# The quest for leptonic CP violation

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P. Huber – VT-CNP – p. 1

# Status quo

A common framework for all the neutrino data is oscillation of three active neutrinos

- $\Delta m_{21}^2 \sim 8 \cdot 10^{-5} \,\mathrm{eV}^2$  and  $\theta_{12} \sim 1/2$
- $\Delta m^2_{31} \sim 2 \cdot 10^{-3} \,\mathrm{eV}^2$  and  $\theta_{23} \sim \pi/4$
- $\theta_{13} \sim 0.16$

This implies a lower bound on the mass of the heaviest neutrino

$$\sqrt{2 \cdot 10^{-3} \,\mathrm{eV}^2} \sim 0.04 \,\mathrm{eV}$$

but we currently do not know which neutrino is the heaviest.

# **Mixing matrices**

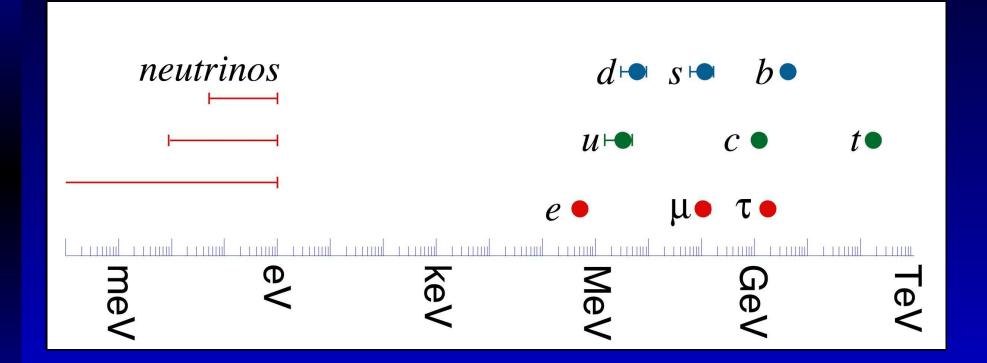
#### Quarks

$$|U_{CKM}| = \begin{pmatrix} 1 & 0.2 & 0.005 \\ 0.2 & 1 & 0.04 \\ 0.005 & 0.04 & 1 \end{pmatrix}$$

Neutrinos

$$|U_{\nu}| = \begin{pmatrix} 0.8 & 0.5 & 0.15 \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix}$$

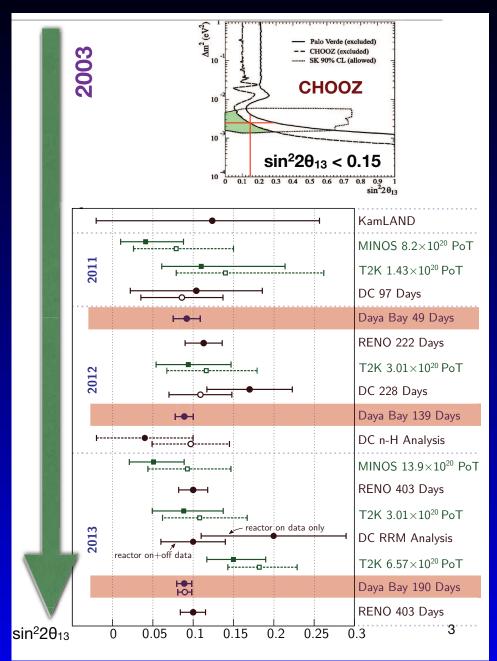
## Fermion masses



# $\theta_{13}$ is large!

- Many results from reactor and beam experiments
- Some single results exceed  $5\sigma$  significance
- All results agree well
- Current Daya Bay result

