Heavy Sterile Phenomonology and the MiniBooNE anomaly

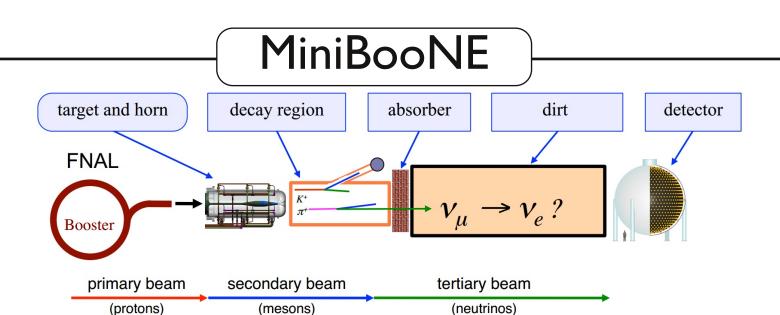
Mark Ross-Lonergan IPPP, Durham University Based on work with Silvia Pascoli and Peter Ballett

Invisibles | 5 **IFT & Thyssen Museum Madrid** 24th June 2015



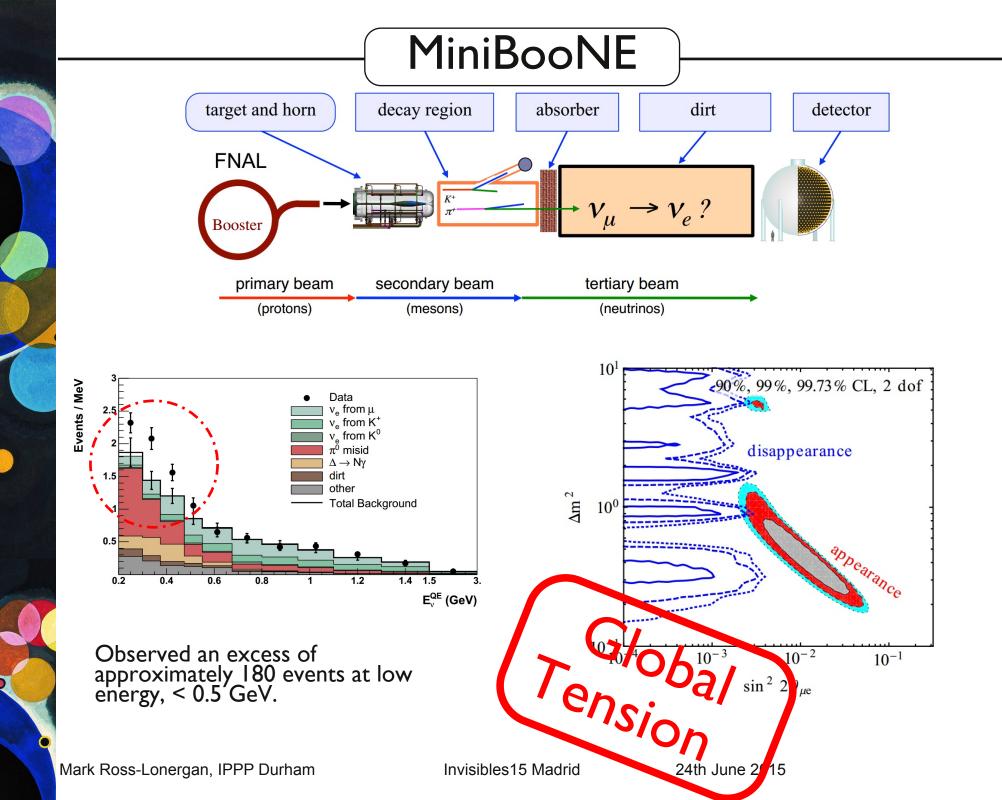


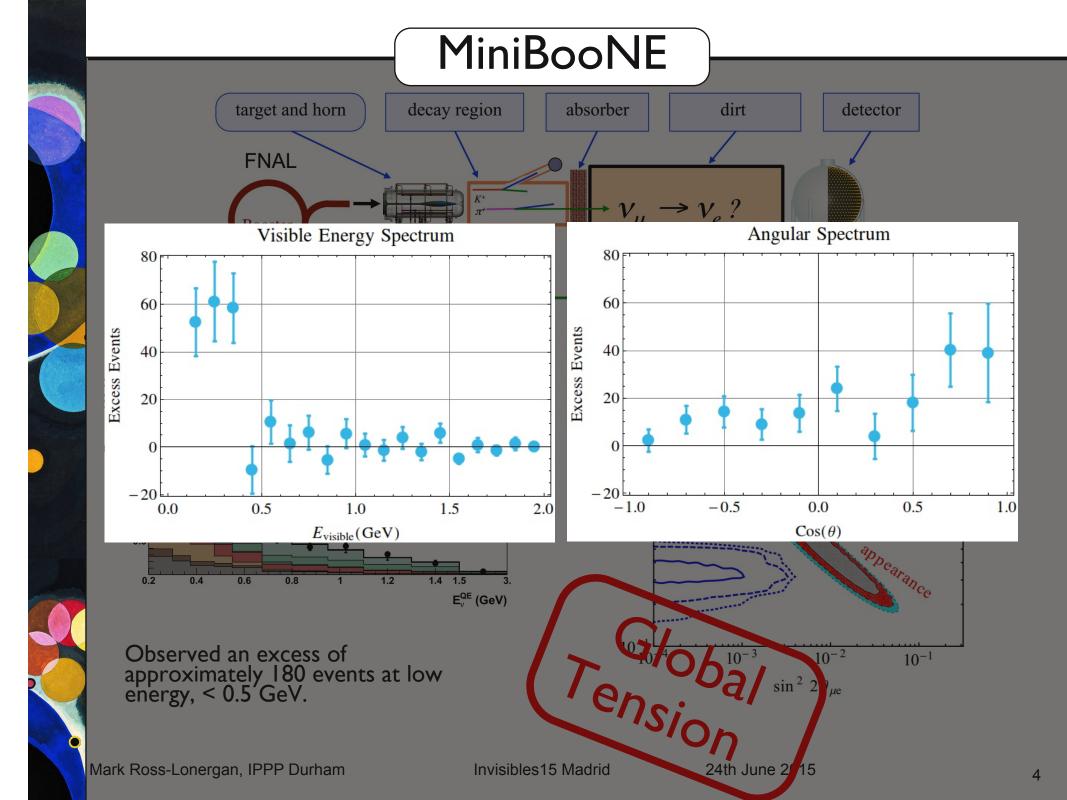




Events / MeV Data v_{o} from μ v_{e} from K⁺ v_{0}^{v} from K^{0} π^{0} misid $\Delta \rightarrow N\gamma$ dirt other **Total Background** 0.5 0.2 0.4 0.8 0.6 1.2 1.4 1.5 1 3. E^{QE}_v (GeV)

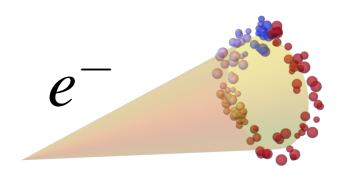
Observed an excess of approximately 180 events at low energy, < 0.5 GeV.



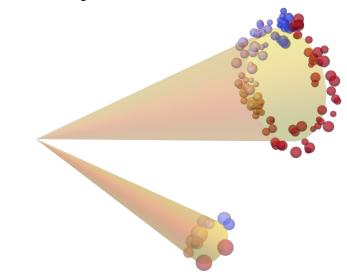


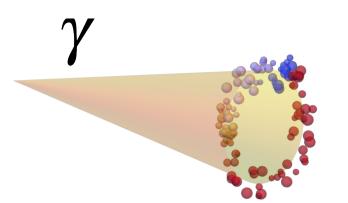
What's the Signal?

