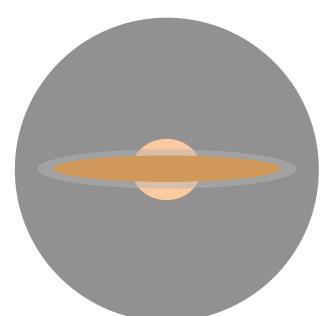
## Dark Disks

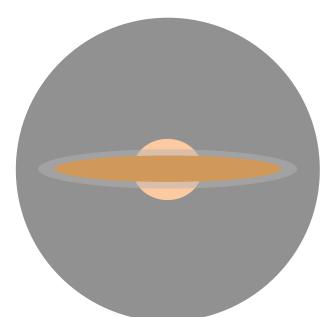
Invisibles, Madrid June 25<sup>th</sup> 2015



Matthew McCullough CERN

# Dark Disks (and other Dark exotica)

Invisibles, Madrid June 25<sup>th</sup> 2015

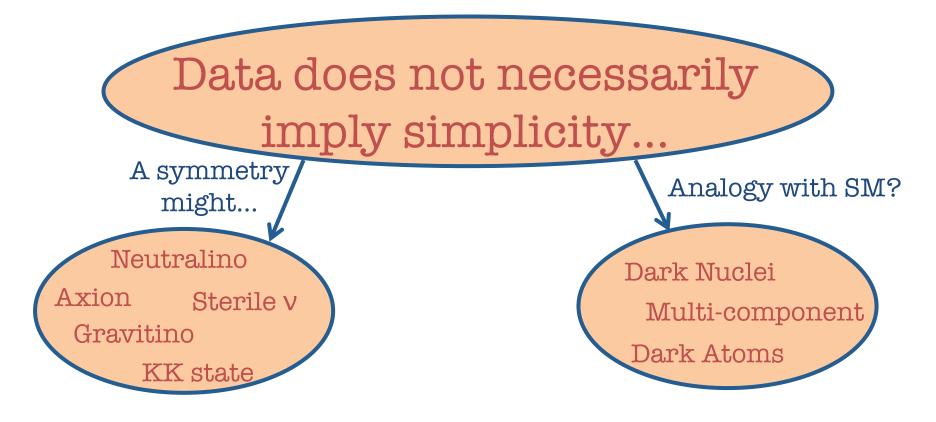


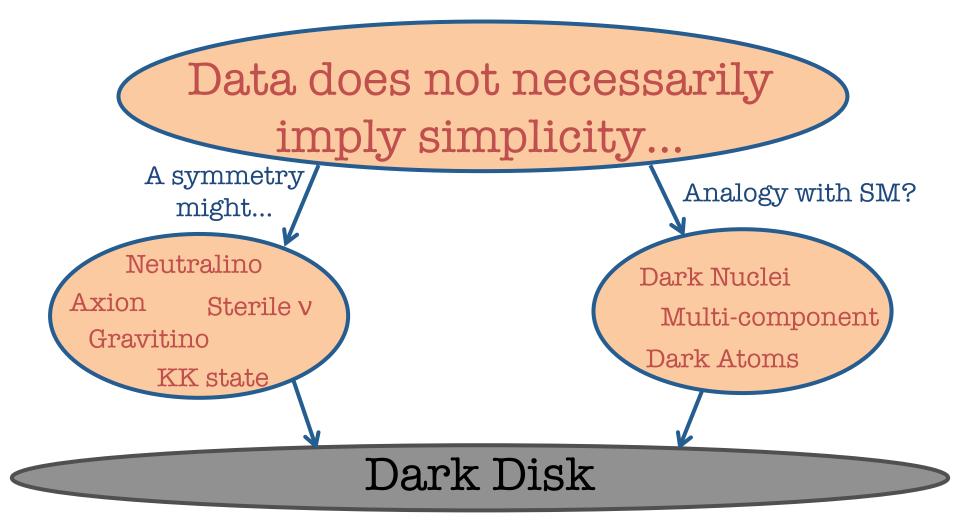
Matthew McCullough CERN

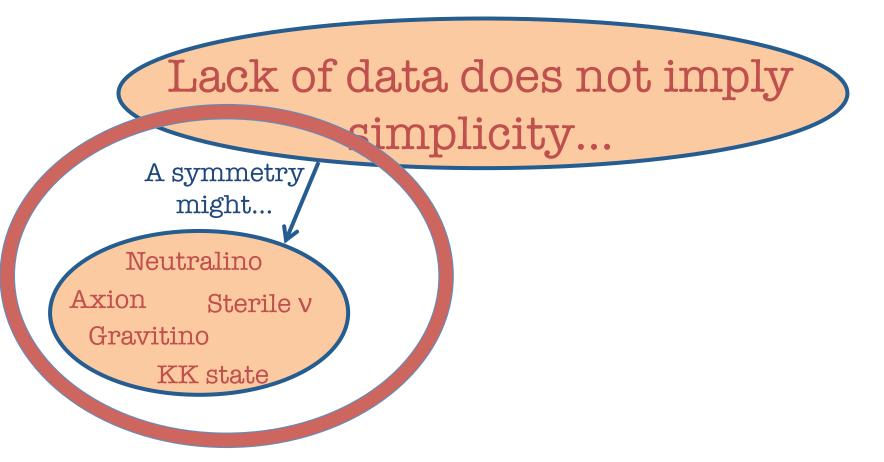
# DarkDarkDarkWIMPsAtomsDisksNuclei

#### Dark Matter Menu

Image: smalljude, Wikimedia commons.







