

inVisibles 15 Workshop 22nd - 26th June, 2015

Place: IFT and Thyssen-Bornemisza Museum, Madrid

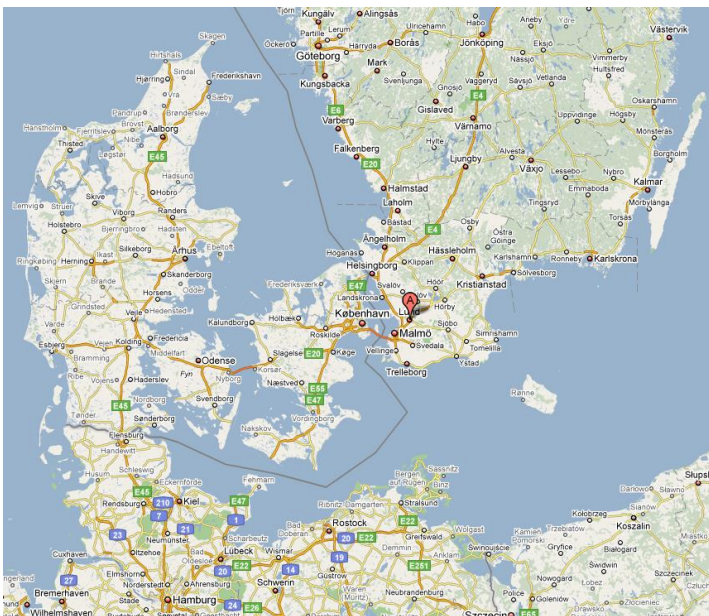
ESSnuSB

An aerial night-time photograph of the ESS (European Spallation Source) facility. The image shows a large, circular, illuminated structure in the center, surrounded by various buildings and infrastructure. A bright, glowing yellow and orange beam of light originates from the center and extends towards the top of the frame, representing the particle beam path. The surrounding area is dark, with some lights from the facility and nearby roads visible.

Elena WILDNER
CERN
for the ESSnuSB Collaboration

The European Spallation Source (ESS)

- ESS is a **neutron spallation source** that is being built by a collaboration of 17 European countries.
- ESS is located in southern Sweden (Lund)



ESS Technical Design Report, April 23, 2013, ESS-doc-274
<http://europeanspallationsource.se/documentation/tdr.pdf>

