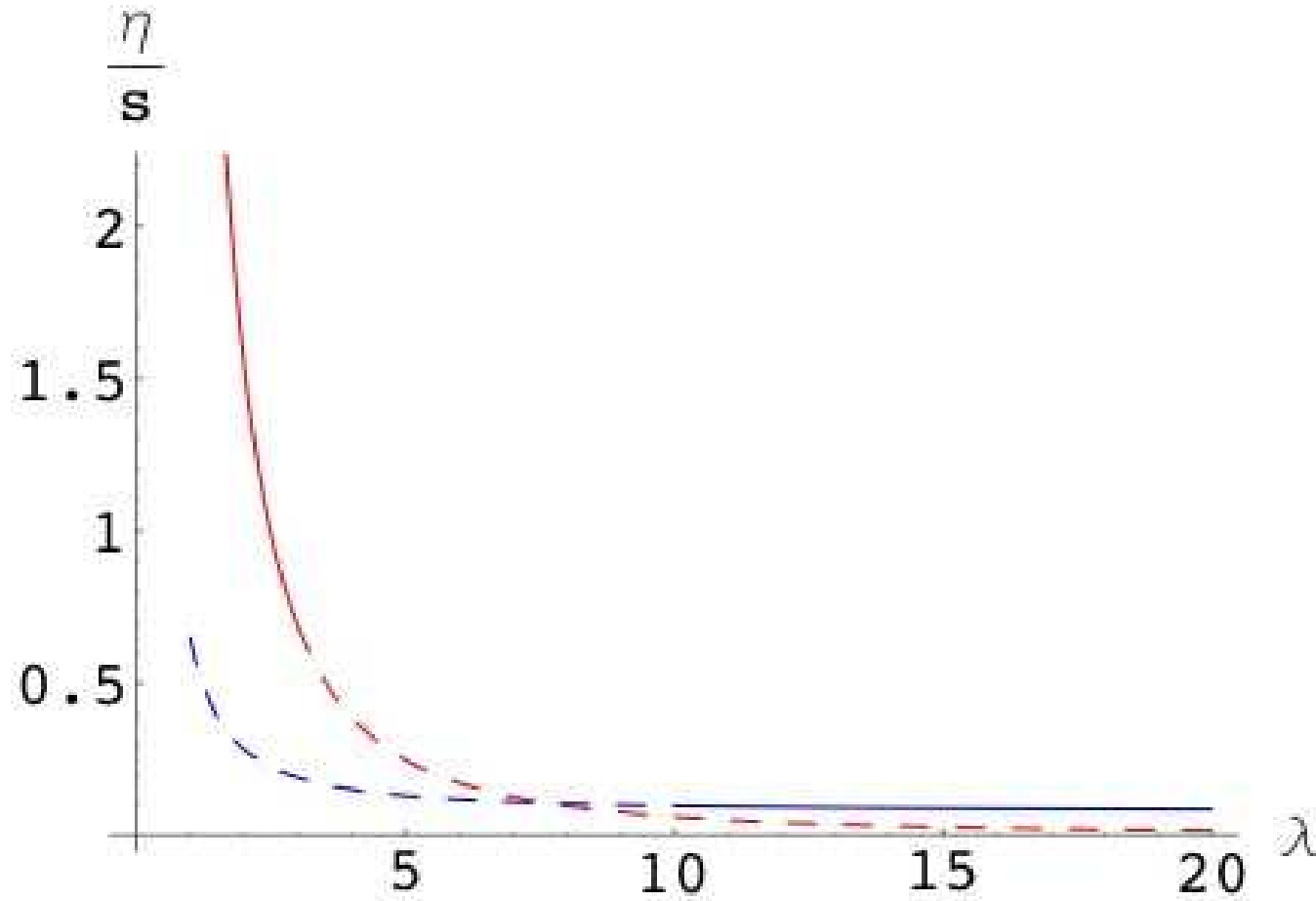
$$S^{classical} = -\frac{\pi^2 N_c^2 T^4}{8} \int \frac{d^4k}{(2\pi)^4} \Phi_0(-k) \left[\frac{(1-u^2)}{u} f'_k(u) f_k(u) \right] \Big|_{u=0} \Phi_0(k)$$

and finally

$$G_R^{xy,xy}(\omega, k) = \frac{\pi N_c^2 T^3}{8} \left[-i\omega + \frac{1 - 2 \ln 2}{2\pi T} \omega^2 - \frac{1}{2\pi T} k^2 \pm \dots \right],$$

$$\eta = \frac{\pi N_c^2 T^3}{8}, \quad \tau_\pi = \frac{2 - \ln 2}{2\pi T}, \quad \kappa = \frac{\eta}{\pi T}$$

$$\eta/s$$



behaviour of η/s as a function of the 't Hooft coupling

shear viscosity η

near equilibrium: $\eta \simeq \epsilon \bar{v} \lambda_f$, *entropy density* $s \simeq \epsilon/m \rightarrow$

$$\frac{\eta}{s} \simeq m \bar{v} \lambda_f \simeq \hbar \frac{\text{mean free path}}{\text{deBroglie wavelength}}$$