Considered in context of standard Cold Dark Matter WIMP paradigm for some time...

Lake 1989: "Must the Disk and Halo Dark Matter be Different?" Stellar Disk Central Stellar Bulge

Considered in context of standard Cold Dark Matter WIMP paradigm for some time...

Lake 1989: "Must the Disk and Halo Dark Matter be Different?" Stellar Disk Central Stellar Bulge

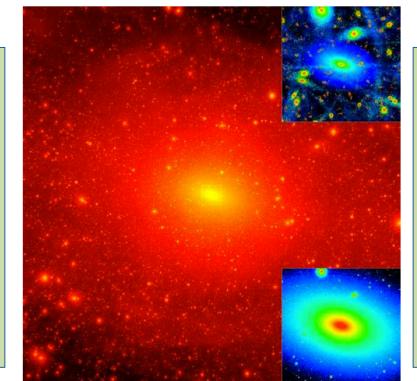
Can show up in modern simulations:

• Read, Lake, Agertz, Debattista (0803.2714)

How could they form? Lake 1989: "Must the Disk and Halo Dark Matter be Different?"

This image not a dark disk, but pretty. 800kpc and 40 kpc.

From Diemand, Kuhlen, et al. 0805.1244



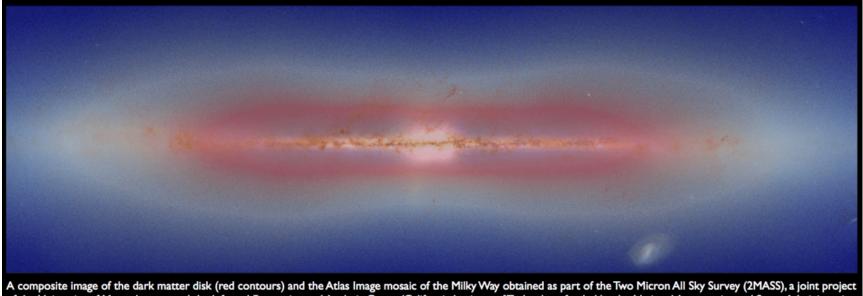
Can result from hierarchical structure formation leading to mergers of subhalos into halo.

Can show up in modern simulations:

- Read, Lake, Agertz, Debattista (0803.2714)
- However, see Cooper, Cole, Frenk, White, et al. (0910.3211)

What do they look like?

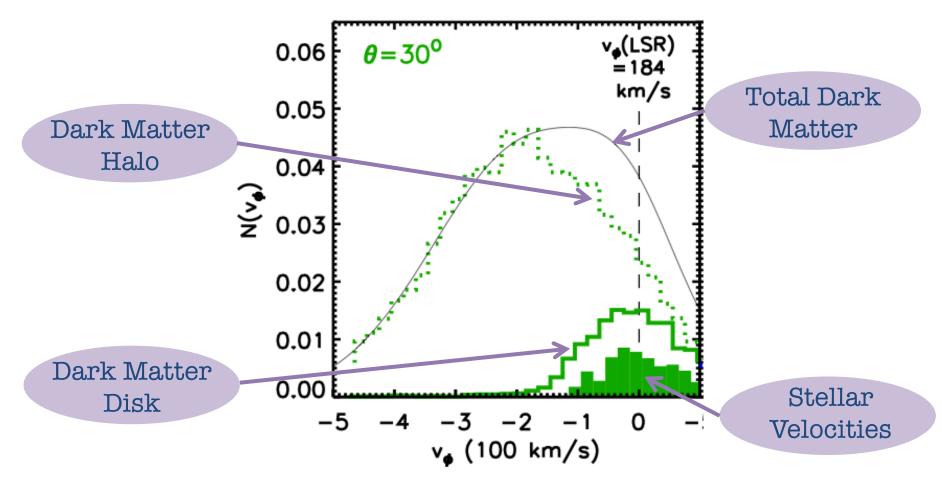
#### In a magazine they look like this:



A composite image of the dark matter disk (red contours) and the Atlas Image mosaic of the Milky Way obtained as part of the Two Micron All Sky Survey (2MASS), a joint project of the University of Massachusetts and the Infrared Processing and Analysis Center/California Institute of Technology, funded by the National Aeronautics and Space Administration and the National Science Foundation. Credit: J. Read & O.Agertz.