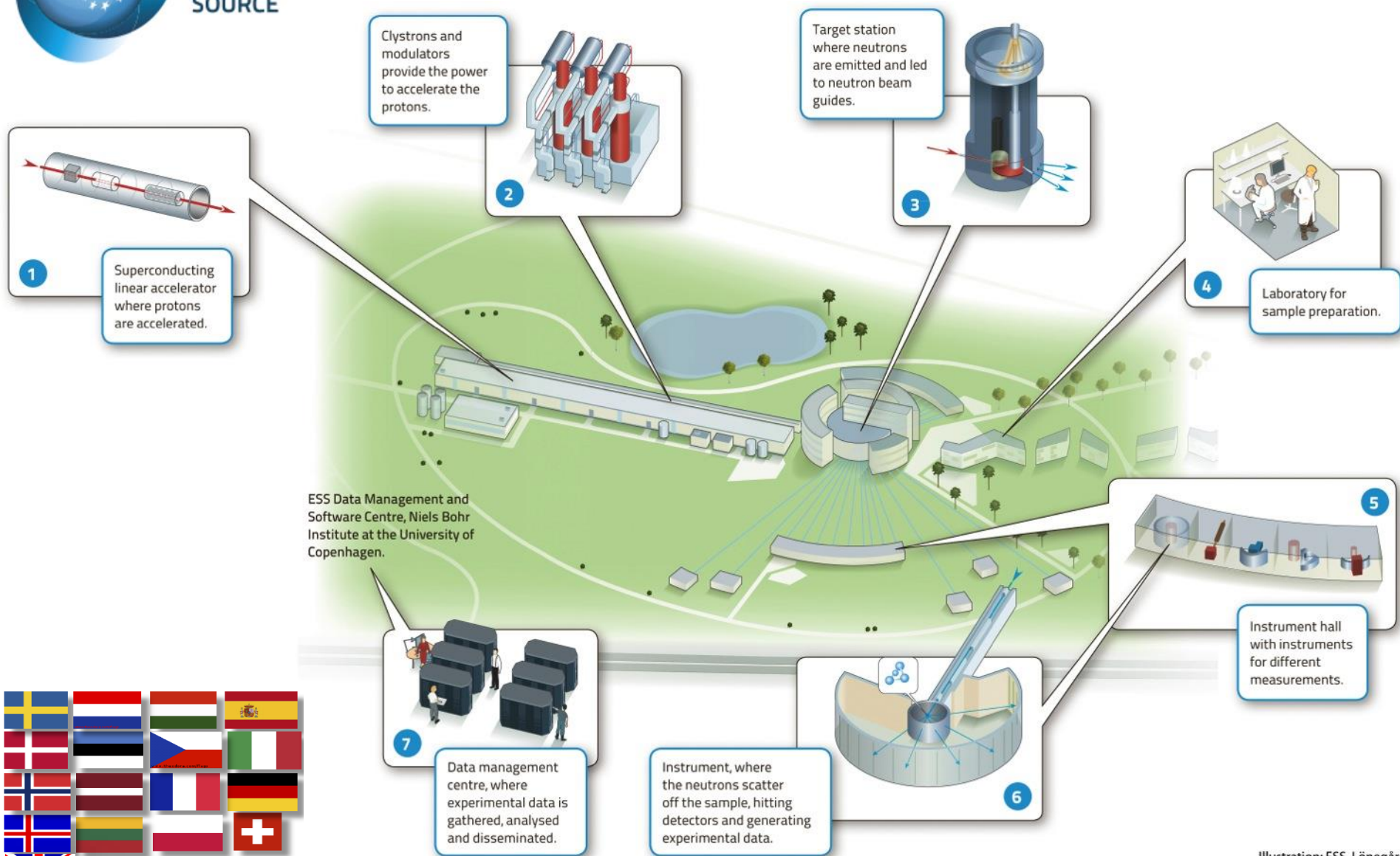
The linac tunnel on 10 april 2015



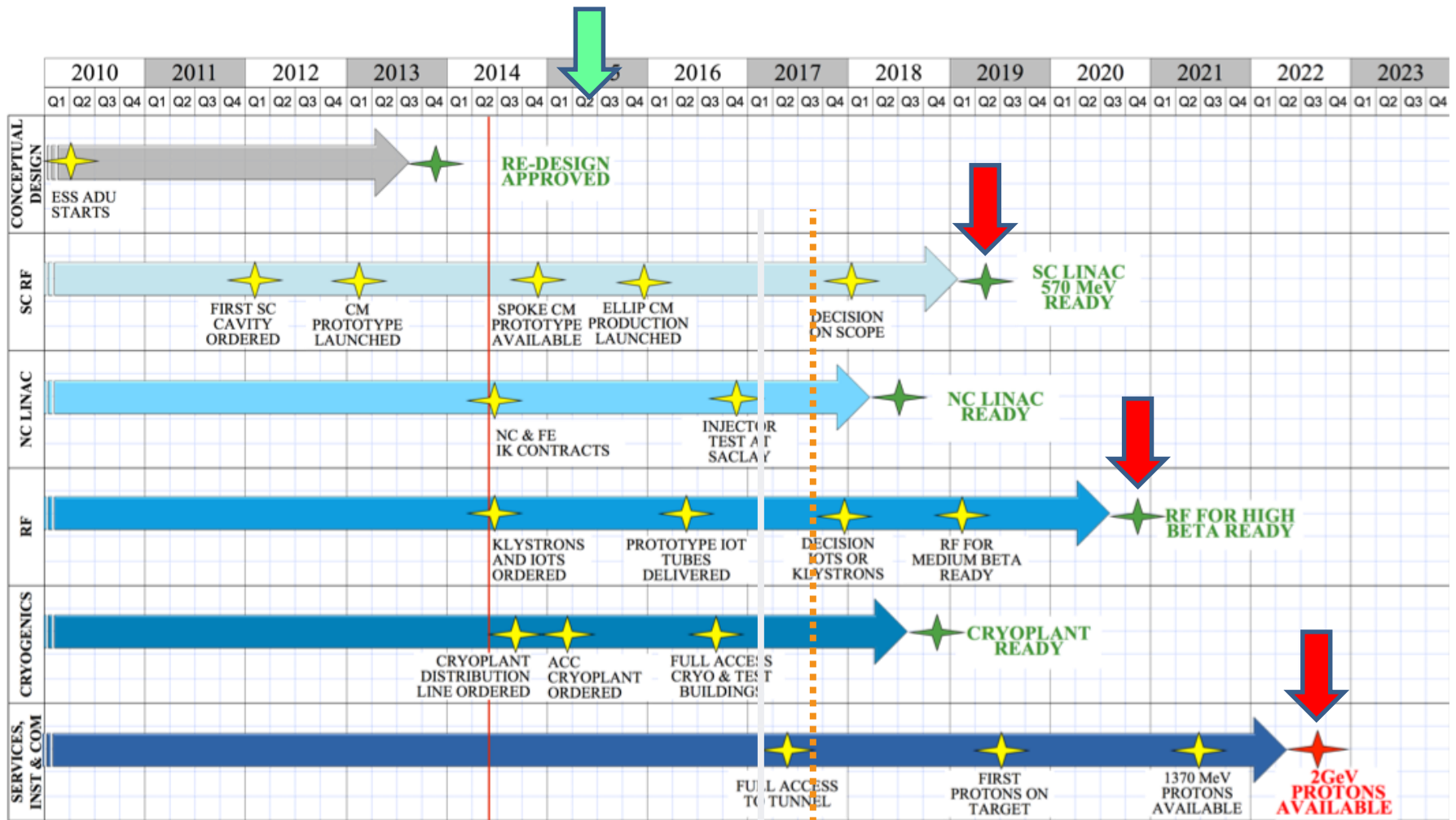


EUROPEAN
SPALLATION
SOURCE

ESS as proton driver for spallation



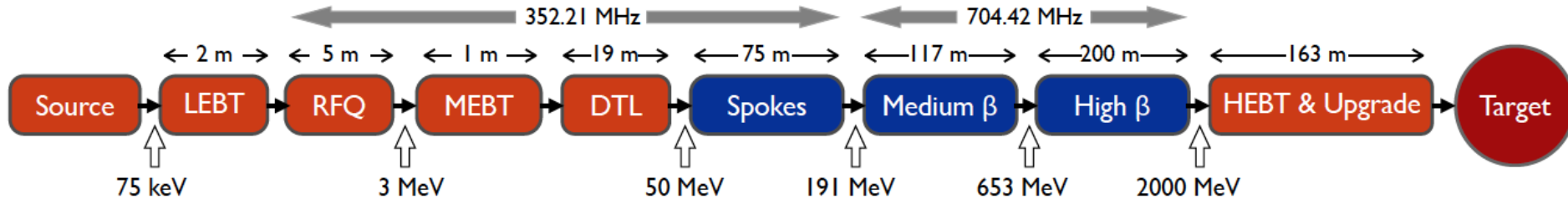
ESS LINAC PROJECT SCHEDULE



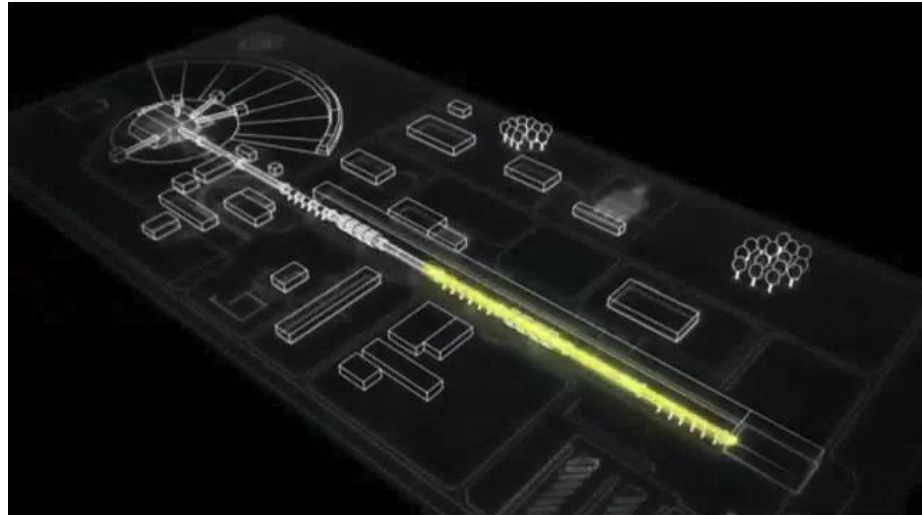
Seminar at Karlsruhe KIT 2015-04-28

Tord Ekelöf Uppsala University

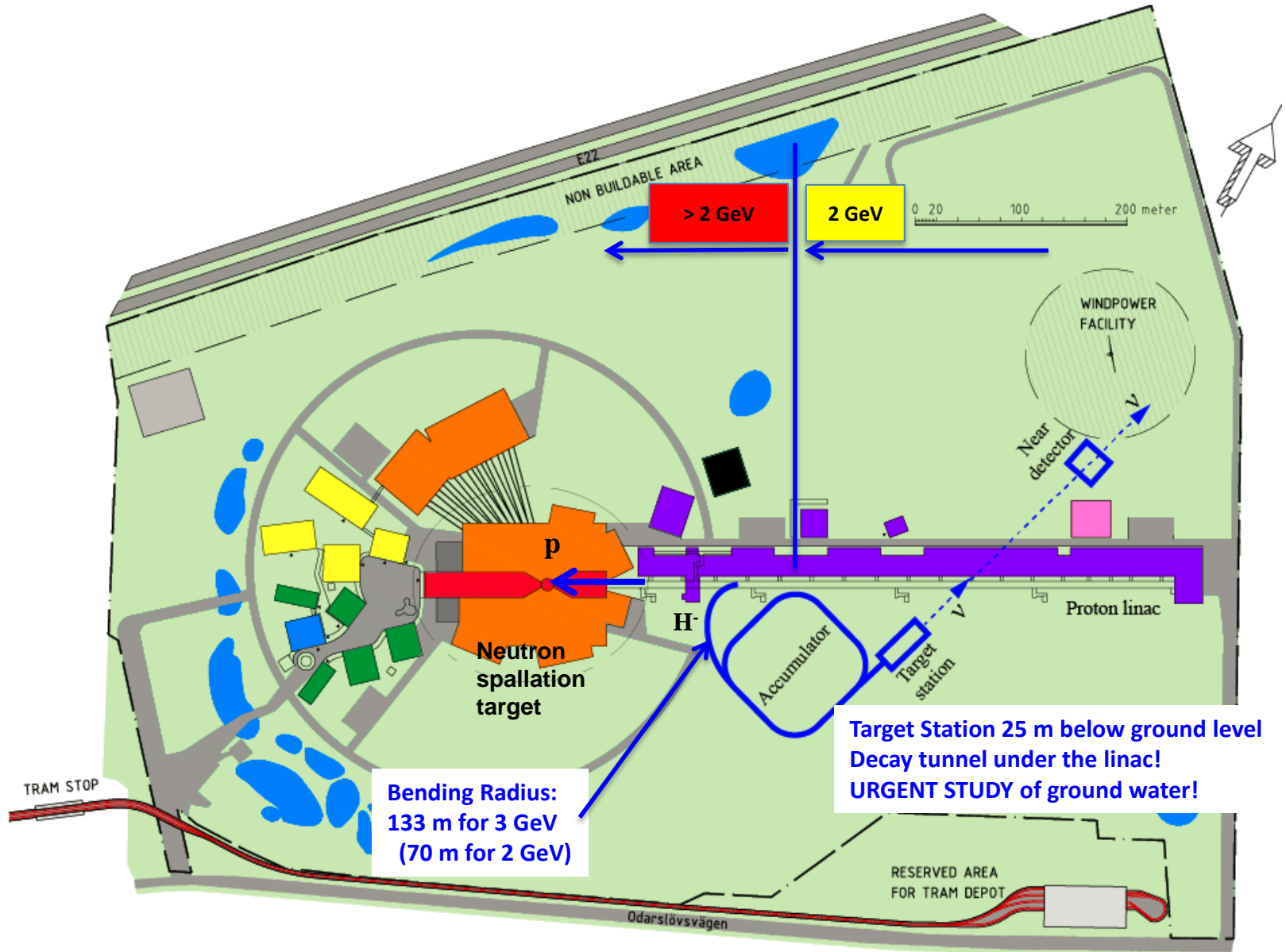
ESS 5MW proton linac



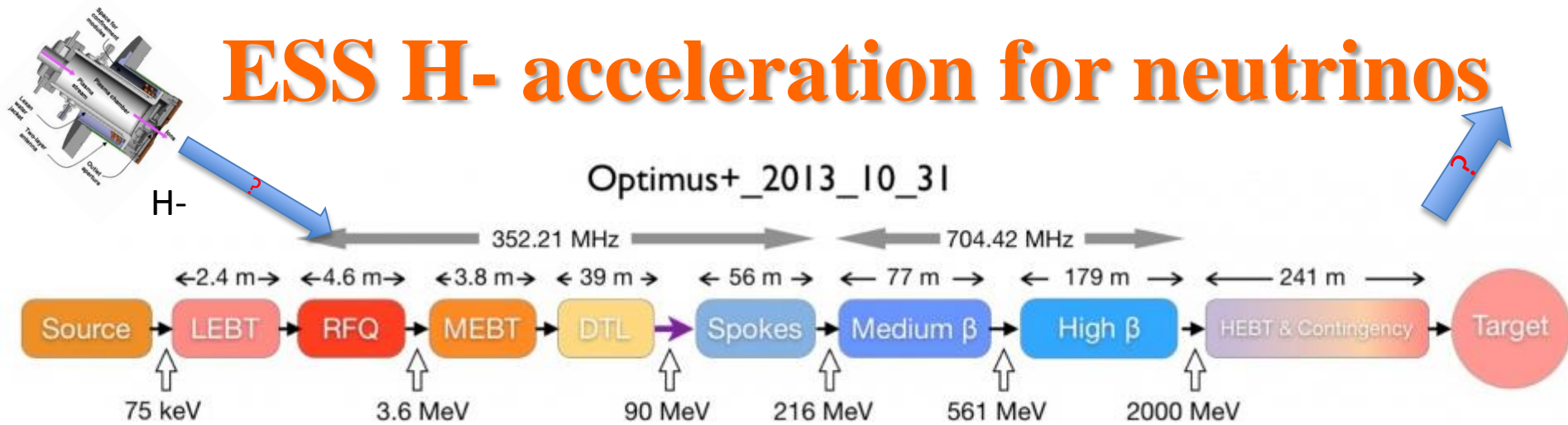
- The ESS will be a copious source of spallation neutrons
- 5 MW average beam power
- 125 MW peak power
- 14 Hz repetition rate (2.86 ms pulse duration, 10^{15} protons)
- 2.0 GeV protons (up to 3.5 GeV with linac upgrades)



ESSnuSB on ESS site



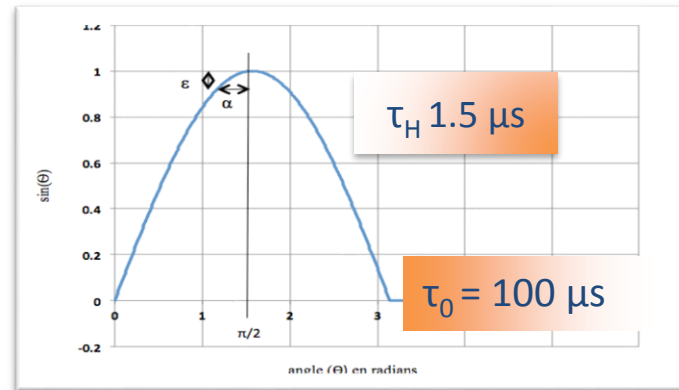
ESS H- acceleration for neutrinos



- The ESS linac for neutron spallation: **proton acceleration 14 Hz**
