

Constraining discrete flavour symmetries with neutrino oscillation experiments

Peter Ballett
IPPP, Durham University

Invisibles 15 Workshop – *IFT/Thyssen-Bornemisza Museum, Madrid*

23rd June 2015



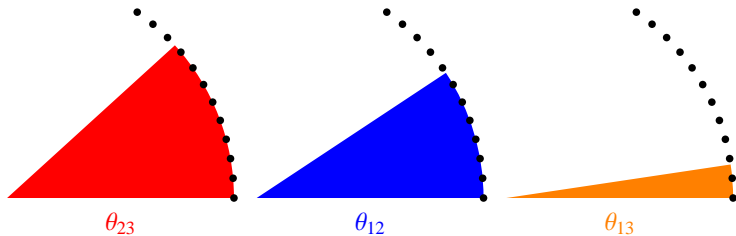
Outline of talk

- 1 What do we know about leptonic flavour?
- 2 Parameter correlations from discrete symmetries
 - Semi-direct A_5 with generalised CP
- 3 Conclusions

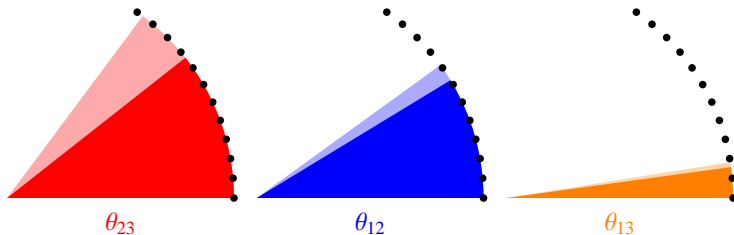
Outline of talk

- 1 What do we know about leptonic flavour?
- 2 Parameter correlations from discrete symmetries
 - Semi-direct A_5 with generalised CP
- 3 Conclusions

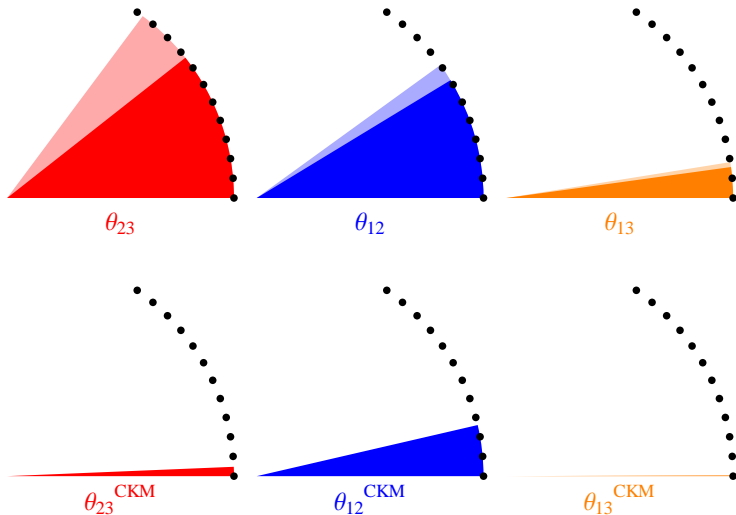
Leptonic flavour



Leptonic flavour: current knowledge



Leptonic flavour: some explanation?



Outline of talk

- 1 What do we know about leptonic flavour?
- 2 **Parameter correlations from discrete symmetries**
 - Semi-direct A_5 with generalised CP
- 3 Conclusions

Leptonic flavour symmetries

- Extend the SM symmetry group by a discrete non-Abelian factor. Assign fields to representations of this group.

$$G_{\text{SM}} \rightarrow G_{\text{SM}} \times G_{\text{F}} \quad \text{and} \quad \begin{pmatrix} L_e \\ L_\mu \\ L_\tau \end{pmatrix} \sim \underline{\mathbf{3}}$$

- The symmetry restricts the most general mass terms of the theory:

$$T^\dagger \left(m_e^\dagger m_e \right) T = m_e^\dagger m_e \quad \text{and} \quad S^T m_\nu S = m_\nu.$$

- The full symmetry cannot be exact at low-energies whilst maintaining distinct neutrino masses. However, a set of **residual symmetries** may enforce a given mixing pattern or, most generally, specific correlations between observables.