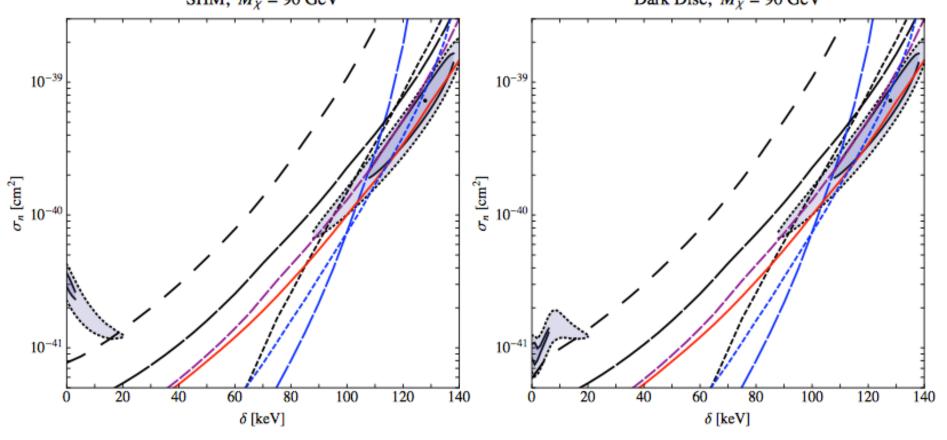
# <u>www.universetoday.com</u> (Nancy Atkinson) "Dark Matter Halos? How About Disks, Too"

In a paper they look like this:

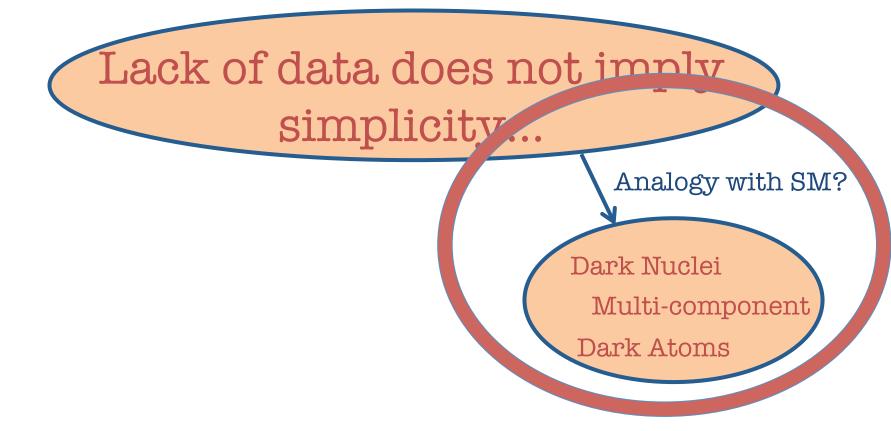


• Purcell, Bullock, Kaplinghat (0906.5348)

Modifications to exotic scattering candidates as well. Old paper on inelastic dark matter: SHM,  $M_{\chi} = 90 \text{ GeV}$ Dark Disc,  $M_{\chi} = 90 \text{ GeV}$ 

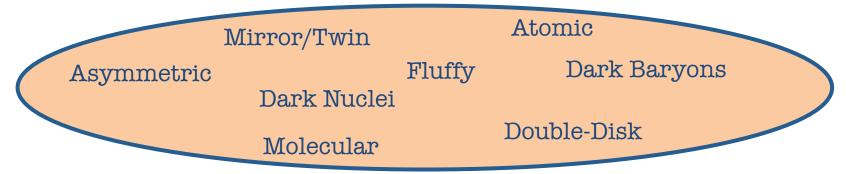


• March-Russell, McCabe, MM (0812.1931)



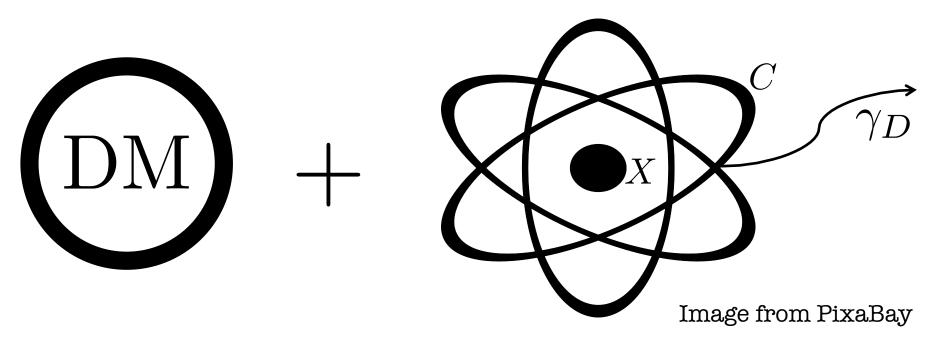
WIMPs are interesting. However...

• The dark sector may be a cocktail of states



- Subcomponent with self-interactions, "Partially Interacting Dark Matter"
- Possibilities rich: Fan, Katz, Randall, Reece.
- Focus here: "Double Disk Dark Matter"

Consider a simple subcomponent model:



Could have a DM subcomponent like a dark atom. Plausible possibility. Phenomenology potentially rich.

Dark Atoms: Kaplan, Krnjaic, Rehermann, Wells. Cline, Liu, Xue. Dark Atom Subcomp: Fan, Katz, Randall, Reece.