- Duty factor low (4%); **some additional capacity is available**
- Repetition rate can be increased to permit extra acceleration cycles (total power of extra cycles 5 MW)
- **ESS would accelerate in total 10 MW!**
- **Charge in the accumulator $1.1 \cdot 10^{15}$ protons 2 GeV, big challenge**
- **$>2.7 \times 10^{23}$ 2 GeV p.o.t/year**

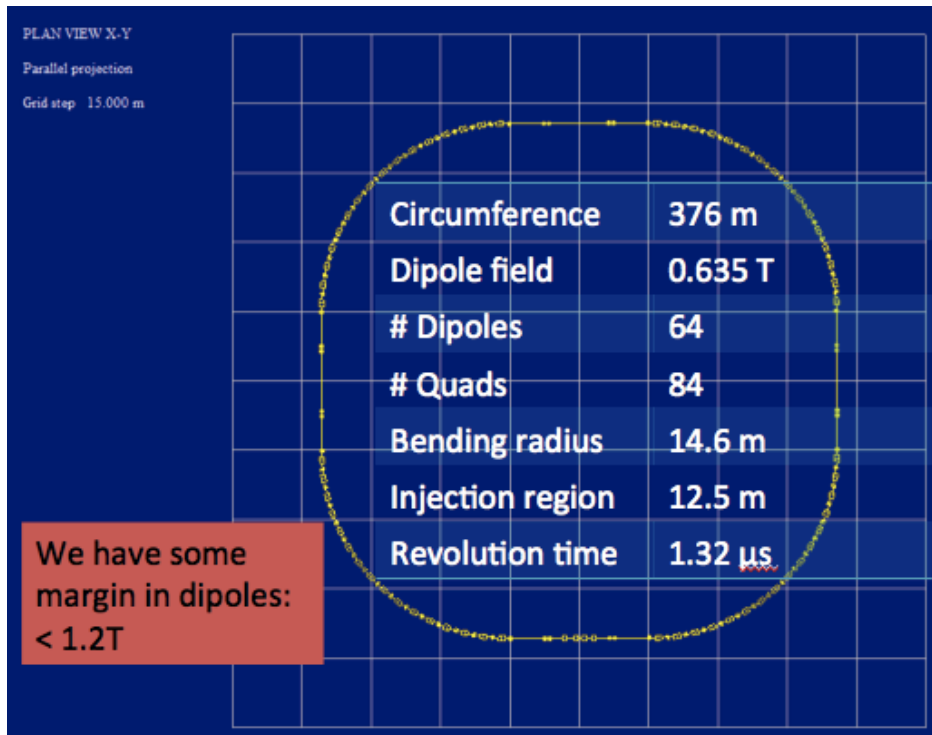
ESS Accumulator for neutrino production

- **Constraints on present neutrino target focusing system: short pulses**



- **Solution: reduce beam pulse length:**
 - Shorten the linac pulse by accumulating the linac beam
- **Accumulator constraints**
 - **Reasonably-sized** accumulator ring circumference and apertures
 - **Multiturn injection** of high intensity linac beam: **we need H-**
 - **High intensities** may cause collective effects and beam loss

ESS Accumulator for neutrino production (II)



