Two-loop snail diagrams: relating neutrino masses to dark matter

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■ Neutrino mass<< Electron mass</p>

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Who cares?!

Neutrino mass<< Electron mass

Who cares?!

Let's try to explain

new mass scale

Seesaw type I

$$\begin{bmatrix} 0 & m_D \\ m_D & m_M \end{bmatrix}$$

 $m_D \ll m_M$



new mass scale

Seesaw type I

$$\begin{bmatrix} 0 & m_D \\ m_D & m_M \end{bmatrix}$$

$$m_D \ll m_M$$



Popularity can change in time

LHC and other upcoming experiments

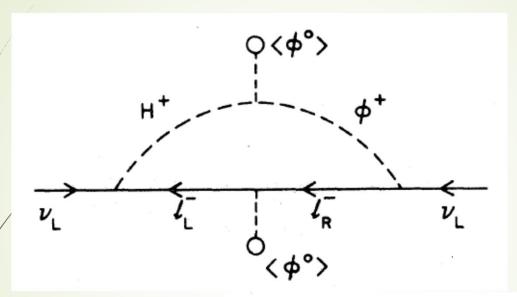


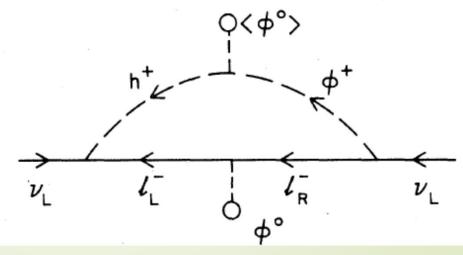
Testable models with low scale mass

Loop suppression

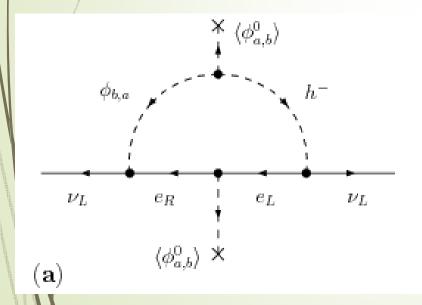
- Cheng-Li, PRD 1980
- Zee model, NPB 1986
- Babu model, PLB, 1988
-

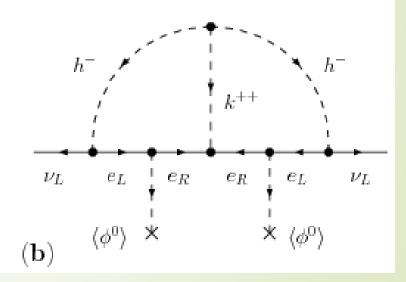
Cheng-Li





Zee-Babu





Radiative neutrino mass production

 Symmetries to forbid lower order (and therefore dominant) loop contribution

If SM fields are invariant under these symmetries, lightest new particle with non-trivial behavior under the symmetry can be stable.

Dark matter

Ma's Scotogenic model and our SLIM scenario

E. Ma, Phys. Rev. **D73**, 077301 (2006), hep-ph/0601225.

C. Boehm, Y. Farzan, T. Hambye, S. Palomares-Ruiz, and S. Pascoli, Phys. Rev. D77, 043516 (2008), hep-ph/0612228.

Proliferation of radiative neutrino mass models in literature

Recipes and Ingredients for Neutrino Mass at Loop Level

Y.F., S. Pascoli and T. Schmidt, JHEP 1303 (2013) 107

Weinberg operator

effective dimension 5 operator, (HL)(HL)

$$(H^{\dagger}H)^{m}(HL)(HL)$$

Weinberg operator

effective dimension 5 operator, (HL)(HL)

$$\mathcal{O}(5) \sim L^T c(i\tau_2) H H^T (i\tau_2) L$$

$$(H^{\dagger}H)^{m}(HL)(HL)$$

General n-loop contribution

$$m_{\nu} \sim \left(\frac{g^2}{16\pi^2}\right)^n \left(\frac{\langle H \rangle^2}{m_{\text{New}}}\right) \left[1, \left(\log \frac{\Lambda}{m_{\text{New}}}\right)^n\right]$$