Cooling allows DDDM to collapse into a disk. Details different from SM as no supernova feedback etc. Can cool to a temperature

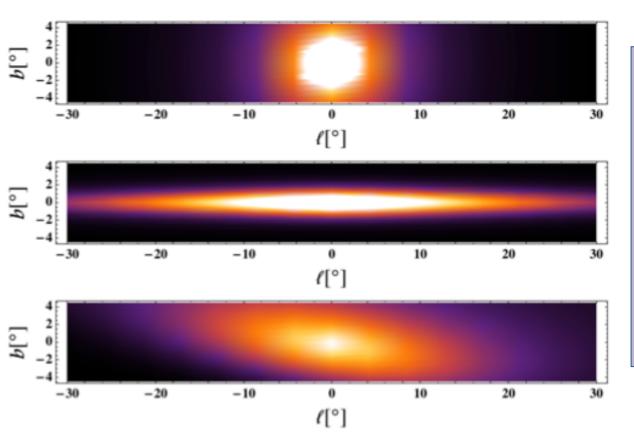
$$T_{cool} \sim 0.1 E_{Binding}$$

And the thickness of the resulting dark disk can be quite small

$$z_d \approx 2.5 \text{pc} \left(\frac{\alpha_D}{0.02}\right)^2 \frac{m_C}{10^{-3} \text{GeV}} \frac{100 \text{GeV}}{m_X}$$

Thus an entirely different mechanism leads to a very different disk from the WIMP disk scenario. Fan, Katz, Randall, Reece

Implications for standard searches could be significant.

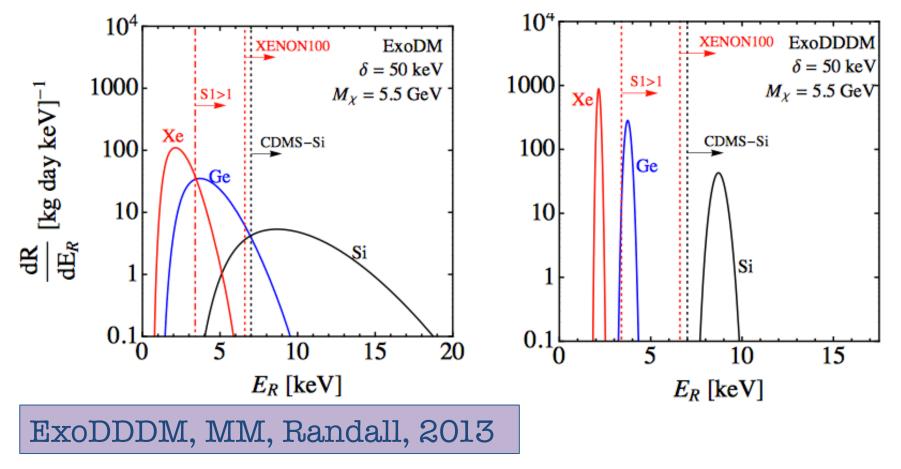


Spectacular modifications of indirect detection signature morphology.

Including possible tilted disks.

Fan, Katz, Randall, Reece

Or for exotic models, such as exothermic scattering, could lead to very exotic direct detection signatures.

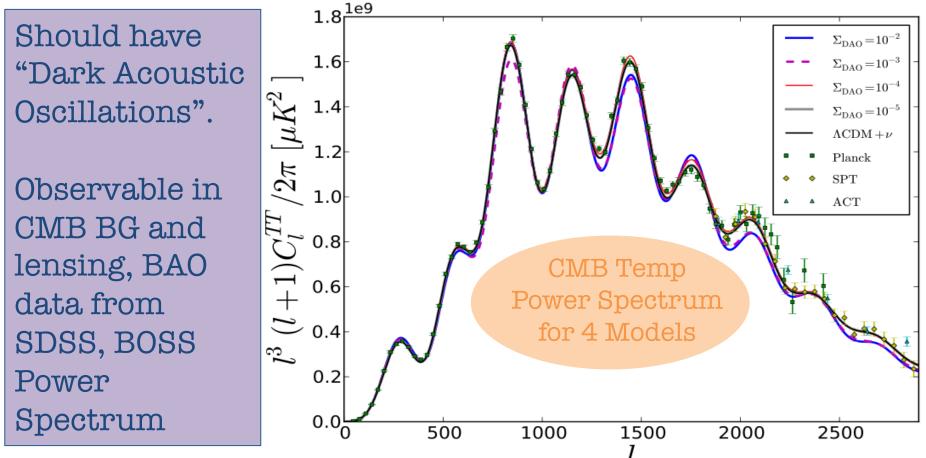


Hipparcos (GAIA Future) local star velocity surveys

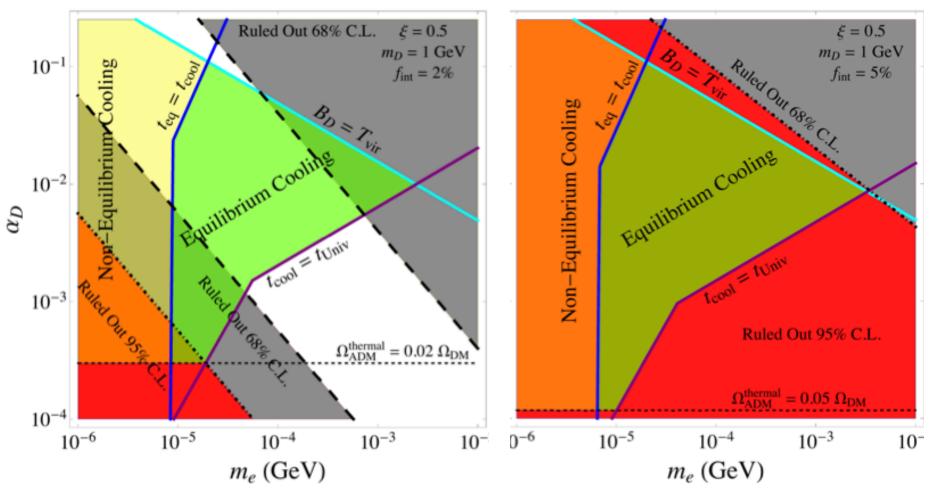
#### SEE POSTER: ERIC KRAMER.

