

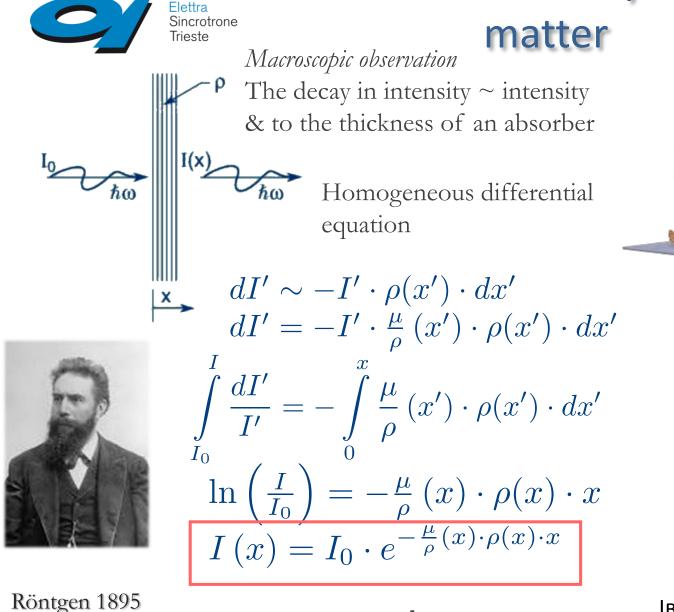
Elettra Sincrotrone Trieste

Introduction to Basics on radiation probing and imaging using x-ray detectors

Ralf Hendrik Menk Elettra Sincrotrone Trieste INFN Trieste

Part 2

Interaction x-rays with



Lamb - Beer's law

Іва́н Па́влович Пулю́й Ivan Pavlovich Puluj 1894



 $I \sim N$

 $N \sim \varphi$

 $\dim \left|\frac{\mu}{\rho}\right| = \frac{cm^2}{qr}$

Interaction x-rays with matter

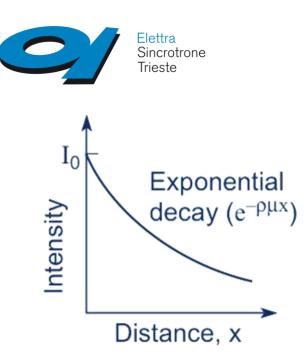
Particle Wave Dualism

 $I = \overrightarrow{E} \cdot \overrightarrow{E^*}$

Wave field

Numbers of measured photons

Flux (photons/time) fluence (photons/ area)



Interaction x-rays with matter

Lamb-Beer Law for one material and one photon energy $I(x) = I_0 \cdot e^{-\left(\frac{\mu}{\rho}\right) \cdot \rho(x) \cdot x}$

$$\frac{I(x)}{I_0} = e^{-\mu \cdot x}$$

 $I(x) \rightarrow 0$ for $x \rightarrow \infty$ Most absorption in the first layers

$$\frac{I(x)}{I_0} = e^{-\mu \cdot x} = e^{-\frac{x}{\lambda}}$$

 λ is called "mean free path"

$$-ln\left(\frac{1}{2}\right) = \frac{x}{\lambda} \qquad \lambda \cdot ln\left(2\right) = x_{1/2}$$