 Λ is the ultraviolet (UV) cut-off scale of the model satisfying $\Lambda \gg m_{\text{New}}$.

$$m_{\text{New}} \sim 1 \text{ TeV}, m_{\nu} \sim 0.1 - 1 \text{ eV}$$

 $\Lambda/m_{\text{New}} \sim 10 \text{ and } n = 2,$

$$\int_{g} 10^{-3}.$$

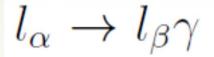
General n-loop contribution

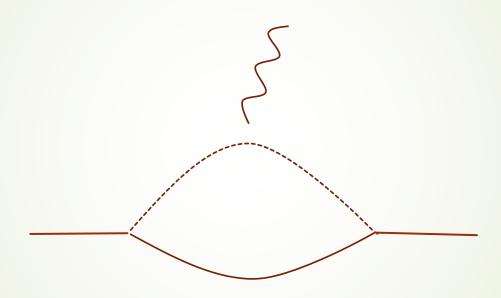
$$m_{\nu} \sim \left(\frac{g^2}{16\pi^2}\right)^n \left(\frac{\langle H \rangle^2}{m_{\text{New}}}\right) \left[1, \left(\log \frac{\Lambda}{m_{\text{New}}}\right)^n\right]$$

$$g \mathbb{Z}$$

$$n=3$$

$$g \sim 0.01 - 0.1$$





$$\Gamma(l_{\alpha} \to l_{\beta} \gamma) \sim \frac{g_{\alpha}^2 g_{\beta}^2}{16\pi (16\pi^2)^2} \frac{m_{\alpha}^5}{Max[m_S^4, m_F^4]}$$

$$Br(\mu \to e\gamma) < 5.7 \times 10^{-13} ,$$

$$Br(\tau \to e\gamma) < 3.3 \times 10^{-8}$$

$$Br(\tau \to \mu\gamma) < 4.4 \times 10^{-8}$$

$$g_e g_\mu \lesssim 10^{-3} \frac{\text{Max}(m_S^2, m_{F_1^-}^2)}{\text{TeV}^2}$$

$$g_e g_\tau, g_\mu g_\tau \lesssim \frac{\operatorname{Max}(m_S^2, m_{F_1^-}^2)}{\operatorname{TeV}^2}.$$

$$Br(\mu \to e\gamma) < 5.7 \times 10^{-13} ,$$

$$Br(\tau \to e\gamma) < 3.3 \times 10^{-8}$$

$$Br(\tau \to \mu\gamma) < 4.4 \times 10^{-8}$$

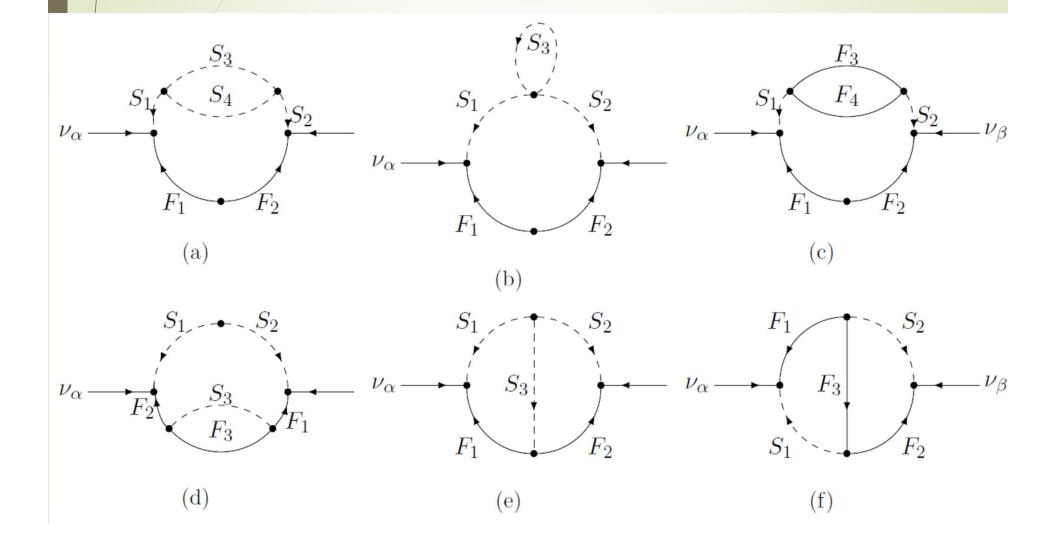
$$g_e g_\mu \lesssim 10^{-3} \frac{\text{Max}(m_S^2, m_{F_1^-}^2)}{\text{TeV}^2}$$