- dilute system (QFT \rightarrow kinetic theory \rightarrow hydro):
scale $\lambda_f \rightarrow \frac{\eta}{s} \gg \hbar$, e.g. **pQCD** ($N_f = 0$)

$$\frac{\eta}{s} \simeq 3.8 \frac{1}{g^4 \ln(2.8/g)} \simeq O(1) \text{ for } g = 2.5$$

BUT with $\ln(2.8/g) \simeq O(1) : \frac{\eta}{s} \simeq 0.1 \rightarrow$ **sensitive to constant under the log !**

[Arnold, Moore and Yaffe 03]

- strongly coupled system (QFT \rightarrow hydro):
scale $1/T \rightarrow \frac{\eta}{s} = \frac{\hbar}{4\pi} \simeq 0.08$

[Policastro, Son and Starinets 01]

sound mode

consistency requirement from sound channel:

$$\omega_{1,2} = \pm c_s k - i\Gamma k^2 \pm \frac{\Gamma}{c_s} \left(c_s^2 \tau_\pi - \frac{\Gamma}{2} \right) k^3 + \mathcal{O}(k^3),$$

with $\Gamma = \frac{2}{3} \frac{\eta}{\epsilon + p}$

from gravity perturbation $g_{tz} = h_{tz} \ll 1$

$$\omega_{1,2} \pm \frac{1}{\sqrt{3}} k - i \frac{1}{6\pi T} k^2 \pm \frac{3 - 2 \ln 2}{24\pi^2 T^2 \sqrt{3}} k^3$$

gives sound velocity $c_s = \frac{1}{\sqrt{3}}$ and

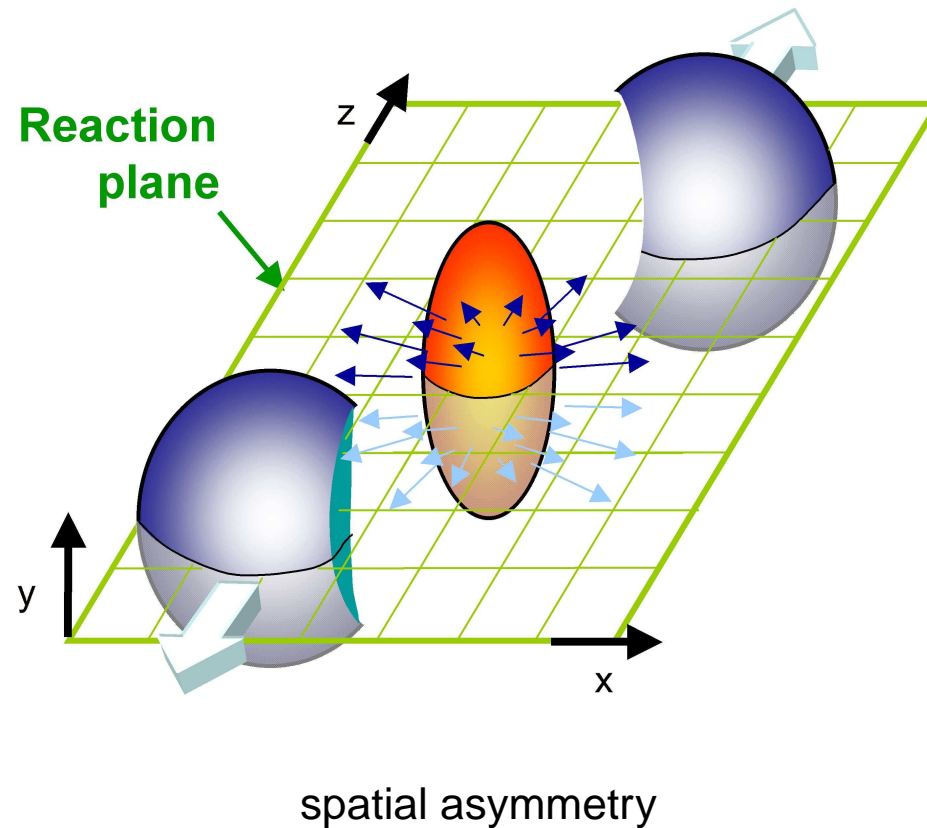
$$\frac{\eta}{s} = \frac{1}{4\pi}, \quad \tau_\pi = \frac{2 - \ln 2}{2\pi T}$$

hydro + RHIC

heavy-ion collisions require **beyond hydrodynamics**:

- hydrodynamics = differential equations
initial conditions !
- initial = equilibration time
- distribution of energy density (Glauber? CGC ?)
- QCD equation of state
- hadronisation prescription (Cooper-Frye?)