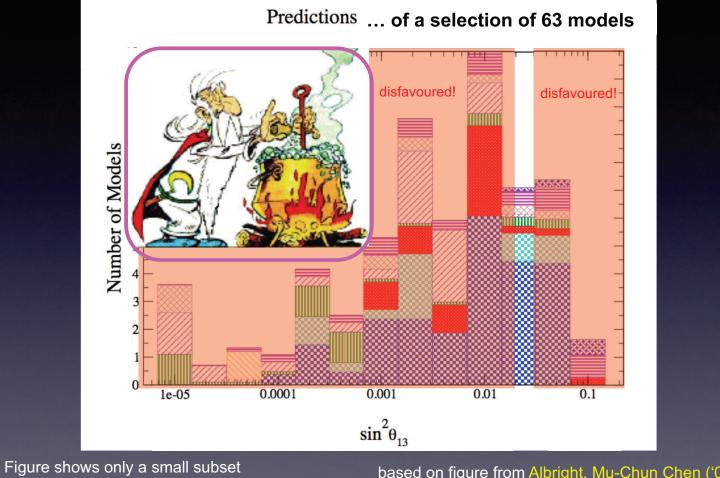
 $\sin^2 2\theta_{13} = 0.084 \pm 0.005$ 



Zhang, Neutrino 2014

# **Model selection**

#### ... a large fraction has been excluded!



of the existing models ... !

based on figure from Albright, Mu-Chun Chen ('06)

Antusch, 2012

# **Measuring leptonic CPV**

In order to measure CP violation we need to reconstruct one out of these

$$P(\nu_{\mu} \to \nu_{e}) \text{ or } P(\nu_{e} \to \nu_{\mu})$$

and one out of these

$$P(\bar{\nu}_{\mu} \to \bar{\nu}_{e}) \text{ or } P(\bar{\nu}_{e} \to \bar{\nu}_{\mu})$$

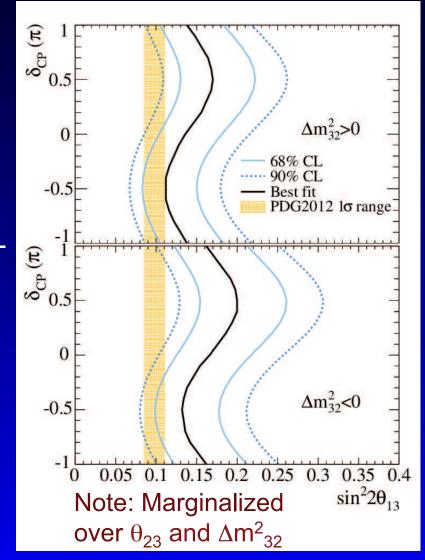
and we'd like to do that at percent level accuracy. Note,

$$\frac{\bar{P} - P}{\bar{P} + P} \propto \frac{1}{\sin \theta_{13}}$$

# **First hints for CP violation?**

Latest T2K results combined with  $\theta_{13}$  constraint from Daya Bay

Hint for  $\delta = -\pi/2$ ?



Walters, Neutrino 2014 P. Huber – VT-CNP – p. 8

#### **Neutrinos are massive – so what?**

Neutrinos in the Standard Model (SM) are strictly massless, therefore the discovery of neutrino oscillation, which implies non-zero neutrino masses requires the addition of new degrees of freedom.

#### We always knew they are ...

The SM, likely, is an effective field theory, *i.e.* at some high scale  $\Lambda$  new degrees of freedom will appear

$$\mathcal{L}_{SM} + rac{1}{\Lambda}\mathcal{L}_5 + rac{1}{\Lambda^2}\mathcal{L}_6 + \dots$$

The first operators sensitive to new physics have dimension 5. It turns out there is only one dimension 5 operator

$$\mathcal{L}_5 = \frac{1}{\Lambda} (LH)(LH) \rightarrow \frac{1}{\Lambda} (L\langle H \rangle)(L\langle H \rangle) = m_{\nu} \nu \nu$$
  
Thus studying neutrino masses is, in principle, the nost sensitive probe for new physics at high scales  
Weinberg

## **Effective theories**

The problem in effective theories is, that there are *a priori* unknown pre-factors for each operator

$$\mathcal{L}_{SM} + \frac{\#}{\Lambda}\mathcal{L}_5 + \frac{\#}{\Lambda^2}\mathcal{L}_6 + \dots$$

Typically, one has  $\# = \mathcal{O}(1)$ , but there may be reasons for this being wrong

- lepton number may be conserved  $\rightarrow$  no Majorana mass term
- lepton number may be approximately conserved  $\rightarrow$  small pre-factor for  $\mathcal{L}_5$

Therefore, we do not know the scale of new physics responsible for neutrino masses – anywhere from keV to the Planck scale is possible.