Overlapping e^+e^-



Highly Asymmetric e^+e^-

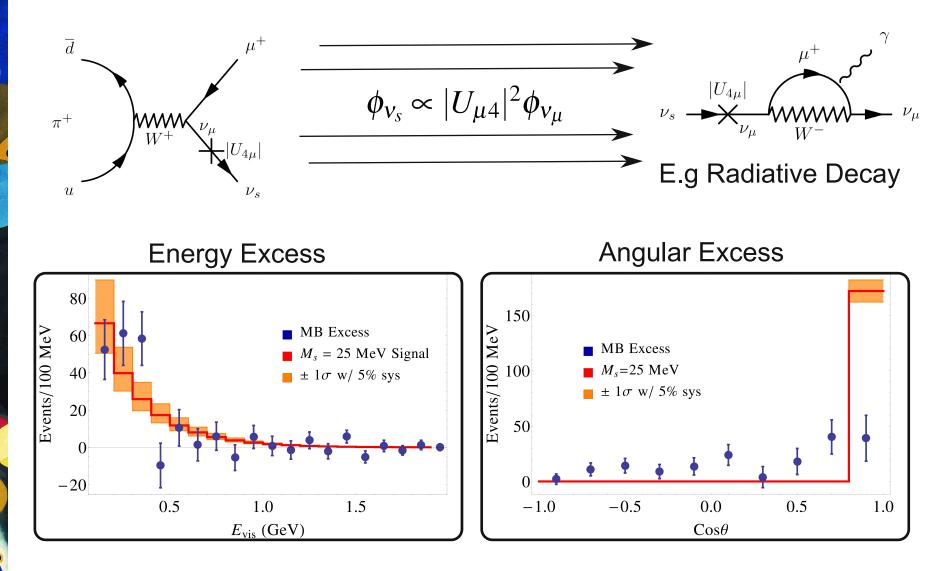




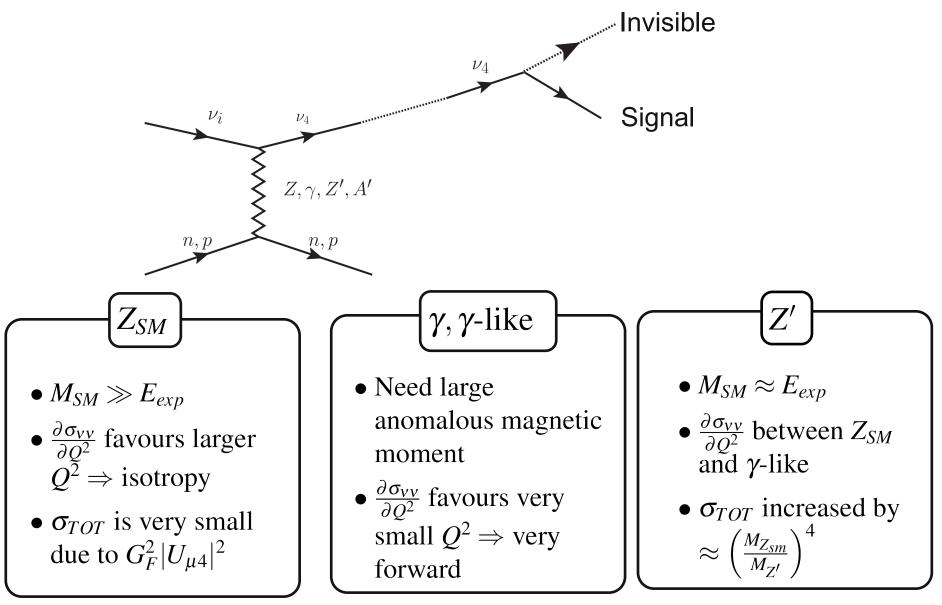
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Alternative Scenarios

Sterile Neutrinos in the MeV-GeV Range have a very rich phenomonology in SBL facilities. A flux of sterile neutrinos created alongside the standard neutrino beam can reach the detector and subsequently decay.



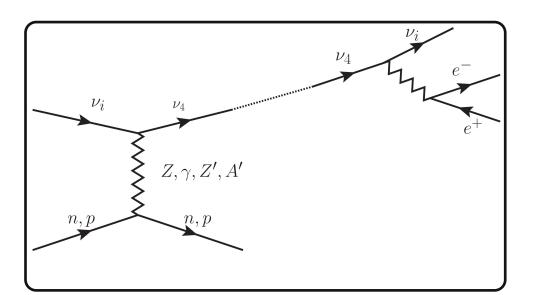
Angular spectrum is *key* tool in excluding large classes of sterile scenarios. Need to produce more isotropic daughter events from sterile decay.