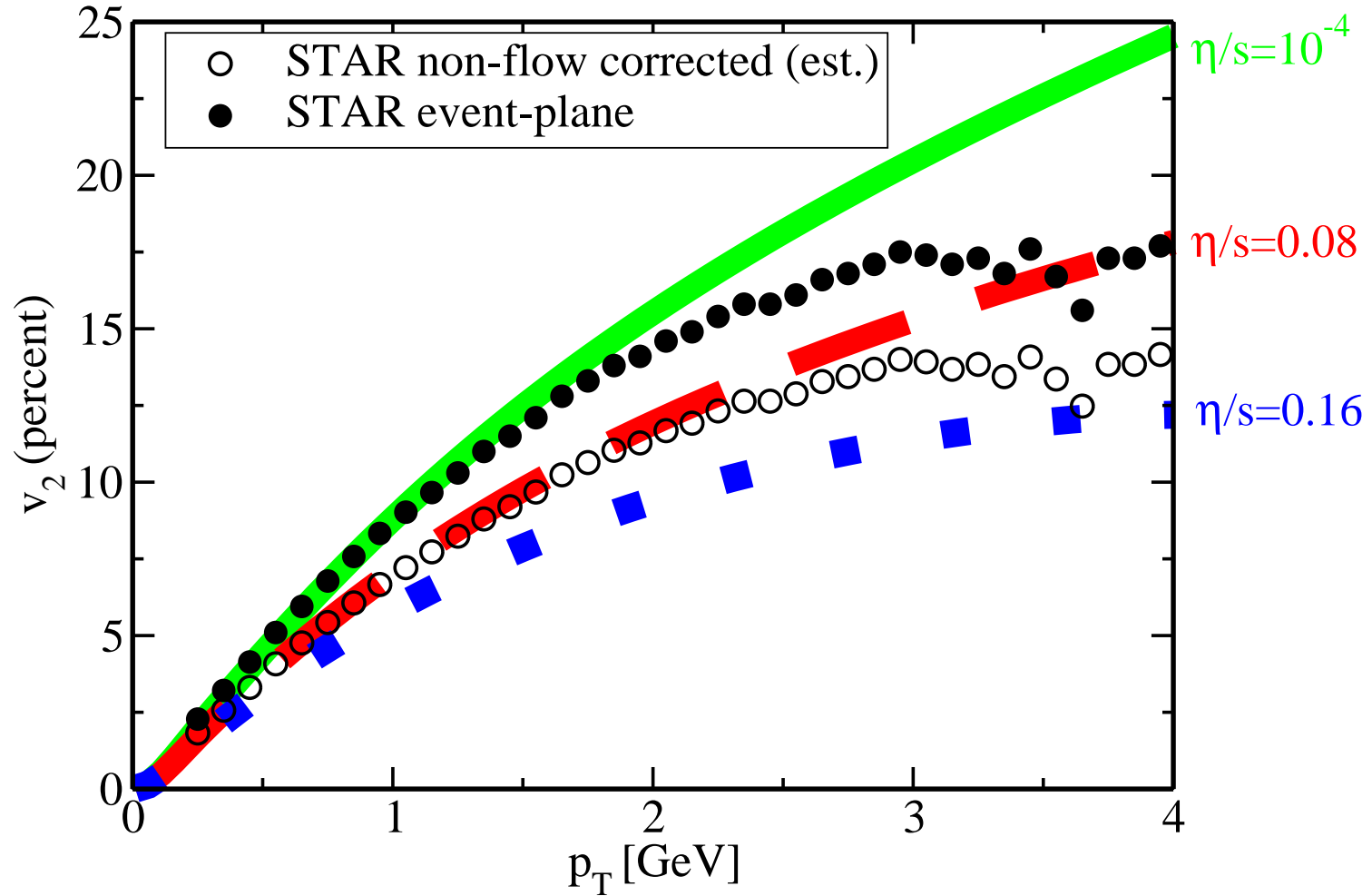
as important example



$$\frac{dN}{dy dp_{\perp} d\phi} = \left\langle \frac{dN}{dy dp_{\perp} d\phi} \right\rangle_{\phi} (1 + 2v_2(p_{\perp}) \cos(2\phi) + \dots)$$