### **Complex Subcomponents**

If subcomponent behaves like matter, similar cosmological effects. Cyr-Racine, de Putter, Raccanelli, Sigurdson 2013



Planck+WP+High-l+BAO+Lens constraints:



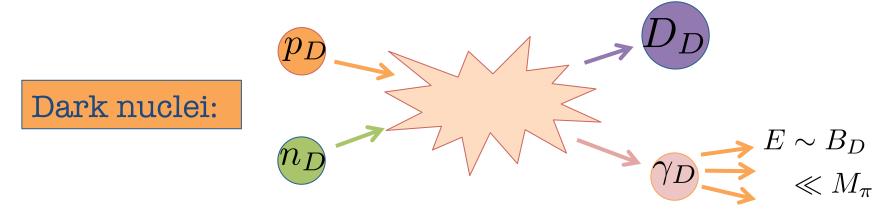
Also: Buckley, Zavalla, Cyr-Racine, Sigurdson, Vogelsberger, 2014





#### Dark Nuclei

Usual lore for asymmetric DM: no annihilation means no indirect detection!



Dark nucleosynthesis preserves dark baryonnumber and can lead to indirect detection signatures! (SM analogy useful here)

"Dark Nuclei" Detmold, MM, Pochinsky 1406.2276

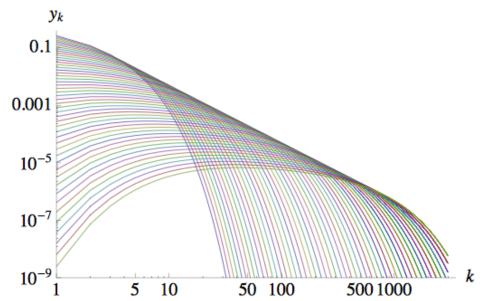
### Large Dark Bound States

A number of works have considered bound states of elementary and composite dark sector states.

• Wise & Zhang. Krjaic & Sigurdson. Hardy, Lasenby, March-Russell, & West.

Distribution of states with dark nuclei number k.

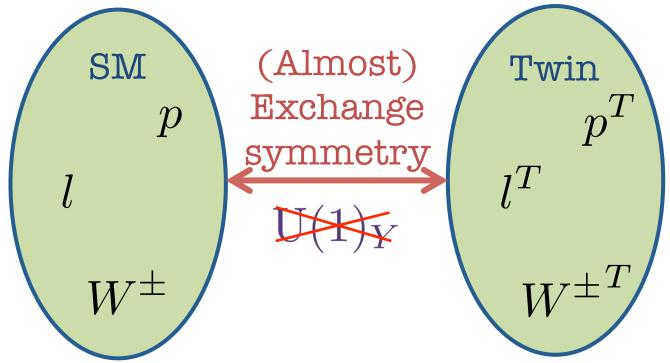
Significant modifications of standard signals.



"Nuggets" of dark matter possible, total mass far exceeding usual unitarity bounds.

#### Twin Miracle

Take analogy between SM and DM to an extreme level (for a reason: Little Hierarchy Problem)

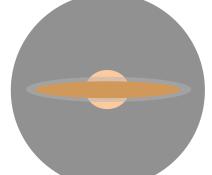


Miracle: Twin weak interactions similar to SM Asymmetry: Mass scale for ADM now motivated! Craig & Katz. Garcia Garcia, Lasenby, & March-Rusell. Farina.

### Conclusions

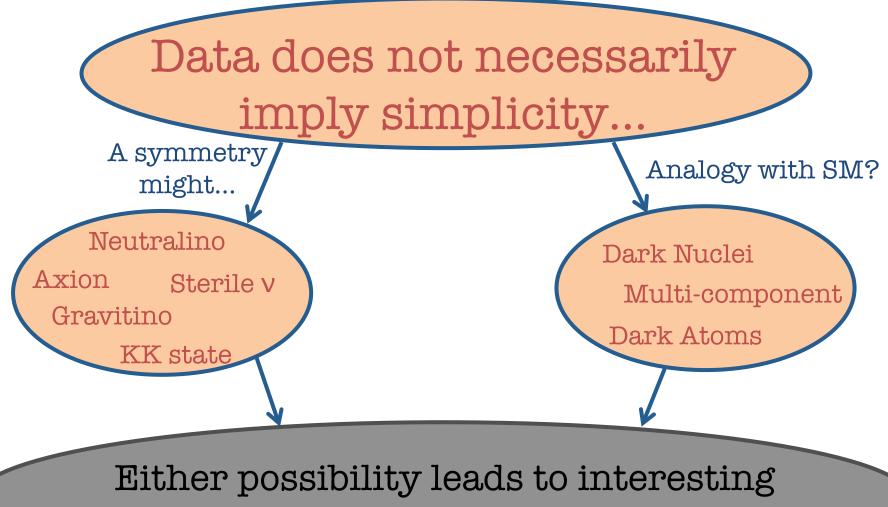
Dark Disks: plausible dark matter substructure

