ENHANCING THE EFFICIENCY OF PLASMA WAKEFIELD ACCELERATION

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OUTLINE

- Introduction
- THEORETICAL BASICS
 - Gaussian Pulses
 - Generalization for Bichromatic Fields
 - Equations of Motion
 - The Presence of an Underdense Plasma
- RESULTS
 - General Remarks
 - Monochromatic Fields
 - Comparison with Experimental Data
 - Bichromatic Fields
- SUMMARY

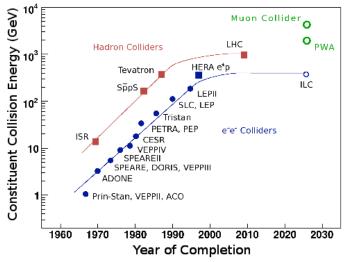


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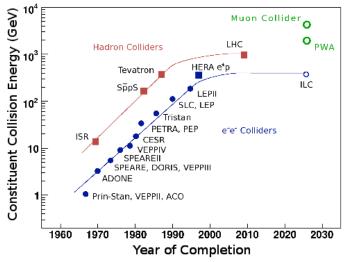
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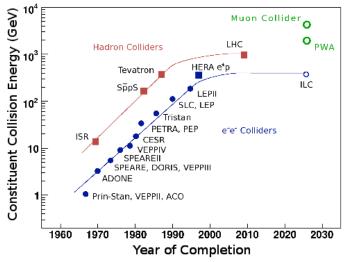
- State-of-the-Art technology: circular accelerator, 8.3 T, 14 TeV
- Limit: accelerating field < 50 MV/m



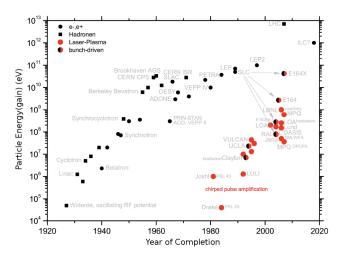
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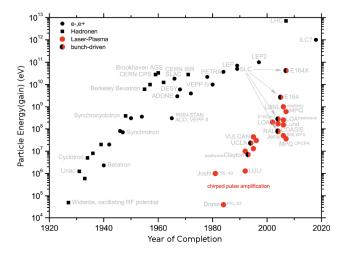


How further → VLHC? Extremely Expensive! (Future Circular Collider Kick Off Meeting, Feb. 2014, Geneva)



Plasma Wakefield Acceleration

How further → VLHC? Extremely Expensive! (Future Circular Collider Kick Off Meeting, Feb. 2014, Geneva) New, cheaper technologies are needed → Plasma based acceleration!



- PWFA: electron/proton bunch drives the wakes.
- **LWFA:** Short (\approx 1 ps), ultra intense $I \ge 10^{18} \mathrm{W \cdot cm^{-2}}$ pulse. $L = c\tau_{D} \approx \lambda_{D} = 2\pi c/\omega_{D}, n = 10^{15} \mathrm{cm^{-3}}.$
- **PBWA:** Two laser pulses, $\omega_1 \omega_2 \sim \omega_p$, $n = 10^{16} 10^{17} \, \mathrm{cm}^{-3}$. An alternative for LWFA.
- **SMLWFA:** LWFA on higher plasma densities. $n=10^{19}\,\mathrm{cm^{-3}}$, $I\approx 10^{19}\,\mathrm{W\cdot cm^{-2}}$, $L>\lambda_p$. The plasma "chops up" the long laser pulse. The length of the equidistently spaced train of smaller pulses mathces the plasma wavelength. This train of pulses resonantly excites the plasma.
- Multiple bunches or pulses: larger amplitude plasma waves.