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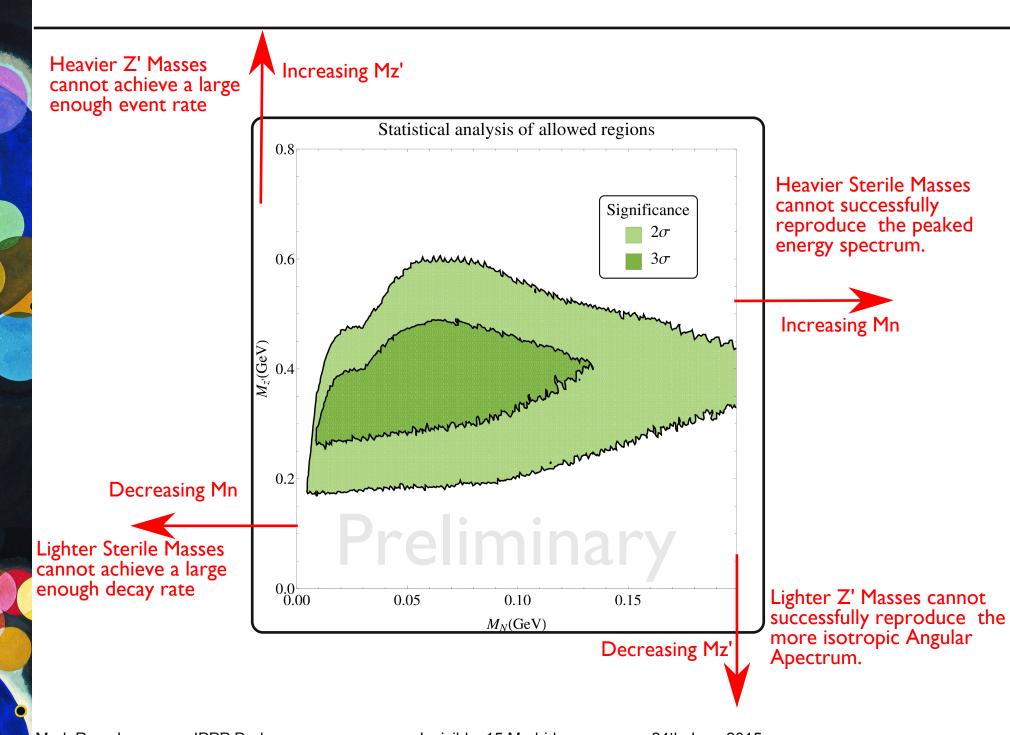
• As a concrete model we intoduce a single MeV scale sterile degree of freedom charged under a new U(1)' symmetry. Although no standard model particles initially feel this new force, they obtain a microcharge due to kinetic mixing with the SM hypercharge boson. In the flavour basis

$$\mathscr{L} = L_{sm} - \frac{1}{4} F_{\mu\nu}^{\prime 2} - \chi F_{\mu\nu}^{\prime} B^{\mu\nu} + \frac{M_{Z^{\prime}}}{2} X^{\prime 2} - ig^{\prime} \overline{\nu_s} \gamma^{\mu} \nu_s X_{\mu}^{\prime},$$



- Post EWSB there will exist a tree level coupling between Massive Z', v₄ and v_{1,2,3}.
- Fraction of subsequent decays to e^+e^- mis-identified as CCQE electrons
- Kinetic mixing χ^2 alongside $|U_{e4}|^2, |U_{\mu4}|^2$ and $|U_{\tau4}|^2$ constrained to be below current 90% C.L bounds.

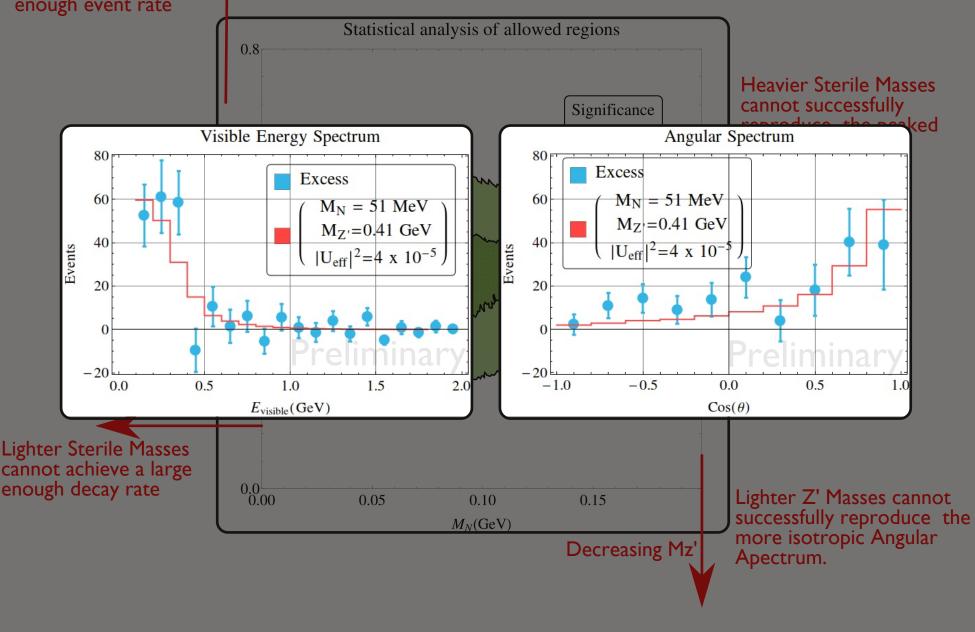
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Sample Points