- WIMP Disks: possible consequence of hierarchical structure formation
  - Modifications to dark matter detection



- Subcomponent disks: plausible prediction of non-minimal dark subcomponent scenarios
  - Dramatic consequences for detection
  - Already driving new signatures, new applications for precision cosmology

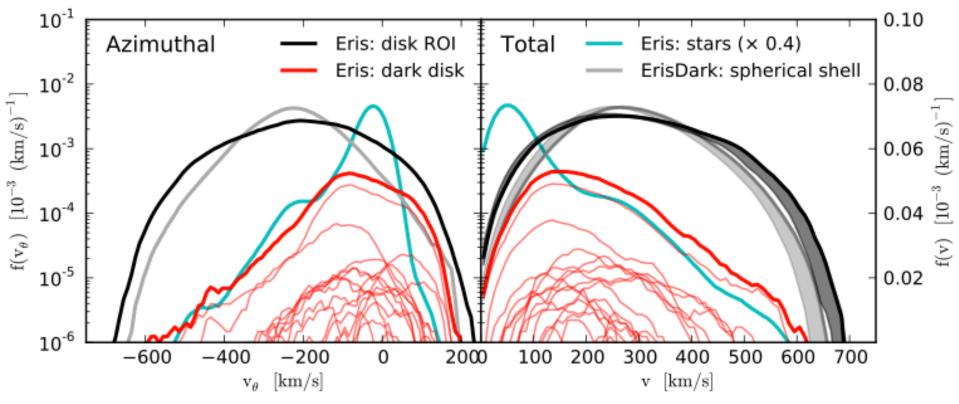
#### Conclusions



and exciting phenomenology.

Backup Slides

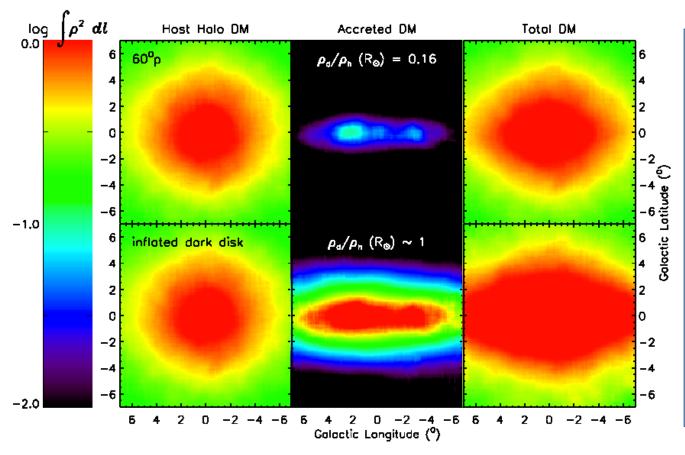
#### In a paper they look like this:



Notice that dark disk does not dominate DM density anywhere.

• Pillepich, Kuhlen, Guedes, Madau (1308.1703)

#### Indirect Detection modified

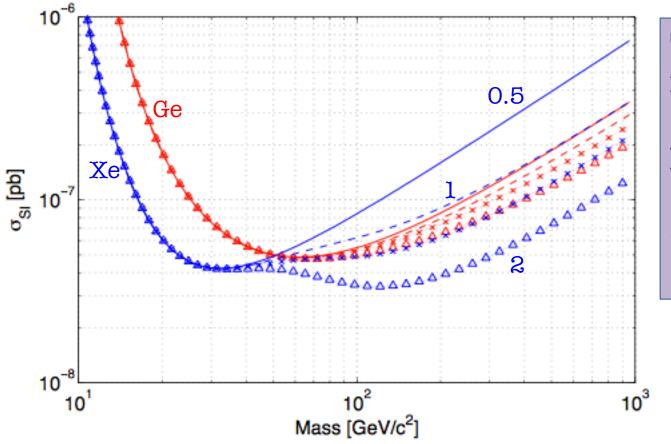


Signatures in gamma rays depend on lineof-sight density integral.

Morphology may be significantly influenced.

• Purcell, Bullock, Kaplinghat (0906.5348)

Direct Detection signatures modified due to enhanced low-velocity component.