Parameter correlations

In some cases, the symmetry alone does not predict the full PMNS matrix. Instead only predicting a constrained form.

For example...

If $G_\nu = \mathbb{Z}_2$ we see that only a single column of the PMNS matrix can be specified:

$$\text{e.g.} \quad U_{\text{PMNS}} = \begin{pmatrix} \sqrt{\frac{2}{3}} & ? & ? \\ \frac{1}{\sqrt{6}} & ? & ? \\ \frac{1}{\sqrt{6}} & ? & ? \end{pmatrix}$$

which leads to **correlations** between mixing parameters:

$$s = \sqrt{1 - \frac{2r^2}{2 - r^2}} \quad \text{and} \quad a \approx r \cos \delta,$$

s.t. $\sin \theta_{23} \equiv \frac{1+a}{\sqrt{2}}$, $\sin \theta_{12} = \frac{1+s}{\sqrt{3}}$, and $\sin \theta_{13} \equiv \frac{r}{\sqrt{2}}$, (see King 0710.0530).

Outline of talk

- 1 What do we know about leptonic flavour?
- 2 **Parameter correlations from discrete symmetries**
 - Semi-direct A_5 with generalised CP
- 3 Conclusions

Generalised CP symmetries

- With three flavours, a **generalised CP symmetry** combines a flavour transformation with a CP transformation:

$$\Psi_\alpha \rightarrow X_{\alpha\beta} \Psi_\beta^{\text{CP}}$$

- The theoretical consistency of this picture is constrained and the form of X must be related to an *automorphism of the flavour group*. See [Holthausen, Lindner & Schmidt 1211.6953](#), [Chen et al. 1402.0507](#).
- This creates novel constraints on the mass matrix, specifically regarding its complex phases,

$$X m_\nu X = m_\nu^*$$

- Taking this symmetry to be a residual symmetry of the lepton mass matrix can lead to predictions of the PMNS matrix including complex phases: δ , α_{21} and α_{31} . See e.g. [Feruglio, Hagedorn & Ziegler 1211.5560](#).

- In [1503.07543](#), we[†] considered the group A₅ and systematically computed all of the predictions from different combinations of residual symmetries with GCP.

$$A_5 = \langle S, T \mid S^2 = T^5 = (ST)^3 \rangle$$

such that,

$$S = \begin{pmatrix} -1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 1 \end{pmatrix} \quad \text{and} \quad T = \frac{1}{2} \begin{pmatrix} 1 & -\varphi & -\varphi_g \\ \varphi & -\varphi_g & -1 \\ -\varphi_g & 1 & \varphi \end{pmatrix}$$

- We focus on the case when $G_\nu = \mathbb{Z}_2$, which produces parameter correlations similar to the [atmospheric sum rules](#) mentioned previously; however, the phases are more constrained.

[†] [PB, Pascoli & Turner](#).

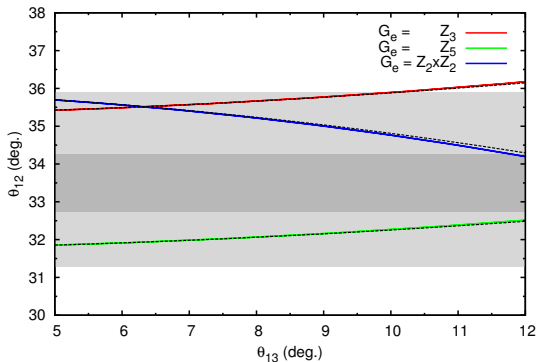
A_5 with GCP: predictions

- Taking $G = A_5$ with GCP, breaking into $G_\nu = \mathbb{Z}_2$ with GCP and a cyclic group for G_e , we find 10 distinct sets of observables.