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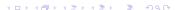
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Short term solution: **PWFA** (CERN AWAKE Experiment)

Long term solution: LWFA



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Plasma Wakefield Acceleration



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Gaussian beams can be derived from the paraxial approximation. For a Gaussian pulse, the electric field has the following form:

$$E_{x} = E_{0} \frac{W_{0}}{W(z)} \exp\left[-\frac{r^{2}}{W^{2}(z)}\right] \exp\left(-\frac{\Theta^{2}}{T^{2}}\right) \times \cos\left[\frac{kr^{2}}{2R(z)} - \Phi(z) + \omega\Theta + \sigma\Theta^{2} + \varphi\right]$$

$$E_{y} = 0$$

$$(1a)$$

$$E_{y} = 0$$

$$(1b)$$

$$E_{z} = -\frac{x}{R(z)}E_{x} + E_{0}\frac{2x}{kW^{2}(z)} \cdot \frac{W_{0}}{W(z)} \exp\left[-\frac{r^{2}}{W^{2}(z)}\right] \times \exp\left[-\frac{\Theta^{2}}{T^{2}}\right] \sin\left[\frac{kr^{2}}{2R(z)} - \Phi(z) + \omega\Theta + \sigma\Theta^{2} + \varphi\right]$$
(1c)

For details, see L.W. Davis: Phys. Rev. A 19 (1979), 1177



Gaussian beams can be derived from the paraxial approximation. For a Gaussian pulse, the magnetic field has the following form:

$$B_{x}=0 (2a)$$

$$B_{y} = \frac{E_{x}}{c} \tag{2b}$$

$$B_{z} = \frac{y}{cR(z)}E_{x} + \frac{1}{c}E_{0}\frac{2y}{kW^{2}(z)} \cdot \frac{W_{0}}{W(z)} \exp\left[-\frac{r^{2}}{W^{2}(z)}\right] \times \exp\left[-\frac{\Theta^{2}}{T^{2}}\right] \sin\left[\frac{kr^{2}}{2R(z)} - \Phi(z) + \omega\Theta + \sigma\Theta^{2} + \varphi\right]$$
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The parameters of the Gaussian pulse are the following:

$$W(z) = W_0 \left[1 + \left(\frac{z}{z_R} \right)^2 \right]^{1/2}$$
 the spot size, (3a)

$$R(z) = z \left[1 + \left(\frac{z_R}{z} \right)^2 \right]$$
 the radius of curvature, (3b)

$$\Phi(z) = \tan^{-1} \frac{z}{z_R} \quad \text{the Gouy phase, and}$$
 (3c)

$$W_0 = \left(\frac{\lambda Z_R}{\pi}\right)^{1/2}$$
 the beam waist. (3d)

and z_R being the Rayleigh-length. The wavenumber has the form of

$$k = \frac{\omega_0}{c} \left(1 + \sigma \Theta \right) \tag{4}$$

with ω_0 being the initial frequency.



At the Rayleigh-length, the area of the beam spot is twice as the minimal size:

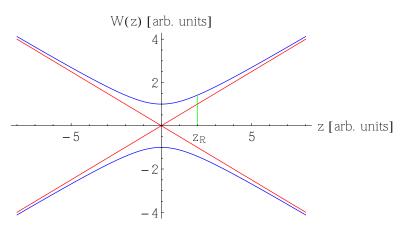
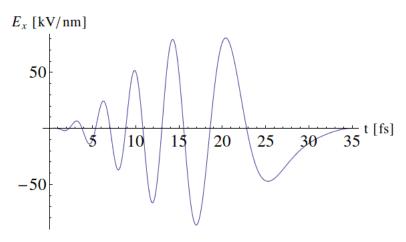


FIGURE: The width of a Gaussian beam as a function of distance along the direction of propagation.

A chirped laser pulse looks like:





- Gaussian beams may have higher order TEM modes! See:
 H. Kogelik and T. Li: Appl. Opt. 5 (1966), 1550.
- For a system with cylindrical symmetry, it is useful to express the beam in cylindrical coordinates. Cylindrical waves also may have higher order TEM modes.
- Is it possible, based on L.W. Davis: Phys. Rev. A 19 (1979), 1177 to obtain an exact solution of Maxwell's equations?
- If so, does it worth the work?
- Exact beam solutions of the Maxwell's equations can be produced using the Hertz-vector! See: P. Varga and P. and Török, Opt. Commun. 152 (1998), 108–118.