Accumulator lattice based on Spallation Neutron Source, Oak Ridge

Layout optimization and beam dynamics

Transfer lines and extraction of beam from linac

Switchyard

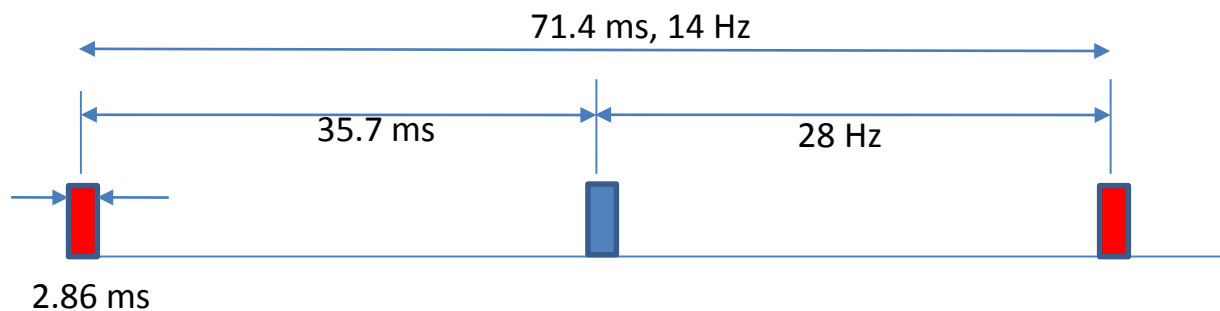
Linac Pulsing, options



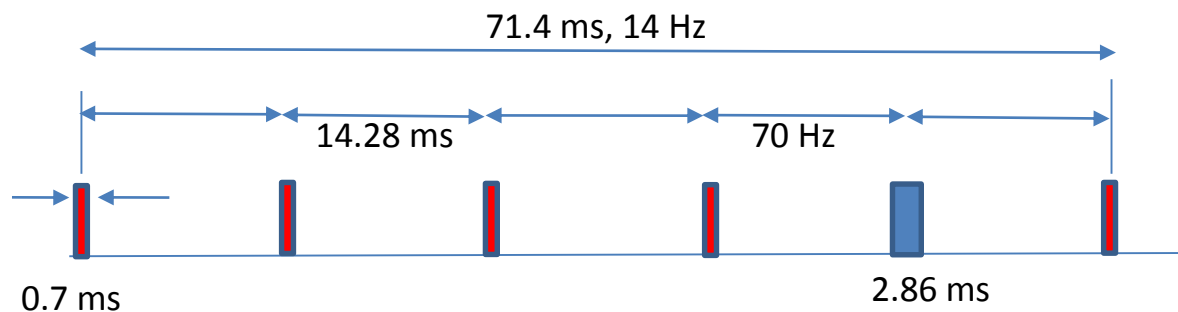
Neutrino (H⁻ pulse)



Neutron (proton pulse)



Insert one H⁻ pulse 28 Hz
May need several accumulators



Mitigation of charges in
accumulator: shorter pulses,
higher pulsing frequency of linac.

Implies some overhead (cavity
filling)

The Linac has to be upgraded

Cost Impact	Per modulator	Total (45 modulators)		Total (45 modulators)
1)- Adding extra capacitor charger modules	+ 60 <u>kEURO</u>	+ 3 MEURO		
2)- Re-winding HVHF transformers and output filter inductors	+ 100 <u>kEURO</u>	+ 4.5 MEURO	Cost Impact	
3)- Labour costs (contract follow-up, testing, etc.)		+ 5 MEURO	1)- Add additional 10kV/600V distribution transformers (doubling the total quantity, keeping existent ones)	+ 2 MEURO
Total cost increase for modulators' upgrade		+ 12.5 MEURO (+ 100%)	2)- Add additional LV power cables and switchboards (doubling the total quantity, keeping existent ones)	+ 1.5 MEURO
			3)- Labour costs (contract follow-up, testing, etc.)	+ 2 MEURO
Footprint Impact	Per modulator		Total cost increase for AC grid upgrade <i>(only for LV grid; new buildings and CF expenses not included)</i>	+ 5.5 MEURO (+ 100%)
Footprint required for additional capacitor chargers	~ 1.2m x 1m			
			Other impacts (to CF)	
			New buildings will be required for extra LV transformers and switchboards; Double footprint required.	
			An additional HV 120kV/10kV power transformer will be required	
			All HV distribution lines and protection devices at 120kV and 10kV levels will be doubled	

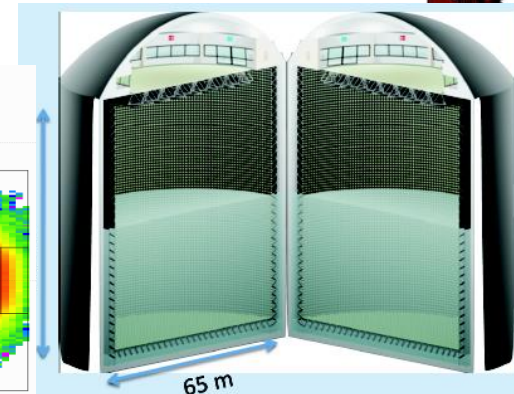
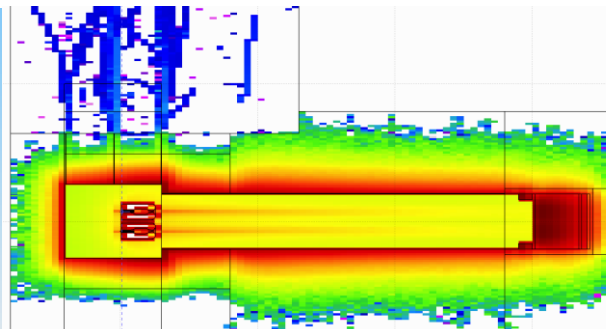
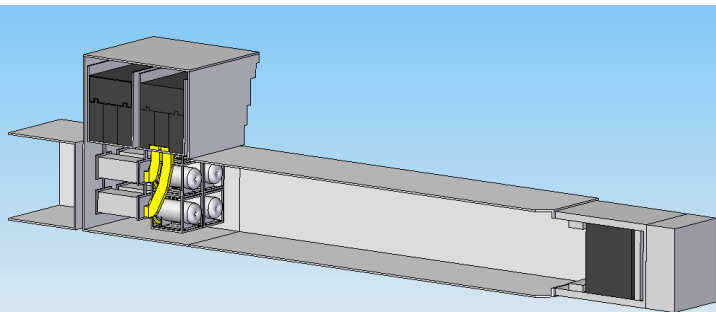
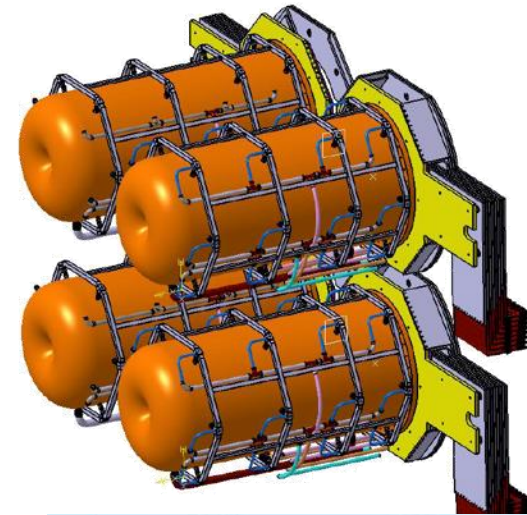
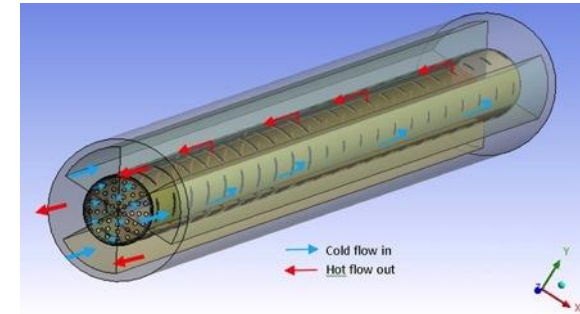
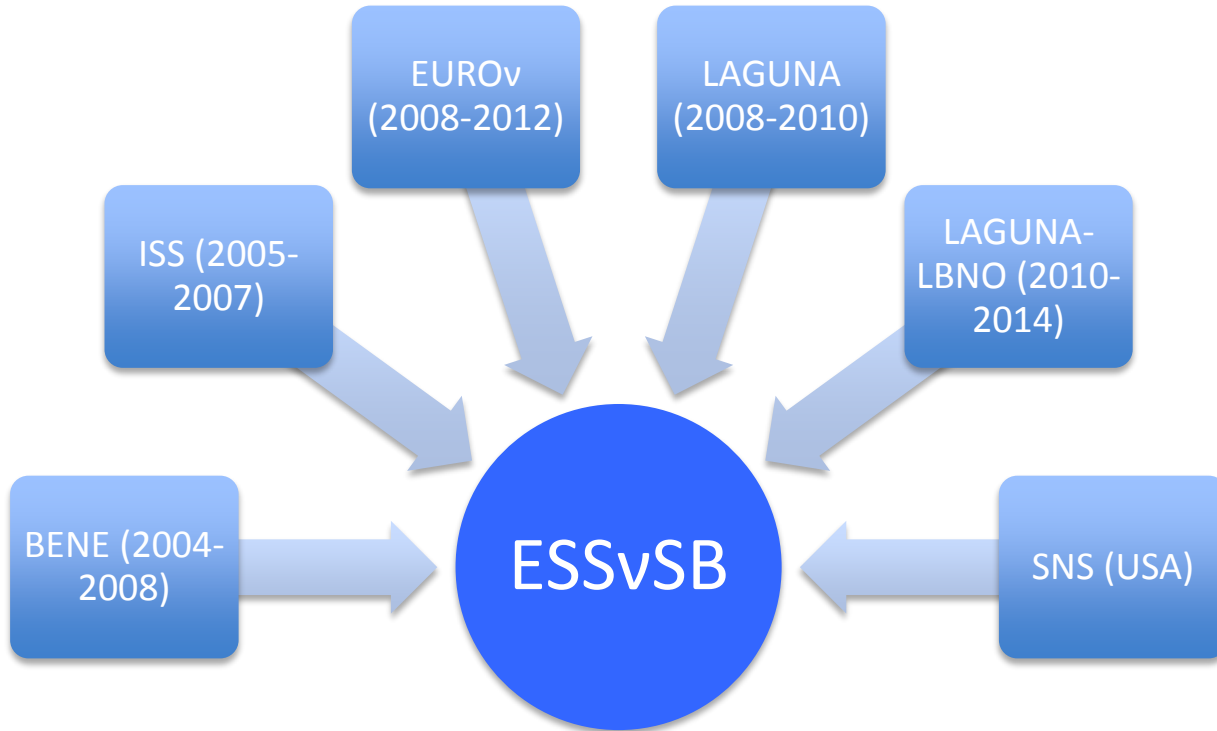
To be prepared for preferably at an early stage