| G_e | θ_{12} | θ_{23} | $\sin \alpha_{ji}$ | δ |
|------------------------------------|---------------------------------|-------------------------------|--------------------|-------------|
| \mathbb{Z}_3 | $35.27^\circ + 10.13^\circ r^2$ | 45° | 0 | 90° |
| | | | | 270° |
| \mathbb{Z}_5 | $31.72^\circ + 8.85^\circ r^2$ | $45^\circ \pm 25.04^\circ r$ | 0 | 0° |
| | | | | 180° |
| | | 45° | 0 | 90° |
| | | | | 270° |
| $\mathbb{Z}_2 \times \mathbb{Z}_2$ | $36.00^\circ - 34.78^\circ r^2$ | $31.72^\circ + 55.76^\circ r$ | 0 | 0° |
| | | | | 180° |
| | | $58.28^\circ - 55.76^\circ r$ | 0 | 0° |
| | | | | 180° |

From PB, Pascoli & Turner [1503.07543](#). See also Li & Ding [1503.03711](#) and De Iura, Hagedorn & Meloni [1503.04140](#).

- Three distinct predictions for θ_{12} in terms of θ_{13} .



- Change in θ_{12} over 3σ range of r only 0.07° , 0.08° and 0.13° for \mathbb{Z}_3 , \mathbb{Z}_5 and $\mathbb{Z}_2 \times \mathbb{Z}_2$. Testing these relationships in themselves will be beyond the capabilities of upcoming reactor experiments.

A_5 with GCP: θ_{23} and δ

- We find an interesting relationship between θ_{23} and δ .

| G_e | θ_{12} | θ_{23} | $\sin \alpha_{jj}$ | δ |
|------------------------------------|---------------------------------|-------------------------------|--------------------|-------------|
| \mathbb{Z}_3 | $35.27^\circ + 10.13^\circ r^2$ | 45° | 0 | 90° |
| | | | | 270° |
| \mathbb{Z}_5 | $31.72^\circ + 8.85^\circ r^2$ | $45^\circ \pm 25.04^\circ r$ | 0 | 0° |
| | | | | 180° |
| | | 45° | 0 | 90° |
| | | | | 270° |
| $\mathbb{Z}_2 \times \mathbb{Z}_2$ | $36.00^\circ - 34.78^\circ r^2$ | $31.72^\circ + 55.76^\circ r$ | 0 | 0° |
| | | | | 180° |
| | | $58.28^\circ - 55.76^\circ r$ | 0 | 0° |
| | | | | 180° |

A_5 with GCP: θ_{23} and δ

- We find an interesting relationship between θ_{23} and δ .

| G_e | θ_{12} | θ_{23} | $\sin \alpha_{jj}$ | δ |
|------------------------------------|---------------------------------|-------------------------------|--------------------|-------------|
| \mathbb{Z}_3 | $35.27^\circ + 10.13^\circ r^2$ | 45° | 0 | 90° |
| | | | | 270° |
| \mathbb{Z}_5 | $31.72^\circ + 8.85^\circ r^2$ | $45^\circ \pm 25.04^\circ r$ | 0 | 0° |
| | | | | 180° |
| | | 45° | 0 | 90° |
| | | | | 270° |
| $\mathbb{Z}_2 \times \mathbb{Z}_2$ | $36.00^\circ - 34.78^\circ r^2$ | $31.72^\circ + 55.76^\circ r$ | 0 | 0° |
| | | | | 180° |
| | | $58.28^\circ - 55.76^\circ r$ | 0 | 0° |
| | | | | 180° |

