

# Physics mit Schmaltz

## Lecture II

Locality: can expand  $\mathcal{L}$  in power series in  $\phi, \psi, \partial_\mu$

$$\mathcal{L} = c_0 M^4 + c_2 M^2 H^\dagger H + M \mathcal{L}_3 + \mathcal{L}_4 + \frac{\mathcal{L}_5}{M} + \frac{\mathcal{L}_6}{M^2} + \dots$$

mass dimensions:  $[\partial_\mu] = 1$      $[\mathcal{L}] = 4$

$[\phi] = 1$      $[\psi] = 3/2$

$[M] = 1$   some UV mass scale

Examine dimensionless couplings first  $\mathcal{L}_4$

$$\mathcal{L}_4^{\text{toy}} \sim (\partial\phi)^2 + \lambda^2 \phi^4 + \bar{\psi} \not{\partial} \psi + \lambda_t \phi \bar{\psi} \psi$$

What is the theoretically natural/expected size for

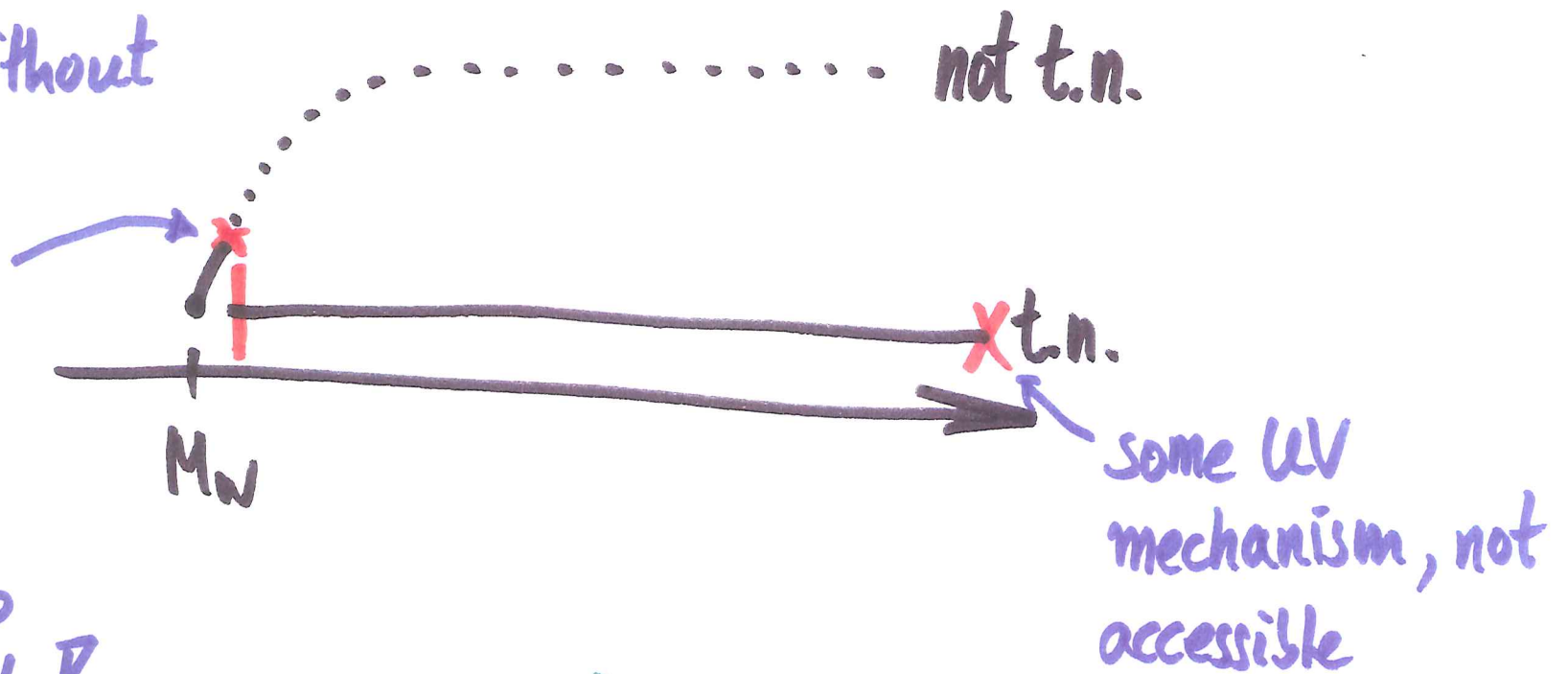
the couplings  $\lambda, \lambda_t$  ?