List of need prepared and taken into consideration by the accelerator division

Some options already built in

Infrastructure upgrades are important to plan for

Existing Expertise



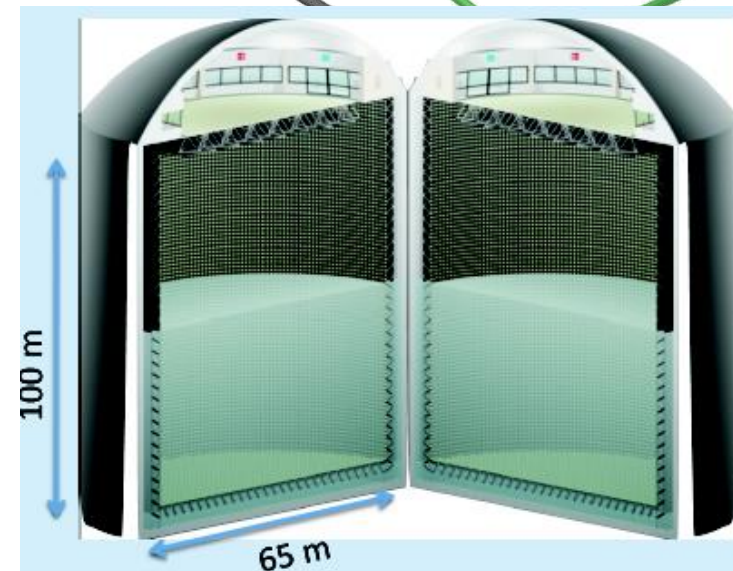
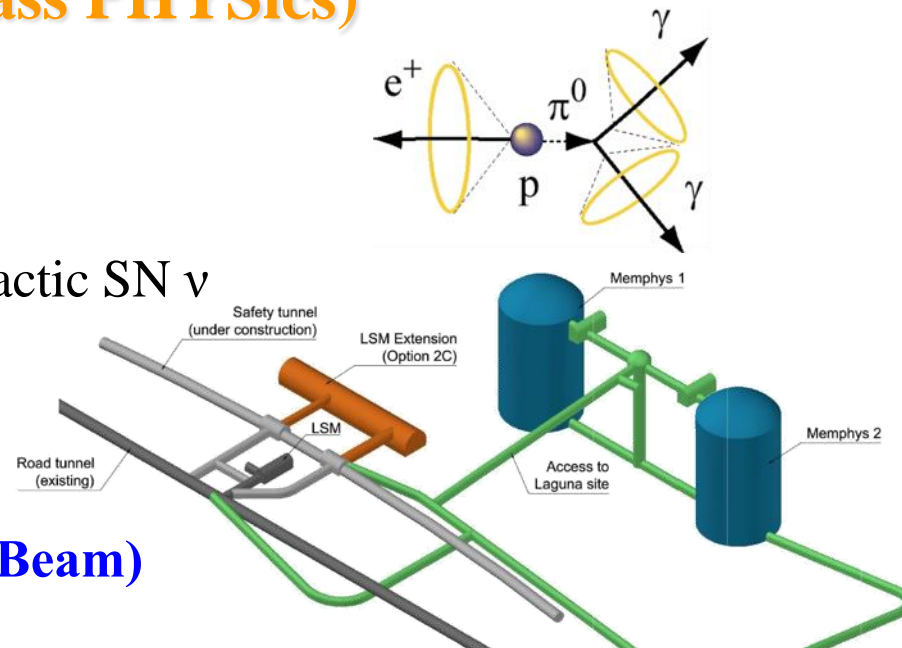
The MEMPHYS WC Detector

(MEgaton Mass PHYSics)