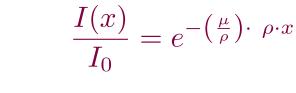


ħω

Interaction x-rays with matter

Lamb-Beer Law for one material and one photon energy

$$I(x) = I_0 \cdot e^{-\left(\frac{\mu}{\rho}\right) \cdot \rho(x) \cdot x}$$

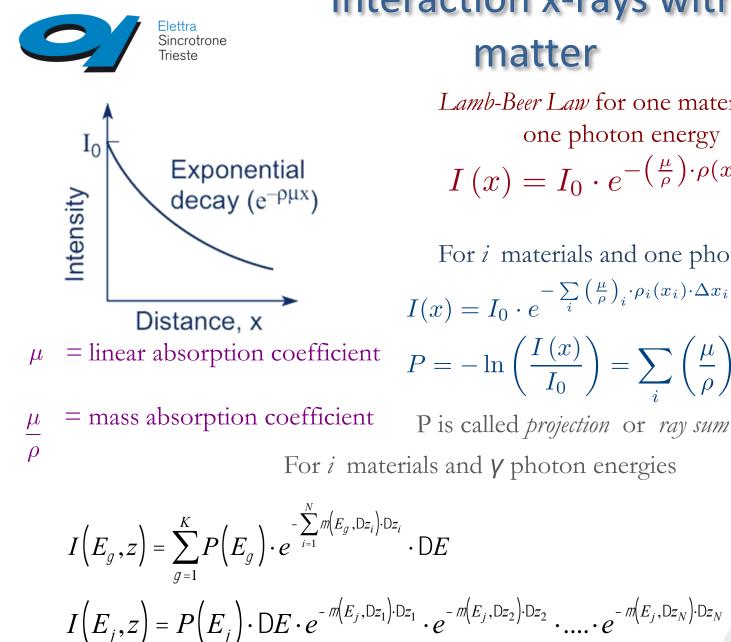


L(II)

 $\frac{I(x)}{I_0} = e^{-\left(\frac{\mu}{\rho}\right) \cdot \rho \cdot x}$ is called "**transmission**" This is the fraction of photons not having any interaction in the sample

$$1 - \frac{I(x)}{I_0} = 1 - e^{-\left(\frac{\mu}{\rho}\right) \cdot \rho \cdot x}$$

is called "absorption" This is the fraction of photons having an interaction in the sample



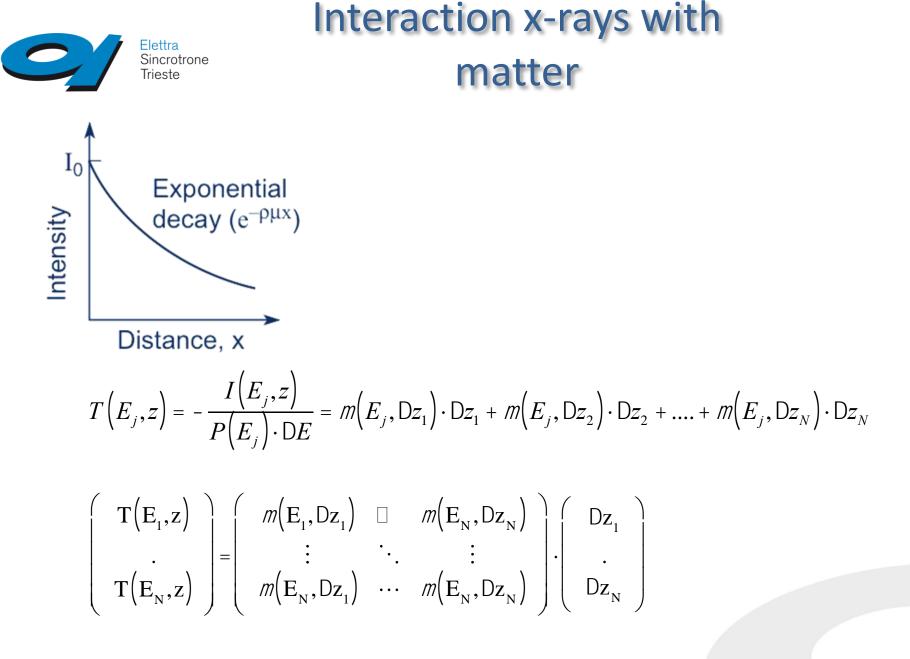
Interaction x-rays with matter

Lamb-Beer Law for one material and one photon energy $I(x) = I_0 \cdot e^{-\left(\frac{\mu}{\rho}\right) \cdot \rho(x) \cdot x}$

For *i* materials and one photon energy

$$I(x) = I_0 \cdot e^{-\sum_i \left(\frac{\mu}{\rho}\right)_i \cdot \rho_i(x_i) \cdot \Delta x_i}$$

$$P = -\ln\left(\frac{I(x)}{I_0}\right) = \sum_i \left(\frac{\mu}{\rho}\right)_i \cdot \rho_i(x_i) \cdot \Delta x_i$$