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The q^{th} harmonic of a Gaussian pulse has the following form:

$$E_{q,x} = E_0 \frac{W_0}{W_q(z)} \exp\left[-\frac{r^2}{W_q^2(z)}\right] \exp\left[-\frac{\Theta^2}{T^2}\right] \times$$

$$\cos\left[\frac{k_q r^2}{2R_q(z)} - \Phi_q(z) + q\Theta + q^2 \sigma \Theta^2 + \varphi_q\right],$$

$$E_{q,y} = 0,$$

$$E_{q,z} = -\frac{x}{R_q(z)} E_{q,x} +$$

$$\frac{2x}{k_q W_q^2(z)} \cdot \frac{W_0}{W_q(z)} \exp\left[-\frac{r^2}{W_q(z)}\right] \exp\left[-\frac{\Theta^2}{(\omega_0 T)^2}\right] \times$$

$$\sin\left[\frac{k_q r^2}{2R_q(z)} - \Phi_q(z) + q\Theta + q^2 \sigma \Theta_q^2 + \varphi_q\right]$$
(5a)
$$(5b)$$

$$B_{q,x}=0, (6a)$$

$$B_{q,y} = \frac{E_{q,x}}{c},\tag{6b}$$

$$B_{q,z} = \frac{y}{cR_q(z)} E_{q,x} +$$

$$\frac{1}{c}E_0 \frac{2y}{k_q W_q^2(z)} \cdot \frac{W_0}{W_q(z)} \exp\left[-\frac{r^2}{W_q^2(z)}\right] \exp\left[-\frac{\Theta^2}{T^2}\right] \times \tag{6c}$$

$$\sin\left[\frac{k_q r^2}{2R_q(z)} - \Phi_q(z) + q\Theta + q^2 \sigma \Theta^2 + \phi_q\right]$$

The parameters of the higher harmonic are:

$$W_q(z) = W_{q,0} \left[1 + \left(\frac{z}{z_{q,R}} \right)^2 \right]^{1/2}$$
 the spot size, (7a)

$$R_q(z) = z \left[1 + \left(\frac{Z_{q,R}}{z} \right)^2 \right]$$
 the radius of curvature, (7b)

$$\Phi_q(z) = \tan^{-1} \frac{z}{z_{q,R}}$$
 the Gouy phase, and (7c)

$$W_{q,0} = \left(\frac{\lambda Z_{q,R}}{\pi}\right)^{1/2}$$
 the beam waist. (7d)

and $z_{a,R}$ being the Rayleigh-length. The wavenumber has the form of

$$k = q \frac{\omega_0}{c} \left(1 + q \sigma_q \Theta \right) \tag{8}$$

with $q\omega_0$ being the initial frequency of the $q^{ ext{th}}$ harmonic.

Finally, the general form of a bichromatic field:

$$\mathbf{E} = \mathbf{E}_1 + \frac{A}{q} \mathbf{E}_q, \tag{9a}$$

$$\mathbf{B} = \mathbf{B}_1 + \frac{A}{q} \mathbf{B}_q \tag{9b}$$

with \mathbf{E}_1 and \mathbf{B}_1 being the electric and magnetic fields of the main harmonic and $0 \le A \le 1$ the relative amplitude of the harmonics.

A bichromatic pulse looks like:

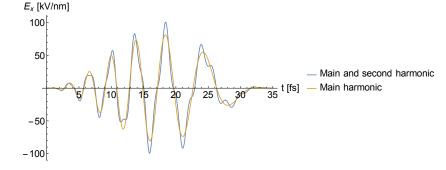


FIGURE: A bichromatic (main and second harmonic) pulse, compared with the corresponding monochromatic (main harmonic) component.