# Neutrino masses are different

The crucial difference between neutrinos and other fermions is the possibility of a Majorana mass term

 $m_L \bar{\psi}_L \psi_R^C + m_R \bar{\psi}_R \psi_L^C$ 

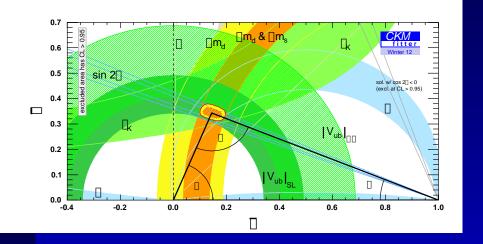
on top of the usual Dirac mass term

 $m_D \bar{\psi}_L \psi_R$ 

This allows for things like the seesaw mechanism (many versions) and implies that the neutrino flavor sector probes very different physics than the quark sector.

# What did we learn from that?

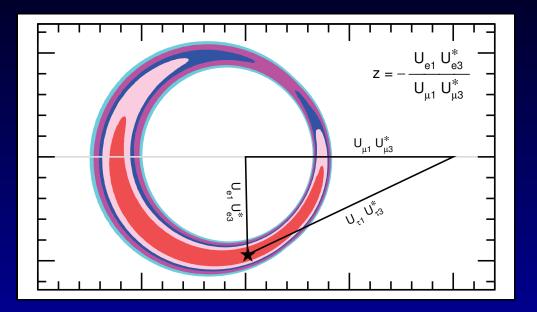
Our expectations where to find BSM physics are driven by models – but we should not confuse the number of models with the likelihood for discovery.



- CKM describes all flavor effects
- SM baryogenesis difficult
- New Physics at a TeV
  - has a special flavor structure
  - or does not exist...

and a vast number of parameter and model space excluded. Neutrinos are very different from quarks, therefore precision measurements will yield very different answers  $\Rightarrow$  complementary to collider searches

# **Unitarity triangles**



#### 0.7 0.95 CKM fitter ٦, 0.6 0.5 sin 2∏ sol. w/ cos 2[] < 0 (excl. at CL > 0.95) 0.4 0.3 0.2 0.1 V<sub>ub</sub> 0.0 ⊾ -0.4 0.2 -0.2 0.0 0.4 0.6 0.8 1.0

Neutrino sector Gonzalez-Garcia, Maltoni, Schwetz, 2014

Quark sector

# **CP** violation

There are only very few parameters in the  $\nu$ SM which can violate CP

- CKM phase measured to be  $\gamma \simeq 70^\circ$
- $\theta$  of the QCD vacuum measured to be  $< 10^{-10}$
- Dirac phase of neutrino mixing
- Possibly: 2 Majorana phases of neutrinos

At the same time we know that the CKM phase is not responsible for the Baryon Asymmetry of the Universe...

#### **Flavor models**

Simplest un-model – anarchy Murayama, Naba, DeGouvea

$$dU = ds_{12}^2 \, dc_{13}^4 \, ds_{23}^2 \, d\delta_{CP} \, d\chi_1 \, d\chi_2$$