(at the weak scale)

for a technically natural coupling the mechanism which explains smallness can be in the far UV

in the case without tech.nat. the mechanism must be in the IR

⇒ accessible to experiment!



a coupling which is not tech.nat. is good news for experiment.

in  $\mathcal{L}_4^{SM}$  all\* couplings are tech. nat.

⊖ no sign  
of N.P.  
nearby.

$\lambda_{\text{quarks}}, \lambda_{\text{leptons}}$  chiral symmetry

$g_s, g, g'$  gauge symmetry

$\lambda_{\text{Higgs}} \sim 0.1$  no new symmetry in  $\lambda_H \rightarrow 0$  limit but  
 $\lambda_H$  has natural size in SM

$$\delta \lambda_{\text{Higgs}} \sim 0.1$$

the rest of the SM Lagrangian ...

$$C_0 M^4 + C_2 M^2 H^\dagger H + \mathcal{L}_4 + \underbrace{\frac{\mathcal{L}_5}{M} + \frac{\mathcal{L}_6}{M^2} + \dots}_{\text{tomorrow}}$$

cosmo  
const.

Higgs mass  
term

tomorrow



is  $m^2 H^\dagger H$  tech. nat.?

symmetry:  $H \rightarrow H + \text{const}$  can forbid  $m^2 H^\dagger H$

but: broken badly by other couplings

$$\lambda_t Q H U^c$$

$$\lambda^2 H^4$$

$$g^2 W_\mu W^\mu H^2$$

how about scale invariance symmetry?

$$\begin{array}{l} \phi \rightarrow s \phi(sx) \\ \psi \rightarrow s^{3/2} \psi(sx) \end{array} \quad \begin{array}{l} \text{equivalent} \\ \text{to} \end{array} \quad \begin{array}{l} \phi \rightarrow s \phi(x) \\ \psi \rightarrow s^{3/2} \psi(x) \\ d^4x \rightarrow \frac{1}{s^4} d^4x \\ \partial \rightarrow s \partial \end{array}$$



# Scale invariance and the Higgs

SM approximately  
Scale inv't near  $M_{\text{weak}}$

broken by

$$m_H^2 H^\dagger H$$
$$g(\mu/M)$$

$$\frac{d}{d \ln \mu} m_H^2 \propto m_H^2 \frac{g^2}{16\pi^2} \dots$$

technically natural  
in IR effective theory

problem: SI must be strongly broken (in UV)

( $\Rightarrow$  cannot be used to argue that  $m_H$  is natural)

why? assume no new scale...



this problem does not occur for  $\lambda_{up}$

ie. there is no coupling getting large and  
breaking  $u^c \rightarrow e^{i\theta} u^c$  symmetry