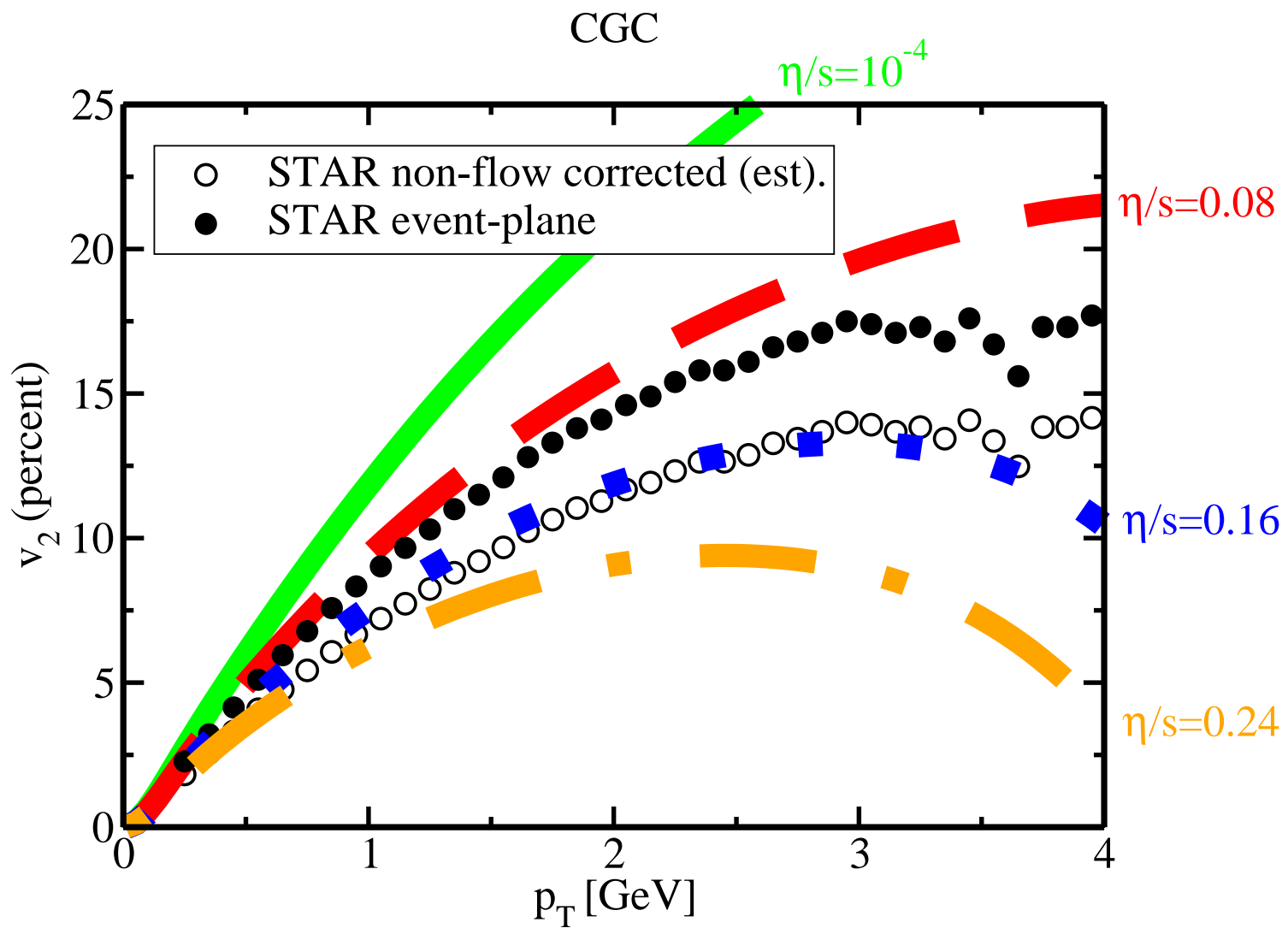
● elliptic flow: $v_2(p_{\perp})$

Glauber



elliptic flow

[Luzum and Romatschke 08]



elliptic flow

[Luzum and Romatschke 08]

main results [Luzum and Romatschke 08]