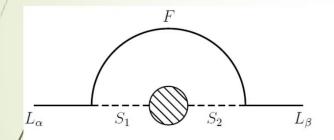
$$g_e g_\tau, g_\mu g_\tau \lesssim \frac{\operatorname{Max}(m_S^2, m_{F_1^-}^2)}{\operatorname{TeV}^2}.$$

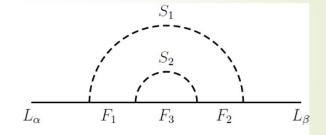
If neutrino mass in obtained at two-loop level, these bounds are Readily satisfied.

Two-loop might be preferred.

 See however, Ahriche, McDonald and Nasri, 1505.04320 which advocates three-loop.

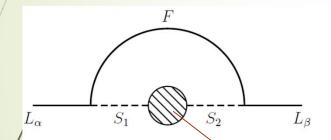


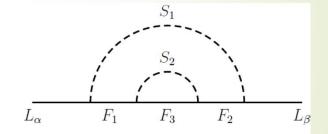




SSHH

FFHH

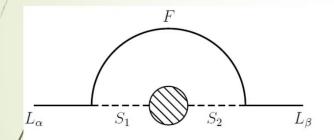


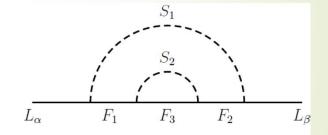


SSHH

FFHH

If there is no symmetry to forbid the bulb, the vertex can also exist.

