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The latest results of the soft+hard model in heavy ion collisions

Motivation

- Non-extensive statistical approarch
- Why to use Tsallis Pareto-like distributions?

• Fit to experimental data

- Can Tsallis Pareto fit spectra of HIC?
- The soft+hard model and its applications
- Spectra fit and extraction of q and T

Connecting spectra and v₂

Summary





Motivation



- ALICE: A Large Ion Colliding Experiment
- Analysing QGP: only through observable quantities
- Hadron spectra: a break around 4-6 GeV/c

Motivation

Hadron spectra in *pp* collisions can be described by the Tsallis distribution

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 π spectra in pp collisions depends similarly on \sqrt{s} and on the multiplicity N



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• Extensive Boltzmann – Gibbs statistics $S_{12} = S_1 + S_2$ $E_{12} = E_1 + E_2$ \longrightarrow $S_B = -\sum_i p_i \ln p_i$

Non-extensivity
 → generalized entropy

$$\hat{L}_{12} = \hat{L}_1(S_1) + \hat{L}_2(S_2)$$

$$L_{12} = L_1(E_1) + L_2(E_2)$$

$$S_T = \frac{1}{1-q} \sum_i (p_i^q - p_i)$$

$$S_{12} = S_1 + S_2 + (q-1)S_1S_2 \implies \hat{L}(S) = \frac{1}{q-1}\ln\left(1 + (q-1)S\right)$$

Tsallis entropy from here: Tsallis – Pareto distribution

$$f(\varepsilon) = \left[1 + (q-1)\frac{\varepsilon}{T}\right]^{-\frac{1}{q-1}}$$

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Eur. Phys. J. A49 (2013) 110 Physica A 392 (2013) 3132

• For more details see KeMings lecture!

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Tsallis – Pareto distribution:



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Motivation Why use Tsallis – Pareto distribution? Tsallis – Pareto: whole mumentum range

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Power-law: Boltzmann – Gibbs: Sectior high momenta low momenta Incl. Cross It was very good for **pp**, but in **AA** and **pA** collision there is soft a significant break! hard 4-6 GeV/c

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Soft + Hard

Simplest approximation: soft ('bulk') + hard ('jet') contribution

$$p^{0}\frac{dN}{d^{3}p} = p^{0}\frac{dN}{d^{3}p}^{hard} + p^{0}\frac{dN}{d^{3}p}^{soft}$$

Identified hadron spectra is given by double Tsallis–Pareto:

$$\frac{dN}{2\pi p_T dp_T dy}\Big|_{y=0} = f_{hard} + f_{soft} \qquad f_i = A_i \left[1 + \frac{(q_i - 1)}{T_i} \left[\gamma_i (m_T - v_i p_T) - m \right] \right]^{-1/(q_i - 1)}$$

in where parameters are given by

• Lorentz factor $\gamma_i = 1/\sqrt{1-v_i^2}$

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- Transverse mass $m_T = \sqrt{p_T^2 + m^2}$
- Doppler temperature $T_i^{Dopp} = T_i \sqrt{\frac{1+v_i}{1-v_i}}$

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Soft + Hard

1.Two single Tsallis fit:

1) 1st Tsallis fit: the range $i[p_0 - \varepsilon; p_{max}]$,

the fitted parameters are:

 $\boldsymbol{q}_{hard'} \; \boldsymbol{T}_{hard'} \; \boldsymbol{A}_{hard'}(\boldsymbol{v}_{hard})$

2) **2nd Tsallis fit**: the range is $[p_{min}; p_0 + \varepsilon]$

the fitted parameters are:

$$\boldsymbol{q}_{soft}, \boldsymbol{T}_{soft}, \boldsymbol{A}_{soft}, (\boldsymbol{v}_{soft})$$

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- **2.Double Tsallis** fit, the range is $[p_0 \varepsilon; p_{max}]$, the fitted parameters are the **hard** parameters
- **3.Double Tsallis** fit, the range is $[p_{min}; p_0 + \varepsilon]$, the fitted parameters are the **soft** parameters

4.Double Tsallis fit, the range is $[p_{min}; p_{max}]$, all of the parameters are fitted

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Results

Results

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Results – Id. particles



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Results



$$f_i = A_i \left[1 + \frac{(\mathbf{q}_i - 1)}{T_i} \left[\gamma_i (m_T - v_i p_T) - m \right] \right]^{-1/(\mathbf{q}_i - 1)}$$

Reminder:
$$q = 1 + \frac{\Delta T^2}{T^2} - \frac{1}{C}$$

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Results - charged



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Results - Pion



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Results - Kaon



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Results - (Anti)Proton



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Results - Lambda



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Results



$$f_i = A_i \left[1 + \frac{(q_i - 1)}{T_i} \left[\gamma_i (m_T - v_i p_T) - m \right] \right]^{-1/(q_i - 1)}$$

Reminder:

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$$\frac{E}{\langle n \rangle} = \frac{\int \varepsilon f_{TS}(\varepsilon)}{\int f_{TS}(\varepsilon)} = \frac{DT}{1 - (q-1)(D+1)}$$

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Results - charged



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Results - Pion



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Results - Kaon

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Results - (Anti)Proton



Results - Lambda



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Results





Results – Summary

Clear separation of the soft and hard components

Strong centrality dependence

 ${}_{\bullet}$ There is a significant ${\rm N}_{\rm \tiny part}$ and \sqrt{s} ${}_{\rm dependence}$ too