The Lorentz-Force acting on the electron:

$$\mathbf{F} = e(\mathbf{E} + \mathbf{v} \times \mathbf{B}) \tag{10}$$

Equations of Motion for a relativistic electron

$$\frac{\mathrm{d}\gamma}{\mathrm{d}t} = \frac{1}{m_e c^2} \mathbf{F} \cdot \mathbf{v} \tag{11a}$$

$$\frac{\mathrm{d}\mathbf{p}}{\mathrm{d}t} = e\left(\mathbf{E} + \frac{\mathbf{p}}{m_{\mathrm{e}}\gamma} \times \mathbf{B}\right) \tag{11b}$$

 $\mathbf{E}(t,\mathbf{r}) = \mathbf{E}(\Theta(t,\mathbf{r}))$ and $\mathbf{B}(t,\mathbf{r}) = \mathbf{B}(\Theta(t,\mathbf{r}))$, respectively, with

$$\Theta(t,\mathbf{r}) := t - \mathbf{n} \cdot \frac{\mathbf{r}}{c}.\tag{12}$$

being the retarded time



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The presence of an Underdense Plasma can be taken into account via it's n_m index of refraction!

$$n_m = \sqrt{1 - \frac{\omega_p^2}{\omega_L^2}}$$
 and $\omega_p^2 = \frac{n_e e^2}{\varepsilon_0 m_e}$ (13)

$$\Theta(t, \mathbf{r}, n_m) := t - n_m \mathbf{n} \cdot \frac{\mathbf{r}}{c}. \tag{14}$$

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The retarded time, including the index of refraction:

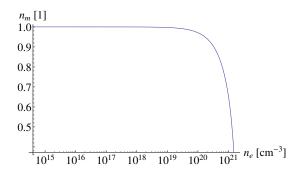
$$\Theta(t,\mathbf{r},n_m):=t-n_m\mathbf{n}\cdot\frac{\mathbf{r}}{c}.$$
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The relevant plasma densities are far below the critical density. At $\lambda = 800 \, \text{nm}, \, n_c = 1.74196 \cdot 10^{21} \, \text{cm}^{-3}.$



 $n_m(10^{15}\,\mathrm{cm}^3) \approx n_m(0) \Rightarrow \Theta(t,\mathbf{r},n_m) \approx \Theta(t,\mathbf{r})$

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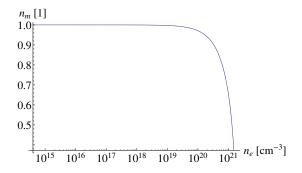


FIGURE: The index of refraction as a function of plasma electron density.

 $n_m(10^{15}\,\mathrm{cm}^3) \approx n_m(0) \Rightarrow \Theta(t,\mathbf{r},n_m) \approx \Theta(t,\mathbf{r})$ ACCELERATION IN UNDERDENSE PLASMAS CAN BE WELL APPROXIMATED BY ACCELERATION IN VACUUM!

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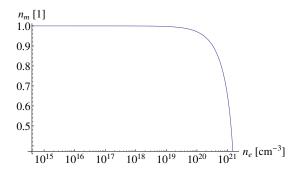


FIGURE: The index of refraction as a function of plasma electron density.

$$n_m(10^{15}\,\mathrm{cm}^3) \approx n_m(0) \Rightarrow \Theta(t,\mathbf{r},n_m) \approx \Theta(t,\mathbf{r})$$

ACCELERATION IN UNDERDENSE PLASMAS CAN BE WELL



The relevant plasma densities are far below the critical density. At $\lambda = 800 \, \mathrm{nm}$, $n_c = 1.74196 \cdot 10^{21} \, \mathrm{cm}^{-3}$.

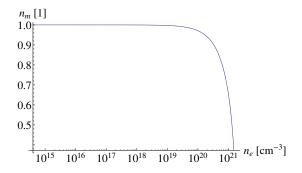


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ACCELERATION IN UNDERDENSE PLASMAS CAN BE WELL APPROXIMATED BY ACCELERATION IN VACUUM!

- Only chirped pulses provide non-negligible acceleration.
- The electron has to be on-axis and propagate parallel with the pulse.
- Larger beam waists provide more energy gain.
- Shorter pulses provide more energy gain.
- With a Gaussian pulse of $\lambda=800\,\mathrm{nm}$ wavelength, $T=30\,\mathrm{fs}$ pulse duration, $I=10^{21}\,\mathrm{W\cdot cm^{-2}}$ intensity and $W_0=100\lambda$ beam waist, an energy gain of 270 MeV pro pulse can be achieved.