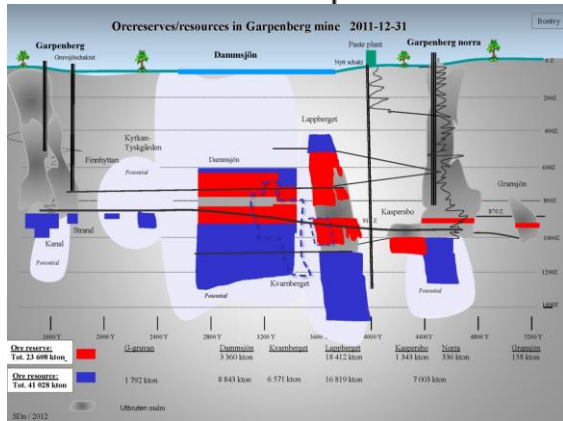
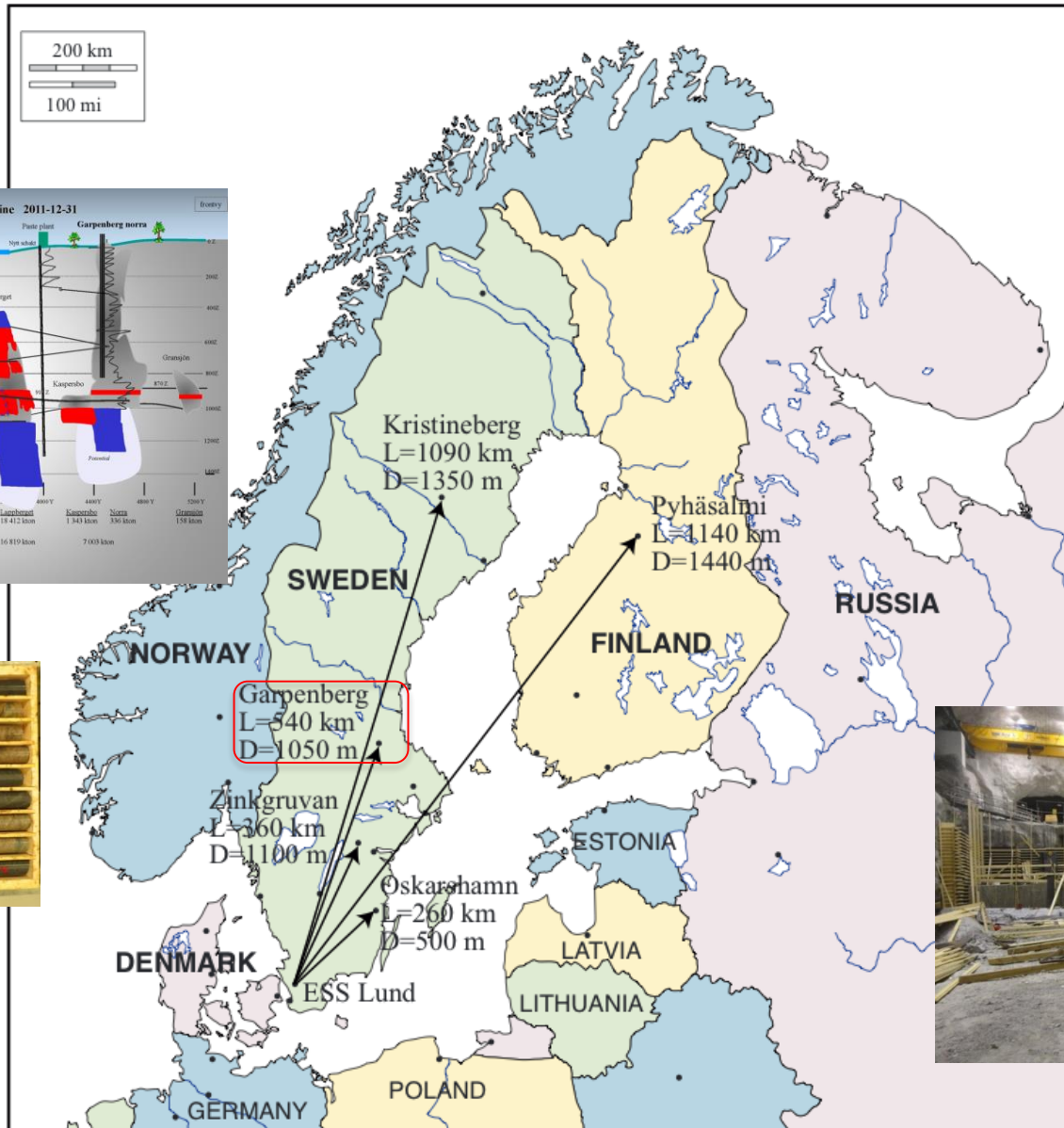
- Proton decay
- Astroparticles
- Understand the gravitational collapsing: galactic SN ν
- Supernovae "relics"
- Solar Neutrinos
- Neutrino Oscillations (Super Beam, Beta Beam)
- Atmospheric Neutrinos

- 500 kt fiducial volume ($\sim 20 \times$ SuperK)
- Readout: $\sim 240k$ 8" PMTs
- 30% optical coverage

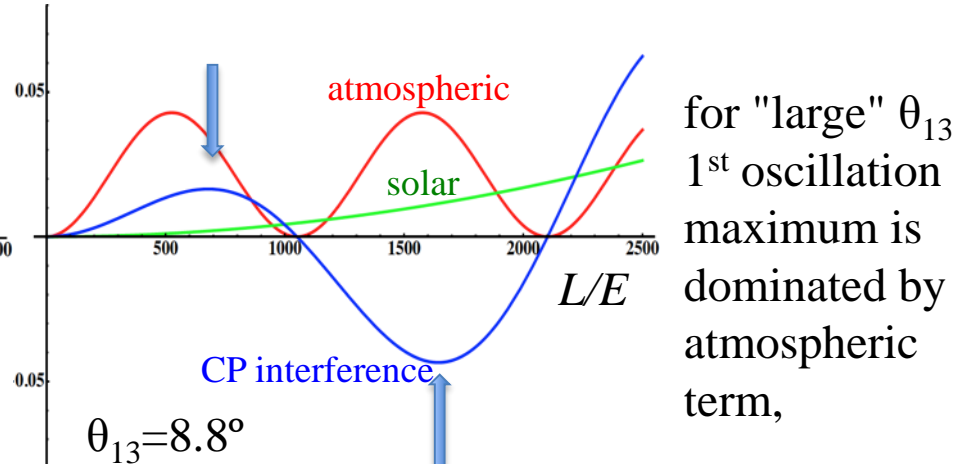
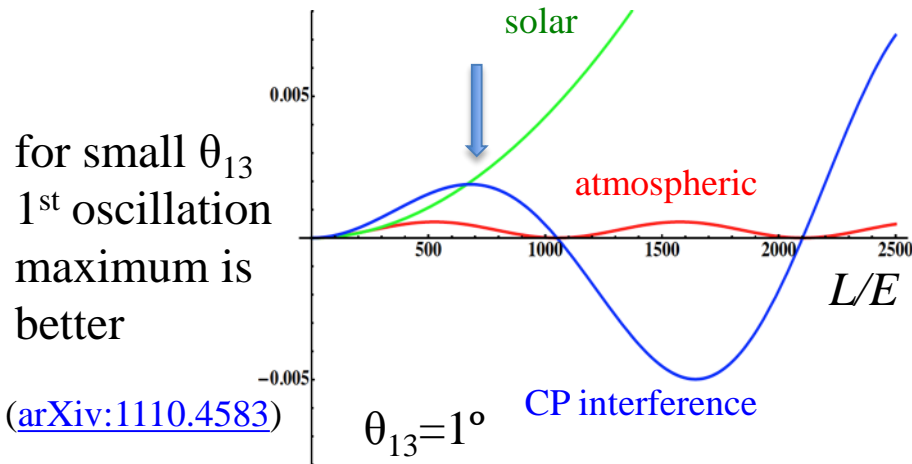
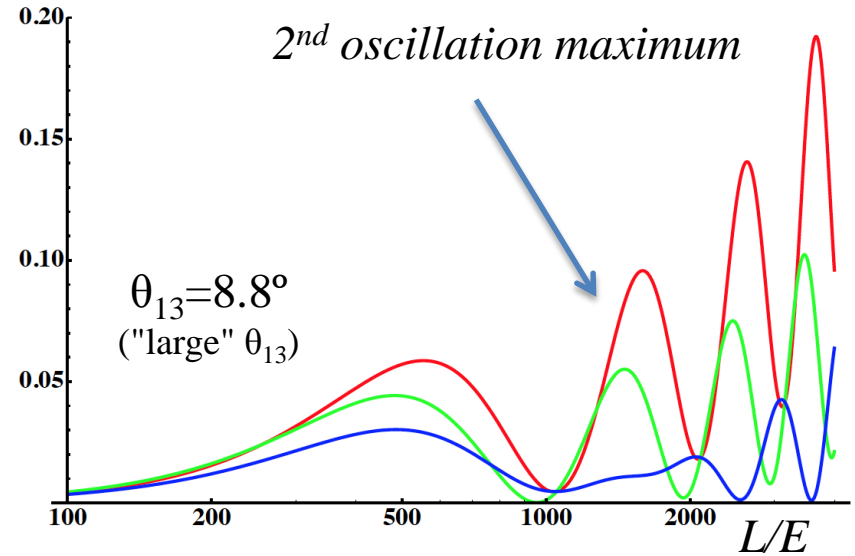
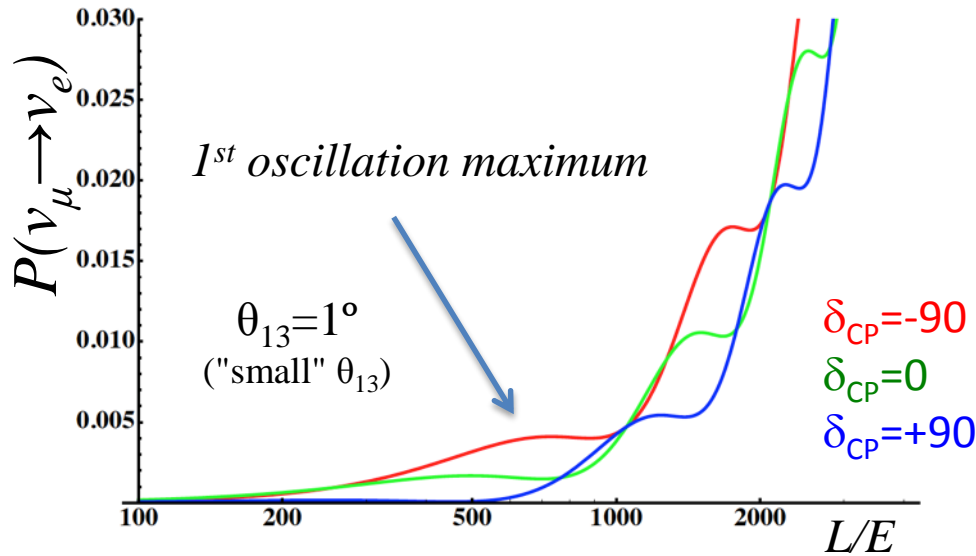
(arXiv: hep-ex/0607026)