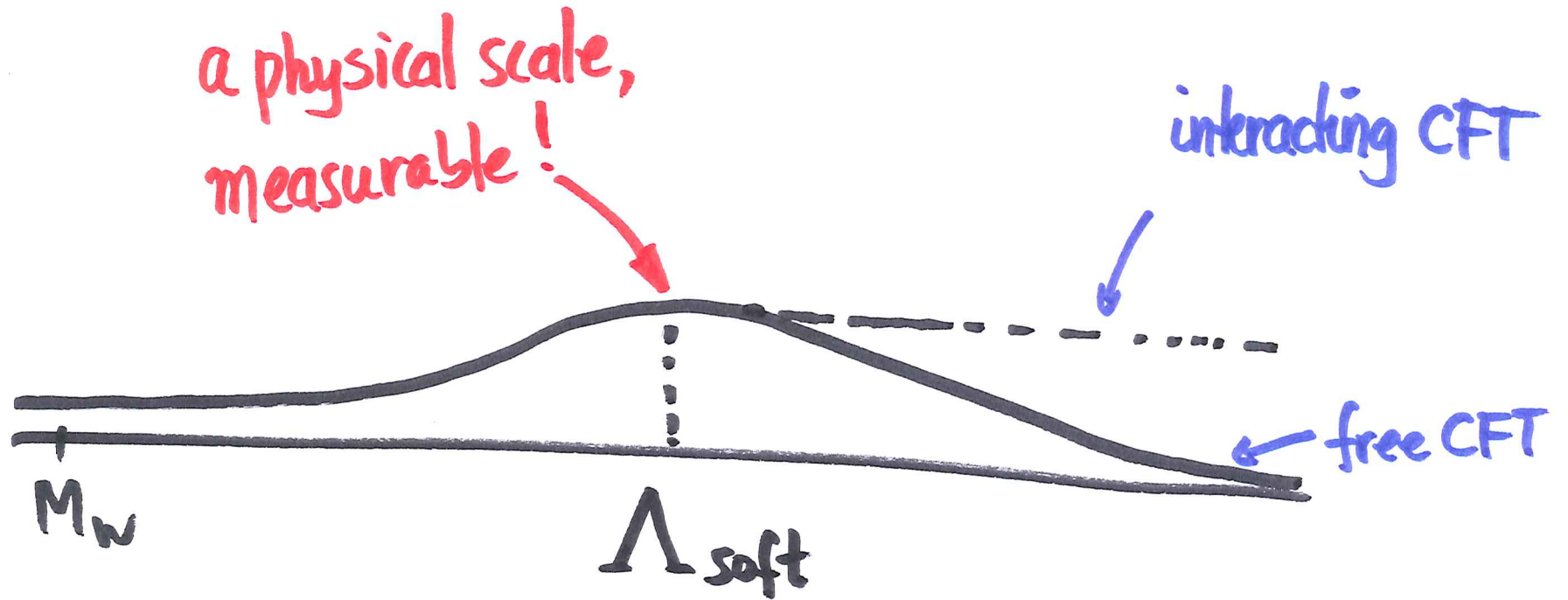


$\lambda_{up} QH u^c$

consistent to assume no large chiral  
symmetry violation in UV

not consistent to assume no scale  
violation in UV

must "soften" interactions



$$m_H^2 \propto \Lambda_{\text{soft}}^2 \Rightarrow \Lambda_{\text{soft}} \lesssim \frac{4\pi}{g'} m_H \rightarrow \text{LHC}$$

or fine-tuning ☹️

final points:

$\Lambda_{\text{soft}}$ ? embed  $U(1)_Y$  in non-Abelian group

$M_{\text{pl}}$ ? perhaps QFT does not apply to gravity ... desperate ... I know...



other symmetries that can make Higgs mass  
(technically) natural:

+ SUSY  $H \leftrightarrow \psi$

+ Shift-(Goldstone)-sy.  $H \rightarrow H + \text{const}$

+ Combinations (twin,...)



# "partner" symmetries



UV-sensitive

symmetry  
↔

cancel



UV-sensitive

$$\delta m_H^2 \sim m_{\tilde{t}}^2 \frac{\lambda_t^2}{16\pi^2}$$

e.g. SUSY



$$\Rightarrow \delta m_H^2 = -\frac{3}{4\pi^2} \lambda_t^2 m_{\tilde{t}}^2 \log \frac{\Lambda_{UV}}{m_{\tilde{t}}}$$

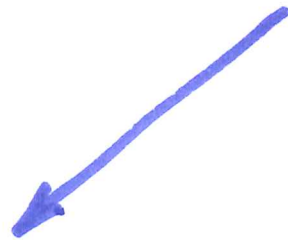
$$\left| \frac{\delta m_H^2}{m_H^2} \right| \approx 40 \left[ \frac{m_{\tilde{t}}}{\text{TeV}} \right]^2 \left[ \frac{\log(M_{UV}/m_{\tilde{t}})}{10} \right]$$