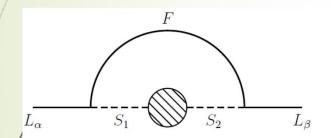


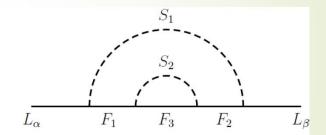


SSHH

FFHH

Accompanied by a one-loop dominant contribution



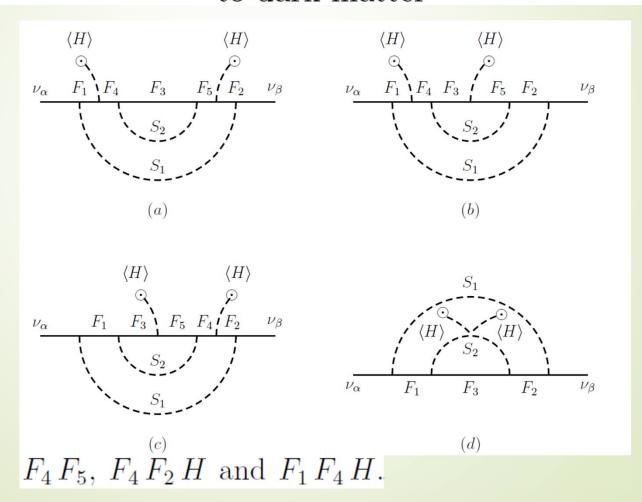


SSHH

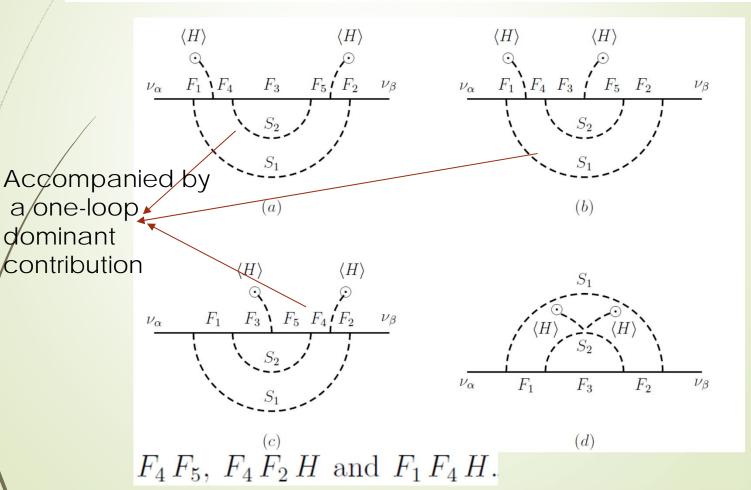
FFHH

Leading loop contribution

Two-loop snail diagrams: relating neutrino masses to dark matter

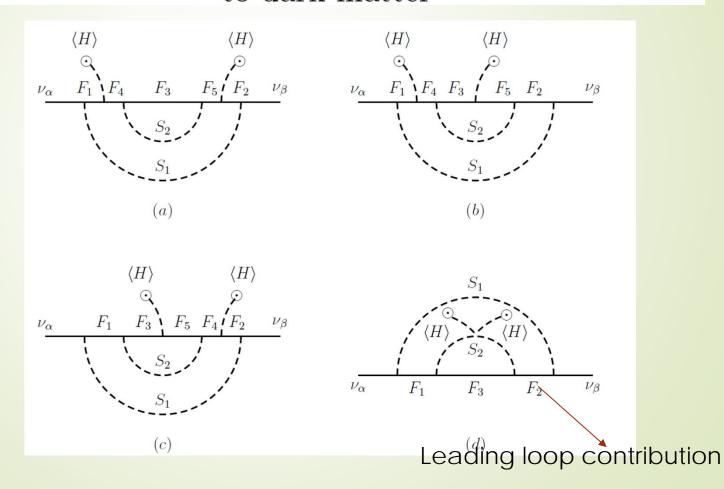


Two-loop snail diagrams: relating neutrino masses to dark matter



dominant

Two-loop snail diagrams: relating neutrino masses to dark matter



Two-loop snail diagrams: relating neutrino masses to dark matter

JHEP 05 (2015) 029

	SU(2)	$U(1)_Y$	$U(1)_L$	$U(1)_{NEW}$	Z_2
F_1	d	-1	1	1	+
F_2	d	- 1	1	-1	+
F_3	d	1	1	1	+
ψ	\mathbf{S}	0	1	1	-
S	S	0	0	-1	+
Φ	d	-1	0	0	-
Φ'	d	-1	0	-1	-

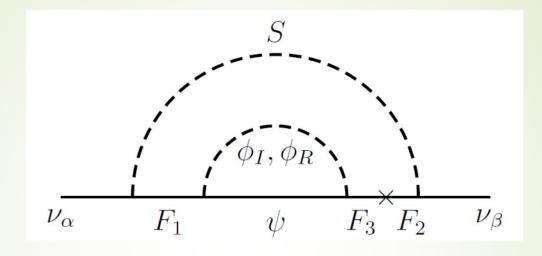
 $m_M(F_{2R}^a)^T c F_{3R}^b \epsilon_{ab} + m'_M(F_{2L}^a)^T c F_{3L}^b \epsilon_{ab} + \text{H.c.}$

$$\mathcal{L}_{Yukawa} = g_{\alpha}S^{\dagger}F_{1R}^{\dagger}L_{\alpha} + h_{\alpha}SF_{2R}^{\dagger}L_{\alpha} + Y_{R\alpha}\Phi'^{\dagger}\psi_{R}^{\dagger}L_{\alpha} +$$

$$Y_{1}\Phi^{\dagger}\psi_{L}^{\dagger}F_{1R} + Y_{2}\epsilon_{ab}\Phi^{a}\psi_{L}^{\dagger}F_{3R}^{b} + Y_{1}'\Phi^{\dagger}\psi_{R}^{\dagger}F_{1L} + Y_{2}'\epsilon_{ab}\Phi^{a}\psi_{R}^{\dagger}F_{3L}^{b} + \text{H.c.}$$