Spectral imaging



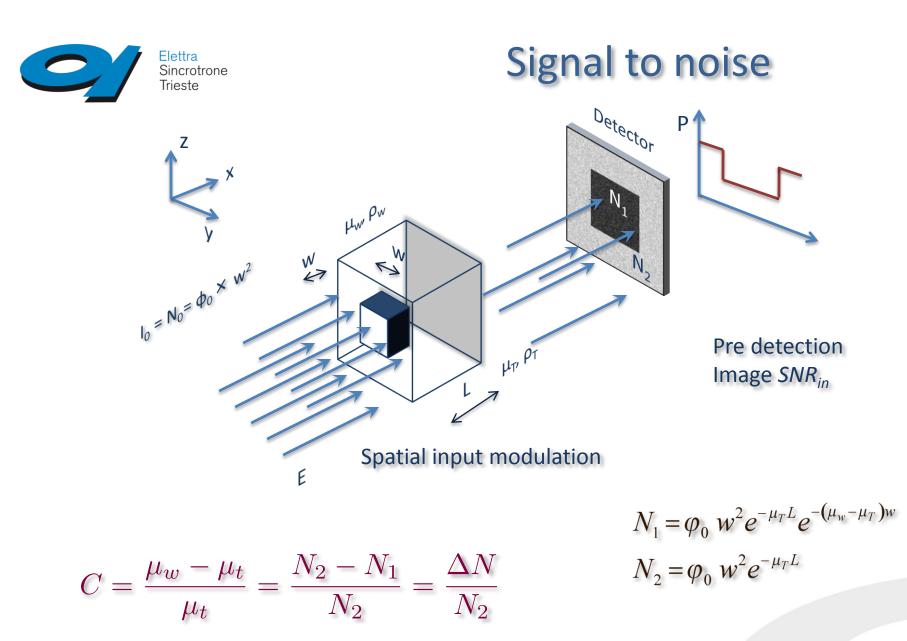
Interaction x-rays with matter

For *i* materials and one photon energy

$$I(x) = I_0 \cdot e^{-\sum_i \left(\frac{\mu}{\rho}\right)_i \cdot \rho_i(x_i) \cdot \Delta x_i}$$
$$P = -\ln\left(\frac{I(x)}{I_0}\right) = \sum_i \left(\frac{\mu}{\rho}\right)_i \cdot \rho_i(x_i) \cdot \Delta x_i$$



For x-rays the order of materials is exchangeable without changing the projection – Principle of Superposition





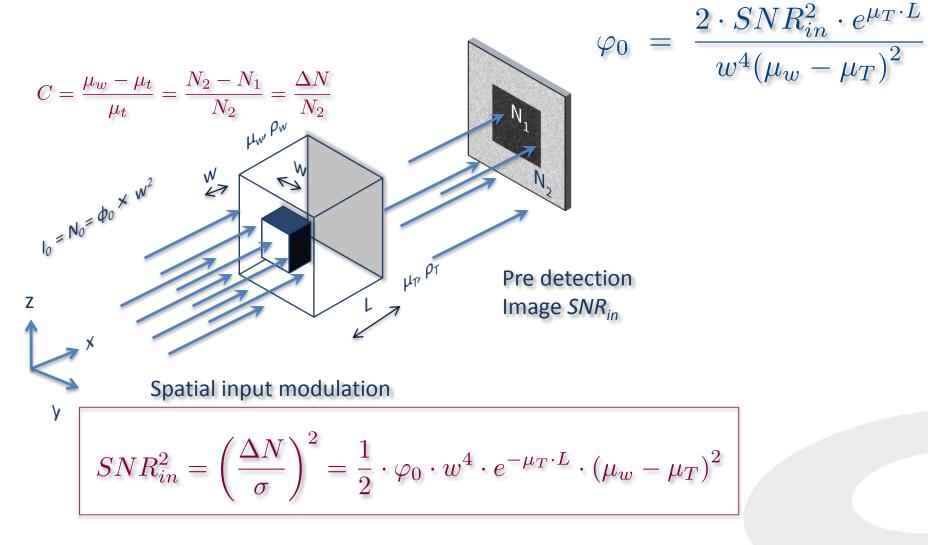


Photon statistics – Poisson noise <N> mean value √N noise $S = \sqrt{\left(\frac{\partial DN}{\partial N}\right)^{2}} \cdot \left(D(DN)\right)^{2} + \left(\frac{\partial DN}{\partial N}\right)^{2} \cdot \left(D(DN)\right)^{2}; \quad D(DN) = \sqrt{DN}$ $= \sqrt{j_{0} w^{2} e^{-m_{T}L} \left(1 + e^{-(m_{w} - m_{T})w}\right)} = \sqrt{j_{0} w^{2} e^{-m_{T}L} \left(2 - (m_{w} - m_{T})w\right)}$ $\approx \sqrt{2j} w^2 e^{-m_T L}$ Signal to noise ratio $SNR_{in}^2 = \left(\frac{\Delta N}{\sigma}\right)^2 = \frac{1}{2} \cdot \varphi_0 \cdot w^4 \cdot e^{-\mu_T \cdot L} \cdot \left(\mu_w - \mu_T\right)^2$ $\varphi_0 = \frac{2 \cdot SNR_{in}^2 \cdot e^{\mu_T \cdot L}}{w^4 (\mu_m - \mu_T)^2} \quad \text{Photon flux}$