125 GeV

Somewhat tuned ... similar in all other known partner models

# the SM Lagrangian

$$c_0 M^4 + c_2 M^2 H^\dagger H + \mathcal{L}_4 + \frac{\mathcal{L}_5}{M} + \frac{\mathcal{L}_6}{M^2} + \dots$$




only 1 term allowed:  $\frac{(HL)^2}{10^{14} \text{ GeV}}$  or Dirac neutrino mass  $\lambda H L N^c$

Optimistic: new physics @ few TeV  $\equiv M$

(probably coupled to Higgs  $\leftrightarrow$  naturally)

integrate out ...

$$\mathcal{L}_6 \sim \frac{H^\dagger \not{D}_\mu H H^\dagger \not{D}^\mu H}{M^2} + \frac{H^\dagger \not{D}_\mu H \bar{e}_R \gamma^\mu e_R}{M^2} + \frac{H^\dagger \sigma^a H W_{\mu\nu}^a B^{a\mu\nu}}{M^2} + \mathcal{L}_7 + \mathcal{L}_8 + \dots$$
$$+ \frac{H^\dagger H B^{\mu\nu} B_{\mu\nu}}{M^2} + \frac{H^\dagger H W_{\mu\nu}^a W^{a\mu\nu}}{M^2} + > 100 \text{ more at dim 6}$$

predictive?   $\sim \frac{1}{q^2} + \frac{1}{M^2} = \frac{1}{q^2} \left(1 + \frac{q^2}{M^2}\right)$  expansion in  $q^2/M^2$

# most violate flavor symmetries

e.g.  $(\bar{d}_L \gamma_\mu S_L)^2 + \text{h.c.}$   $\frac{1}{M^2}$  complex  $\epsilon_K: M \gtrsim 10^4 \text{ TeV}$

real  $\Delta m_K: \gtrsim 10^3 \text{ TeV}$

$(\bar{c}u)^2$

$D-\bar{D}: \gtrsim 10^3 \text{ TeV}$

$(\bar{d}b)$

$B-\bar{B}: \gtrsim 3 \cdot 10^2 \text{ TeV}$

$(\bar{s}b)$

$B_s-\bar{B}_s: \gtrsim 10^2 \text{ TeV}$

assume minimal flavor violation - tech. nat.

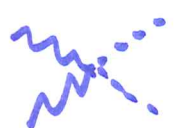


# A Higgs physics example

$$\mathcal{L}_6 \sim c_1 g'^2 \frac{H^\dagger H B_{\mu\nu} B^{\mu\nu}}{M^2} + c_2 g^2 \frac{H^\dagger H W_{\mu\nu}^a W^{a\mu\nu}}{M^2} + c_{12} g g' \frac{H^\dagger \sigma^a H W_{\mu\nu}^a B^{\mu\nu}}{M^2}$$

plug in  $\langle H \rangle$ :  $H = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v+h \end{pmatrix}$

expand ...  $v^2 F_{\mu\nu}^2 + v h F_{\mu\nu}^2 + h^2 F_{\mu\nu}^2$

PEW  $h \rightarrow v V_{\mu\nu}$  

$$\rightarrow c_1 g'^2 \frac{v^2}{2M^2} B_{\mu\nu} B^{\mu\nu} + c_2 g^2 \frac{v^2}{2M^2} W_{\mu\nu}^a W_a^{\mu\nu} + c_{12} g g' \frac{v^2}{2M^2} W_{\mu\nu}^3 B^{\mu\nu}$$

Corrections to  $SU(2) \times U(1)$  kinetic terms,  
rescale  $B_\mu, W_\mu^a \Rightarrow$  redefine  $g, g'$

$SU(2)$  breaking  $Z$ - $\gamma$  mixing

$$"S" \equiv 16\pi c_{12} \frac{v^2}{M^2}$$

$$+ c_1 g'^2 \frac{v^2}{M^2} \frac{\hbar}{v} B_{\mu\nu}^2 + c_2 g^2 \frac{v^2}{M^2} \frac{\hbar}{v} (W_{\mu\nu}^a)^2 + c_{12} g g' \frac{v^2}{M^2} \frac{\hbar}{v} W_{\mu\nu}^3 B^{\mu\nu}$$

$$= \underbrace{4e^2 \frac{v^2}{M^2} (c_1 + c_2 + c_{12}) \frac{\hbar}{v} \frac{F_{\mu\nu}^2}{4}}_{\equiv C_{\gamma\gamma} e^2} + \dots \frac{\hbar}{v} F_{\mu\nu} Z^{\mu\nu} + \dots \frac{\hbar}{v} Z_{\mu\nu} Z^{\mu\nu}$$