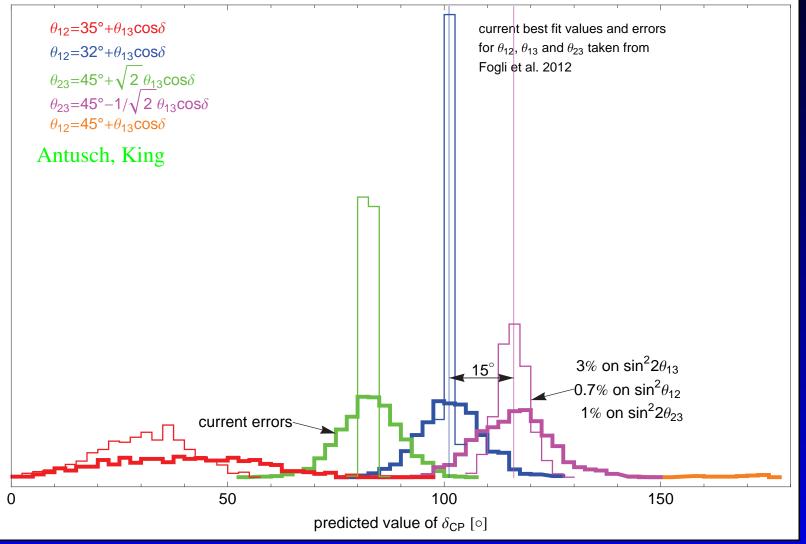
predicts flat distribution in  $\delta_{CP}$ 

Simplest model – Tri-bimaximal mixing Harrison, Perkins, Scott

$$\begin{pmatrix} \sqrt{\frac{1}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

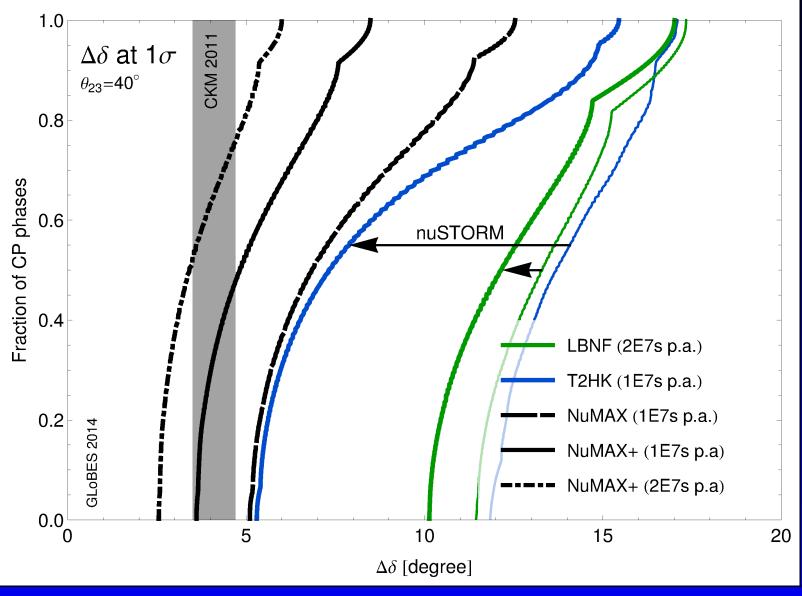
to still fit data, obviously corrections are needed – predictivity?

# **Sum rules**



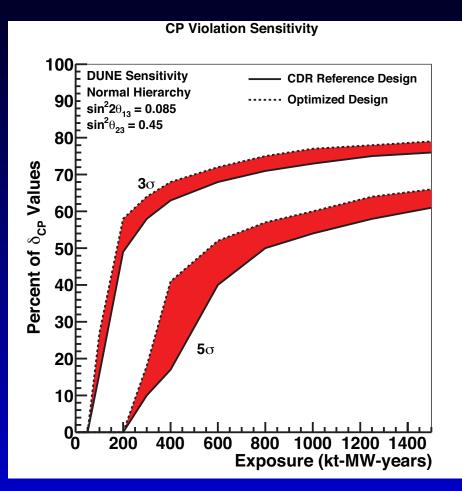
 $3\sigma$  resolution of 15° distance requires 5° error. NB – smaller error on  $\theta_{12}$  requires dedicated experiment like JUNO