 $D_{entrance} = \frac{\Delta E}{\Delta M} = \frac{\Delta E}{\rho \cdot \Delta V} = \frac{\Delta \left(N_0 \cdot E_{\gamma} \cdot \left(1 - e^{-\mu_T \cdot x}\right)\right)}{\rho \cdot A \cdot \Delta x} \sim \frac{\varphi_0 \cdot E_{\gamma}}{\rho}$



Signal to noise



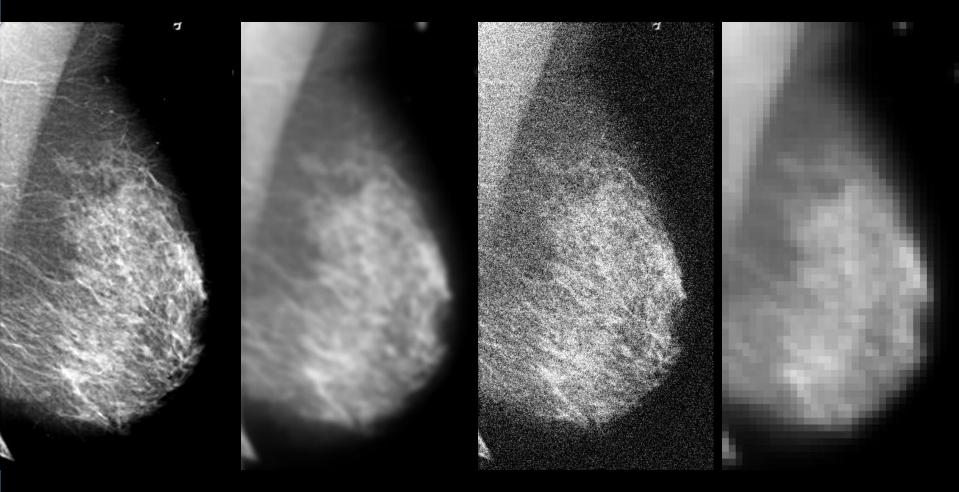
Detective quantum efficiency & Detector problems