$\uparrow$   $C_{\gamma Z}$                        $\uparrow$   $C_{ZZ}$

Higgs decays:

$\gamma Z$

$Z Z^*$



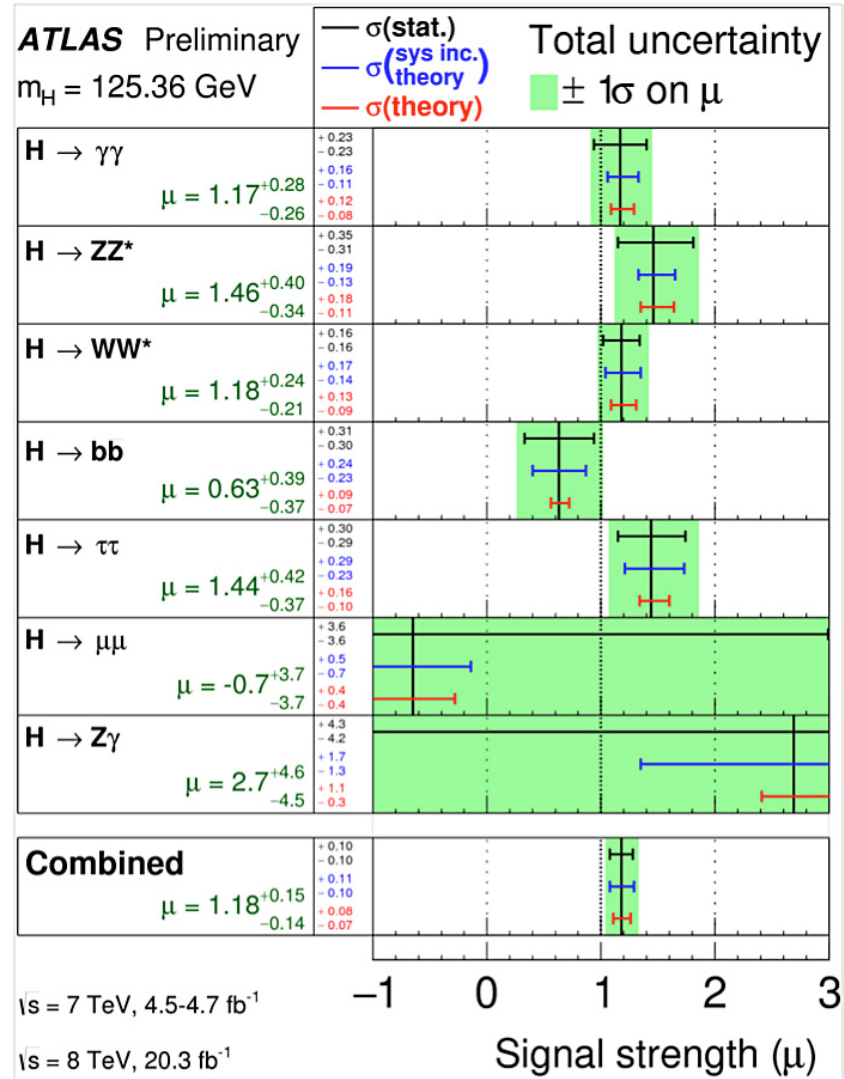
PDG 2015:  
Z-pole

Quantity	Value	Standard Model	Pull
$M_Z$ [GeV]	$91.1876 \pm 0.0021$	$91.1880 \pm 0.0020$	-0.2
$\Gamma_Z$ [GeV]	$2.4952 \pm 0.0023$	$2.4955 \pm 0.0009$	-0.1
$\Gamma(\text{had})$ [GeV]	$1.7444 \pm 0.0020$	$1.7420 \pm 0.0008$	—
$\Gamma(\text{inv})$ [MeV]	$499.0 \pm 1.5$	$501.66 \pm 0.05$	—
$\Gamma(\ell^+\ell^-)$ [MeV]	$83.984 \pm 0.086$	$83.995 \pm 0.010$	—
$\sigma_{\text{had}}$ [nb]	$41.541 \pm 0.037$	$41.479 \pm 0.008$	1.7
$R_e$	$20.804 \pm 0.050$	$20.740 \pm 0.010$	1.3
$R_\mu$	$20.785 \pm 0.033$	$20.740 \pm 0.010$	1.4
$R_\tau$	$20.764 \pm 0.045$	$20.785 \pm 0.010$	-0.5
$R_b$	$0.21629 \pm 0.00066$	$0.21576 \pm 0.00003$	0.8
$R_c$	$0.1721 \pm 0.0030$	$0.17226 \pm 0.00003$	-0.1
$A_{FB}^{(0,e)}$	$0.0145 \pm 0.0025$	$0.01616 \pm 0.00008$	-0.7
$A_{FB}^{(0,\mu)}$	$0.0169 \pm 0.0013$		0.6
$A_{FB}^{(0,\tau)}$	$0.0188 \pm 0.0017$		1.6
$A_{FB}^{(0,b)}$	$0.0992 \pm 0.0016$	$0.1029 \pm 0.0003$	-2.3
$A_{FB}^{(0,c)}$	$0.0707 \pm 0.0035$	$0.0735 \pm 0.0002$	-0.8
$A_{FB}^{(0,s)}$	$0.0976 \pm 0.0114$	$0.1030 \pm 0.0003$	-0.5
$\bar{s}_\ell^2$	$0.2324 \pm 0.0012$	$0.23155 \pm 0.00005$	0.7
	$0.23176 \pm 0.00060$		0.3
	$0.2297 \pm 0.0010$		-1.9
$A_e$	$0.15138 \pm 0.00216$	$0.1468 \pm 0.0004$	2.1