- viscous hydrodynamics simulation give a good description of RHIC data for

$$\frac{\eta}{s} = 0.1 \pm 0.1(\text{theory}) \pm 0.08(\text{experiment})$$

- modest estimate:

$$\frac{\eta}{s} < 0.5$$

- weak dependence on the values of the second-order parameters $\tau_\pi, \lambda_1, \dots$

- early thermalisation time is questioned, but

$$\tau_0 < 2 \text{ fm}$$

transport coefficients

kinetic QCD theory (weak coupling):

$$\frac{\eta/s}{\tau_\pi T} \simeq 0.17-0.2, \quad \lambda_1 \simeq -(2.0-2.2) \frac{\eta^2}{sT}, \quad \lambda_2 = -2\tau_\pi\eta, \quad \lambda_3 = \kappa = 0$$

[York and Moore 08]

finite 't Hooft coupling $\lambda \equiv g_{YM}^2 N_c$, $\lambda \gg 1$ corrections to coefficients by gauge/gravity duality, e.g.:

[Buchel and Paulos 08]

$$\frac{\eta/s}{\tau_\pi T} = 0.383 (1 - 3.52 \lambda^{-3/2} + \dots),$$

$$\kappa = \frac{\eta}{\pi T} \left(1 - \frac{145\zeta(3)}{8} \lambda^{-3/2} + \dots \right)$$

excitement about gauge/gravity correspondence:
mainly to gain intuition into **STRONG COUPLING**

ASKING FOR MORE:

Is there an experiment whose outcome could cast strong doubts on the relevance of AdS/CFT to understand QCD ?

[P. Jacobs 08]

JET PHYSICS ?

[Hatta, Iancu and Mueller 07 - 08]

MOST CHALLENGING TASK of the theory is to find the microscopic mechanism for the rather **RAPID EQUILIBRATION** of matter in RHIC collisions

EXTRAS

COMPARISON

	QCD	$\mathcal{N}=4$ SYM
$T=0$	$N_c=3=N_f$, confinement, discrete spectrum, scattering,	N_c large, N_f/N_c small, deconfined, conformal, supersymmetric,
	very different !!	
$T > T_c$	strongly-coupled plasma of gluons & fundamental matter deconfined, screening, finite corr. lengths, . . .	strongly-coupled plasma of gluons & adjoint and fundamental matter deconfined, screening, finite corr. lengths, . . .
	very similar !!	
$T \gg T_c$	runs to weak coupling	remains strongly-coupled
	very different !!	

QCD and $\mathcal{N} = 4$ SYM as a function of temperature

[from Myers and Vazquez 08]

entropy current

Israel - Stewart 79: $s^\mu = (s - \frac{\tau_\pi}{4\eta T} \Pi_{\alpha\beta} \Pi^{\alpha\beta}) u^\mu$

\Rightarrow **second law:**

$$\nabla_\mu s^\mu = \frac{\Pi_{\alpha\beta} \Pi^{\alpha\beta}}{2\eta T} \geq 0$$

instead in **causal viscous and conformal hydrodynamics**
a more general current in terms of u^μ and its derivatives:

$$s^\mu = s u^\mu + (\# \sigma^2 + \# \Omega^2) u^\mu + O(u \nabla^2 u)$$

with

$$\nabla_\mu s^\mu = \frac{\eta}{2T} \sigma^{\mu\nu} \sigma_{\mu\nu} + \frac{1}{4T} (\kappa - 2\lambda_1) \sigma_\nu^\mu \sigma_\lambda^\nu \sigma_\mu^\lambda$$

in $\mathcal{N} = 4$ SYM: $\kappa = 2\lambda_1$

[Loganayagam 08]

Bjorken flow

boost-invariant (irrotational) 1 + 1 flow [Bjorken 83]
second-order equations (proper time τ , Φ ... viscous flow):

$$\partial_\tau \epsilon = -\frac{4\epsilon}{3\tau} + \frac{\Phi}{\tau}$$

$$\tau_\pi \partial_\tau \Phi = \frac{4\eta}{3\tau} - \Phi - \frac{4\tau_\pi}{3\tau} \Phi - \frac{\lambda_1}{2\eta^2} \Phi^2$$

non-linear term NOT in MIS theory!