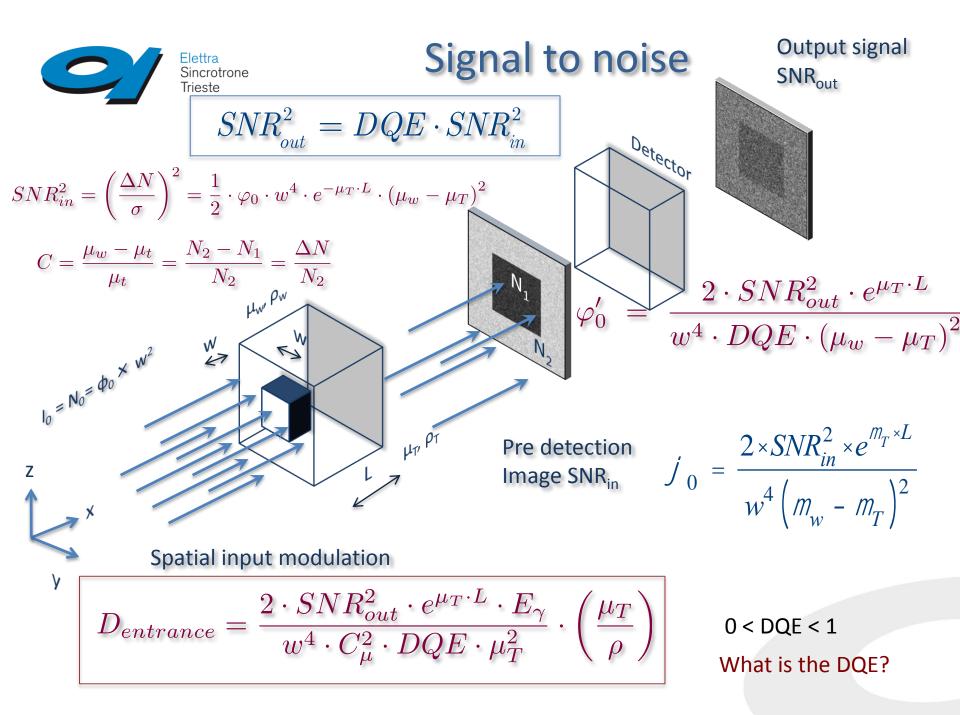
Original image

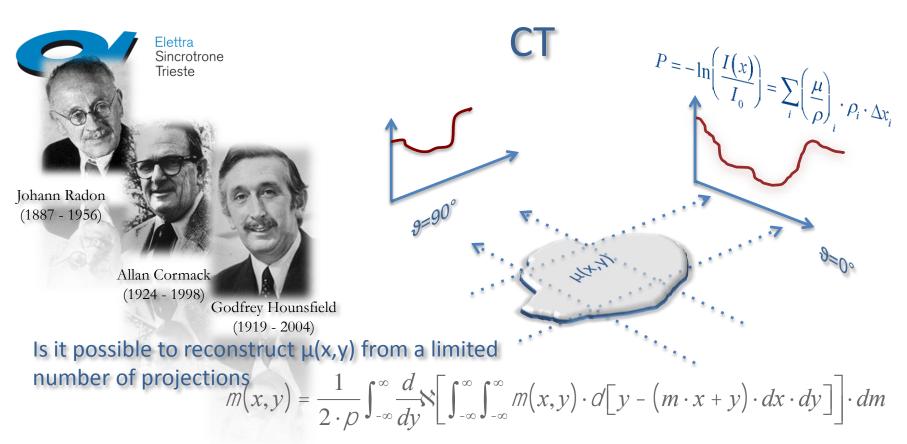
Image blurring

Detector noise

Fixed pattern

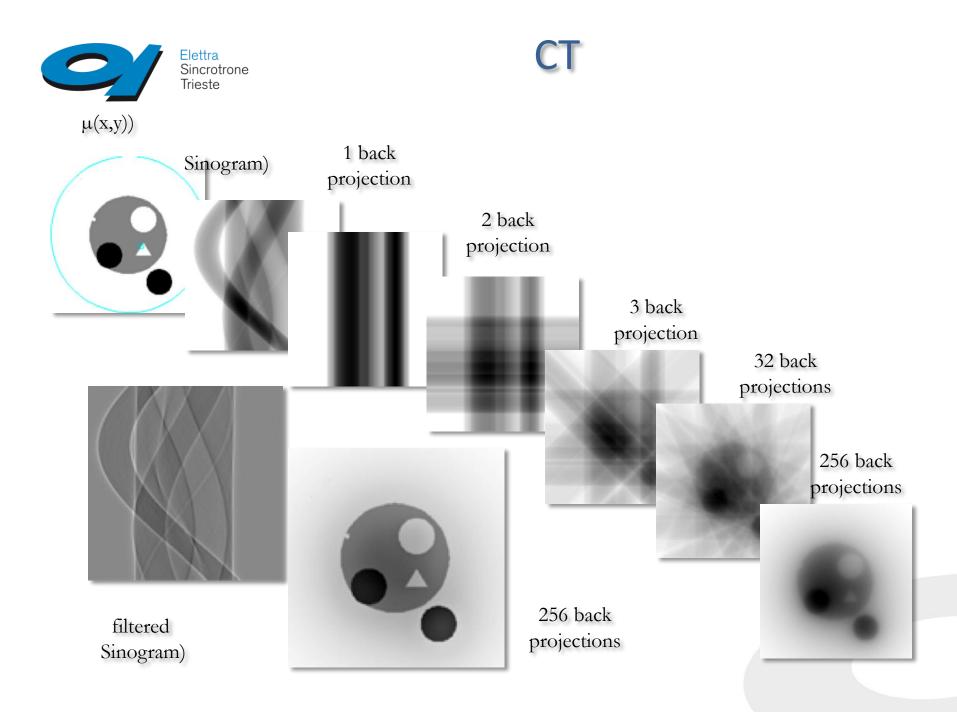






11	15	10				3.66	5	3.33]
3	4	2		10]	2.46			
4	3	4		10			₽ .16 ₽ .16	3.33	
4	8	3		15		3.66	4.10 5	3. 3 9	
]]	4.00	5	9.10	

3.33
3.66
5







$$D_{entrance} = \frac{2 \cdot SNR_{out}^2 \cdot e^{\mu_T \cdot L} \cdot E_{\gamma}}{w^4 \cdot C_{\mu}^2 \cdot DQE \cdot \mu_T^2} \cdot \left(\frac{\mu_T}{\rho}\right)$$

What is the DQE?

0 < DQE < 1