	$0.1544 \pm 0.0060$		1.3
	$0.1498 \pm 0.0049$		0.6
$A_\mu$	$0.142 \pm 0.015$		-0.3
$A_\tau$	$0.136 \pm 0.015$		-0.7
	$0.1439 \pm 0.0043$		-0.7
$A_b$	$0.923 \pm 0.020$	0.9347	-0.6
$A_c$	$0.670 \pm 0.027$	$0.6676 \pm 0.0002$	0.1
$A_s$	$0.895 \pm 0.091$	0.9356	-0.4

PDG 2015:  
non-Z-pole

Quantity	Value	Standard Model	Pull
$m_t$ [GeV]	$173.24 \pm 0.95$	$173.87 \pm 0.87$	-0.7
$M_W$ [GeV]	$80.387 \pm 0.016$	$80.363 \pm 0.006$	1.5
	$80.376 \pm 0.033$		0.4
$\Gamma_W$ [GeV]	$2.046 \pm 0.049$	$2.090 \pm 0.001$	-0.9
	$2.196 \pm 0.083$		1.3
$M_H$ [GeV]	$125.6 \pm 0.4$	$125.5 \pm 0.4$	0.1
$\rho_{\gamma W}$	$0.45 \pm 0.31$	$0.01 \pm 0.03$	1.4
	$0.12 \pm 0.43$	$0.00 \pm 0.03$	0.3
$\rho_{\gamma Z}$	$0.08 \pm 0.28$	$0.01 \pm 0.04$	0.2
$\rho_{ZW}$	$0.30 \pm 0.39$	$0.00 \pm 0.01$	0.8
$g_V^{\nu e}$	$-0.040 \pm 0.015$	$-0.0397 \pm 0.0001$	0.0
$g_A^{\nu e}$	$-0.507 \pm 0.014$	$-0.5064$	0.0
$Q_W(e)$	$-0.0403 \pm 0.0053$	$-0.0473 \pm 0.0003$	1.3
$Q_W(p)$	$0.064 \pm 0.012$	$0.0708 \pm 0.0003$	-0.6
$Q_W(\text{Cs})$	$-72.62 \pm 0.43$	$-73.25 \pm 0.01$	1.5
$Q_W(\text{Tl})$	$-116.4 \pm 3.6$	$-116.90 \pm 0.02$	0.1
$\hat{s}_Z^2(\text{eDIS})$	$0.2299 \pm 0.0043$	$0.23126 \pm 0.00005$	-0.3
$\tau_\tau$ [fs]	$291.13 \pm 0.43$	$291.19 \pm 2.41$	0.0
$\frac{1}{2}(g_\mu - 2 - \frac{\alpha}{\pi})$	$(4511.07 \pm 0.79) \times 10^{-9}$	$(4508.68 \pm 0.08) \times 10^{-9}$	3.0