Possible locations for far detector



Neutrino Oscillations with "large" θ_{13}



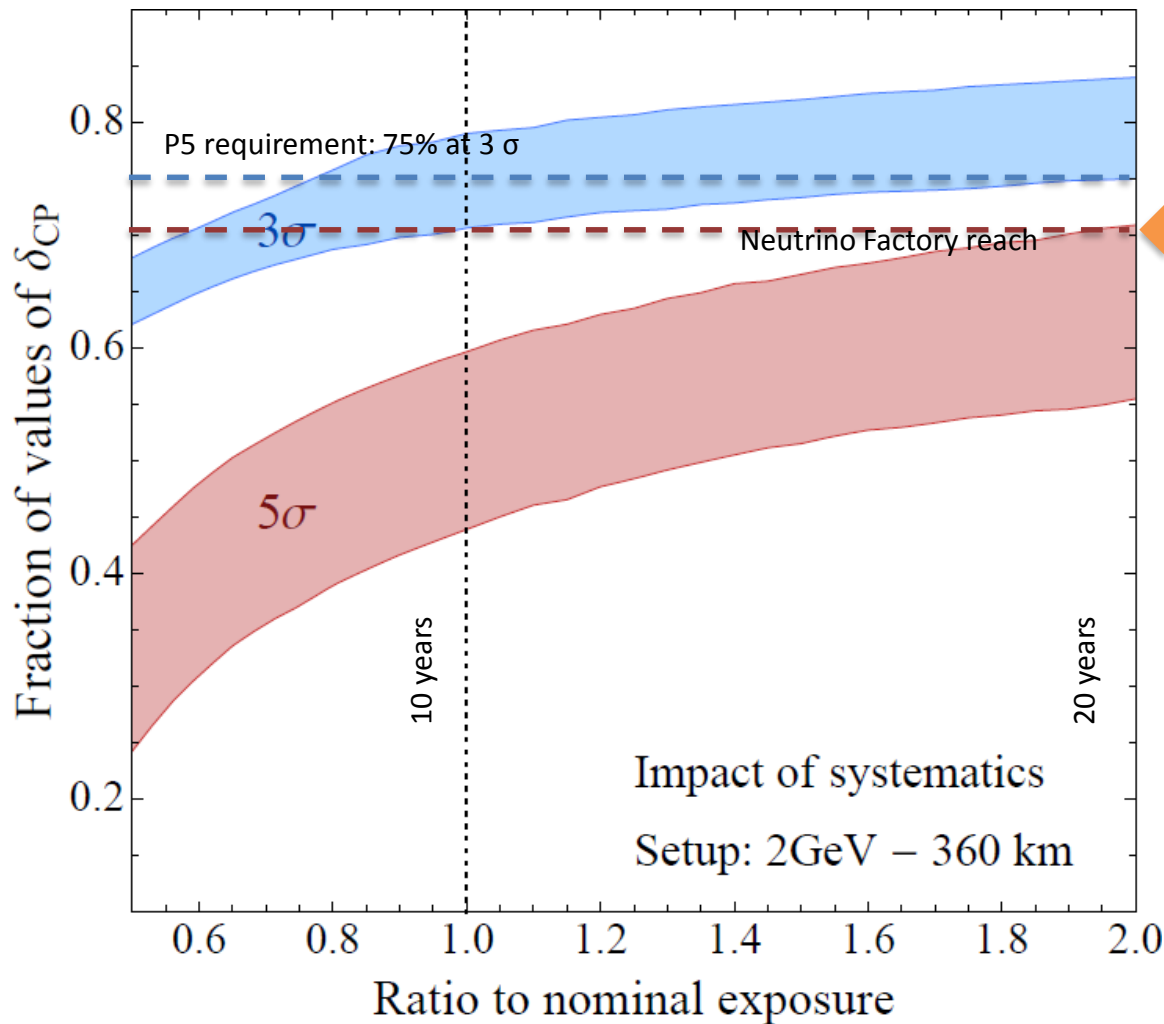
- 1^{st} oscillation max.: $A=0.3\sin\delta_{CP}$
- 2^{nd} oscillation max.: $A=0.75\sin\delta_{CP}$



more sensitivity at 2^{nd} oscillation max.
(see [arXiv:1310.5992](https://arxiv.org/abs/1310.5992) and [arXiv:0710.0554](https://arxiv.org/abs/0710.0554))

Systematic errors and exposure

for ESSnuSB systematic errors see 1209.5973 [hep-ph] (lower limit "default" case, upper limit "optimistic" case)



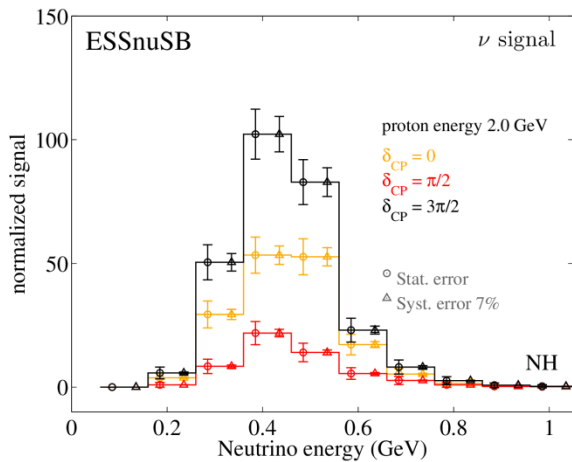
High potentiality

60% of the δ_{CP} range could be reached to at least 5σ

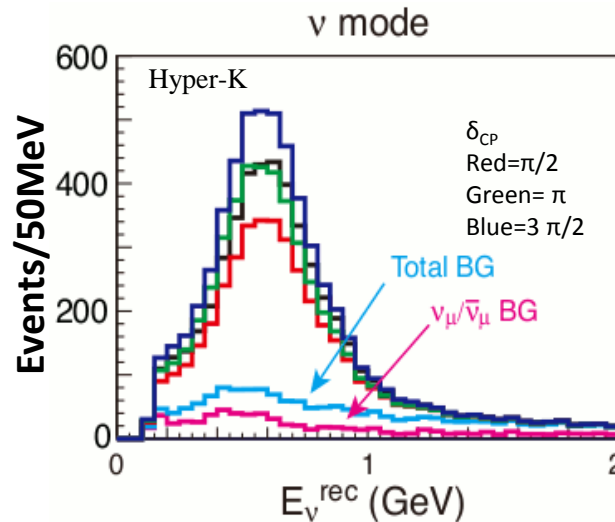
(courtesy P. Coloma)