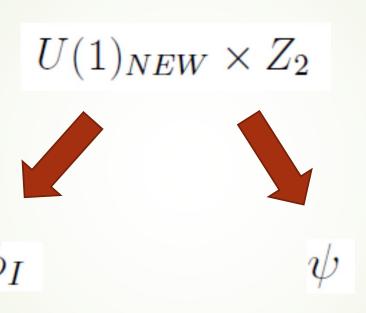
$$(\lambda (H^a \Phi^b \epsilon_{ab})^2 + \text{H.c})$$
 and $\lambda' |H^{\dagger} \Phi|^2$.

$$\Phi^0 \equiv (\phi_R + i\phi_I)/\sqrt{2}. \qquad m_R^2 - m_I^2 = \lambda \langle H^0 \rangle^2.$$



$$m_{\nu} \sim (0.01-0.1~{\rm eV}) Y_1 Y_2 \frac{g \times h}{10^{-1} \times 10^{-2}} \frac{m_M}{5~{\rm GeV}} \frac{(m_R^2 - m_I^2)/m_{max}^2}{1/20}.$$

Dark matter candidate



Thermal freeze-out scenario

$$\Omega_{DM} = \frac{nm}{\rho_c} \propto \frac{m/T_f}{\langle \sigma v \rangle}$$

$$\langle \sigma_{tot} v \rangle = 3 \times 10^{-26} \text{ cm}^3 \text{sec}^{-1}$$

Coannihilation

$$m_R^2 - m_I^2 = \lambda \langle H^0 \rangle^2$$

 $(m_R - m_I)/m_R$ should be smaller than ~ 0.05

$$\phi_I \phi_R \to Z^* \to SM$$

$$\langle \sigma(\phi_I + \phi_R \to Z^* \to f + \bar{f})v \rangle = \frac{16}{3\pi} N_C G_F^2 \frac{(a_L^2 + a_R^2)(m_I v)^2}{(1 - 4m_I^2/m_Z^2)^2}$$

Main dark matter component

$$\langle \sigma(\psi \bar{\psi} \to \ell_{\alpha} \bar{\ell}_{\alpha}) v \rangle = \frac{|Y_{R\alpha}|^4}{32\pi} \frac{m_{\psi}^2}{(m_{\psi}^2 + (m_{\phi'}^2)^2)^2}.$$

$$\langle \sigma_{tot} v \rangle = 3 \times 10^{-26} \text{ cm}^3 \text{sec}^{-1}$$

$$m_{\phi'^{-}}, \ m_{\phi'^{0}} \le 1.4 Y_{R\alpha}^{2} \text{TeV}$$

LHC signals

- Mono-lepton plus missing energy signal through $u\bar{d} \to \phi'^+\phi'^0 \to (l^+\psi)(\nu\bar{\psi})$ and the charge conjugate processes.
- Two-lepton plus missing energy signal through $u\bar{u}, d\bar{d} \to \phi'^+\phi'^- \to (l^+\psi)(l^-\bar{\psi})$.
- Missing energy through $u\bar{u}, d\bar{d} \to \phi'^0 \bar{\phi}'^0 \to (\bar{\nu}\psi)(\nu\bar{\psi}).$

G. Aad et al. [ATLAS Collaboration], JHEP **1405**, 071 (2014)

G. Aad et al. [ATLAS Collaboration], JHEP **1410**, 96 (2014)

Muon and electron: >325 GeV

Tau: >90 GeV

ILC signals

$$g_{\alpha}S^{\dagger}\bar{F}_{1R}L_{\alpha} + h_{\alpha}S\bar{F}_{2R}L_{\alpha}$$

$$e^-e^+ \to S\bar{S}$$
.