[BRSSS 07]

compare with AdS/CFT calculation:

$$\frac{\lambda_1 T}{\eta} = \frac{1}{2\pi} \left[1 + \frac{215 \zeta(3)}{8} \lambda^{-3/2} + \dots \right]$$

[Janik, Peschanski, Heller 06; Buchel, Paulos 08]

Müller - Israel - Stewart theory

keeping essentially **one term** in the derivative expansion
up to **second order**

$$\Pi^{\mu\nu} = -2\eta\sigma^{\mu\nu} + 2\eta \tau_\pi \langle D\sigma^{\mu\nu} \rangle, \quad D = u \cdot \nabla$$

remark: does not match with AdS/CFT $\mathcal{N} = 4$ SYM

$$\text{sound} : \tau_\pi = \frac{2 - \ln 2}{2\pi T}$$

$$\text{Bjorken flow} : \tau_\pi = \frac{1 - \ln 2}{2\pi T}$$

\Rightarrow **all second-order terms** consistent with
conformal symmetry have to be included !

equilibration time

REMINDER:

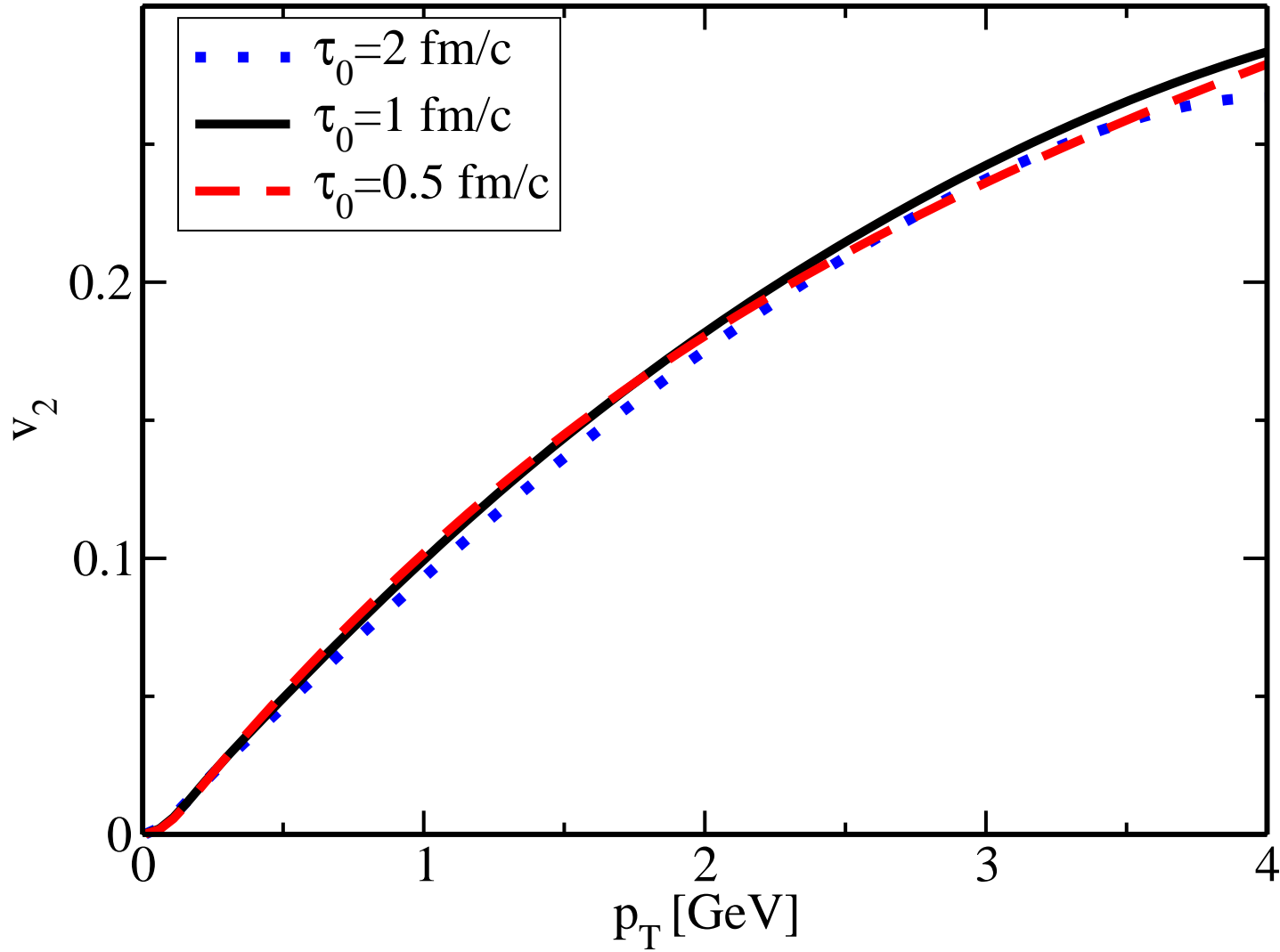
claim of short $\tau_0 \equiv \tau_{eq} \leq 0.5$ fm at RHIC

OPEN QUESTION:

CGC ($\alpha_s \ll 1$) \rightarrow *sQGP* ($\alpha_s > O(1)$)

within a very short time < 0.5 fm ?