# ATLAS Run1: Signal strength



the data :  
(@ 95%)

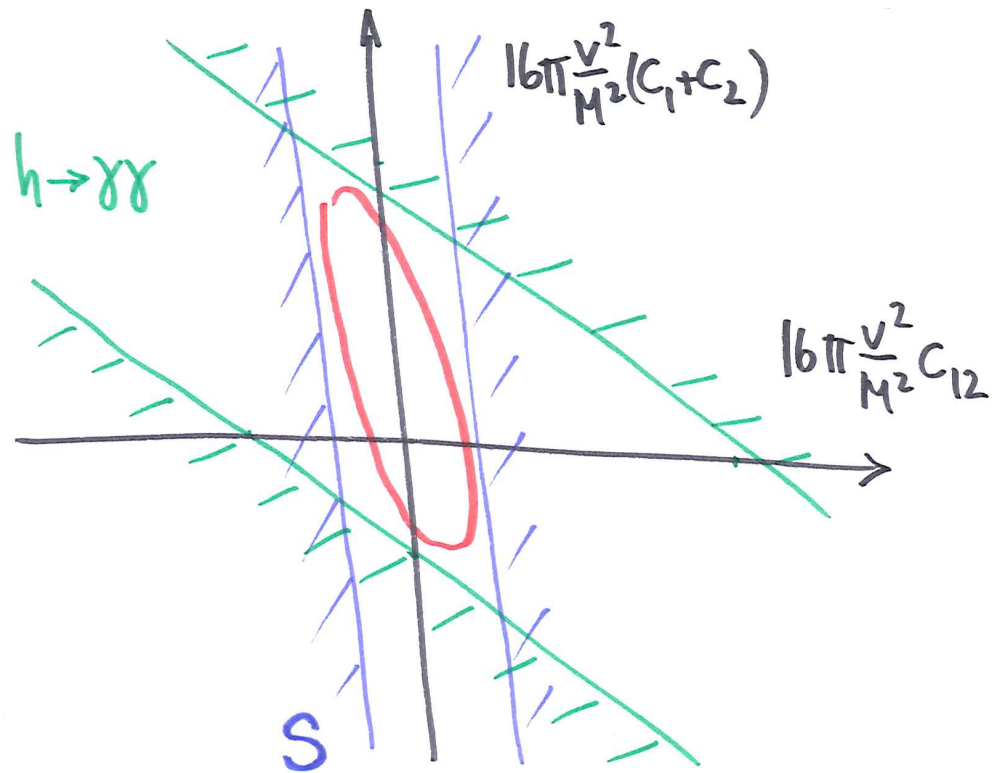
• Precision electroweak  $S = -0.03 \pm 0.10$  PDG

•  $h \rightarrow \gamma\gamma$

$$C_{\gamma\gamma} = 0.014 \pm 0.058$$

( $h \rightarrow \gamma Z$   $C_{\gamma Z} < 0.2$ )

Falkowski et al hep-ph/  
1505.00046



Bound on  $M$ ?

$$C_i = 1 \Rightarrow M \gtrsim 6 \text{ TeV}$$

$$C_i = \frac{1}{16\pi^2} \Rightarrow M \gtrsim 500 \text{ GeV}$$

We are pleased to announce the **fourth Higgs Effective Field Theories** workshop (**HEFT2016**) will take place at the Niels Bohr Institute at the University of Copenhagen from Wednesday 26th to Friday 28th October 2016. Pervious installments of the workshop were held at CERN (2013), Madrid (2014) and Chicago in 2015.