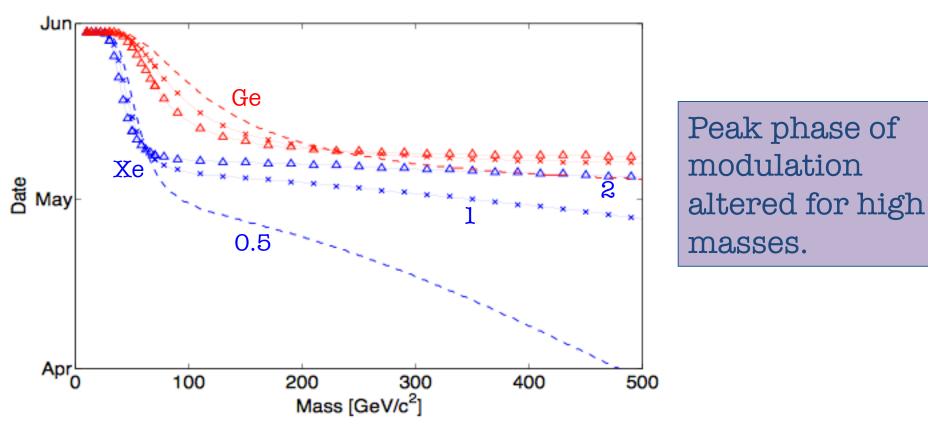


• Bruch, Read, Baudis, Lake (0804.2896)

Total scattering rate enhanced, especially for heavier DM.

Changes cross section limits.

Direct Detection signatures modified due to enhanced low-velocity component.



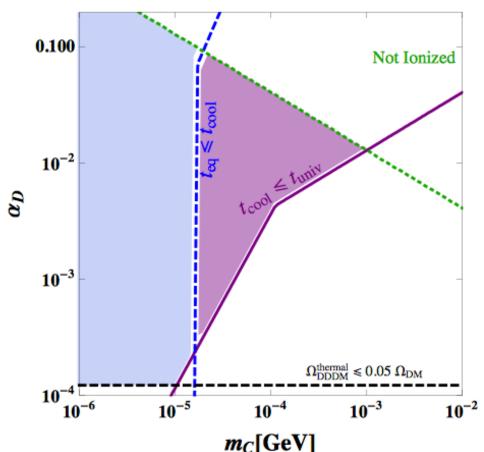
• Bruch, Read, Baudis, Lake (0804.2896)

Annihilation into dark photons efficient, so imagine an asymmetric scenario (like SM!).

 $\epsilon = 0.05, m_X = 1 \text{ GeV}, n_X = n_C = 3.3 \times 10^{-4} \text{ cm}^{-3}$ 

DDDM can cool adiabatically via Compton and Bremsstrahlung.

This allows for different structure within the DM halo and a significantly modified cosmology.



Fan, Katz, Randall, Reece

Wolf-Rayet Stars and Gamma Ray Bursts



In galaxy merger events Wolf-Rayet Stars form (think  $20 M_{sun}!$ ).

Wolf-Rayet believed to be progenitors of Black Holes.

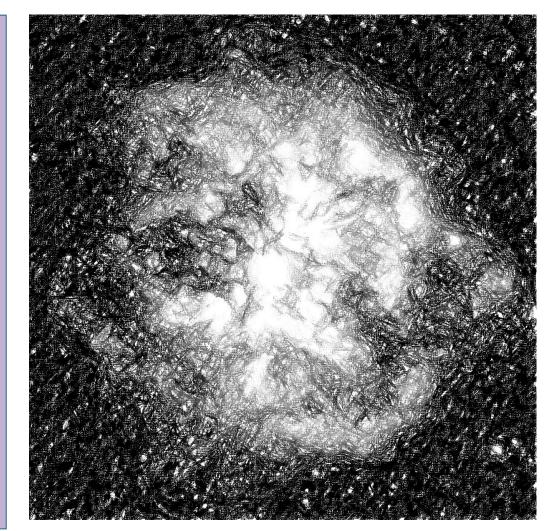
Blandford-Znajek mechanism: Magnetic flux threads BH, spindown via emission of radiation jets.

#### Fischler, J&D Lorshbough 2015

If merger involved significant amounts of DM, subcomponents would also be in merger region.

Dark magnetic field threads BH and... Dark Gamma Ray Burst!

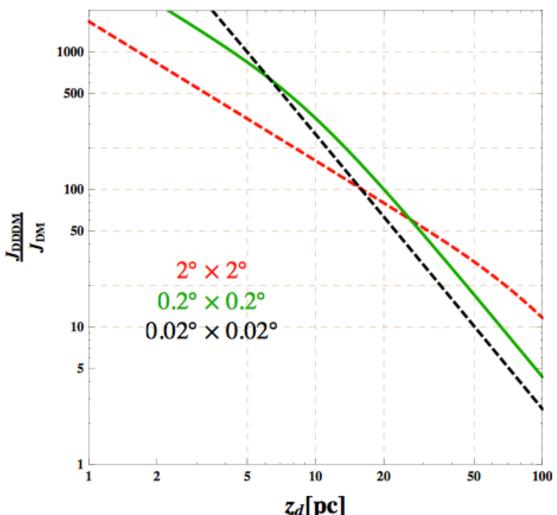
Mismatch between spindown and observed GRB energy outflow.



Implications for standard searches could be significant.

Large boost factors for indirect detection due to high density of disk.

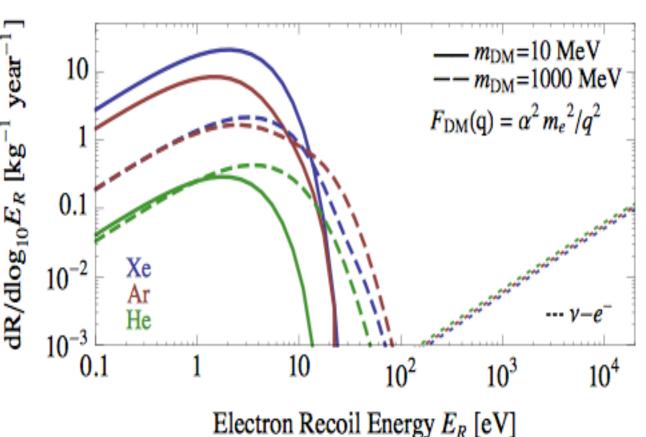
Also possible Sommerfeld enhancement.