$$\Gamma(S \to l_{\alpha}^- F_1^+) \propto g_{\alpha}^2$$
 and $\Gamma(S \to l_{\alpha}^+ F_{2,3}^-) \propto h_{\alpha}^2$

 F_i^+ can decay into $\phi^+\psi$ and $\phi^+\to (W^+)^*\phi_I\to \nu l^+\phi_I, q\bar{q}\phi_I$.

$$\Gamma(S \to l_{\alpha}^- F_1^+) \propto g_{\alpha}^2 \text{ and } \Gamma(S \to l_{\alpha}^+ F_{2,3}^-) \propto h_{\alpha}^2$$
.

$$e^+e^- \rightarrow l_{\alpha}^+ + l_{\beta}^- + l_{\gamma}^+ + l_{\theta}^- + \text{missing energy};$$

 $e^+e^- \rightarrow l_{\alpha}^+ + l_{\beta}^- + l_{\gamma}^+ + 2 \text{ jets} + + \text{missing energy};$
 $e^+e^- \rightarrow l_{\alpha}^+ + l_{\beta}^- + l_{\gamma}^- + 2 \text{ jets} + \text{missing energy};$
 $e^+e^- \rightarrow l_{\alpha}^+ + l_{\beta}^- + 4 \text{ jets} + \text{missing energy},$

$$h_{\alpha}^2 h_{\beta}^2$$

$$g_{\alpha}^2 g_{\beta}^2$$

$$\Gamma(S \to l_{\alpha}^- F_1^+) \propto g_{\alpha}^2 \quad \Gamma(\bar{S} \to l_{\beta}^- F_{2,3}^+) \propto h_{\beta}^2$$

 $e^+e^- \rightarrow l_{\alpha}^- + l_{\beta}^- + 4 \text{ jets} + \text{missing energy}$

Summary

We presented a model that contribute to neutrino mass via "two-loop snail diagram."

Phenomenological consequences are rich.

$$\langle \sigma(\psi \bar{\psi} \to l \bar{l}) v \rangle = 0.86 \times 10^{-26} \text{ cm}^3 \text{sec}^{-1} \text{ and } m_{\psi} \sim 10 \text{ GeV},$$

$$Y_R = 0.5(m_{\phi'^-}/100 \text{ GeV})(10 \text{ GeV}/m_{\psi})^{1/2}$$

taking $m_I < m_W$, we find that $\langle \sigma_{\rm tot} v \rangle \sim 40 \ {\rm pb}(m_I/70 \ {\rm GeV})^2$

$$/\mathcal{O}(1 \ pb/\langle \sigma_{\rm tot} v \rangle) \sim 2.5\% (70 \ {\rm GeV}/m_I)^2$$

For $m_W < \phi_I < 200 \,\mathrm{GeV}$, new annihilation modes will open.

$$G_F^2 m_Z^2 / m_{\phi_I}^2$$

the λ' coupling $\phi_I \phi_I$ and $\phi_R \phi_R$ pair annihilation

$$m_I \sim \text{few TeV}$$
.

$$Y_1, Y_2 \gg g, h$$
.

$$F_i^- \to \psi \phi^-$$
 and $F_i^0 \to \psi \phi_{I(R)}^0$

Via tree-level Z^* exchange, $\phi_R \to \phi_I \nu \bar{\nu}$, $\phi_I l l$ and $\phi_I q \bar{q}$.

$$e^+e^- \to l_{\alpha}^- + l_{\beta}^- + 4$$
 jets + missing energy .

In general by studying these modes information on flavor structure of g_{α} and h_{α} can be extracted and cross checked against the information from neutrino mass matrix and LFV.

Notice that the signature of the present model at colliders is completely different from those of SLIM model or of Ma's Scotogenic model [5–13] which both lack doublet fermions. The present model is also distinguishable from models in which neutrino mass is produced via a one loop diagram in which a fermion doublet and scalar singlet propagate as in such a model fermion doublet will decay into leptons rather than $\psi\phi$.