A_5 with GCP: θ_{23} and δ

- Taking an order-5 generator in the charged-lepton sector, we find two choices depending on X :

$$\left(\begin{array}{ccc} \frac{\varphi}{\sqrt{2+\varphi}} & \frac{1}{\sqrt{2+\varphi}} & 0 \\ -\frac{1}{\sqrt{4+2\varphi}} & \frac{\varphi}{\sqrt{4+2\varphi}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{4+2\varphi}} & -\frac{\varphi}{\sqrt{4+2\varphi}} & \frac{1}{\sqrt{2}} \end{array} \right) R_{13}(\theta) \quad \left| \quad \left(\begin{array}{ccc} \frac{\varphi}{\sqrt{2+\varphi}} & -\frac{i}{\sqrt{2+\varphi}} & 0 \\ -\frac{i}{\sqrt{4+2\varphi}} & \frac{\varphi}{\sqrt{4+2\varphi}} & \frac{1}{\sqrt{2}} \\ \frac{i}{\sqrt{4+2\varphi}} & -\frac{\varphi}{\sqrt{4+2\varphi}} & \frac{1}{\sqrt{2}} \end{array} \right) R_{13}(\theta)$$

$$\sin^2 \theta_{23} = \frac{1}{2} \frac{(\sin \theta + \sqrt{1+\varphi^2} \cos \theta)^2}{1+\varphi^2 \cos^2 \theta}$$

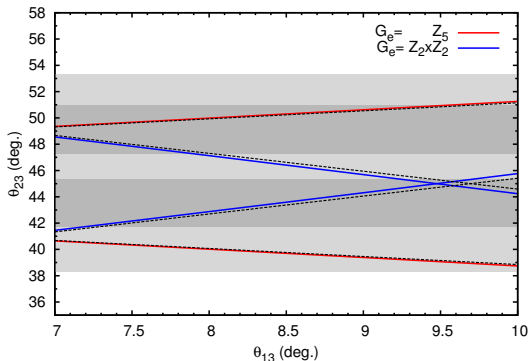
$$\delta = \begin{cases} 0 & \theta \in (0, \frac{\pi}{2}) \\ \pi & \theta \in (\frac{\pi}{2}, \pi) \end{cases}$$

$$\sin^2 \theta_{23} = \frac{1}{2} \quad \forall \theta$$

$$\delta = \begin{cases} \frac{3\pi}{2} & \theta \in (0, \frac{\pi}{2}) \\ \frac{\pi}{2} & \theta \in (\frac{\pi}{2}, \pi) \end{cases}$$

Non-maximal θ_{23} & Dirac CP conservation

- If δ is unobserved at T2K and $\text{NO}\nu\text{A}$, then we can look for information in the (possible) deviation from $\theta_{23} = 45^\circ$.



- For $\mathbb{Z}_2 \times \mathbb{Z}_2$, the deviation from maximality is around 1° . However, a greater than 2σ exclusion of maximality would be possible for \mathbb{Z}_5 with T2K and $\text{NO}\nu\text{A}$.

In conclusion...

- **Parameter correlations**, predicted in some models of discrete symmetries, are attractive (and accessible) signatures of fairly generic model-building ideas.
- I discussed a recent study on **A_5 with a generalised CP**. We found 10 distinct sets of parameter correlations.
- High-precision θ_{12} measurements (JUNO, RENO-50) will be able to distinguish between these candidates; although, the correlations only weakly depend on θ_{13} .
- θ_{23} - δ correlations are interesting in these models:
 - **maximal-maximal** predictions for θ_{23} and δ (T2K, NO ν A)
 - CP conservation with θ_{23} - θ_{13} measurements (T2HK, DUNE).