Fan, Katz, Randall, Reece

Dark Disk likely co-rotating with stellar disk: Small relative velocity and dispersion suppress direct detection.

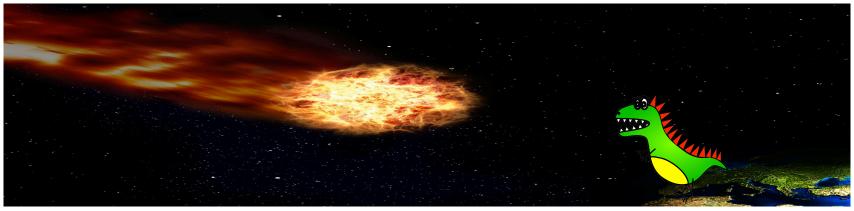
Perhaps show up in low-threshold experiments, such as proposed by Essig, Mardon, Volansky.



Fan, Katz, Randall, Reece

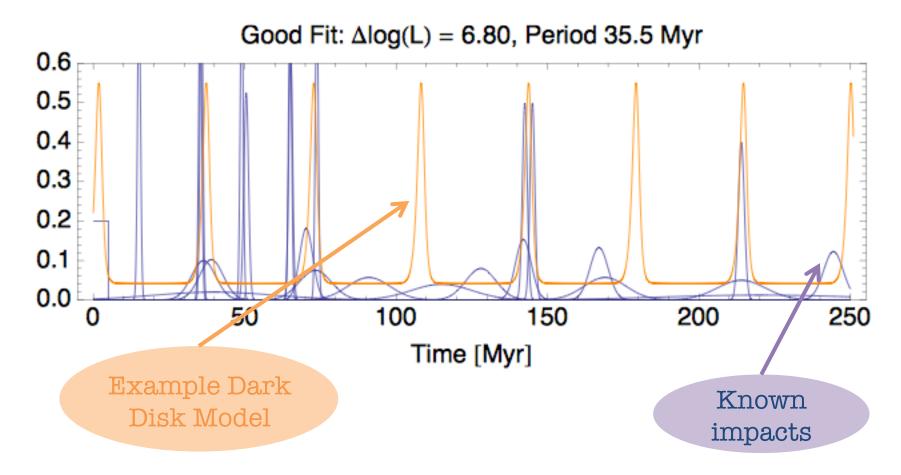
#### Periodic Comet Impacts

• Possible 35 Myr period in comet impacts, noted in 1984. Statistical significance small.



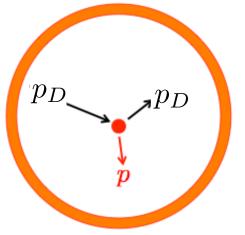
- Some possible explanations...
  - Nearby star "Nemesis Hypothesis"
  - Tidal stripping of Oort cloud by periodic potential of motion through galactic disk
- Latter taken seriously by serious people

If stellar disk could do it, then why not a dark disk? Randall and Reece 2014.



### Dark Nuclei

Components may be captured in Sun via scattering:



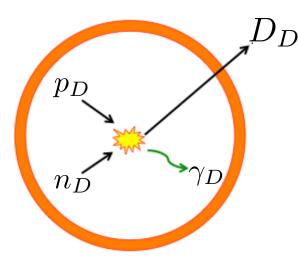
Standard scenario:

- Symmetric: May annihilate leading to detectable neutrinos
- Asymmetric: No annihilation, DM just builds up!

Dark nuclei:

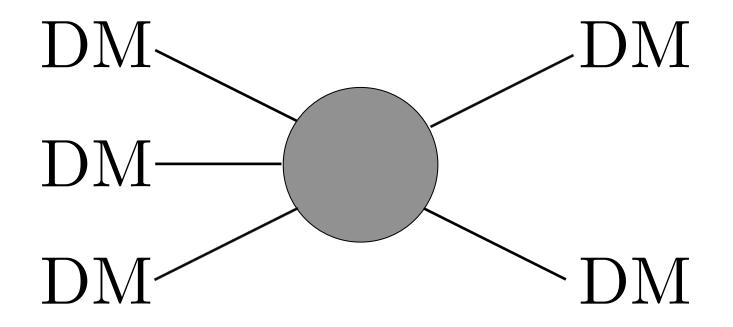
- Symmetric: May also annihilate leading to detectable neutrinos
- Symm and Asymm: Dark nucleo may eject dark nucleus from star!

 $\beta_D \sim B_D / M_D \gtrsim 0.01 > \beta_{\text{Escape}}$ 



#### SIMP Miracle

"Strongly Interacting" abundance miracle: Hochberg, Kuflik, Murayama, Volanksy, Wacker.



Emerges naturally in scenarios where the DM is a bound state of strongly coupled sector.