## Is 5° feasible?



PH, Bross, Palmer

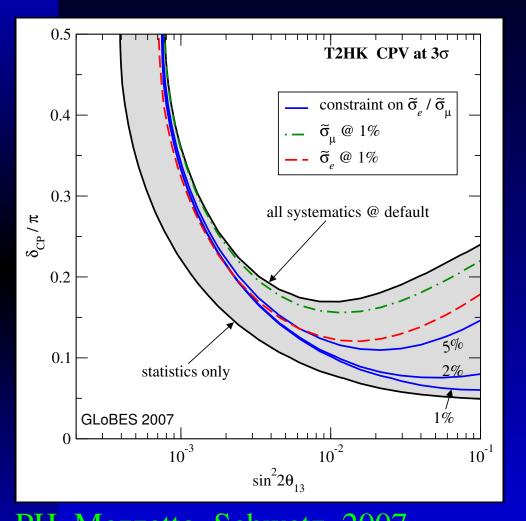
### DUNE



One year at 1.2 MW and 40 kt corresponds to 48 kt MW y Beam upgrade to 2.3 MW foreseen At 1000 MW kt y reaches 8-12° CP phase accuracy

#### Scheduled start of data taking 2026

# **Neutrino cross sections**

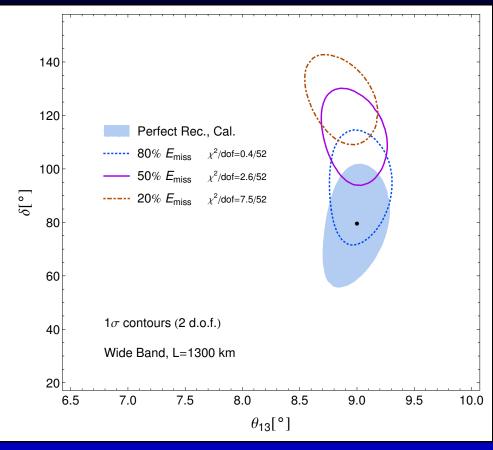


Using current cross section uncertainties and a perfect near detector.

Appearance experiments using a (nearly) flavor pure beam can **not** rely on a near detector to predict the signal at the far site!

PH, Mezzetto, Schwetz, 2007 Differences between  $\nu_e$  and  $\nu_{\mu}$  are significant below 1 GeV, see e.g. Day, McFarland, 2012

# Nuclear effects – example



In elastic scattering a certain number of neutrons is made

Neutrons will be largely invisible even in a liquid argon TPC ⇒ missing energy

Ankowski *et al.*, in preparation We can correct for the missing energy **IF** we know the mean neutron number and energy made in the event...

## **Theory and cross sections**

Theory is cheap, but multi-nucleon systems and their dynamic response are a hard problem. Currently, there are two major approaches

Greens function Monte Carlo: numerically "exact" solutions for light nuclei (A $\leq$ 12) and non-relativistic kinematics.

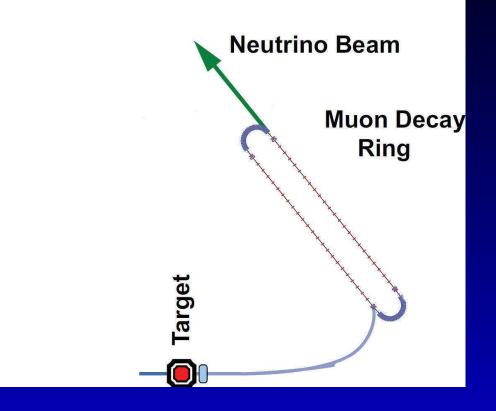
Spectral functions: use information on the initial state from electron-scattering data.

Both techniques are not controlled approximations and thus to trust theory at x% we have to experimentally test the theory at x% – ultimately, precision cross section measurements are unavoidable.

# **Towards precise cross sections**

#### Needs better neutrino sources

- Sub-percent beam flux normalization
- Very high statistics needed to map phase space
- Neutrinos and antineutrinos
- $\nu_{\mu}$  and  $\nu_{e}$



One (the only?) source which can deliver all that is a muon storage ring, aka nuSTORM.

# Summary

- Neutrino oscillation is solid evidence for new physics
- Current data allows  $\mathcal{O}(1)$  corrections to three flavor framework
- Precision measurements have the best potential to uncover even "newer" physics
- Sterile neutrinos?

Neutrinos have provided us with many surprises and neutrinos are still largely unexplored !