(b)



early thermalisation ?

[Luzum and Romatschke 08]

parametric pQCD estimate

for thermalisation in an expanding gluonic medium

near equilibrium at $T(\tau)$: **Knudsen number Kn**

$$\frac{1}{Kn} = \frac{\text{longitudinal expansion time}}{\text{mean free path}} = \frac{\tau}{\lambda_f} \gg 1$$

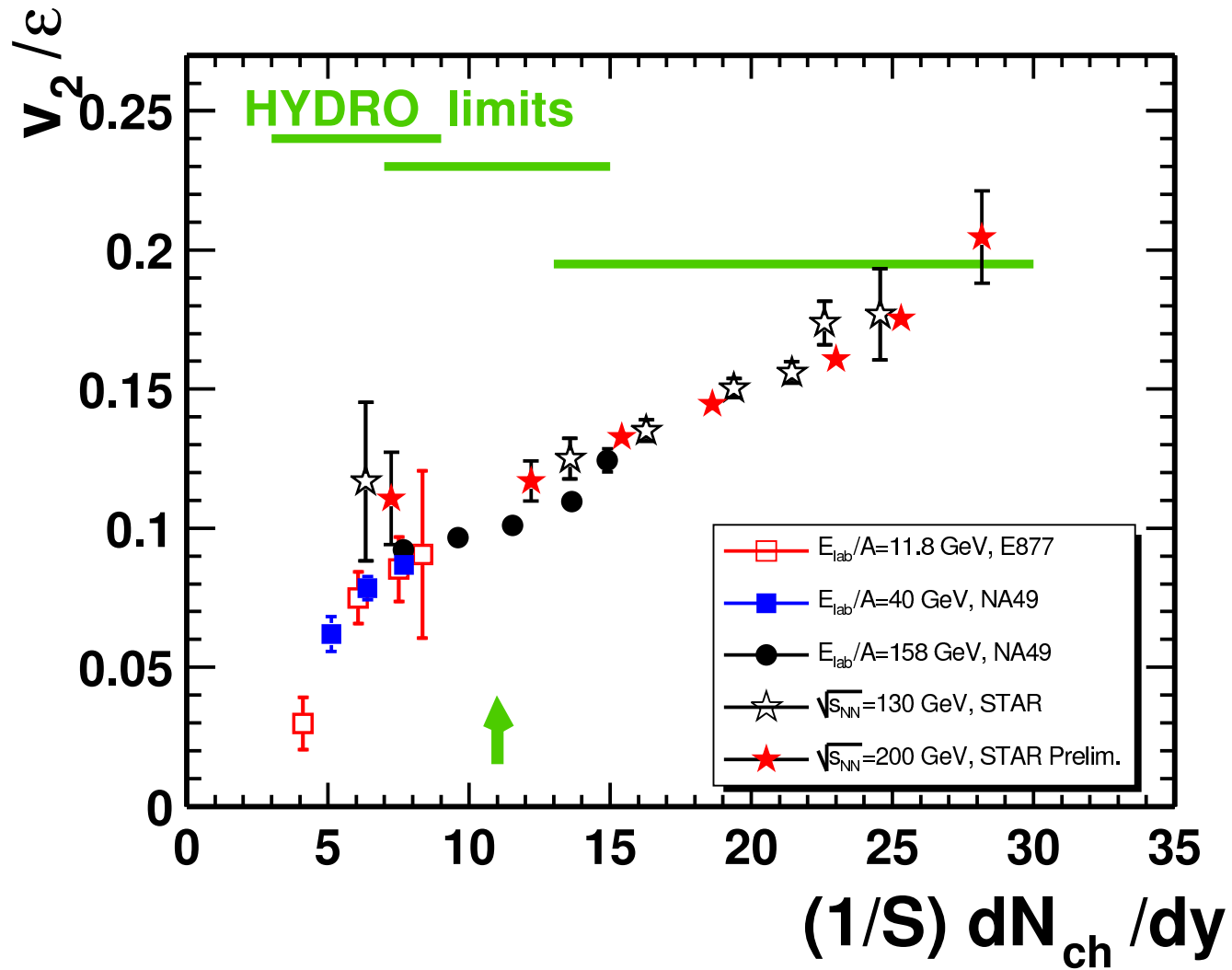
from many gluon interactions (including saturation):