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Connecting spectra and v₂



Spectra originating from hadronic sources

$$p^{0} \left. \frac{dN}{d^{3}p} \right|_{y=0} = \int_{-\infty}^{+\infty} d\zeta \int_{0}^{2\pi} d\alpha f[u_{\mu}p^{\mu}] \longrightarrow \frac{dN}{2\pi p_{T} dp_{T} dy} \bigg|_{y=0} = \int_{0}^{2\pi} \frac{d\varphi}{2\pi} p^{0} \left. \frac{dN}{d^{3}p} \right|_{y=0}$$

where we used parameters and assumptions:

- Hadron momenum: $p^{\mu} = (m_T \cosh y, m_T \sinh y, p_T \cos \varphi, p_T \sin \varphi)$
- Cylindric symmetry: $u^{\mu} = (\gamma \cosh \zeta, \gamma \sinh \zeta, \gamma v \cos \alpha, \gamma v \sin \alpha)$

> where
$$\zeta = rac{1}{2} \ln \left[(t+z)/(t-z)
ight]$$
 and $\gamma = 1/\sqrt{1-v^2}$

- > Co-moving energy: $u_{\mu}p^{\mu}|_{y=0} = \gamma \left[m_T \cosh \zeta v p_T \cos (\varphi \alpha)\right]$
- Transverse flow:

Taylor expansion:

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$$v(\alpha) = v_0 + \sum_{m=1}^{\infty} \delta v_m \cos(m\alpha) \equiv v_0 + \delta v(\alpha)$$
$$f[u_\mu p^\mu]|_{y=0} = \sum_{m=0}^{\infty} \frac{[\delta v(\alpha)]^m}{m!} \frac{\partial^m}{\partial v_0^m} f[u_\mu p^\mu]|_{y=0}^{v(\alpha)=v_0}$$

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Spectra originating from hadronic sources

$$\frac{dN}{2\pi p_T dp_T dy}\Big|_{y=0} = \int_0^{2\pi} \frac{d\varphi}{2\pi} p^0 \left. \frac{dN}{d^3 p} \right|_{y=0} = \sum_{m=0}^{\infty} \frac{a_m}{m!} \frac{\partial^m}{\partial v_0^m} f[E(v_0)] \approx f[E(v_0)] + O(\delta v^2)$$

where $E(v_0) = \gamma_0 (m_T - v_0 p_T)$ and $a_m = \int_0^{2\pi} d\alpha [\delta v(\alpha)]^m$

Azimuthal anisotropy:



• Spectra originating from hadronic sources

$$\frac{dN}{2\pi p_T dp_T dy}\Big|_{y=0} = \int_0^{2\pi} \frac{d\varphi}{2\pi} p^0 \left. \frac{dN}{d^3 p} \right|_{y=0} = \sum_{m=0}^{\infty} \frac{a_m}{m!} \frac{\partial^m}{\partial v_0^m} f[E(v_0)] \approx f[E(v_0)] + O(\delta v^2)$$

where $E(v_0) = \gamma_0 (m_T - v_0 p_T)$ and $a_m = \int_0^{2\pi} d\alpha [\delta v(\alpha)]^m$

• Azimuthal anisotropy:

$$v_n = \frac{\int_{0}^{2\pi} d\varphi \cos(n\varphi) p^0 \left. \frac{dN}{d^3 p} \right|_{y=0}}{\int_{0}^{2\pi} d\varphi p^0 \left. \frac{dN}{d^3 p} \right|_{y=0}} \approx \frac{\delta v_n \gamma_0^3}{2} \frac{(v_0 m_T - p_T) f'[E(v_0)]}{f[E(v_0)]} + O(\delta v^2)$$

• Using the soft+hard model:

$$v_2 = \frac{w_{hard}f_{hard} + w_{soft}f_{soft}}{f_{hard} + f_{soft}} \quad \text{with the coefficient } w_i = \frac{\delta v_i \gamma_i^3}{2T_i} \frac{p_T - v_i m_T}{1 + \frac{q_i - 1}{T_i} [\gamma_i (m_T - v_i p_T) - m]}$$



• Using the soft+hard model:

$$v_2 = rac{w_{hard}f_{hard} + w_{soft}f_{soft}}{f_{hard} + f_{soft}}$$
 with the coefficient $w_i = rac{\delta v_i \gamma_i^3}{2T_i} rac{p_T - v_i m_T}{1 + rac{q_i - 1}{T_i} [\gamma_i (m_T - v_i p_T) - m]}$



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Summary

Non-extensive statistical approach HIC $q = 1 - \frac{1}{C} + \frac{\Delta T^2}{T^2}$

- Providing phyiscal meaning of q:
- Boltzmann Gibbs limit: $C \to \infty$, $\frac{\Delta T^2}{T^2} \to 0$ $(q \to 1)$
- Tsallis Pareto fits on spectra in e⁺e⁻, pp
- Doesn't work for larger system, like pA, AA

Application of 'soft+hard' model in AA

- Double Tsallis Pareto measures non-extensitivity
- SOFT: $q \rightarrow 1$, suggest Boltzmann Gibbs (QGP)
- HARD: q > 1.1, Tsallis Pareto like
- Asimuthal anisotropy can be obtained too

Future plans

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Investigation of multiplicity-, centrality- and energy dependency







Thank you for your attention!

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Backup







N_{part} scaling



Results – q trends



Results – T trend

Results - charged

Results - Pion

Results - Kaon

Results - (Anti)Proton

Results - Lambda