Increasing Mz'

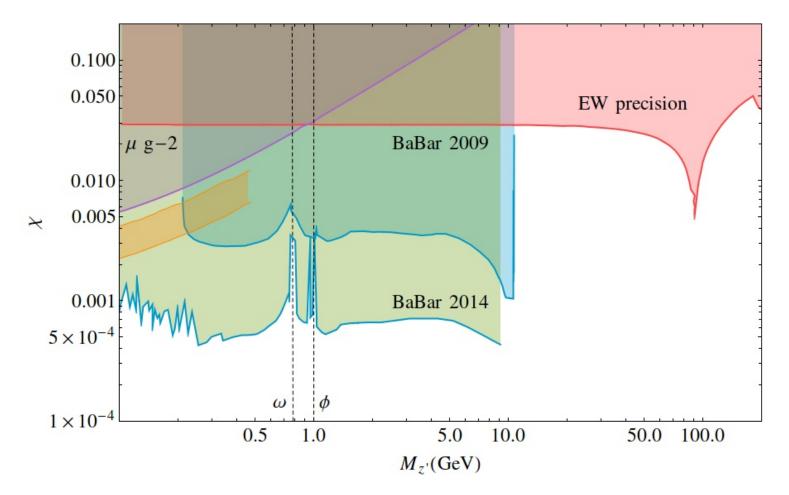
Heavier Z' Masses cannot achieve a large enough event rate

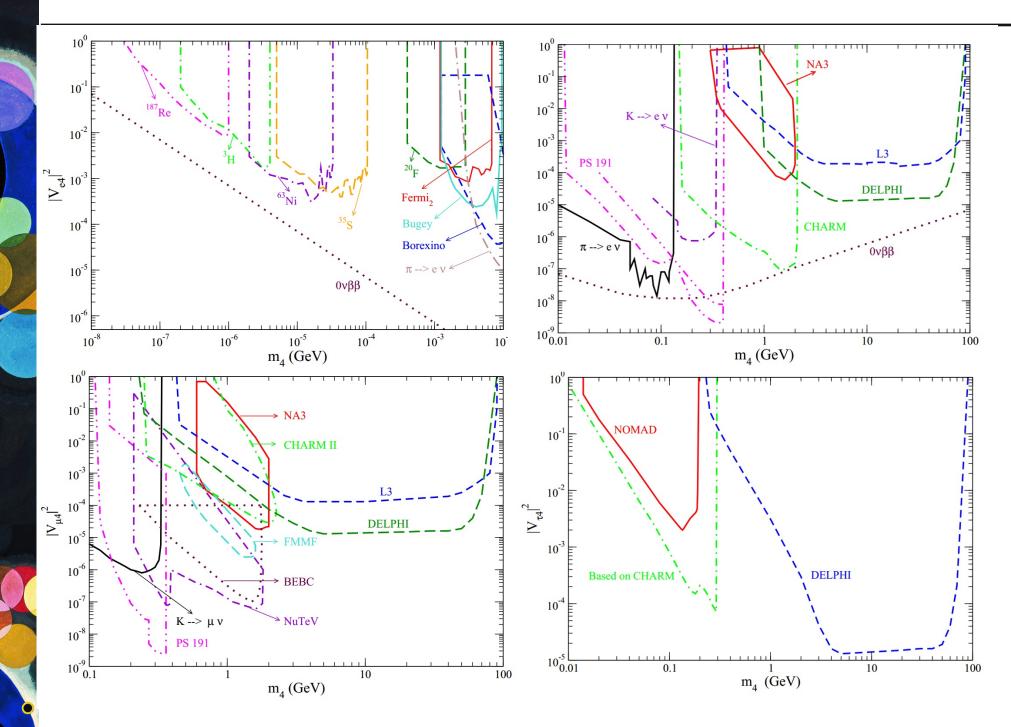


Thank You!

Mark Ross-Lonergan, IPPP Durham

U(1)' Kinetic Mixing Bounds





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