The sensitivity of the neutrino energy distribution to δ_{CP}

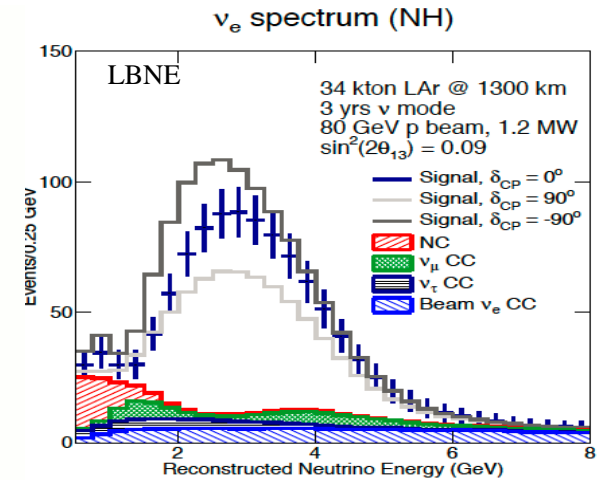
ESSnuSB second maximum



Hyper-K first maximum



LBNE first maximum



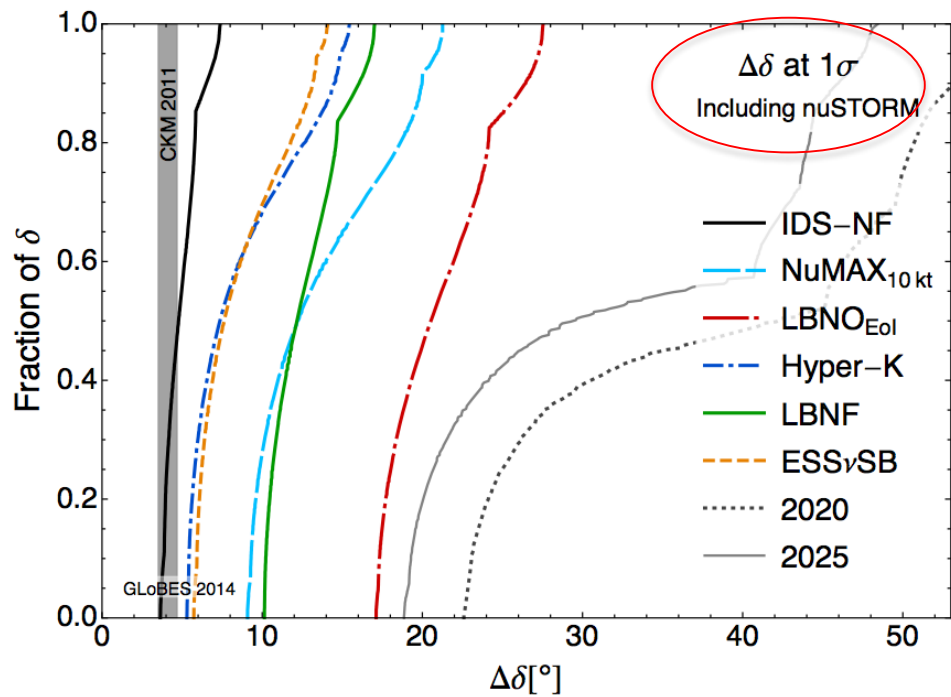
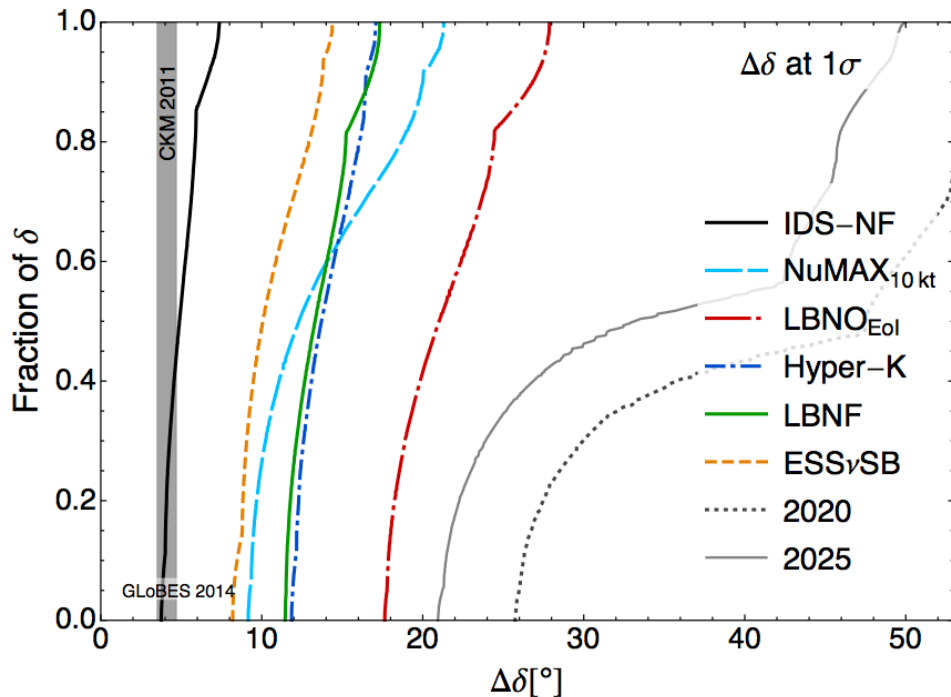
Relative difference in counts at maximum between $\delta_{CP} = 3\pi/2$ and $\pi/2$:

$$105/22 = 4.8$$

$$510/340 = 1.5$$

$$110/65 = 1.5$$

Results including nuSTORM



LBNF (40 kton detector, 1.2MW beam power)

Pilar Coloma et al.

Systematics are basically the same as before. nuSTORM plots assume that cross sections are determined at the 1% level of precision, removing the constraint between the cross sections for different flavors, they are all allowed to vary independently in the fit.

ESS Neutrino Super Beam



Available online at www.sciencedirect.com

ScienceDirect



Nuclear Physics B 885 (2014) 127–149

www.elsevier.com/locate/nuclphysb

arXiv:1212.5048

arXiv:1309.7022

A very intense neutrino super beam experiment for leptonic CP violation discovery based on the European spallation source linac

14 participating institutes
from 10 different countries,
among them ESS and CERN

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H. Danared^g, T. Davenne^c, C. Densham^c, M. Dracos^{m,*}, T. Ekelöf^{n,*},
M. Eshraqi^g, E. Fernandez Martinez^h, G. Gaudiot^m, R. Hall-Wilton^g,
J.-P. Koutchouk^{n,d}, M. Lindroos^g, P. Loveridge^c, R. Matev^k,
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ⁿ Department of Physics and Astronomy, Uppsala University, Box 516, SE-75120 Uppsala, Sweden

EU H2020 Design Study application
has been submitted recently !

When and to what price ?

- ✓ Total price of ESSnuSB including the detector is 1.2 BEUR:
 - ✓ 100 MEUR linac
 - ✓ 200 MEUR accumulator
 - ✓ 200 MEUR target station
 - ✓ 700 MEUR the far detector
- ✓ If we have our CDR in 2018 and if we convince everybody to build this facility, we could start construction at the moment when the neutron facility will be ready, i.e., 2023. The construction of the neutrino facility could be completed 2030 when we will be able to start data taking.

Conclusions

- CP Violation intense neutrino beams are needed
 - Profit of the worlds most powerful linac ar ESS
- Better CPV sensitivity at the 2nd oscillation maximum (theta13 rel. large)
 - Mines available at a good distance
- ESS will have enough protons to go to the 2nd oscillation maximum increasing the sensitivity to CPV.
- The European Spallation Source Linac will be ready in less than 10 years
- CPV: 5 σ could be reached over 60% of δ_{CP} range (ESSvSB)
- Full complementarity with a long baseline experiment on the 1st oscillation maximum using a LAr detector.
- A feasibility Design Study for ESSnuSB has now to be made
 - Collaborators are welcome: experiment, detector, physics, accelerator