Thank you.

backup slides

Solar sum rules

- The PMNS matrix measures the discrepancy between charged lepton and neutrino mass bases w.r.t a common flavour basis.

$$U_{\text{PMNS}} = U_e^\dagger U_\nu$$

such that

$$U_e^\dagger m_e^\dagger m_e U_e = \text{Diag}(m_e^2, m_\mu^2, m_\tau^2) \quad U_\nu^T m_\nu U_\nu = \text{Diag}(m_1, m_2, m_3).$$

- Leading-order mass matrices may be predominately symmetric but with higher-order symmetry violating terms. In the lepton sector, this would lead to

$$U_e = U_e^{\text{Symmetric}} V^{\text{Correction}} \quad \Longrightarrow \quad U_{\text{PMNS}} = V^{\text{Correction}} U_0$$

- We considered PMNS matrices of the form

$$U_{\text{PMNS}} = U_{23}^{e\tau} U_{12}^{e\tau} R_{23}^{\nu} R_{12}^{\nu} P$$

where R_{ij} denotes a real rotation from the neutrino sector, U_{ij} a unitary transformation in the ij -plane, and P a diagonal matrix of phases.

- There is a simple derivation for a prediction of $\cos \delta$. Multiplying out the PMNS matrix above,

$$|U_{\tau 1}/U_{\tau 2}| = \tan \theta_{12}^{\nu}$$

which, in terms of observable mixing parameters, can be expressed as

$$\cos \delta = \frac{t_{23} s_{12}^2 + s_{13}^2 c_{12}^2 / t_{23} - s_{12}^{\nu 2} (t_{23} + s_{13}^2 / t_{23})}{\sin 2\theta_{12} s_{13}}.$$

Solar sum rules: leading-order patterns

Survey of possibilities finds 4, all with $\theta_{13}^\nu = 0$ and $\theta_{23}^\nu = \frac{\pi}{4}$:

$$\theta_{12}^\nu = 45^\circ$$
$$\begin{pmatrix} \frac{1}{\sqrt{2}} & \frac{1}{\sqrt{2}} & 0 \\ -\frac{1}{2} & \frac{1}{2} & \frac{1}{\sqrt{2}} \\ \frac{1}{2} & -\frac{1}{2} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

$$\theta_{12}^\nu = 35.3^\circ$$
$$\begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

$$\theta_{12}^\nu = 31.7^\circ$$
$$\begin{pmatrix} \frac{\varphi}{\sqrt{2+\varphi}} & \frac{1}{\sqrt{2+\varphi}} & 0 \\ -\frac{1}{\sqrt{4+2\varphi}} & \frac{\varphi}{\sqrt{4+2\varphi}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{4+2\varphi}} & -\frac{\varphi}{\sqrt{4+2\varphi}} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

$$\theta_{12}^\nu = 20.9^\circ$$
$$\begin{pmatrix} \frac{\varphi}{\sqrt{3}} & \frac{\varphi_g}{\sqrt{3}} & 0 \\ -\frac{\varphi_g}{\sqrt{6}} & \frac{\varphi}{\sqrt{6}} & \frac{1}{\sqrt{2}} \\ \frac{\varphi_g}{\sqrt{6}} & -\frac{\varphi}{\sqrt{6}} & \frac{1}{\sqrt{2}} \end{pmatrix}$$

Solar sum rules: predictions for δ

- We can use global data to make statistical predictions for $\cos \delta$.

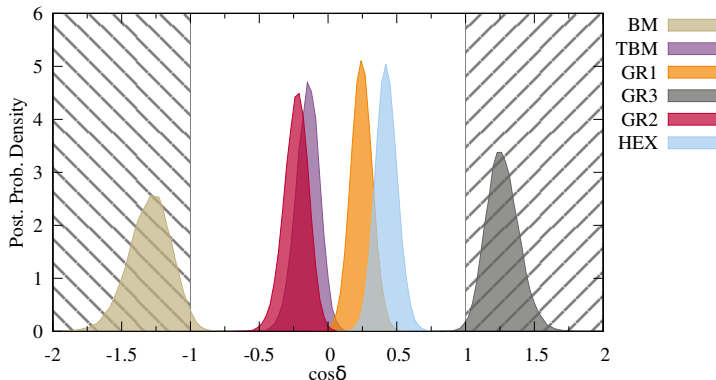


Figure from PB, King, Luhn, Pascoli, Schmidt [1410.7573](#).