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	Our Results	Kneip <i>et. al</i> Phys. Rev. Lett. 103 (2009), 035002
Wavelength	800 nm	800 nm
Pulse Duration	30 fs	55 fs
Intensity	10 ²¹ W ⋅ cm ⁻²	10 ¹⁹ W ⋅ cm ⁻²
Beam Waist	100λ	10 mm
Total Pulse Energy	9.6 J	10 J
Average Power	320 TW	180 TW
Energy gain	275 MeV (on 5 mm)	420 MeV (on 5 mm)
		800 MeV (on 10 mm)
Accelerating Gradient	58 GVm ⁻¹	80 GVm ⁻¹

OUR RESULTS AGREE WITHIN A FACTOR OF TWO WITH THE **EXPERIMENTAL DATA!**



Summarizing the results (see: doi: 10.1016/j.nimb.2015.10.013; (M.A. Pocsai, S. Varró, I.F. Barna, article in press))

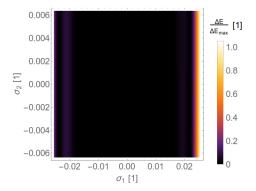


FIGURE: The relative energy gain as a function of the chirp parameters. ΔE depends very weakly on σ_2 . $\lambda = 800 \, \mathrm{nm}$, $T = 5 \, \mathrm{fs}$, $I = 8.544 \cdot 10^{18} \, \mathrm{W} \cdot \mathrm{cm}^2$, $W_0 = 0.8 \, \mu \mathrm{m}$, A = 0.1, $\varphi_1 = 2.058$.



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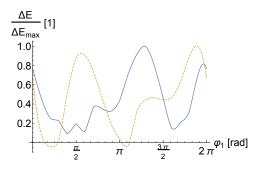


FIGURE : The relative energy gain as a function of the carrier–envelope phase. $\lambda = 800 \, \mathrm{nm}, \ T = 5 \, \mathrm{fs}, \ \sigma_1 = 0.0471 \, \mathrm{fs}^{-2}, \ \sigma_2 = 0, \ A = 0.24.$ $I = 10^{20} \, \mathrm{W} \cdot \mathrm{cm}^{-2}, \ W_0 = 8.976 \, \mu \mathrm{m}$ (solid line), $a_0 = 3 \cdot 10^{20} \, \mathrm{W} \cdot \mathrm{cm}^{-2}, \ W_0 = 8 \, \mu \mathrm{m}$ (dashed line).

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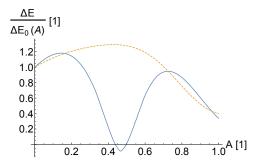


FIGURE : The relative energy gain as a function of the relative amplitude of the harmonics. The second harmonic enhanced the energy gain by about 20 - 30 %, compared to the monochromatic case. $\lambda=800$ nm, T=5 fs, $\sigma_1=0.0471$ fs $^{-2},\,\sigma_2=0.$ $I=10^{20}$ W \cdot cm $^{-2},\,W_0=8$ μ m, $\varphi_1=0$ (solid line), $a_0=3\cdot10^{20}$ W \cdot cm $^{-2},\,W_0=8.976$ μ m, $\varphi_1=0$ (dashed line).

OUTLINE

- Introduction
- 2 THEORETICAL BASICS
 - Gaussian Pulses
 - Generalization for Bichromatic Fields
 - Equations of Motion
 - The Presence of an Underdense Plasma
- 3 RESULTS
 - General Remarks
 - Monochromatic Fields
 - Comparison with Experimental Data
 - Bichromatic Fields
- SUMMARY



- A simple but (computationally) efficient model has been presented.
- Chirped Gaussian pulses can transfer up to 270 MeV energy to a single electron.
- The addition of the second harmonic boosts the energy transfer to the electron by even 30 %—it is tempting to use a bichromatic driver pulse for electron acceleration.
- The results obtained with our simple model agree quite well with the experimental data.

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THANK YOU FOR YOUR ATTENTION!