Arnold et al.: $\tau_{eq} Q_s \geq \alpha_s^{-7/3}$

‘bottom-up’ [Baier, Son, Mueller and Schiff 01]

$$\tau_{eq} Q_s \geq \alpha_s^{-13/5}$$

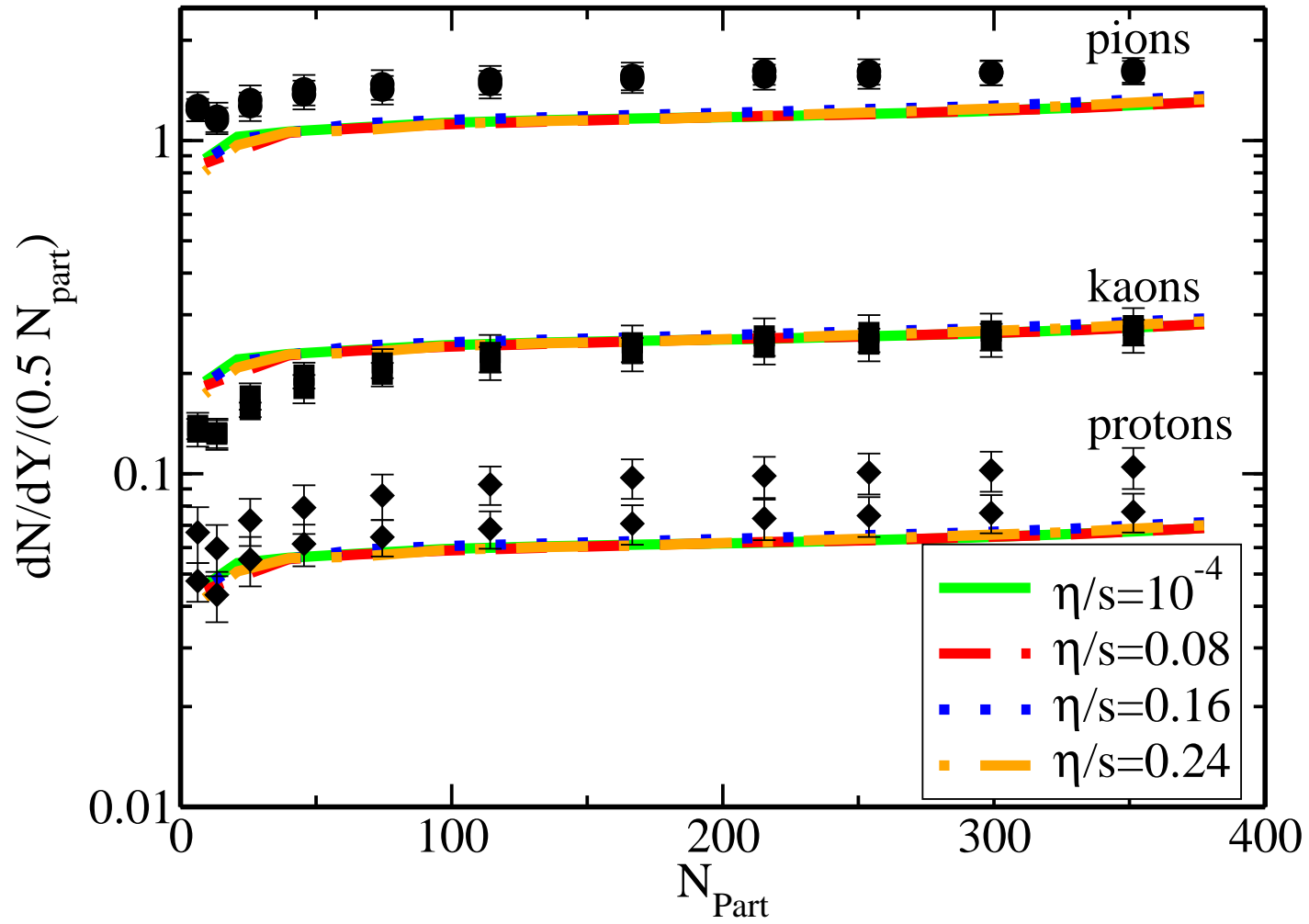
$$\text{RHIC: } \tau_{eq} \geq 2 - 3 \text{ fm}$$



v_2 : experiment vs. perfect hydro

[Heinz 04]

CGC



multiplicity

[Luzum and Romatschke 08]